



The Status of the DAMPE BGO calorimeter in space

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Outline

- DAMPE Experiment
- BGO Calorimeter
- In-flight status
- Summary

DAMPE Experiment

- DAMPE (DArk Matter Particle Explorer) is a spaceborne highenergy cosmic ray and gamma-ray detector
- •Launch: December 17, 2015
 - sun-synchronous orbit at the attitude of 500 km
 - Total weight ~1850 kg, power consumption ~640 W
 - Scientific payload ~1400 kg, 400 W
 - Operation time > 6 years

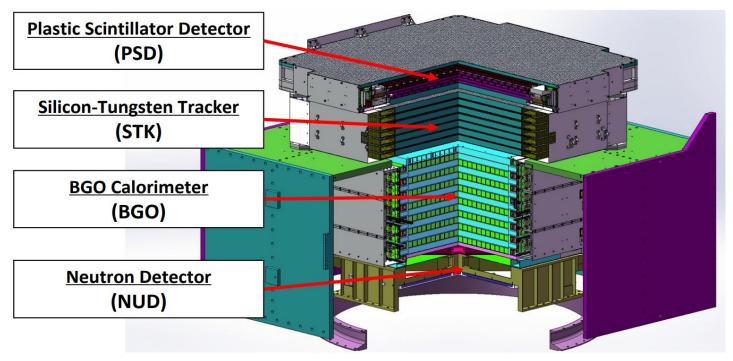
Scientific goals

- Search for dark matter particles in electrons and γ-rays
- Study of cosmic ray acceleration, propagation, and interaction
- \blacklozenge Study of high-energy γ -ray astronomy



DAMPE Experiment

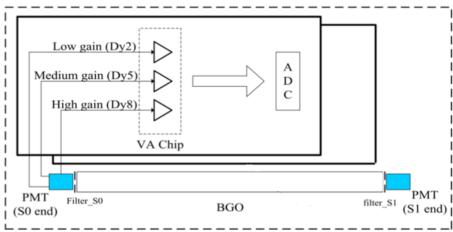
DAMPE detector

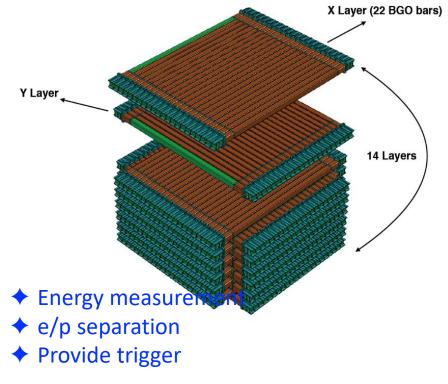


- Charge measurement (dE/dx in PSD, STK and BGO)
- Pair production and precise tracking (STK and BGO)
- Precise energy measurement (BGO)
- Electron/hadron identification (BGO and NUD)

BGO Calorimeter

- 308 BGO bars (2.5×2.5×60 cm³)
- 14 layers, 22 bars per layer
- Hodoscopic stacking alternating orthogonal layers
- Depth: 32 *X*₀, 1.6 λ_I
- Energy resolution: 1.5% @800 GeV for e/γ



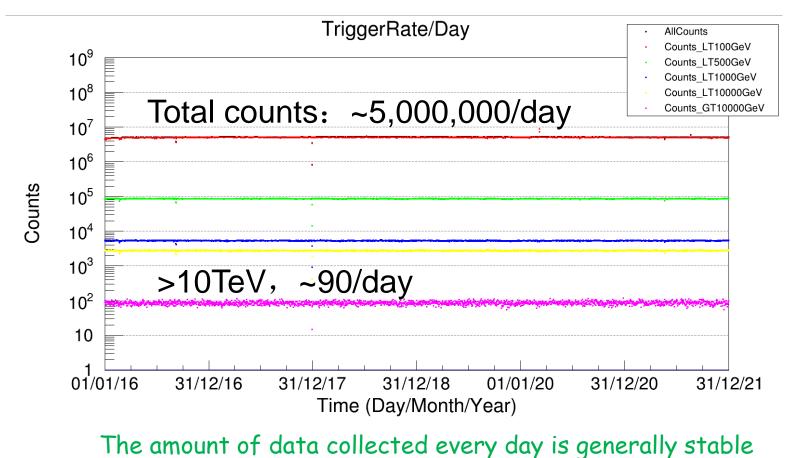


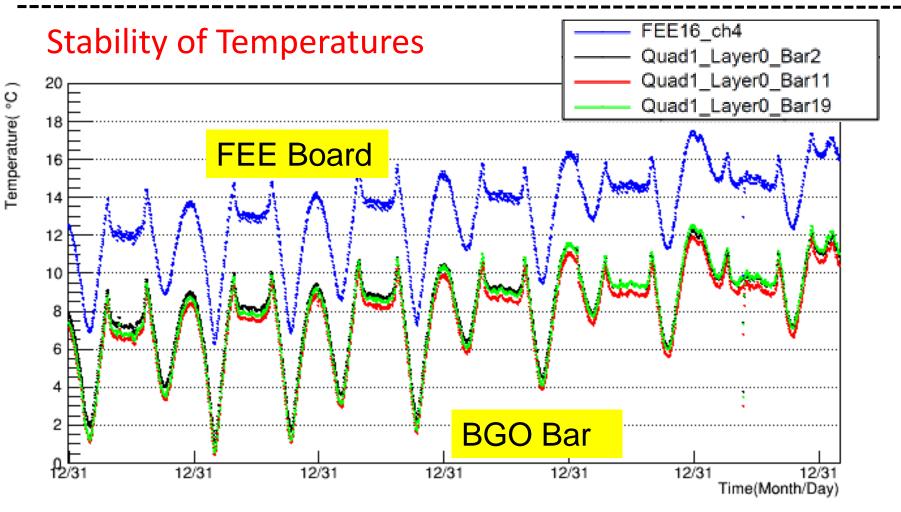
Large dynamic range (5 GeV-10 TeV)

-Two PMTs with different attenuation filters coupled with each BGO crystal bar in two ends -Multi-dynode readout of each PMT

Data acquired during six years

Time interval: 2016/01/01 - 2021/12/31

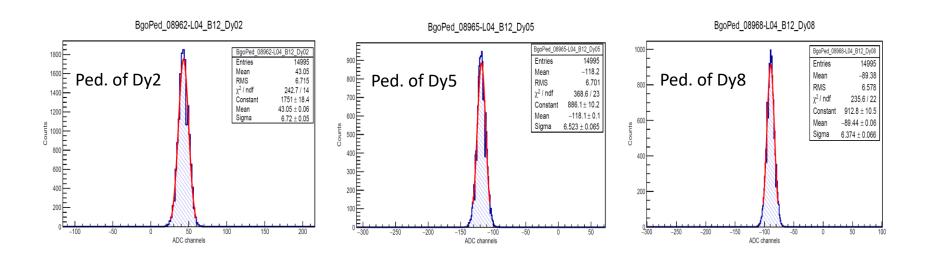




Temperature varies periodically and steadily

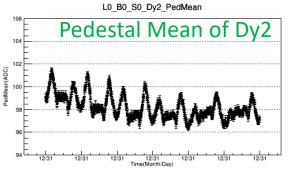
Pedestal

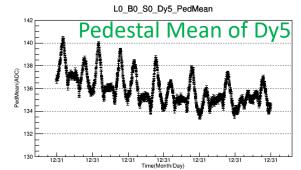
• The pedestal reflects the baseline and noise level of electronics.

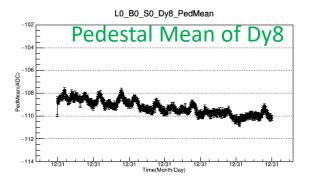


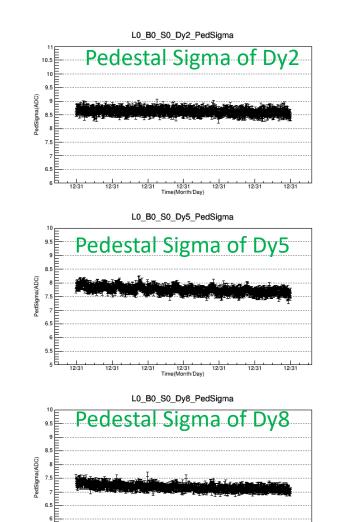
In order to measure the energy from 5GeV to 10TeV, a multi-dynode readout structure of PMT is designed.

Pedestal stabilitv









12/31

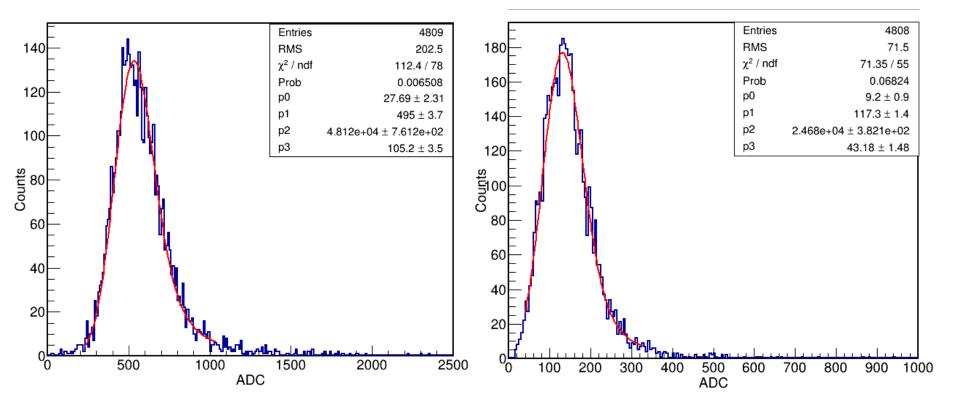
12/31

Time(Month/Day)

12/31

5.5

The MIP Spectra of a BGO bar

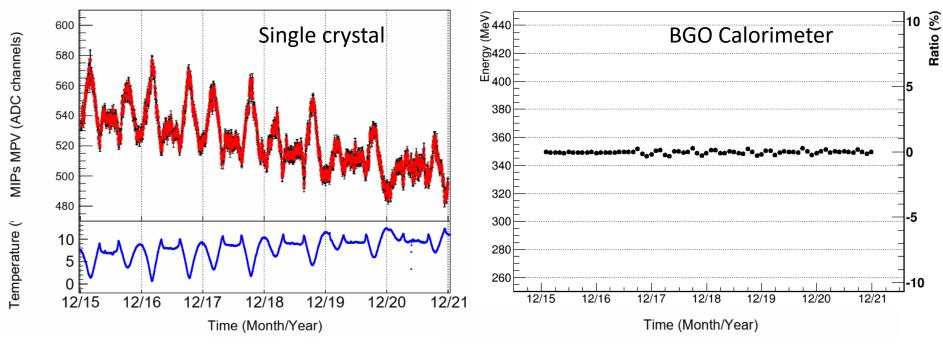


The response of the BGO Calorimeter to MIPs is the reference of energy reconstruction, and the precision of the MIP calibration influences the energy reconstruction directly.

Stability of MPV for MIP proton events

before temperature correction

after temperature correction

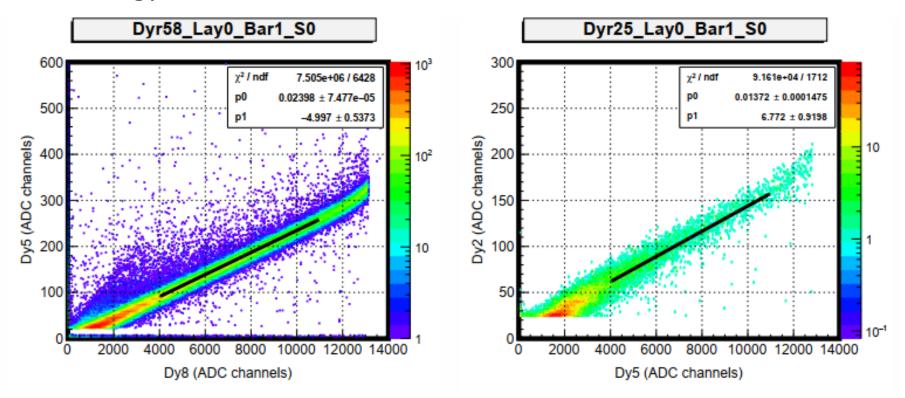


The MIPs MPV value and the temperature are roughly anti-correlated.

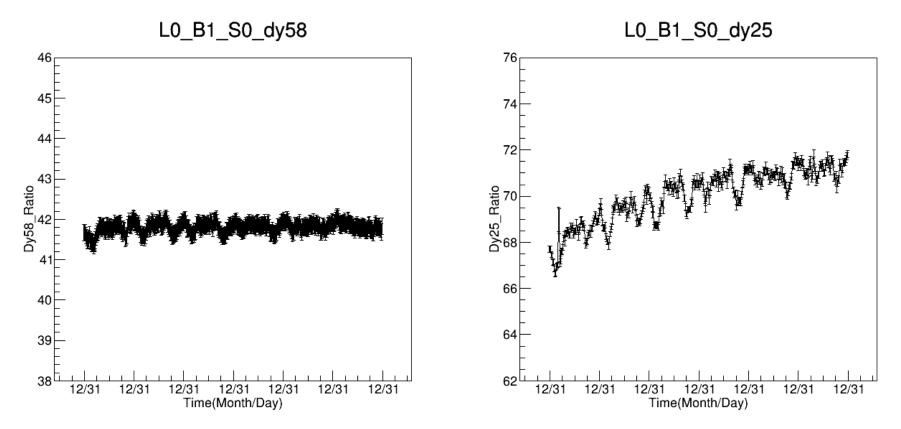
After temperature correction, Proton MIPs energy reconstruction stability is better than 1%

Dynode ratio

The ratios between the three dynodes are key parameters for energy reconstruction

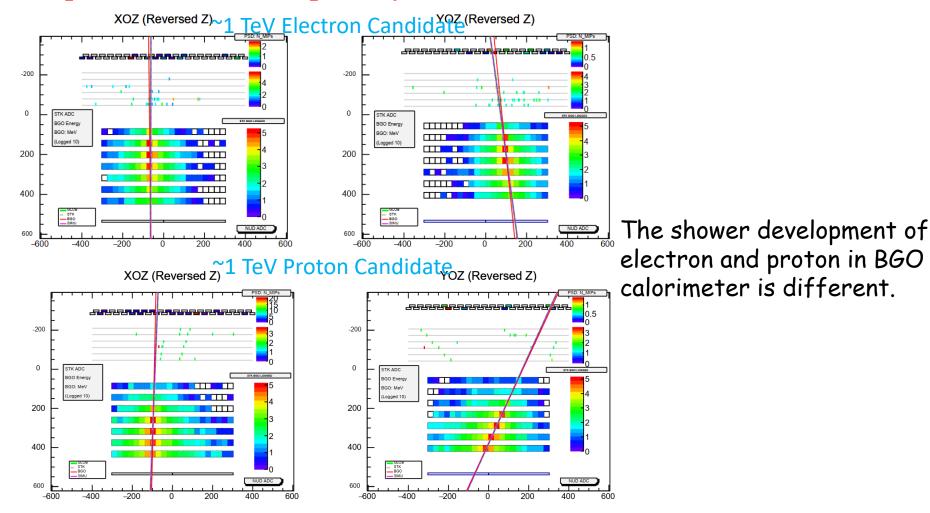


Dynode ratio stability



Most channels of PMT gains have slightly increased/decreased for several percent in whole operation time

e/p discrimination capability



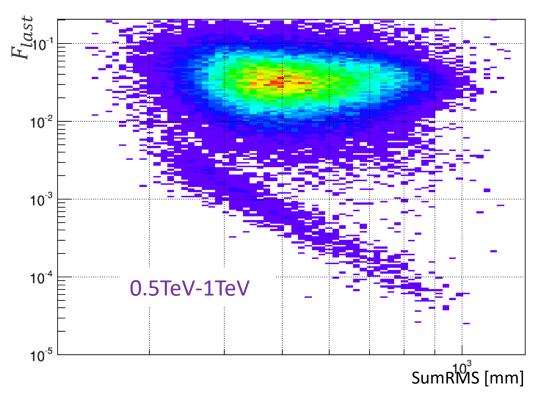
e/p discrimination capability

 F_{last} represents the ratio of energy deposited in the last BGO layer that has energy to the total energy deposited in the BGO calorimeter.

SumRMS (shower spread) is defined as the summation of the energy-weighted shower dispersion of each layer.

$$\zeta = (SumRMS)^4 \times F_{last} / (8 \times 10^6)$$

$$SumRM = \sum_{i} RMS_{i}$$
$$RMS_{i} = \sqrt{\frac{\sum_{j} (x_{i,j} - COG_{i})^{2} * Energy_{i,j}}{\sum_{j} Energy_{i,j}}}$$

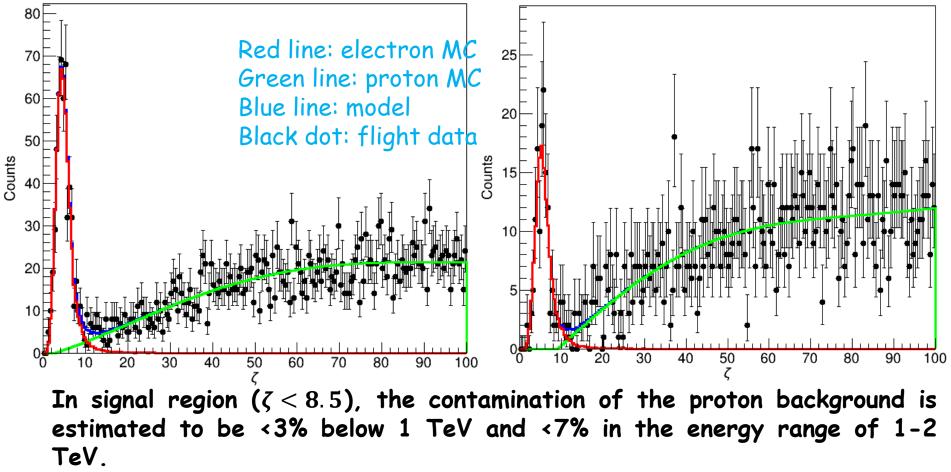


The "relative gap" between the electron and proton distributions in the right figure, it is clear that a curved 'cut line' is required.

Excellent electron/proton discrimination capability with ζ variable

e/p discrimination capability 575.4 GeV – 660.7 GeV

1000.0 GeV - 1148.2 GeV



Summary



- DArk Matter Particle Explorer (DAMPE) has been working very well since launched successfully on Dec. 17th, 2015. 20,000 Science Advances 5,9: eaax3793 (2019)
- Status of BGO Calorimeter is very stable on orbit.
 - Pedestal

Thank you

- **MIP** response
- PMT Dynode ratio
- e/p discrimination capability

E³ × Flux (m⁻² s⁻¹ sr¹ GeV²)

250

200

150

100

50

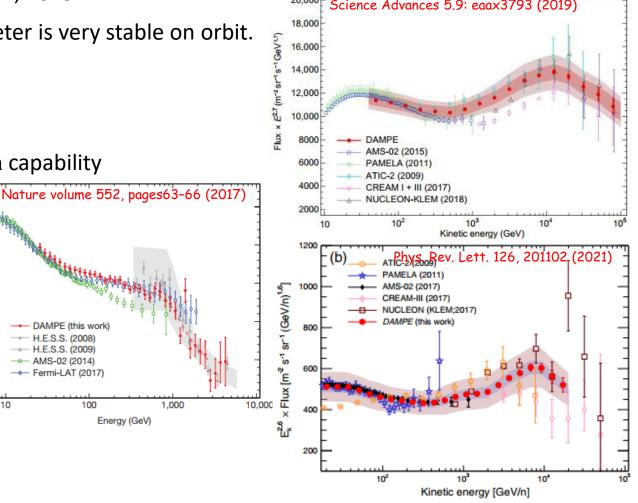
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DAMPE (this work) H.E.S.S. (2008) H.E.S.S. (2009)

AMS-02 (2014)

Fermi-LAT (2017)

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