

Axion Dark Matter Experiment in Korea

정우현 (Woohyun Chung)

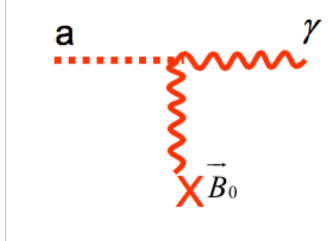
Center for Axion and Precision Physics Research (CAPP)

Institute for Basic Science (IBS)

OUTLINE

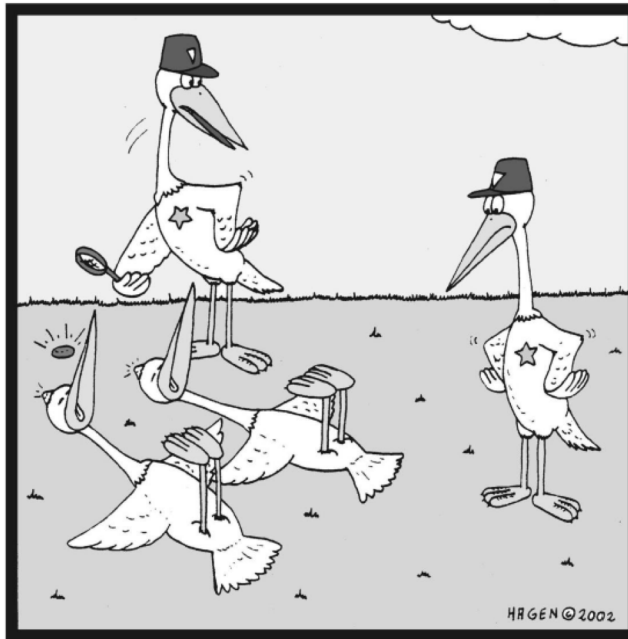
- **Introduction**
 - Dark Matter Axion
 - CULTASK (CAPP's Ultra Low Temperature Axion Search in Korea)
- **CAPP-PACE (Pilot Axion Cavity Exp.)**
 - First complete axion experiment in Korea
 - Physics data ($10 \times$ KSVZ and KSVZ runs) in 2018
- **Improvements**
 - High Field Magnets
 - Quantum Amplifiers
 - High Q-factor (superconducting) and dielectric cavity
- **Summary**

- Peccei and Quinn (1977) postulated an elegant solution by adding a new global symmetry to resolve the **Strong CP Problem** in Standard Model
- **Axion is an excellent (and attractive) dark matter candidate**
 - Pseudo Goldstone Boson
 - Small Mass ($1\mu\text{eV} < m_a < 10\text{meV}$)
 - **Extremely Weakly Interacting**
 - Local Halo Density of $0.45 \text{ GeV}/\text{cm}^3$
 - $\beta \sim 10^{-3} \rightarrow Q_a \sim 10^6$
- Detection scheme by P. Sikivie (PRL 51:1415 1983) : Haloscapy
 - Axions will convert to photons in a strong magnetic field



$$L_{a\gamma\gamma} = g_\gamma \frac{\alpha}{\pi} \frac{a}{f_a} \vec{E} \cdot \vec{B}$$

Killing Two Birds With One Stone



Unbelievable! It looks like they've
both been killed by the same stone...

Peccei-Quinn mechanism

- Solves strong CP problem
- Provides dark matter in the form of axions

Axion dark matter search

- The axion mass is unknown, like any number in a phone book. The way we look for it:



- Once it's discovered, anyone will be able to dial in... and talk to it.

- Cosmic Axion Search

- Haloscopes (Microwave Cavity)

- Dish Antenna

- Dielectric Haloscope

- LC Circuit

- NMR techniques

- Atomic Transitions

- Solar axion search

- Axion Helioscopes

- Bragg Diffraction Scattering

- Geomagnetic Conversion

- Laboratory Axion Search

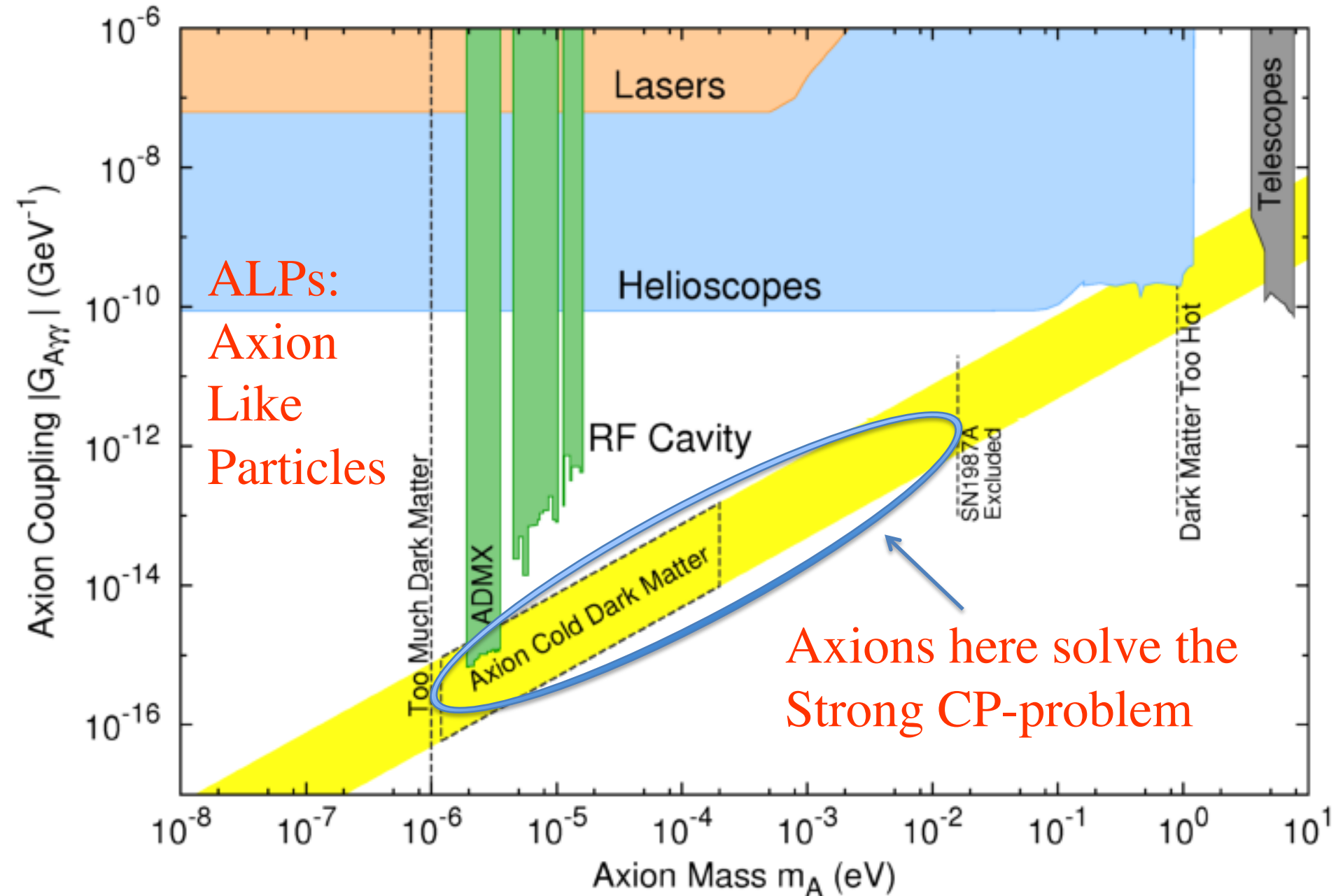
- Light Shining through Wall

- Polarization Experiment

- 5th Force

← Most sensitive
so far

Axion Landscape



Axion Detection Scheme (CULTASK)

P. Sikivie's Haloscope:

Axion Conversion Power ($\sim 10^{-24} \text{W}$):
$$P_{a \rightarrow \gamma\gamma} = g_{a\gamma\gamma}^2 \frac{\rho_a}{m_a} B^2 V C_{mnp} \min(Q_L, Q_a)$$

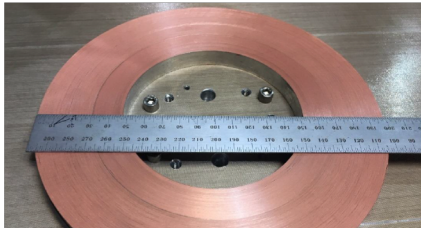
Signal to Noise Ratio:
$$SNR \equiv \frac{P_{\text{signal}}}{P_{\text{noise}}} = \frac{P_{a \rightarrow \gamma\gamma}}{k_B T_{\text{sys}}} \sqrt{\frac{t_{\text{int}}}{\Delta f_a}}$$

Scan rate:
$$\frac{df}{dt} \sim B^4 V^2 C^2 Q_L T_{\text{sys}}^{-2}$$

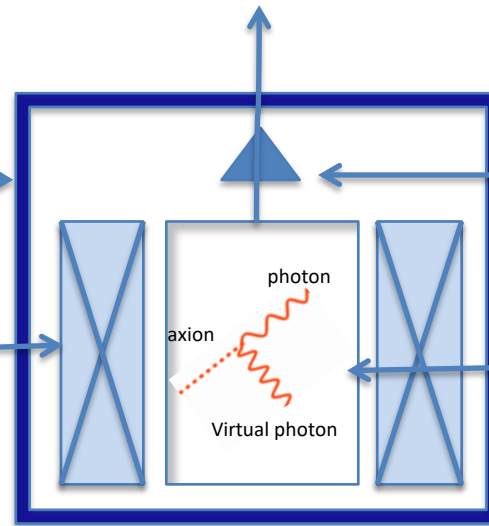
Cryogenics (<50mK)



High Field SC Magnet
25T BNL (HTS Technology) Design



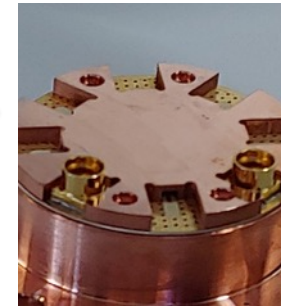
To RF Receiver



Quantum Amplifier

SQUID and/or JPA

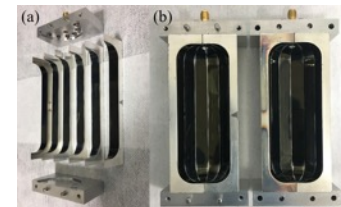
Dr. Matlashov and Uchaikin



High Q Tunable Cavity

Superconducting Coating

Prof. Dojun Youm

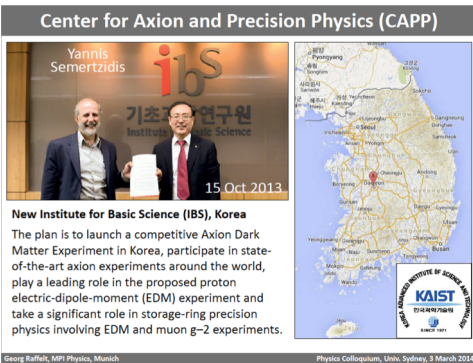


First 5 years...

Axion Laboratory with 7 Low Vibration Pads in KAIST Munji campus

Founded in 2013 (Oct.)

Center for Axion and Precision Physics (CAPP)



Yannis Semertzidis
15 Oct 2013

New Institute for Basic Science (IBS), Korea
The plan is to launch a competitive Axion Dark Matter Experiment in Korea, participate in state-of-the-art axion experiments around the world, play a leading role in the proposed proton electric-dipole-moment (EDM) experiment and take a significant role in storage-ring precision physics involving EDM and muon $g-2$ experiments.

Georg Raffelt, MPI Physics, Munich
Physics Colloquium, Uho, Sydney, 3 March 2016

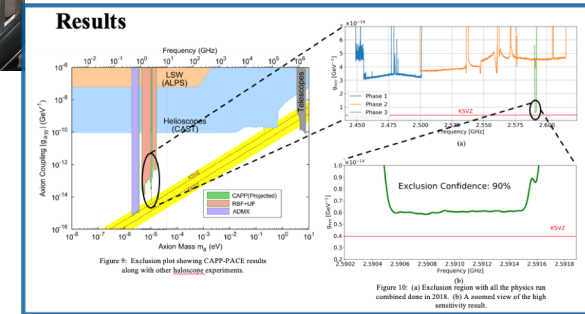


BF3 & BF4

BF5 & BF6

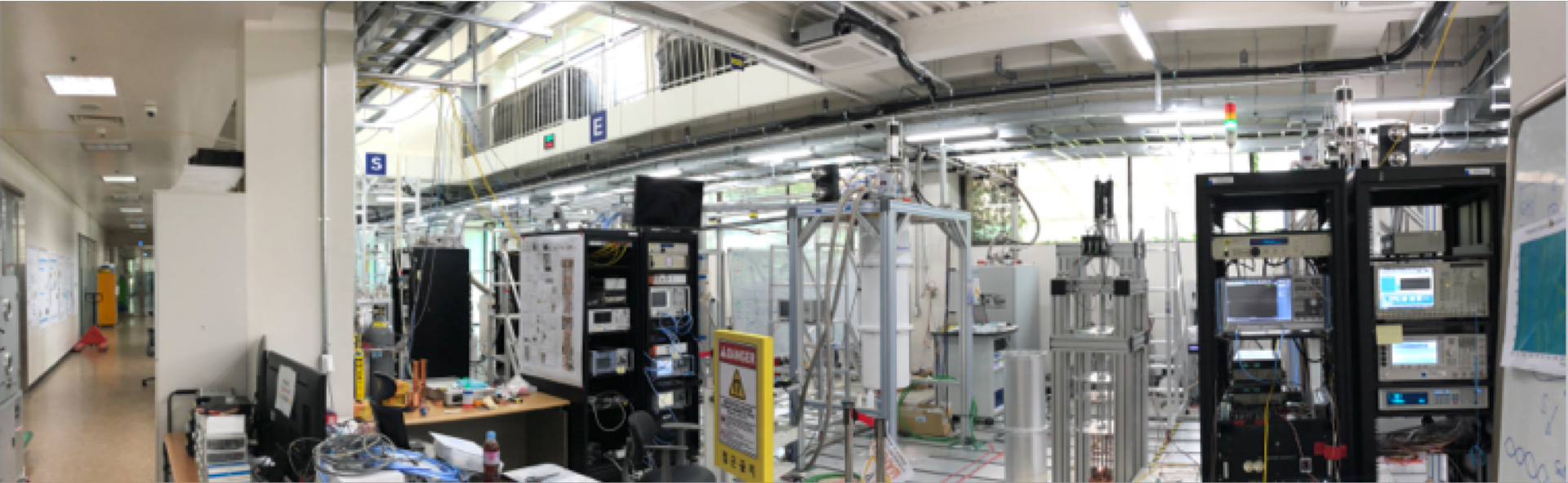
LVP

First Axion DM data!



2013	2014	2015	2016	2017	2018	2019
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CAPP Experimental Hall (LVP) in 2018



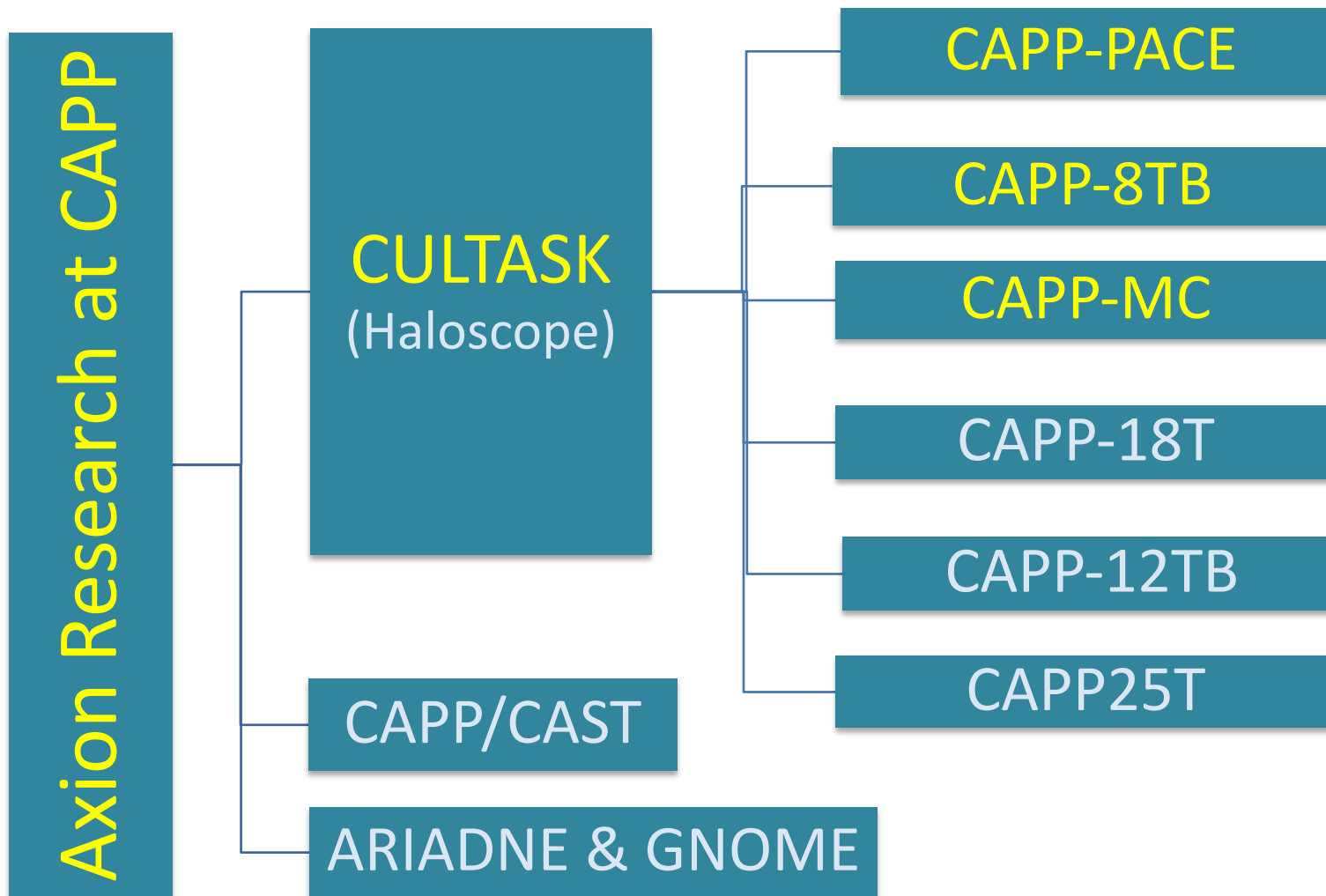
Sep 20th 2019

Long term strategy of IBS in Korea

Woohyun Chung

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CAPP's Axion Research



Refrigerators and Magnets

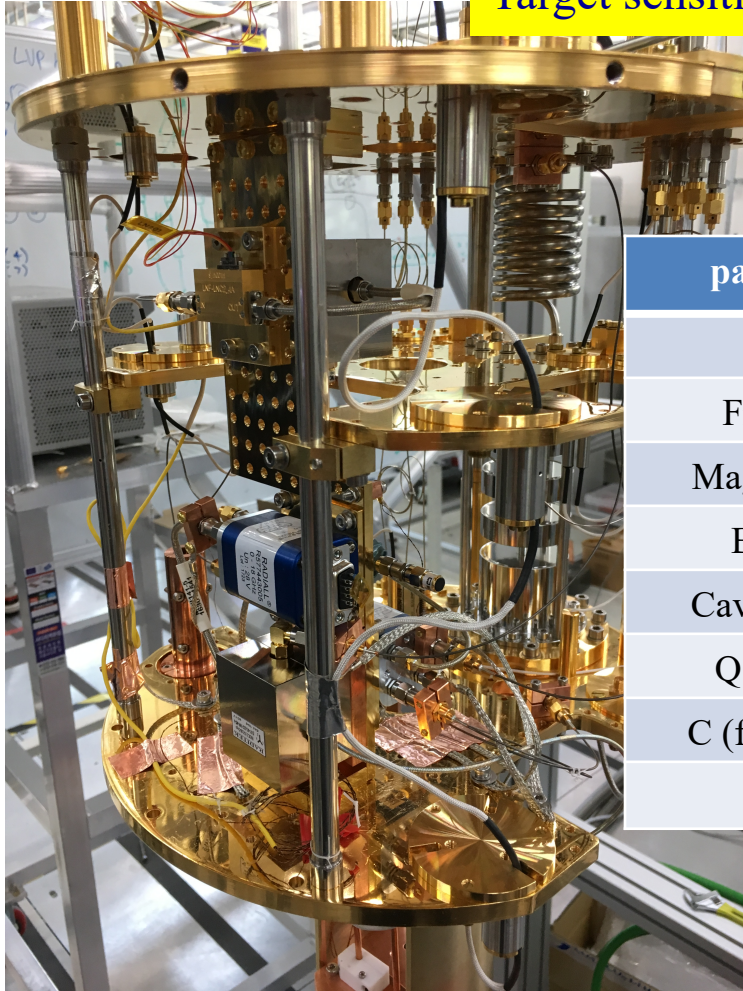
Refrigerators					Magnets					EXP
Vendor	Model	Base T (mK)	Cooling power	Install	B field	Bore (cm)	Material	Vendor	Delivery	
BlueFors (BF3)	LD400	10	18 μ W@20mK 580 μ W@100mK	2016	26T	3.5	HTS	SUNAM	2016	BF3 & BF4 for testing RF, QA and cavities
BlueFors (BF4)	LD400	10	18 μ W@20mK 580 μ W@100mK	2016	18T	7	HTS	SUNAM	2017	
Janis	HE3	300	25 μ W@300mK	2017	9T	12	NbTi	Cryo-Magnetics	2017	CAPP-MC
BlueFors (BF5)	LD400	10	18 μ W@20mK 580 μ W@100mK	2017	8T	12	NbTi	AMI	2016	CAPP-PACE
BlueFors (BF6)	LD400	10	18 μ W@20mK 580 μ W@100mK	2017	8T	16.5	NbTi	AMI	2017	CAPP-8TB
Oxford	Kelvinox	<30	400 μ W@120mK	2017	25T	10	HTS	BNL/CAPP	2020	Preparing for CAPP-12TB and CAPP-25T
Leiden	DRS1000	100	1.3mW @120mK	2019	12T	32	Nb ₃ Sn	Oxford	2020	

CAPP-PACE (Pilot Axion Cavity Experiment)

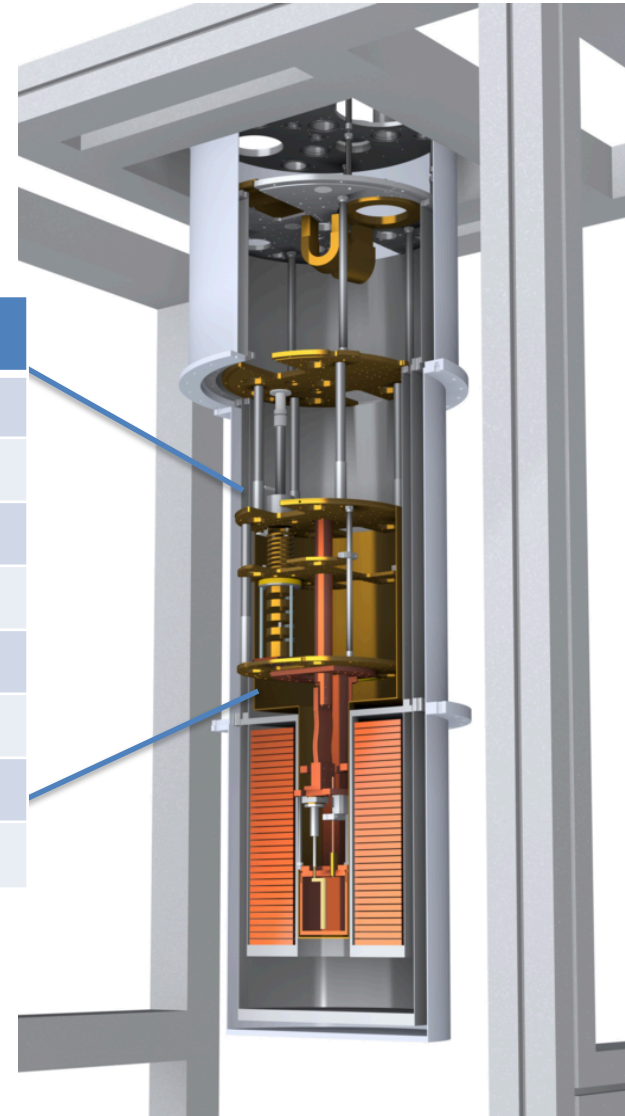
- **Originally**, R&D Project and testbed for
 - Cavity development
 - Frequency Tuning System (FTS)
 - RF receiver components (Optimization with cryo-RF)
 - DAQ and Controls
- **Has grown into** the first complete axion experiment in Korea
 - Achieved cavity physical temperature below 40 mK
 - Flawless operation of FTS w/ Piezo actuators (sapphire and Cu rod)
 - System noise temperature below 1.2 K
 - Complete DAQ and Controls including automatic Safety Warnings
- **Physics Data in 2018**
 - 10*KSVZ runs: 2.45 – 2.70 GHz scanned
 - KSVZ run: around 2.59 GHz, ~ 1 MHz scanned

CAPP-PAACE

Target sensitivity: **10*KSVZ, KSVZ**

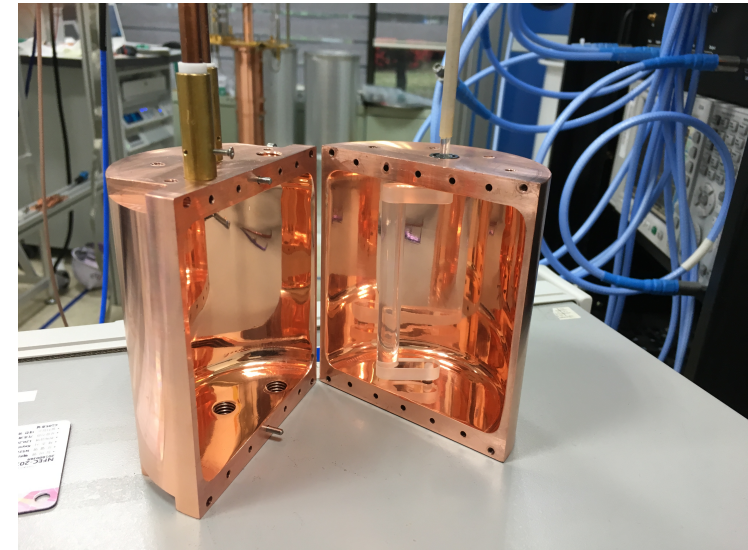


parameters	value
T_{cavity}	<40mK (WR)
Frequency	2.45~2.70GHz
Magnetic field	8 T
Bore size	11.8 cm
Cavity volume	1.12 L
Q unloaded	~ 100,000
C (form factor)	~ 0.6
T_{system}	< 1.5 K

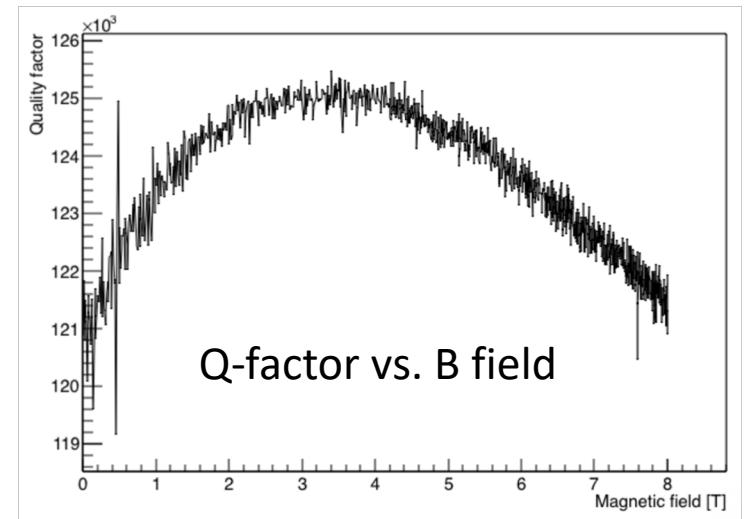


Cavity: OFHC Cu “split” type
Unloaded Q-factor of $\sim 100,000$

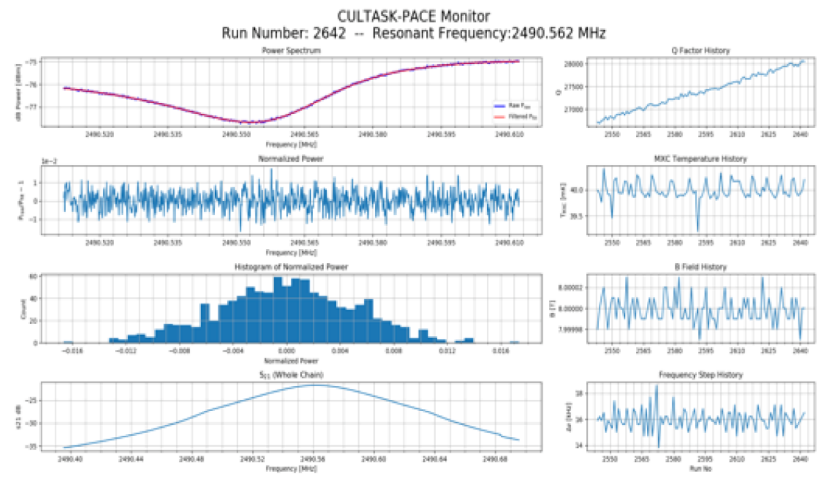
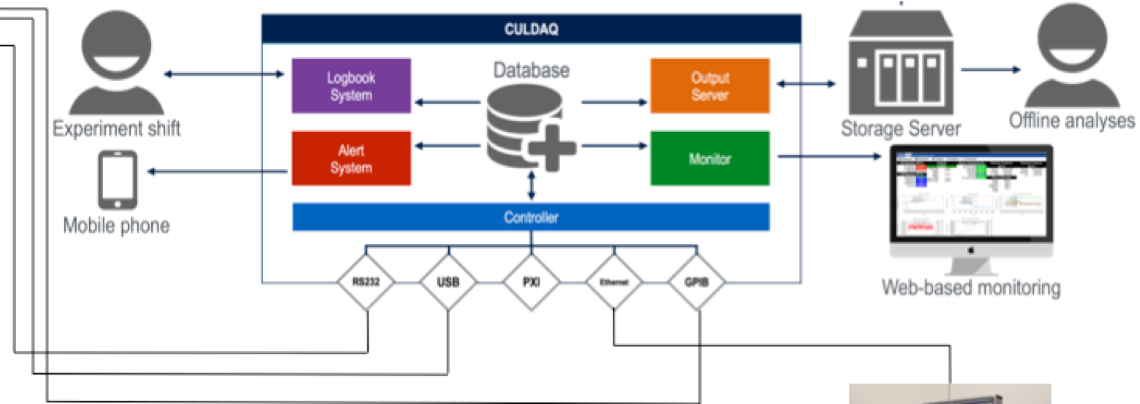
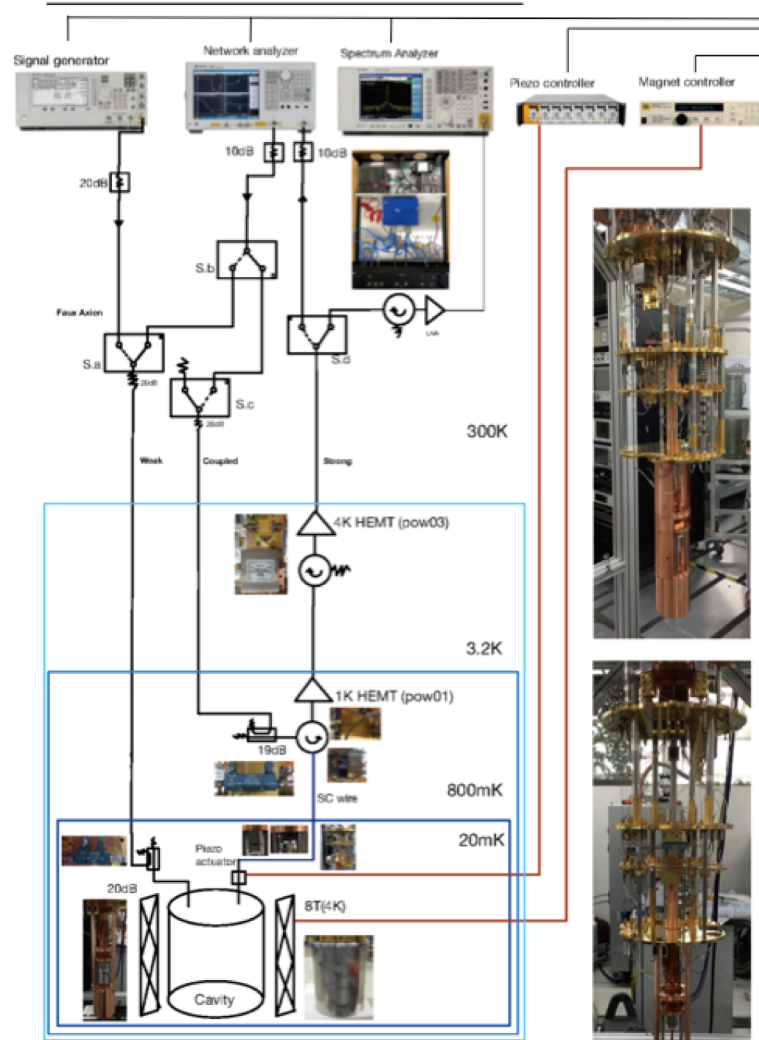
Tuning: Piezoelectric actuators (Attocube)
Thermal link to 1K plate
Sapphire rod to cavity by cryo bearing
Rotator resolution of $1/1000$ deg \rightarrow 16 kHz/step
Vibration free: w/ ball and spring



Linear and Rotational Piezo Actuators



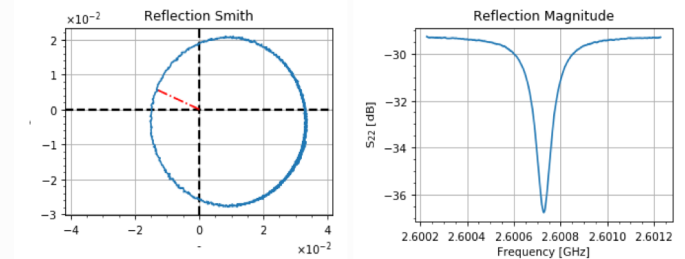
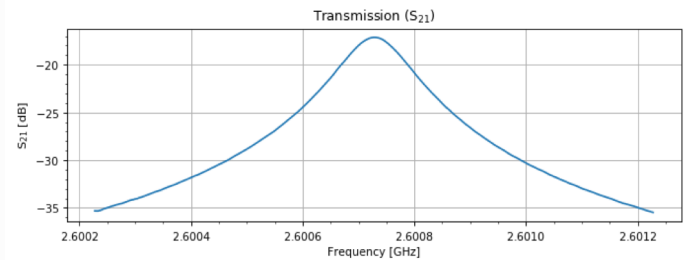
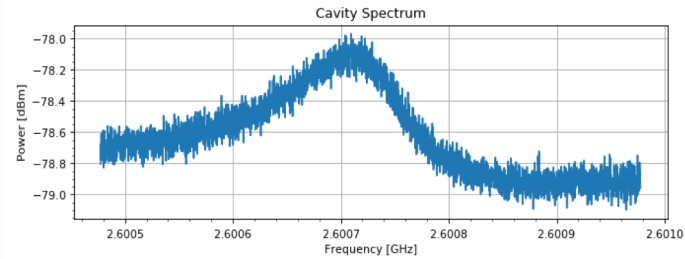
RF read-out chain & Controls



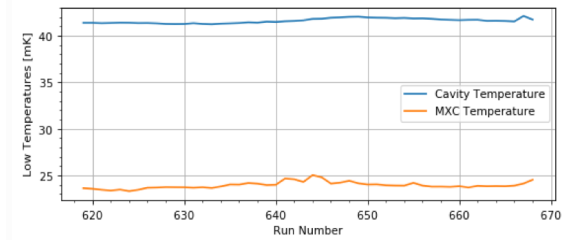
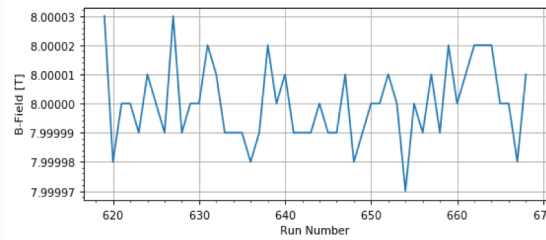
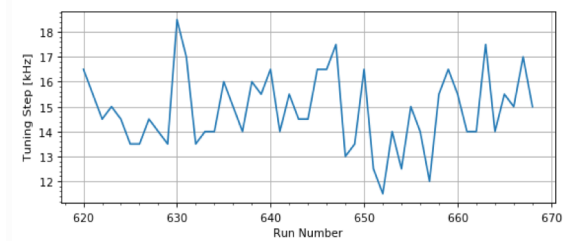
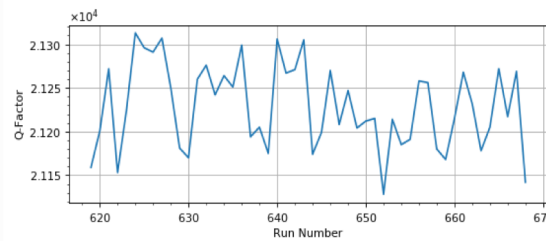
Exp#: 450 **Run#:** 668 **Res. Freq.:** 2600.7275 MHz **Recorded:** 18/08/21-18:47:32 **Minutes passed:** 2.95

Sensitivity: 10 KSVZ **SNR:** 5 **Target Range:** 2.5908 - 2.7 GHz

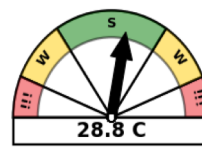
Current Run



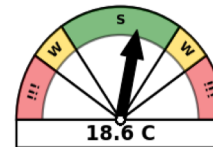
History



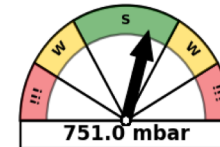
Oil Temperature



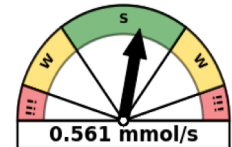
Water In Temperature



Pressure: 3



Flow Rate



CAPP-PACE (Online Monitor)



CAPP-PACE axion data

In 2018

	10*KSVZ (1)	10*KSVZ (2)	10*KSVZ (3)	KSVZ	10*KSVZ (4)
Date	1/19 – 2/13	7/23 – 8/01	8/14 – 8/23	9/01 – 10/26	11/1 – 11/24
Frequency [GHz]	2.450 – 2.500	2.500 – 2.548	2.547 – 2.613	2.5905 – 2.5915	2.613 – 2.710
Volume [liter]	0.59	0.59	1.12	1.12	1.12
T_{system} [K]	1.05	1.05	1.14	1.16	1.16
$\langle B_0 \rangle$ [T]	7.0	7.0	7.2	7.2	7.2
coupling	1.9	1.9	2.0	1.9	2.0
C (form factor)	.50	.50	0.55	0.66	0.55

Results (Mr. Caglar Kutlu's)

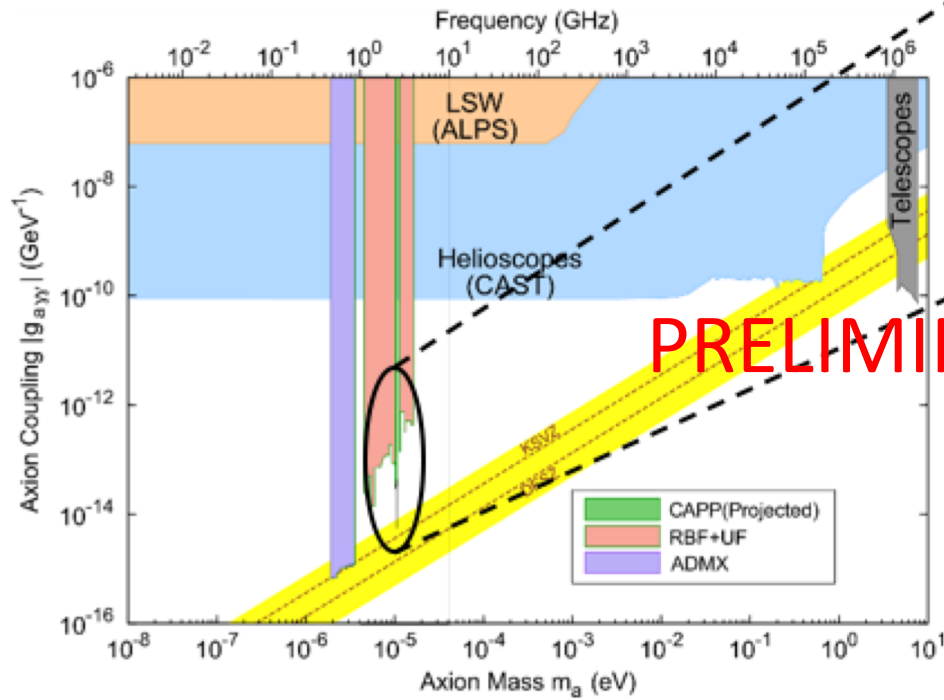
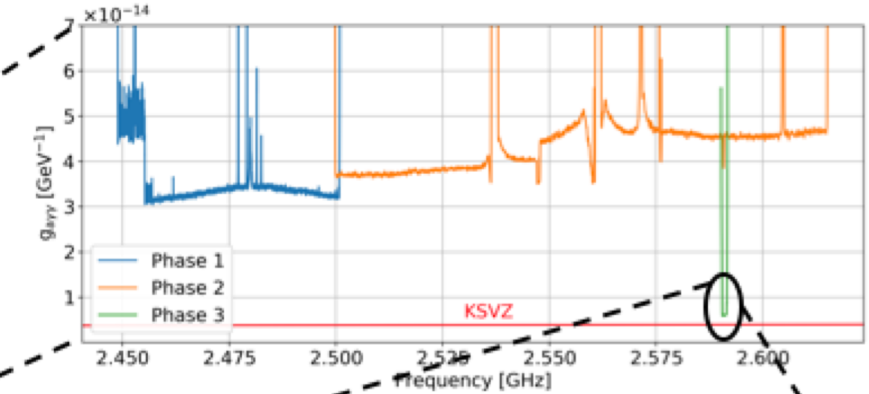
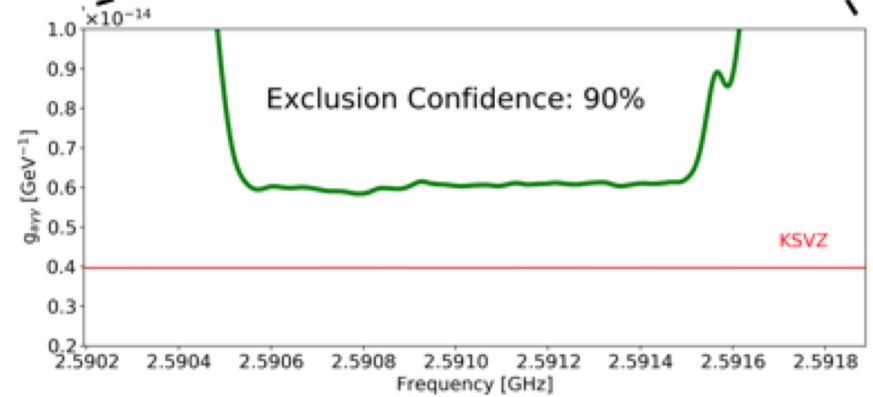


Figure 9: Exclusion plot showing CAPP-pace results along with other haloscope experiments.



(a)



(b)

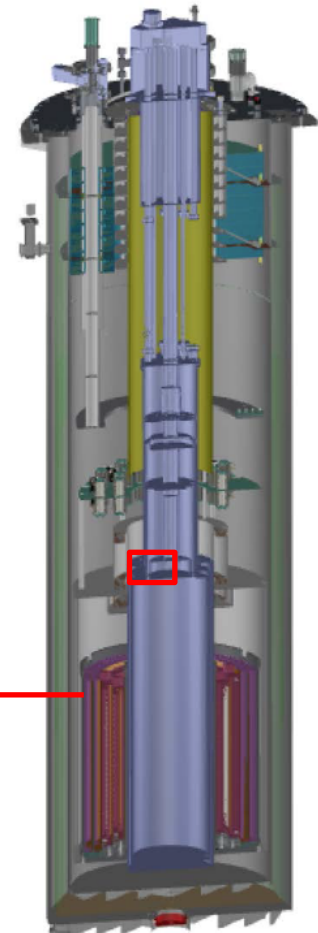
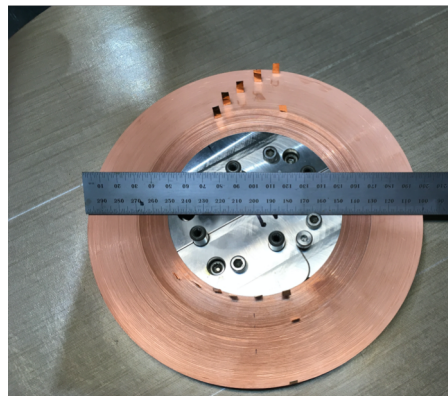
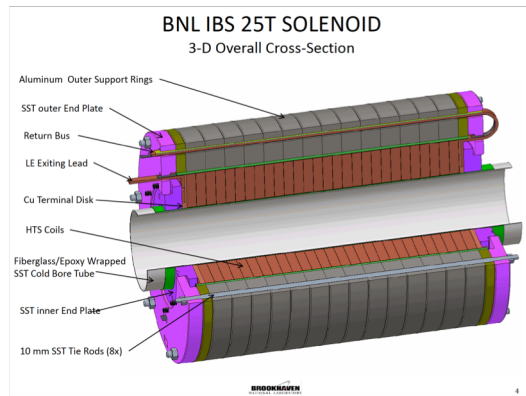
Figure 10: (a) Exclusion region with all the physics run combined done in 2018. (b) A zoomed view of the high sensitivity result.

How to improve?

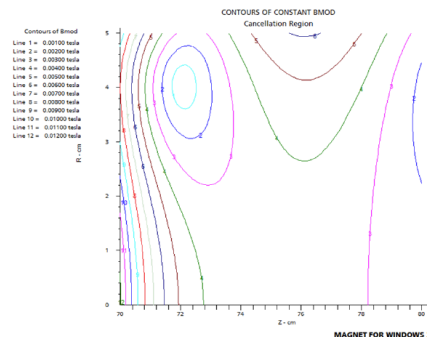
- **Maximize Signal (B^2VQ)**
 - 25T 10cm bore HTS magnet by BNL (2021)
 - 12T 32cm bore LTS magnet by Oxford (2019)
 - Higher frequencies without shrinking volume
 - Pizza Cavity (S. Youn)
 - Dielectric rings (TM_{030} and TM_{050}) (O. Kwon)
 - Improve Q-factor of cavity – YBCO cavity (D. Ahn)
- **Minimize Noise ($T_{\text{system}} = T_{\text{physical}} + T_{\text{amp}}$)**
 - Quantum Amplifier - SQUID and/or JPA
 - Optimize cryo-RF receiver chain
- **Others (DAQ efficiency)**
 - Dead-time-less DAQ

High Field & Big Bore Magnets

- 25T 10cm bore HTS magnet by BNL (2021) – Funding limited!
 - The first 16 (of 28) pancakes wound!
 - No-insulation coil design (ReBCO tapes)
 - 5 km of SC tape will be delivered

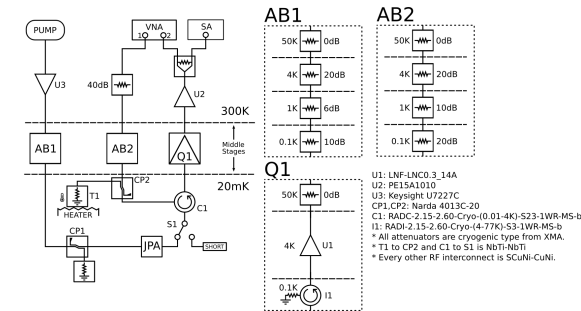
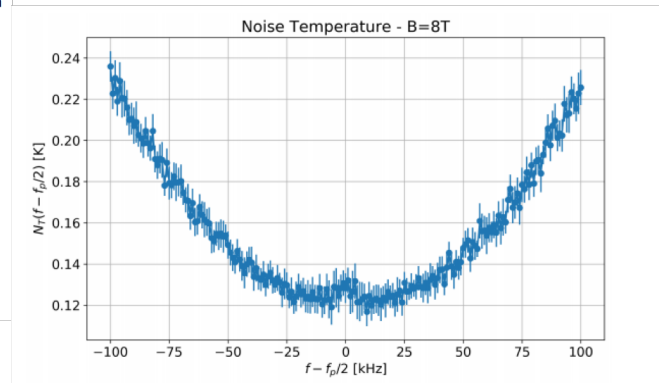
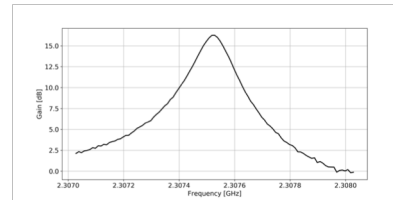
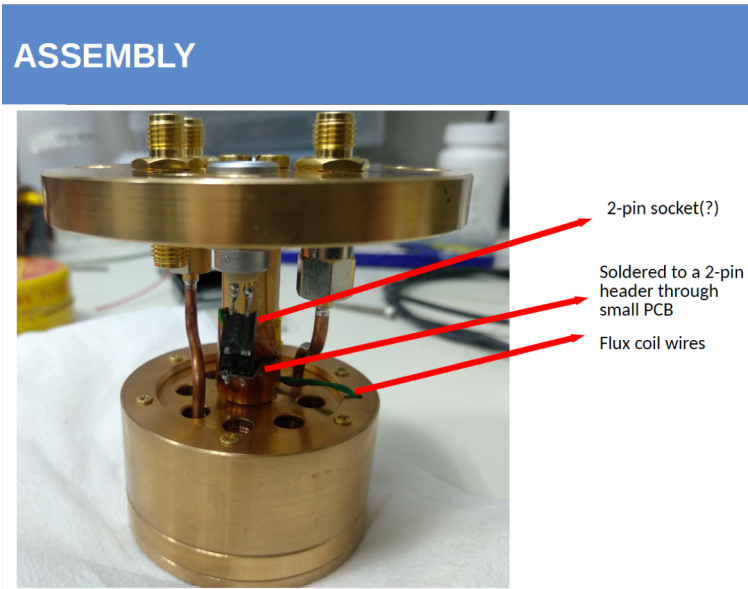


- 12 T 32 cm bore LTS magnet by Oxford Inst. (end of 2019)
 - Nb_3Sn
 - Powerful Leiden DRS1000



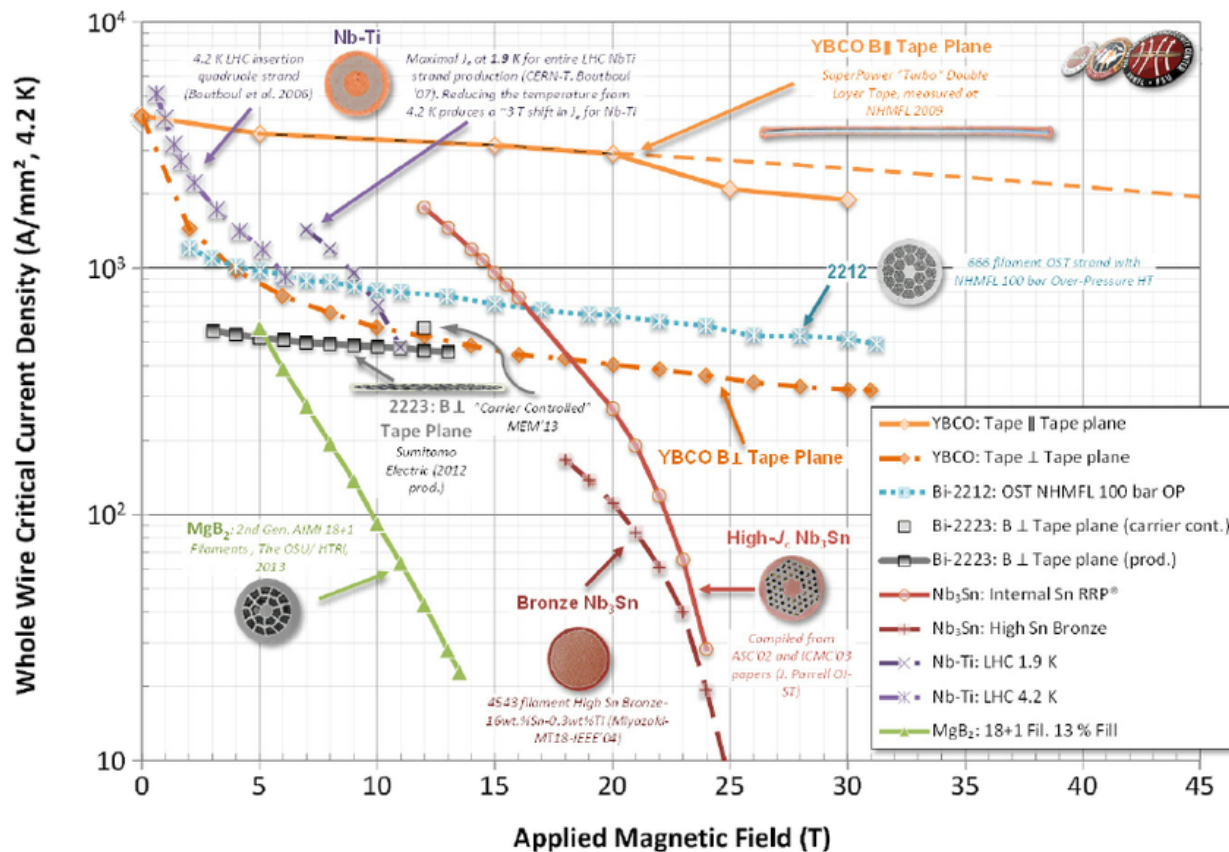
R&D Projects (JPA)

- Led by A. Matlashov and S. Uchaikin (from D-Wave)
- First batch of JPAs for PACE frequency range (2.4 GHz) from U. of Tokyo (Nakamura's group): taking adv. of their know-hows
- Noise measurement in test bench: **< 200 mK** and keeps shrinking...
- Implemented into CAPP-PACE in Aug. (2019)
- Crucial to speed up the search (**20~100 times**) w/ squeezing



R&D Projects (Superconducting cavity)

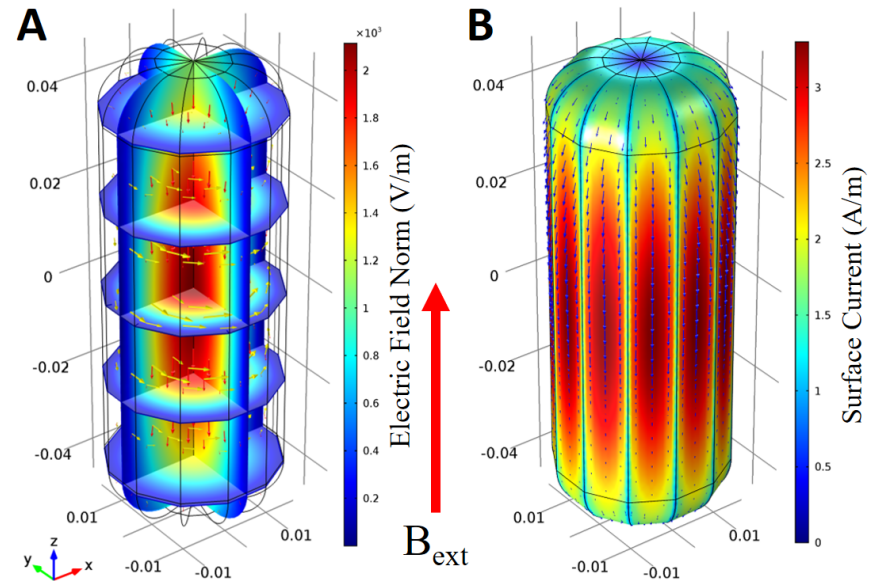
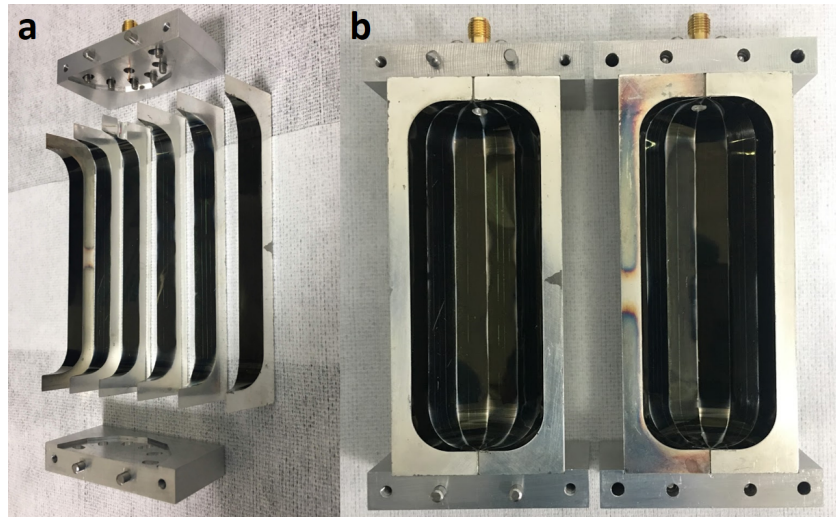
Superconductivity disappears in high magnetic field!



R&D Projects (YBCO cavity)

Superconducting cavity with YBCO tapes (grain alignment)

➤ 12 piece polygon cavity concept works!



TM_{010} mode

➤ Maintains Q-factor up to 8 T!

Maintaining high Q-factor of superconducting $\text{YBa}_2\text{Cu}_3\text{O}_{7-x}$ microwave cavity in a high magnetic field

Danho Ahn,^{1,2} Ohjoon Kwon,¹ Woohyun Chung,^{1,*} Wonjun Jang,³ Doyu Lee,^{1,2} Jinhwan Lee,⁴ Sung Woo Youn,¹ Dojun Youm,² and Yannis K. Semertzidis^{1,2}

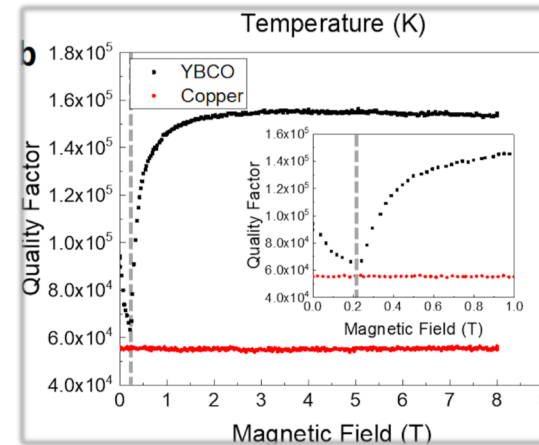
¹Center for Axion and Precision Physics Research, Institute for Basic Science, Daejeon 34051, Republic of Korea

²Department of Physics, Korea Advanced Institute of Science and Technology (KAIST), Daejeon 34141, Republic of Korea

³Center for Quantum Nanoscience, Institute for Basic Science, Seoul 33760, Republic of Korea

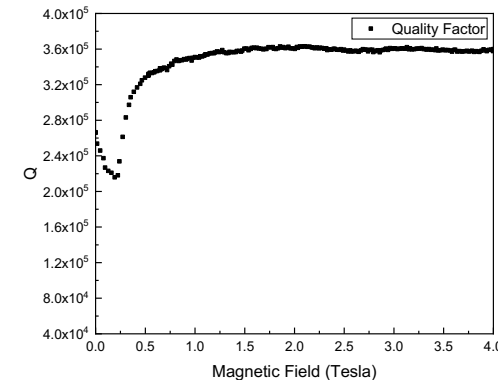
⁴Center for Artificial Low Dimensional Electronic Systems, Institute for Basic Science, Pohang 37673, Republic of Korea
(Dated: May 23, 2019)

A high Q-factor microwave resonator in a high magnetic field could be of great use in a wide range of fields, from accelerator design to axion dark matter search. The natural choice of material for the superconducting cavity to be placed in a high field is a high temperature superconductor (HTS) with a high critical field (>100 T) and depinning frequency (>10 GHz). The deposition, however, of a high-quality, grain-aligned HTS film on a three-dimensional surface is technically challenging. As a technical solution for the problem, we introduce a polygon-shaped resonant cavity with commercial YBCO (YBaCuO) thin film on a copper substrate.



Q ~ 155,000

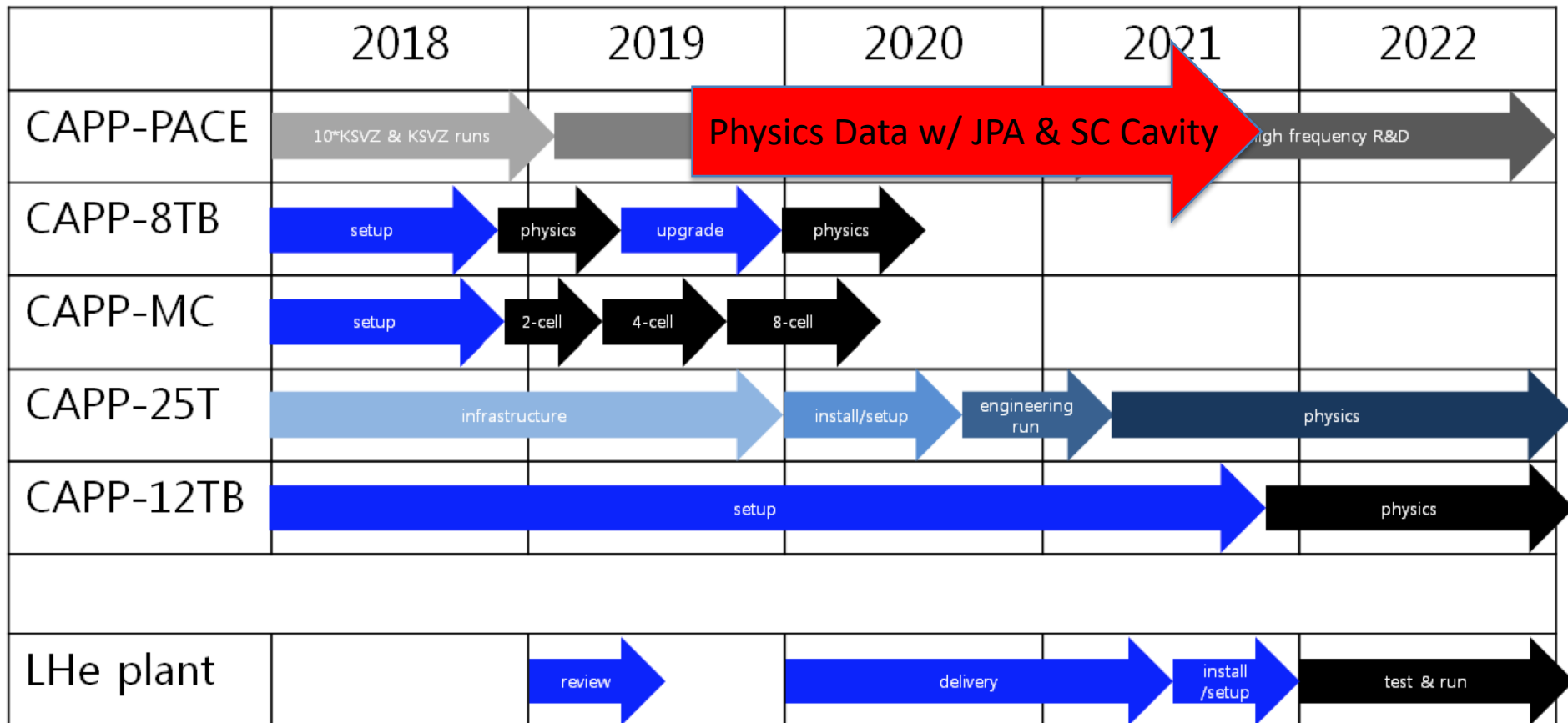
Q ~ 350,000



- It was only our first (and 2nd) Try...
- Improvement in YBCO surface is in order

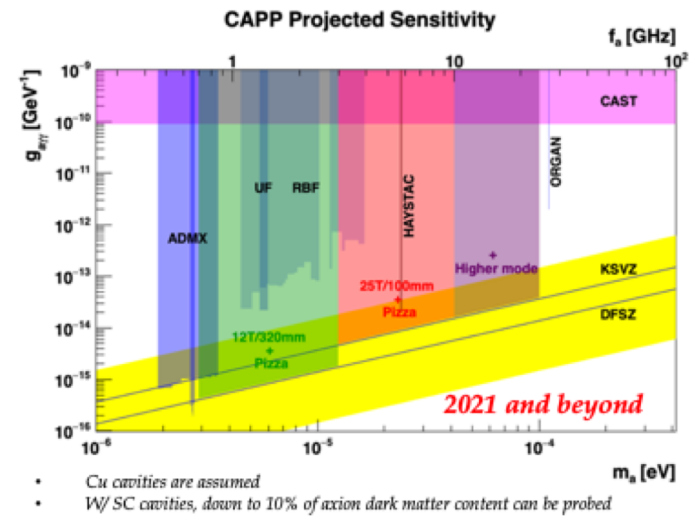
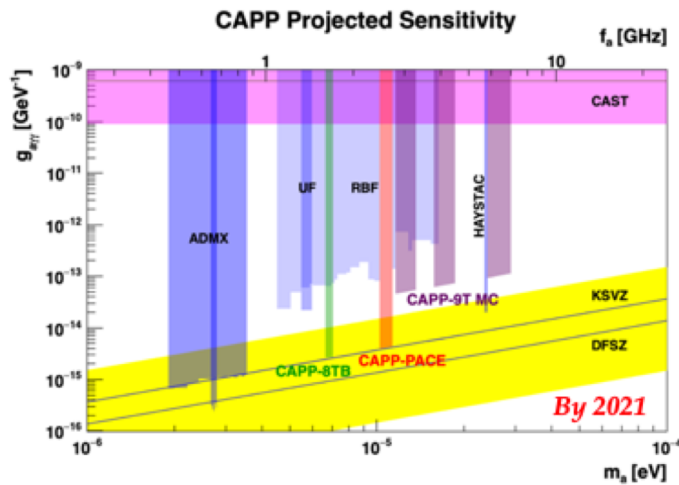
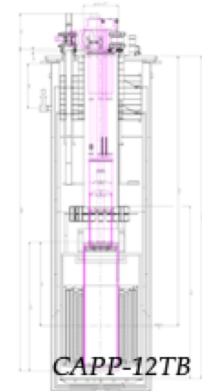
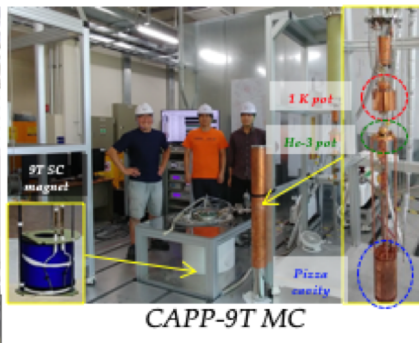
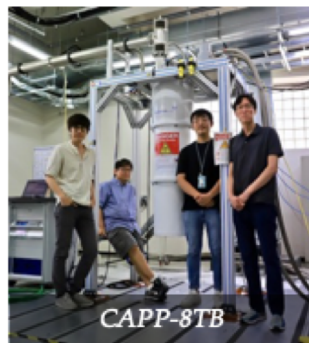
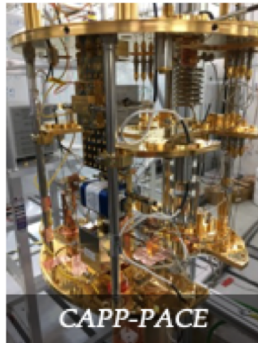
Timeline

CAPP Axion Dark Matter Search Timeline



CULTASK Prospects

- All the ingredients together, we will reach the DFSZ sensitivity even for 10% axion content in the local dark matter halo.



Summary

- **CAPP has successfully established multiple haloscope axion dark matter experiments in Korea.**
- **CAPP's pilot experiment, CAPP-PACE started to take physics data in 2018 (10*KSVZ and KSVZ runs).**
- **2 more experiments, CAPP-8TB and CAPP-MC, are ready to take data soon.**
- **CAPP will focus on taking data with JPA and YBCO cavity for axion search in 2019.**
- **Major improvement is expected with big bore (12 T, 32 cm bore) magnet (end of 2019) and high field (25 T, 10 cm bore) HTS magnet delivery.**
- **R&D on superconducting cavity looks promising!**

Upcoming Publications in 2019

- **Design and Operation of a Microwave Cavity Axion Detector for the 10 – 20 μeV**
For PRD
- **First results from the CAPP-PACE microwave cavity axion experiment**
For Physical Review Letters
- **A superconducting microwave cavity made of YBCO tapes in a high magnetic field**
For Nature (rapid communication) or PRR

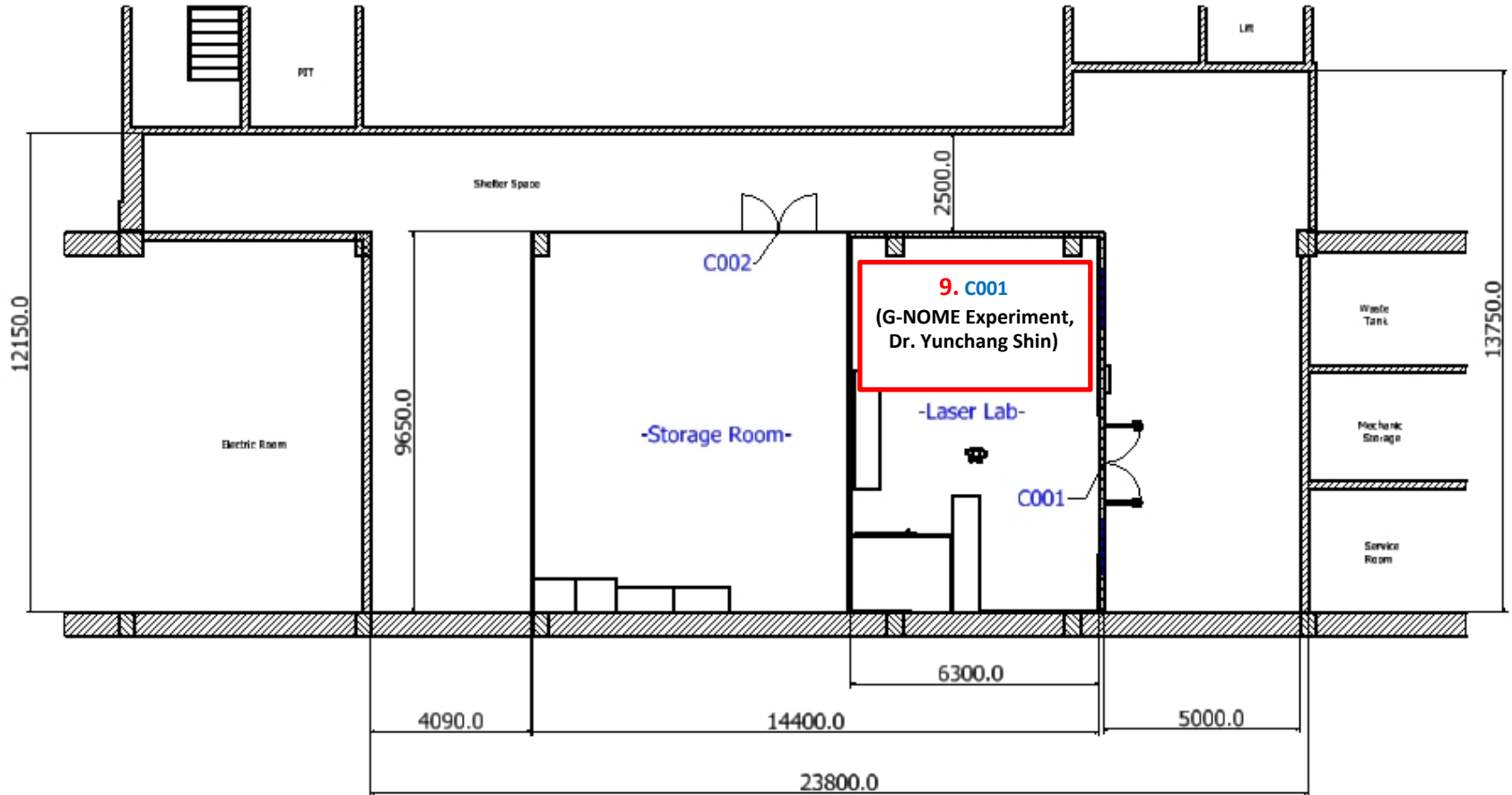
- **And Many More on...**
 - **SQUID and/or JPA test results**
 - **LVP**
 - **Physics results from CAPP-8TB**
 - **Results from CAPP-MC**
 - **Another from SC cavity development**
 - **Dielectric cavity for high frequency results**
 -

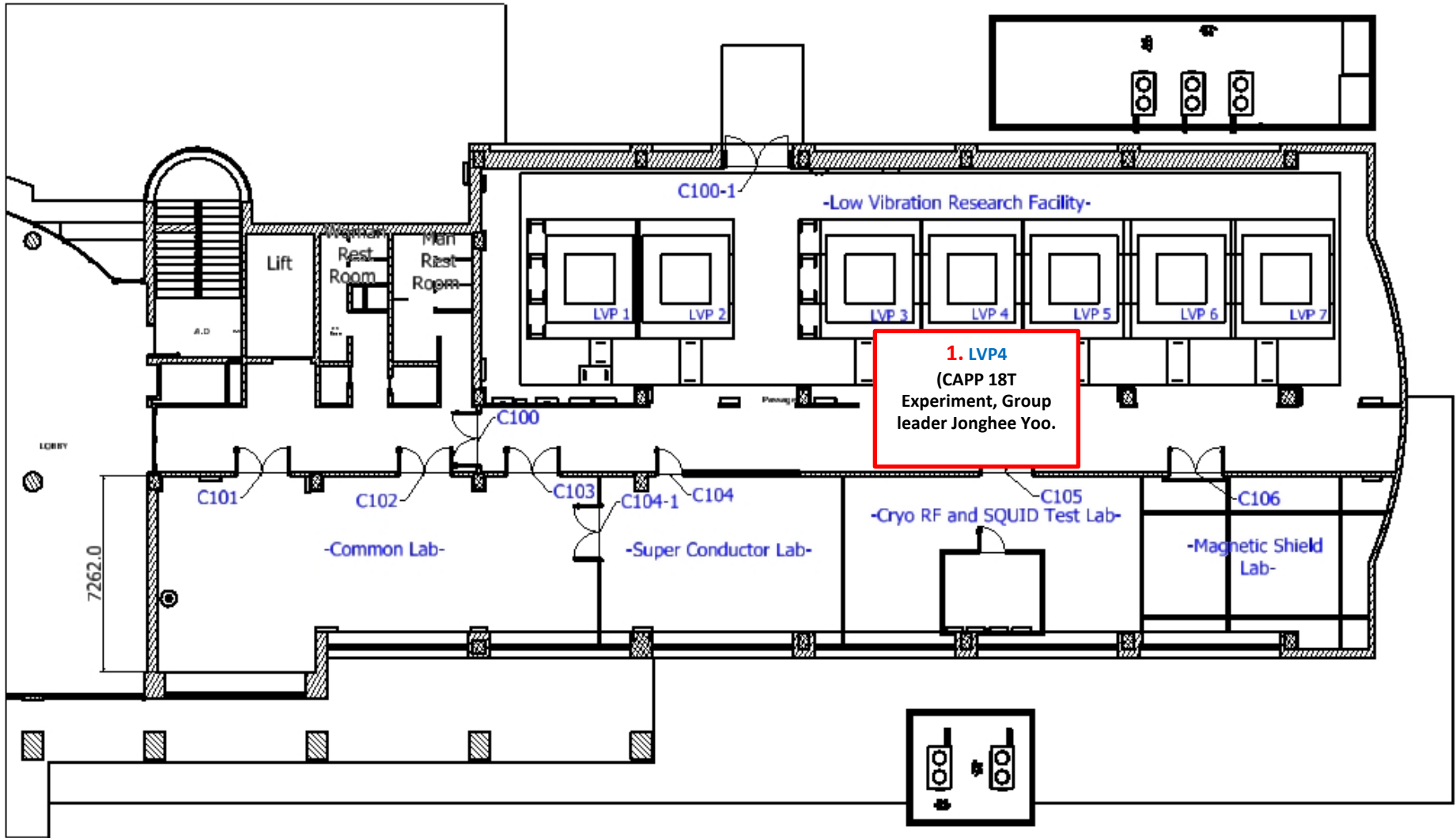
Thank You For Your Attention!

Center for Axion and Precision Physics Research (CAPP)

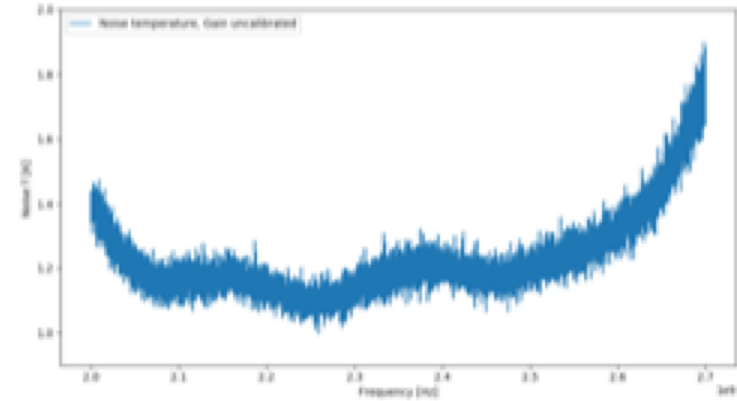
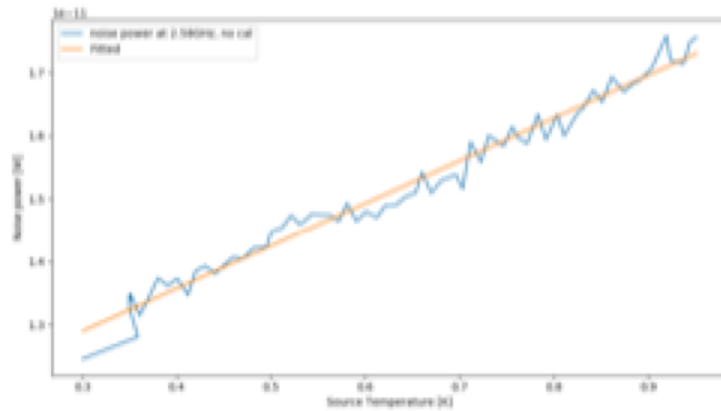
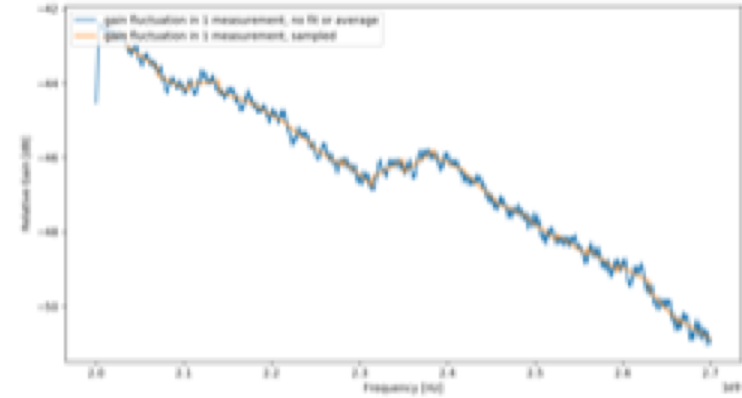
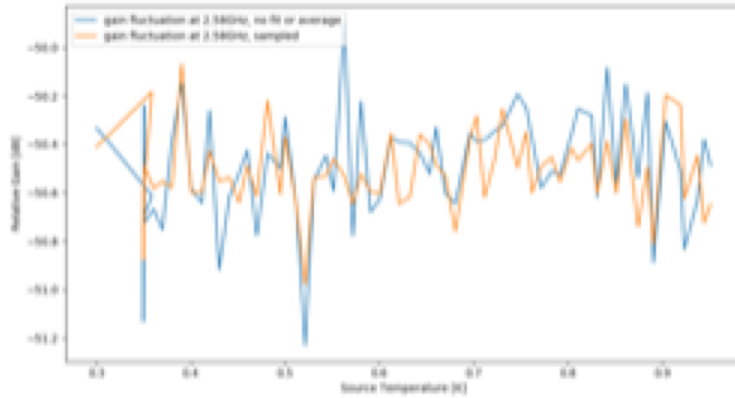
Funded by the Institute for Basic Science (IBS)

- 6 years old in Oct.
- Led by Director, Yannis Semertzidis (first gen. axion hunter)
- Physics at CAPP:
 - Dark Matter Axion Search (Cosmic Frontier)
 - Storage Ring Proton EDM (Strong CP Problem, BAU)
 - Muon $g-2$, J-PARC, COMET, CAST, ARIADNE
- Located at and working with KAIST (Korea Advanced Institute of Science and Technology)
- ~50 members

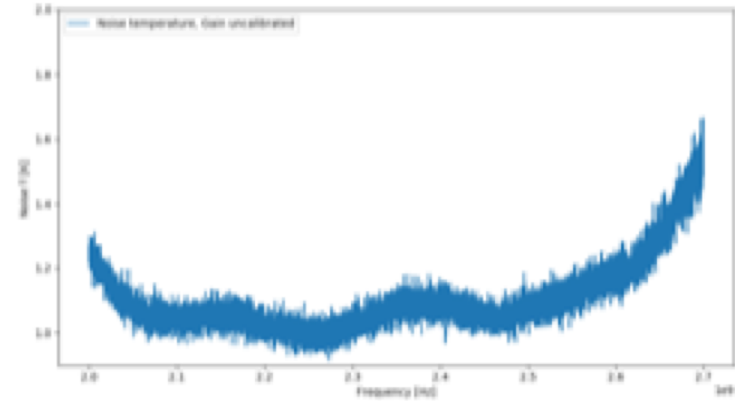
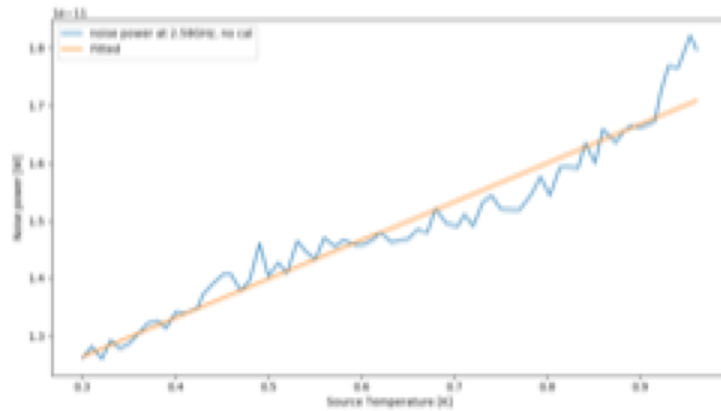
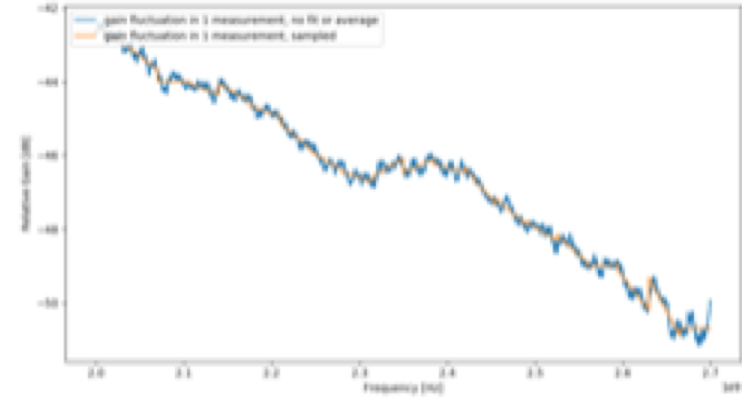
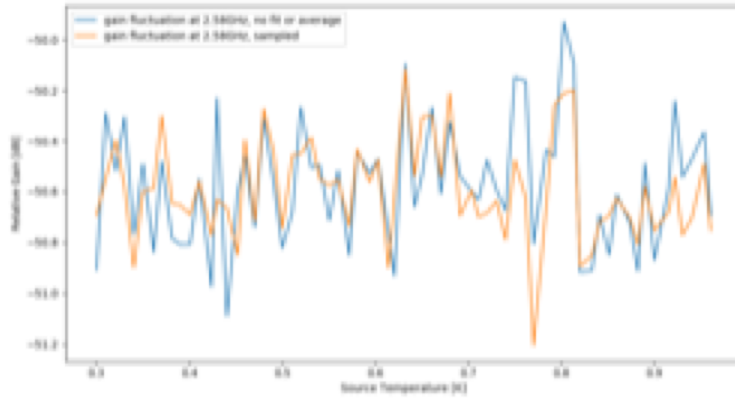




Backup Slides

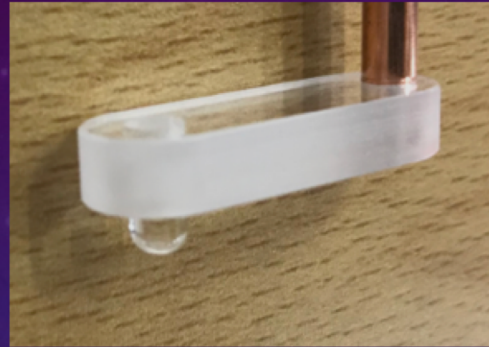


Backup Slides

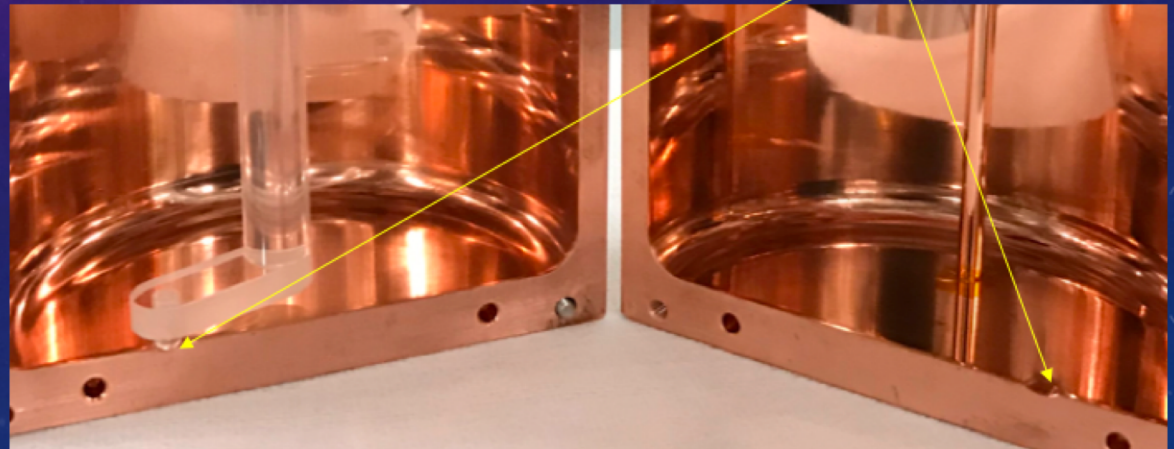


– direct touch between tuning rod and cavity wall

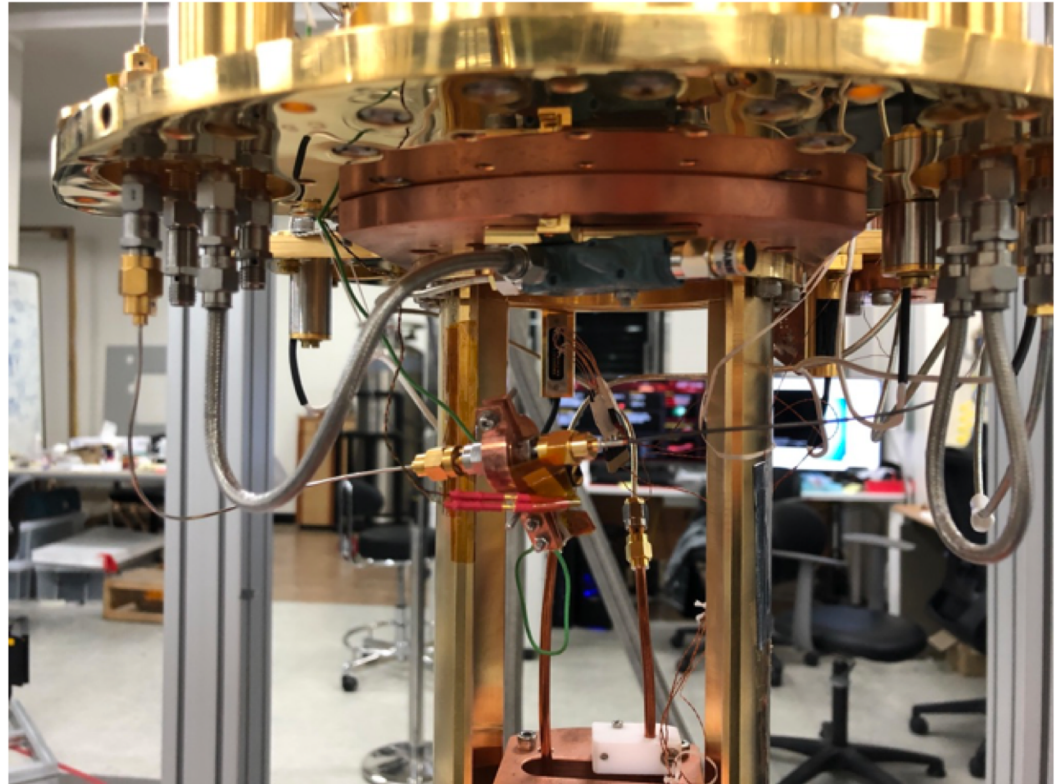
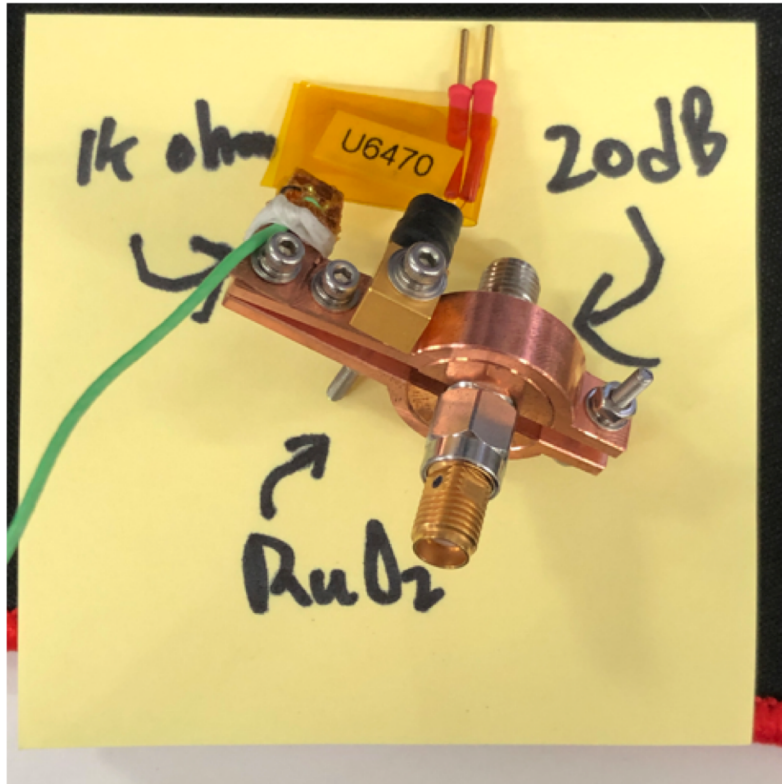
VIBRATION FREE DESIGN



Slightly bigger radius of well

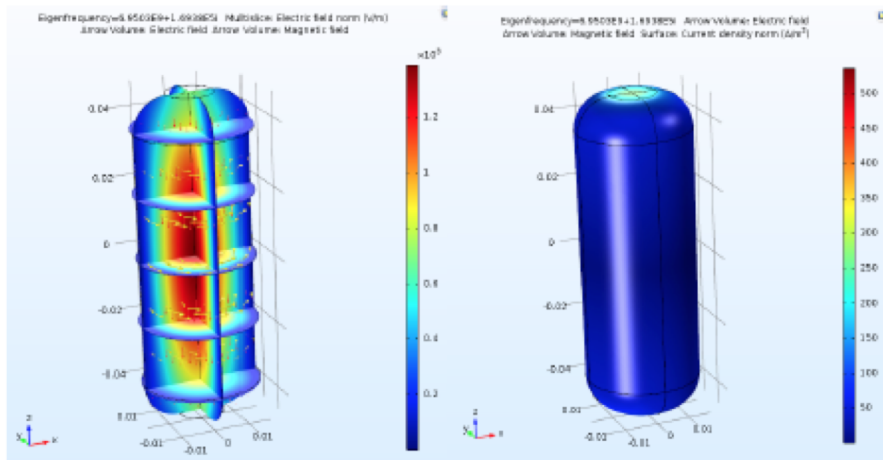


Setup - Photos

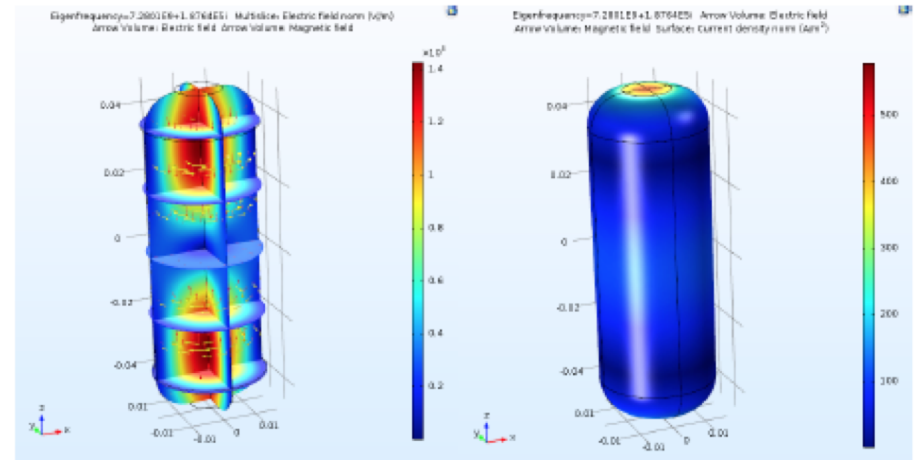


TM010 & TM011 modes

TM010



TM011



- TM010: Current density is large at the middle wall.
- TM011: Current density is large at the top and bottom.

Backup Slides(YBCO Cavity)

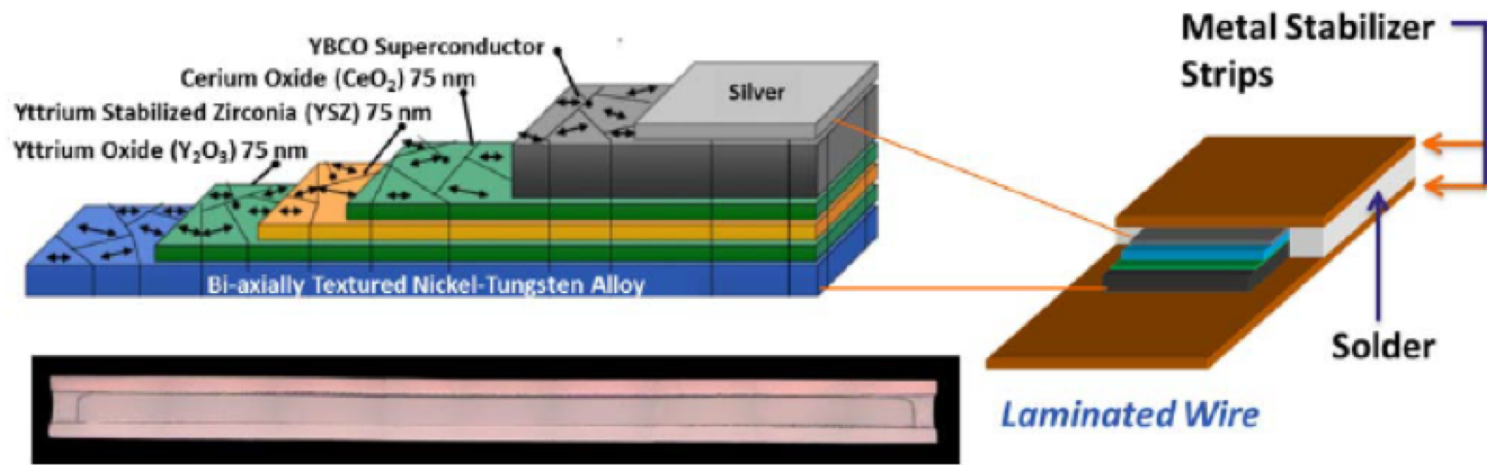


Figure 1 The architecture of the AMSC tape [19]

Backup Slides(YBCO Cavity)

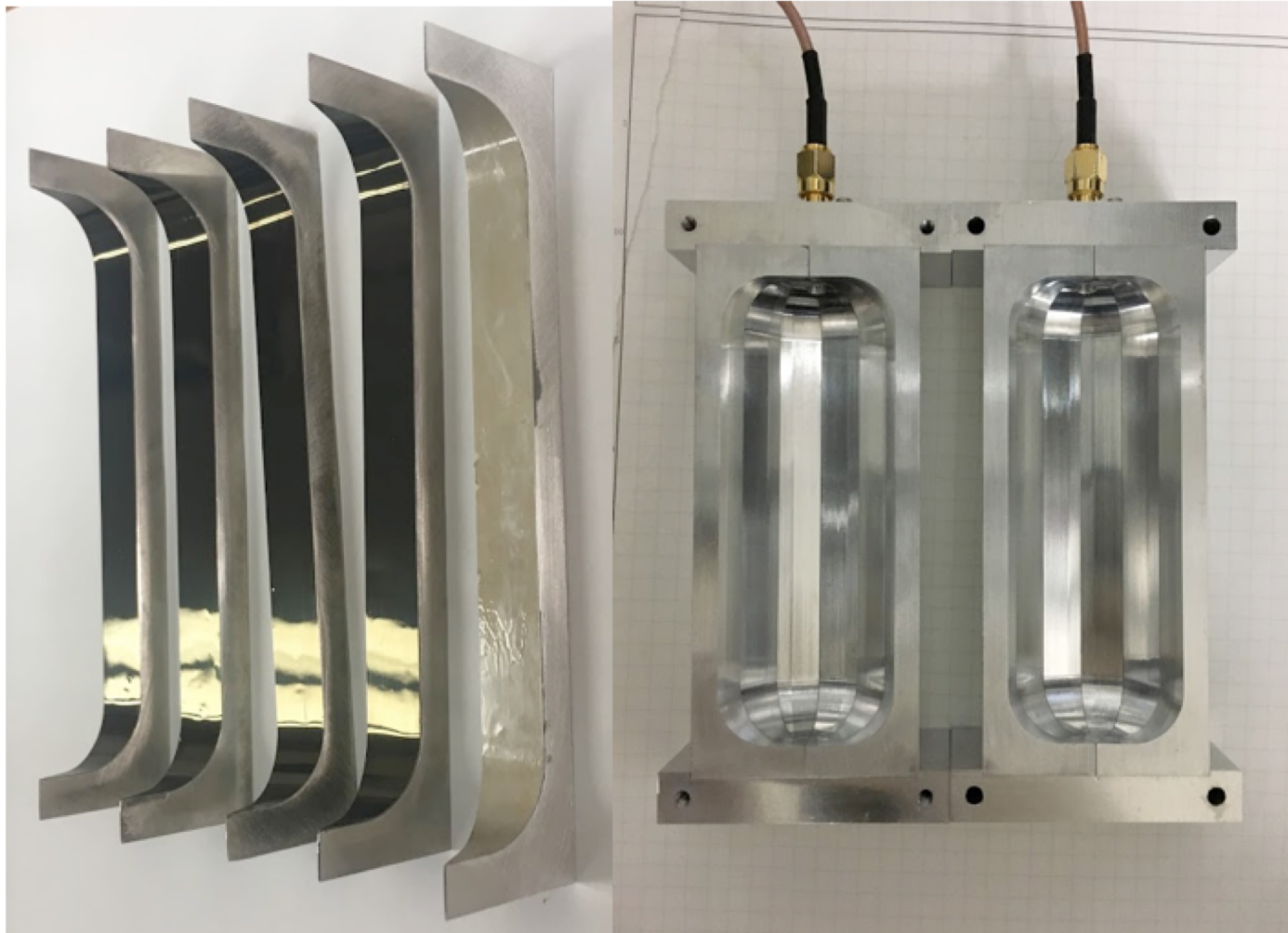
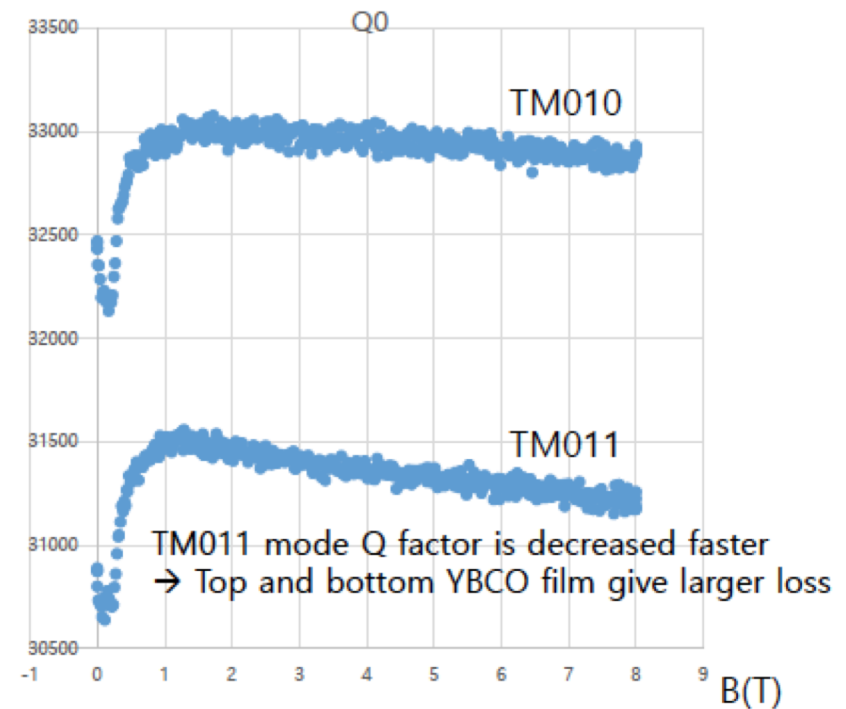
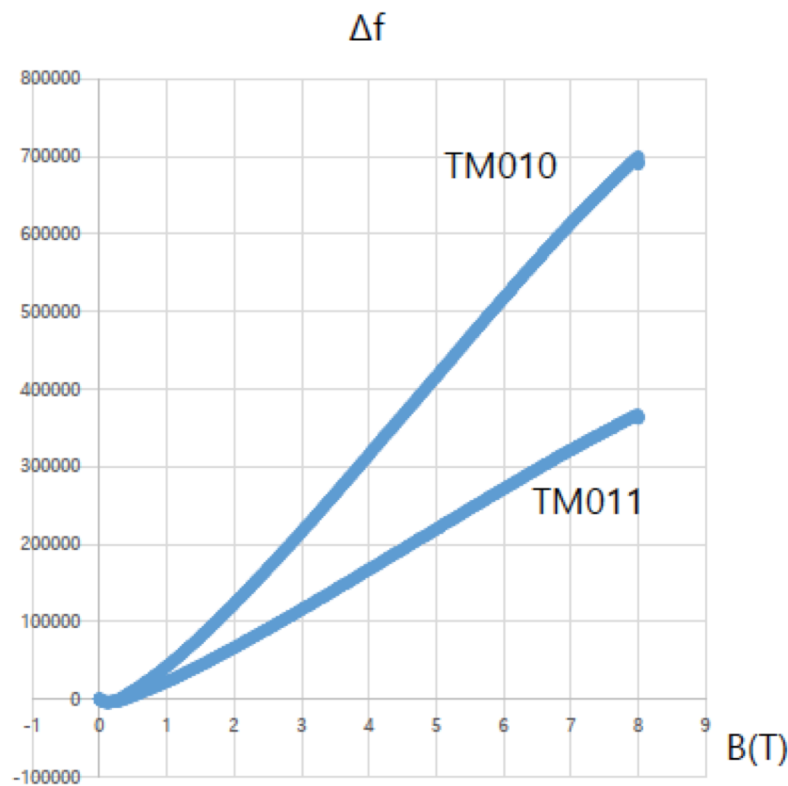


Figure 2 The structure of polygon cavity.

Magnetic Field Dependency



Backup Slides