





# Ab Initio Coupling of Jets to Collective Flow in the Opacity Expansion Approach

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# **Jet Tomography**

- Jets are main available probes of the matter in HIC;
- The hot nuclear matter in HIC undergoes multi-phase evolution and its details are hard to access through the soft sector. In turn, jets see the matter at multiple scales, and essentially X-ray it;
- However, most approaches to the jet-medium interaction are either empirical or based on multiple simplifying assumptions – static matter, no fluctuations, etc;
- In what follows I will highlight our recent progress on the medium motion effects in the QCD calculations for jet broadening and gluon emission;
- The developed formalism can be also applied to include orbital motion of nucleons and some of in-medium fluctuations (e.g. spatial inhomogeneities) in the DIS context;





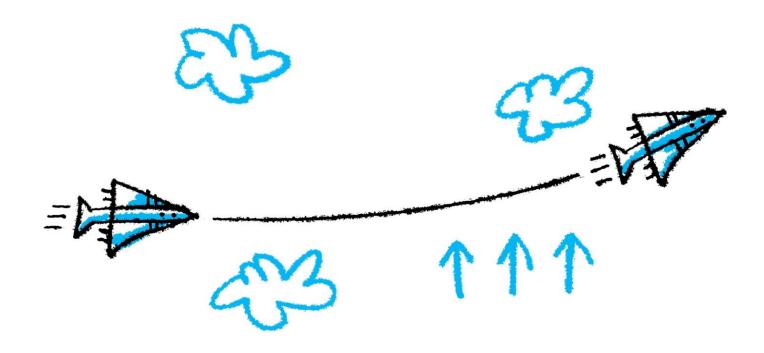






### **Jets**

Does a jet feel the flow?





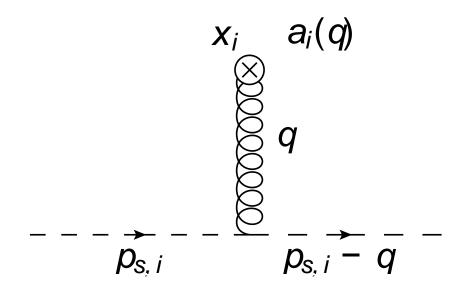








### **Color Potential**



$$a_i^{\mu a}(q) = (ig\ t_i^a)\ (2p_{s\,i} - q)_\nu \left(\frac{-ig^{\mu\nu}}{q^2 - \mu^2 + i\epsilon}\right) \ (2\pi)\ \delta\Big((p_{s\,i} - q)^2 - M^2\Big)$$
 
$$v(q^2) \text{ - the Gyulassy-Wang potential}$$

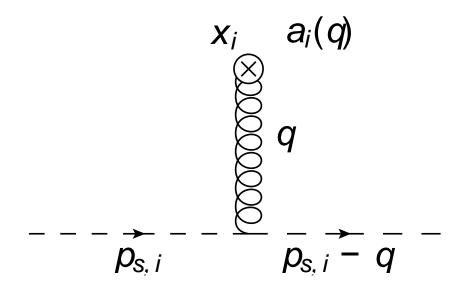








### **Color Potential**



$$a_i^{\mu a}(q) = g \, t_i^a \, \left( \frac{u_i^\mu}{q^2 - \mu^2 + i\epsilon} \right) \, (2\pi) \, \delta \left( q^0 - \vec{u}_i \cdot \vec{q} \right)$$

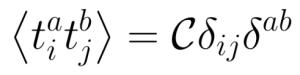
finite energy transfer!!!





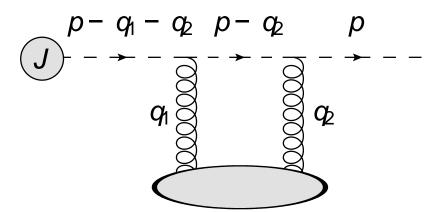






color neutrality

### **Jet Broadening**



#### the static case:

$$E\frac{dN^{(1)}}{d^3p} = \int dz\,d^2q_\perp\,\rho(z)\,\frac{d\sigma}{d^2q_\perp} \left[\,\left(E\frac{dN^{(0)}}{d^2(p-q)_\perp\,dE}\right) - \left(E\frac{dN^{(0)}}{d^2p_\perp\,dE}\right)\,\right]$$

elastic scattering cross section









#### with flow:

$$E\frac{dN^{(1)}}{d^{3}p} = \int dz \, d^{2}q_{\perp} \, \rho(z) \, \frac{d\sigma}{d^{2}q_{\perp}} \left[ \left( E\frac{dN^{(0)}}{d^{2}(p-q)_{\perp} \, dE} \right) \left( 1 + \vec{u}_{\perp}(z) \cdot \vec{\Gamma}(\vec{q}_{\perp}) \right) - \left( E\frac{dN^{(0)}}{d^{2}p_{\perp} \, dE} \right) \left( 1 + \vec{u}_{\perp}(z) \cdot \vec{\Gamma}_{DB}(\vec{q}_{\perp}) \right) \right] + \mathcal{O}\left(\partial_{\perp}\right)$$

$$\vec{\Gamma}(\vec{q}_{\perp}) = -2\frac{\vec{p}_{\perp} - \vec{q}_{\perp}}{(1 - u_{iz})E} + \frac{2\vec{q}_{\perp}}{(1 - u_{iz})E} \left(\frac{(p - q)_{\perp}^2 - p_{\perp}^2}{v(q_{\perp}^2)}\right) \frac{\partial v}{\partial q_{\perp}^2} - \frac{\vec{q}_{\perp}}{1 - u_z} \left(\frac{1}{\bar{N}_0(E, \vec{p}_{\perp} - \vec{q}_{\perp})} \frac{\partial \bar{N}_0}{\partial E}\right)$$









#### An illustration:

assuming a model source

$$E rac{dN^{(0)}}{d^3p} \propto \delta^{(2)}(\vec{p}_\perp)$$

and ignoring z-dependence, we find

the distribution in energies

$$\langle \boldsymbol{p}_{\perp}(p_{\perp}^2)^k \rangle = -\frac{\boldsymbol{u}_{\perp}}{(1-u_z)} \frac{L}{2\lambda} (\mu^2)^{k+1} \bigg( -\frac{2}{E} \int_0^{\infty} d\xi \frac{\xi^{k+2}}{(1+\xi)^3} + \frac{1}{f(E)} \frac{\partial f}{\partial E} \int_0^{\infty} d\xi \frac{\xi^{k+1}}{(1+\xi)^2} \bigg)$$

while even moments are unmodified

A family of velocity-sensitive observables!









#### An illustration:

assuming a model source

$$E \frac{dN^{(0)}}{d^3p} \propto E^{-4} \, \delta^{(2)}(\vec{p}_{\perp})$$

and ignoring z-dependence, we find

$$\left\langle \frac{\vec{p}_{\perp}}{p_{\perp}^{2}} \right\rangle = \frac{5}{2} \frac{\vec{u}_{\perp}}{(1 - u_{z})} \frac{L}{\lambda} \frac{1}{E} \quad , \quad \frac{\left\langle \vec{p}_{\perp} \right\rangle}{\mu^{2}} = 3 \frac{\vec{u}_{\perp}}{(1 - u_{z})} \frac{L}{\lambda} \frac{1}{E} \ln \frac{E}{\mu}$$

while even moments are unmodified

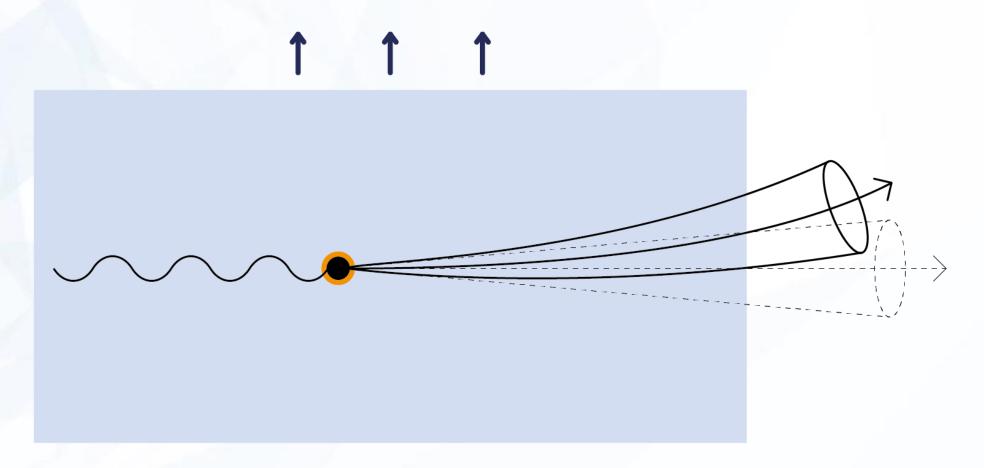
A family of velocity-sensitive observables!



















gradient expansion

$$E\frac{dN^{(0)}}{d^{3}p} = \frac{1}{2(2\pi)^{3}}|J(p)|^{2} = \frac{f(E)}{2\pi w^{2}}e^{-\frac{p_{\perp}^{2}}{2w^{2}}}$$

$$\downarrow \qquad \qquad \qquad \downarrow$$

$$\langle \vec{p}_{\perp} \, p_{\perp}^{2} \rangle^{(linear)} \simeq \frac{L}{\lambda} \frac{L}{E} \, w^{2} \mu^{2} \frac{\vec{\nabla}_{\perp} \rho}{\rho} \, \ln \frac{E}{\mu}$$

the leading gradient correction at the first order in opacity









# Summary

- We have constructed a generalization of the GLV approach which includes the medium motion effects. With this tool one can study general flow, temperature, and source density profiles in the context of HIC;
- It is shown that the odd moments of the jet momentum are modified by the medium motion, and the jet is bended by the flow and gradients;
- We have also derived the gluon emission spectrum in the case of uniformly flowing matter, which cannot be discussed in details in a short talk (see the paper);
- In the context of DIS our formalism can be used to study nucleon orbital motion and spatial inhomogeneities in the system (a relation to GPDs and TMDs?);



