

# Prospects and new viewpoints of Future Collider Probes

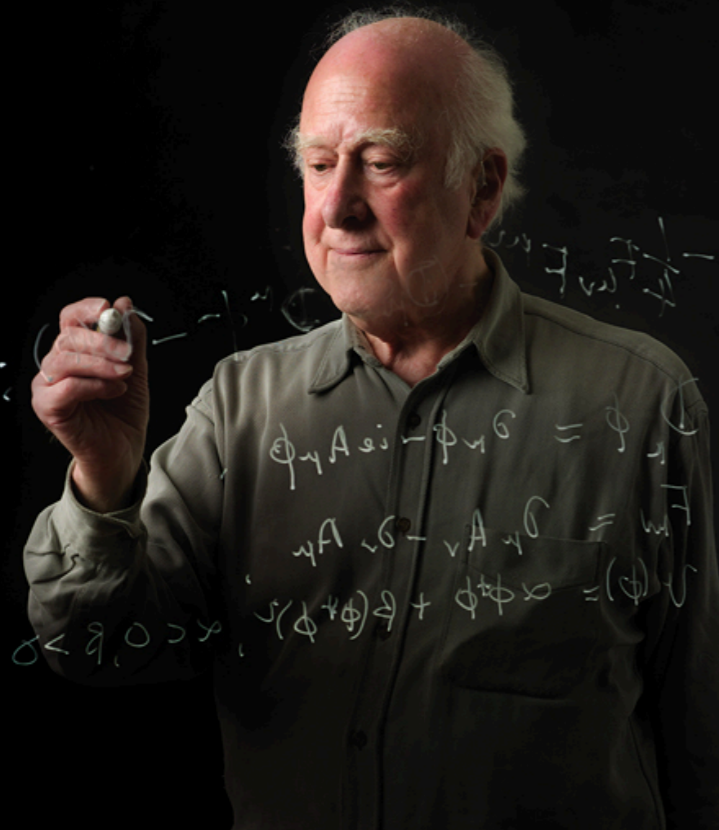
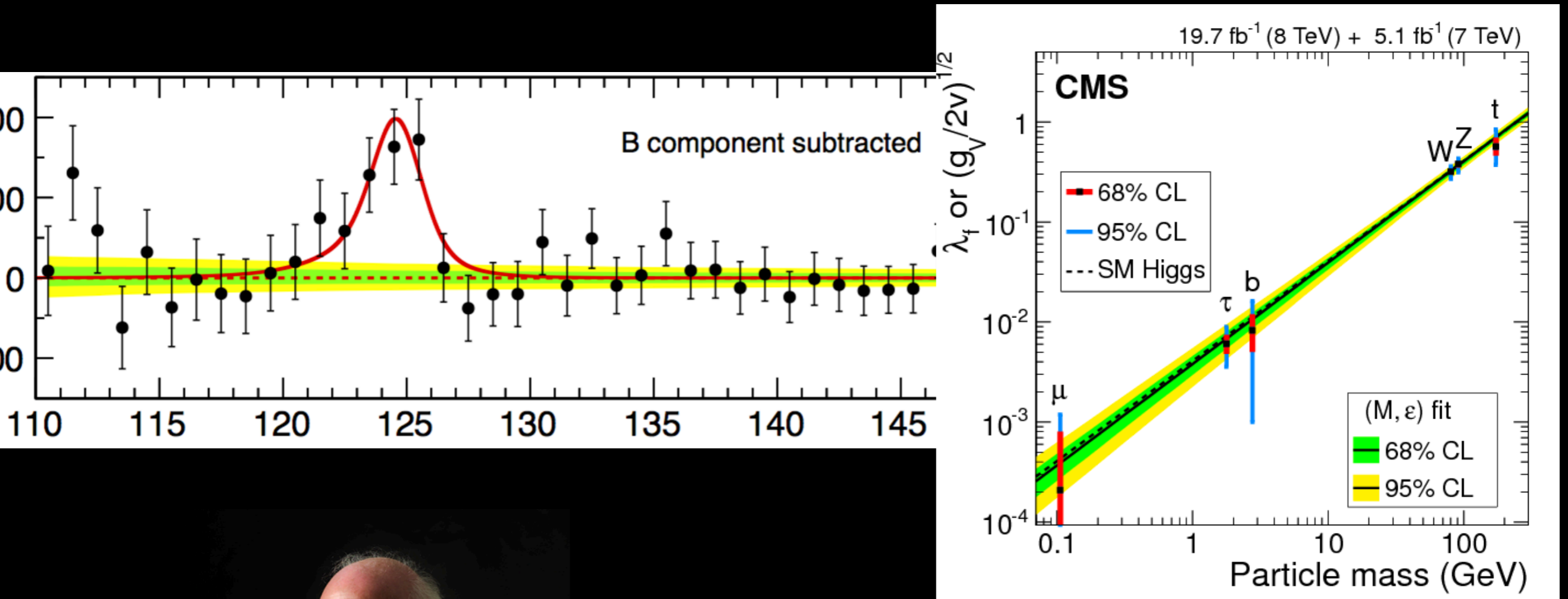


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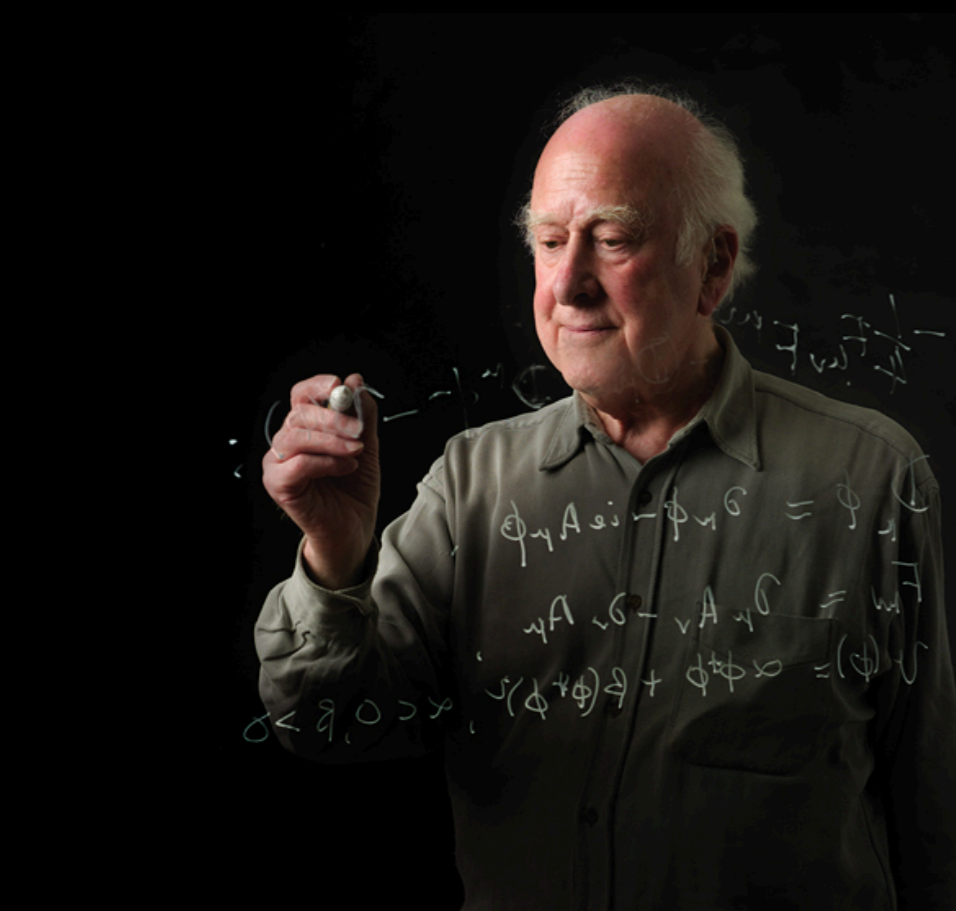
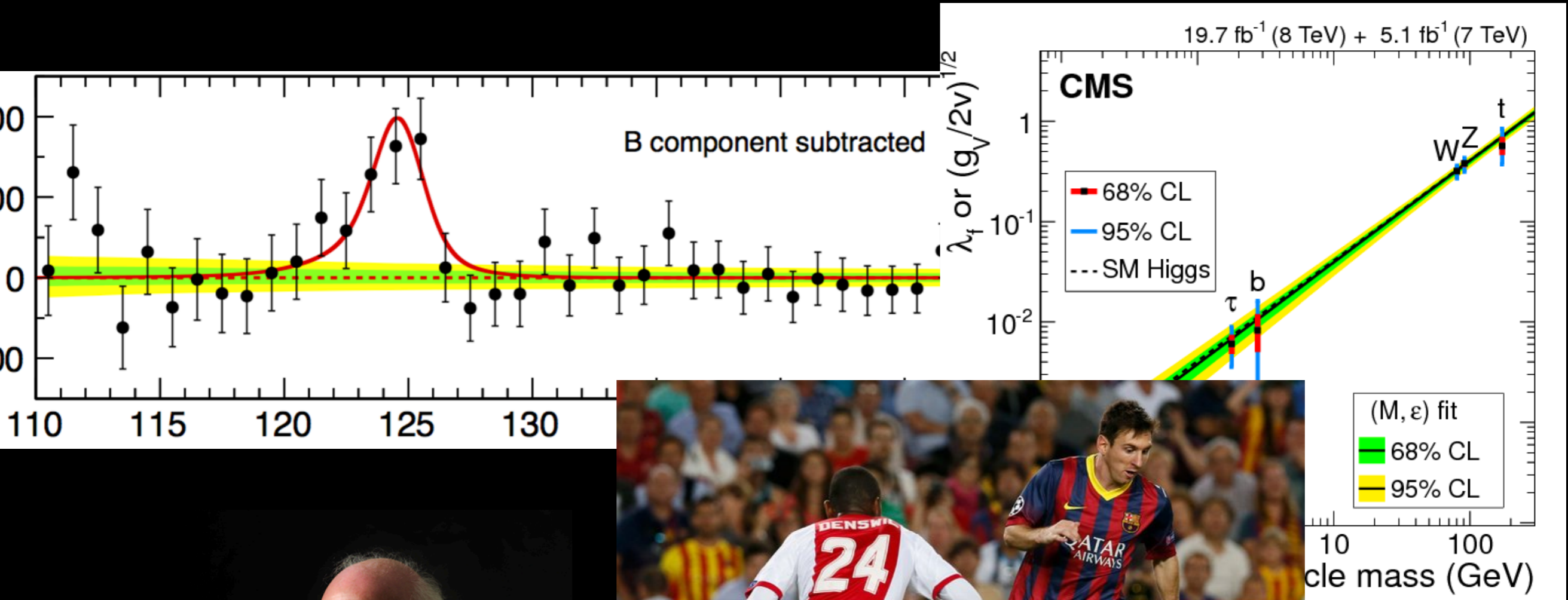
Sunghoon Jung  
Seoul National University

입자분과 General Meeting, 2019/09/21

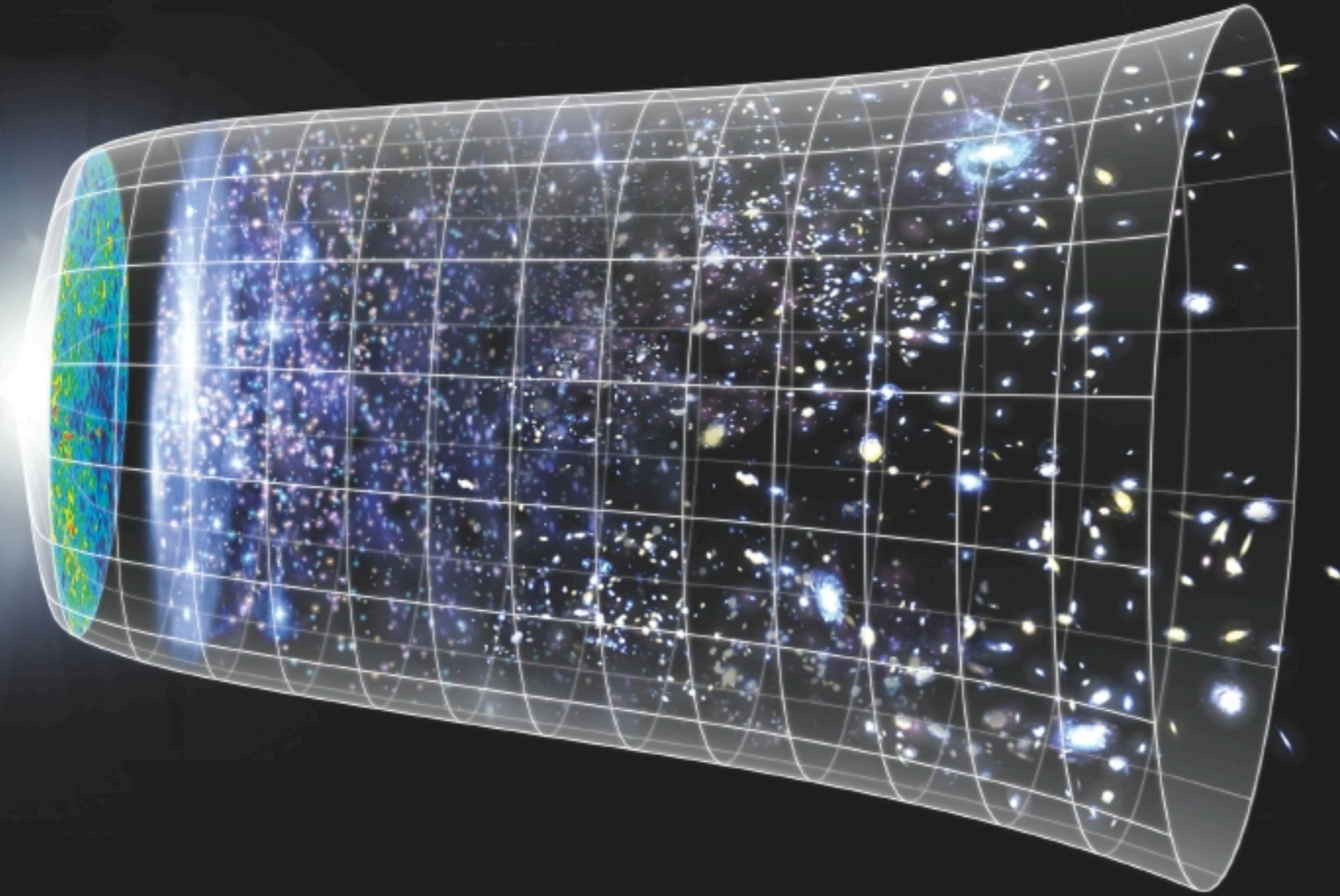
# It all starts with the Higgs boson



# Post Higgsism

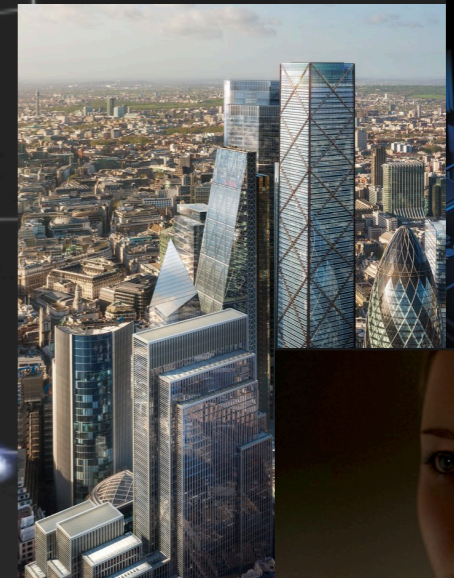
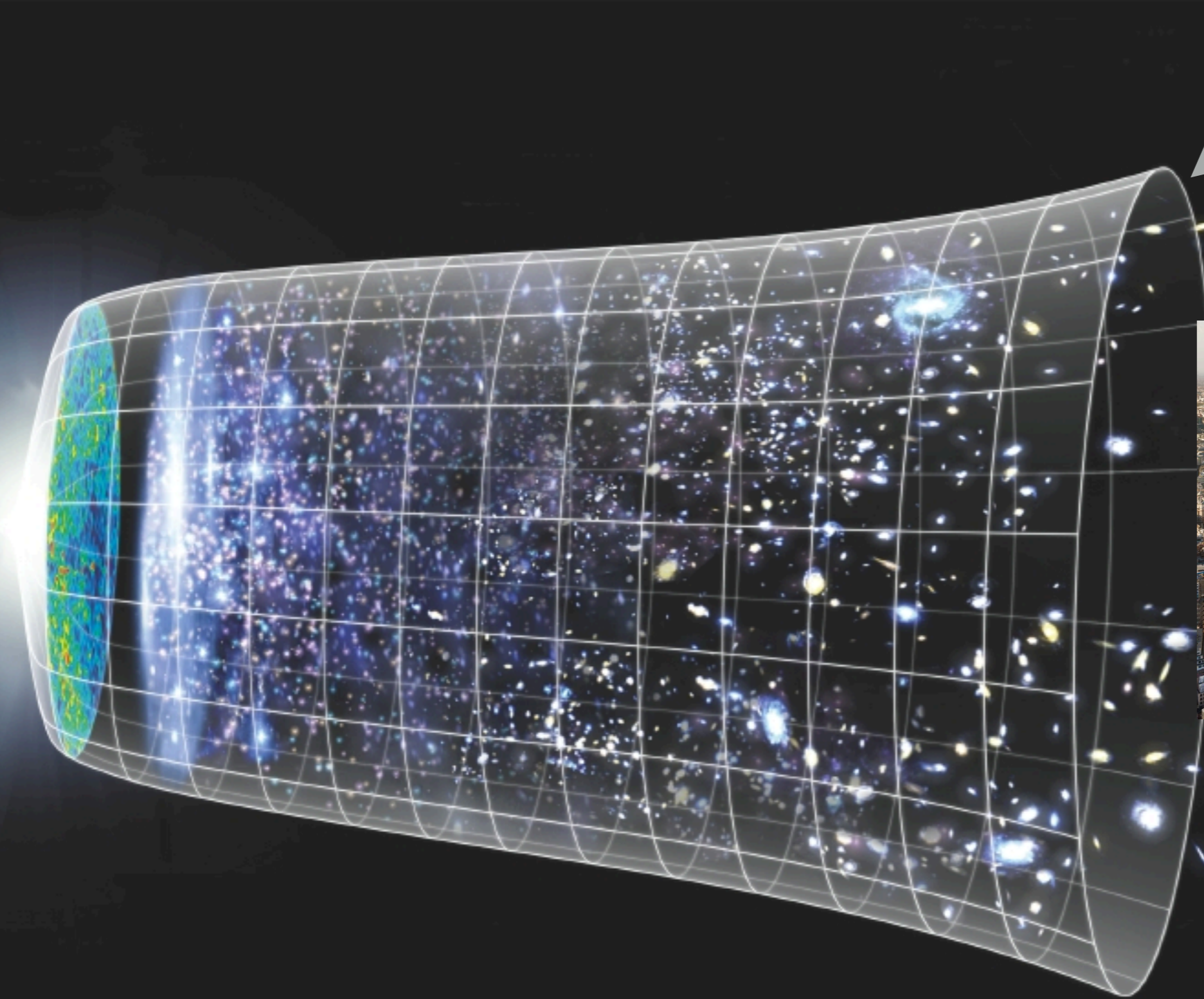


Questions now became real too.



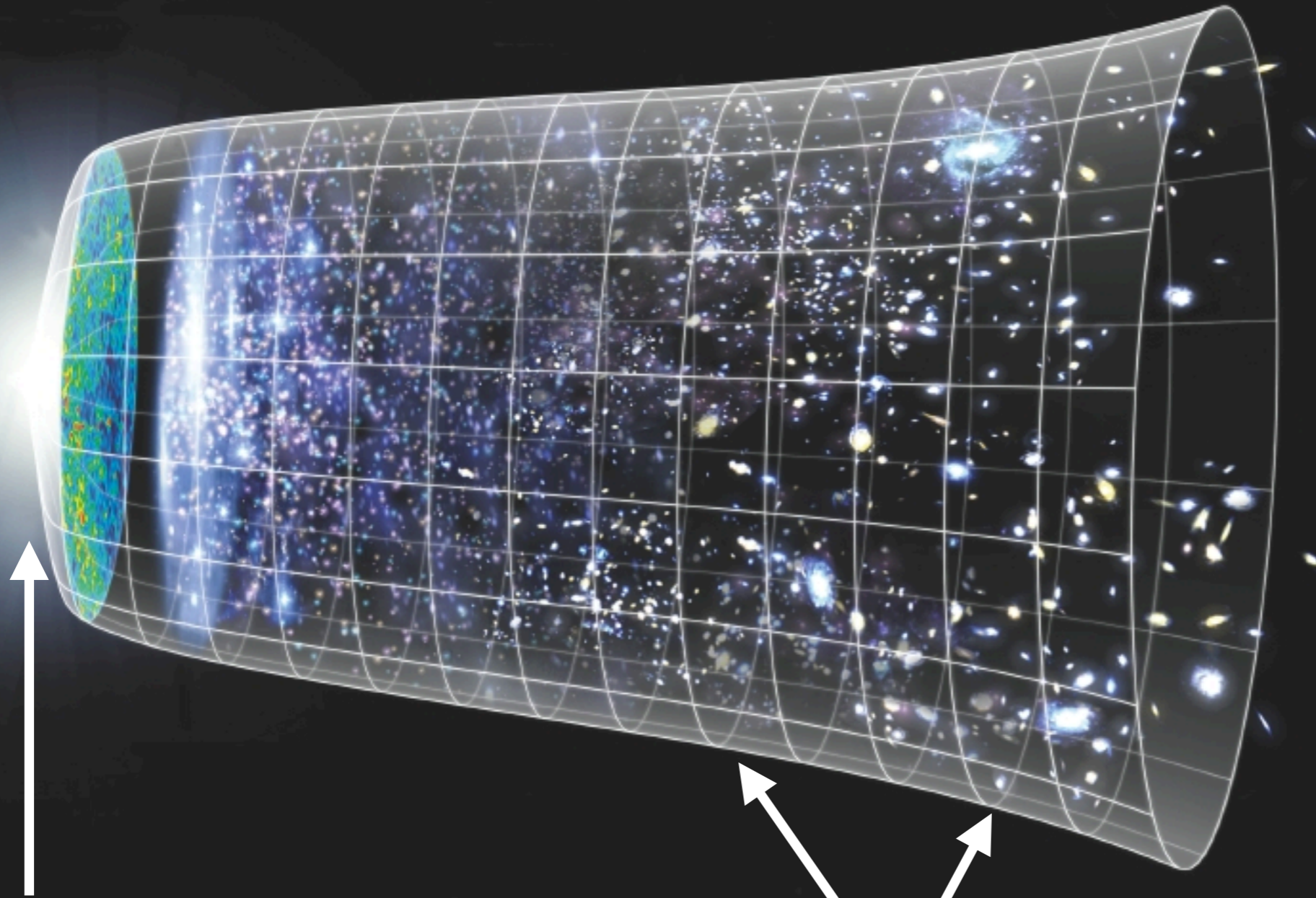
# Questions

Why do we exist as we do today?



Our existence and shape today all depend on the history of the Universe and Higgs properties.

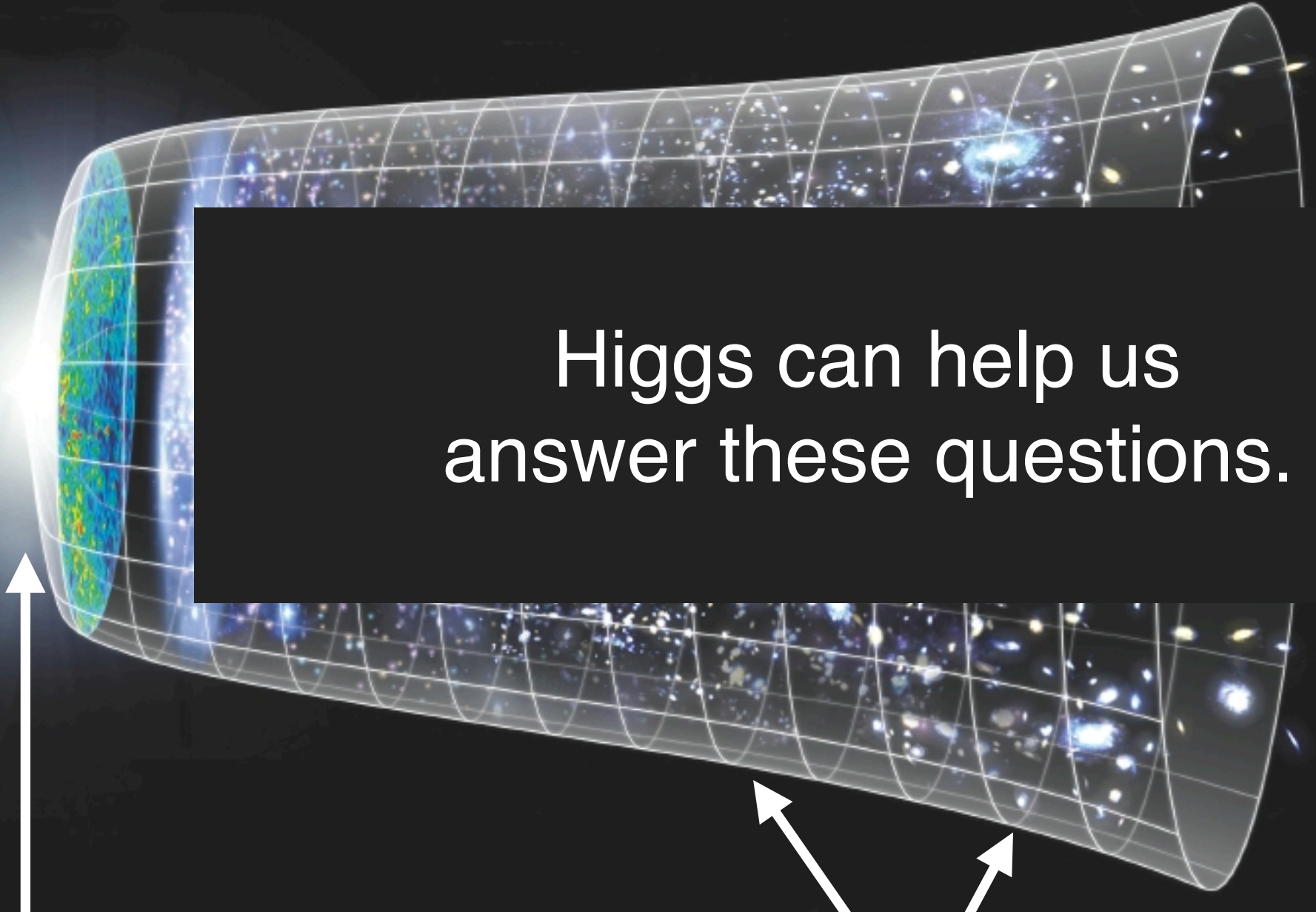
# We are so eager to see the Early Universe



What is here?!

How did Universe evolve?  
What causes that?

# Higgs as a new probe



Higgs can help us  
answer these questions.

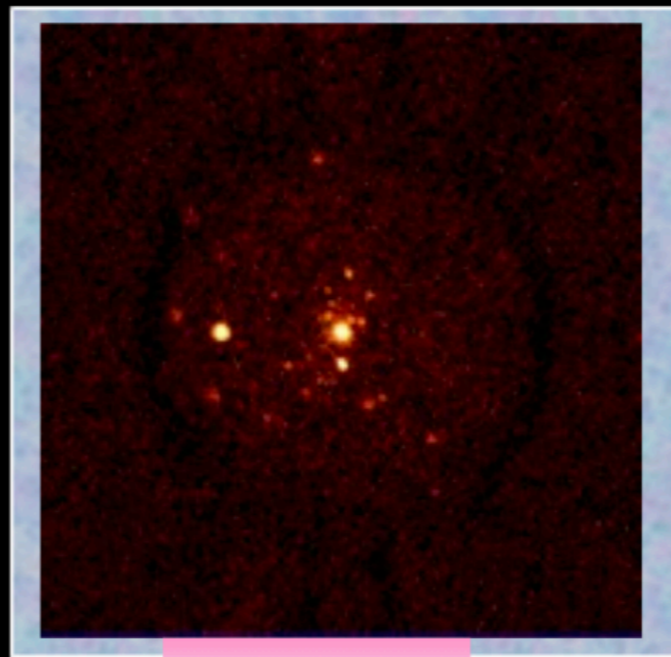
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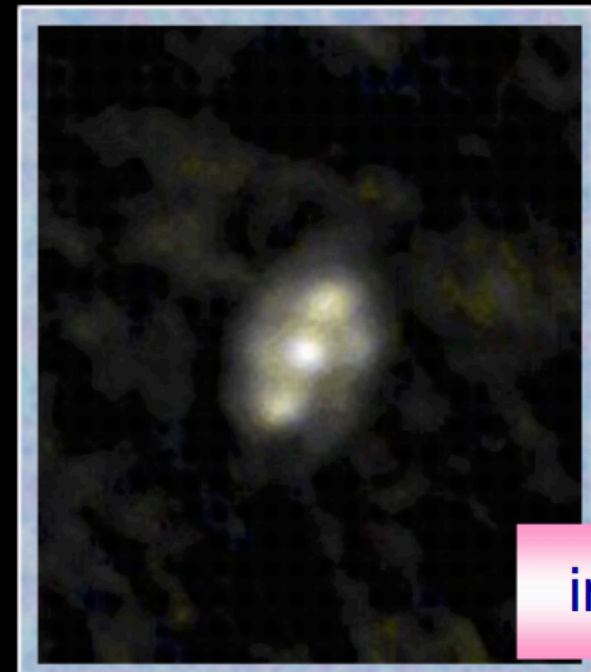
# Higgs can open up whole new ways to see and understand the Universe



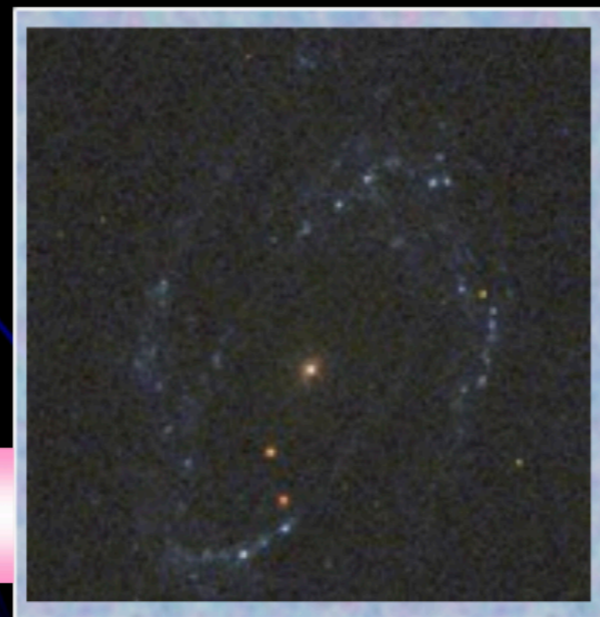
optical



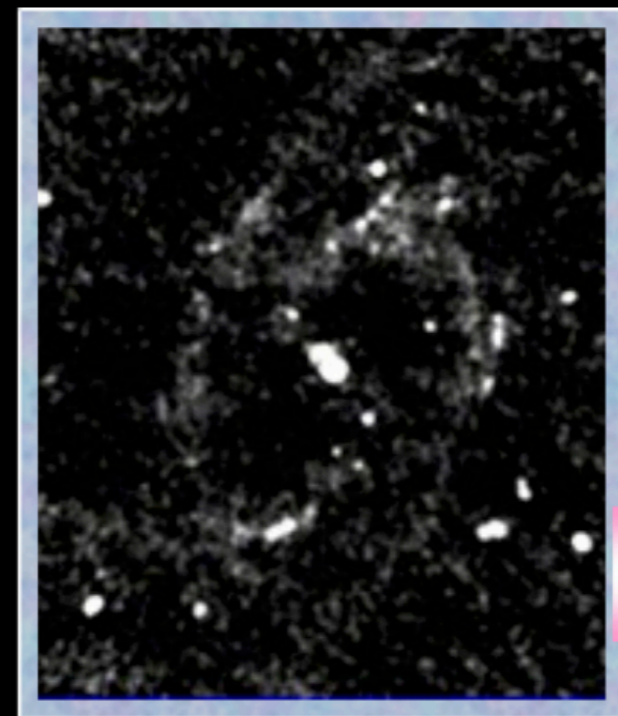
X-ray



infrared



ultraviolet

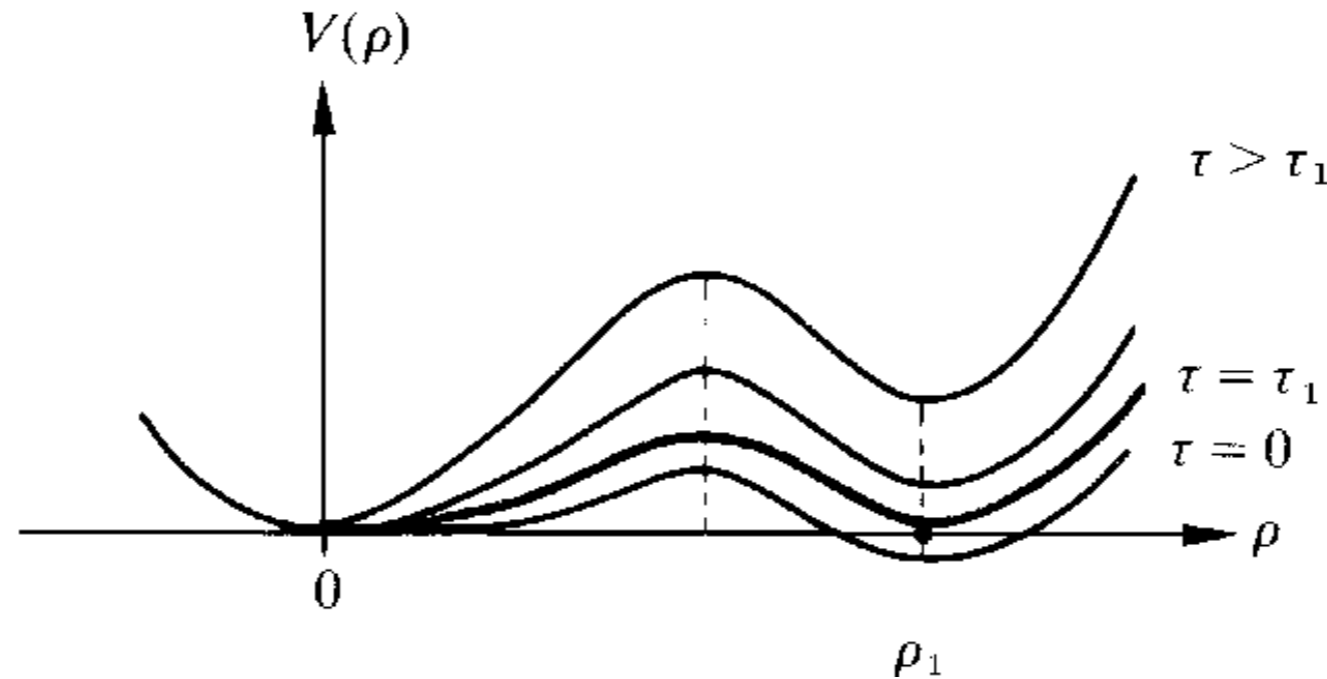


radio wave

**GW**  
neutrinos



# Higgs potential has evolved over a long time with Universe



- A long way from the Big Bang ( $10^{19}$  GeV) down to the subatomic scale (GeV).
- Must be sensitive to the initial condition and evolution history of the Early Universe.

# Energy scales of the Universe

<b>M_planck</b>	<b><math>10^{19}</math> GeV</b>	<b>initial condition</b>
<b>GUT</b>	<b><math>10^{16}</math> GeV</b>	
<b>electroweak</b>	<b>246 GeV</b>	<b>Higgs</b>
<b>QCD</b>	<b>200 MeV</b>	<b>fm</b>
<b>Supernova T</b>	<b>60 MeV</b>	<b>heavy nucleus</b>
<b>nuclear BE</b>	<b>10 MeV</b>	<b>BBN</b>
<b>Sun's core</b>	<b>1 keV</b>	
<b>atomic</b>	<b>1 eV</b>	<b>CMB</b>
<b>room T</b>	<b><math>3 \cdot 10^{-2}</math> eV</b>	
<b>m_neutrino</b>	<b><math>10^{-2}</math> eV</b>	<b><math>\ll</math> CMB</b>
<b>DE</b>	<b><math>10^{-3}</math> eV</b>	
<b>H0</b>	<b><math>10^{-33}</math> eV</b>	<b><math>10^{18}</math> sec size/age of Univ.</b>

**The Higgs cannot withstand this huge hierarchy with QM.**

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# In all, the big question is

What causes the electroweak symmetry breaking at such a low energy scale?

# More specific questions relevant to collider physics are

Is there naturalon? Which naturalon is realized in nature?

(SUSY, composite Higgs, neutral naturalness, and even relaxions!)

How did the Higgs look like in the Early Universe?

(Higgs potential/phase transition, relaxion, DM)

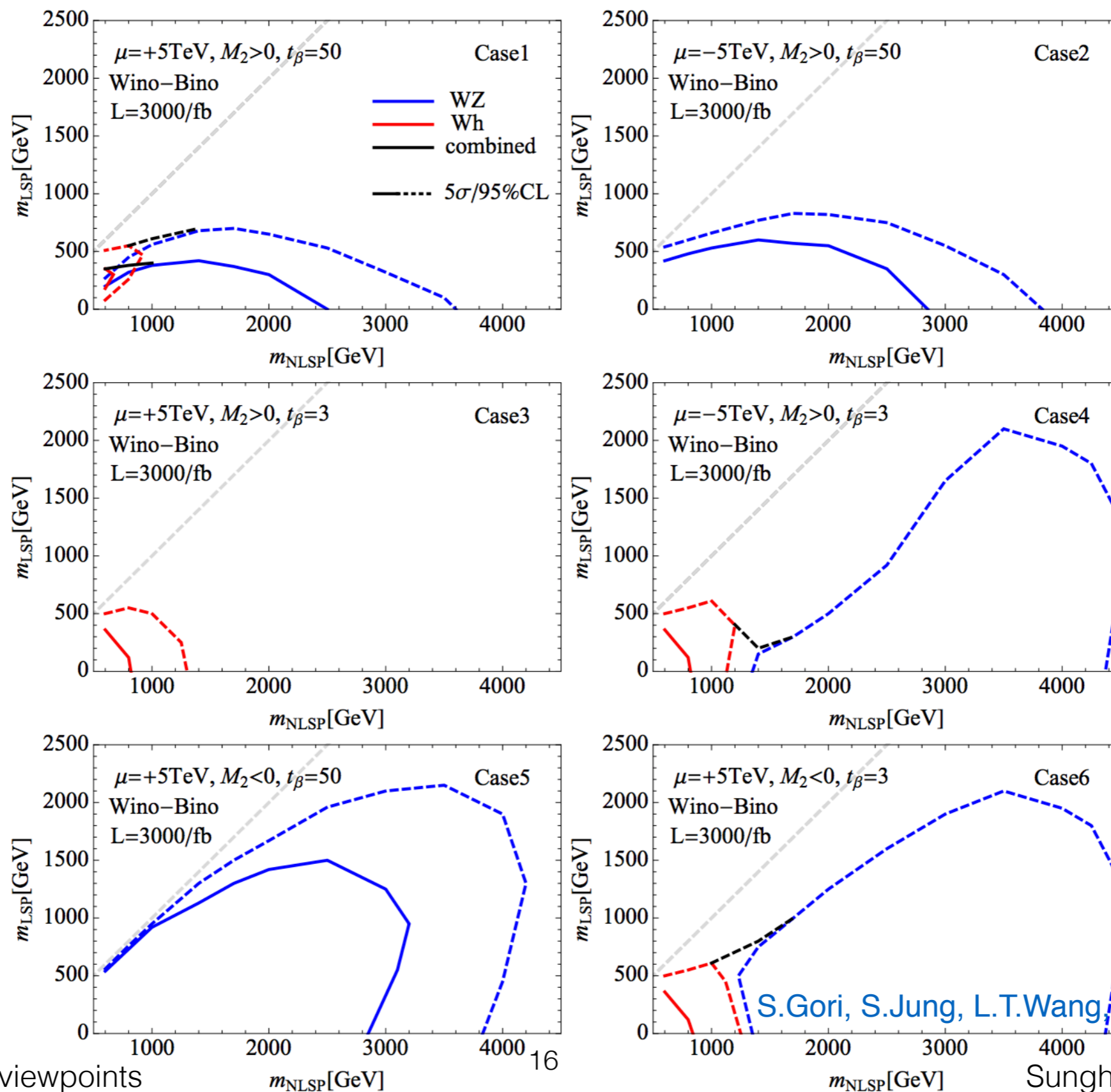
# Takehome messages

- **Future TeV-scale SUSY** is *qualitatively different*, requiring new viewpoints, insights and works.
- **Higgs precision** largely prefers to *polarization!* (of linear  $e^+e^-$  colliders.)
- **Relaxion** solutions are also *uniquely testable* at colliders.
- “**Alpha Go** sensei never speaks.”

# 1. Future SUSY

Higgsino vs. gaugino  
What is a resonance?

# Even for TeV SUSY, pure Wino-Bino exhibit so much varieties of results!

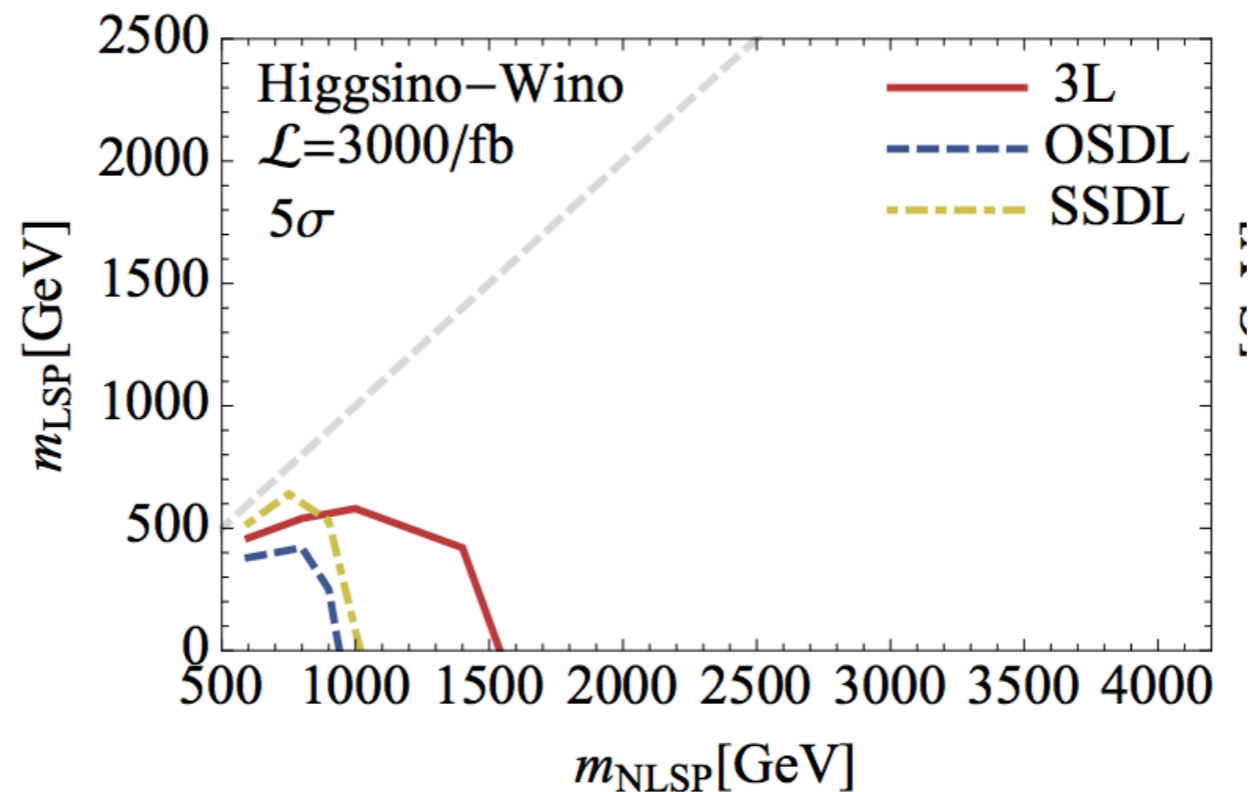


S.Gori, S.Jung, L.T.Wang, J.D.Wells, 1410.6287



# But Higgsino is different; One plot is enough!

One plot is enough with Higgsino (if either LSP or NLSP).

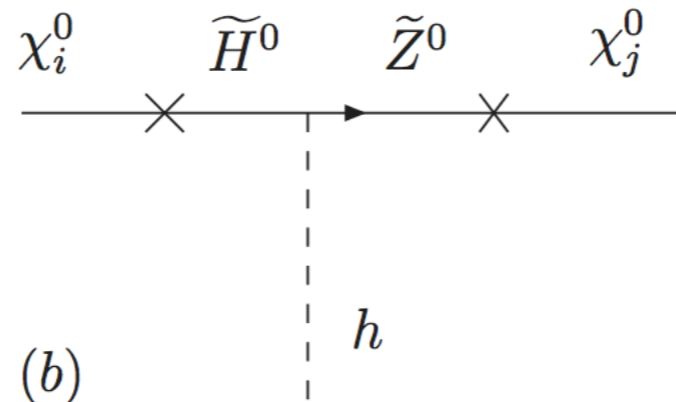
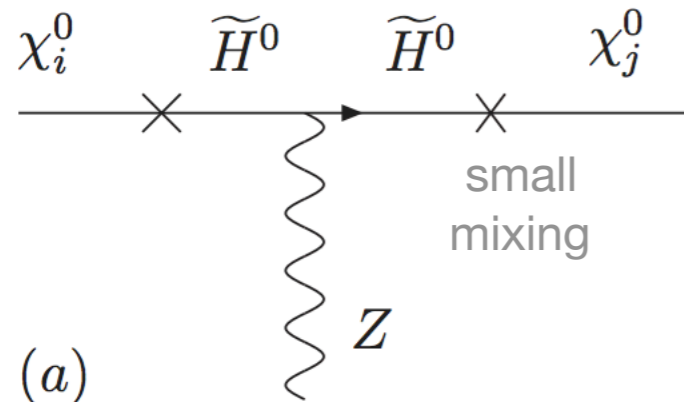


$\text{BR}(h) = \text{BR}(Z)$  ( $\sim 0.5$  instead of 1), always!

# The reason is that Higgsino's couplings to $h$ and $Z$ are inherited by the same one.

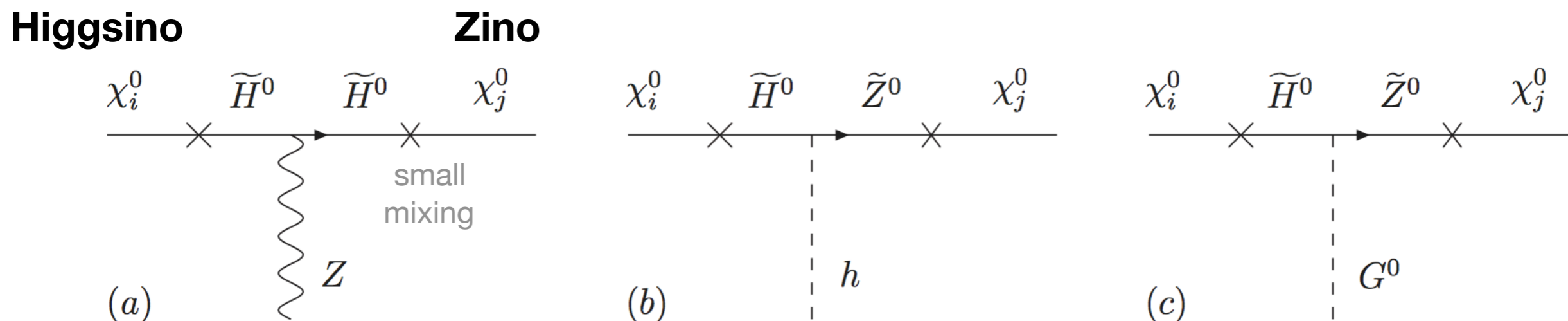
Higgsino

Zino



(a)  $\ll$  (b) ?

# The reason is that Higgsino's couplings to $h$ and $Z$ are inherited by the same one.



(a)  $\ll$  (b) ?

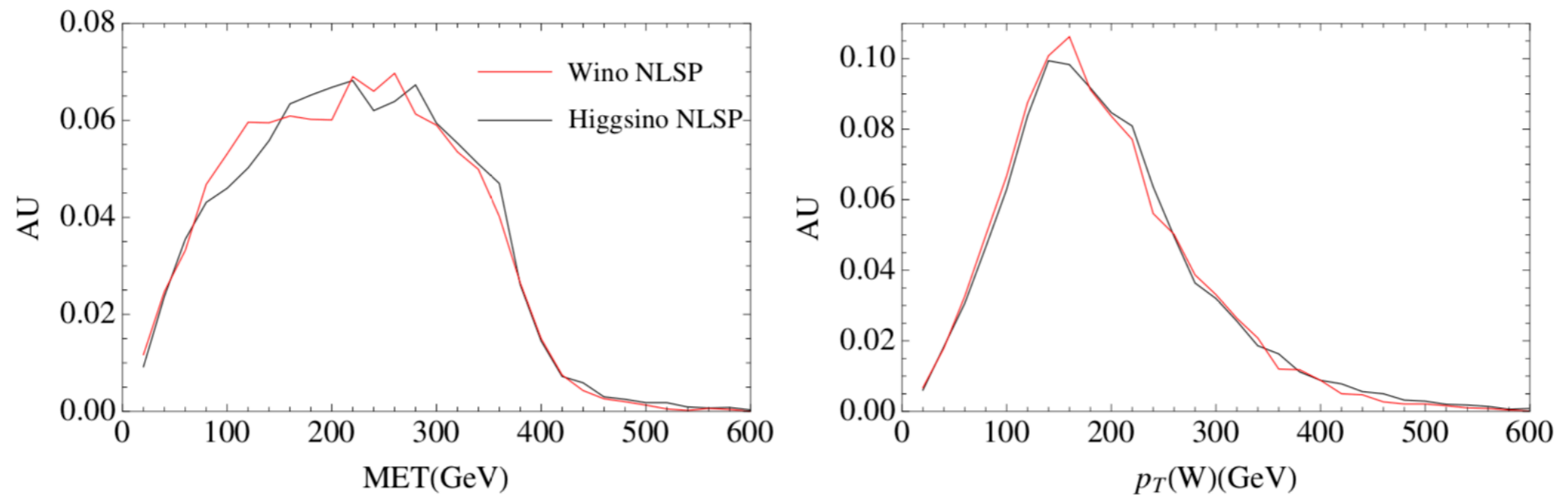
No. (a) = (c)  $\sim$  (b).

“Goldstone equivalence theorem”  
inherently relates the Higgsino couplings to  $Z$  and  $h$ .

This becomes more and more accurate at higher energy,  
hence more relevant at future collider experiments.

# Implication to LHC Inverse Problem

Typical LHC Inverse Problem.

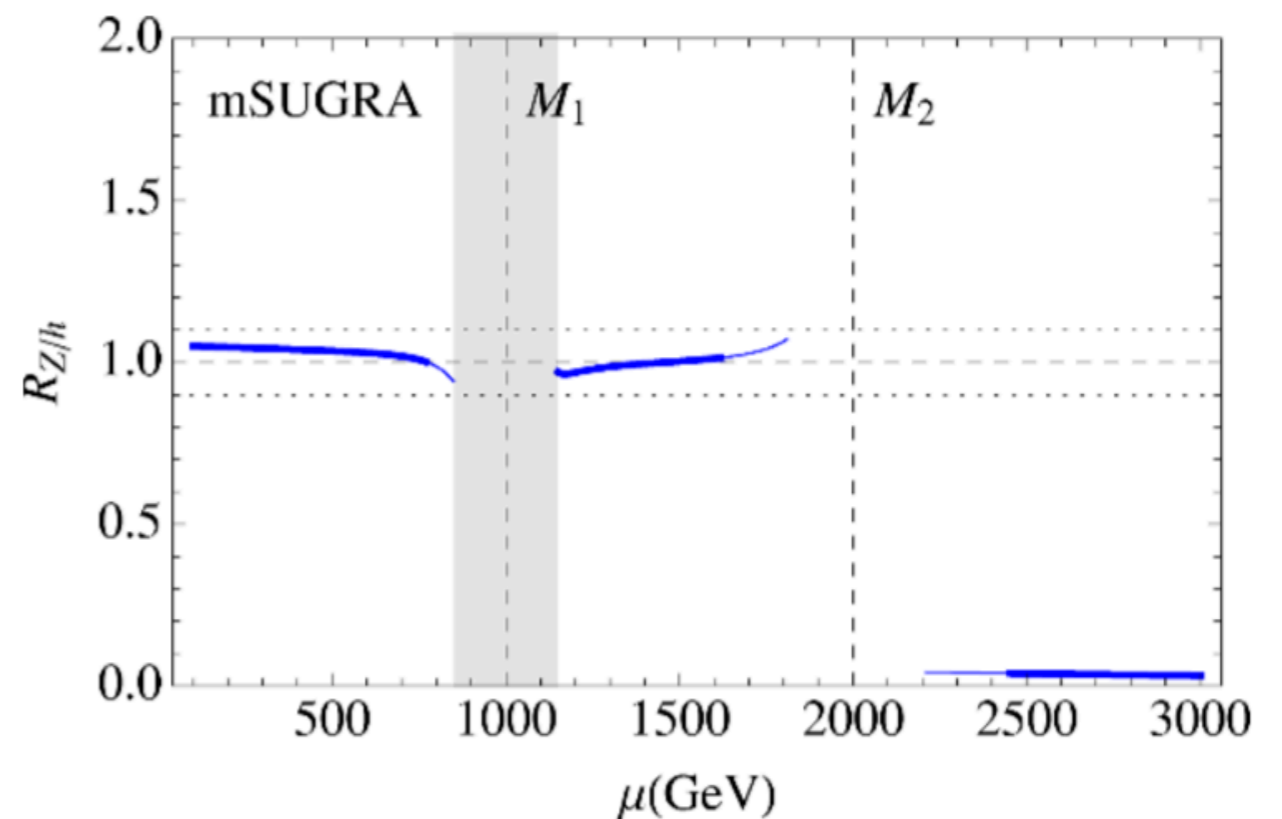
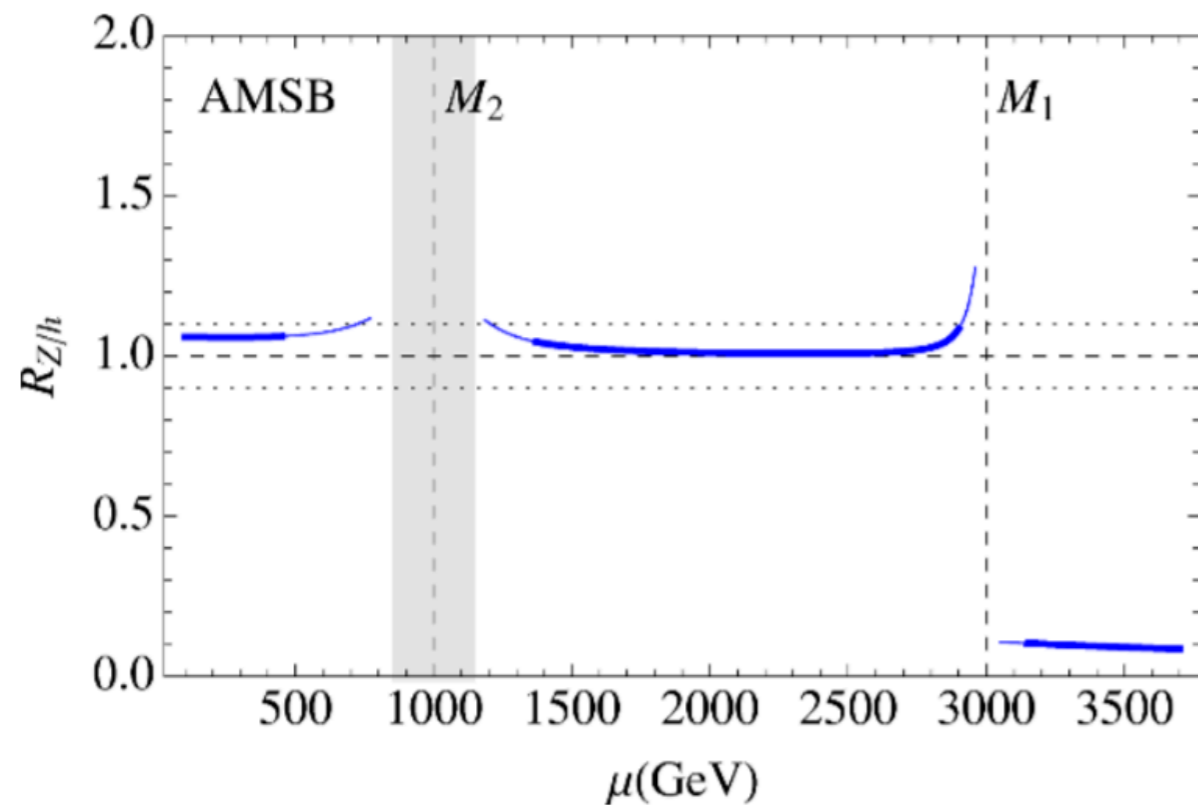


Model	parameters ( $M_1, M_2, \mu, t_\beta$ )	$\sigma(W^+W^-)$	$\sigma(W^\pm Z)$	$\sigma(ZZ)$
Wino-NLSP	0.5 TeV, 1.0 TeV, -2.0 TeV, 4.3	0.60 fb	1.1 fb	0 fb
Higgsino-NLSP	0.2 TeV, 2.0 TeV, 0.8 TeV, 2.0	0.61 fb	1.1 fb	0.02 fb

S.Jung, 1404.2691

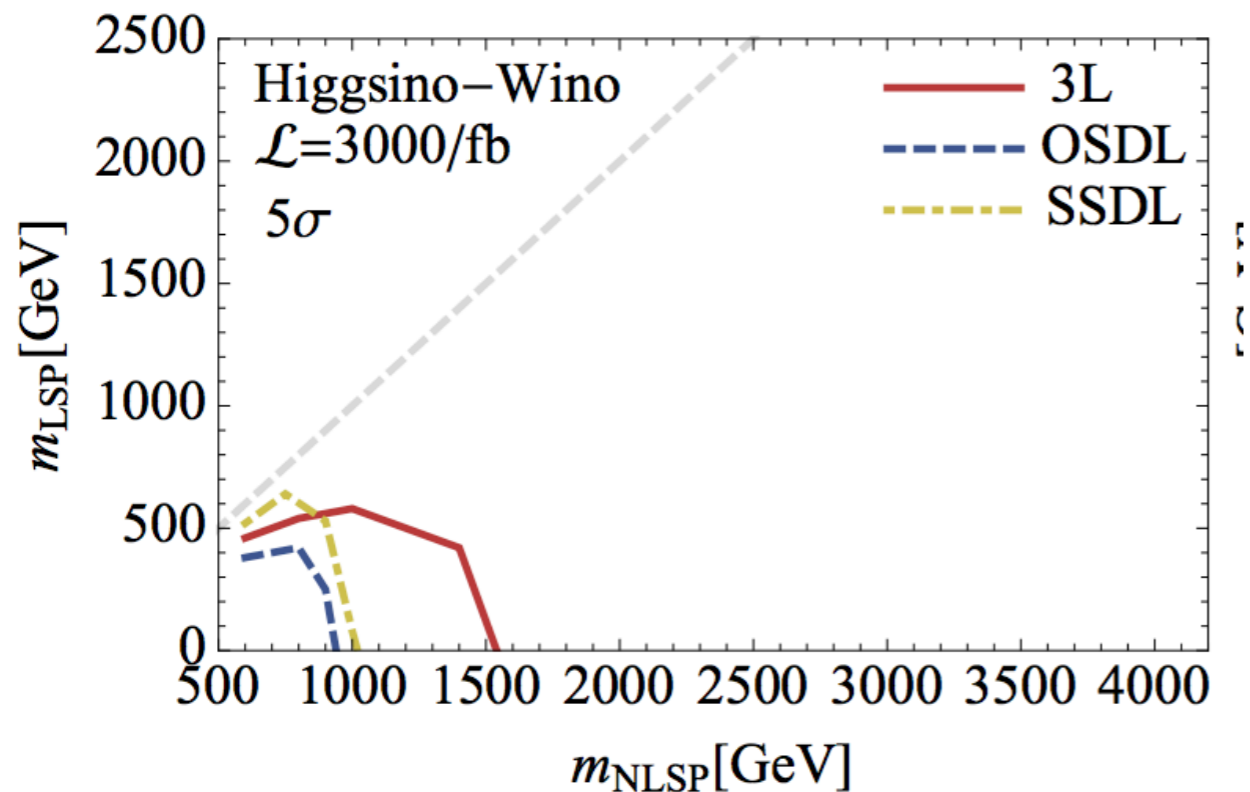
# LHC Inverse Problem with Higgsinos can be resolved.

But the existence of Higgsinos lead to  
the same number of  $h$  and  $Z$ !



# TeV SUSY comes with the electroweak symmetry restored.

making Higgsino simplified and unique,

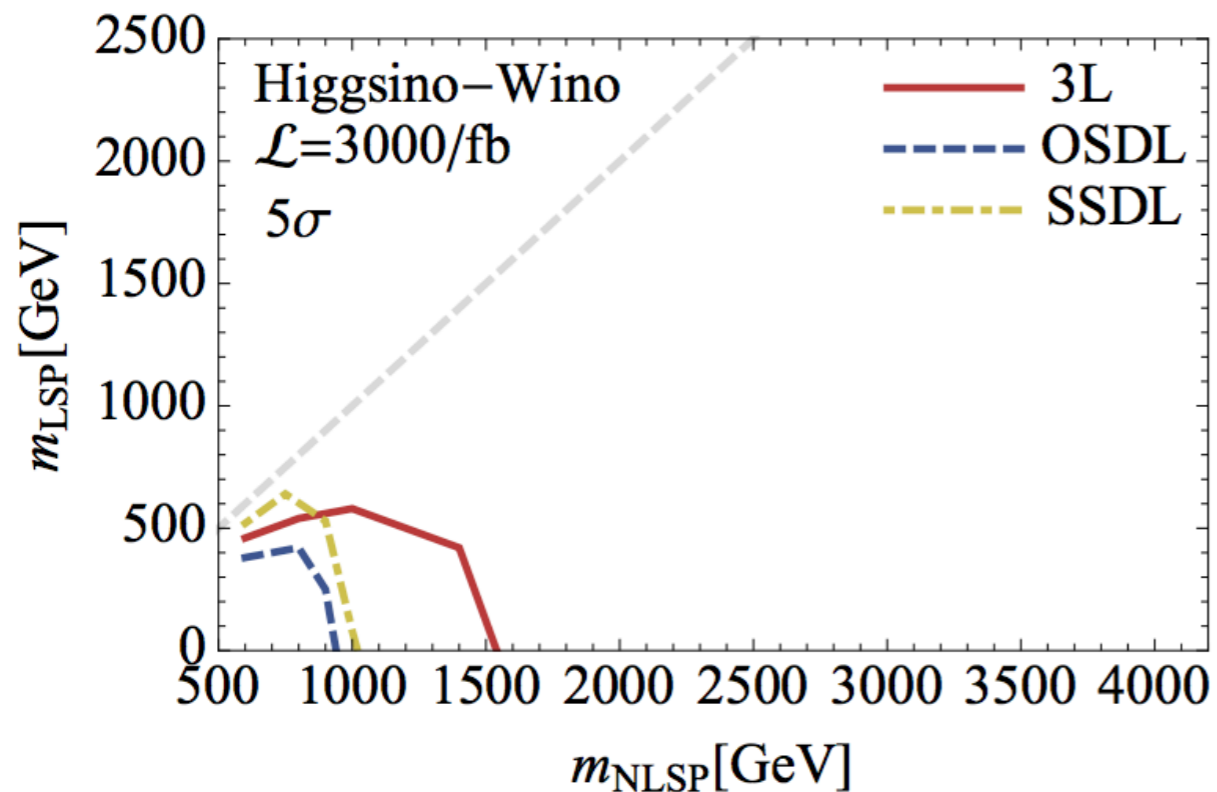


Always,  $\text{BR}(h) = \text{BR}(Z)$ .

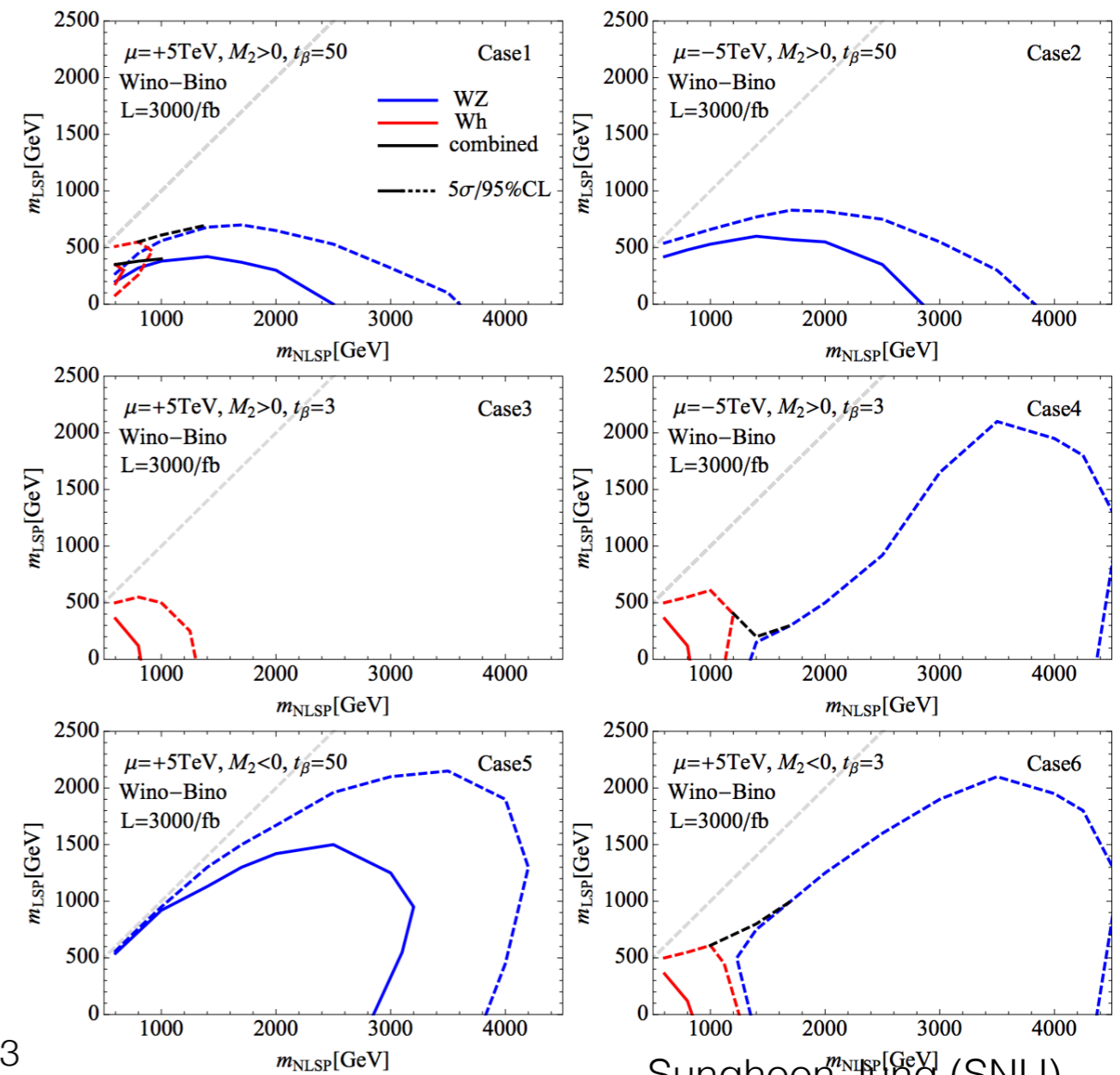
# TeV SUSY comes with the electroweak symmetry (almost) restored.

making Higgsino simplified and unique,

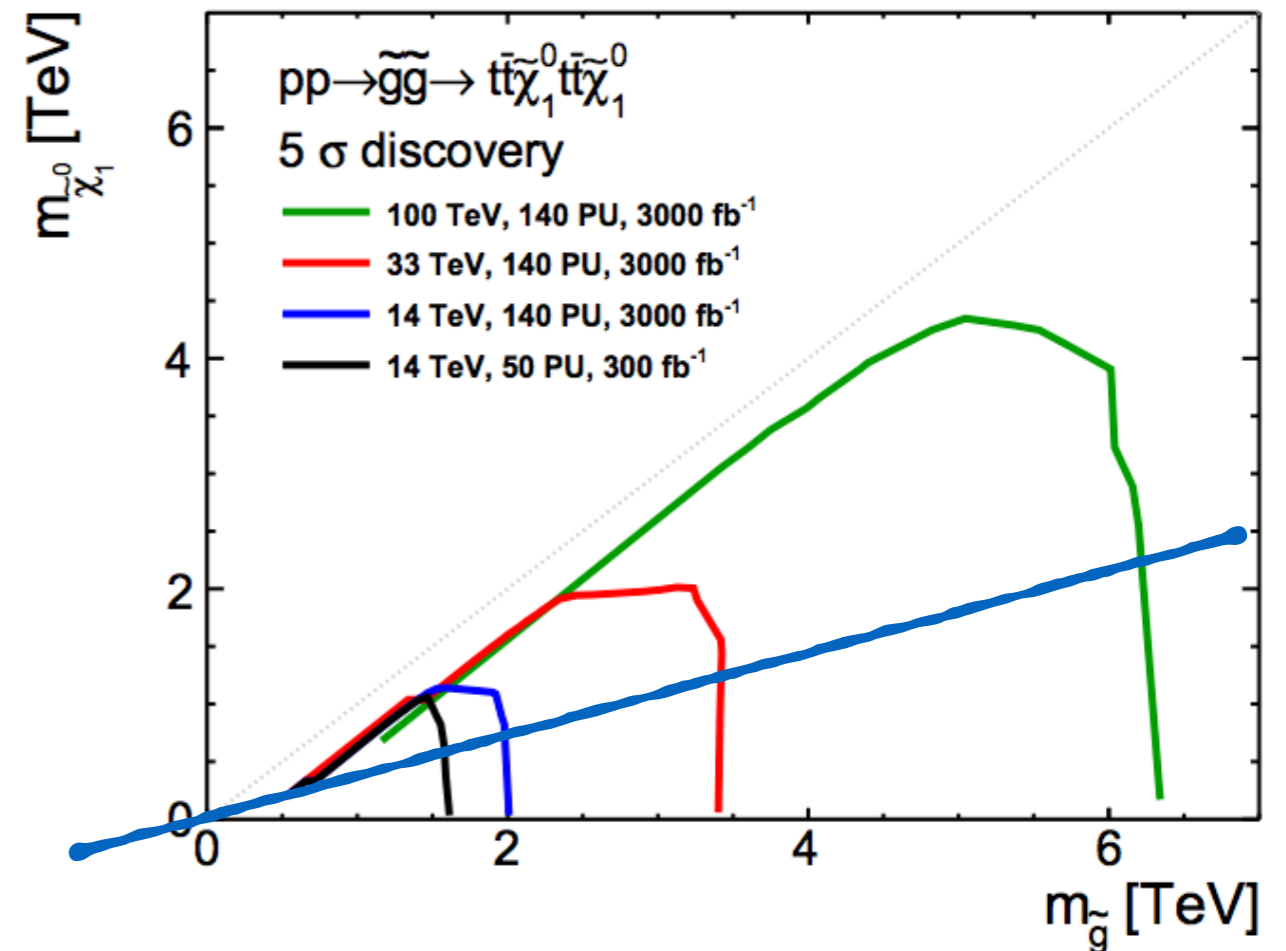
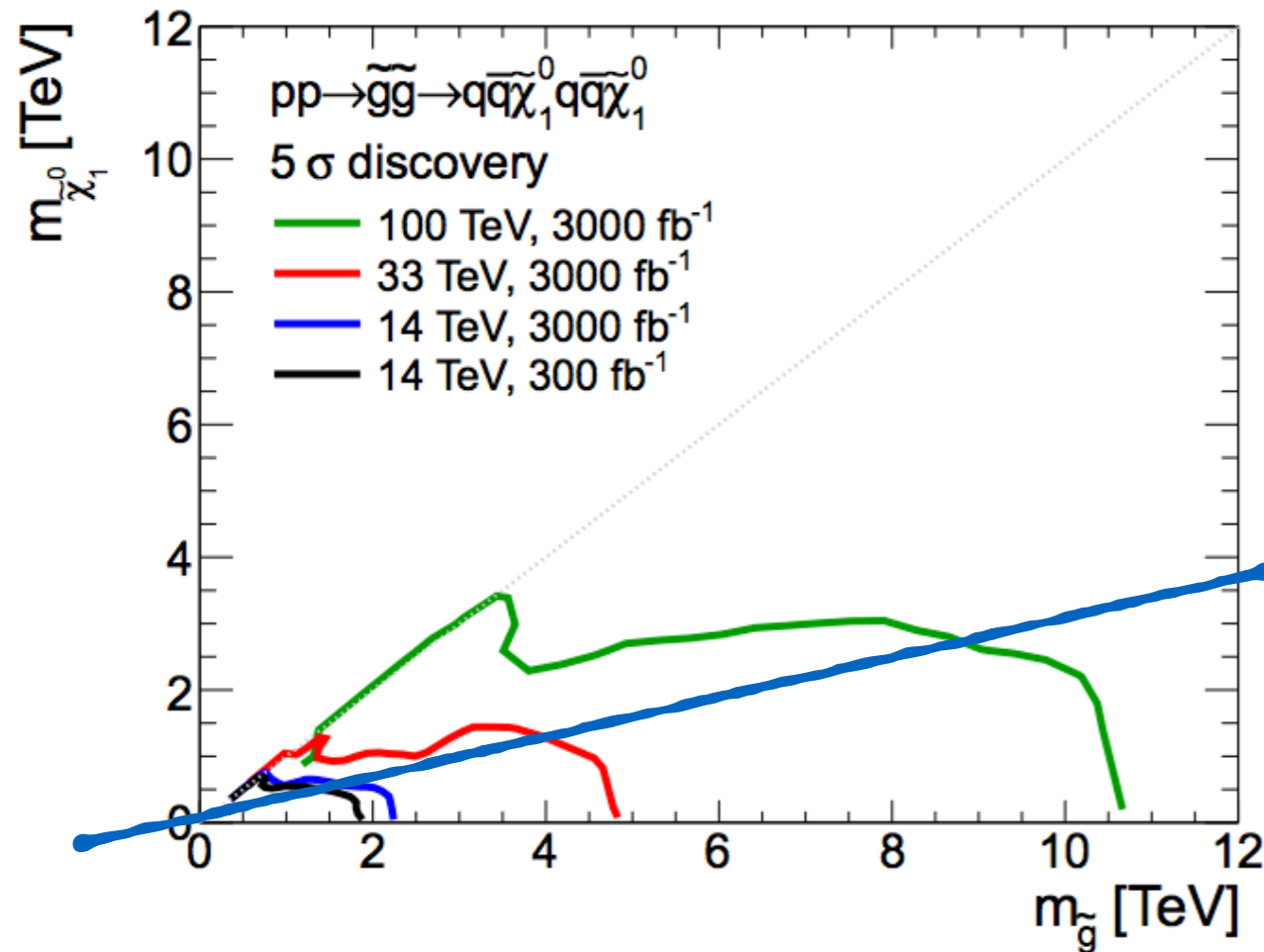
but retaining the trace of its breaking as small mixing.



Always,  $\text{BR}(h) = \text{BR}(Z)$ .



# Back to gauginos, in TeV SUSY, what matters is the $m(\text{gluino})$ .

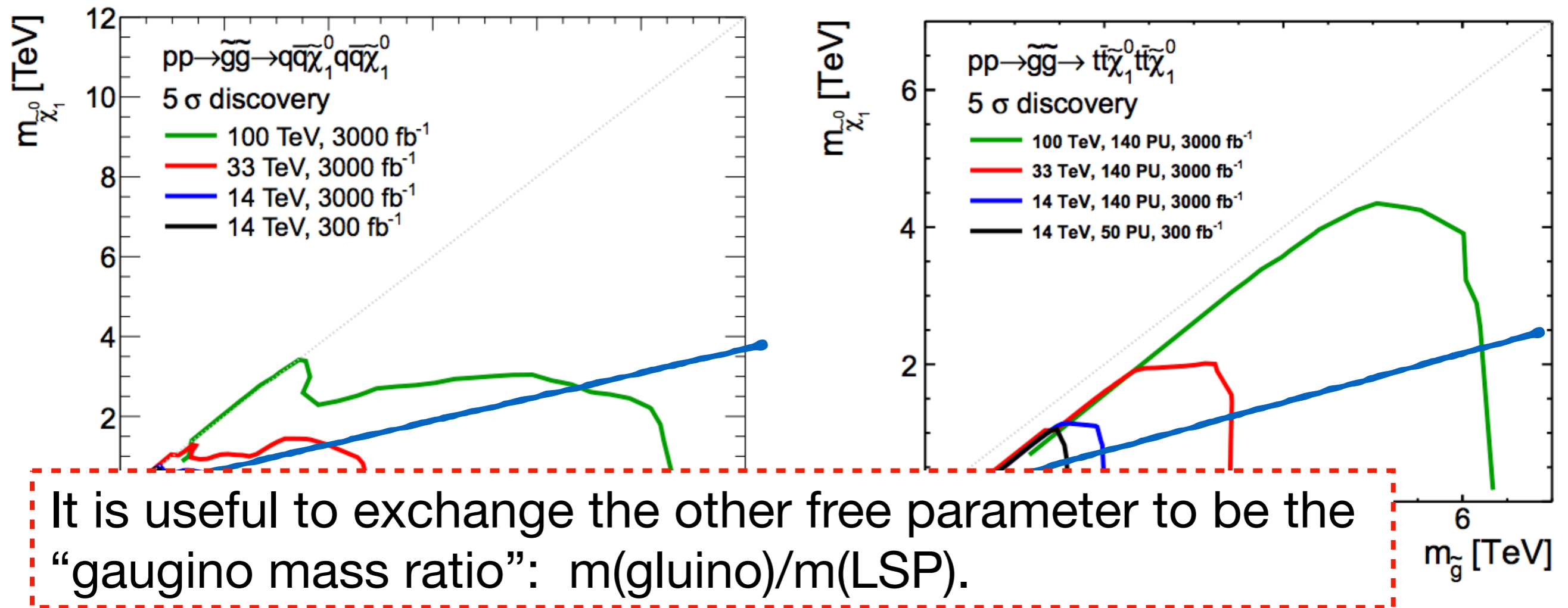


T. Cohen et al.

As long as  $m(\text{gluino}) > (2\sim 3) m(\text{LSP})$ ,  
only  $m(\text{gluino})$  matters in the search kinematics.



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As long as  $m(\text{gluino}) > (2\sim 3) m(\text{LSP})$ ,  
 only  $m(\text{gluino})$  matters in the search kinematics.

# Gaugino code

mSUGRA pattern :  $M_a \propto \frac{\alpha_a}{4\pi} \Lambda$

K.Choi, H.P.Nilles, 0702146

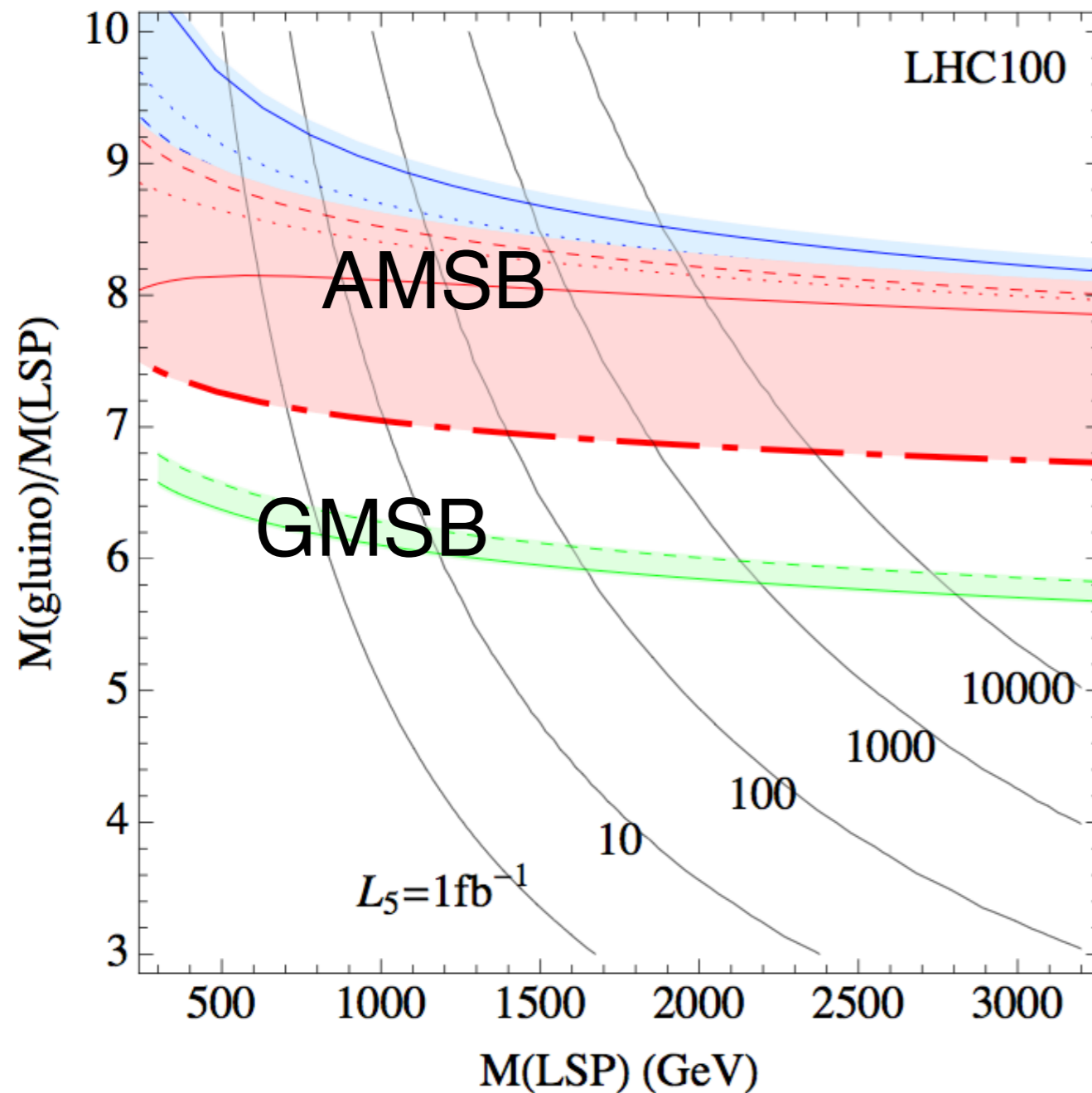
AMSB pattern :  $M_a \propto \frac{b_a \alpha_a}{4\pi} m_{3/2}$

mirage pattern :  $M_a \propto \frac{\alpha_a}{4\pi} \left( b_a + \frac{1}{0.1\alpha} \right) m_{3/2}$

Gaugino mass ratios are characteristic variables of ~~SUSY~~ mediation.

Gauginos are most model-independent fields.

# A useful presentation, probing ~~SUSY~~ mediation

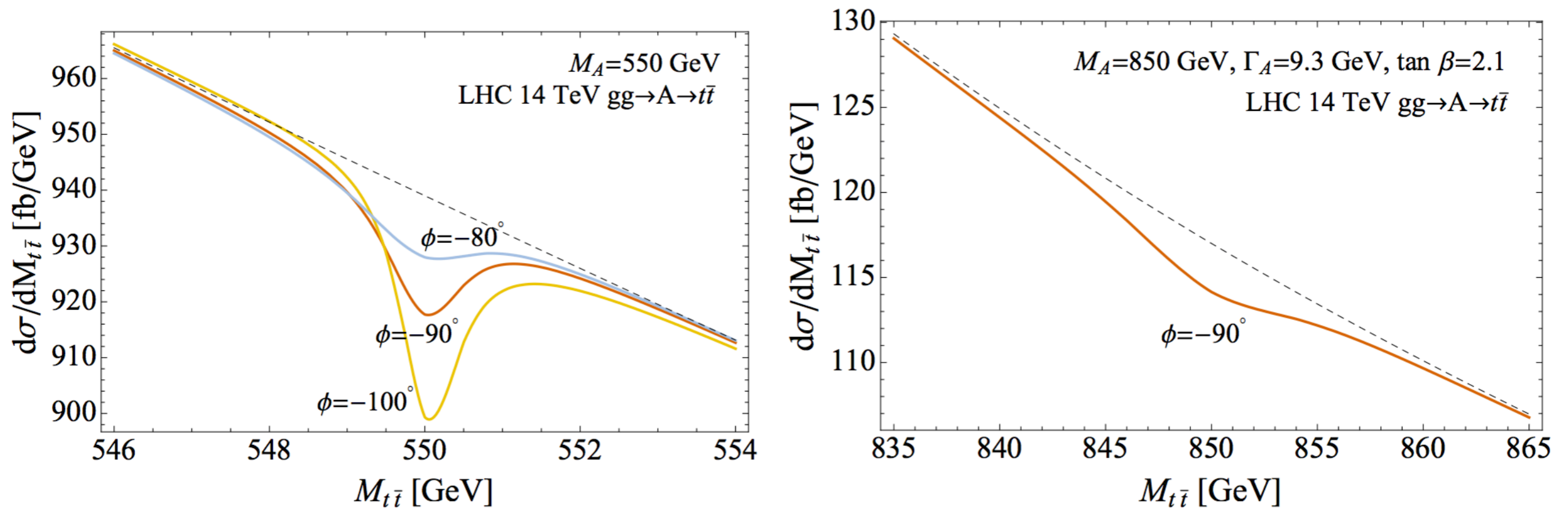


This is a useful way  
to present SUSY search  
results  
(gluino pair in this case).

S.Jung and J.D.Wells, 1312.1802

# In TeV SUSY, heavy Higgs bosons are generically NOT resonance peaks.

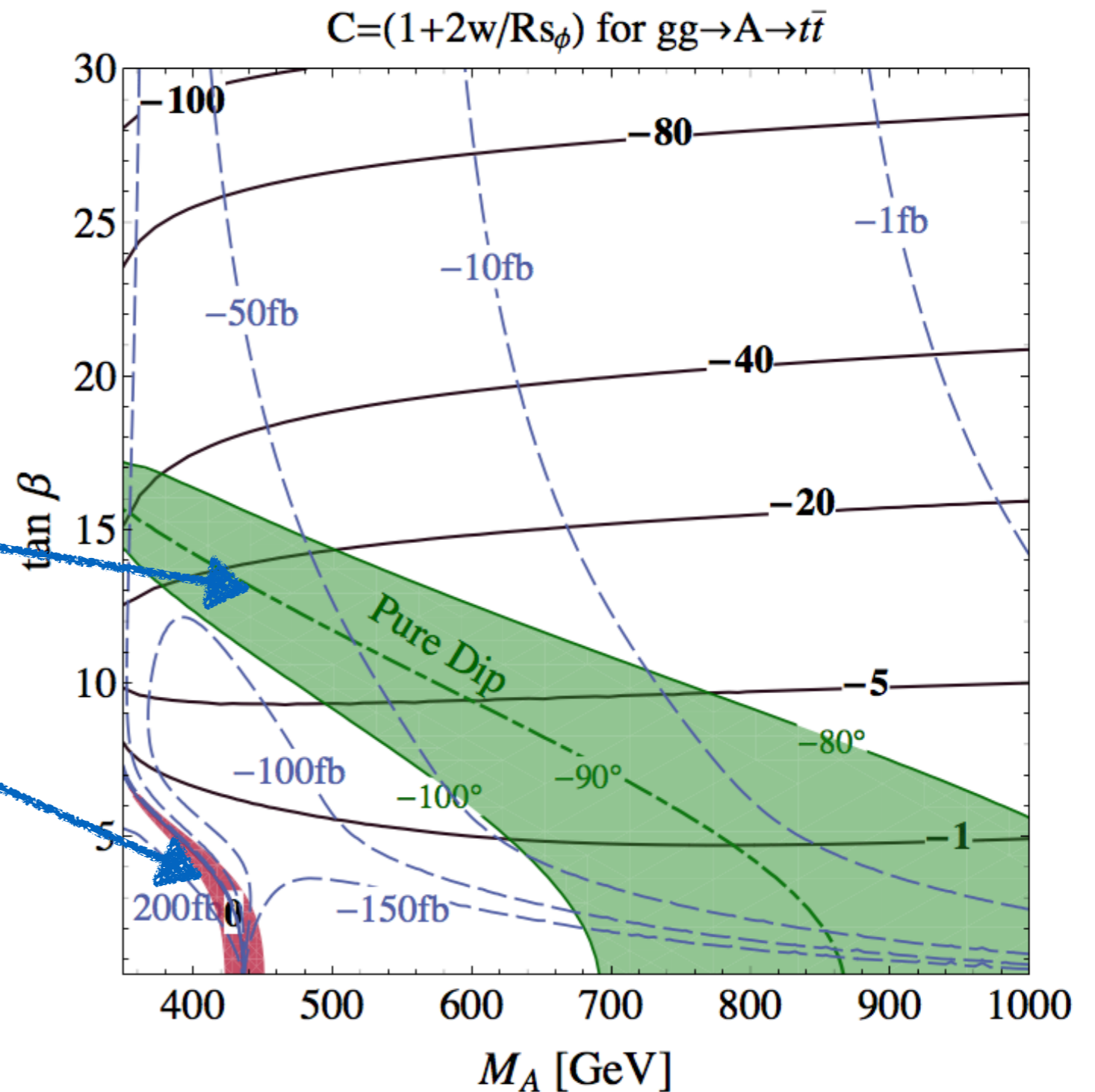
$gg > H, A > t\bar{t}$



Even a pure dip or nothingness is not only possible but generic.

# In TeV SUSY, they are generically NOT resonance peaks.

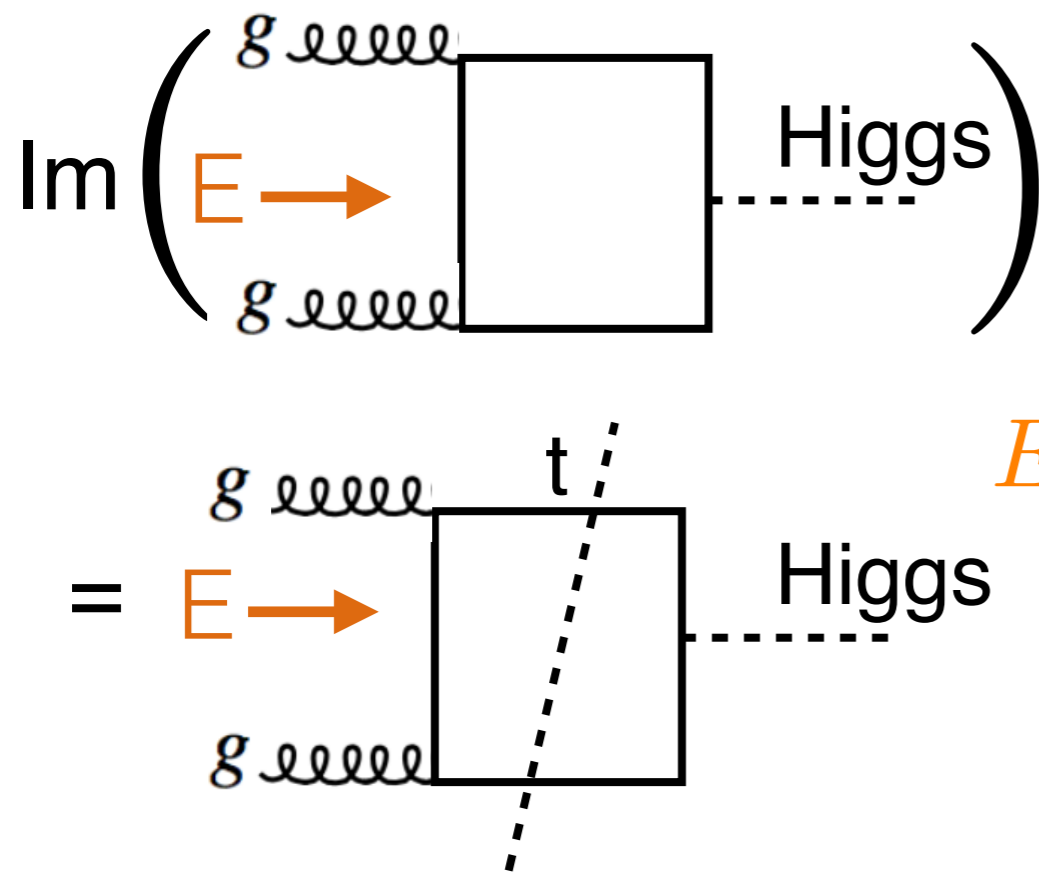
- No pure peak anywhere!
- Pure dip and nothingness regions.
- Generic in energy-frontier future colliders.



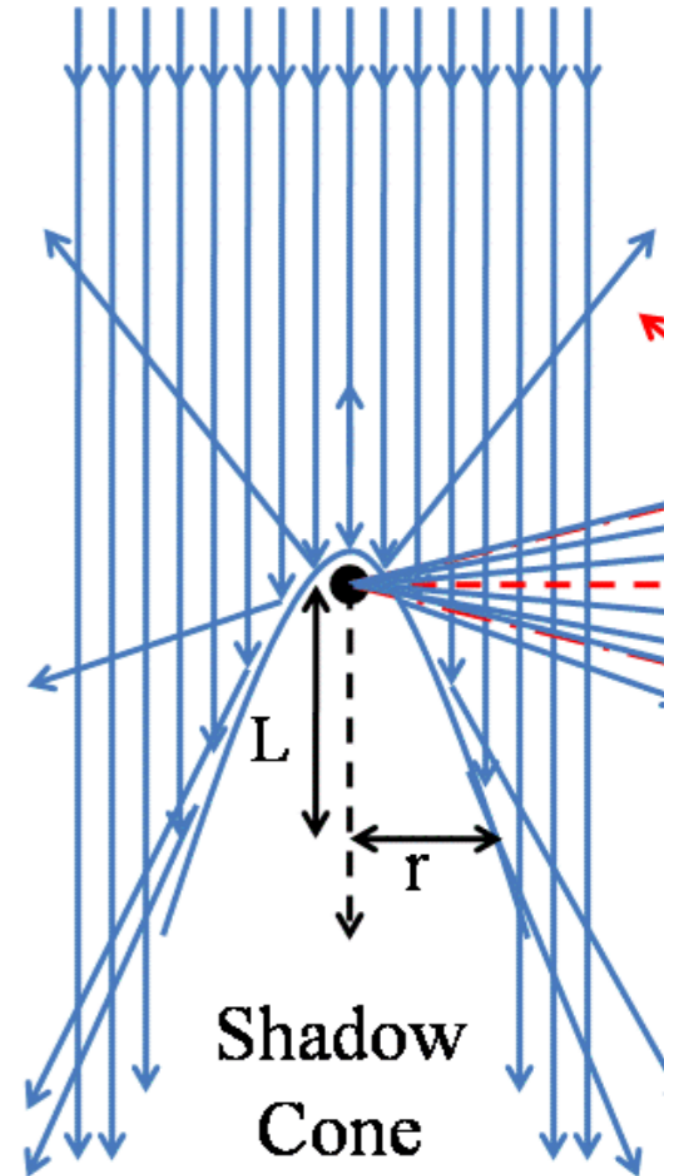
S.Jung, J.Song, Y.W.Yoon, 1505.00291

# Analogous to Shadow scattering

Attenuation of forward-going wave (shadow)  
 = Imaginary part of forward-scatt. amplitude



$$E = m_{\text{Higgs}} > 2m_{\text{top}}$$



# Above all, new search method is needed.

1. Resonance shape is not Breit-Wigner any more.
2. Peak and dip may cancel in the binning of invariant mass.
3. Narrow-width approximation (NWA) is wrong

$\sigma(H/A) \cdot BR(H/A \rightarrow t\bar{t})$  cannot be used anymore!

# Why haven't seen so far?

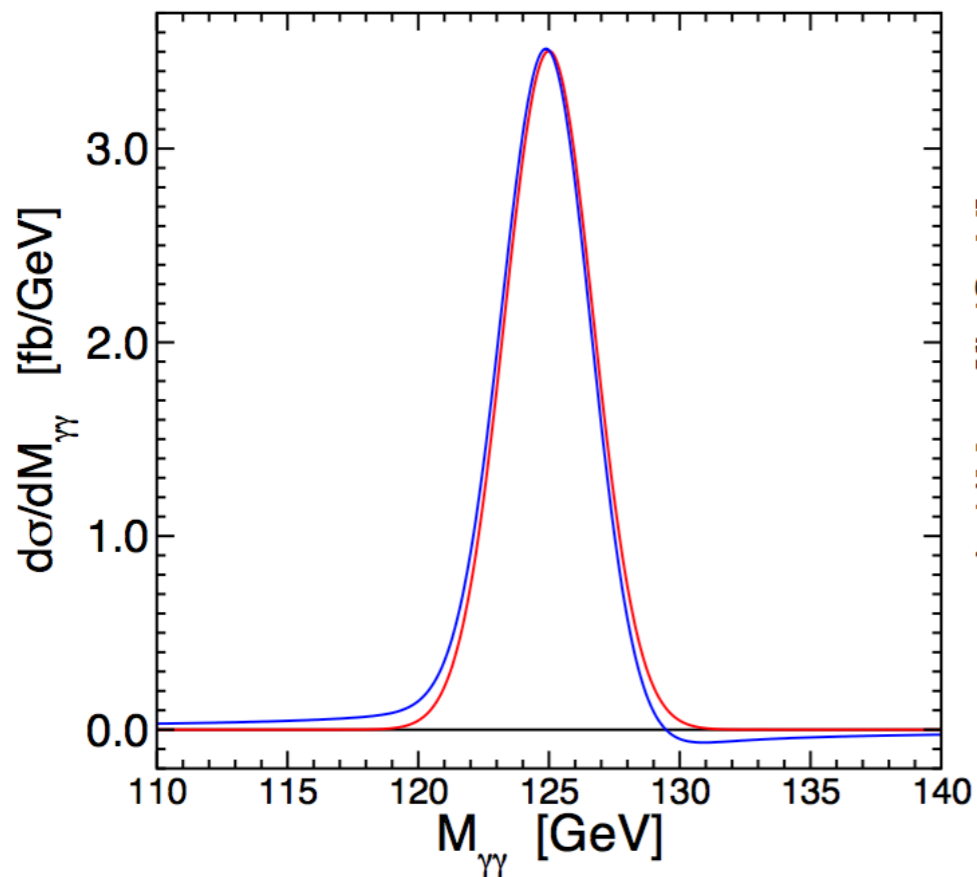
- 1. Narrow resonances:
  - Complex interference is linear in width.
  - Small width from small mass, small number of decay modes.
- 2. Small complex phase:
  - Either tree-level decays or no Cutkosky cuts.

Thus, will be generic for future heavy new particles!



# Milder interference is also important

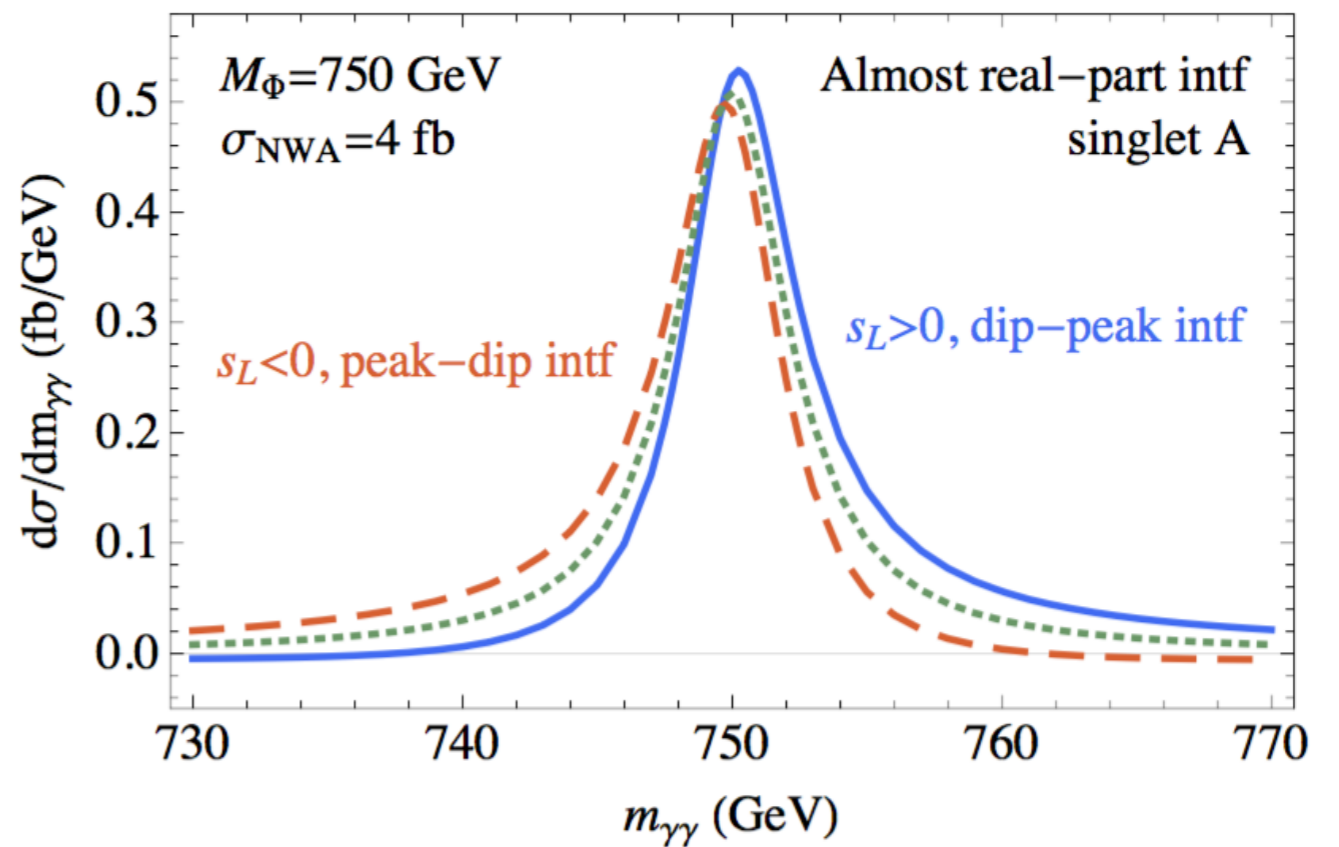
125 GeV SM Higgs



~70 MeV shift to the mass of the (narrow) Higgs.

L.Dixon et al, PRL(2013)

750 GeV diphoton resonance was thought to be *broad*.

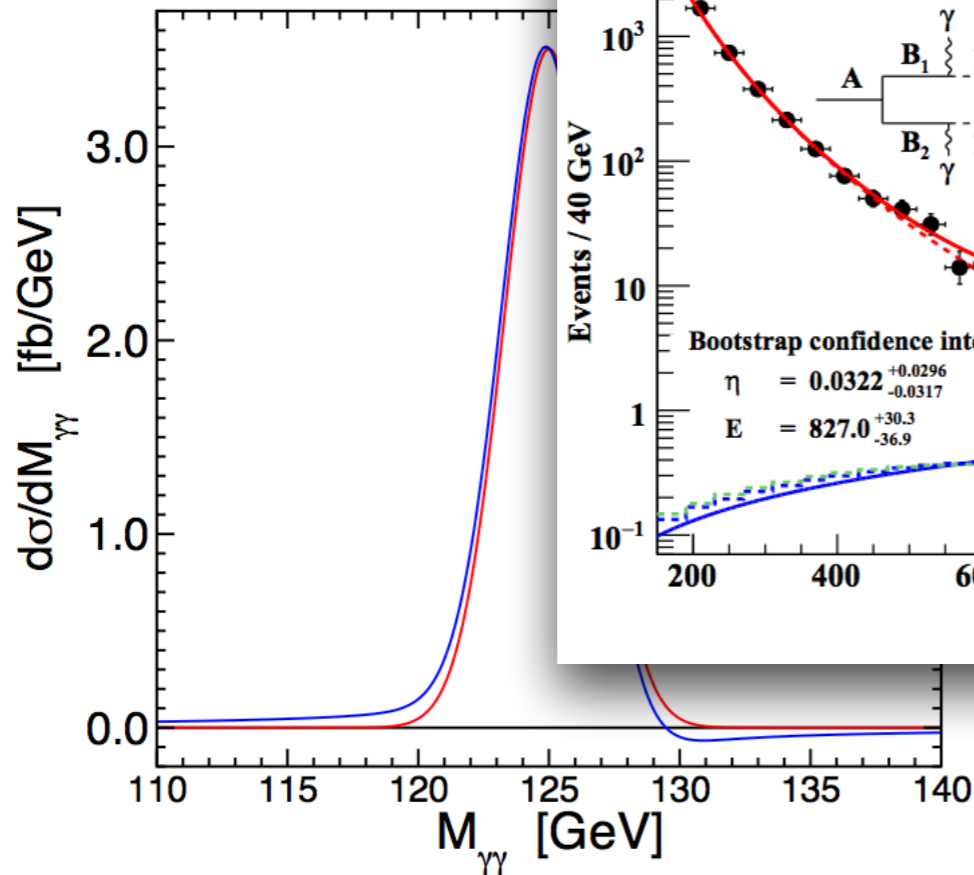


Pole shift and peak enhancement had to be accounted, but none did.

S.Jung, J.Song, Y.W.Yoon, 1601.00006

# Milder interference is also important

125 GeV SM Higgs

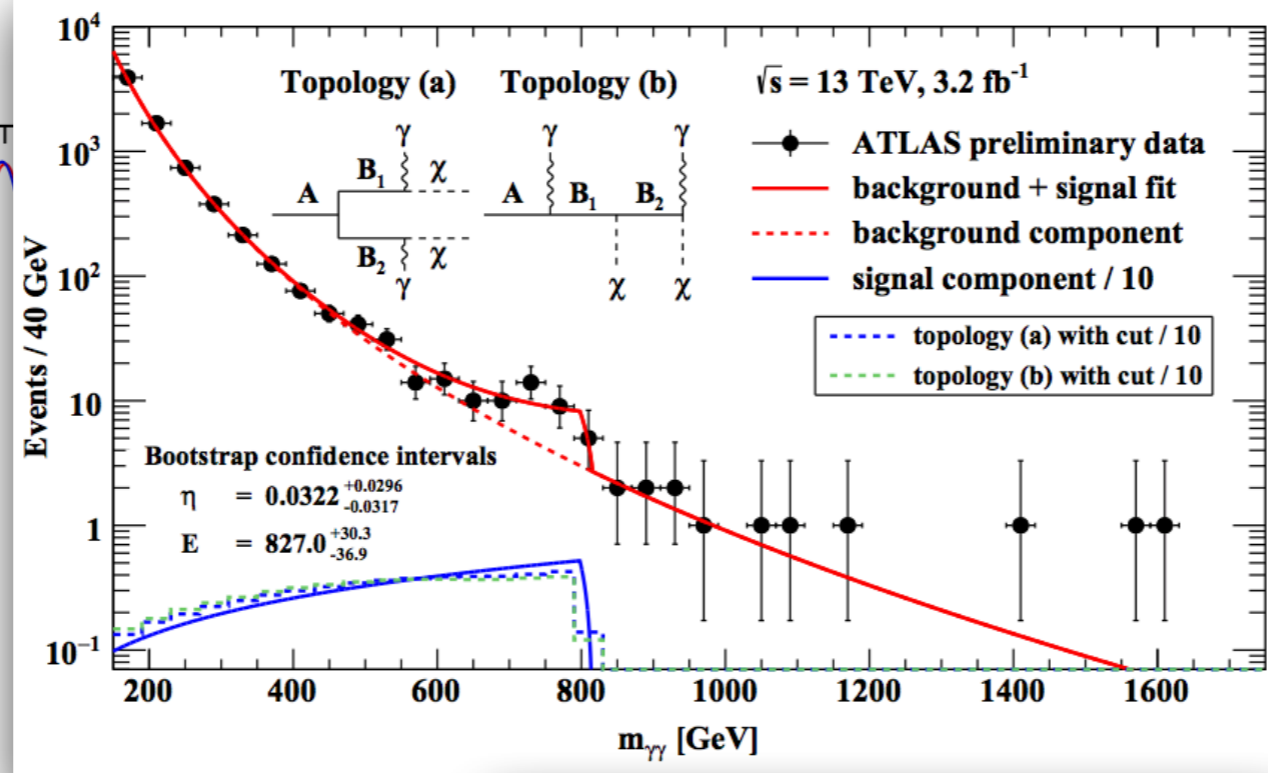


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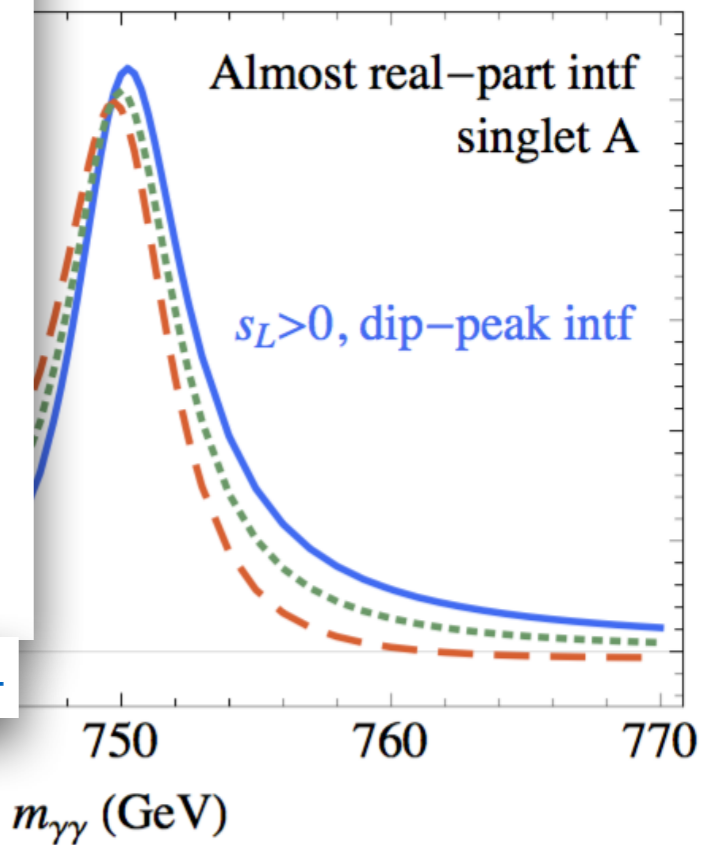
L.Dixon et al, PRL(2013)

750 GeV diphoton resonance

to be *broad*.



W.S.Cho, J.C.Park et al, 1512.06824

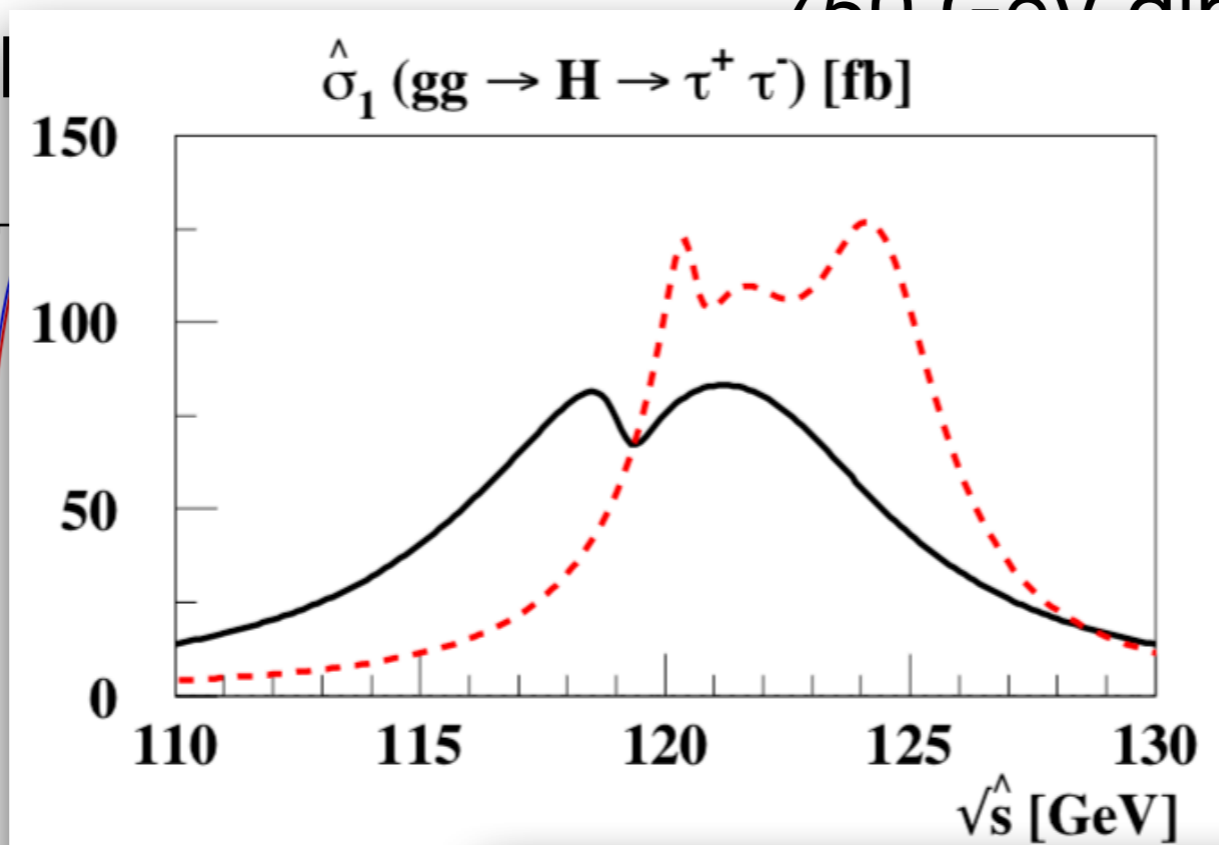
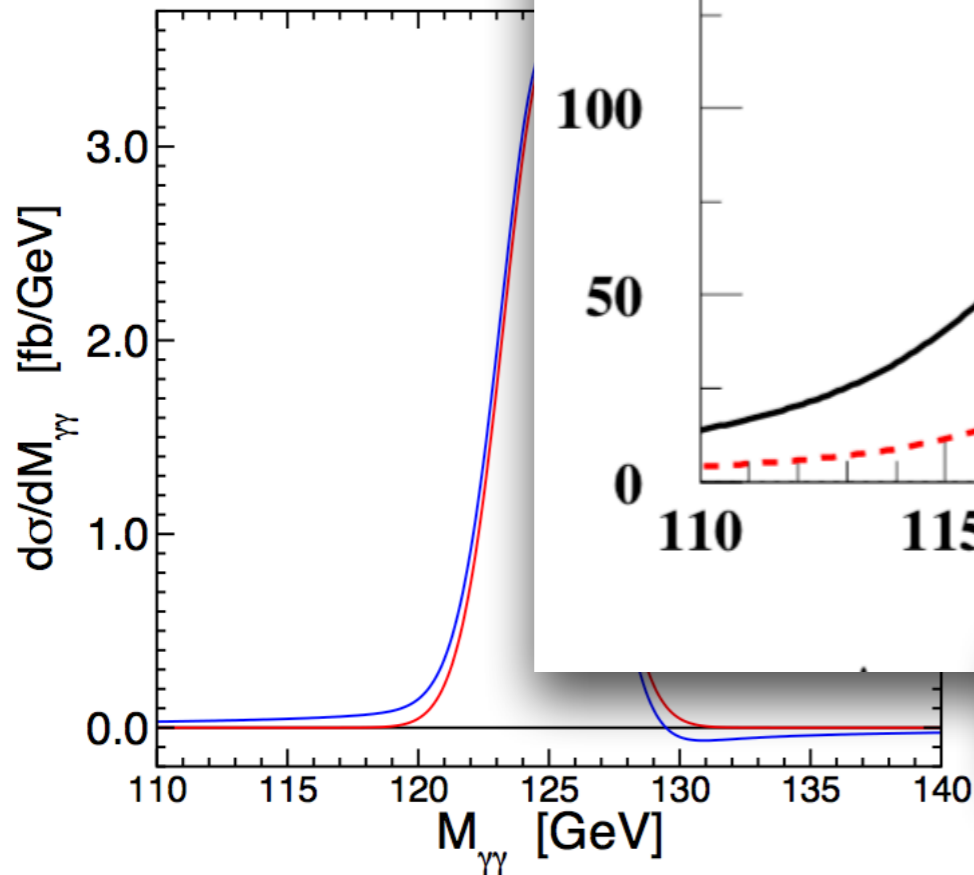


Pole shift and peak enhancement had to be accounted, but none did.

S.Jung, J.Song, Y.W.Yoon, 1601.00006

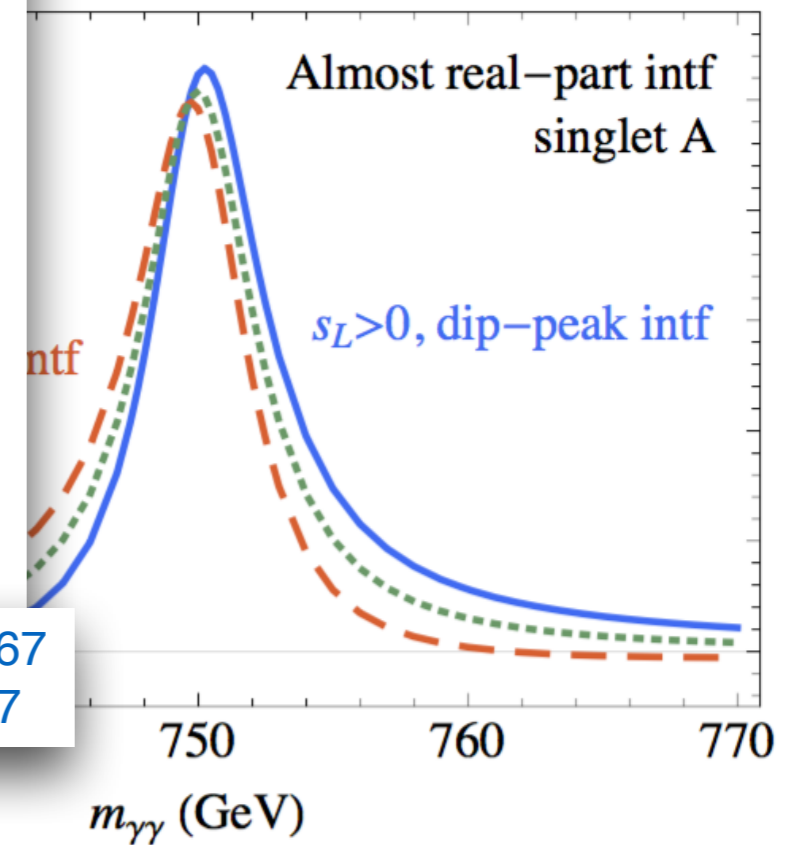
# Milder interference is also important

125 GeV S



J.S.Lee, J.Ellis, A.Pilaftsis, 0404167  
S.Y.Choi, P.Zerwas, et al, 0407347

750 GeV diphoton resonance  
to be *broad*.



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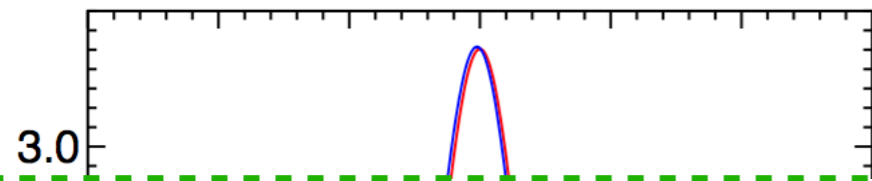
L.Dixon et al, PRL(2013)

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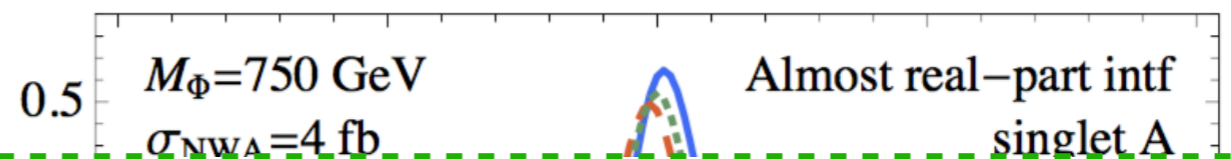
S.Jung, J.Song, Y.W.Yoon, 1601.00006

# Milder interference is also important

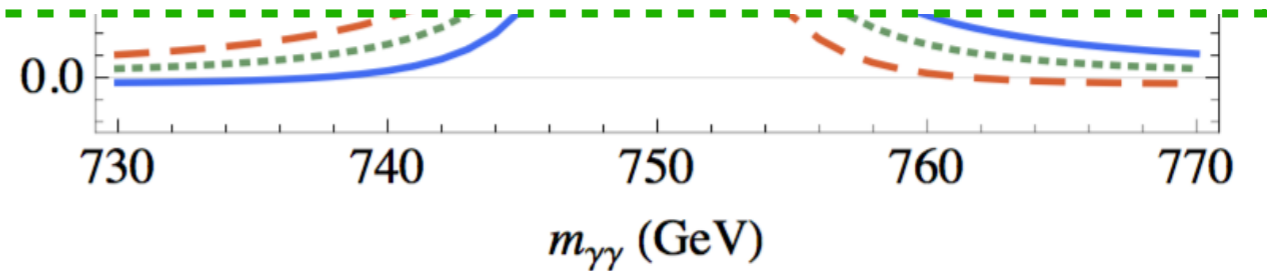
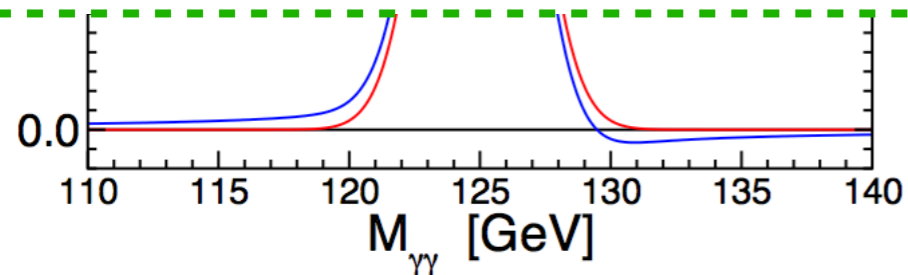
125 GeV SM Higgs



750 GeV diphoton resonance was thought to be *broad*.



Will be more and more relevant in future collider experiments, both for SUSY heavy Higgs and other heavy resonances.



~70 MeV shift to the mass of the (narrow) Higgs.

[L.Dixon et al, PRL\(2013\)](#)

Pole shift and peak enhancement had to be accounted, but none did.

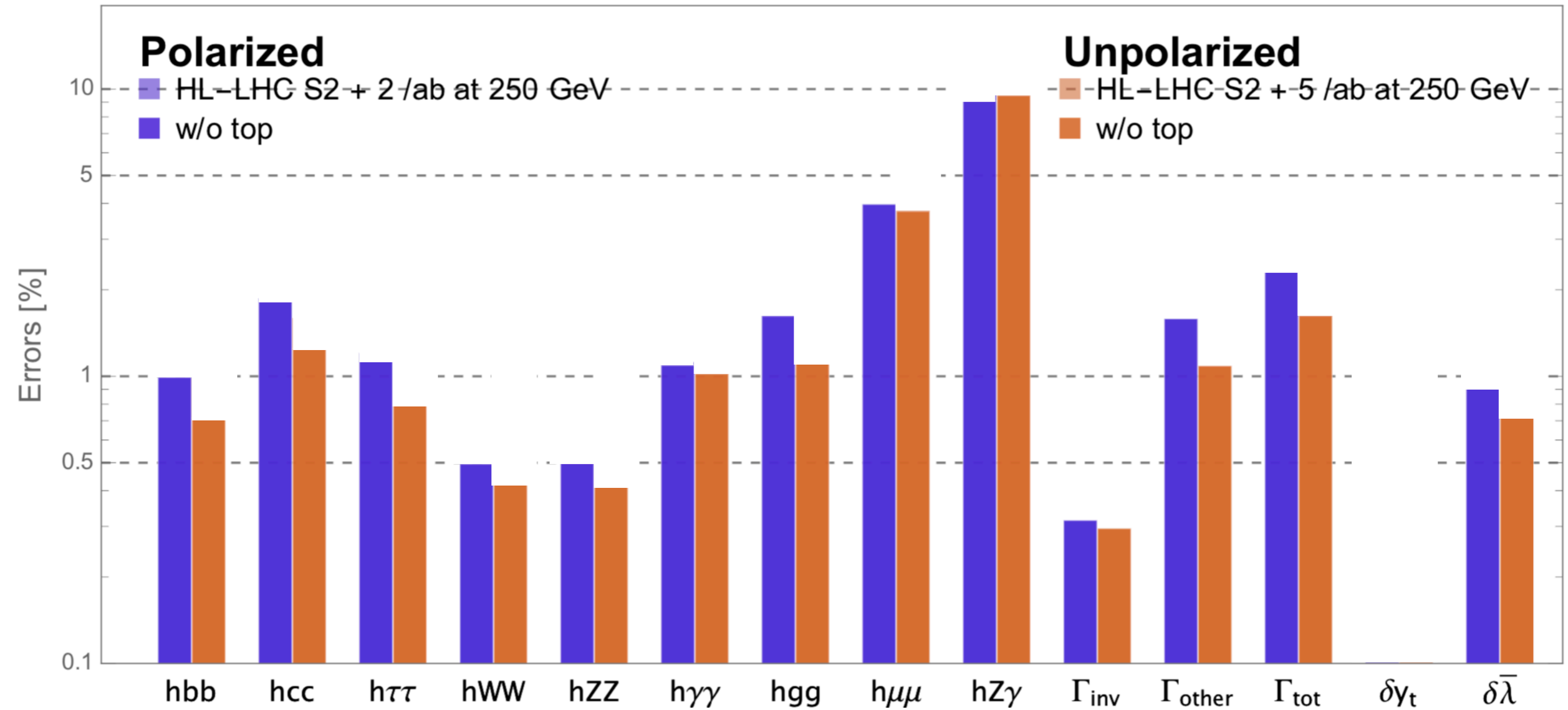
[S.Jung, J.Song, Y.W.Yoon, 1601.00006](#)

# **2. Higgs precision**

**Polarization!**

**Model-independent formalism**

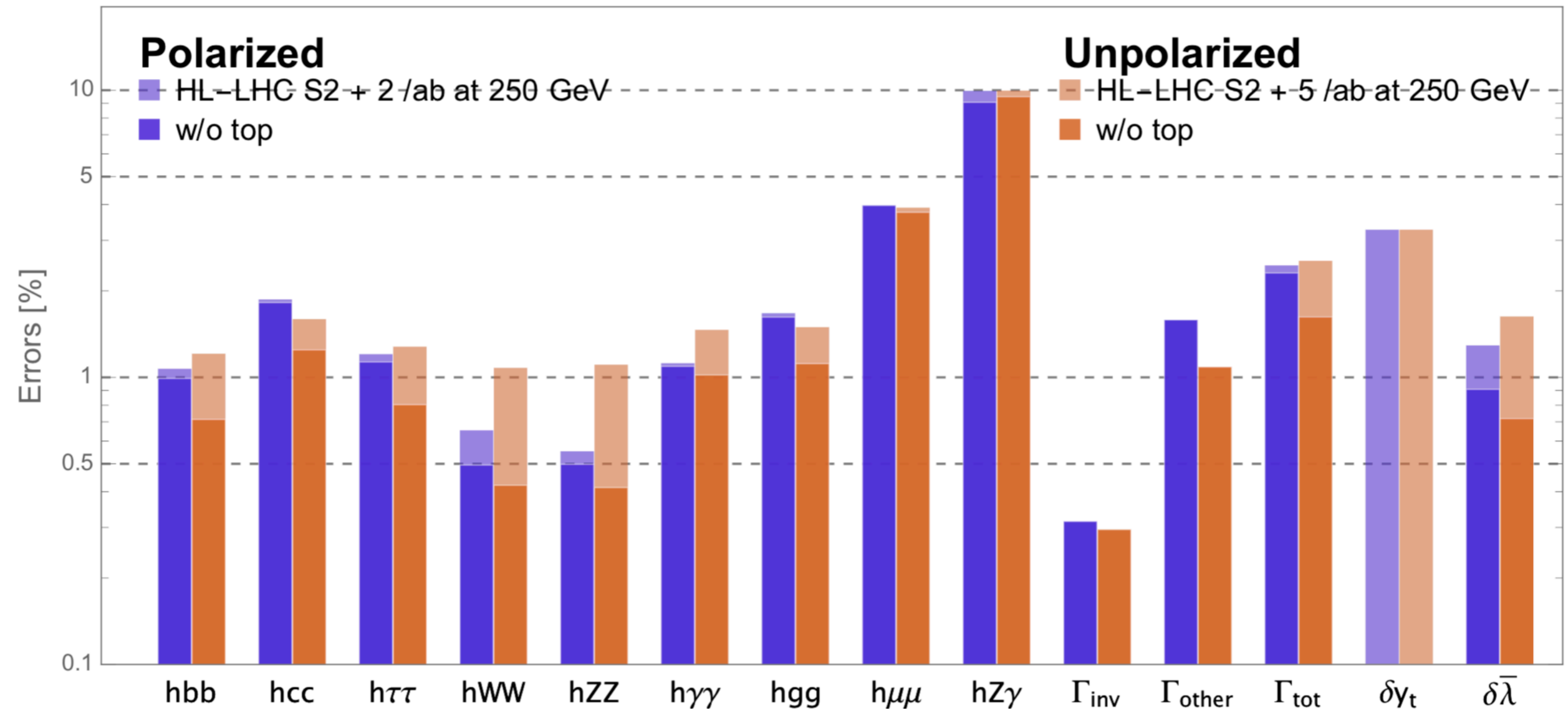
# Higgs coupling precision



ILC polarization is thought to be compensated by FCC-ee/CEPC Z-pole and luminosity.

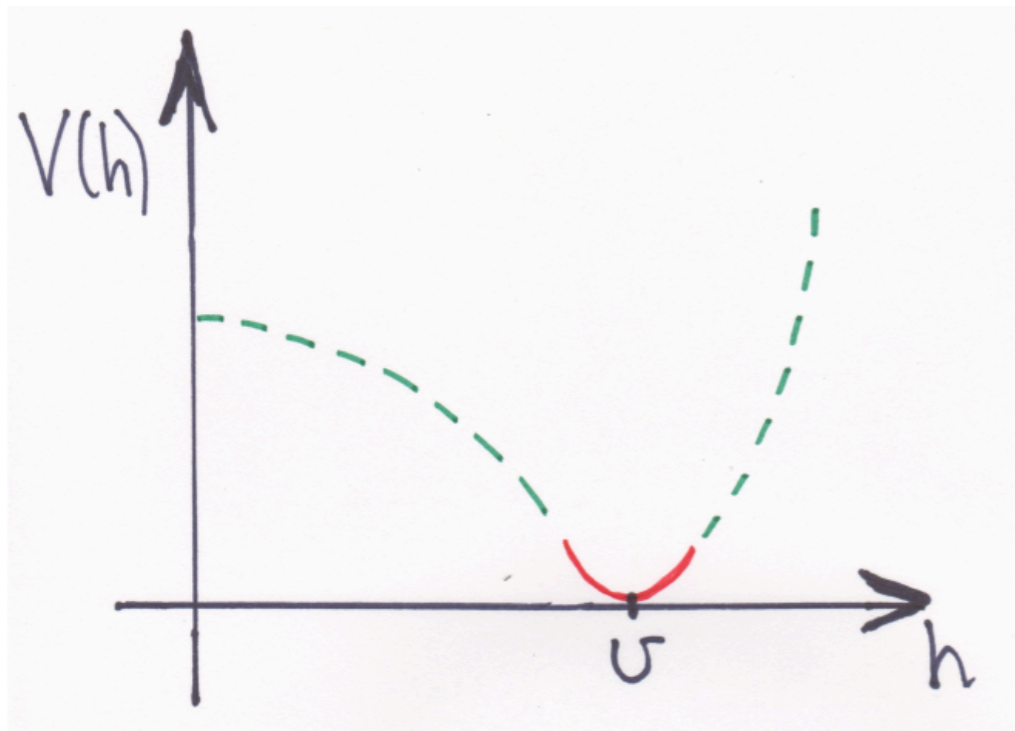
S.Jung, J.Lee, J.Tian, M.Vos, 1909.xxxxx

# Higgs coupling precision



But once top-quark quantum effects are included, polarization is needed to distinguish Higgs vs. top effects.

# Unknown Higgs self coupling & phase transition



by M.Perelstein

- What we know now:

$$V'(h) = 0 @ h = v \approx 250 \text{ GeV}$$

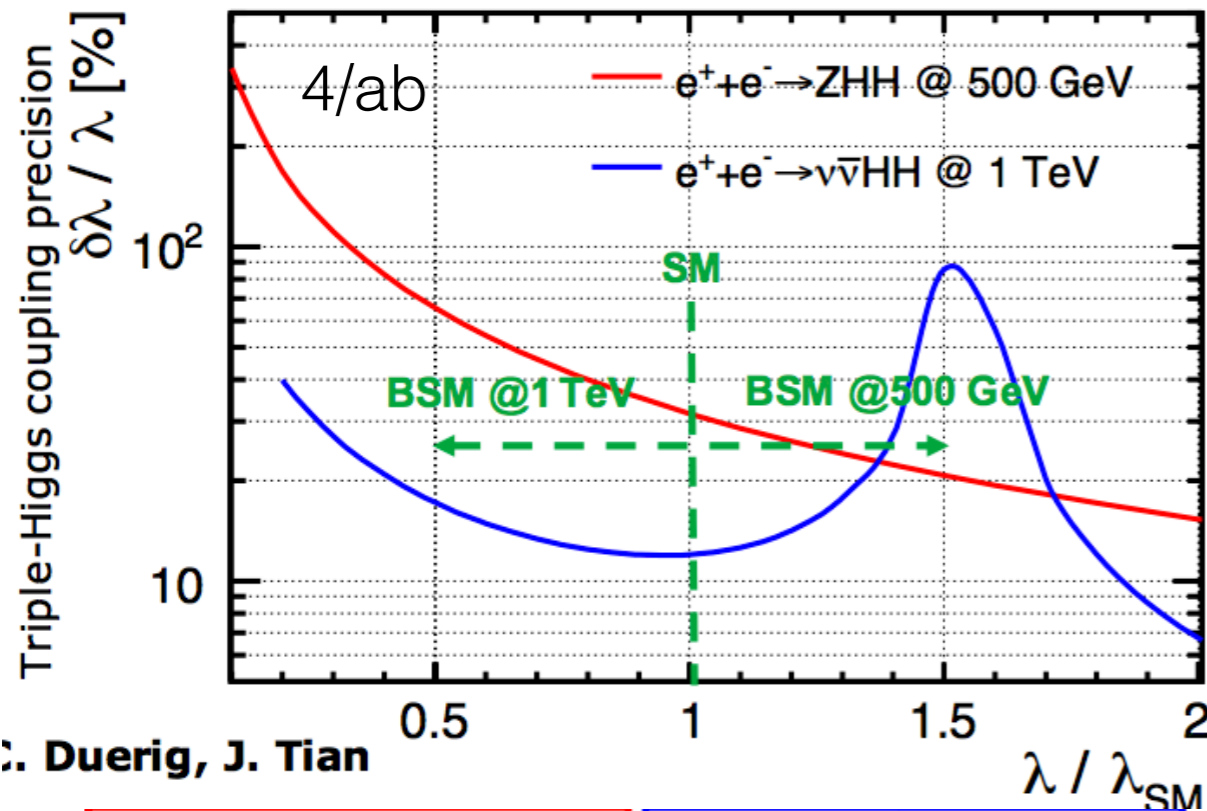
$$m_h^2 = V''(v), m_h \approx 125 \text{ GeV}$$

- Measuring Higgs cubic coupling is the next step in extending our knowledge of the shape of V:

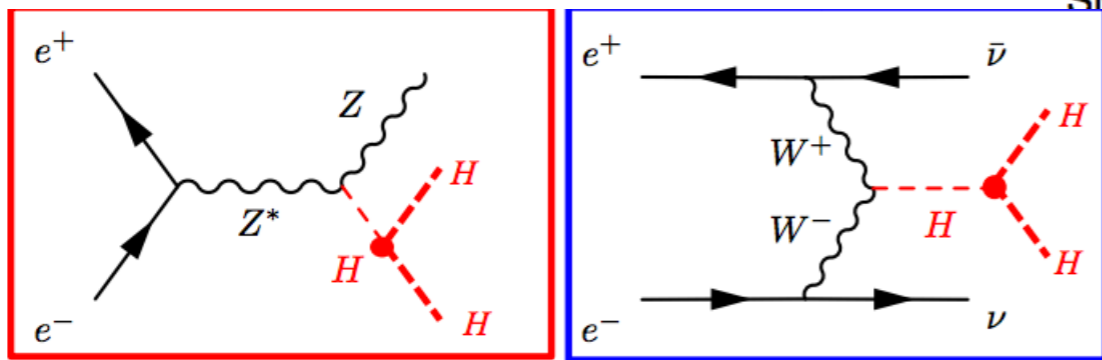
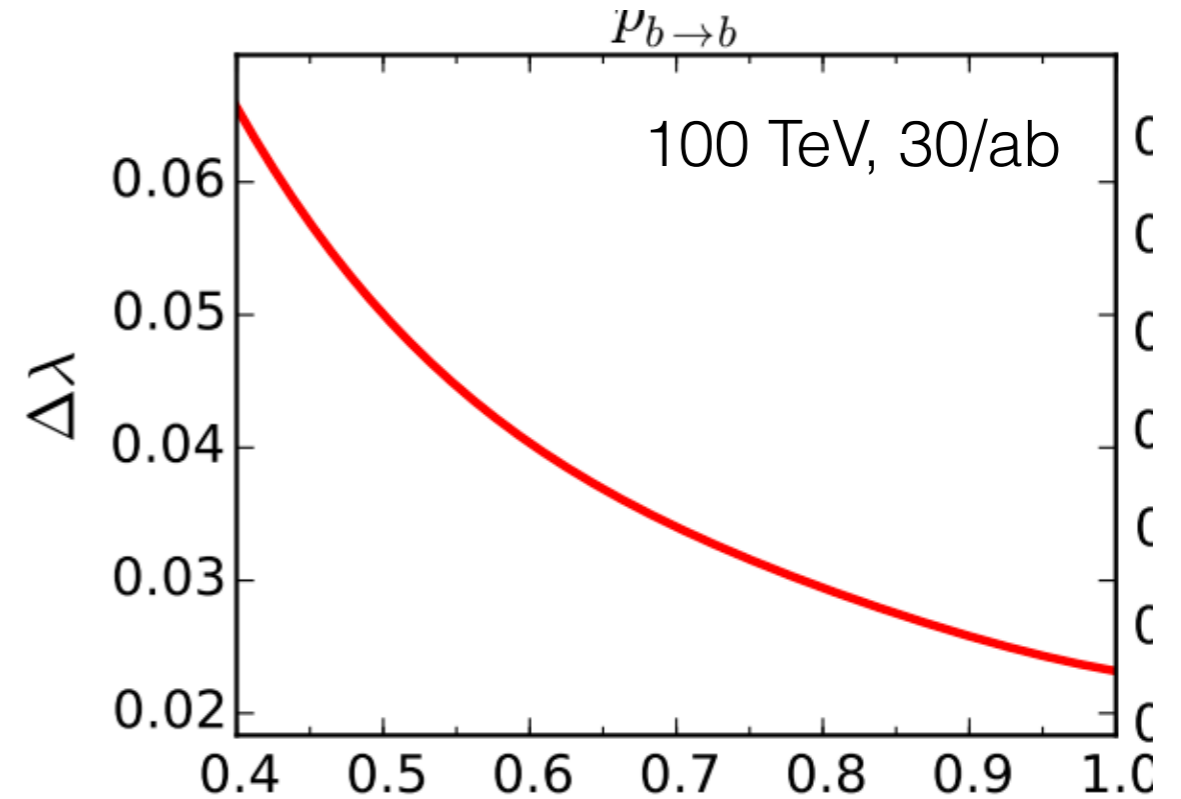
$$\lambda_3 = \frac{1}{6} V'''(v)$$



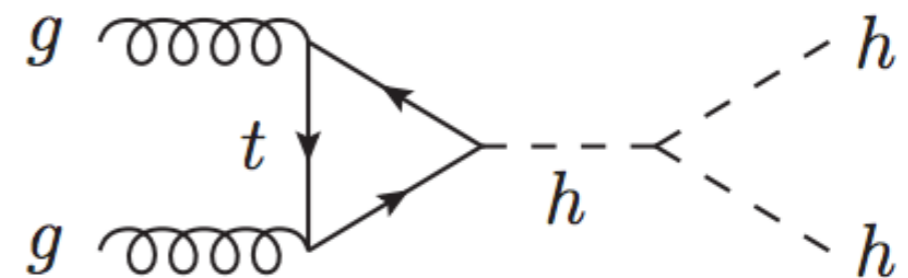
# How we usually think about triple-Higgs measurement



∴ Duerig, J. Tian



Diagrams with triple-Higgs coupling

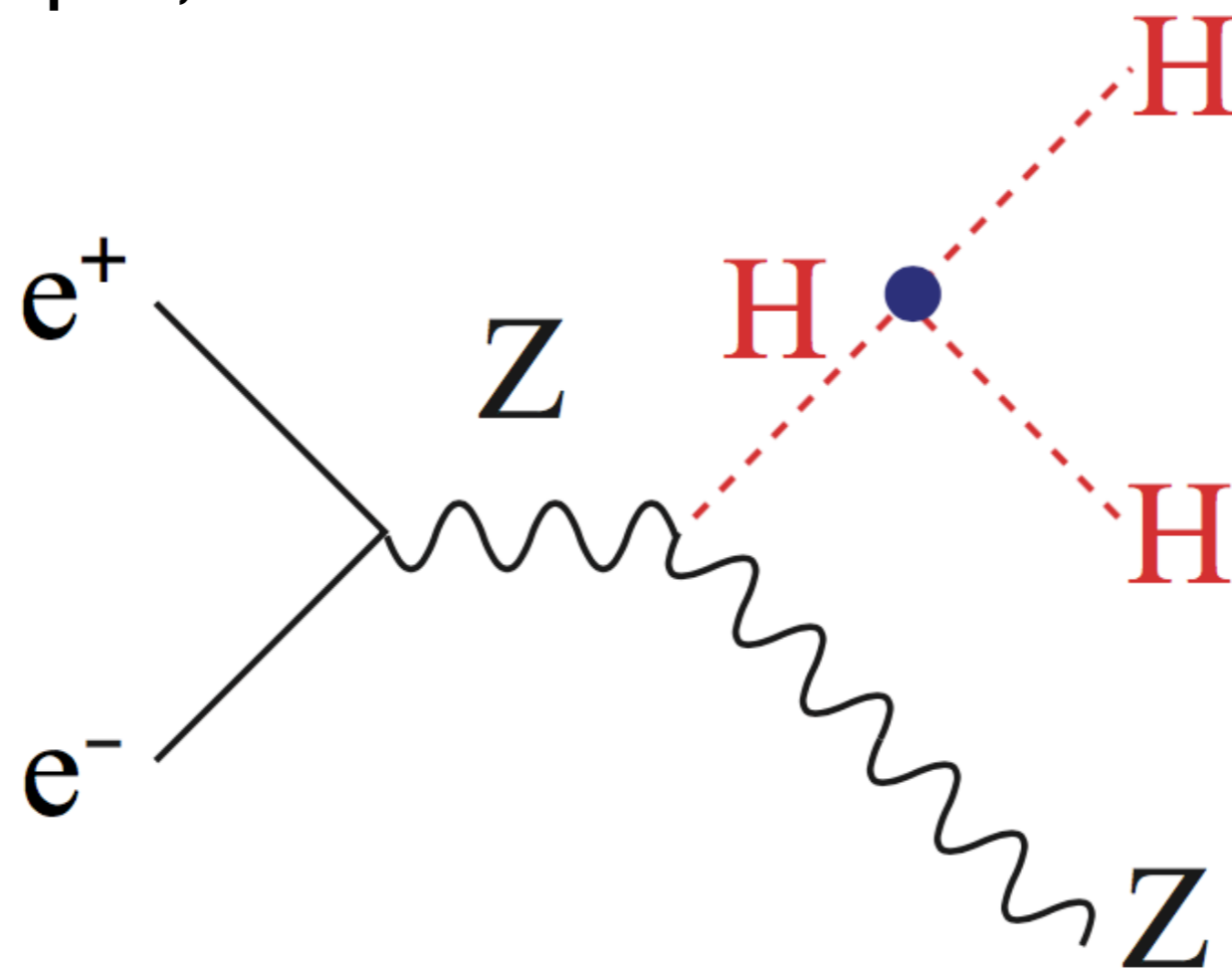


# Extracting triple Higgs

- These results of  $\Delta\lambda$  might be good enough if the only question is to test the SM.
- If there's a deviation, there's a new physics! Not only  $\lambda$ , but many others will be non-SM.
- How do we separate the desired deviations in the Higgs triple coupling from those of others? We don't even know what are 'others'.

# How shall we extract Higgs potential?

For example,



# HEFT as a model-independent framework

- In the HEFT, the deviation of the Higgs potential (triple Higgs coupling in particular) is associated with

$$\Delta\mathcal{L} = -\frac{c_6\lambda}{v^2}|\Phi^\dagger\Phi|^3$$

# 10 d=6 operators

$$\begin{aligned}
 \Delta\mathcal{L} = & \frac{c_H}{2v^2} \partial^\mu (\Phi^\dagger \Phi) \partial_\mu (\Phi^\dagger \Phi) + \frac{c_T}{2v^2} (\Phi^\dagger \overleftrightarrow{D}^\mu \Phi) (\Phi^\dagger \overleftrightarrow{D}_\mu \Phi) - \frac{c_6 \lambda}{v^2} (\Phi^\dagger \Phi)^3 \\
 & + \frac{g^2 c_{WW}}{m_W^2} \Phi^\dagger \Phi W_{\mu\nu}^a W^{a\mu\nu} + \frac{4gg' c_{WB}}{m_W^2} \Phi^\dagger t^a \Phi W_{\mu\nu}^a B^{\mu\nu} \\
 & + \frac{g'^2 c_{BB}}{m_W^2} \Phi^\dagger \Phi B_{\mu\nu} B^{\mu\nu} + \frac{g^3 c_{3W}}{m_W^2} \epsilon_{abc} W_{\mu\nu}^a W^{b\nu\rho} W^{c\rho\mu} \\
 & + i \frac{c_{HL}}{v^2} (\Phi^\dagger \overleftrightarrow{D}^\mu \Phi) (\bar{L} \gamma_\mu L) + 4i \frac{c'_{HL}}{v^2} (\Phi^\dagger t^a \overleftrightarrow{D}^\mu \Phi) (\bar{L} \gamma_\mu t^a L) \\
 & + i \frac{c_{HE}}{v^2} (\Phi^\dagger \overleftrightarrow{D}^\mu \Phi) (\bar{e} \gamma_\mu e) .
 \end{aligned}$$

These 10 HEFT ops consist of:

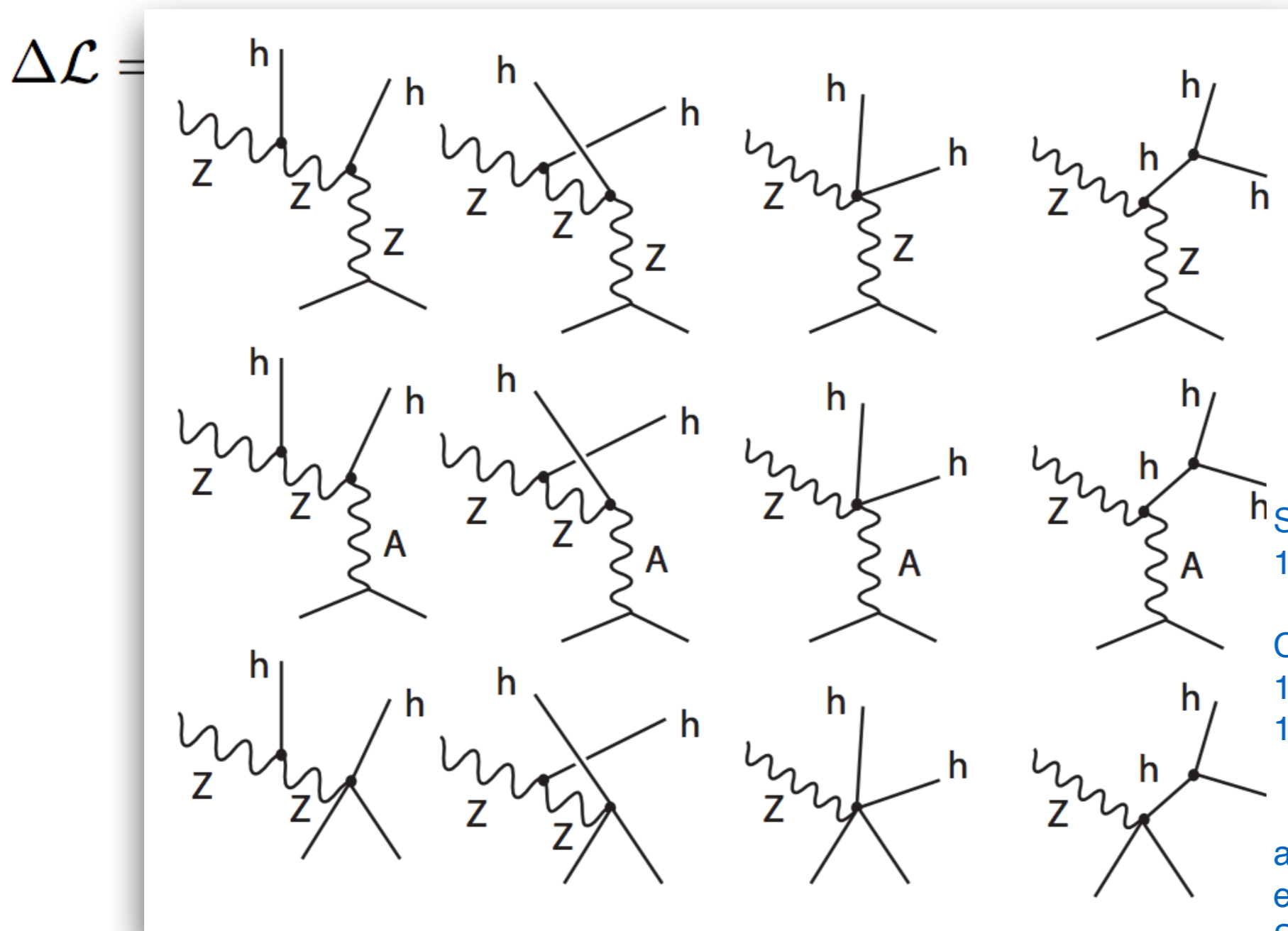
- (1) at least one Higgs or EW gauge,
- (2) only Higgs, EW gauge and electrons

# All 10 ops contribute!

$$\Delta\mathcal{L} = \frac{c_H}{2v^2} \left[ \begin{array}{l} + g^2 (Z_\mu Z^\mu)^2 \\ + g^2 (Z_\mu A^\mu)^2 \\ + i c_{\text{EW}} (Z_\mu A^\mu)^2 \end{array} \right] + \frac{c_6 \lambda}{v^2} (\Phi^\dagger \Phi)^3 + \dots$$

- (1) at least one Higgs or EW gauge,
- (2) only Higgs, EW gauge and electrons

# This was the state-of-the art since 2yrs ago.



$$\mathcal{L}(\Phi) = -\frac{c_6\lambda}{v^2}(\Phi^\dagger\Phi)^3$$

$B^{\mu\nu}$

$\epsilon^{\rho\mu}$

S.Jung, M.Peskin, J.Tian, et al.  
1708.09070, 1708.08912

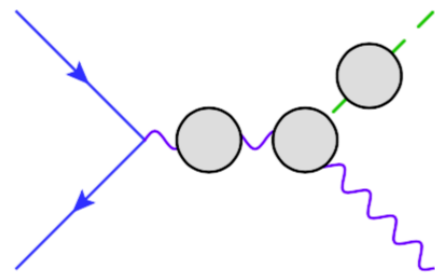
C.Grojean, J.Gu, G.Durieux, et al.  
1704.02333, 1711.03978,  
1704.01953, 1907.04311

as well as earlier, later works by  
e.g. J.S.Lee, M.H.Park, P.Ko,  
S.Y.Choi ...

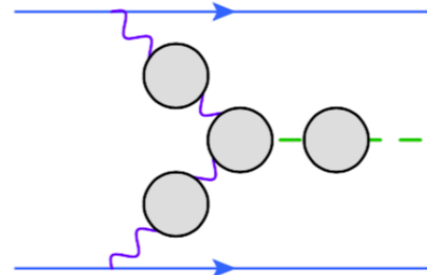
(2) only Higgs, EW gauge and elec

# Top quark quantum effects

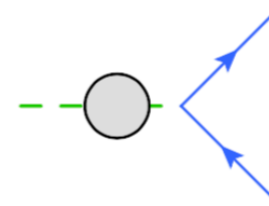
- At one-loop, top quarks also contribute.
- Small?  
Absolutely can be larger than the desired Higgs precision!



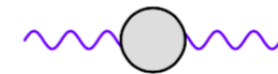
WH,ZH



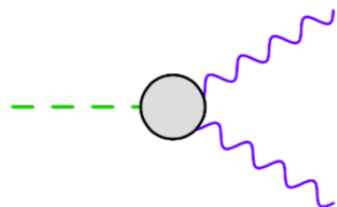
VBF



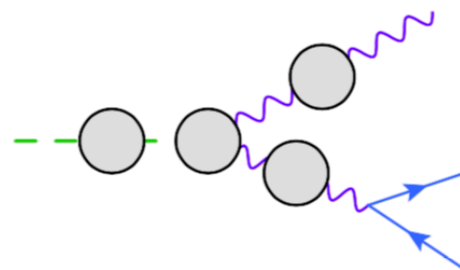
$H \rightarrow \mu\mu, \tau\tau$



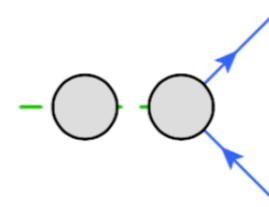
W,Z masses, oblique parameters



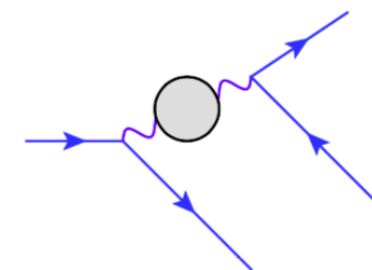
$H \rightarrow \gamma\gamma, \gamma Z$



$H \rightarrow Z ll, W l\nu$



$H \rightarrow bb$



$\mu$  decay



# RG operator mixing

- Renormalization group evolution of 10 Higgs and 7 top ops can account for dominant one-loop corrections.

$$\dot{c}_i \equiv 16\pi^2 \frac{dc_i}{d \ln \mu} = \gamma_{ij} c_j,$$

$$\dot{c}_H = (12y_t^2 N_c - 4g^2 N_c) c_{Hq}^{(3)} - 12y_t y_b N_c c_{Htb}, \quad (2.17)$$

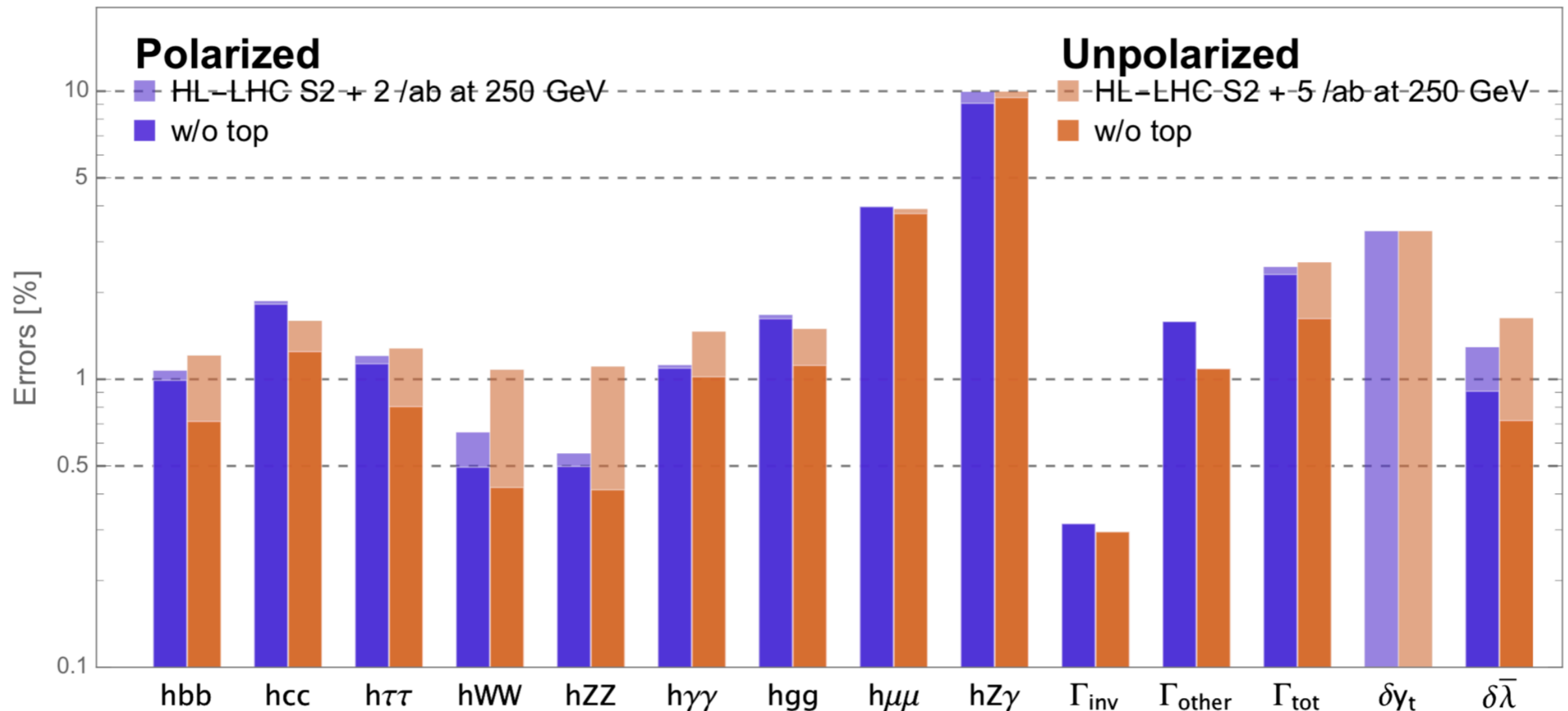
$$\dot{c}_T = (4y_t^2 N_c - \frac{8}{3} g'^2 Y_h Y_u N_c) c_{Ht} - (4y_t^2 N_c + \frac{8}{3} g'^2 Y_h Y_q N_c) c_{Hq}^{(1)} + 4y_t y_b N_c c_{Htb} \quad (2.18)$$

$$\dot{c}_{WW} = \frac{1}{4} (-2gy_t N_c c_{tW}), \quad (2.19)$$

$$\dot{c}_{BB} = \frac{1}{4t_W^2} (-4g' y_t (Y_q + Y_u) N_c c_{tB}), \quad (2.20)$$

S.Jung, J.Lee, J.Tian, M.Vos, 1909.xxxxx  
C.Zhang et al. 1804.09766, 1809.03520

# Top-loop now changes quite significantly

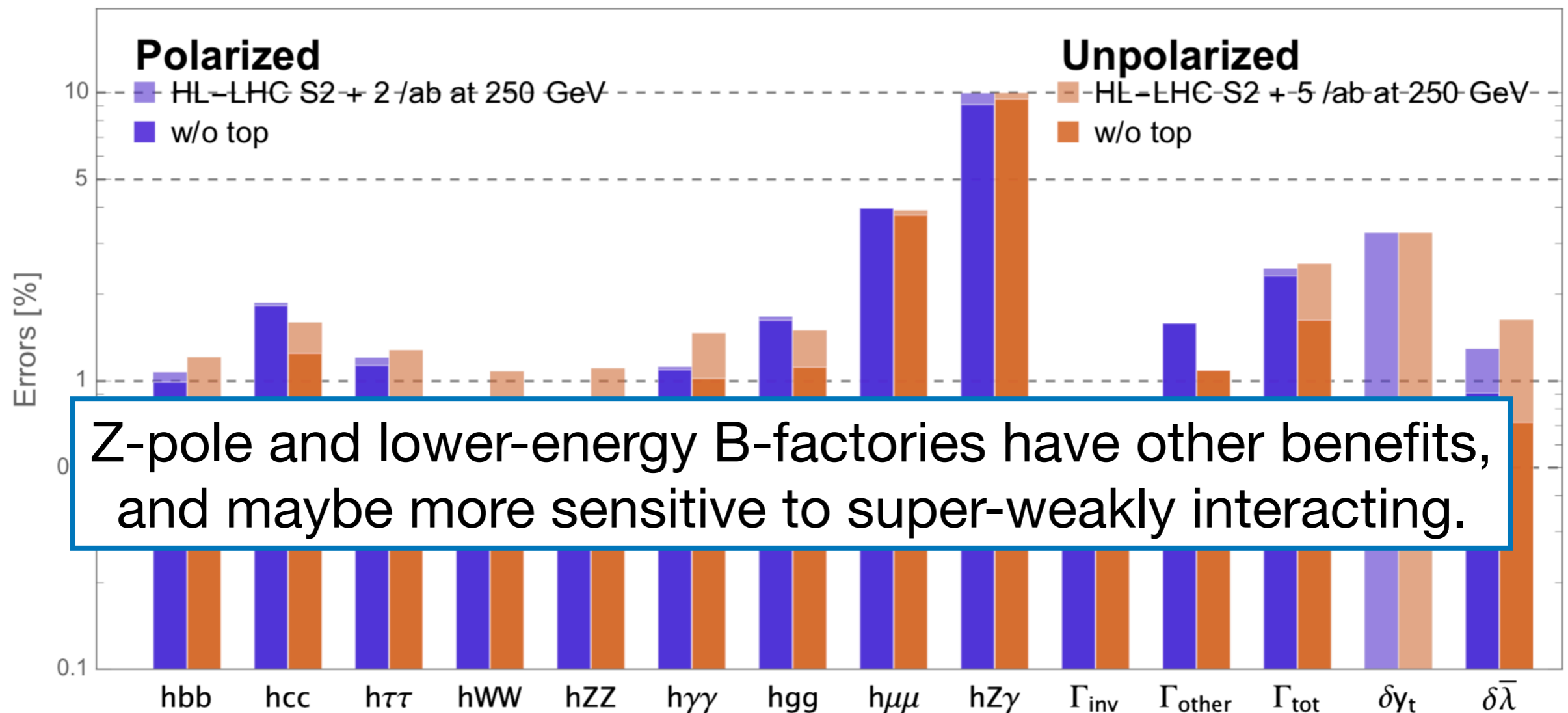


Unexpected benefits of polarization.

Model-indep formalism still needs improvements.

LHC precision really needs ingenuities. A lot to learn!

# Top-loop now changes quite significantly



Unexpected benefits of polarization.

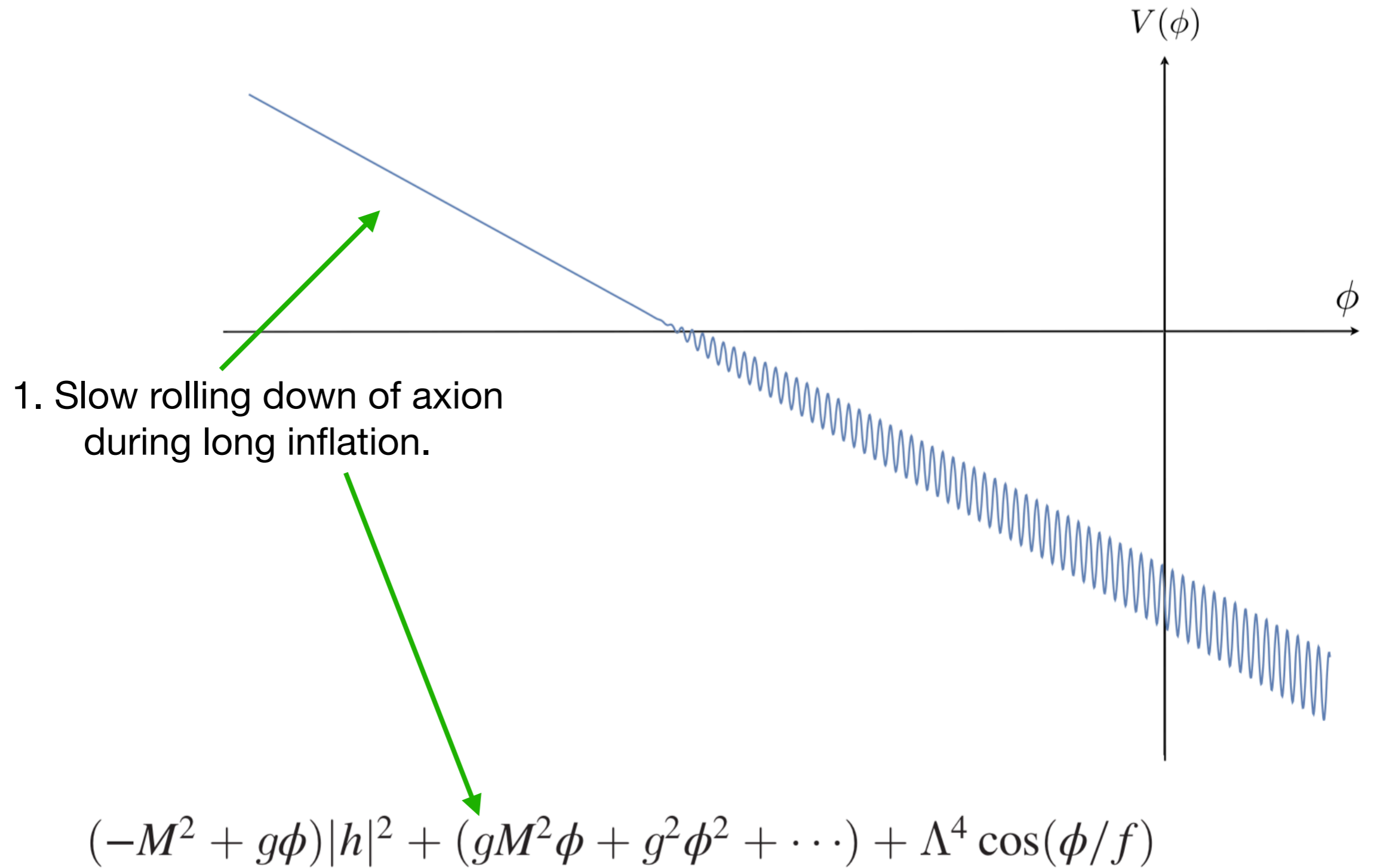
Model-indep formalism still needs improvements.

LHC precision really needs ingenuities. A lot to learn!

# **3. Relaxion test**

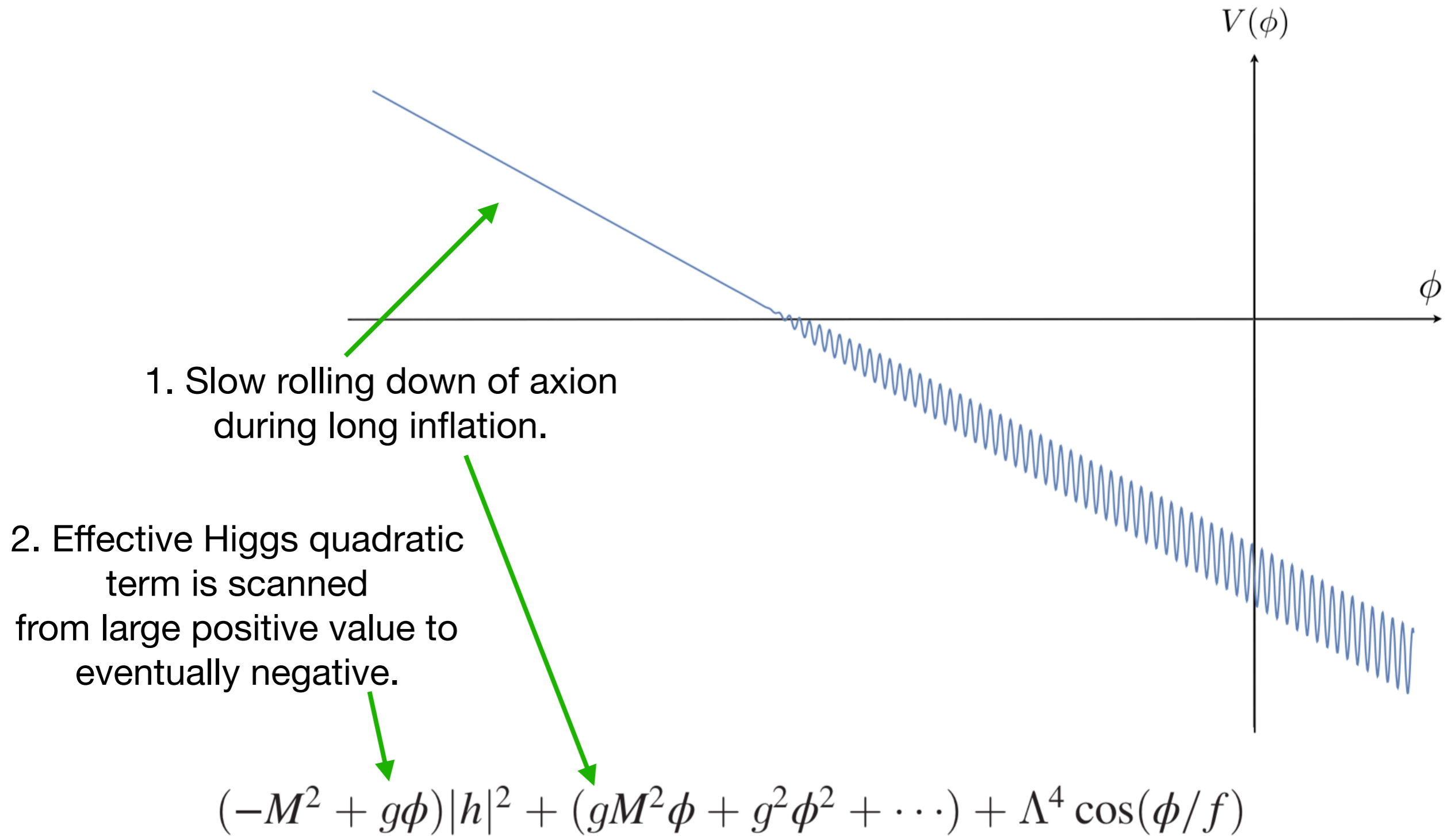
**hh+MET with hidden Nc**

# Relaxation of Higgs vev

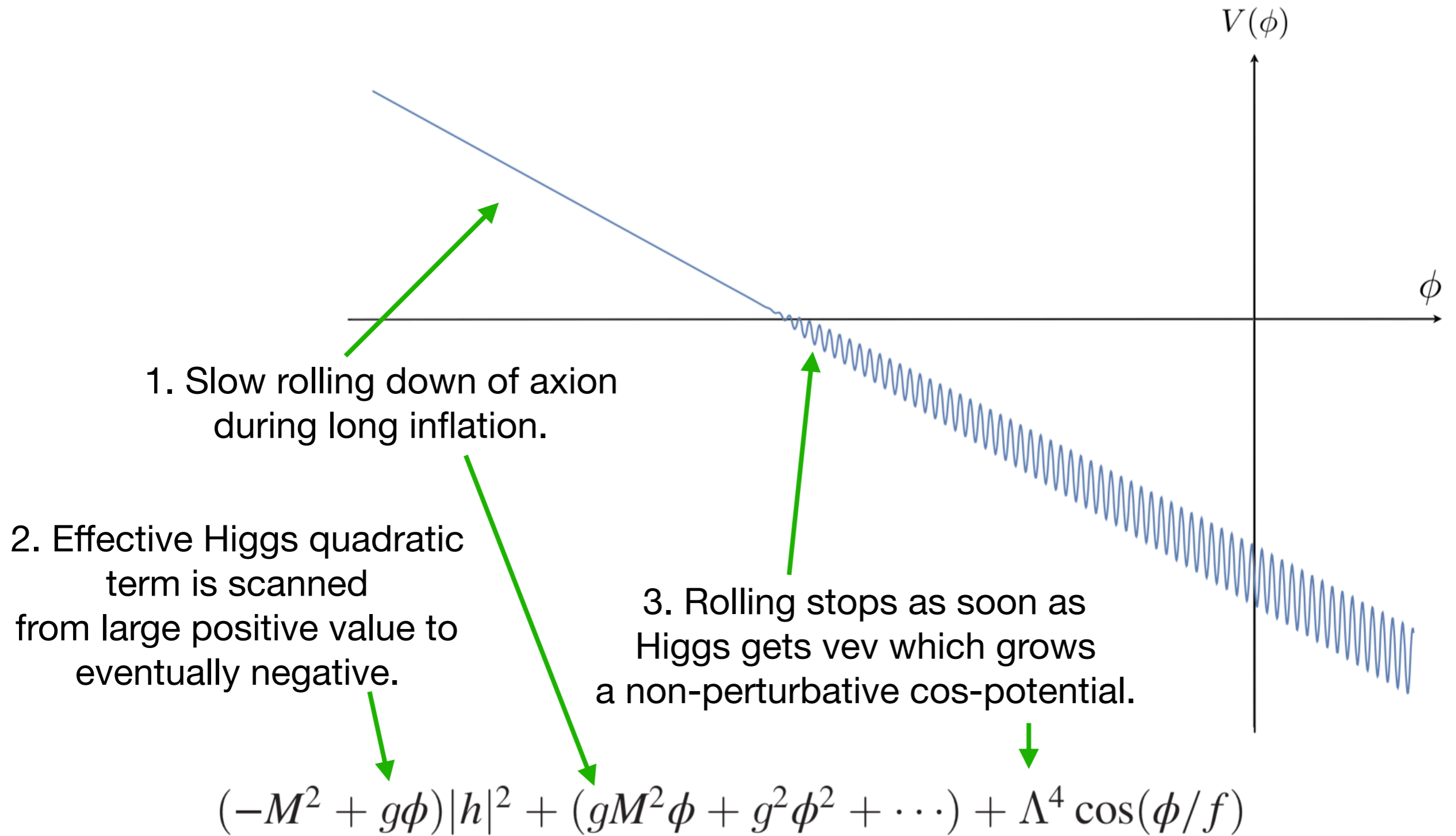


P.Graham, D.Kaplan, S.Rajendran, PRL2015

# Relaxation of Higgs vev



# Relaxation of Higgs vev



# Where does relaxion-Higgs coupling come from?

A natural example is

Vector-like lepton (with a hidden confining  $SU(N_c)$ )

$$y H L N^c + y' H^\dagger L^c N + m_L L L^c + m_N N N^c$$

& relaxion

$$\frac{1}{16\pi^2} \frac{\phi}{f} G_H \tilde{G}_H$$

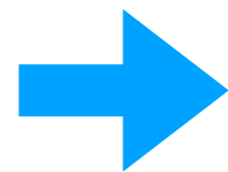
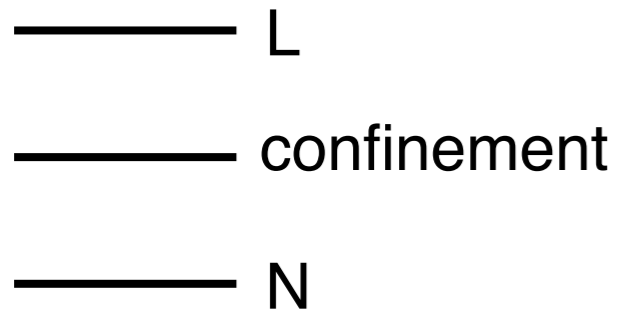


# Where does relaxion-Higgs coupling come from?

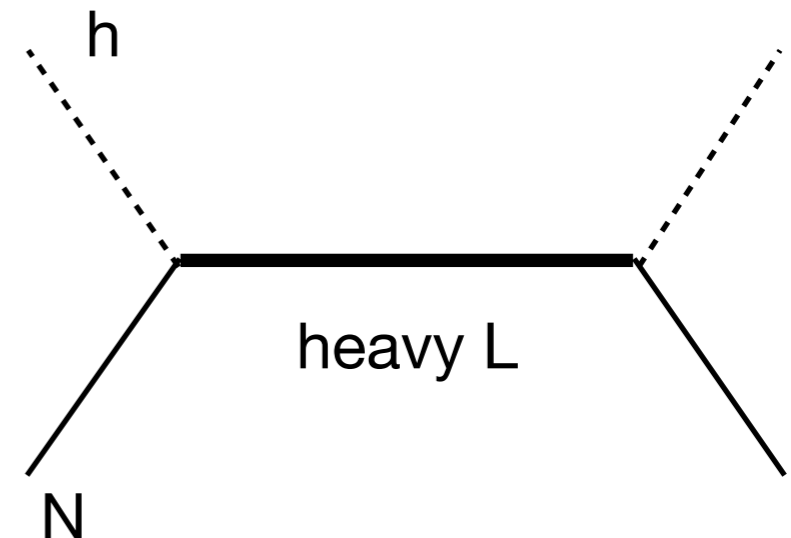
A natural example is

Vector-like lepton (with a hidden confining  $SU(N_c)$ )

$$yHLN^c + y'H^\dagger L^c N + m_L LL^c + m_N NN^c$$



$$\frac{yy'}{m_L} |H|^2 NN^c.$$



# Where does relaxion-Higgs coupling come from?

A natural example is

After NNC condensation,  
we obtain the desired relaxion-Higgs coupling.

$$\Delta V = -M^2 \cos\left(\frac{\phi}{f} + \alpha\right) |H|^2 - \Lambda^4 \cos\left(\frac{\phi}{f}\right)$$

—— L

—— confinement

—— N

# Unique collider prediction

At leading approximation, L-Z mixing is zero.

Dominant decay mode is thought to be

$$L^0 \rightarrow hN, \quad L^\pm \rightarrow L^0 + \text{soft},$$

so that the collider signal is

$$hh + \text{MET} \quad \longrightarrow \quad 4b + \text{MET}$$

without corresponding  $ZZ + \text{MET}$  nor  $hZ + \text{MET}$ .

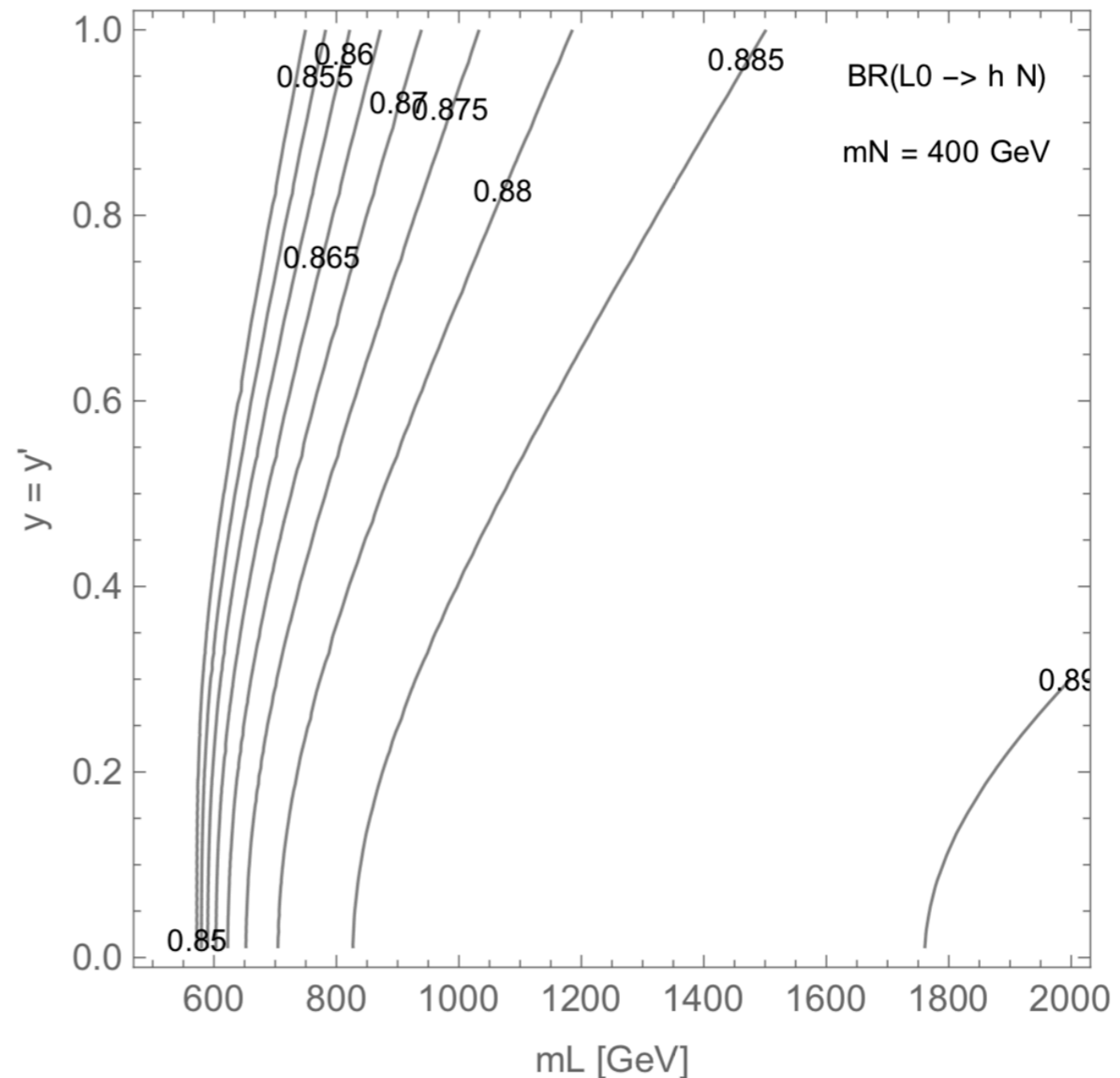
# Unique collider prediction

But small L-Z mixing induces

$$ZZ + \text{MET} \quad hZ + \text{MET}$$

Interestingly,

$\text{BR}(h) \approx 0.88$  is almost fixed!



S.Jung, K.S.Jeong, J.C.Park, S.Shin, 1910.xxxxx

# Unique collider prediction

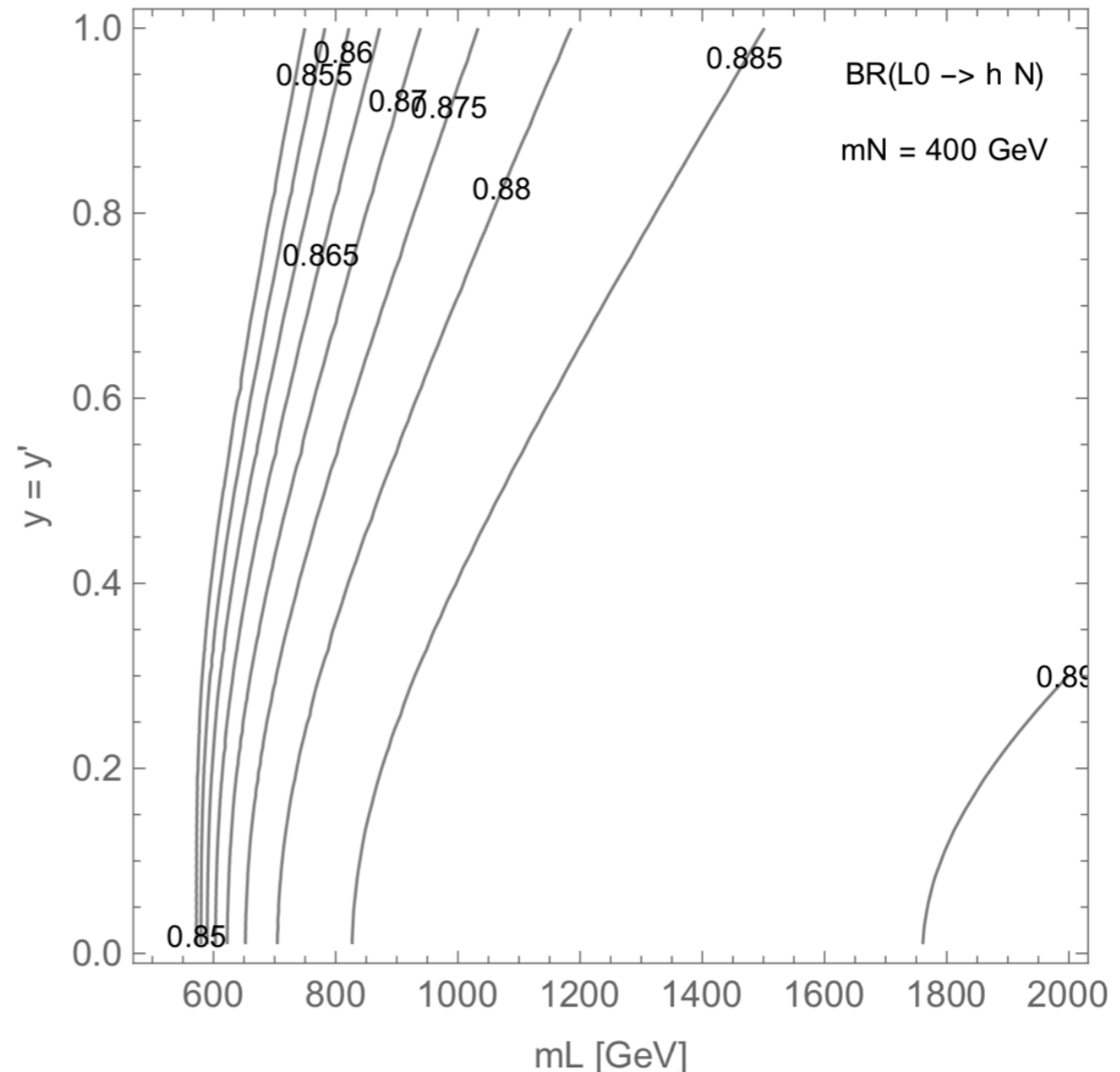
But small L-Z mixing induces

$$ZZ + \text{MET} \quad hZ + \text{MET}$$

Interestingly,

$\text{BR}(h) \approx 0.88$  is almost fixed!

It's again due to the Goldstone eq thm, that overcomes small mixing!



S.Jung, K.S.Jeong, J.C.Park, S.Shin, 1910.xxxxx

# Predictions and limits on relaxion

Prediction 1:

Dominant  $hh + \text{MET}$  .

Prediction 2:

Hidden  $N_c$ -enhanced Drell-Yan.

4b+MET at current LHC14 30/fb leads to 1 TeV lower limit.

**4. “Alpha Go sensei  
never speaks”**

# Before Alpha Go vs. Lee Sedol:

Machine is much better at calculation than human.

But machine does not have intuition and abstraction.

→ “Human will be better in the opening part of a game,  
while machine will do better at the ending part.”



# After Alpha Go vs. Lee Sedol:

Alpha Go turned out to be much much better than human in the BEGINNING.

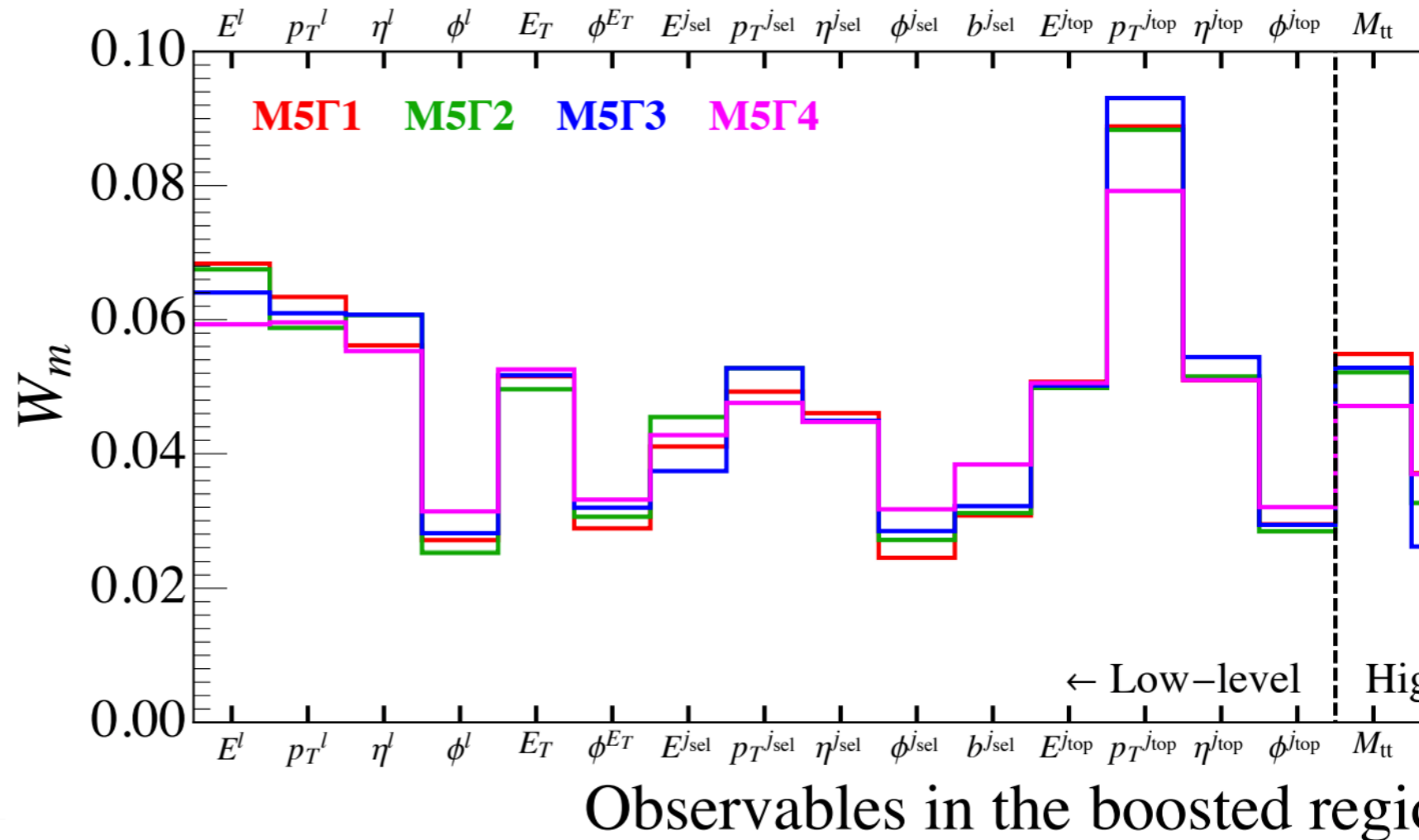
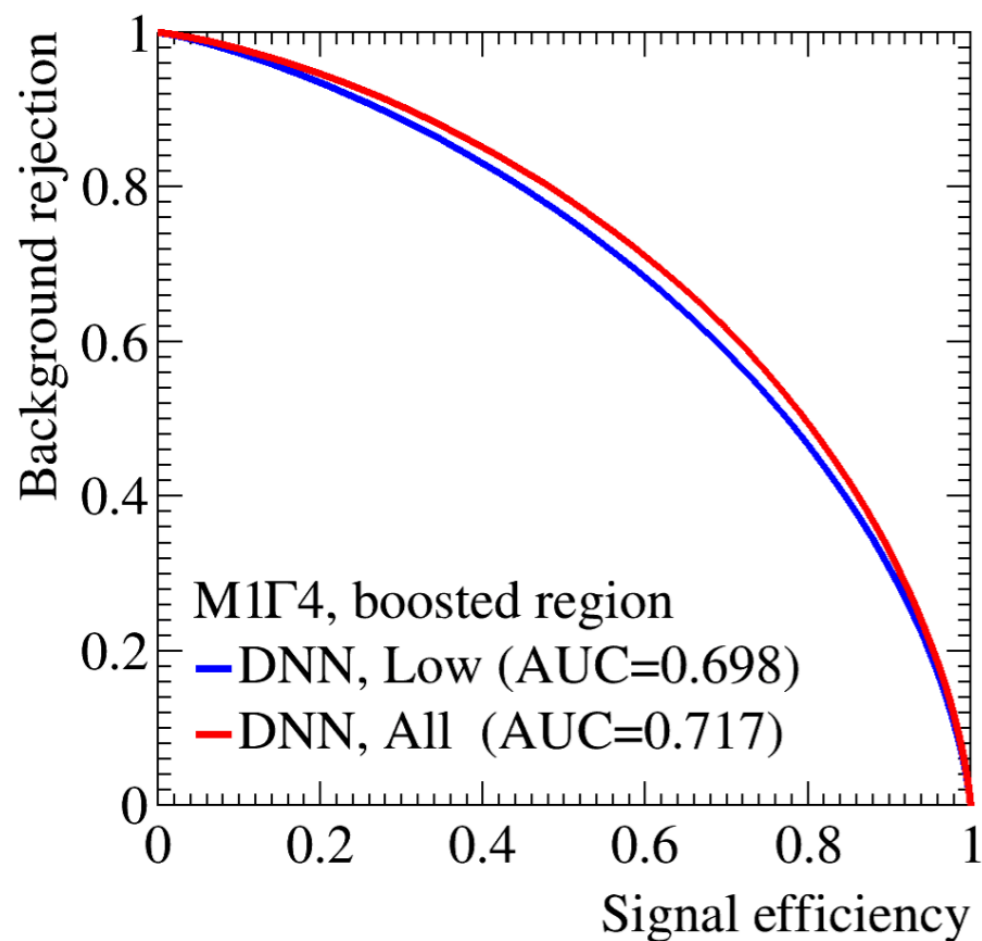
Intuition and abstraction are nothing but the pattern recognition based on experiences!  
→ This is one great lesson that changed my mind.

Machine can do the *unexpected!*

# Asking to a machine

specialized at signal vs. background,

“Did you learn these observables?  
How important are these observables?”



S.Jung, D.Lee, K.Xie, 1906.02810

# Remark

But there could still be unknown (to us) info that are not identified in our analysis!!

It's difficult to find out something that we don't know even though the machine had learned it.

# Lee Changho

“Alpha Go teacher never speaks.”

Being able to communicate with networks will enable better explorations of the Nature, beyond what we know.

# Takehome messages

- **Future TeV-scale SUSY** is *qualitatively different*, requiring new viewpoints, insights and works.
  - Higgsino vs. gaugino exhibit underlying symmetry.
  - What is a resonance?
- **Higgs precision** largely prefers to *polarization!* (of linear  $e^+e^-$  colliders.)
  - Z-pole or B-factories for super-weakly interacting.
  - Truly model-independent formalism not yet.
- **Relaxion** solutions are also *uniquely testable* at colliders.
- “**Alpha Go** sensei never speaks.”