Cross-correlating the GW astrophysical background with galaxy clustering

Guadalupe Cañas-Herrera, Omar Contigiani, Valeri Vardanyan

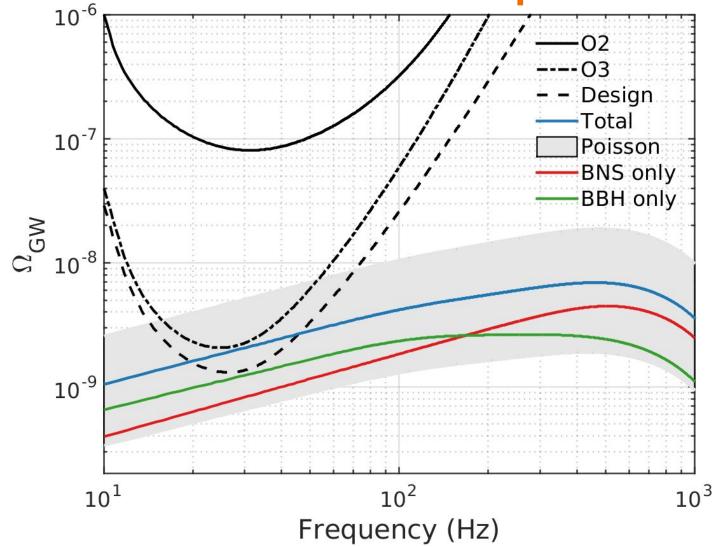
Gravitational Wave Probes of Fundamental Physics





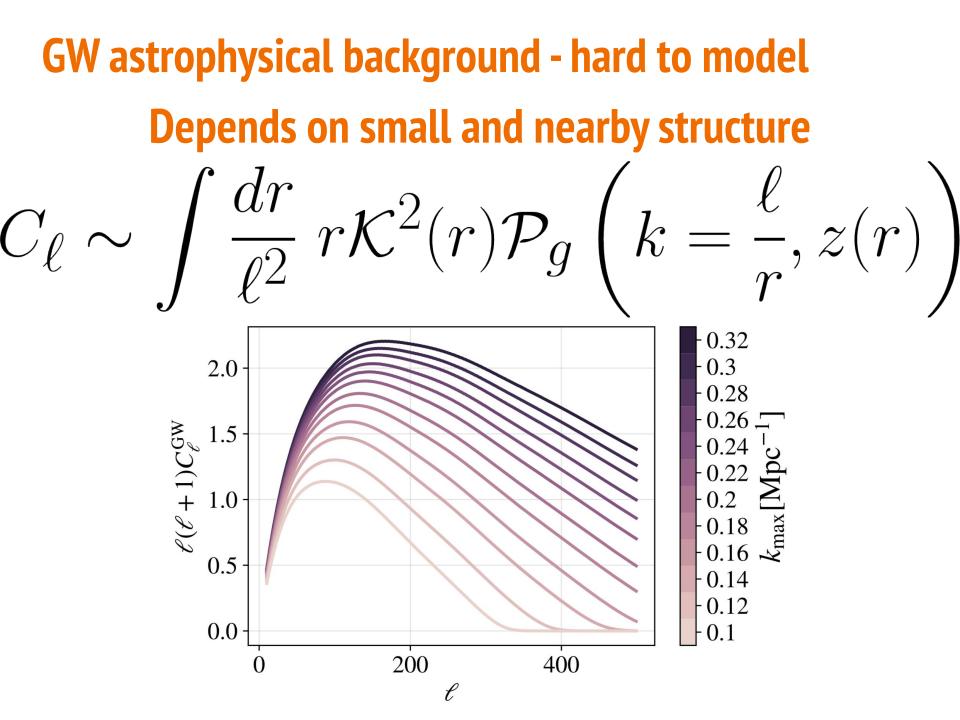
$$\Omega_{\rm GW}(\nu_0, \hat{\mathbf{r}}) = \frac{\nu_0}{\rho_{\rm c}} \frac{d\rho_{\rm GW}(\nu_0, \hat{\mathbf{r}})}{d\nu_0 d^2 \hat{\mathbf{r}}},$$

GW astrophysical background - hard to detect LIGO & the monopole



Abbott+ 2018

GW astrophysical background - hard to model **Depends on small and nearby structure** $C_{\ell} \sim \int \frac{dr}{\ell^2} r \mathcal{K}^2(r) \mathcal{P}_g\left(k = \frac{\ell}{r}, z(r)\right)$

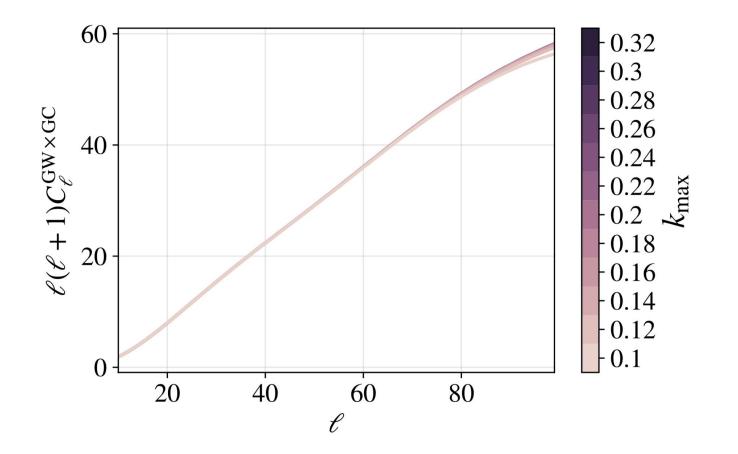


Cross-correlating <u>GW background</u> with <u>galaxy</u> <u>catalogs</u>

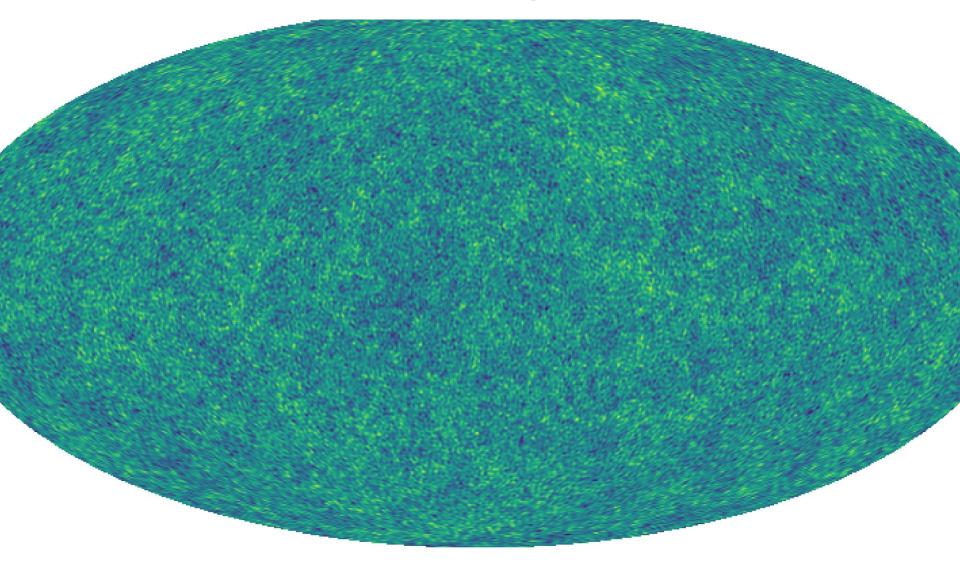
$$\Omega_{\rm GW}(\hat{\mathbf{r}}) \equiv \int dr \; r^2 \mathcal{K}(r) n(\vec{\mathbf{r}}),$$

GW-GC cross-correlation - no noise, easy to model

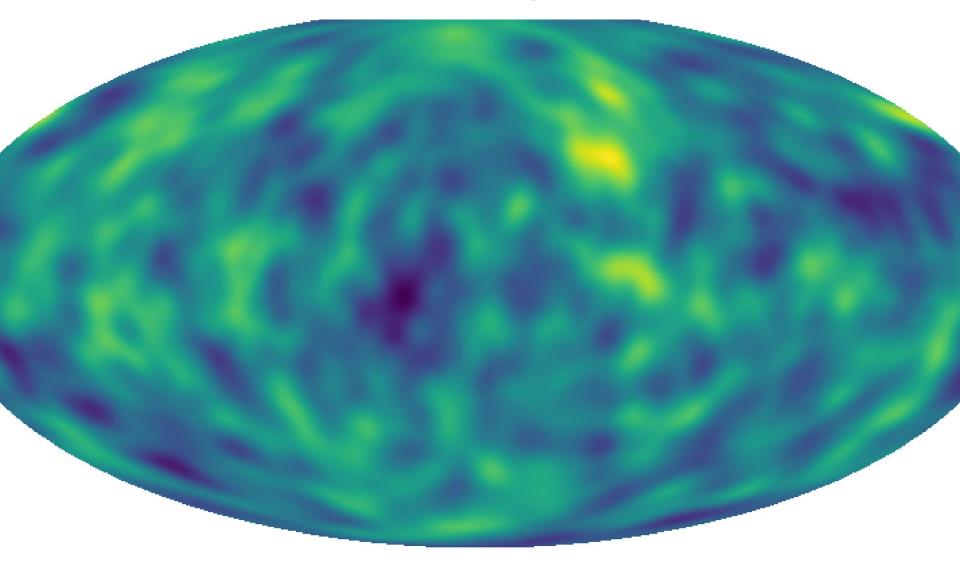
$$C_{\ell} \sim \int \frac{dr}{\ell^2} r \mathcal{K}(r) \mathcal{W}(r) \mathcal{P}_g\left(k = \frac{\ell}{r}, z(r)\right)$$



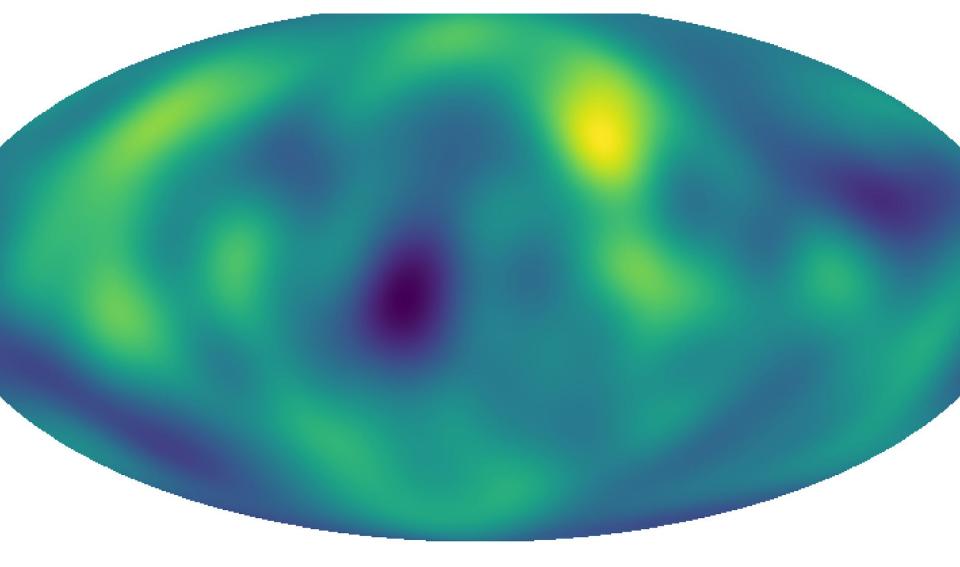
GW-GC cross-correlation - angular resolution



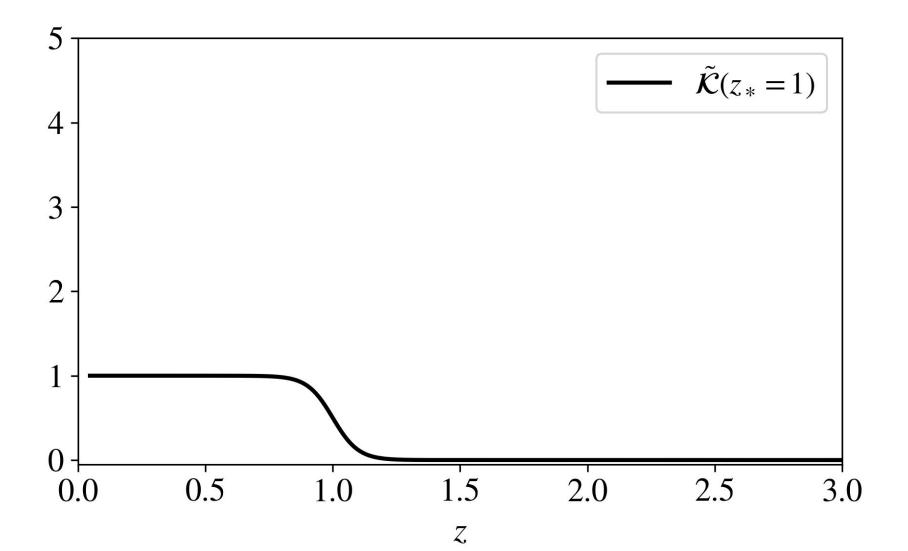
GW-GC cross-correlation - angular resolution



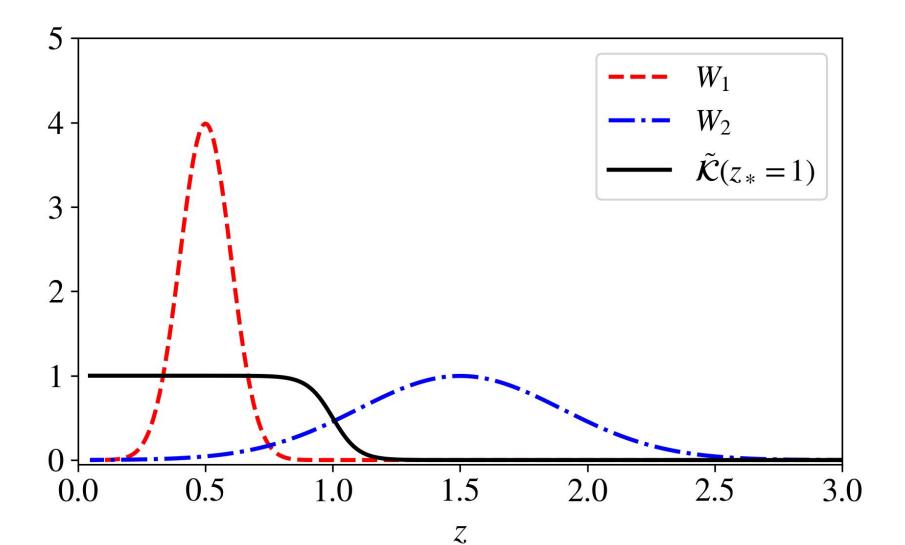
GW-GC cross-correlation - angular resolution

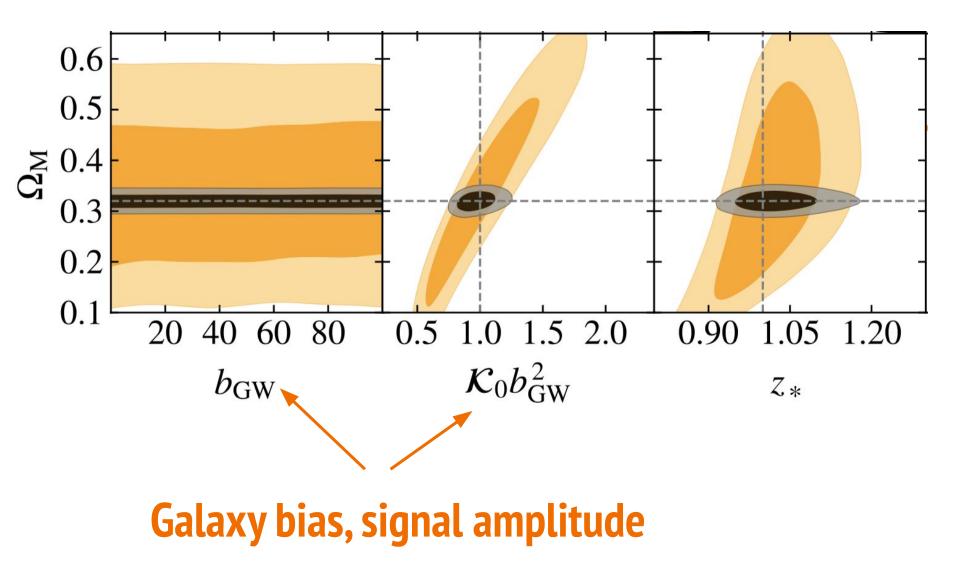


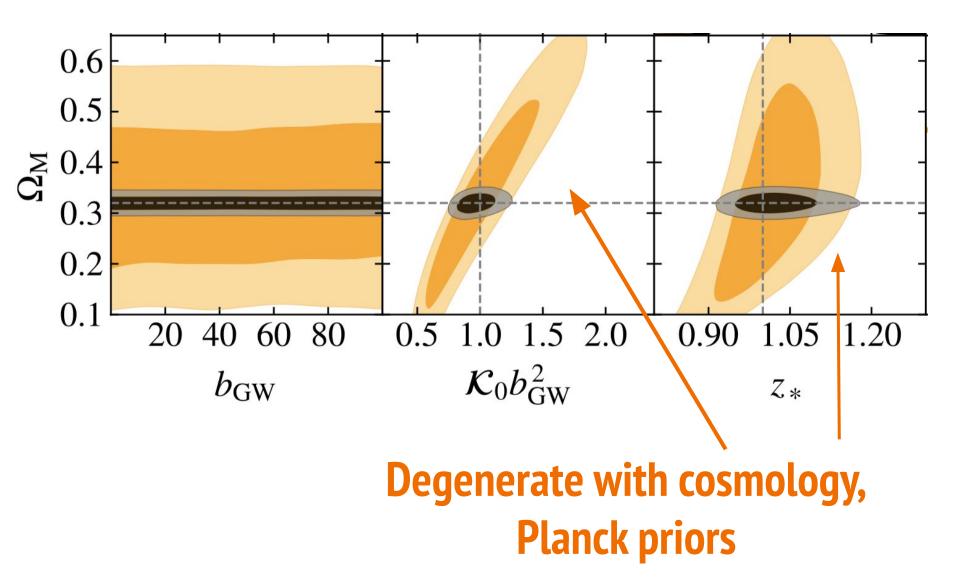
GW-GC cross-correlation - toy model setup

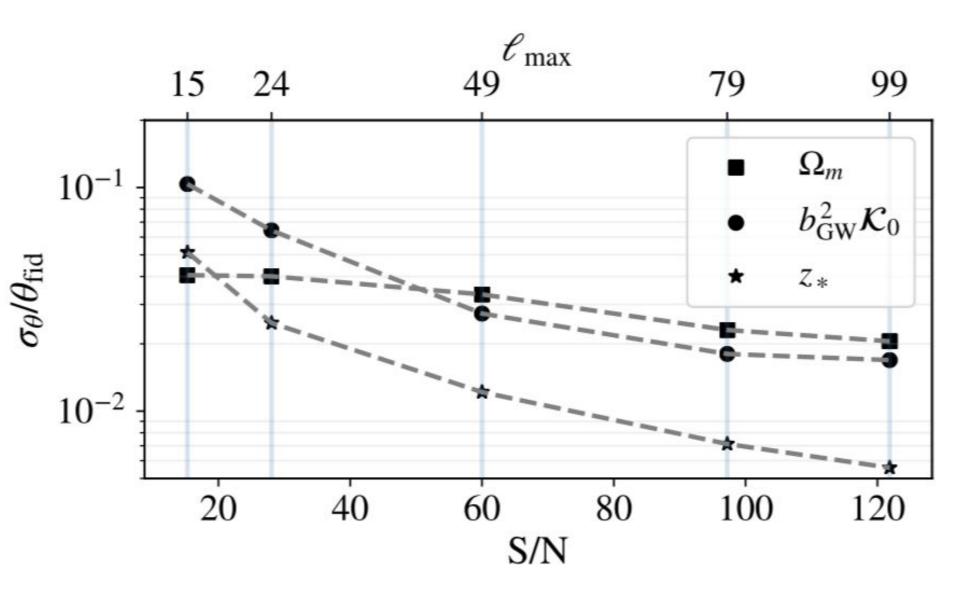


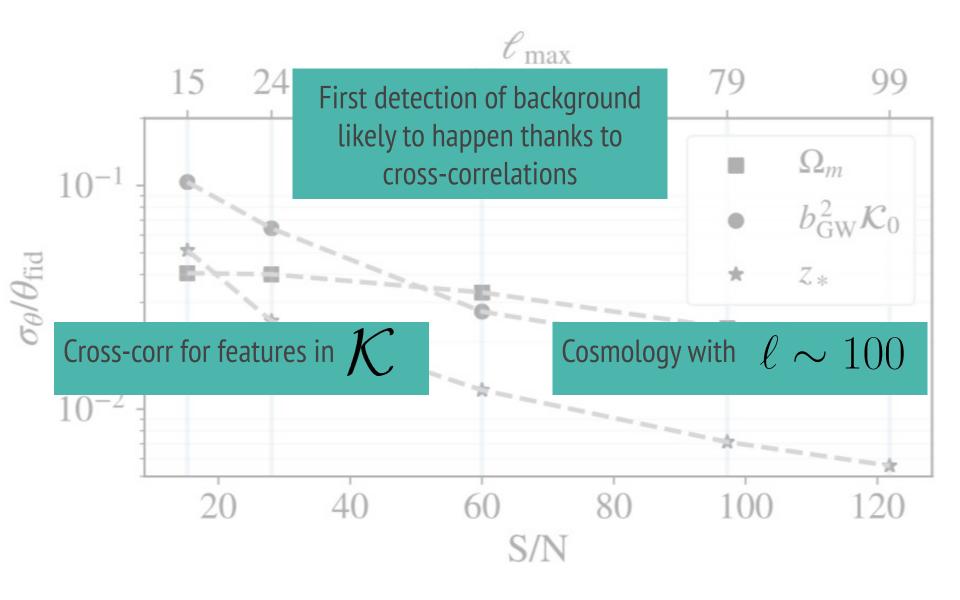
GW-GC cross-correlation - toy model setup









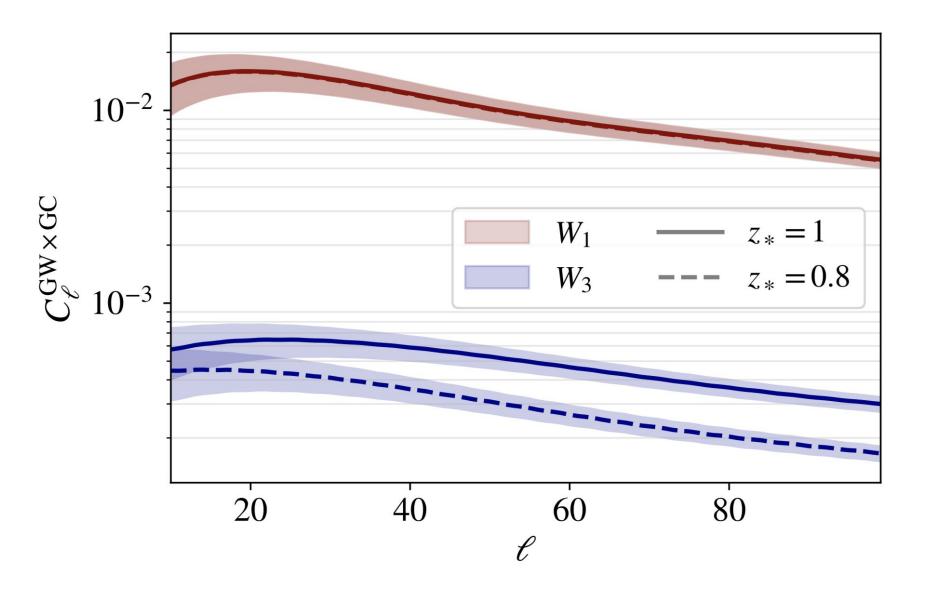


GW astrophysical background - hard to detect

shot-noise

Original calculation in Jenkins+ 2019

GW-GC cross-correlation - playing with toy model



GW-GC cross-correlation - playing with toy model

