

Directed flow v1



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Third flow component as QGP signal

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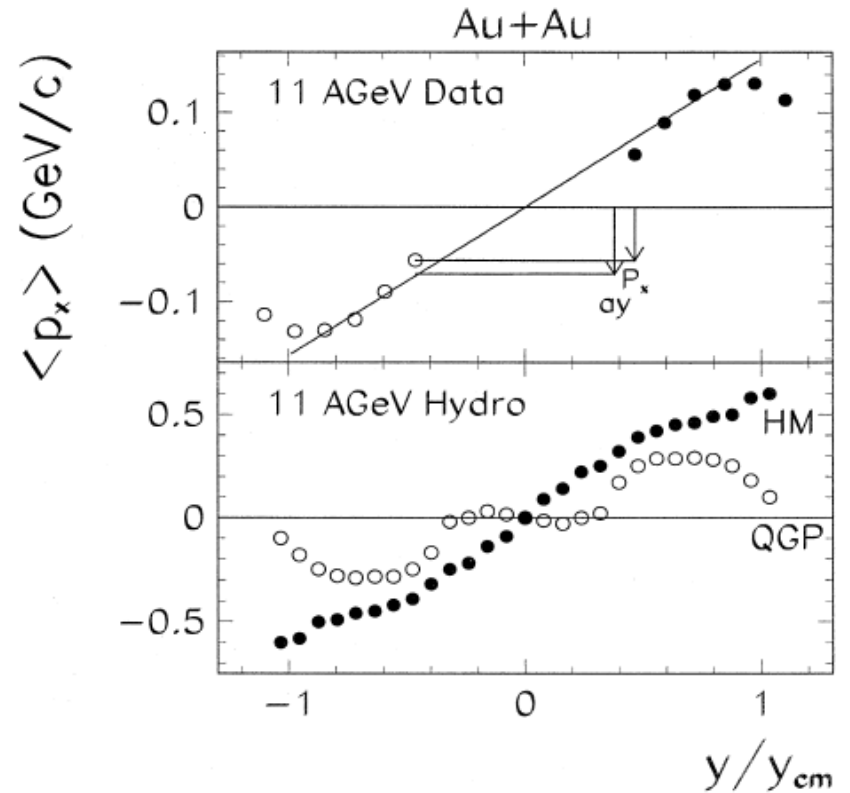
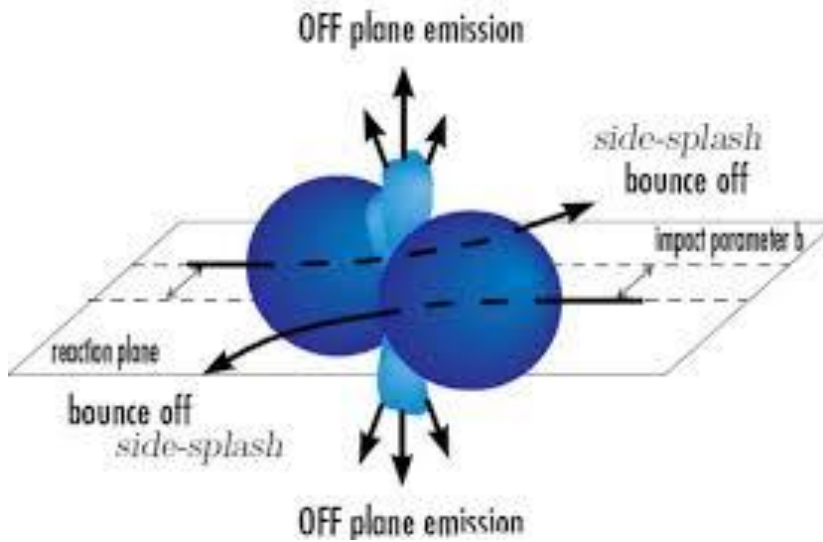
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Abstract

A review of earlier fluid dynamical calculations with QGP show a *softening* of the directed flow while with hadronic matter this effect is absent. The effect shows up in the reaction plane as enhanced emission which is orthogonal to the directed flow. Thus, it is not shadowed by the deflected projectile and target. As both of these flow components are in the reaction plane they form an enhanced ‘elliptic flow’ pattern. Recent experimental data at 11 AGeV and above show the same *softening*, hinting at QGP formation. © 1999 Published by Elsevier Science B.V. All rights reserved.

Directed flow v1

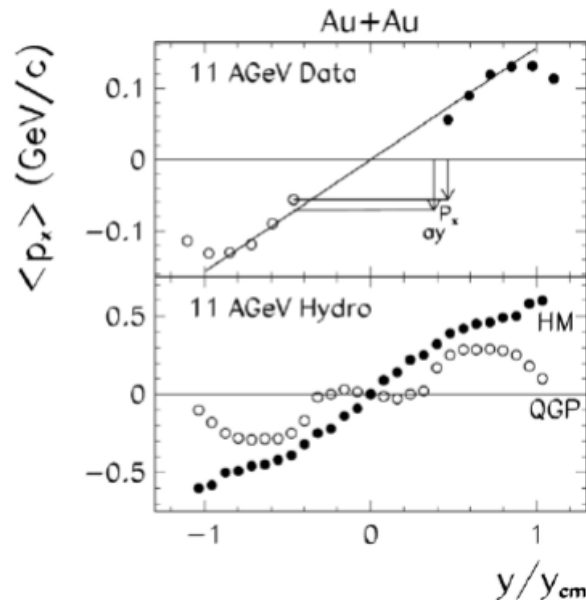


Directed flow v_1 – NA61

Directed flow and the onset of deconfinement

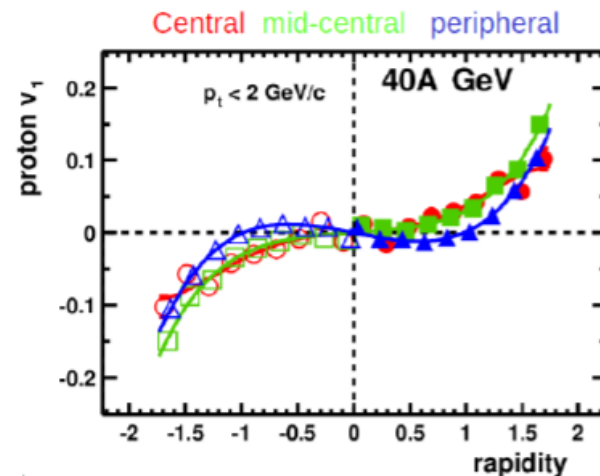
Directed flow v_1 is considered to be sensitive to 1st order phase transition (softening of EOS). Expected: non-monotonic behavior (positive → negative → positive) of proton dv_1/dy as a function of beam energy - “collapse of proton flow”

Predictions of hydrodynamical model:



$$v_1 = \left\langle \frac{p_x}{p_T} \right\rangle$$

Directed flow measured by NA49 at middle SPS energy (“anti-flow” of protons at mid-rapidity):

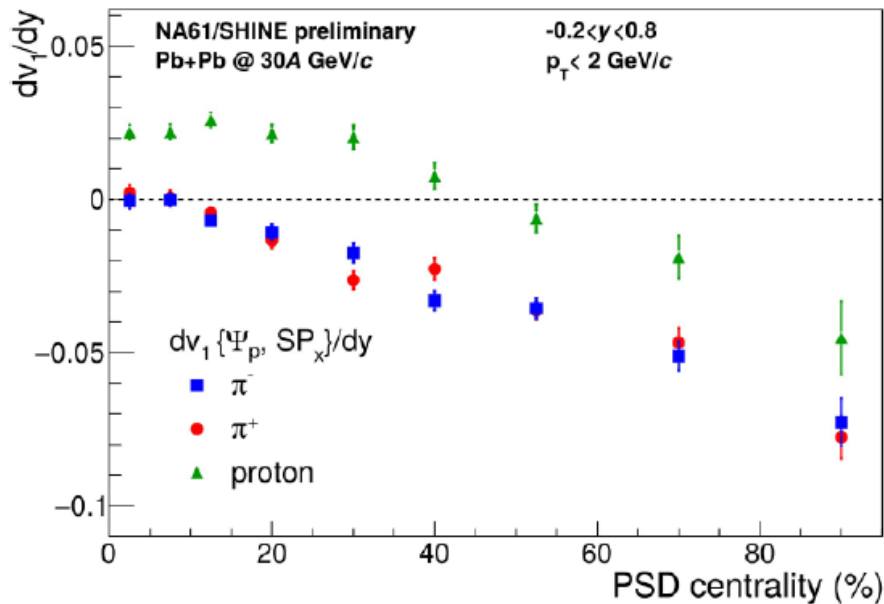


Directed flow v_1 – NA61

Centrality dependence of dv_1/dy in Pb+Pb at $\sqrt{s_{NN}} = 7.6$ GeV

NA61/SHINE fixed target setup → tracking and particle identification over wide rapidity range

Flow coefficients are measured relative to the **spectator plane estimated with Projectile Spectator Detector (PSD)** → unique for NA61/SHINE



Close to mid-rapidity (-0.2 < y < 0.8)

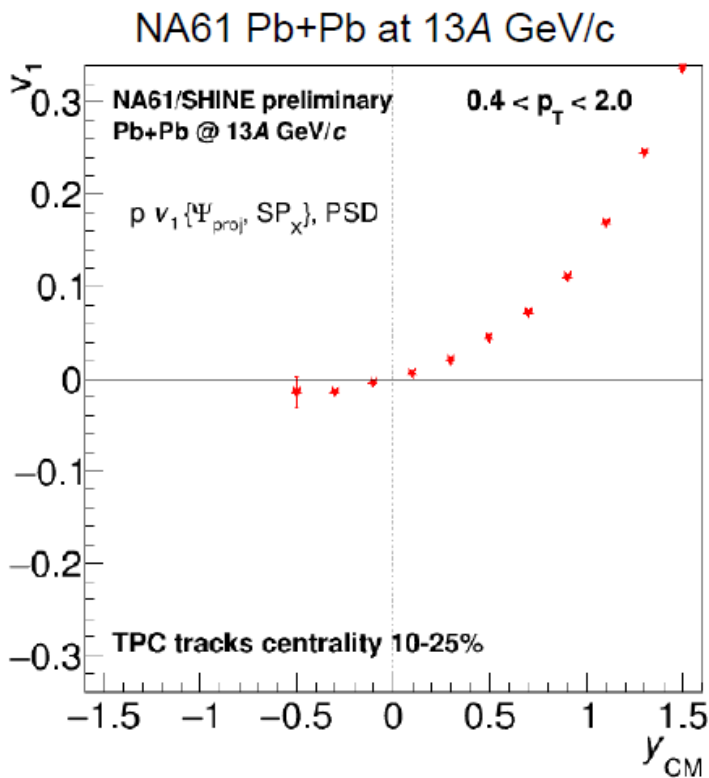
- slope of pion v_1 is negative for all centralities

- slope of proton v_1 changes sign at centrality of about 50%

More NA61/SHINE flow results:
Klochkov, Selyuzhenkov (QM2018 talk)

Directed flow v_1 – NA61

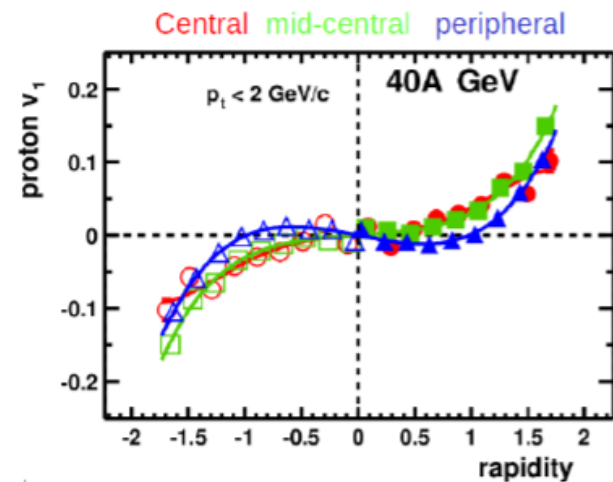
Proton directed flow vs rapidity



No evidence for the collapse of
proton directed flow in Pb+Pb
at 13A GeV/c



Directed flow measured by NA49
at middle SPS energy (“anti-flow”
of protons at mid-rapidity):

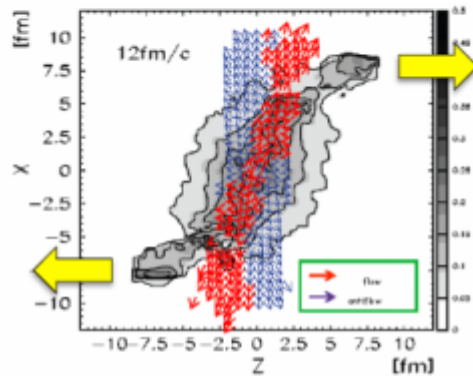


Directed flow v_1 – STAR

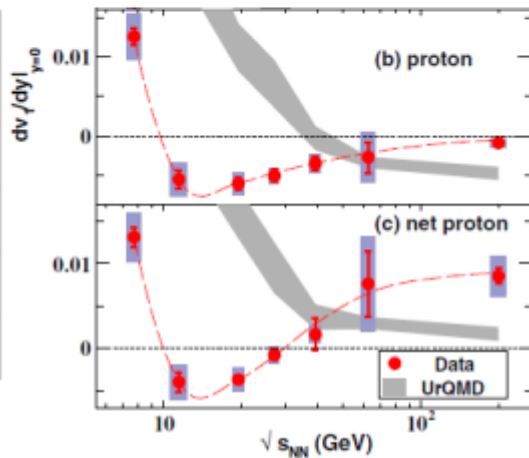
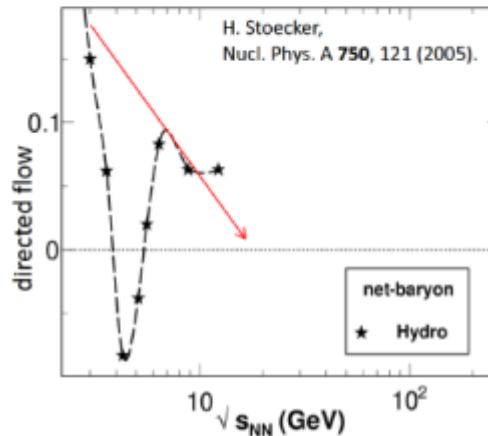
STAR

Directed Flow (v_1)

STAR, Phys. Rev. Lett. 112 (2014) 162301



J. Brachmann et al., PRC 61, 24909 (2000).

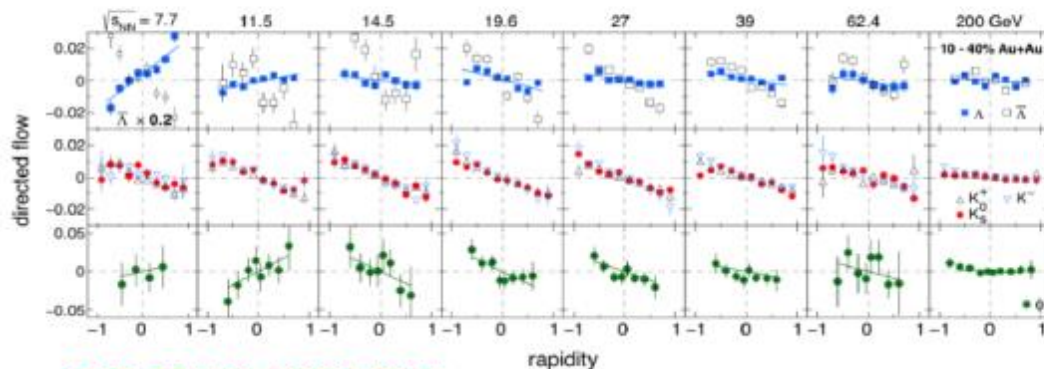


Probe of the softening of the Equation of State

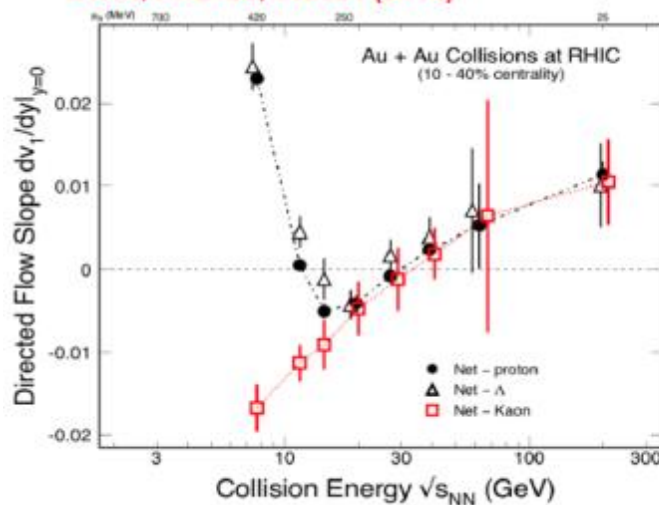
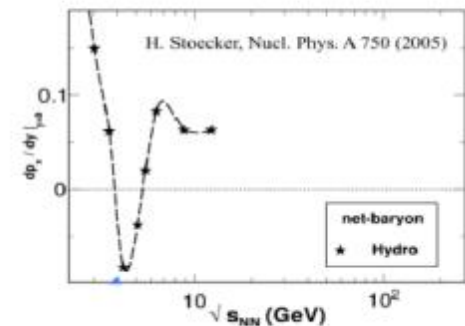
- strong softening: consistent with the 1st-order phase transition
- weaker softening: more likely due to crossover

Directed flow v_1 – STAR

Directed flow (STAR BES-I)



STAR, PRL112, 162301 (2014)
 STAR, PRL120, 062301 (2018)

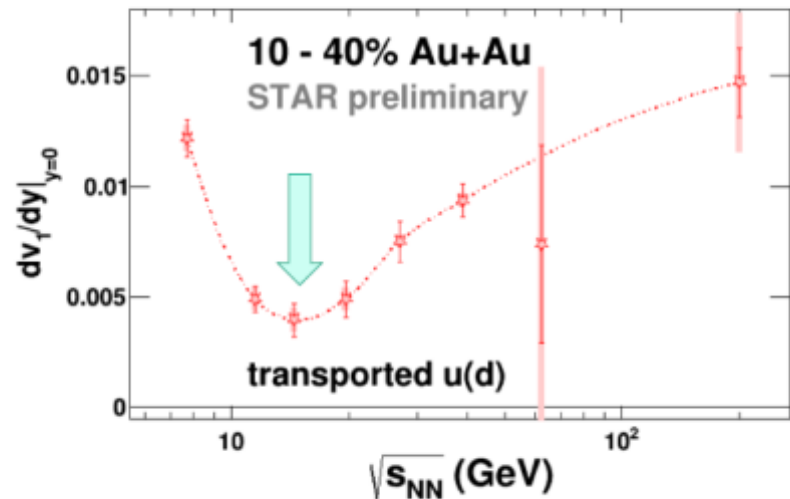
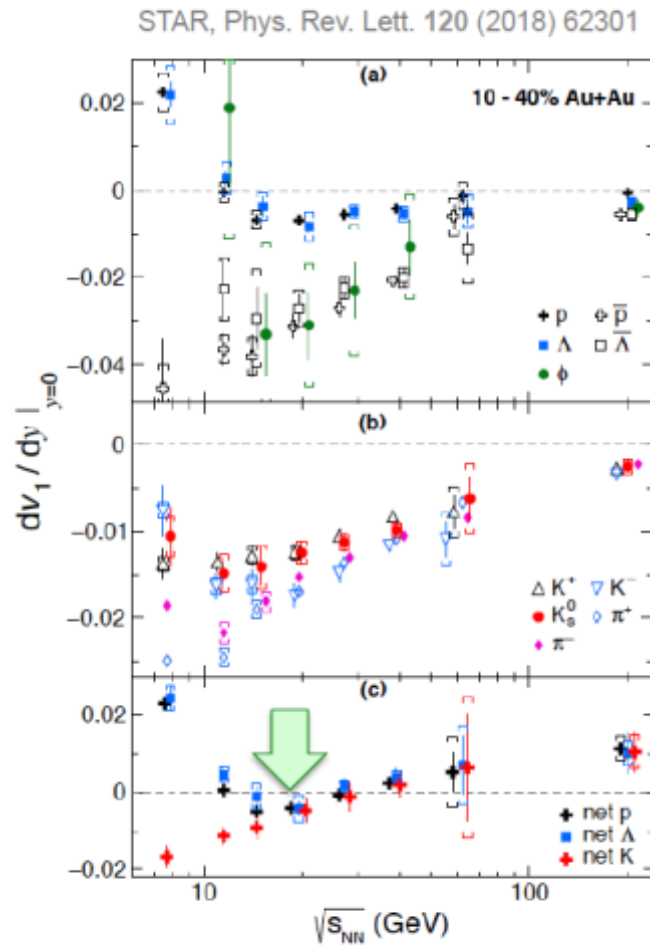


- Sign change of proton dv_1/dy , softening of EOS, first-order phase transition
- Double sign change seen in net-proton, net- Λ , not seen in net-kaon
- Need theory to explain

Directed flow v_1 – STAR



Directed Flow of Identified Particles



$$(v_1)_{trans.u(d)} = [(v_1)_{net p} - (3 - N_{trans.u(d)})(v_1)_{\bar{u}(\bar{d})}] / N_{trans.u(d)}$$

$$N_{trans.u(d)} = 3[1 - \exp(-2\mu_{u(d)}/T_{ch})] / (1 - r_{\bar{p}/p})$$

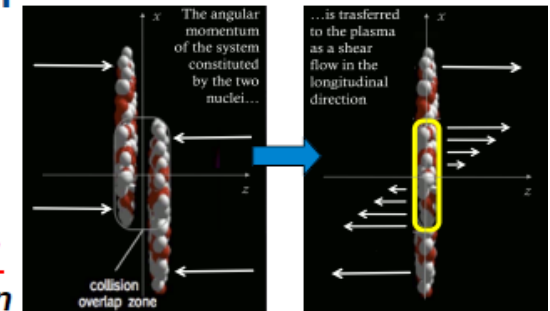
- 10 species & 8 energies allow a detailed study of constituent-quark v_1 . In most cases, the coalescence picture works for both “produced” particles and “net” particles
- “Transported quark” v_1 has a local minimum at ~ 14.5 GeV

Heavy quarks directed flow v_1

Heavy Flavour dynamics: sources of v_1 for charm quarks

- Vorticity due to the large orbital angular momentum in uRHIC $J \approx 10^5 - 10^6 \hbar$

Becattini, Piccinini e Rizzo, PRC 77, 024906 (2008)
 Csernai, Magas and Wang - Phys. Rev. C 87 (2013) 034906
 Becattini et al, EPJ C 75, 406 (2015)
 Deng and Huang, PRC 93, 064907 (2016)
 Jiang, Lin and Liao, PRC 94, 044910 (2016); PRC 95, 049904 (2017)



- **Are HQ affected by the initial vorticity of the QGP?**
 Solving the relativistic Langevin eq. with tilted initial distribution in the reaction plane produce a v_1 of D meson several times larger than that of charged particle.

S. Chatterjee, P. Božek PRL 120 (2018) no.19, 192301

- Intense magnetic field B:

created on Earth $\approx 10^7$ Gauss

in Neutron Star $\approx 10^{13}$ Gauss

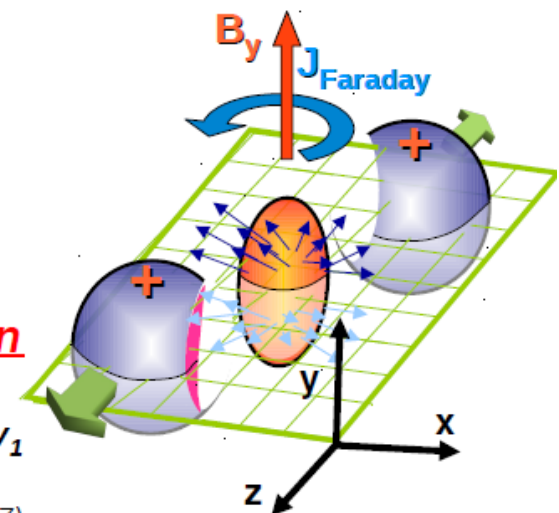
in uRHIC $\approx 10^{19}$ Gauss $\approx 10 m_\pi^2$

A. Bzdak, V. Skokov, PLB 710 (2012) 171-174
 K. Tuchin, PRC 88, 024911 (2013).
 K. Tuchin, Adv. High Energy Phys. 2013, 1 (2013).
 K. Hattori, X.-G. Huang Nucl.Sci.Tech. 28 (2017) no.2, 26.

- **Are HQ affected by the initial EM field produced in a HIC?**

Solving the relativistic Langevin eq. with Lorentz force a sizeable v_1 for charm (anti-charm) quarks is produced.

S.K. Das, S. Plumari, S. Chatterjee, J. Alam, F. Scardina, V. Greco, PLB 768 (2017) 260.

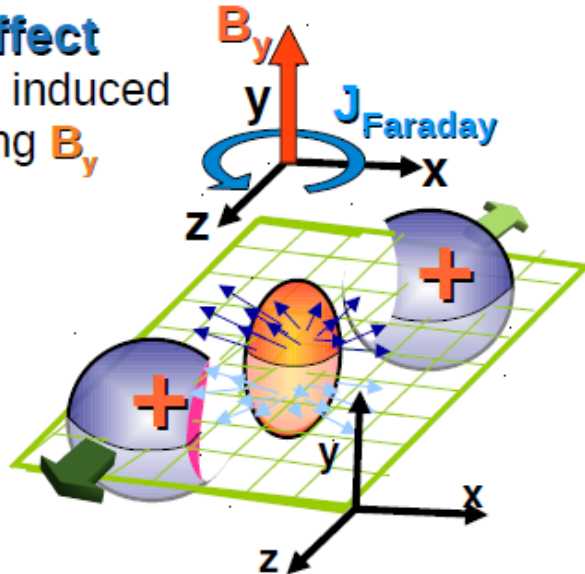


Heavy quarks directed flow v_1

The direct flow v_1 originates from two competing effects:

Faraday effect

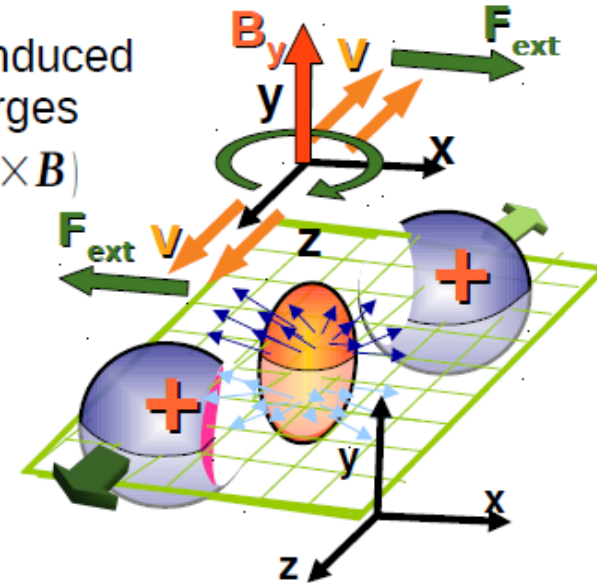
Electric field induced by decreasing B_y



Hall effect

Lorentz force induced by moving charges

$$F_{ext} = qE + \frac{q}{E_p} (p \times B)$$



Transport properties of Heavy Quarks and their correlations to the bulk dynamics and the initial Electromagnetic field

S. Plumari

Università degli Studi di Catania

INFN-LNS

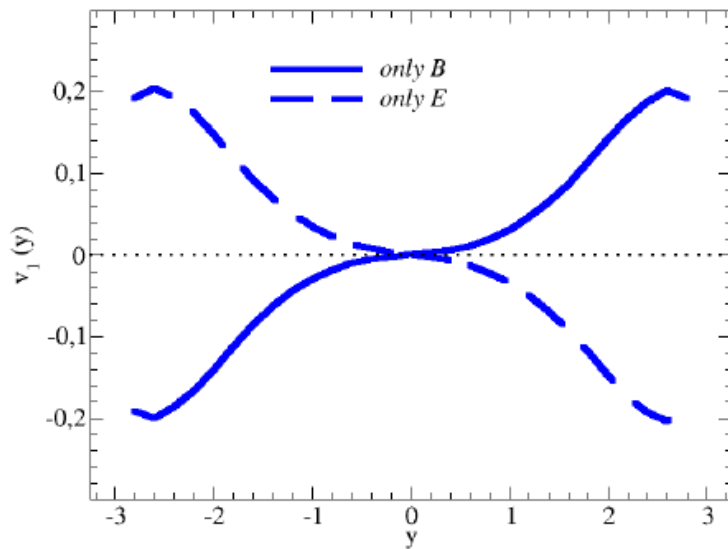
IN COLLABORATION WITH:

V. Minissale, G. Coci, L. Oliva, S. K. Das, V. Greco

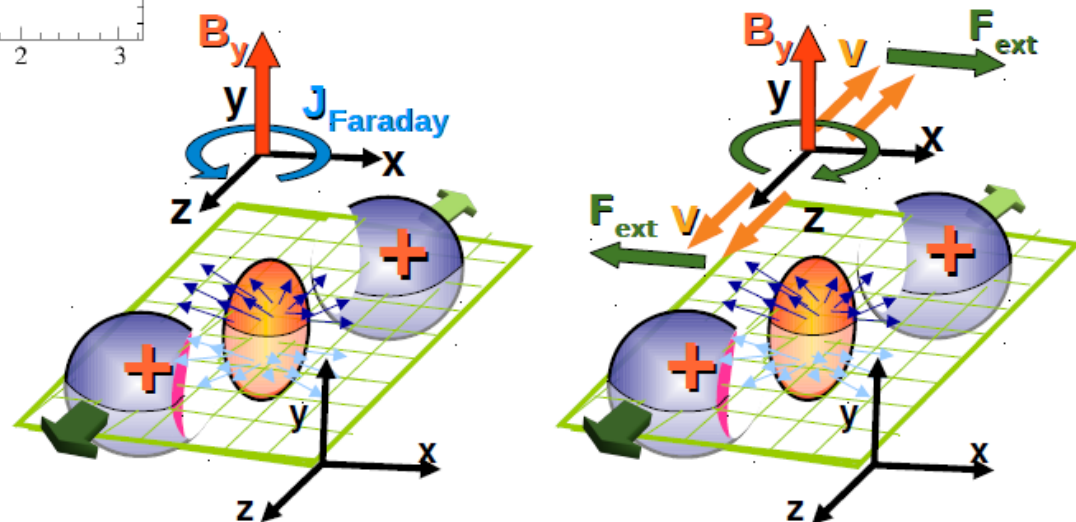


Heavy quarks directed flow v_1

Balance between Magnetic and Electric fields

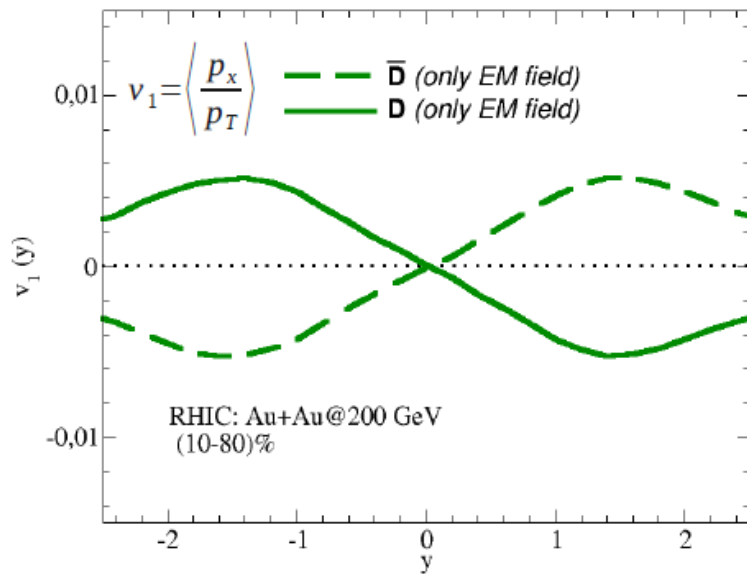


- Decreasing magnetic field B_y creates E_x that induces a current in opposite direction w.r.t. to the Magnetic Hall drift: delicate balance!
- Larger initial ($t < 1$ fm/c) field important to determine a sizeable v_1 flow
S.K. Das et al., PLB 768 (2017)



Heavy quarks directed flow v_1

Direct Flow v_1 of charm quarks

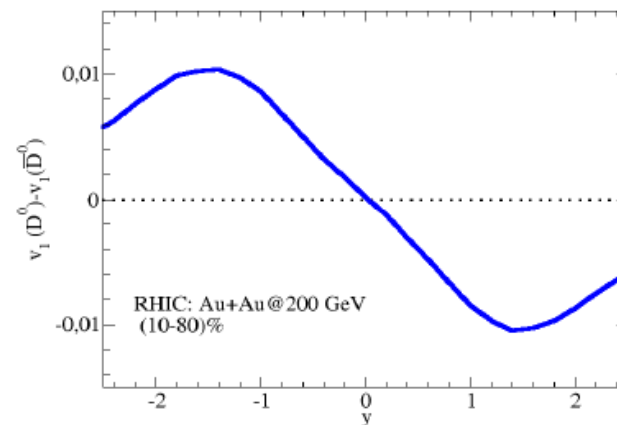


For light quarks was predicted $v_1 \approx 10^{-3} - 10^{-4}$
 U. Gürsoy, D. Kharzeev, K. Rajagopal PRC 89, 054905 (2014).

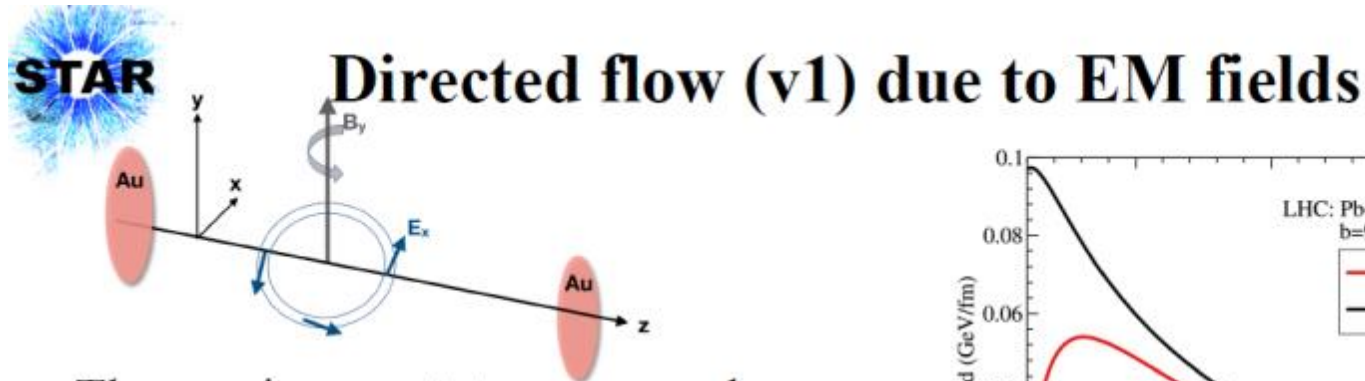
For charm quarks due to early production we find a sizeable v_1 with the same E-B evolution
 S. K. Das, S. Plumari, S. Chatterjee, J. Alam, F. Scardina, V. Greco, PLB768 (2017) 260-264.

HQ best probe for v_1 from e.m. field:

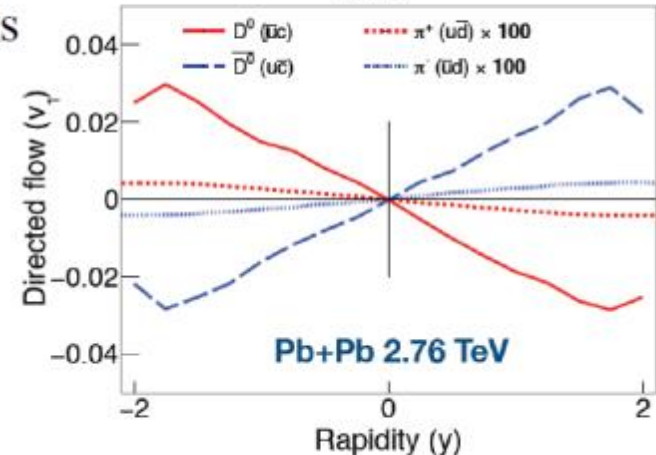
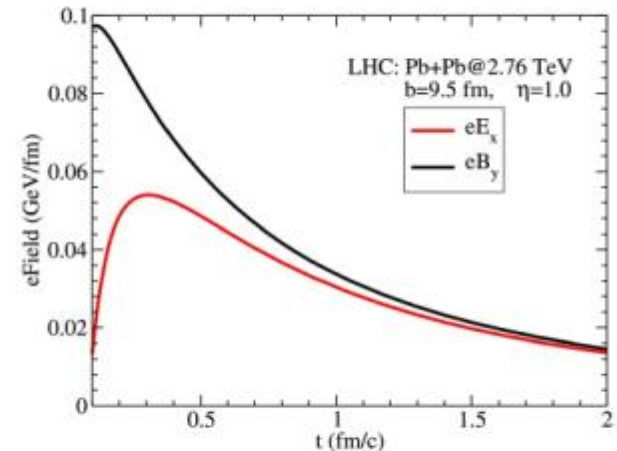
- $t_{\text{form}} \approx 0.1 \text{ fm/c}$
- $\tau_{\text{th}}(c) \approx \tau_{\text{QGP}} \gg \tau_{\text{e.m}}$
- do not mix vorticity [Odd- parity]



Heavy quarks directed flow v_1 - STAR



- The moving spectators can produce enormously large electromagnetic field ($eB \sim 10^{18}$ G at RHIC)
- Due to early production of heavy quarks ($\tau_{CQ} \sim 0.1$ fm/c) positive and negative charm quarks (CQs) can get deflected by the initial EM force
- D^0 and \bar{D}^0 v_1 can offer insight into the early time EM fields



Das et al., Phys Lett B 768, 260 (2017)

Heavy quarks directed flow v_1 - STAR

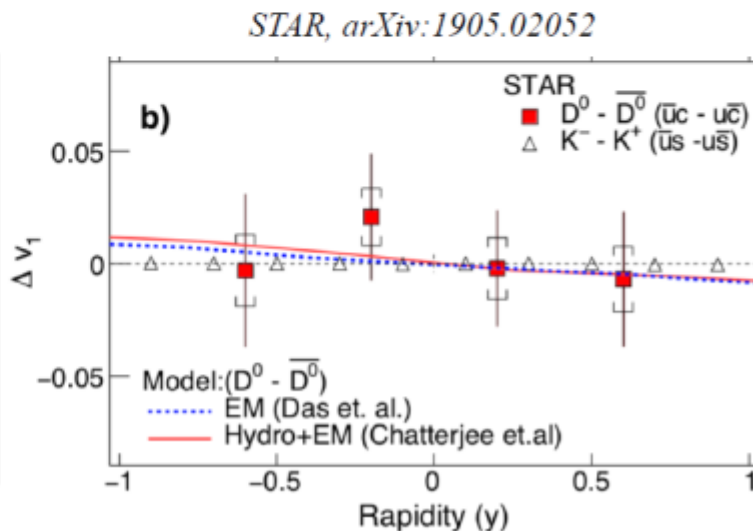
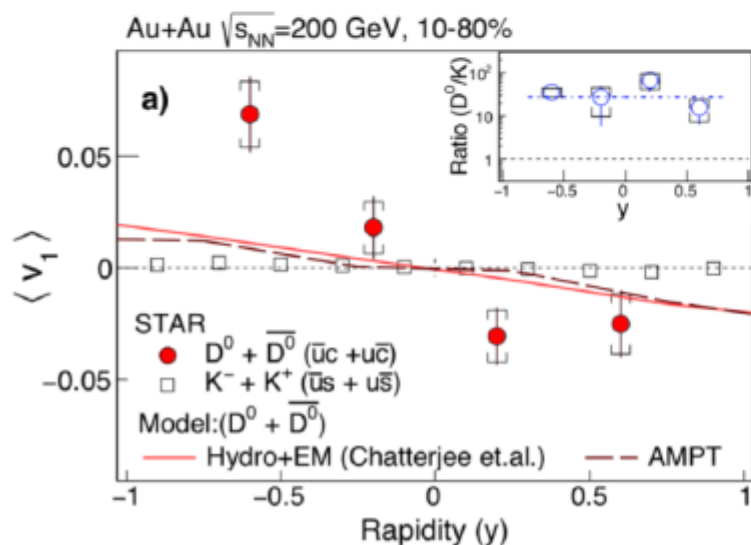


D^0 Directed Flow (v_1)

S.K. Das et al, PLB 768 (2017) 260

S. Chatterjee and P. Bozek, PRL 120 (2018) 192301

- Charm quarks interact with bulk medium $\rightarrow D^0 v_1$ sensitive to the initial tilt of the source (bulk)
- Charm and anti-charm quarks can be deflected differently by the initial EM field \rightarrow difference between D^0 and $\overline{D^0} v_1$ sensitive to EM field
- First observation of non-zero (negative) $D^0(\overline{D^0}) v_1$ slope, much larger than that of kaons
 $D^0 + \overline{D^0} dv_1/dy = -0.081 \pm 0.021(stat) \pm 0.017(sys)$
- More precise data are needed for Δv_1 $d\Delta v_1/dy = -0.041 \pm 0.041(stat) \pm 0.020(sys)$



Heavy quarks directed flow v_1 - ALICE

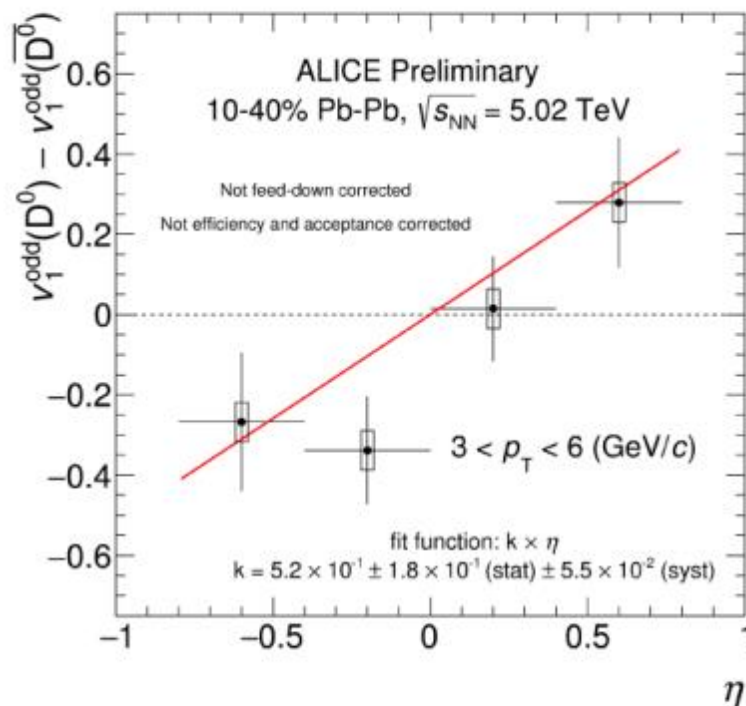


Universiteit Utrecht

Delta v_1 low p_T for D^0 meson



ALICE



- The difference of the directed flow, $\Delta v_1 \rightarrow$ quantify the effect of the charge separation due to the presence of an electromagnetic field
- Rapidity dependence of the charge difference Δv_1 is fitted using a linear function with slope k
- $k = 0.52 \pm 0.18(\text{stat}) \pm 0.05(\text{syst})$
- The significance of the measurement 2.7σ .

