

# HIGGS PHYSICS AT FUTURE COLLIDERS

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Based mostly on:  
1811.00017 D. Egaña, S. Homiller, PM  
1908.11376 D. Egaña, S. Homiller, PM  
+ work in progress 1910.????x2

# OUTLINE

- Motivation\*
- BSM Flavor physics - Spontaneous Flavor Violation (SFV)
- Higgs and Flavor Physics
- Di-Higgs
- Other BSM SFV examples

LATELY MANY PHENO SEMINARS  
START WITH SOMETHING LIKE...



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START WITH SOMETHING LIKE...

ATLAS SUSY Searches* - 95% CL Lower Limits									
July 2019									
Model		Signature		$\int \mathcal{L} dt$ [fb <sup>-1</sup> ]		Mass limit		Reference	
Inclusive Searches	$\tilde{q}\tilde{q}, \tilde{q} \rightarrow q\tilde{\chi}_1^0$	0 $e, \mu$	2-6 jets	$E_T^{\text{miss}}$	36.1	$\tilde{q}$ [2x, 8x Degen.]	0.9	$m(\tilde{\chi}_1^0) < 100$ GeV	1712.02332
		mono-jet	1-3 jets	$E_T^{\text{miss}}$	36.1	$\tilde{q}$ [1x, 8x Degen.]	0.43 0.71	$m(\tilde{q}) - m(\tilde{\chi}_1^0) = 5$ GeV	1711.03301
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\tilde{q}\tilde{\chi}_1^0$	0 $e, \mu$	2-6 jets	$E_T^{\text{miss}}$	36.1	$\tilde{g}$	2.0	$m(\tilde{\chi}_1^0) < 200$ GeV	1712.02332
						$\tilde{g}$	Forbidden 0.95-1.6	$m(\tilde{\chi}_1^0) = 900$ GeV	1712.02332
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\tilde{q}(\ell\ell)\tilde{\chi}_1^0$	3 $e, \mu$	4 jets	$E_T^{\text{miss}}$	36.1	$\tilde{g}$	1.85	$m(\tilde{\chi}_1^0) < 800$ GeV	1706.03731
		$ee, \mu\mu$	2 jets	$E_T^{\text{miss}}$	36.1	$\tilde{g}$	1.2	$m(\tilde{g}) - m(\tilde{\chi}_1^0) = 50$ GeV	1805.11381
$\tilde{\chi}_1^0$ gen. squarks direct production	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\tilde{q}WZ\tilde{\chi}_1^0$	0 $e, \mu$	7-11 jets	$E_T^{\text{miss}}$	36.1	$\tilde{g}$	1.8	$m(\tilde{\chi}_1^0) < 400$ GeV	1708.02794
		SS $e, \mu$	6 jets	$E_T^{\text{miss}}$	139	$\tilde{g}$	1.15	$m(\tilde{g}) - m(\tilde{\chi}_1^0) = 200$ GeV	ATLAS-CONF-2019-015
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow t\tilde{t}\tilde{\chi}_1^0$	0-1 $e, \mu$	3 $b$	$E_T^{\text{miss}}$	79.8	$\tilde{g}$	2.25	$m(\tilde{\chi}_1^0) < 200$ GeV	ATLAS-CONF-2018-041
		SS $e, \mu$	6 jets	$E_T^{\text{miss}}$	139	$\tilde{g}$	1.25	$m(\tilde{g}) - m(\tilde{\chi}_1^0) = 300$ GeV	ATLAS-CONF-2019-015
	$\tilde{b}_1\tilde{b}_1, \tilde{b}_1 \rightarrow b\tilde{\chi}_1^0 / t\tilde{\chi}_1^\pm$	Multiple	Multiple	$E_T^{\text{miss}}$	36.1	$\tilde{b}_1$	Forbidden 0.9	$m(\tilde{\chi}_1^0) = 300$ GeV, $\text{BR}(b\tilde{\chi}_1^0) = 1$	1708.09266, 1711.03301
		Multiple	Multiple	$E_T^{\text{miss}}$	36.1	$\tilde{b}_1$	Forbidden 0.58-0.82	$m(\tilde{\chi}_1^0) = 300$ GeV, $\text{BR}(b\tilde{\chi}_1^0) = \text{BR}(t\tilde{\chi}_1^\pm) = 0.5$	1708.09266
EW direct		Multiple	Multiple	$E_T^{\text{miss}}$	139	$\tilde{b}_1$	Forbidden 0.74	$m(\tilde{\chi}_1^0) = 200$ GeV, $m(\tilde{\chi}_1^\pm) = 300$ GeV, $\text{BR}(t\tilde{\chi}_1^\pm) = 1$	ATLAS-CONF-2019-015
	$\tilde{b}_1\tilde{b}_1, \tilde{b}_1 \rightarrow b\tilde{\chi}_2^0 \rightarrow b h\tilde{\chi}_1^0$	0 $e, \mu$	6 $b$	$E_T^{\text{miss}}$	139	$\tilde{b}_1$	Forbidden 0.23-1.35	$\Delta m(\tilde{\chi}_2^0, \tilde{\chi}_1^0) = 130$ GeV, $m(\tilde{\chi}_1^0) = 100$ GeV	SUSY-2018-31
						$\tilde{b}_1$	0.23-0.48	$\Delta m(\tilde{\chi}_2^0, \tilde{\chi}_1^0) = 130$ GeV, $m(\tilde{\chi}_1^0) = 0$ GeV	SUSY-2018-31
	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow Wb\tilde{\chi}_1^0$ or $t\tilde{\chi}_1^0$	0-2 $e, \mu$	0-2 jets/1-2 $b$	$E_T^{\text{miss}}$	36.1	$\tilde{t}_1$	1.0	$m(\tilde{\chi}_1^0) = 1$ GeV	1506.08616, 1709.04183, 1711.11520
	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow Wb\tilde{\chi}_1^0$	1 $e, \mu$	3 jets/1 $b$	$E_T^{\text{miss}}$	139	$\tilde{t}_1$	0.44-0.59	$m(\tilde{\chi}_1^0) = 400$ GeV	ATLAS-CONF-2019-017
	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow \tilde{\tau}_1 b\nu, \tilde{\tau}_1 \rightarrow \tau\tilde{G}$	1 $\tau + 1 e, \mu, \tau$	2 jets/1 $b$	$E_T^{\text{miss}}$	36.1	$\tilde{t}_1$	1.16	$m(\tilde{\tau}_1) = 800$ GeV	1803.10178
Long-lived particles	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow c\tilde{\chi}_1^0 / \tilde{c}\tilde{c}, \tilde{c} \rightarrow c\tilde{\chi}_1^0$	0 $e, \mu$	2 $c$	$E_T^{\text{miss}}$	36.1	$\tilde{t}_1$	0.85	$m(\tilde{\chi}_1^0) = 0$ GeV	1805.01649
						$\tilde{t}_1$	0.46	$m(\tilde{t}_1, \tilde{c}) - m(\tilde{\chi}_1^0) = 50$ GeV	1805.01649
						$\tilde{t}_1$	0.43	$m(\tilde{t}_1, \tilde{c}) - m(\tilde{\chi}_1^0) = 5$ GeV	1711.03301
	$\tilde{t}_2\tilde{t}_2, \tilde{t}_2 \rightarrow \tilde{t}_1 + h$	1-2 $e, \mu$	4 $b$	$E_T^{\text{miss}}$	36.1	$\tilde{t}_2$	0.32-0.88	$m(\tilde{\chi}_1^0) = 0$ GeV, $m(\tilde{t}_1) - m(\tilde{\chi}_1^0) = 180$ GeV	1706.03986
	$\tilde{t}_2\tilde{t}_2, \tilde{t}_2 \rightarrow \tilde{t}_1 + Z$	3 $e, \mu$	1 $b$	$E_T^{\text{miss}}$	139	$\tilde{t}_2$	Forbidden 0.86	$m(\tilde{\chi}_1^0) = 360$ GeV, $m(\tilde{t}_1) - m(\tilde{\chi}_1^0) = 40$ GeV	ATLAS-CONF-2019-016
RPV	$\tilde{\chi}_1^\pm \tilde{\chi}_2^0$ via WZ	2-3 $e, \mu$	$\geq 1$	$E_T^{\text{miss}}$	36.1	$\tilde{\chi}_1^\pm / \tilde{\chi}_2^0$	0.6	$m(\tilde{\chi}_1^0) = 0$	1403.5294, 1806.02293
		$ee, \mu\mu$	$\geq 1$	$E_T^{\text{miss}}$	139	$\tilde{\chi}_1^\pm / \tilde{\chi}_2^0$	0.205	$m(\tilde{\chi}_1^\pm) - m(\tilde{\chi}_1^0) = 5$ GeV	ATLAS-CONF-2019-014
	$\tilde{\chi}_1^\pm \tilde{\chi}_1^\mp$ via WW	2 $e, \mu$		$E_T^{\text{miss}}$	139	$\tilde{\chi}_1^\pm$	0.42	$m(\tilde{\chi}_1^0) = 0$	ATLAS-CONF-2019-008
	$\tilde{\chi}_1^\pm \tilde{\chi}_2^0$ via Wh	0-1 $e, \mu$	2 $b/2 \gamma$	$E_T^{\text{miss}}$	139	$\tilde{\chi}_1^\pm / \tilde{\chi}_2^0$	0.74	$m(\tilde{\chi}_1^0) = 70$ GeV	ATLAS-CONF-2019-019, ATLAS-CONF-2019-XYZ
	$\tilde{\chi}_1^\pm \tilde{\chi}_1^\mp$ via $\tilde{\ell}_L/\tilde{\nu}$	2 $e, \mu$		$E_T^{\text{miss}}$	139	$\tilde{\chi}_1^\pm$	1.0	$m(\tilde{\ell}, \tilde{\nu}) = 0.5(m(\tilde{\chi}_1^\pm) + m(\tilde{\chi}_1^0))$	ATLAS-CONF-2019-008
	$\tilde{\tau}\tilde{\tau}, \tilde{\tau} \rightarrow \tau\tilde{\chi}_1^0$	2 $\tau$		$E_T^{\text{miss}}$	139	$\tilde{\tau}$ [ $\tilde{\tau}_L, \tilde{\tau}_{R,L}$ ]	0.16-0.3 0.12-0.39	$m(\tilde{\chi}_1^0) = 0$	ATLAS-CONF-2019-018
RPV	$\tilde{\ell}_{L,R}\tilde{\ell}_{L,R}, \tilde{\ell} \rightarrow \ell\tilde{\chi}_1^0$	2 $e, \mu$	0 jets	$E_T^{\text{miss}}$	139	$\tilde{\ell}$	0.7	$m(\tilde{\chi}_1^0) = 0$	ATLAS-CONF-2019-008
		$\geq 1$		$E_T^{\text{miss}}$	139	$\tilde{\ell}$	0.256	$m(\tilde{\ell}) - m(\tilde{\chi}_1^0) = 10$ GeV	ATLAS-CONF-2019-014
	$\tilde{H}\tilde{H}, \tilde{H} \rightarrow h\tilde{G}/Z\tilde{G}$	0 $e, \mu$	$\geq 3 b$	$E_T^{\text{miss}}$	36.1	$\tilde{H}$	0.13-0.23	$\text{BR}(\tilde{\chi}_1^0 \rightarrow h\tilde{G}) = 1$	1806.04030
		4 $e, \mu$	0 jets	$E_T^{\text{miss}}$	36.1	$\tilde{H}$	0.3	$\text{BR}(\tilde{\chi}_1^0 \rightarrow Z\tilde{G}) = 1$	1804.03602
RPV	Direct $\tilde{\chi}_1^\pm \tilde{\chi}_1^\mp$ prod., long-lived $\tilde{\chi}_1^\pm$	Disapp. trk	1 jet	$E_T^{\text{miss}}$	36.1	$\tilde{\chi}_1^\pm$	0.15 0.46	Pure Wino	1712.02118
						$\tilde{\chi}_1^\pm$		Pure Higgsino	ATL-PHYS-PUB-2017-019
	Stable $\tilde{g}$ R-hadron		Multiple		36.1	$\tilde{g}$	2.0		1902.01636, 1808.04095
	Metastable $\tilde{g}$ R-hadron, $\tilde{g} \rightarrow q\tilde{q}\tilde{\chi}_1^0$		Multiple		36.1	$\tilde{g}$ [ $\tau(\tilde{g}) = 10$ ns, 0.2 ns]	2.05 2.4	$m(\tilde{\chi}_1^0) = 100$ GeV	1710.04901, 1808.04095
RPV	LFV $pp \rightarrow \tilde{\nu}_\tau + X, \tilde{\nu}_\tau \rightarrow e\mu/\tau\mu/\tau\tau$	$e\mu, e\tau, \mu\tau$			3.2	$\tilde{\nu}_\tau$	1.9	$\lambda'_{311} = 0.11, \lambda'_{132/133/233} = 0.07$	1607.08079
	$\tilde{\chi}_1^\pm \tilde{\chi}_1^\mp / \tilde{\chi}_2^0 \rightarrow WW/Z\ell\ell\ell\nu\nu$	4 $e, \mu$	0 jets	$E_T^{\text{miss}}$	36.1	$\tilde{\chi}_1^\pm / \tilde{\chi}_2^0$ [ $\lambda'_{133} \neq 0, \lambda'_{12k} \neq 0$ ]	0.82 1.33	$m(\tilde{\chi}_1^0) = 100$ GeV	1804.03602
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\tilde{q}\tilde{\chi}_1^0, \tilde{\chi}_1^0 \rightarrow q\tilde{q}q$	4-5 large- $R$ jets			36.1	$\tilde{g}$ [ $m(\tilde{\chi}_1^0) = 200$ GeV, 1100 GeV]	1.3 1.9	Large $\lambda'_{12}$	1804.03568
		Multiple			36.1	$\tilde{g}$ [ $\lambda'_{112} = 2e-4, 2e-5$ ]	1.05 2.0	$m(\tilde{\chi}_1^0) = 200$ GeV, bino-like	ATLAS-CONF-2018-003
	$\tilde{t}\tilde{t}, \tilde{t} \rightarrow t\tilde{\chi}_1^0, \tilde{\chi}_1^0 \rightarrow tbs$	Multiple			36.1	$\tilde{g}$ [ $\lambda'_{323} = 2e-4, 1e-2$ ]	0.55 1.05	$m(\tilde{\chi}_1^0) = 200$ GeV, bino-like	ATLAS-CONF-2018-003
	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow bs$	2 jets + 2 $b$			36.7	$\tilde{t}_1$ [ $qq, bs$ ]	0.42 0.61		1710.07171
RPV	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow q\ell$	2 $e, \mu$	2 $b$		36.1	$\tilde{t}_1$	0.4-1.45	$\text{BR}(\tilde{t}_1 \rightarrow be/b\mu) > 20\%$	1710.05544
		1 $\mu$	DV		136	$\tilde{t}_1$ [ $1e-10 < \lambda'_{23k} < 1e-8, 3e-10 < \lambda'_{23k} < 3e-9$ ]	1.0 1.6	$\text{BR}(\tilde{t}_1 \rightarrow q\mu) = 100\%, \cos\theta = 1$	ATLAS-CONF-2019-006

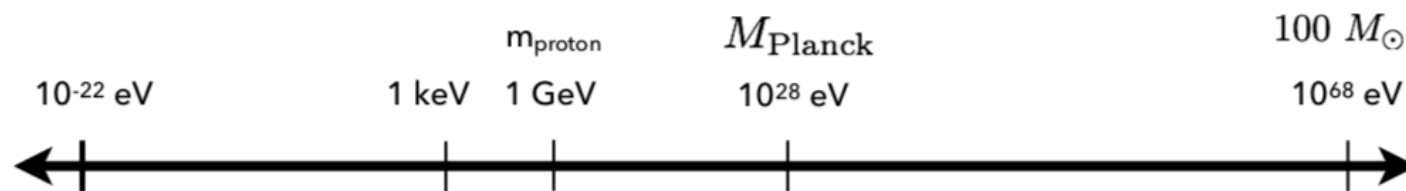
\*Only a selection of the available mass limits on new states or phenomena is shown. Many of the limits are based on simplified models, c.f. refs. for the assumptions made.



AND THEN THERE'S A PIVOT

# AND THEN THERE'S A PIVOT

Where look for DM?



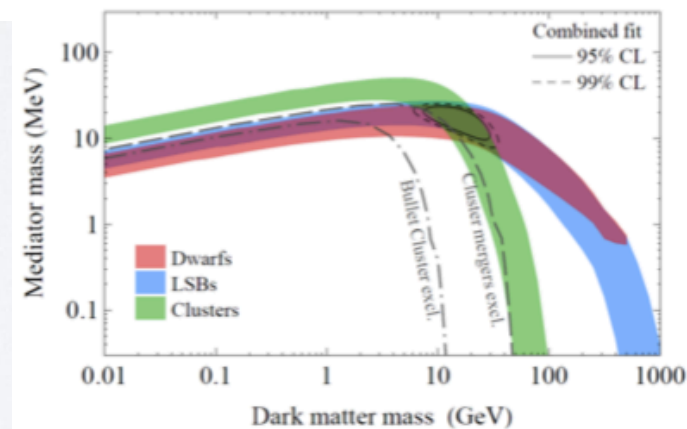
many possibilities

⇒ target motivated areas

## sub GeV-SCALE MEDIATORS & LIGHT DM EXPERIMENTS & OBSERVATIONS

Dark matter self interactions

M. Kaplinghat, S. Tullin, H.-B. Yu, 1508.03339



Search for light mediators at colliders

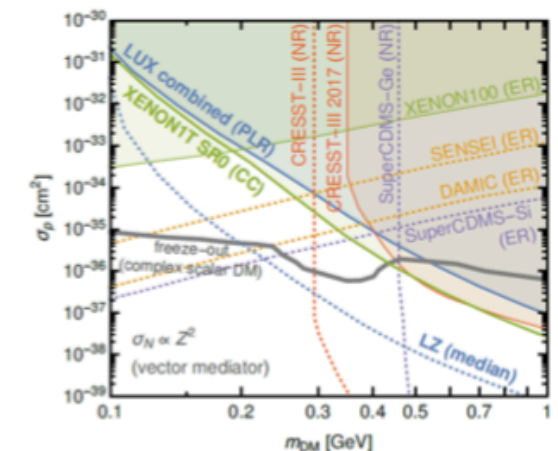


It also e.g. NA62

and many proposed expts e.g. Codex-b, MATHUSLA, SHiP, ...

Light DM direct detection

M. J. Dolan, F. Kahlhoefer, C. McCabe, 1711.09906



& other...

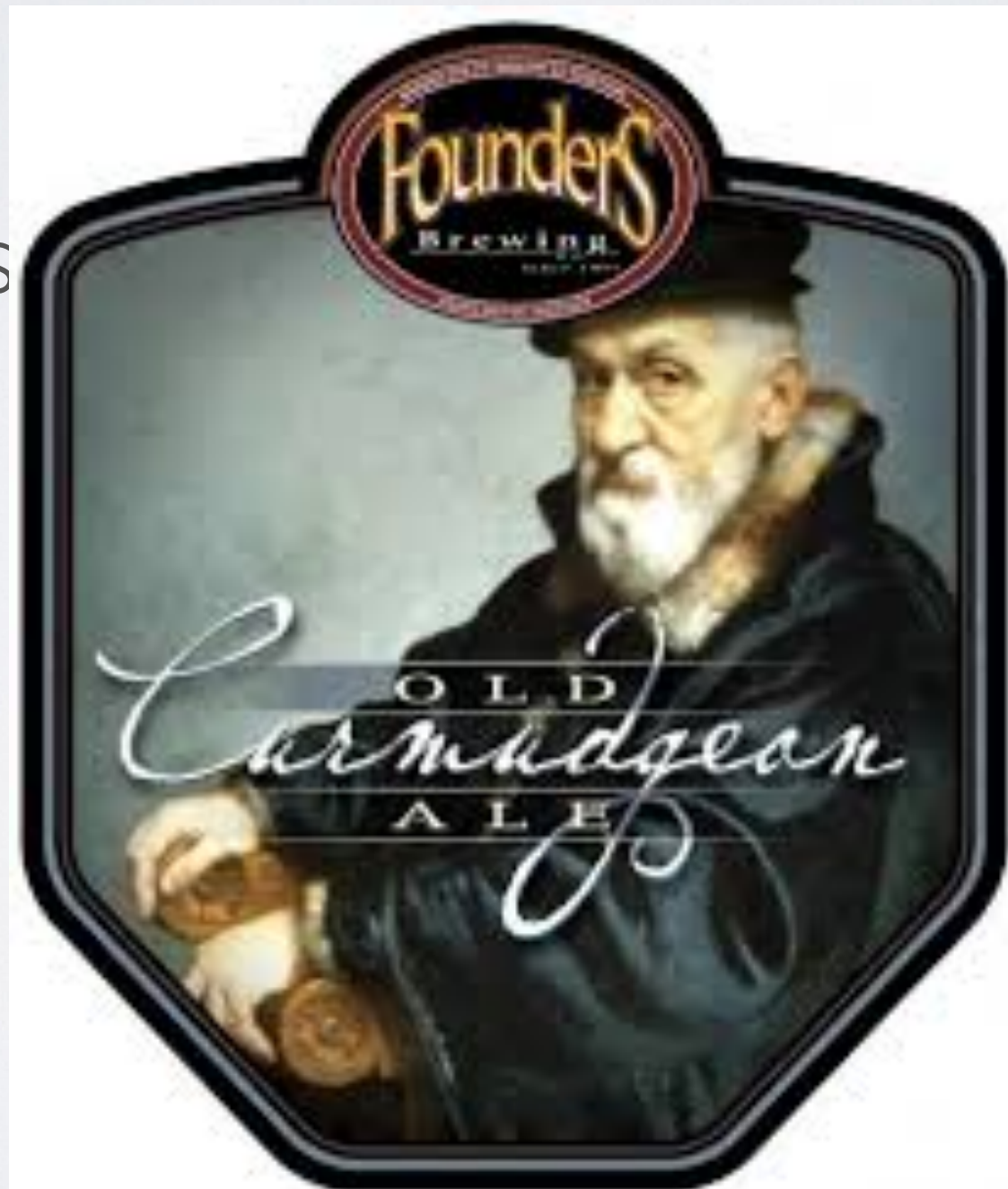
# THERE'S NOTHING WRONG WITH A PIVOT...

- I've spent a large part of my career thinking about crazy things...



# THERE'S NOTHING WRONG WITH A PIVOT...

- I've spent a lot of time thinking about crazy things



# THERE'S NOTHING WRONG WITH A PIVOT...

- I've spent a large part of my career thinking about crazy things...
- In this talk I want to stress that the **Higgs** - something we know is there - really can unlock so many puzzles about the SM qualitatively by studying it with more quantitative precision



**PREACHING TO THE  
CHOIR**



**YOU ARE**



NEVERTHELESS...

# NEVERTHELESS...

- I think it's important to go through these qualitative questions related to the Higgs to understand their implications and where we might be biased
- Despite the fact that many/all of you will be familiar with the qualitative questions, there are *new possibilities* I want to emphasize in this talk

# NEVERTHELESS...

- I think it's important to go through these qualitative questions related to the Higgs to understand their implications and where we might be biased
- Despite the fact that many/all of you will be familiar with the qualitative questions, there are *new possibilities* I want to emphasize in this talk

In the end I'll focus on the **craziest**, but I'm around through tomorrow to talk about any of these...



LET'S START WITH THE BASICS

# THE HIGGS AND SPONTANEOUS SYMMETRY BREAKING

Long before the Higgs was discovered we knew the SM was described by a spontaneously broken gauge symmetry

$$SU(3) \times SU(2) \times U(1)_Y$$



$$SU(3) \times U(1)_Q$$

THE HIGGS WASN'T THE ONLY  
GAME IN TOWN IN PRINCIPLE...

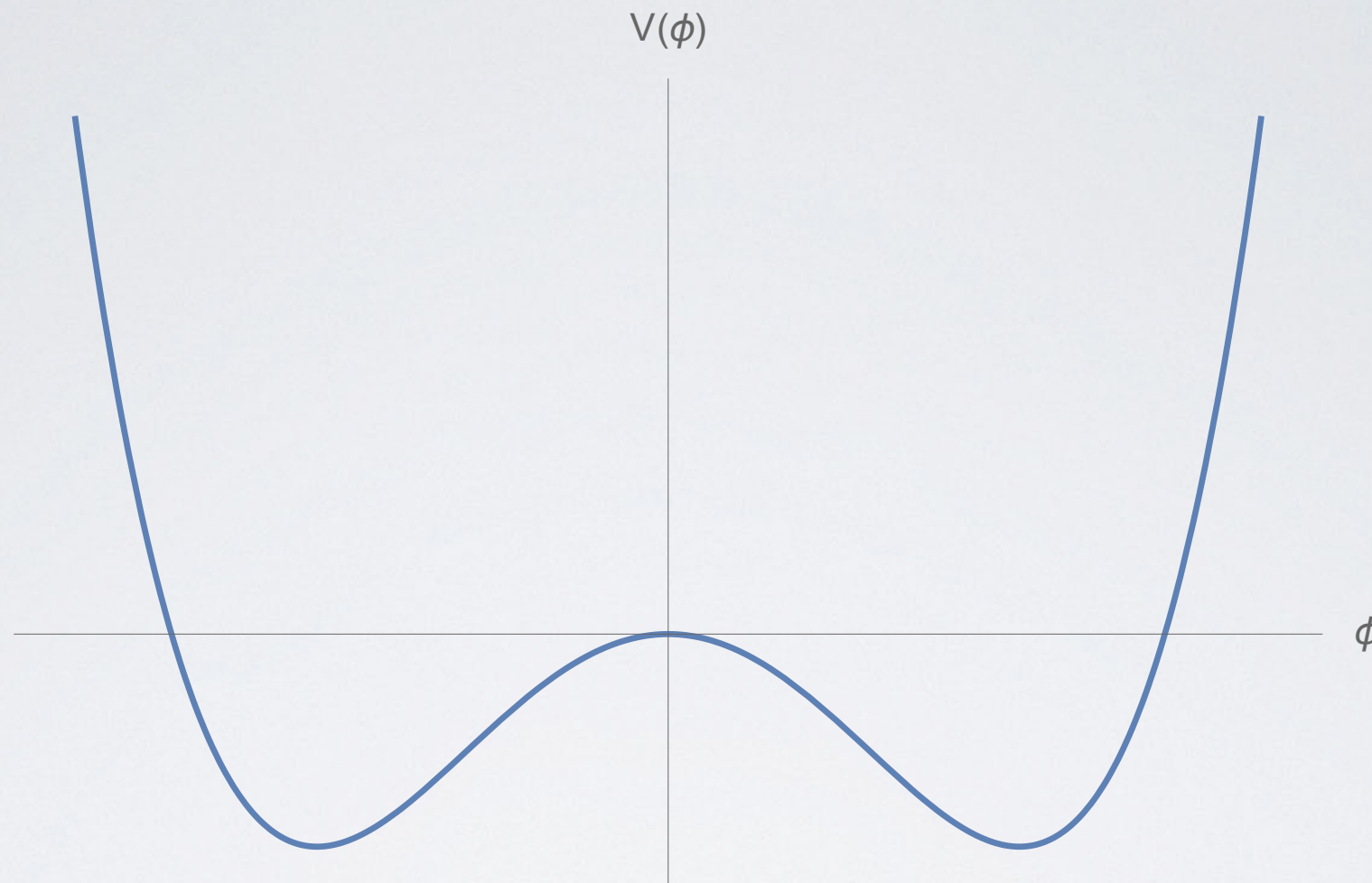


# THE HIGGS WASN'T THE ONLY GAME IN TOWN IN PRINCIPLE...

Dynamical Symmetry Breaking has shown up before...

It explains why the symmetry is broken

# INSTEAD WE GOT...



$$V(\phi) = -\frac{\mu^2}{2}\phi^2 + \frac{\lambda}{4}\phi^4$$

EWSB put in **by hand**

A light fundamental scalar, never seen before in nature without tuning!

# NATURALNESS

$$m_h^2 \sim \Lambda^2$$

$$\Lambda \sim \cancel{100} \text{ GeV}$$

$$\cancel{200} \text{ GeV}$$

$$\cancel{300} \text{ GeV}$$

$$\cancel{500} \text{ GeV}$$

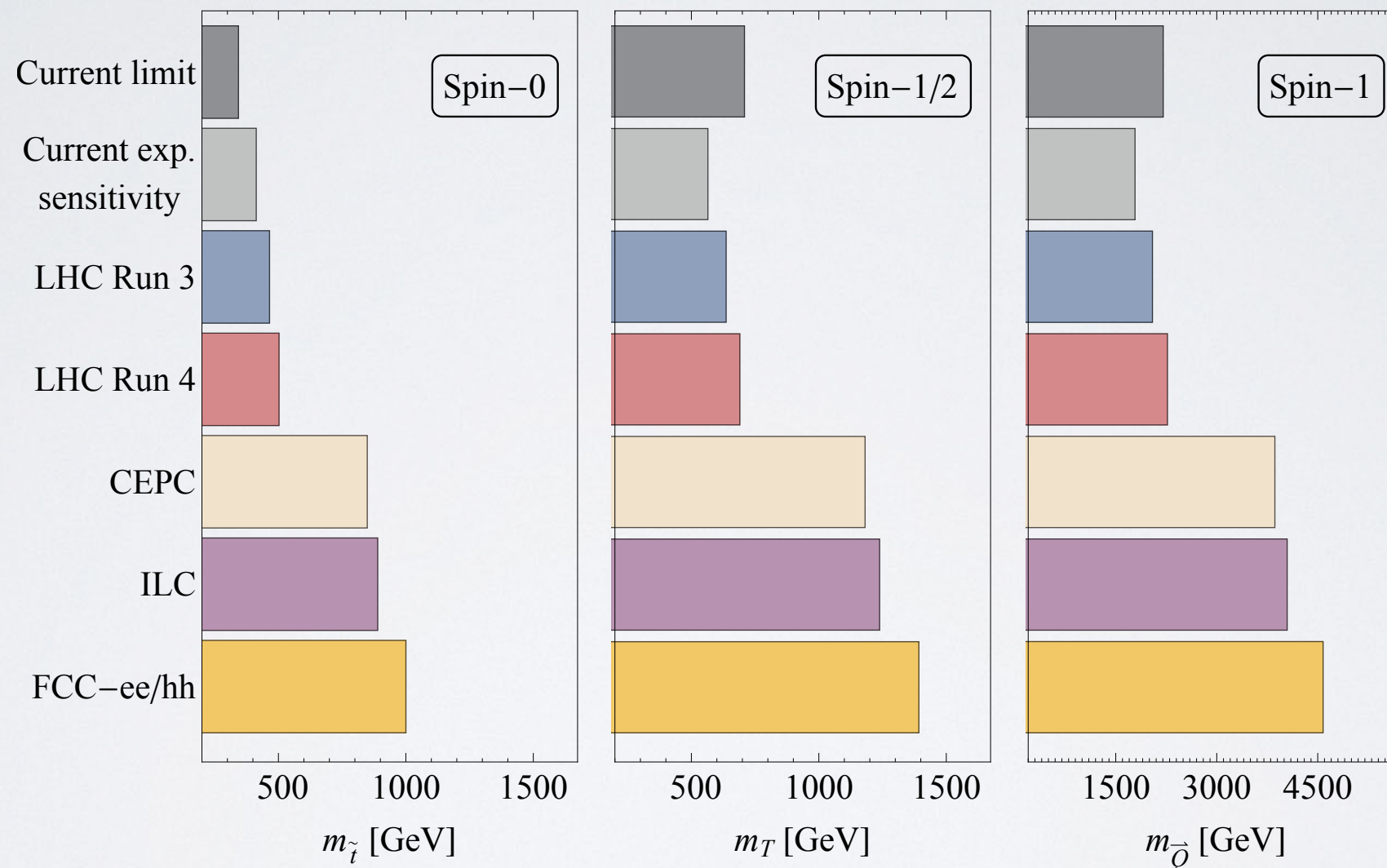
$$\cancel{1000} \text{ GeV}$$

$$5000 \text{ GeV} \quad \text{????}$$

We certainly learn something,  
but what it is telling us isn't as clear



# THIS WILL OBVIOUSLY BE A FOCUS FOR FUTURE COLLIDERS

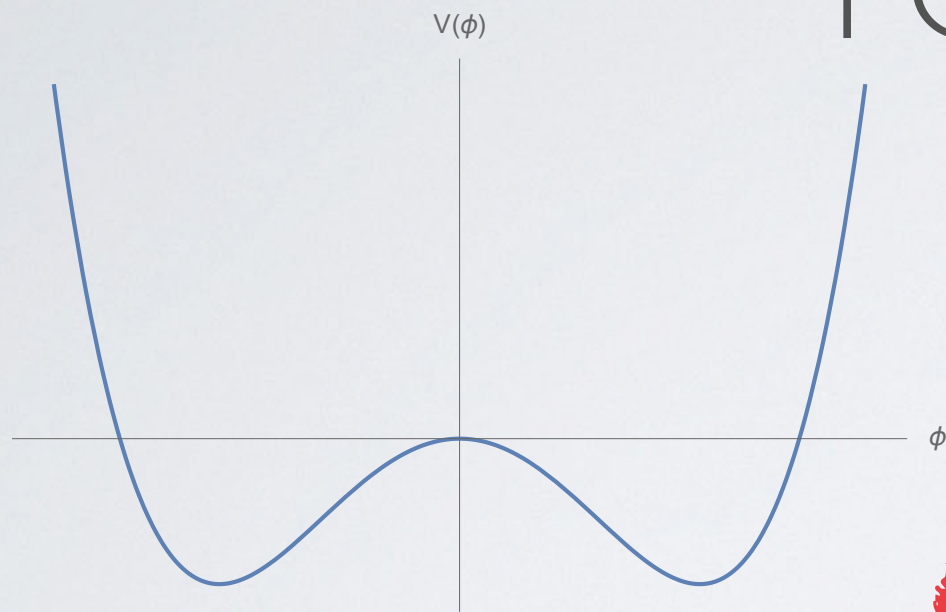


Essig, PM, Ramani, Zhong 1707.03399

IS THE HIGGS PART OF SOME LARGER SECTOR  
WHICH “EXPLAINS” EWSB?

IS THERE SOMETHING MAKING THE HIGGS  
“NATURAL”?

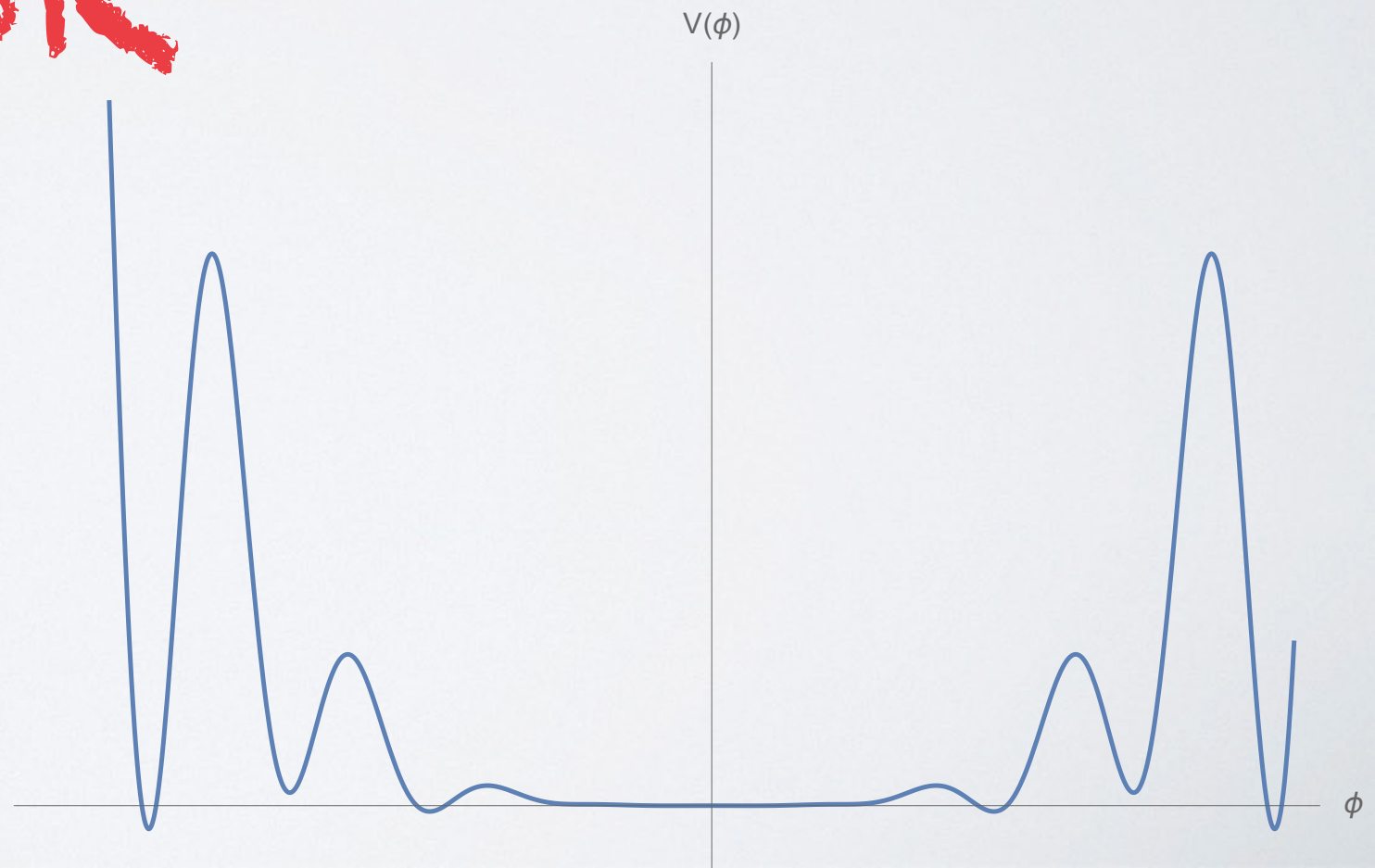
# A MORE FUNDAMENTAL EXPERIMENTAL QUESTION, WHAT IS THE HIGGS POTENTIAL?



$$\left. \frac{\partial V(\phi)}{\partial \phi} \right|_{\phi=v} = 0$$

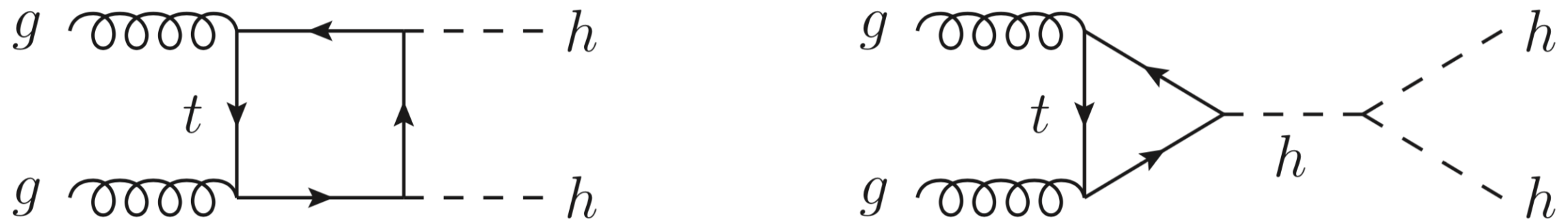
$$\left. \frac{\partial^2 V(\phi)}{\partial \phi^2} \right|_{\phi=v} = m_h^2$$

OR





# NEXT UP IS THE TRIPLE HIGGS COUPLING IN THE SM...

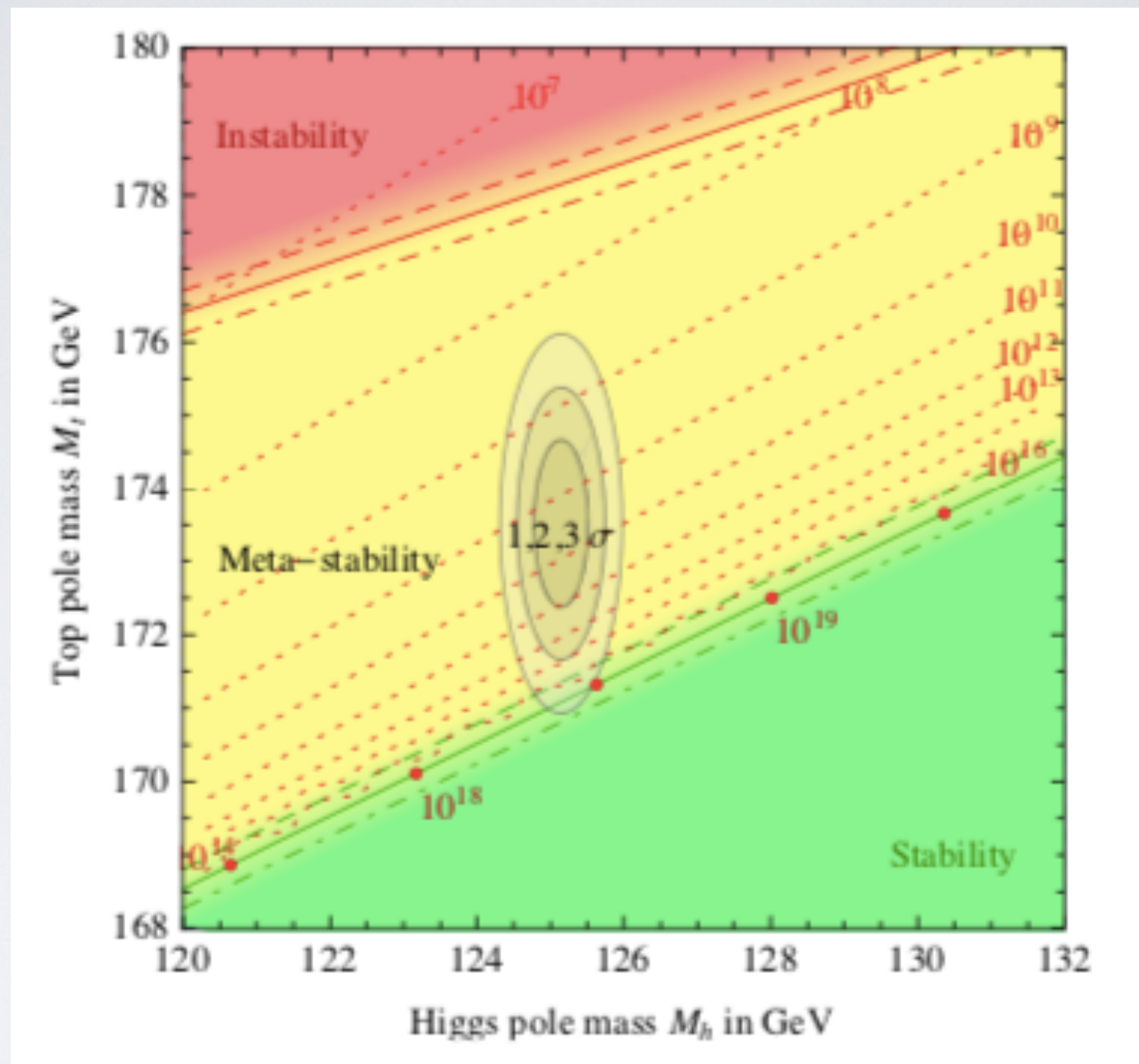


Unfortunately it's very difficult and it interferes with itself

***However, just measuring the SM value  
would be seeing something qualitatively new!***

To go beyond though it though  
how precisely do we need to measure it?

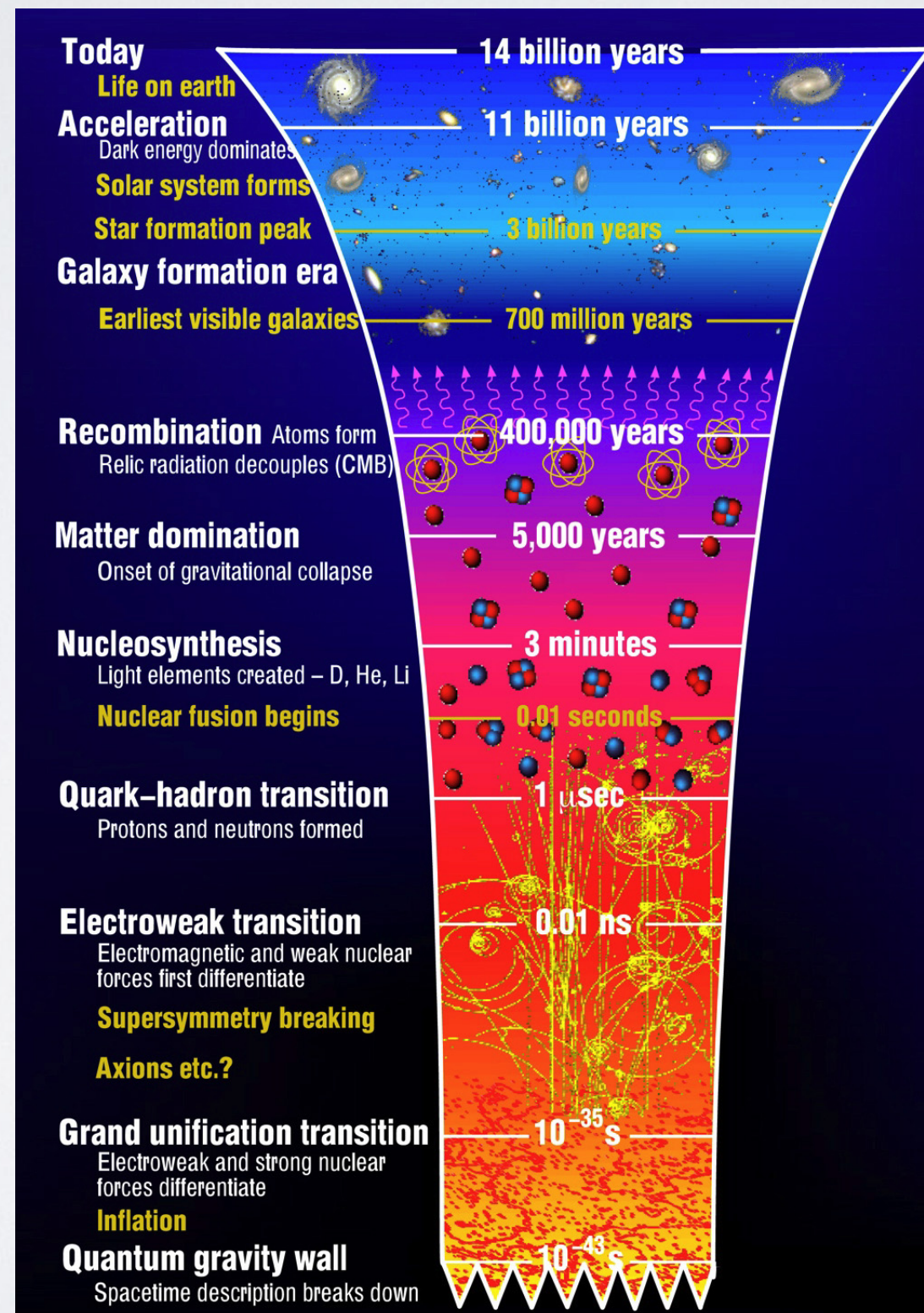
# METASTABILITY?



There's a lot of room between here and the Planck scale  
so maybe this problem is less pressing



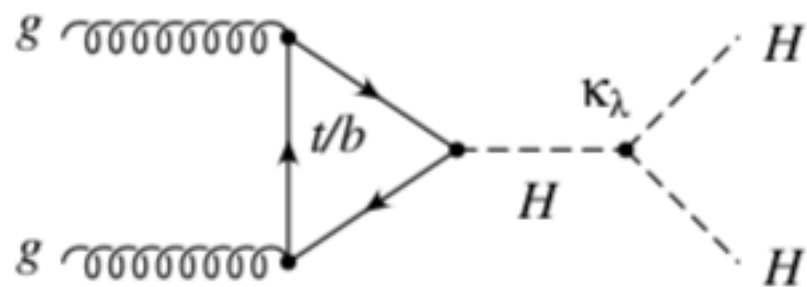
# THE ELECTROWEAK PHASE TRANSITION





# THE ELECTROWEAK PHASE TRANSITION

## Higgs boson pair production at colliders: status and perspectives



Double Higgs production at colliders allows for a direct probe of the couplings in the Higgs potential responsible for strengthening the electroweak phase transition.

Editors:

Biagio Di Micco, Maxime Gouzevitch,  
Javier Mazzitelli and Caterina Vernieri

The Higgs program and open questions in particle physics and cosmology

Beate Heinemann<sup>1,2a</sup>, and Yosef Nir<sup>3b1</sup>

In contrast to the question of electroweak phase transition, for the flavor measurements there is no lower bound on the size of new physics effects. Instead, the better accuracy, the higher the scale of new physics to which there is sensitivity.

# THE ELECTROWEAK PHASE TRANSITION

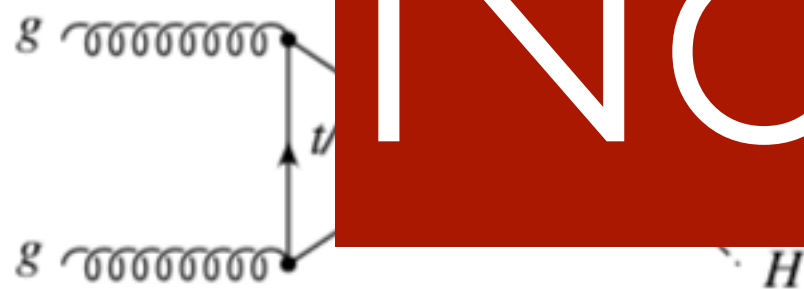
## Higgs boson pair production at colliders:

status and perspectives

The Higgs program and open questions in particle physics and cosmology

Beate Heinemann<sup>1,2a</sup>, and Yosef Nir<sup>3b1</sup>

Not True!



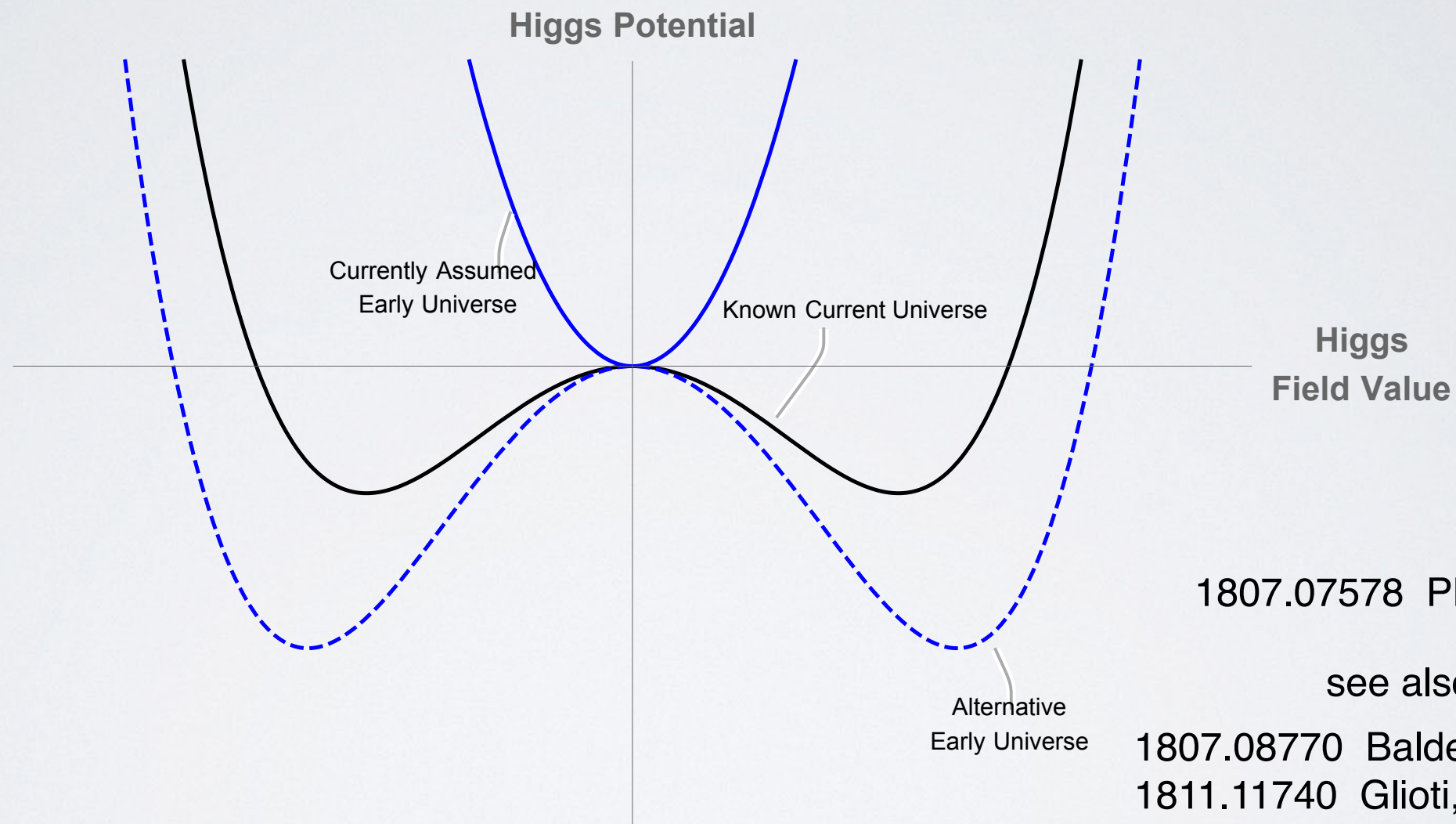
on of electroweak  
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Double Higgs production at colliders allows for a direct probe of the couplings  
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# SYMMETRY NON-RESTORATION OR DELAYED RESTORATION!



1807.07578 PM, H. Ramani

see also

1807.08770 Baldes, Servant

1811.11740 Glioti, Rattazzi, Vecchi

**Need a precise measurement of the triple Higgs coupling to differentiate!**



THE HIGGS CAN TELL US LOTS OF THINGS  
QUALITATIVELY ABOUT OUR UNIVERSE

IF WE MEASURE IT WELL ENOUGH!

Origin of EWSB

naturalness

stability of universe

cosmological history of the universe

All pretty impressive, but let's go next to where the  
Higgs really does the heavy lifting in the SM



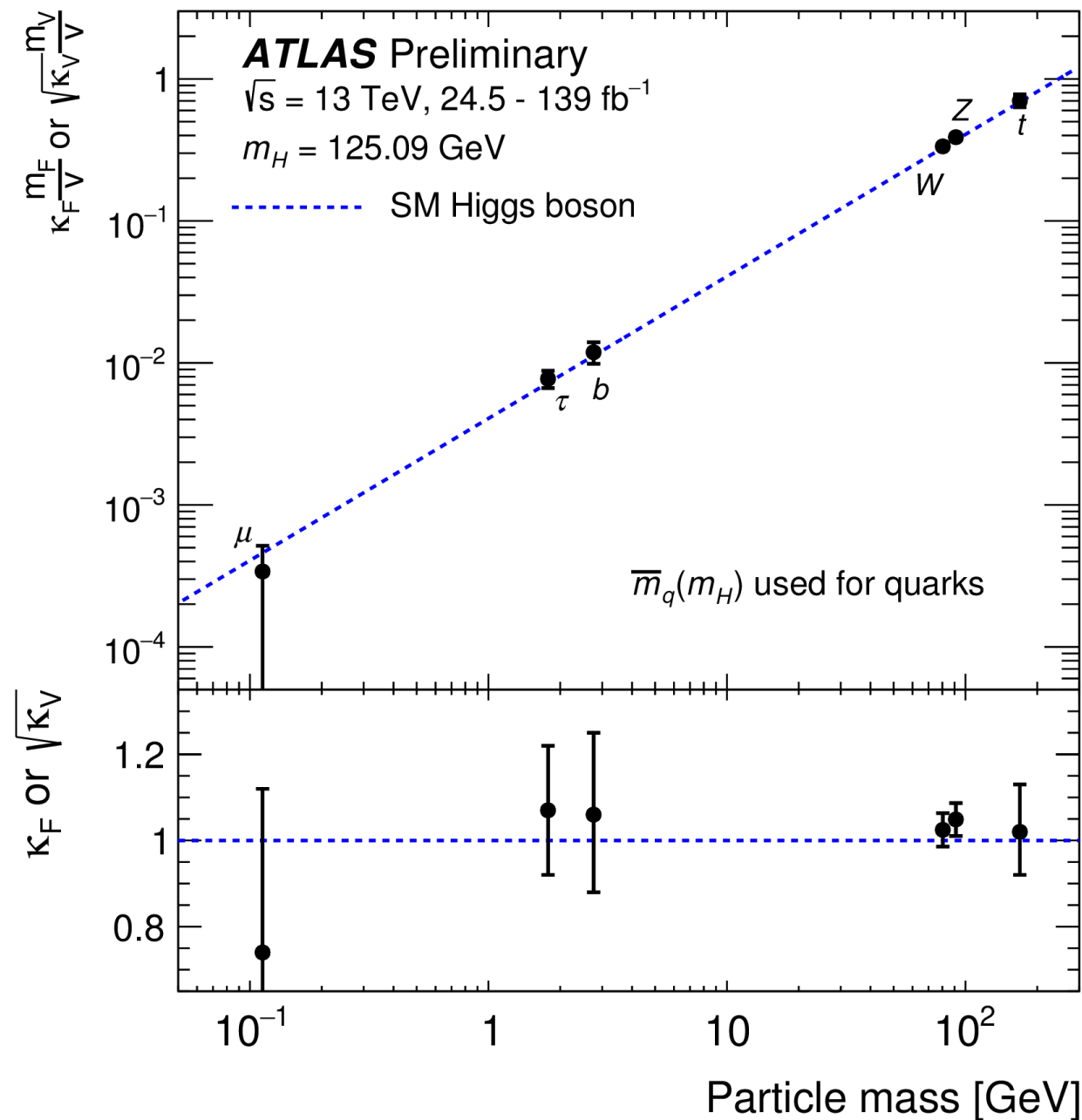


Flavor!!!!!!

Pause (k)



# FLAVOR FOR HIGGS SEEMS STRAIGHTFORWARD



Actually it's responsible for so much more of course...



JUST REMEMBER THE SM WOULD BE  
BEAUTIFUL WITHOUT THE HIGGS...

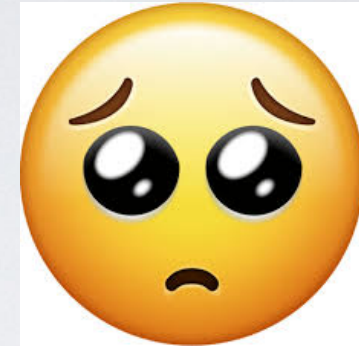
**Gauge theories :**

Gauge group + matter reps + gauge coupling

**Everything is fixed!**

# JUST REMEMBER THE SM WOULD BE BEAUTIFUL WITHOUT THE HIGGS...

Scalars can have Yukawa couplings



$$\mathcal{L} \supset Y_{ij}^d \bar{Q}_{Li} \phi D_{Rj} + Y_{ij}^u \bar{Q}_{Li} \tilde{\phi} U_{Rj}$$

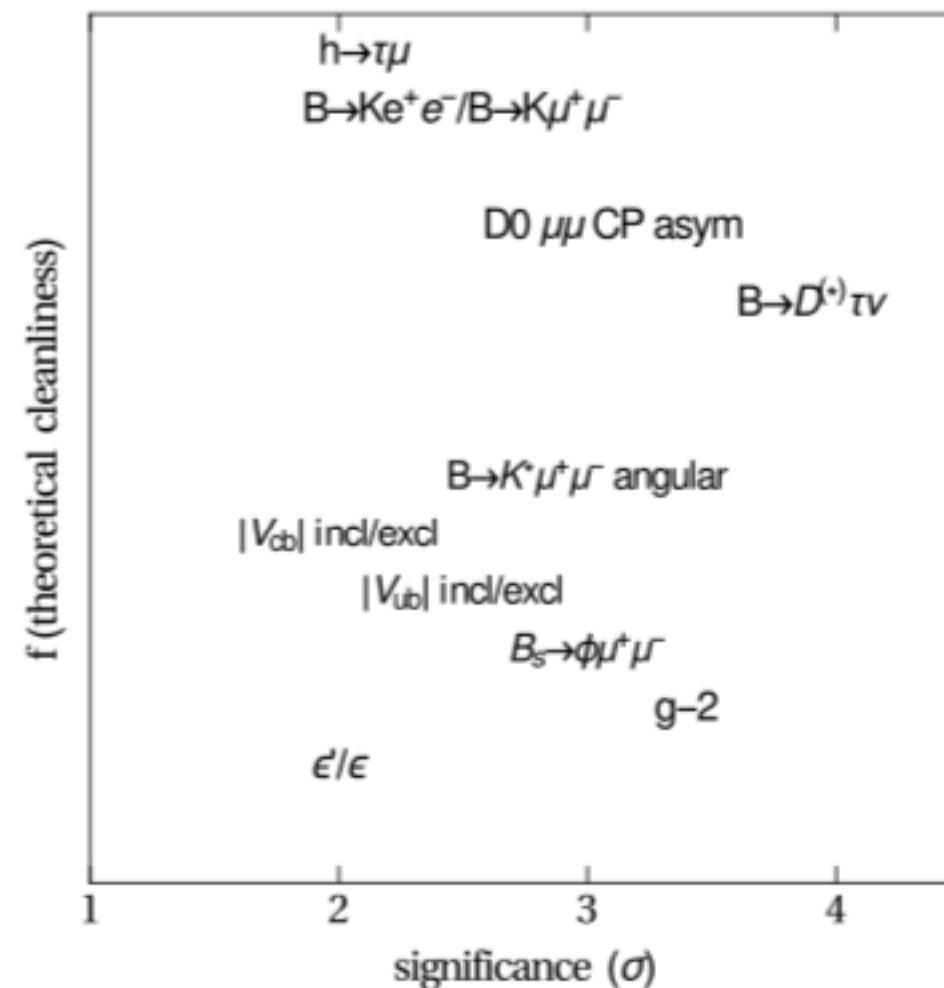
We shove these Yukawa's into masses and CKM matrix in mass basis

**Majority** of SM parameters come from Flavor i.e. couplings of Higgs

WHY DON'T WE TALK ABOUT  
FLAVOR MORE?

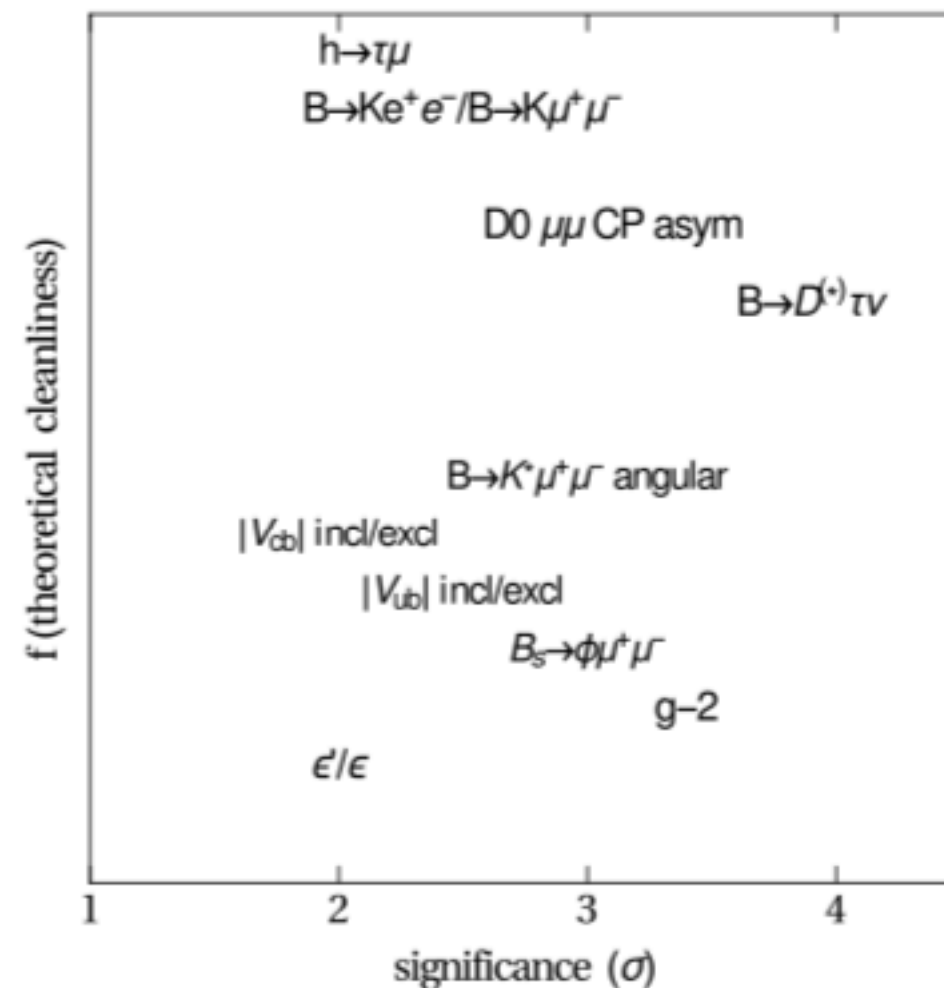


***We talk about it a lot when there are anomalies...***



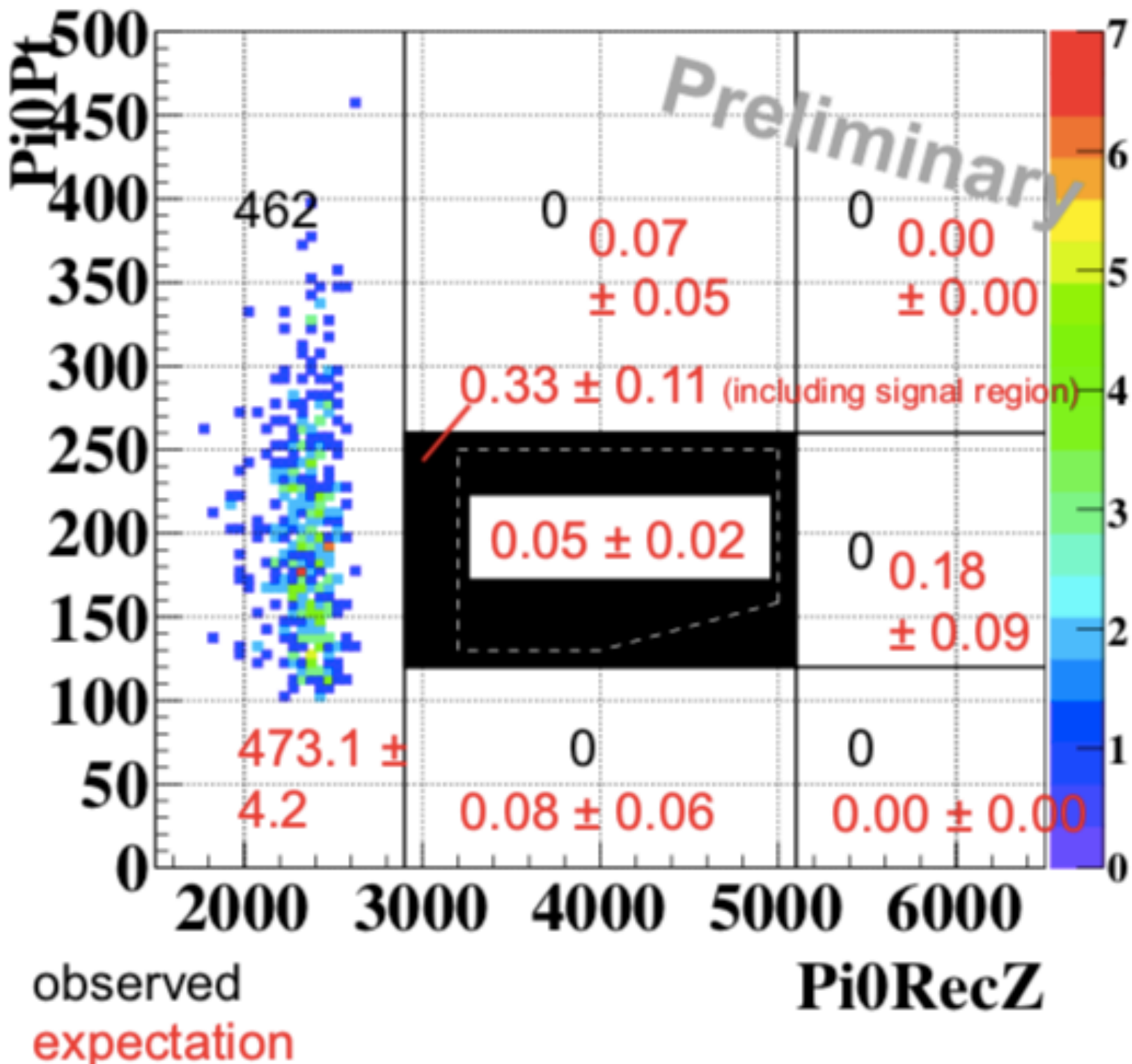
**Figure 3:** Some recent measurements in tension with the SM. The horizontal axis shows the nominal significance. The vertical axis shows (monotonically, in my opinion) an undefined function of an ill-defined variable: the theoretical cleanliness. That is, the level of plausibility that a really conservative estimate of the theory uncertainty of each observable may affect the significance of its deviation from the SM by  $1\sigma$ .

***We talk about it a lot when there are anomalies...***

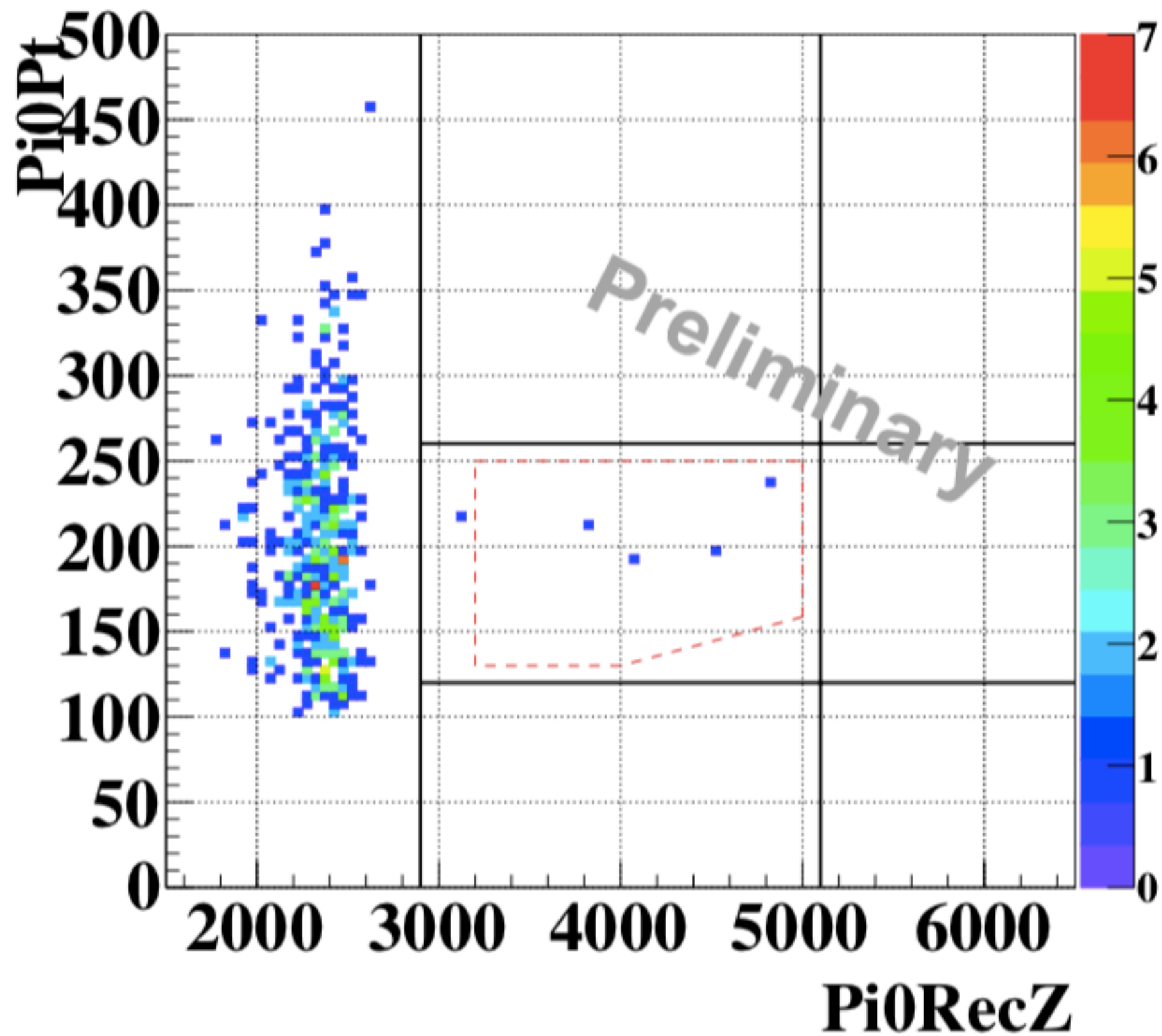


**Figure 3:** Some recent measurements in tension with the SM. The horizontal axis shows the nominal significance. The vertical axis shows (monotonically, in my opinion) an undefined function of an ill-defined variable: the theoretical cleanliness. That is, the level of plausibility that a really conservative estimate of the theory uncertainty of each observable may affect the significance of its deviation from the SM by  $1\sigma$ .

S.E.S :  $6.9 \times 10^{-10}$







unblinded in the end of Aug. 2019

MODELS OF FLAVOR RELATED TO COLLIDER  
PHYSICS GET A LOT LESS DISCUSSION...

IN FACT AN APT ANALOGY IS:



Flavor, the third rail of BSM theory!







BSM Flavor @ TeV Scale



IT'S NOT JUST A MATTER OF TASTE,  
THERE **IS** A PHYSICS REASON

# TYPICAL FLAVOR BOUNDS

Dimension 6 operators suppressing Dirac structure

$$\frac{C_s}{\Lambda^2} s \bar{d} s \bar{d}$$

$$\frac{C_c}{\Lambda^2} c \bar{u} c \bar{u}$$

$$\frac{C_b}{\Lambda^2} b \bar{d} b \bar{d}$$



# TYPICAL FLAVOR BOUNDS

Dimension 6 operators suppressing Dirac structure

$$\frac{C_s}{\Lambda^2} s \bar{d} s \bar{d} \quad \longrightarrow \quad \Lambda \gtrsim 10^4 \text{ TeV}$$

$$\frac{C_c}{\Lambda^2} c \bar{u} c \bar{u} \quad \longrightarrow \quad \Lambda \gtrsim 10^3 \text{ TeV}$$

$$\frac{C_b}{\Lambda^2} b \bar{d} b \bar{d} \quad \longrightarrow \quad \Lambda \gtrsim 10^4 \text{ TeV}$$

**Assuming  $\mathcal{O}(1)$  Wilson Coefficients**

# FLAVOR IS MEASURED REALLY WELL!

- So unless it's a very clean and large deviation it's probably hard to make an anomaly work
- For LHC BSM physics it's even crazier to think we'll see something flavor dependent, right?

# TYPICAL FLAVOR BOUNDS

Dimension 6 operators suppressing Dirac structure

$$\frac{C_s}{\Lambda^2} s \bar{d} s \bar{d} \quad \longrightarrow \quad \Lambda \gtrsim 10^4 \text{ TeV}$$

$$\frac{C_c}{\Lambda^2} c \bar{u} c \bar{u} \quad \longrightarrow \quad \Lambda \gtrsim 10^3 \text{ TeV}$$

$$\frac{C_b}{\Lambda^2} b \bar{d} b \bar{d} \quad \longrightarrow \quad \Lambda \gtrsim 10^4 \text{ TeV}$$

$\Lambda \gg$  **LHC COM Energy**



WHY WOULD ONE EVEN  
BUILD THE LHC?

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**Bounds come from assuming  $O(1)$  Wilson Coefficients**

# WHY WOULD ONE EVEN BUILD THE LHC?

**Bounds come from assuming  $\mathcal{O}(1)$  Wilson Coefficients**

Similar problems for Electroweak Precision before the LHC

$$\frac{C_T}{\Lambda^2} (H^\dagger D_\mu H)^2 \longrightarrow \Lambda \gtrsim 10 \text{ TeV}$$

***If you throw in a weak coupling and a loop factor...***

$$\Lambda \gtrsim 10 \text{ TeV} \longrightarrow \Lambda \gtrsim v$$

Weak scale is no problem!



# WHY WOULD ONE EVEN BUILD THE LHC?

**However...**

$\Lambda \gtrsim 10 \text{ TeV}$  EWPT is much weaker than  $\Lambda \gtrsim 10^4 \text{ TeV}$  FLAVOR

So you have to make ***much stronger assumptions***  
about flavor physics if you want to have any hint of it  
at the LHC

# TYPICAL BSM LHC FLAVOR BIASES



# TYPICAL BSM LHC FLAVOR BIASES

Flavor is super constrained so *new* physics is completely flavor blind

**or**

# TYPICAL BSM LHC FLAVOR BIASES

Flavor is super constrained so *new physics* is completely flavor blind

**or**

**“MFV”**

**M**inimal **F**lavor **V**iolation



# TYPICAL BSM LHC FLAVOR BIASES

Flavor is super constrained so *new* physics is completely flavor blind

or

“MFV”

## 23. Minimal flavor violation: An Effective field theory approach

G. D'Ambrosio (CERN & INFN, Naples), G.F. Giudice (CERN), G. Isidori (CERN & Frascati), A. Strumia (CERN & Pisa U. & INFN, Pisa). Jul 2002. 29 pp.

Published in **Nucl.Phys. B645 (2002) 155-187**

CERN-TH-2002-147, IFUP-TH-2002-17

DOI: [10.1016/S0550-3213\(02\)00836-2](https://doi.org/10.1016/S0550-3213(02)00836-2)

e-Print: [hep-ph/0207036](https://arxiv.org/abs/hep-ph/0207036) | [PDF](#)

[References](#) | [BibTeX](#) | [LaTeX\(US\)](#) | [LaTeX\(EU\)](#) | [Harvmac](#) | [EndNote](#)

[CERN Document Server](#); [ADS Abstract Service](#); [Link to Fulltext](#)

[Detailed record](#) - [Cited by 1442 records](#) 1000+

# MINIMAL FLAVOR VIOLATION

If we ignore SM Yukawa couplings  
SM has a very large global flavor symmetry

$$U(3)^5 = SU(3)_Q \times SU(3)_U \times SU(3)_D \times SU(3)_L \times SU(3)_E \times U(1)^5$$

$$U(1)^5 = U(1)_B \times U(1)_L \times U(1)_Y \times U(1)_{PQ} \times U(1)_E$$

Badly broken by e.g. SM quark Yukawas  
which are arbitrary 3x3 complex matrices

$$\mathcal{L} \supset Y_{ij}^d \overline{Q}_{L\phi} D_{Rj} + Y_{ij}^u \overline{Q}_{Li} \tilde{\phi} U_{Rj}$$

**Seibergology** spurions to the rescue



# MINIMAL FLAVOR VIOLATION

$$\mathcal{L} \supset Y_{ij}^d \overline{Q}_{L\phi} D_{Rj} + Y_{ij}^u \overline{Q}_{Li} \tilde{\phi} U_{Rj}$$

**Assume** Yukawas transform under the global symmetry and they are some background field which spontaneously breaks the symmetry

e.g. under  $SU(3)_Q \times SU(3)_U \times SU(3)_D$

$$Y^u \sim (3, \bar{3}, 1) \quad Y^d \sim (3, 1, \bar{3})$$

Why is this way of thinking useful?

# MINIMAL FLAVOR VIOLATION

**Assume** the only spurions which break the SM global flavor symmetry are the SM Yukawas!

(D' Ambrosio, Giudice, Isidori, Strumia)



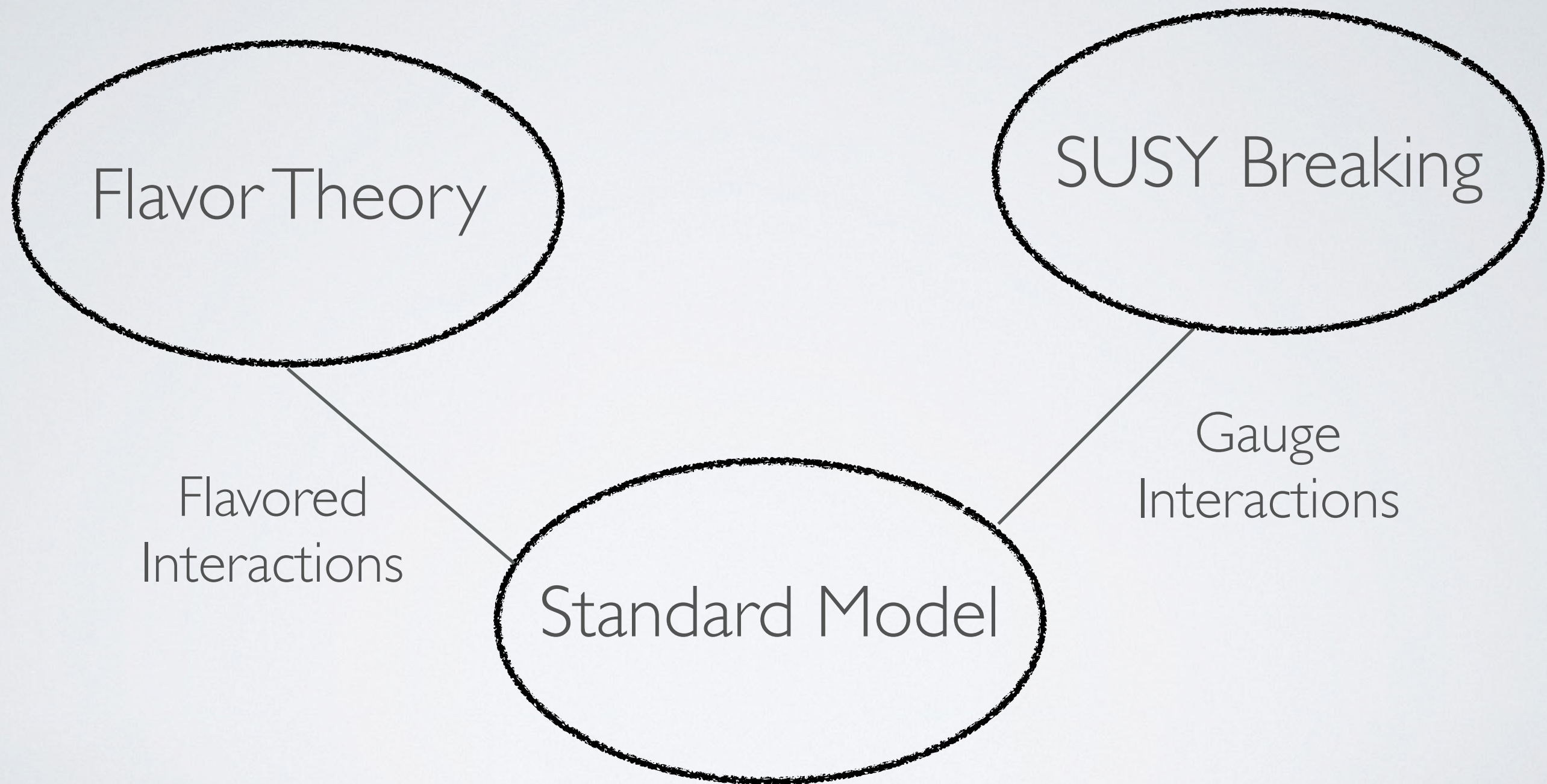
# MINIMAL FLAVOR VIOLATION

**Assume** the only spurions which break the SM global flavor symmetry are the SM Yukawas!

(D' Ambrosio, Giudice, Isidori, Strumia)

This is **NOT** a theory of FLAVOR

# MFV EXAMPLE: GAUGE MEDIATED SUSY BREAKING



In the end of the day SUSY only knows about flavor through SM



# MINIMAL FLAVOR VIOLATION

Assume the only spurions which break the SM global flavor symmetry are the SM Yukawas!

(D' Ambrosio, Giudice, Isidori, Strumia)

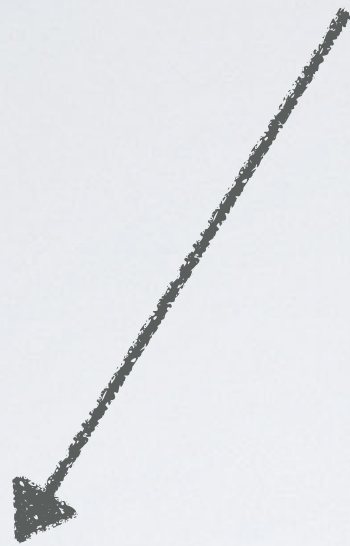
This is **NOT** a theory of FLAVOR

This is an **ansatz**, but one when combined with the symmetry group tells you exactly how any new physics operator transforms

If you **USE** MFV, you can typically bring the flavorful scale of NP down to the few TeV scale rather than  $10^4$  TeV

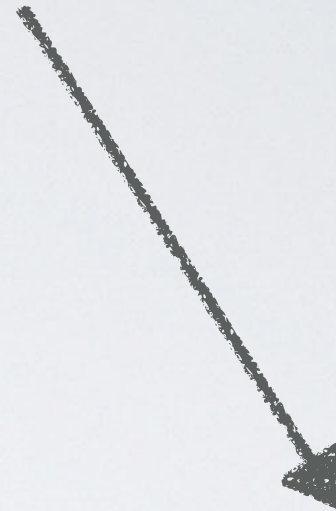
Inherit SM suppression of Yukawa and CKM

# MFV IN PRACTICE



EFT only

$$\frac{1}{\Lambda^2} (\bar{Q}_L \gamma^\mu Q_L)^2$$



Coupling to BSM physics!

$$(\bar{Q}_L \gamma^\mu Q_L) \mathcal{O}_\mu^{BSM}$$

Tells you what to put as  
coefficient in terms of Yukawa's of SM



# IT'S SOMEWHAT “NEEDED” IF YOU WANT TO DO SM EFT

Even at Dimension 6, the number of operators blows up if you have a general flavor structure

Case	CP even	CP odd	WHZ Pole parameters
General SMEFT ( $n_f = 1$ )	53 [10]	23 [10]	$\sim 23$
General SMEFT ( $n_f = 3$ )	1350 [10]	1149 [10]	$\sim 46$
$U(3)^5$ SMEFT	$\sim 52$	$\sim 17$	$\sim 24$
MFV SMEFT	$\sim 108$	-	$\sim 30$

Brivio, Jiang, MT <https://arxiv.org/abs/1709.06492>

Michael Trott

# EVEN THOUGH IT ALLOWS FOR “FLAVORFUL” COUPLINGS

Coupling to BSM physics!

$$(\bar{Q}_L \gamma^\mu Q_L) \mathcal{O}_\mu^{BSM}$$

**It's still kind of boring!**  
**Looks just like the SM: 3rd gen domination**



# EXTENSIONS OF MFV

NMFV: Next to Minimal Flavor Violation (Agashe, Papucci, Perez, Pirjol)

MFV + more flavor violation in 3rd gen

GMFV: General Minimal Flavor Violation (Kagan, Perez, Volansky, Zupan)

MFV + polynomials of Yukawa when needed because of 3rd gen

Amusingly, “flavored” BSM ansatz basically give the same guidance as naturalness

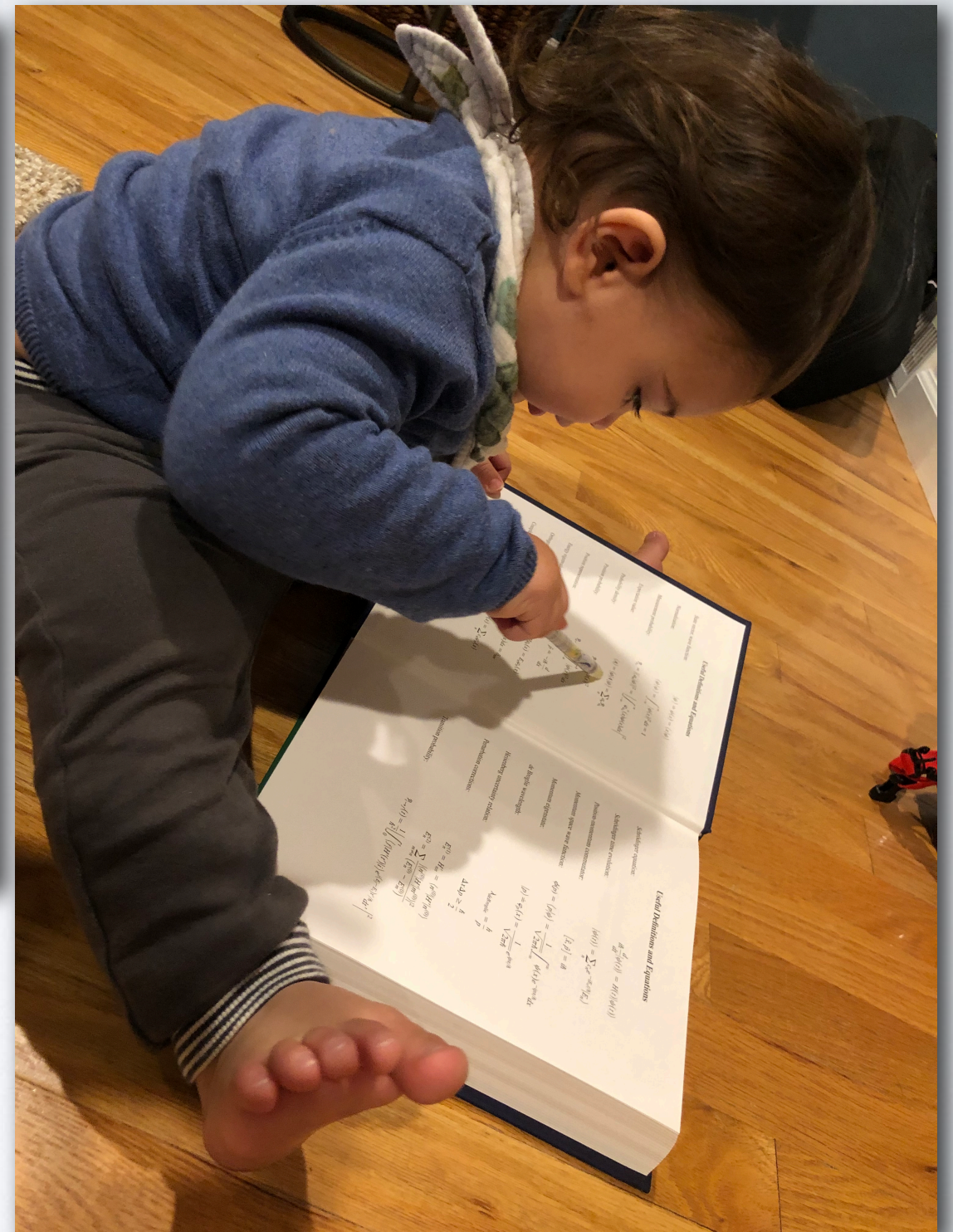
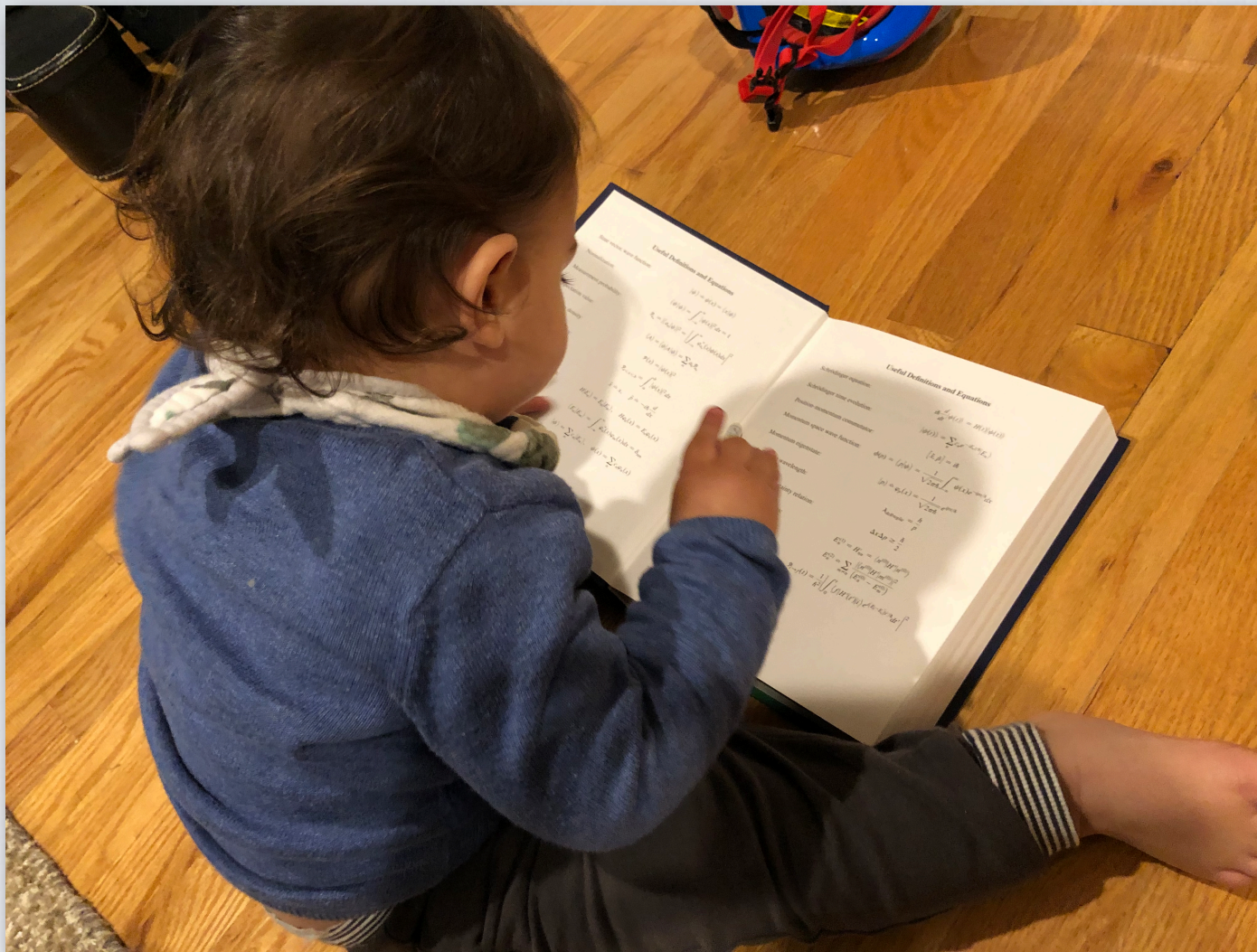
ARE WE BIASING OURSELVES INTO  
LOOKING ONLY IN CERTAIN PLACES?



HOW DO WE GET AROUND  
OUR BIASES?

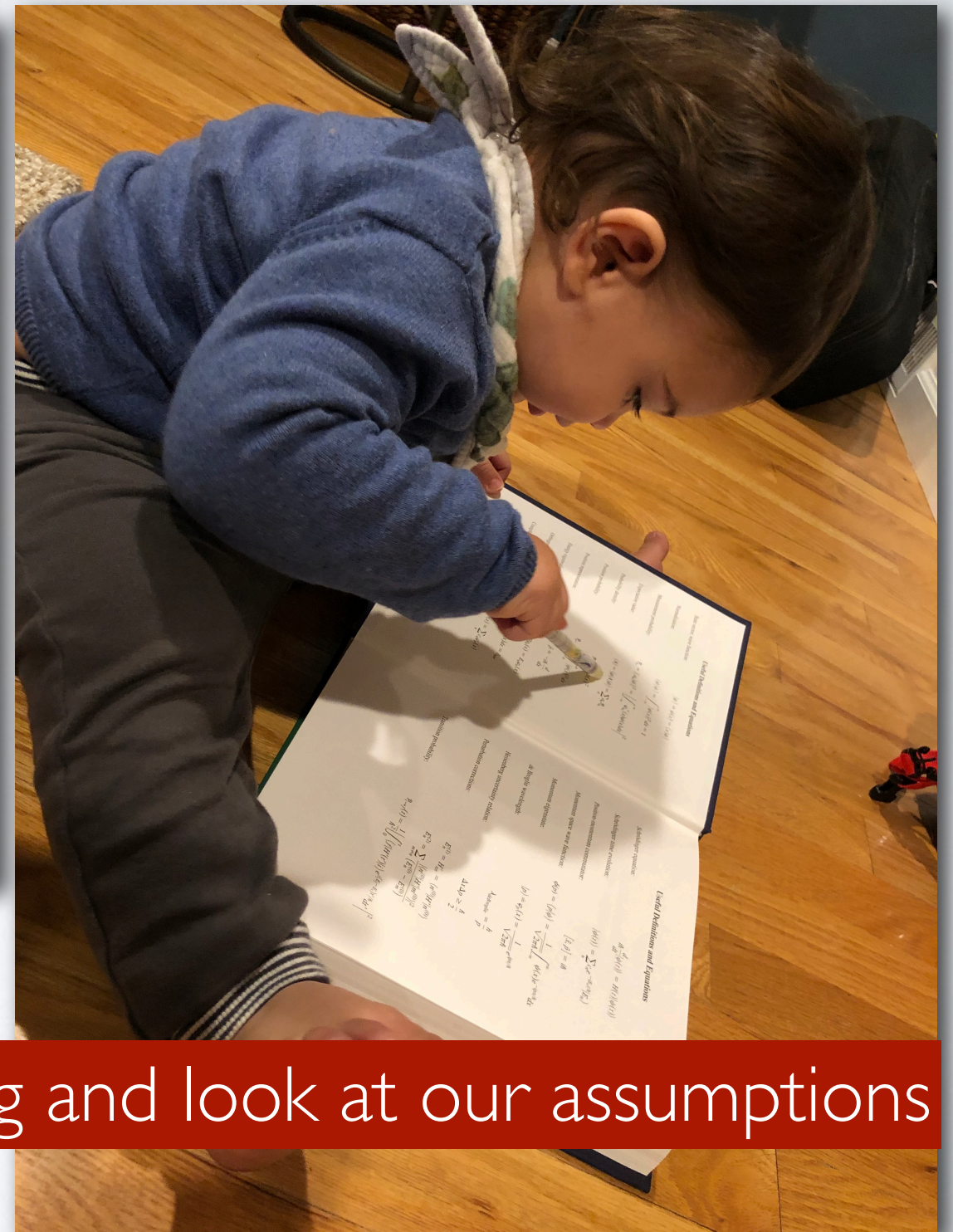
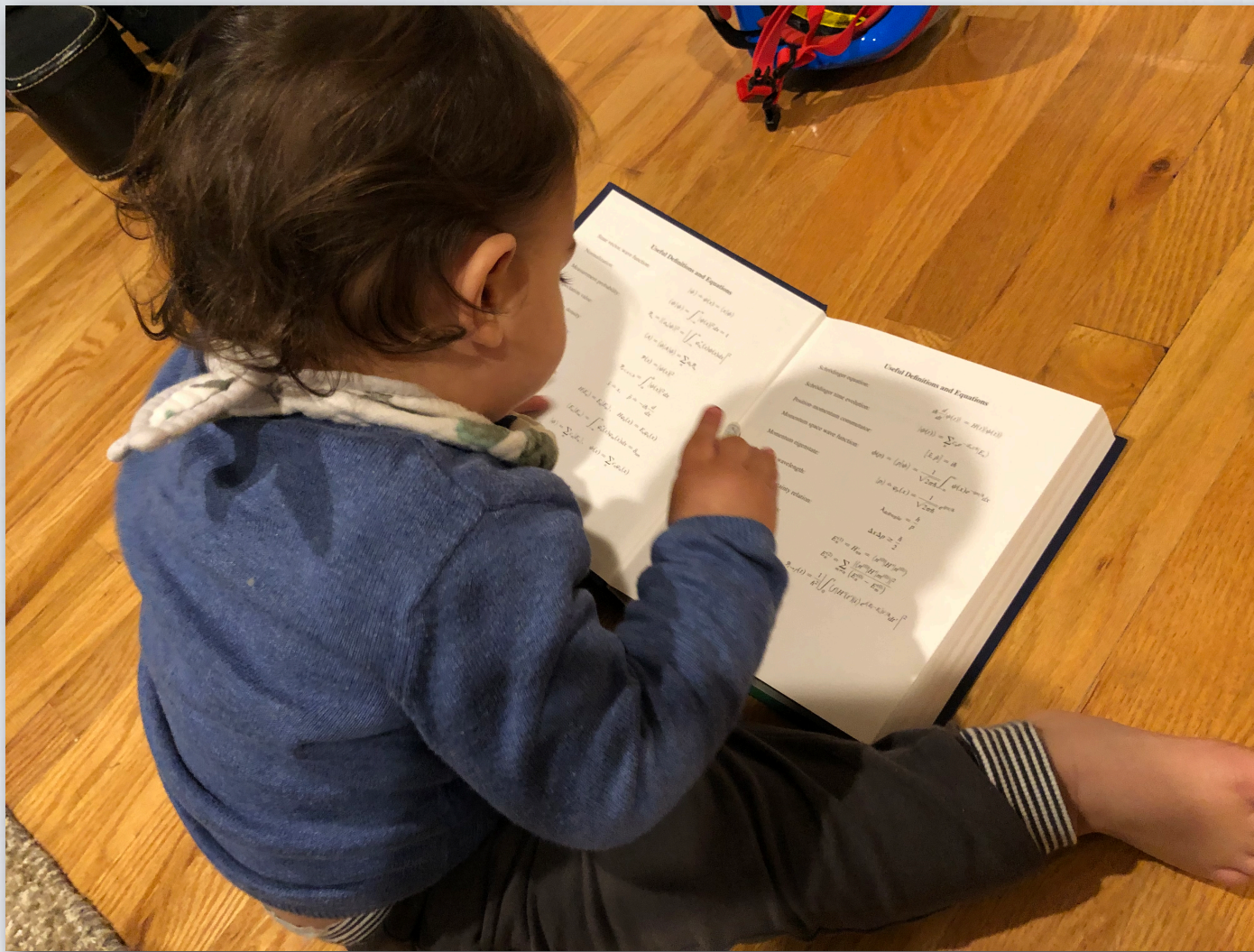


# HOW DO WE GET AROUND OUR BIASES?





# HOW DO WE GET AROUND OUR BIASES?



We have to go back to the beginning and look at our assumptions



# ALIGNED FLAVOR VIOLATION AND SPONTANEOUS FLAVOR VIOLATION



# WE'RE ALL FAMILIAR WITH THE MASS BASIS OF THE SM

All SM Yukawas are **Diagonal** and all non-trivial flavor lives  
in the CKM matrix for charged current interactions

We get there by using all those extra global (symmetry) rotations  
of our matter fields

# WE'RE ALL FAMILIAR WITH THE MASS BASIS OF THE SM

All SM Yukawas are **Diagonal** and all non-trivial flavor lives  
in the CKM matrix for charged current interactions

We get there by using all those extra global (symmetry) rotations  
of our matter fields

In the **gauge** basis we can think of a “**special basis**” where

$$Y_d = \lambda_d \quad \text{and} \quad Y_u = V^\dagger \lambda_u$$

**or**

$$Y_d = V \lambda_d \quad \text{and} \quad Y_u = \lambda_u$$



# IN ONE OF THESE SPECIAL BASIS CHOICES...

$$Y_d = \lambda_d \quad \text{and} \quad Y_u = V^\dagger \lambda_u$$

**or**

$$Y_d = V \lambda_d \quad \text{and} \quad Y_u = \lambda_u$$

For instance the first choice...

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$$Y_d = \lambda_d \quad \text{and} \quad Y_u = V^\dagger \lambda_u$$

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For instance the first choice...

I don't need to do anything more to the down sector to go to the mass basis, it's already diagonal... all non-trivial flavor violation is up-type



# IN ONE OF THESE SPECIAL BASIS CHOICES...

$$\begin{array}{ccc} Y_d = \lambda_d & \text{and} & Y_u = V^\dagger \lambda_u \\ & \mathbf{or} & \\ Y_d = V \lambda_d & \text{and} & Y_u = \lambda_u \end{array}$$

For instance the first choice...

I don't need to do anything more to the down sector to go to the mass basis, it's already diagonal... all non-trivial flavor violation is up-type

If I had some new down-type spurion that coupled BSM physics beyond  $Y_d$ , I'd just need it to be diagonal in this basis and no new tree level FCNCs!!

***Alignment*** (Nir, Seiberg)

# ALIGNMENT

Seiberg and Nir are great physicists, so they wouldn't just make a basis dependent statement...



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If you had a “complete” model of flavor, you could actually make statements about the form of the Yukawas

Implemented Frogatt-Nielsen models for SUSY Flavor Alignment

# ALIGNMENT

Seiberg and Nir are great physicists, so they wouldn't just make a basis dependent statement...

If you had a “complete” model of flavor, you could actually make statements about the form of the Yukawas

Implemented Frogatt-Nielsen models for SUSY Flavor Alignment

**Unfortunately you don't gain a lot in collider pheno**



# CAN WE GET:

- Successes of Flavor Alignment in a Basis Independent way
  - Fancier way of saying: if we introduce a new spurion such as  $\kappa^u, \kappa^d$  transforming like  $y^u, y^d$  is it auto-aligned?
- Can we write down a UV complete model that has parametrically **new** collider phenomenology
- Can we do it in a way that's modular, i.e. can be applied to many BSM theories?

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- Can we do it in a way that's modular, i.e. can be applied to many BSM theories?



# SPONTANEOUS FLAVOR VIOLATION (SFV)

## ***Ansatz to see alignment:***

- No renormalizable breaking of  $U(1)_f^3 \times CP$  other than WF renormalization of RH u **or** d quarks
- No flavor breaking spurions or fields other than SM ones and WF transforming under  $U(3)_{\bar{u}}$  **or**  $U(3)_{\bar{d}}$

# SFV

$$\begin{aligned}\mathcal{L} \supset & iZ_{ij}^u \bar{u}_i^\dagger \bar{\sigma}^\mu D_\mu \bar{u}_j + i\bar{d}_i^\dagger \bar{\sigma}^\mu D_\mu \bar{d}_i + i\bar{Q}_i^\dagger \bar{\sigma}^\mu D_\mu \bar{Q}_i \\ & - [\eta_{ij}^u Q_i H \bar{u}_j - \eta_{ij}^d Q_i H^c \bar{d}_j + \text{h.c.}] + \mathcal{L}_{\text{BSM}}\end{aligned}$$

WLOG  $\eta$  are real diagonal  
go to canonical kinetic terms



# SFV

$$\mathcal{L} \supset i Z_{ij}^u \bar{u}_i^\dagger \bar{\sigma}^\mu D_\mu \bar{u}_j + i \bar{d}_i^\dagger \bar{\sigma}^\mu D_\mu \bar{d}_i + i \bar{Q}_i^\dagger \bar{\sigma}^\mu D_\mu \bar{Q}_i \\ - [\eta_{ij}^u Q_i H \bar{u}_j - \eta_{ij}^d Q_i H^c \bar{d}_j + \text{h.c.}] + \mathcal{L}_{\text{BSM}}$$

WLOG  $\eta$  are real diagonal  
go to canonical kinetic terms

$$\bar{u}'_i = (\sqrt{Z^u})_{ij} \bar{u}_j$$

$$\mathcal{L} \supset i \bar{u}'_i \bar{\sigma}^\mu D_\mu \bar{u}'_i + i \bar{d}'_i \bar{\sigma}^\mu D_\mu \bar{d}'_i + i \bar{Q}'_i \bar{\sigma}^\mu D_\mu \bar{Q}'_i \\ - [y_{ij}^u Q_i H \bar{u}'_j - y_{ij}^{d\dagger} Q_i H^c \bar{d}'_j + \text{h.c.}] + \mathcal{L}_{\text{BSM}}$$

$$y^u = \eta^u (\sqrt{Z^u})^{-1} = V^T Y^u \\ y^{d\dagger} = \eta^d = Y^d .$$

# SFV

If one introduced a new spurion that transformed like the down-type Yukawa, it's **automatically** diagonal if flavor comes from VWF renormalization of the RH up type quarks!

This is **NOT** a basis dependent alignment

$$\kappa^{d\dagger} = K^d \equiv \text{diag}(\kappa_d, \kappa_s, \kappa_b)$$

These kappa don't have to have **anything** to do with SM Yukawas

For example new physics could couple (1,0,0) or (0,1,0) ...

$Z'$ , 2HDM, ...



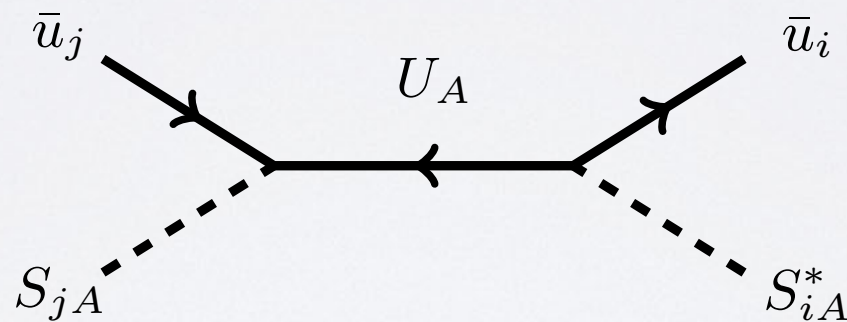
IS THERE A THEORY THAT  
CAN GENERATE SFV ANSATZ?

**YES**

# SFV UV COMPLETION

$$\mathcal{L} \supset M_{AB} U_A \bar{U}_B + \xi S_{iA} \bar{u}_i U_A \\ - [\eta_{ij}^u Q_i H \bar{u}_j - \eta_{ij}^d Q_i H^c \bar{d}_j + \text{h.c.}] + \mathcal{L}_{\text{BSM}}$$

Integrate out vector like quarks U



$$Z_{ij}^u = \delta_{ij} + \frac{\xi^* \xi}{M_A^* M_A} S_{iA}^* S_{jA}$$



# HAVEN'T SPECIFIED BSM...

2HDM Example:

$$D_\mu H_a^\dagger D^\mu H_a - V(H_1, H_2) - \left[ \lambda_{aij}^u Q_i H_a \bar{u}_j - \lambda_{aij}^{d\dagger} Q_i H_a^c \bar{d}_j - \lambda_{aij}^{\ell\dagger} L_i H_a^c \bar{\ell}_j + \text{h.c.} \right]$$

Nothing says the 2nd Higgs has to have the same Yukawas

**Work in the “Higgs” basis where only  
one Higgs gets a VEV**

# SFV 2HDM UP-TYPE

$$\lambda_1^u = V^T Y^u \quad \lambda_1^d = Y^d$$

$$\lambda_2^u = \xi V^T Y^u \quad \lambda_2^d = K^d .$$

$$K^d \equiv \text{diag}(\kappa_d, \kappa_s, \kappa_b)$$

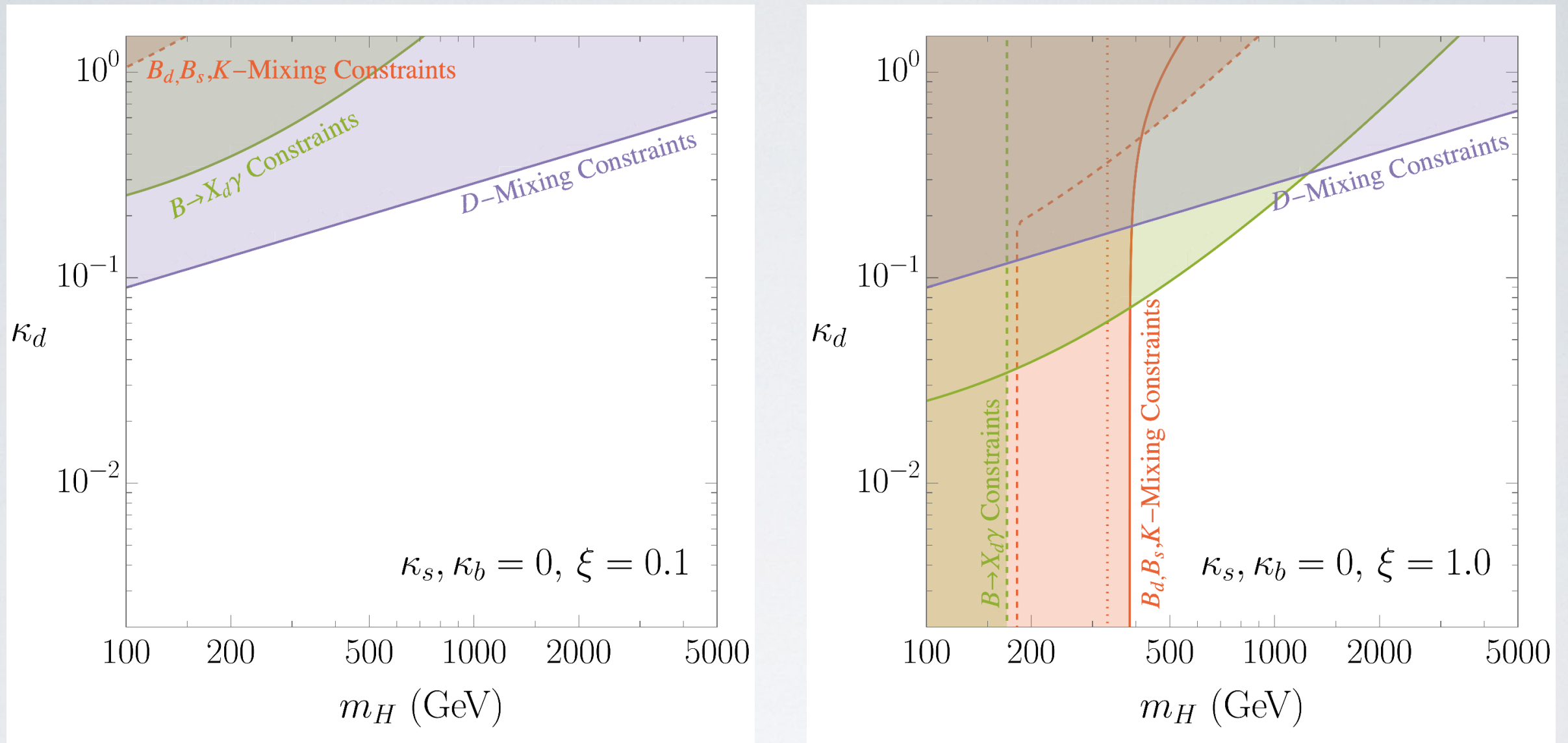
Can just as easily do this for down-type



# THERE WILL BE FLAVOR CONSTRAINTS

- We just got rid of tree-level FCNCs, but we still get to inherit the protections of the SM at loop-level

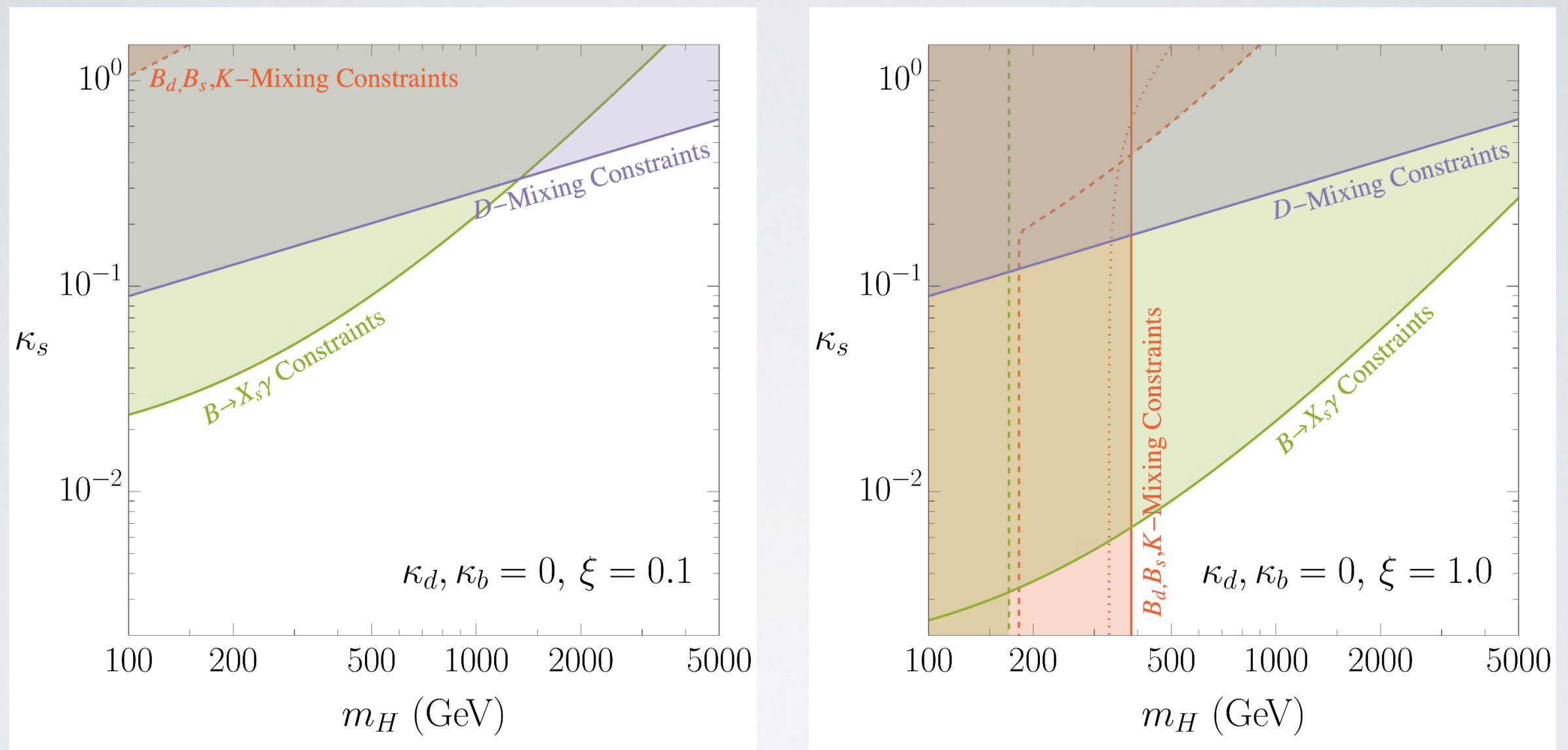
# FLAVOR CONSTRAINTS



**Figure 2:** Constraints on the up-type SFV 2HDM from one-loop FCNC measurements in the  $m_H$  vs.  $\kappa_d$  plane, assuming  $\kappa_s = \kappa_b = 0$ . We show results both for  $\xi = 0.1$  and  $1.0$ . Constraints from  $b \rightarrow s\gamma$  and  $b \rightarrow d\gamma$  transitions are shown in green, with the constraint on  $C_{7'}^{bd}$  ( $C_7^{bs}$ ) indicated by the solid (dashed) line, respectively. Constraints from  $B_d$ ,  $B_s$  and  $K$  mixing are shown as solid, dotted and dashed red lines respectively. The constraint from requiring the absence of fine-tuning in  $D - \bar{D}$  mixing is shown in purple.

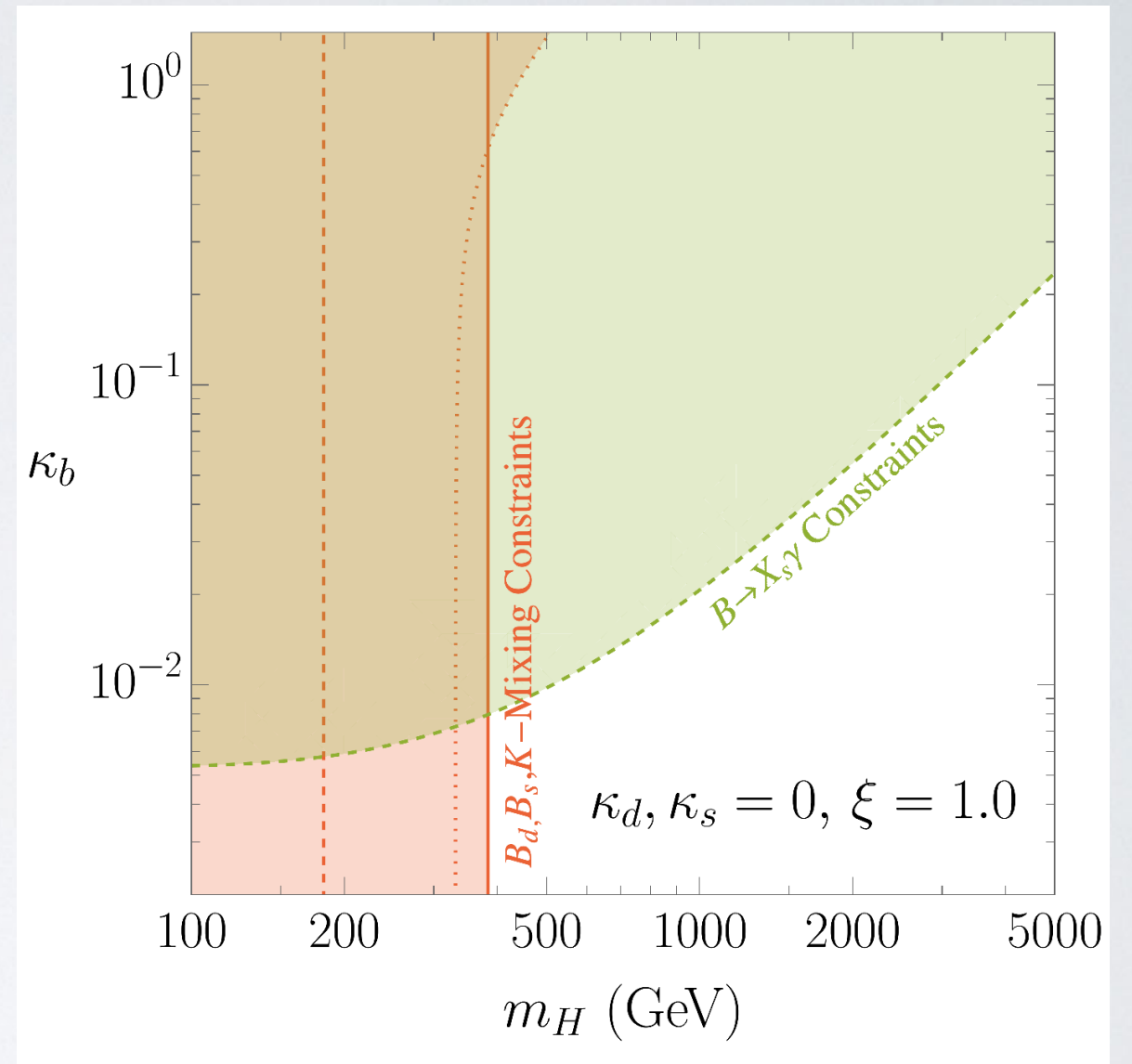
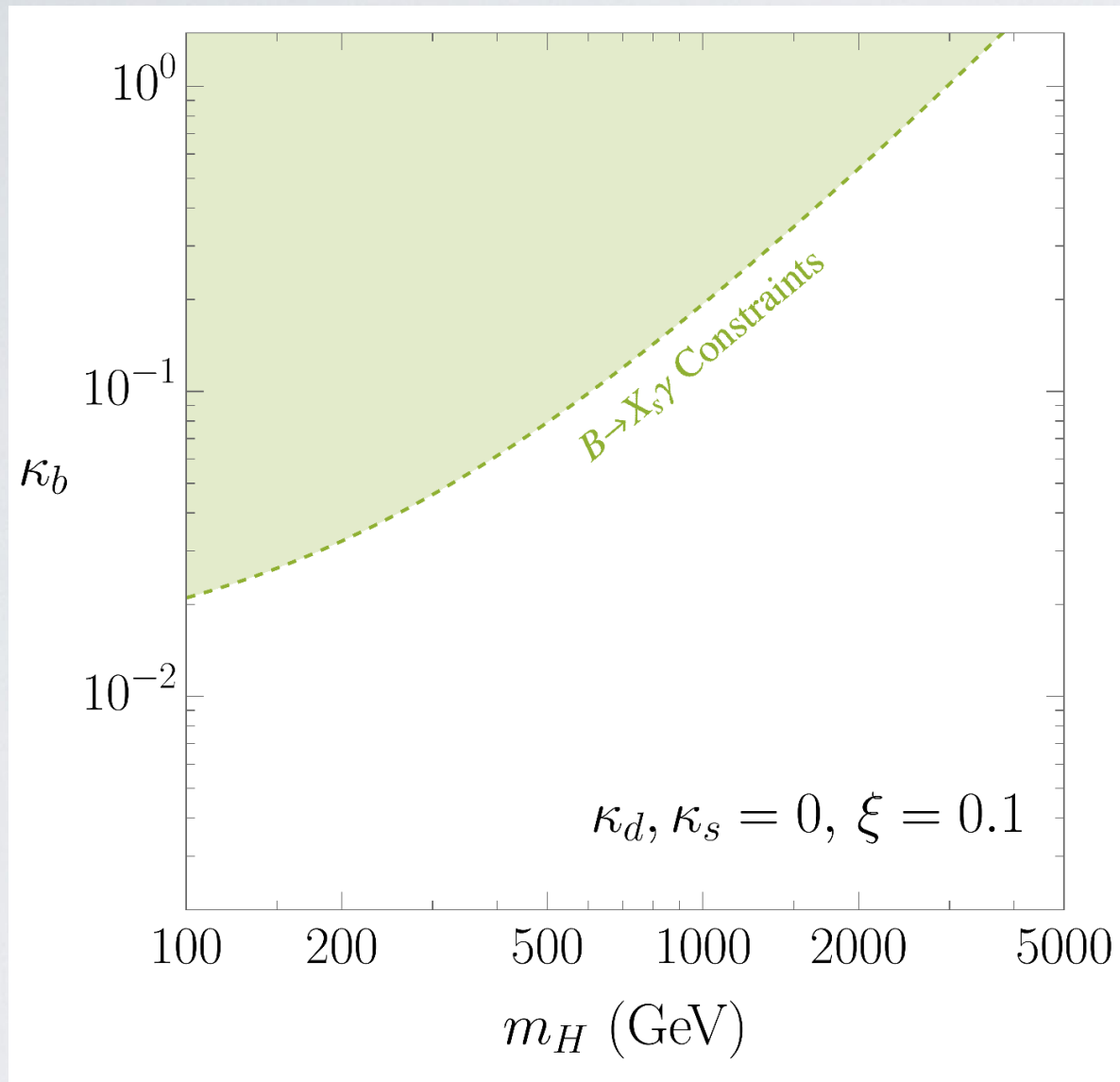


# FLAVOR CONSTRAINTS



**Figure 3:** The same as Fig. 2, but for  $\kappa_s$ , with  $\kappa_d = \kappa_b = 0$ .

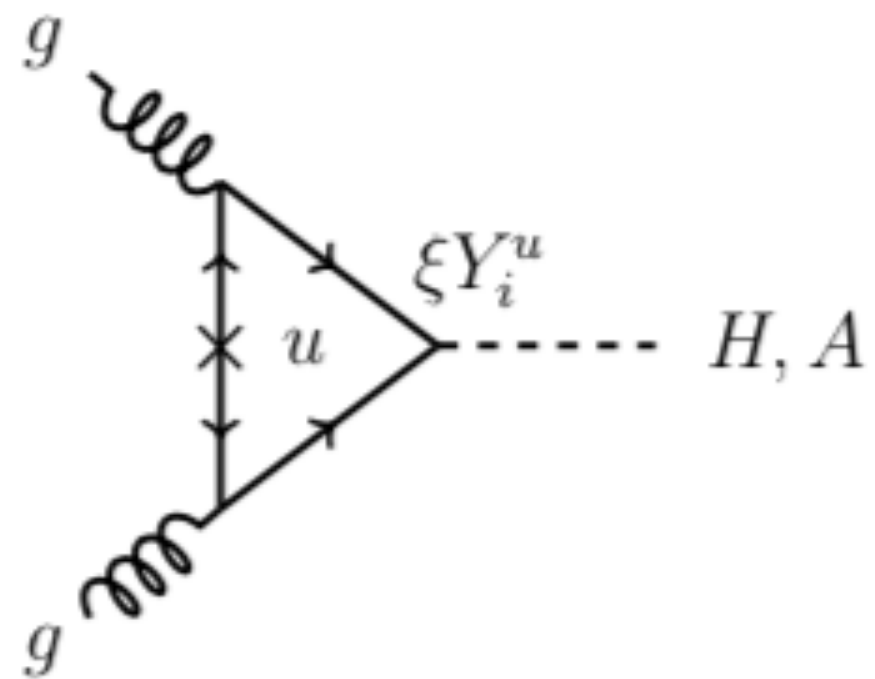
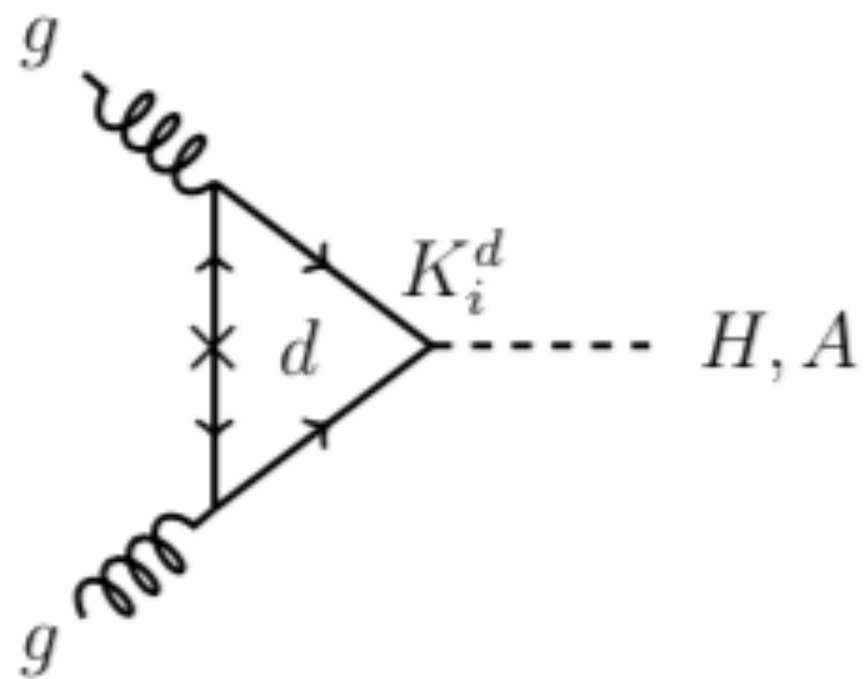
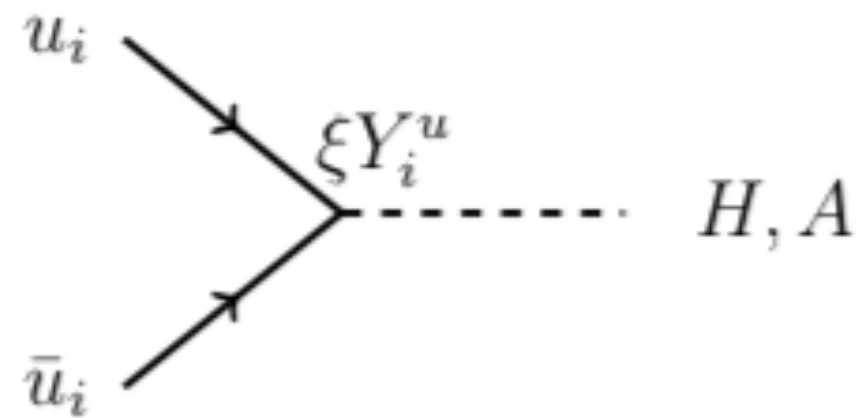
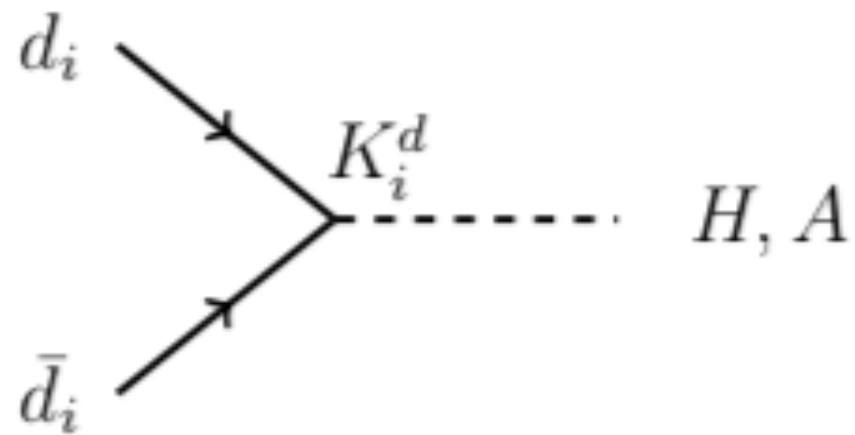
# FLAVOR CONSTRAINTS



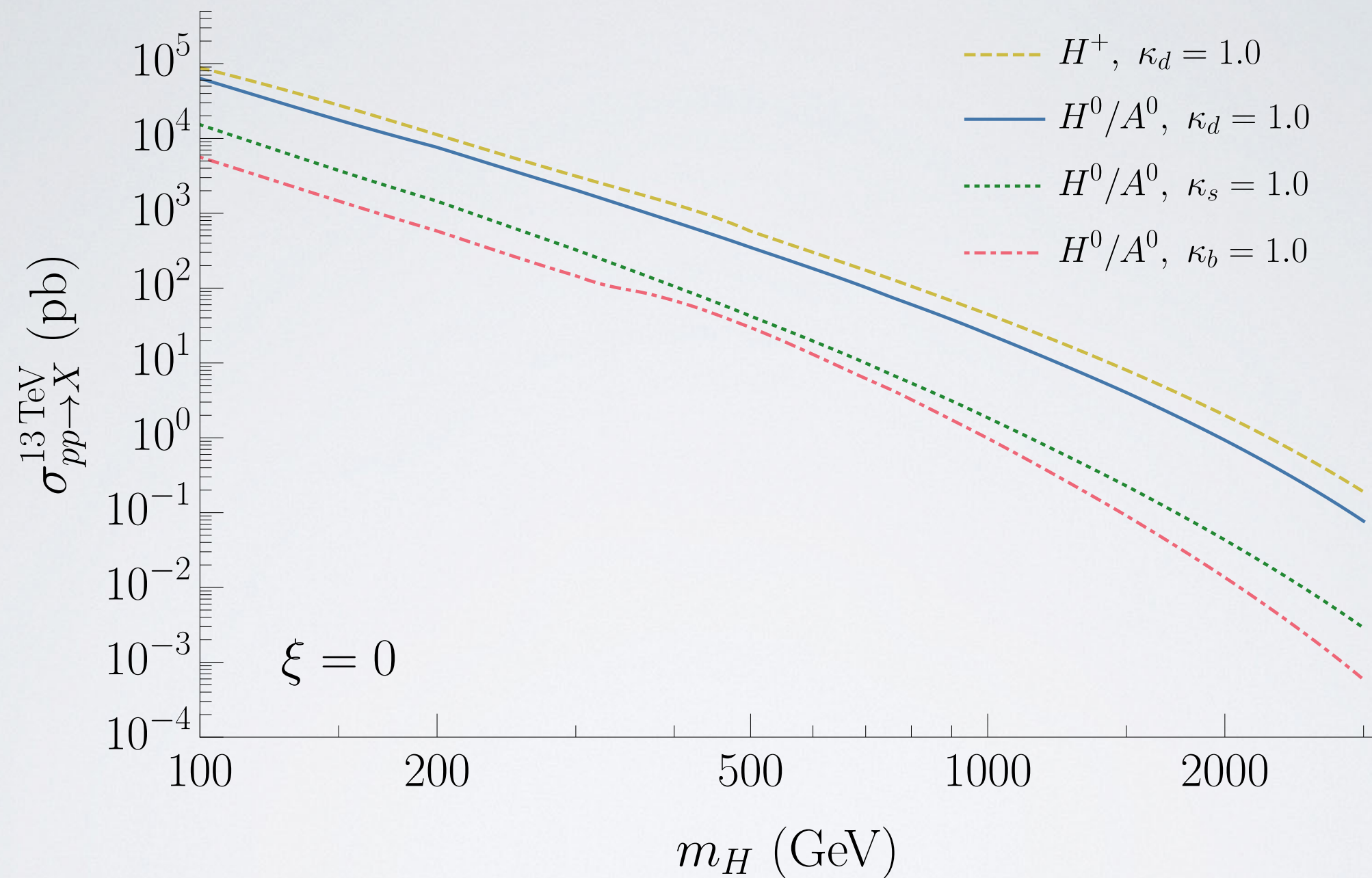
**Figure 4:** The same as Fig. 2 but for  $\kappa_b$ , with  $\kappa_d = \kappa_s = 0$ .



# COLLIDER PHENO



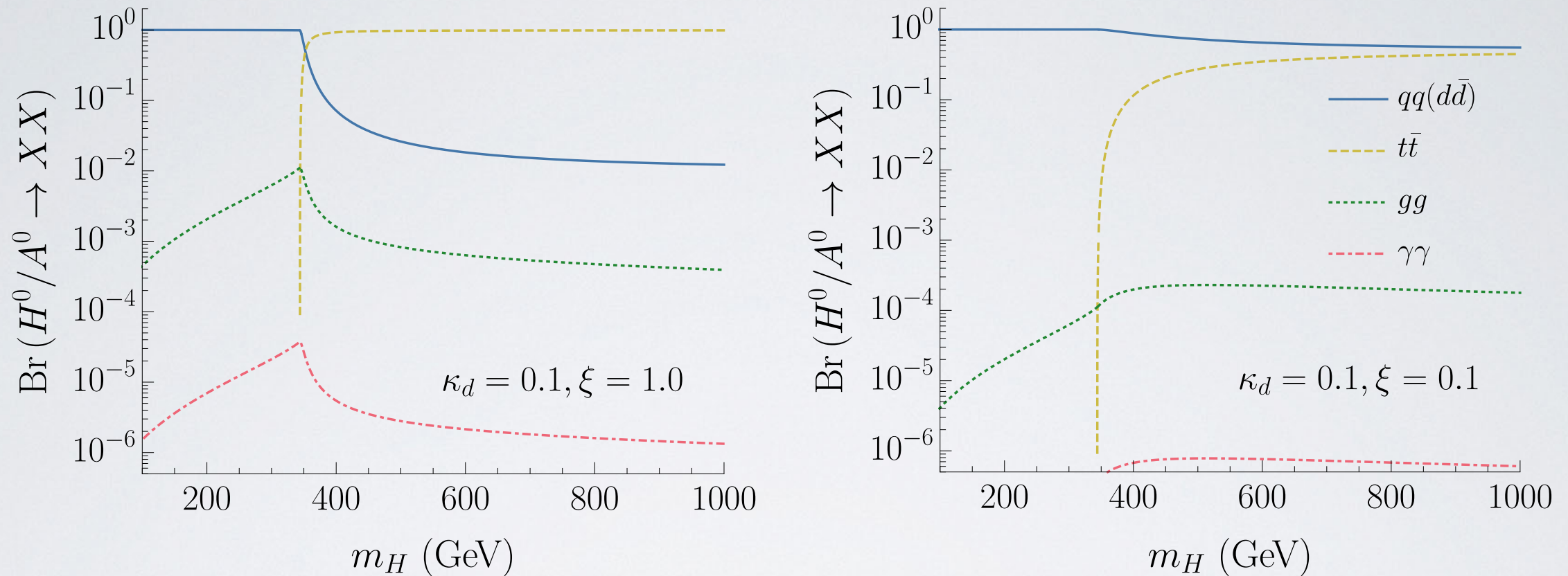
# COLLIDER PHENO



**Remember GGF for SM higgs is  $\sim 49 \text{ pb}$**

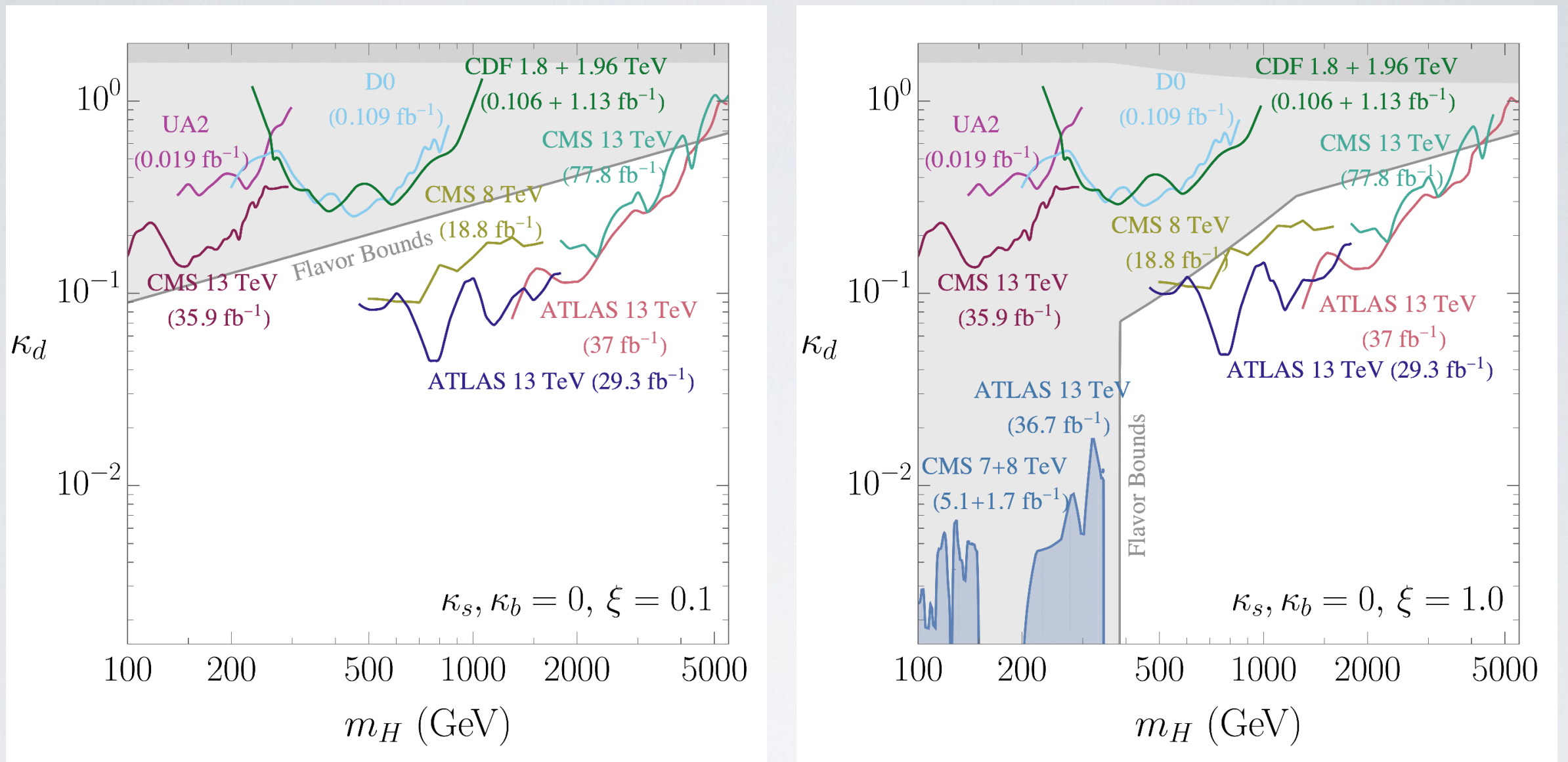


# COLLIDER PHENO



**Figure 8:** Plot of the branching fraction of  $H$  to  $d\bar{d}$  (solid blue),  $t\bar{t}$  (dashed yellow),  $gg$  (dotted green) and  $\gamma\gamma$  (dot-dashed red), as a function of  $m_H$  with  $\kappa_d = 0.1$  for both  $\xi = 1.0$  (left) and  $\xi = 0.1$  (right). In both plots we've taken  $\kappa_s = \kappa_b = 0$ . The behavior when replacing  $\kappa_d$  with either  $\kappa_s$  or  $\kappa_b$  is similar, with the decays to  $d\bar{d}$  replaced by  $s\bar{s}$  or  $b\bar{b}$  correspondingly.

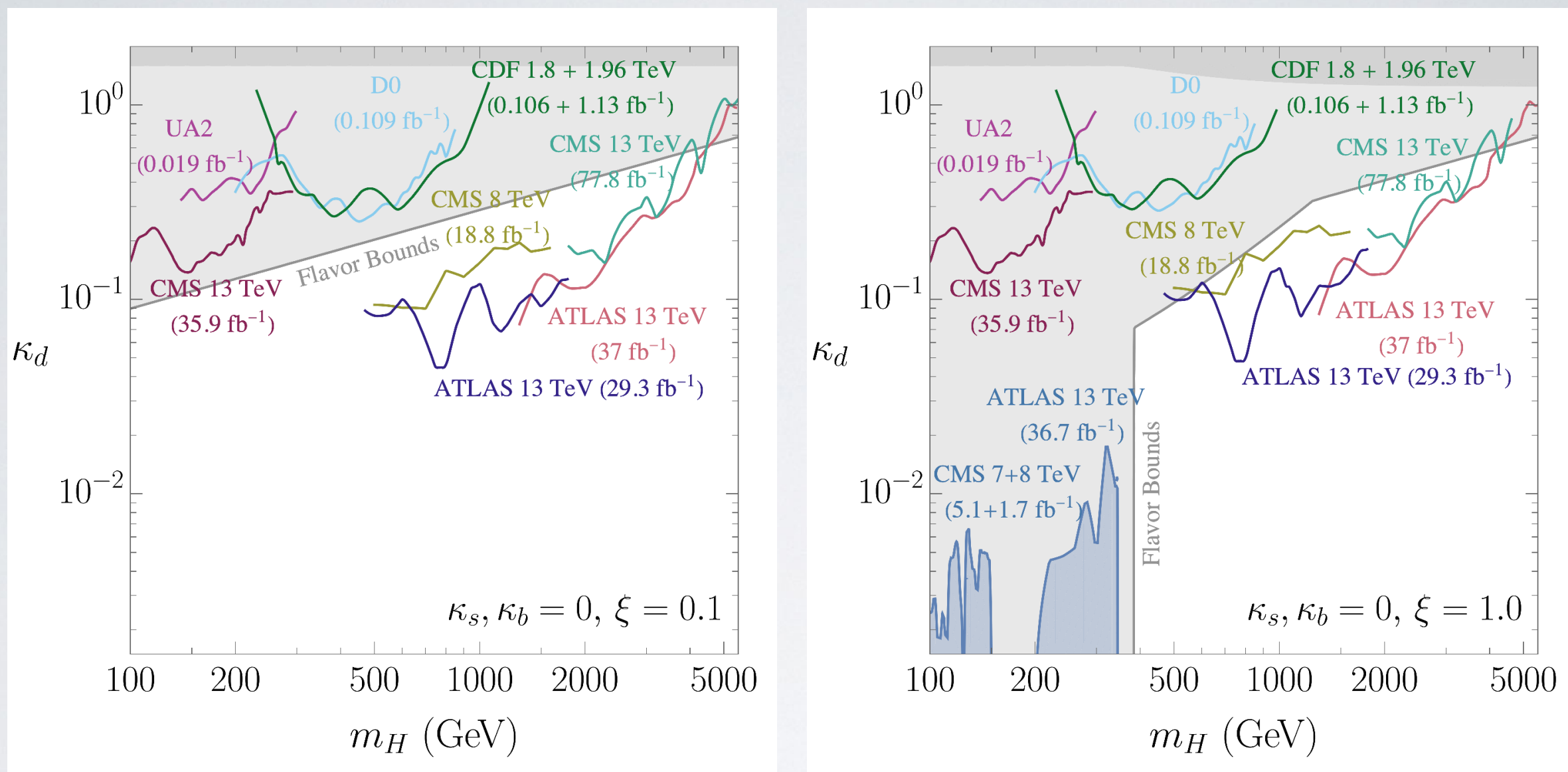
# COLLIDER PHENO



**Figure 10:** Constraints on the up-type SFV 2HDM from dijet and diphoton searches in the  $m_H$  vs.  $\kappa_d$  plane, assuming  $\kappa_s = \kappa_b = 0$ . We show results both for  $\xi = 0.1$  (left) and  $\xi = 1.0$  (right). Constraints from flavor observables, detailed in Fig. 2 are shown as the gray shaded region. The dark gray region above  $\kappa_d \sim 1.0$  indicates values of  $\kappa_d$  for which  $\Gamma/m_H \gtrsim 0.15$  for the heavy neutral Higgs, at which point dijet searches become less reliable and the results should be interpreted with care.



# COLLIDER PHENO



**Figure 10:** Keep in mind analogous SM coupling is  $10^{-5}$  lies in the  $m_H$  vs.  $\kappa_d$  plane, assuming  $\kappa_s = \kappa_b = 0$ . We show results both for  $\xi = 0.1$  (left) and  $\xi = 1.0$  (right). Constraints from flavor observables, detailed in Fig. 2 are shown as the gray shaded region. The dark gray region above  $\kappa_d \sim 1.0$  indicates values of  $\kappa_d$  for which  $\Gamma/m_H \gtrsim 0.15$  for the heavy neutral Higgs, at which point dijet searches become less reliable and the results should be interpreted with care.

# THIS IS AN EXAMPLE OF A HADRONIC - **QUARK DRIVEN** RESONANCE

- Most searches for hadronic resonances either have gluons, or are flavor universal or have leptons i.e.

$$Z' : U(1)_{B-L} \text{ or } U(1)_B$$

e.g. Dobrescu and Yu 1306.2629

- Signal generator!



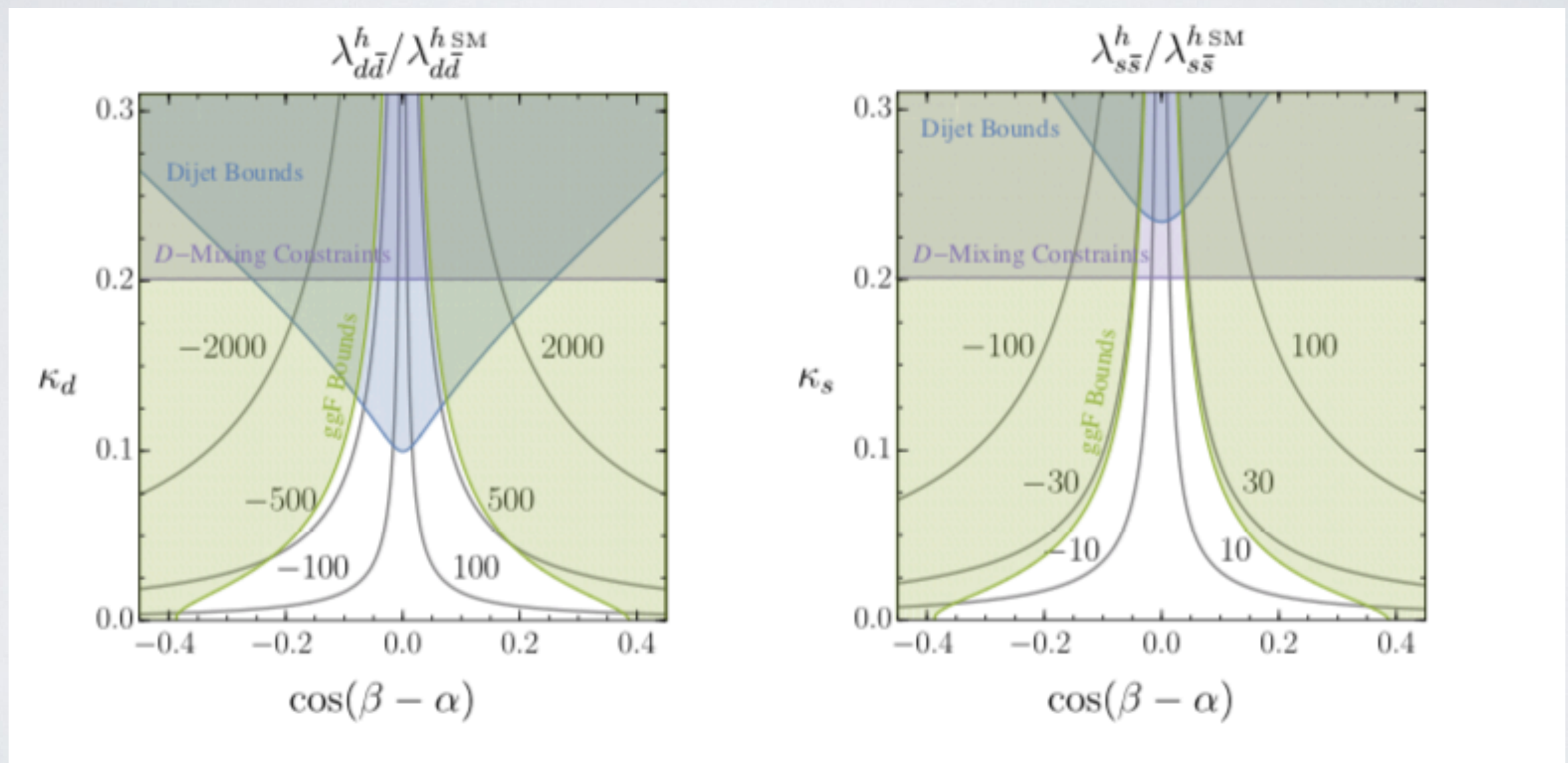
IS THIS JUST FOR NEW BSM  
PARTICLES?

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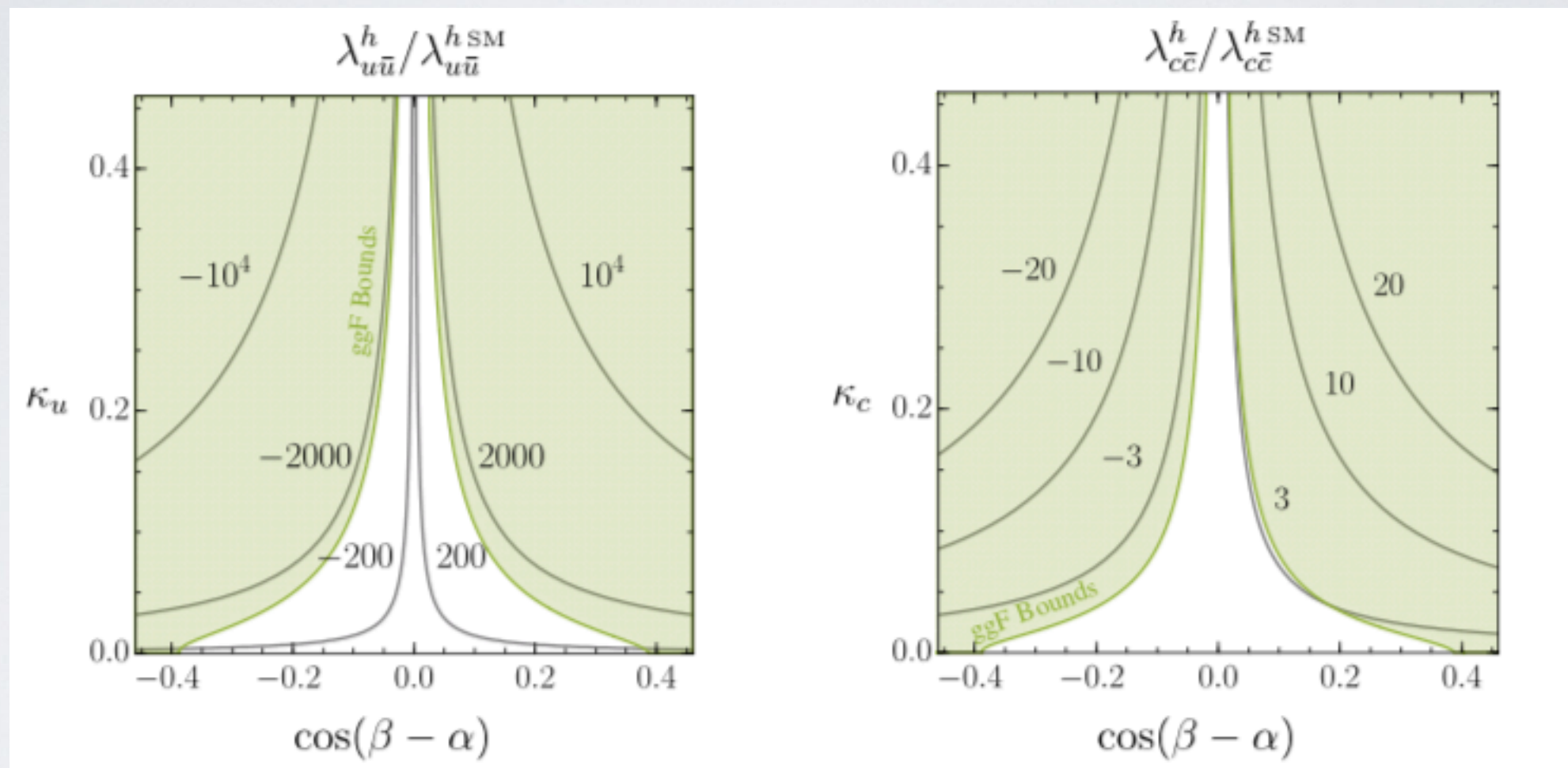
NO! IT CAN CHANGE THE  
PROPERTIES OF OUR HIGGS



# A REASON TO MEASURE LIGHT SM HIGGS YUKAWAS?



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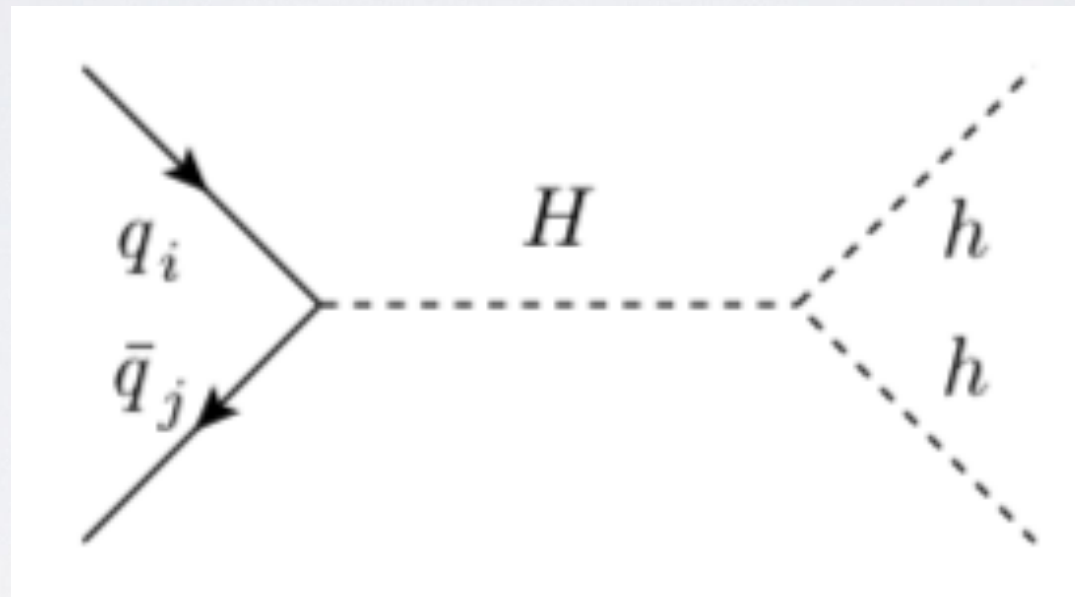




WHAT ELSE CAN IT DO FOR  
**OUR** HIGGS?

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QUARK INITIATED DI-HIGGS PRODUCTION!

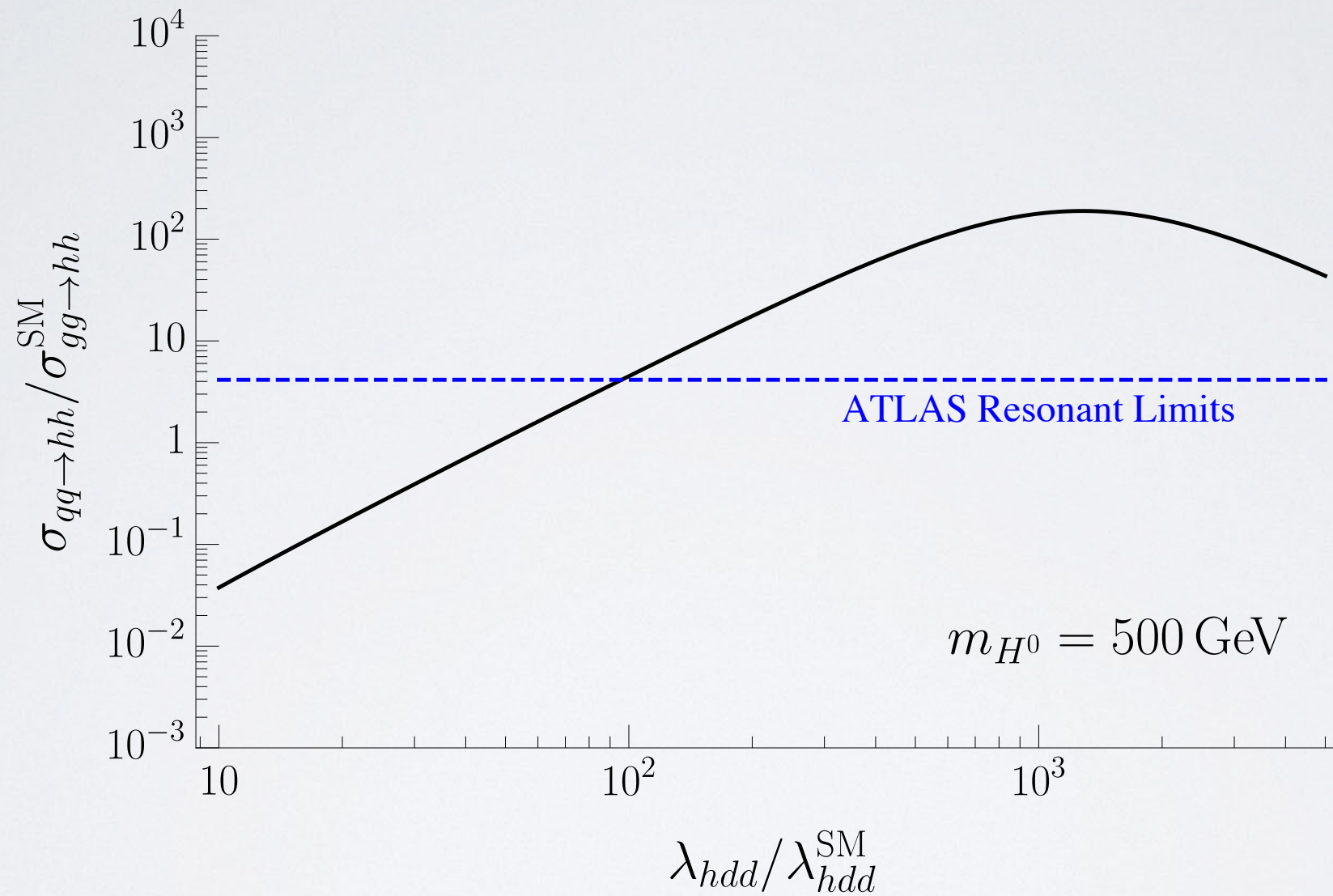


Others have thought about this, e.g:  
1801.00363 Bauer, Carena, Carmona  
1909.05279 Alasfar, Corral Lopez, Grober

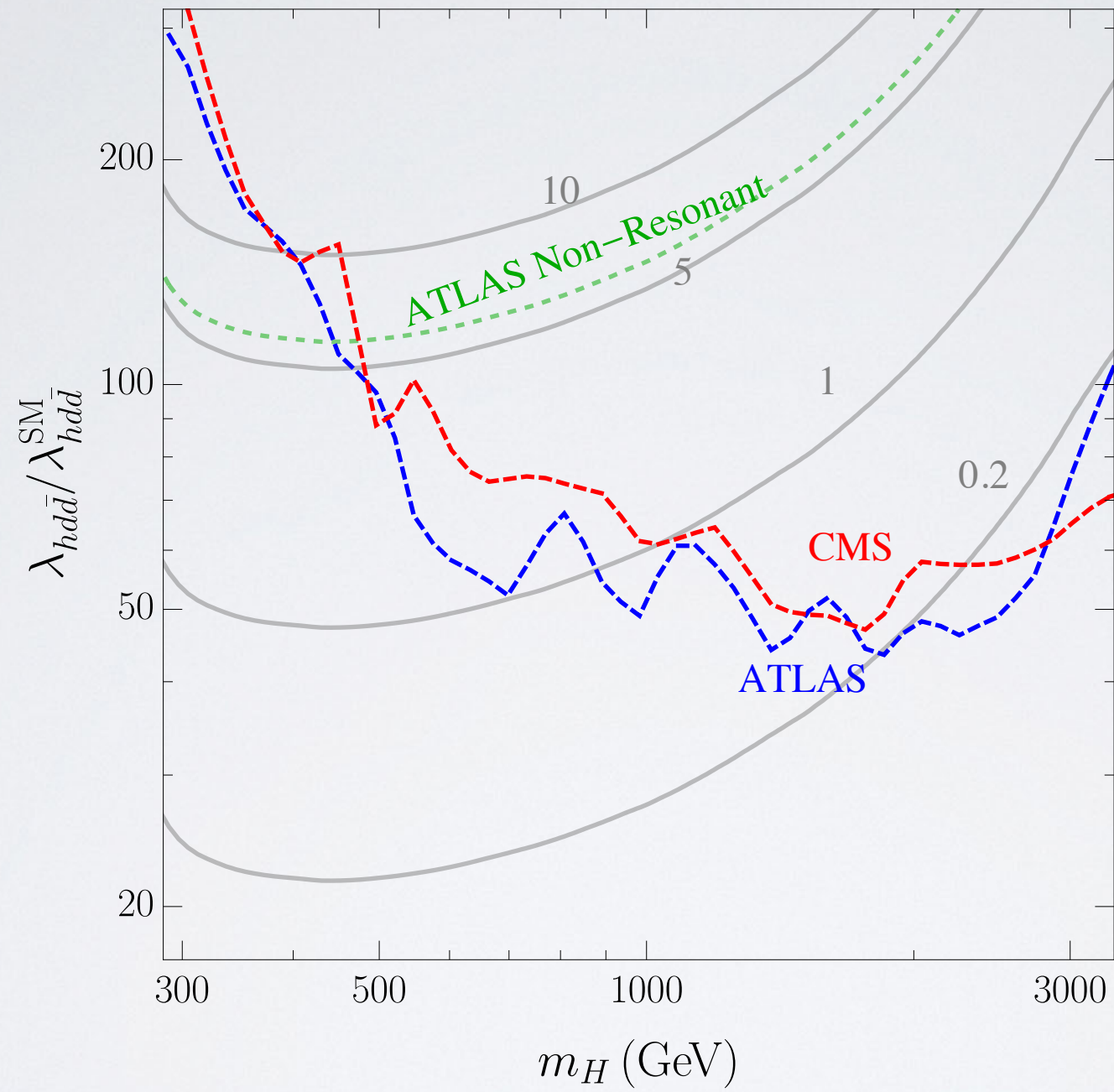
**Can get wider range of effects in SFV**



# DI-HIGGS

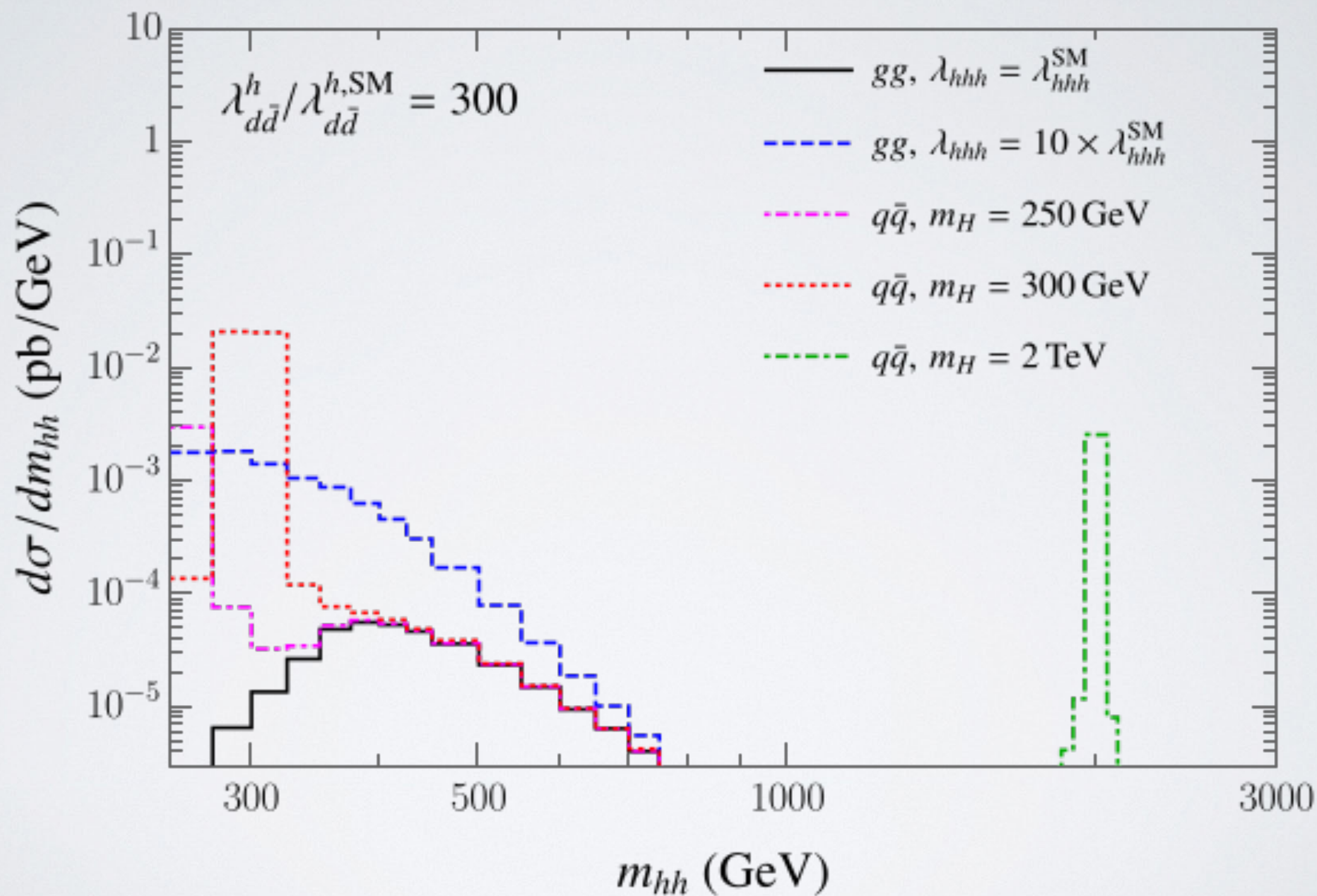


# DI-HIGGS





# DI-HIGGS



IS THIS JUST FOR HEAVY NEW  
PHYSICS OR THE HIGGS?



# VERY LIGHT NEW SCALARS WITH FLAVOR DEPENDENT COUPLINGS

Simply take the model just discussed, and add a singlet  $S$

Allow  $S$  to mix with the heavy Higgs

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From the low energy point of view...

$$\mathcal{L} \supset \frac{S}{M} \left[ c_{Sij}^u Q_i H \bar{u}_j - c_{Sij}^{d\dagger} Q_i H^c \bar{d}_j - c_{Sij}^{\ell\dagger} L_i H^c \bar{\ell}_j \right]$$

**I can inherit large flavor dependent couplings from up-type or down-type SFV!**

VERY LIGHT SCALAR WITH COUPLINGS  
TO THE 1ST GENERATION... E.G.

Light meson decay phenomenology!

$$\eta \rightarrow S\pi^0$$
$$S \rightarrow ee, \mu\mu, \pi\pi$$

Others have thought about flavor dependent light quark couplings e.g:  
1712.10022 Batell, Freitas, Ismail, Mckeen without a UV model



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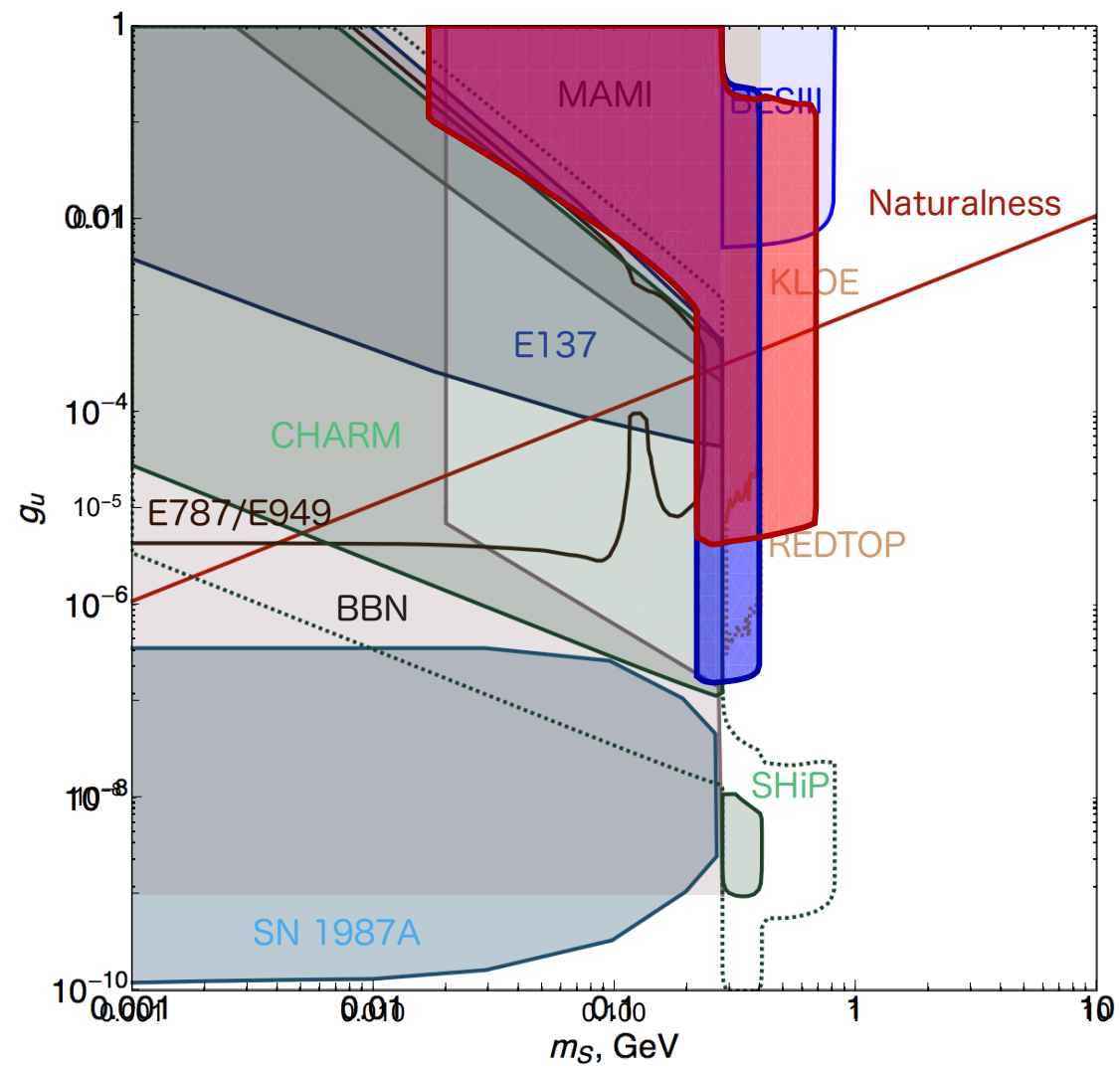
*REDTOP*

*The  $\eta/\eta'$  factory*



# *REDTOP*

*The  $\eta/\eta'$  factory*

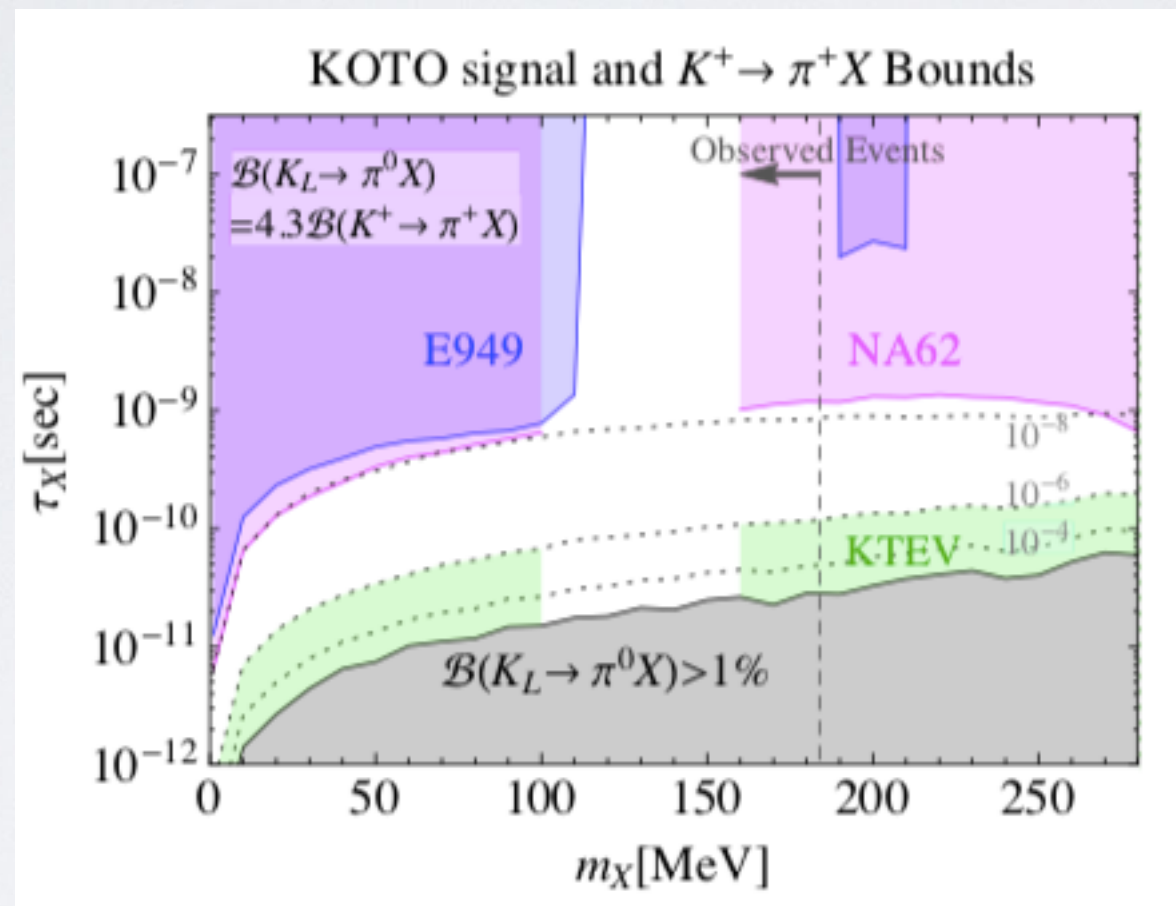




# WHAT COULD BE EVEN CRAZIER...

KOTO  $K \rightarrow \pi^0 \nu \bar{\nu}$

What if it was?  $K \rightarrow \pi^0 S$

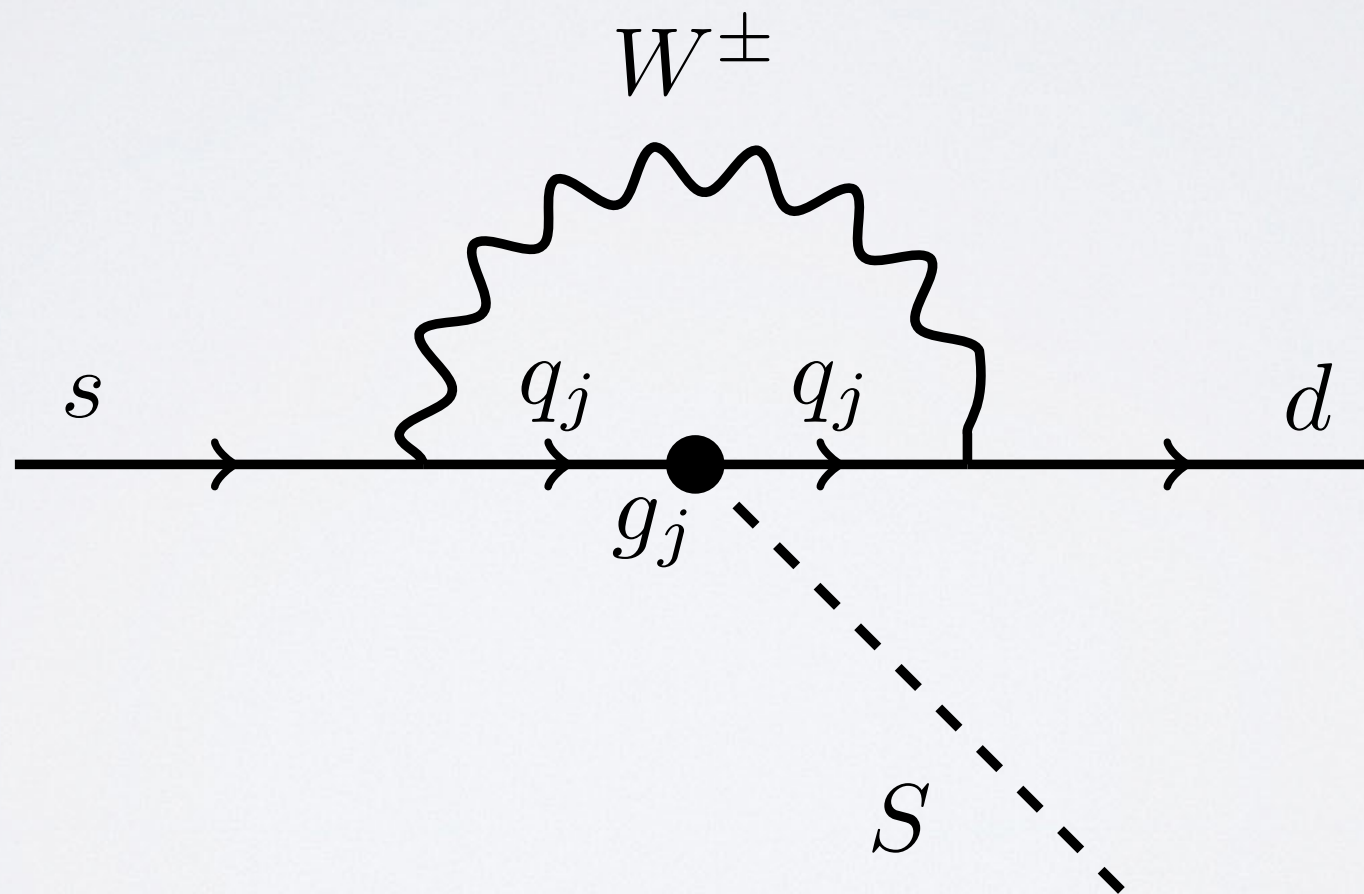


| 909. | | | | | . Kitahara, Okui, Perez, Soreq, Tobioka

didn't have a model just an explanation...

# WHAT COULD BE EVEN CRAZIER...

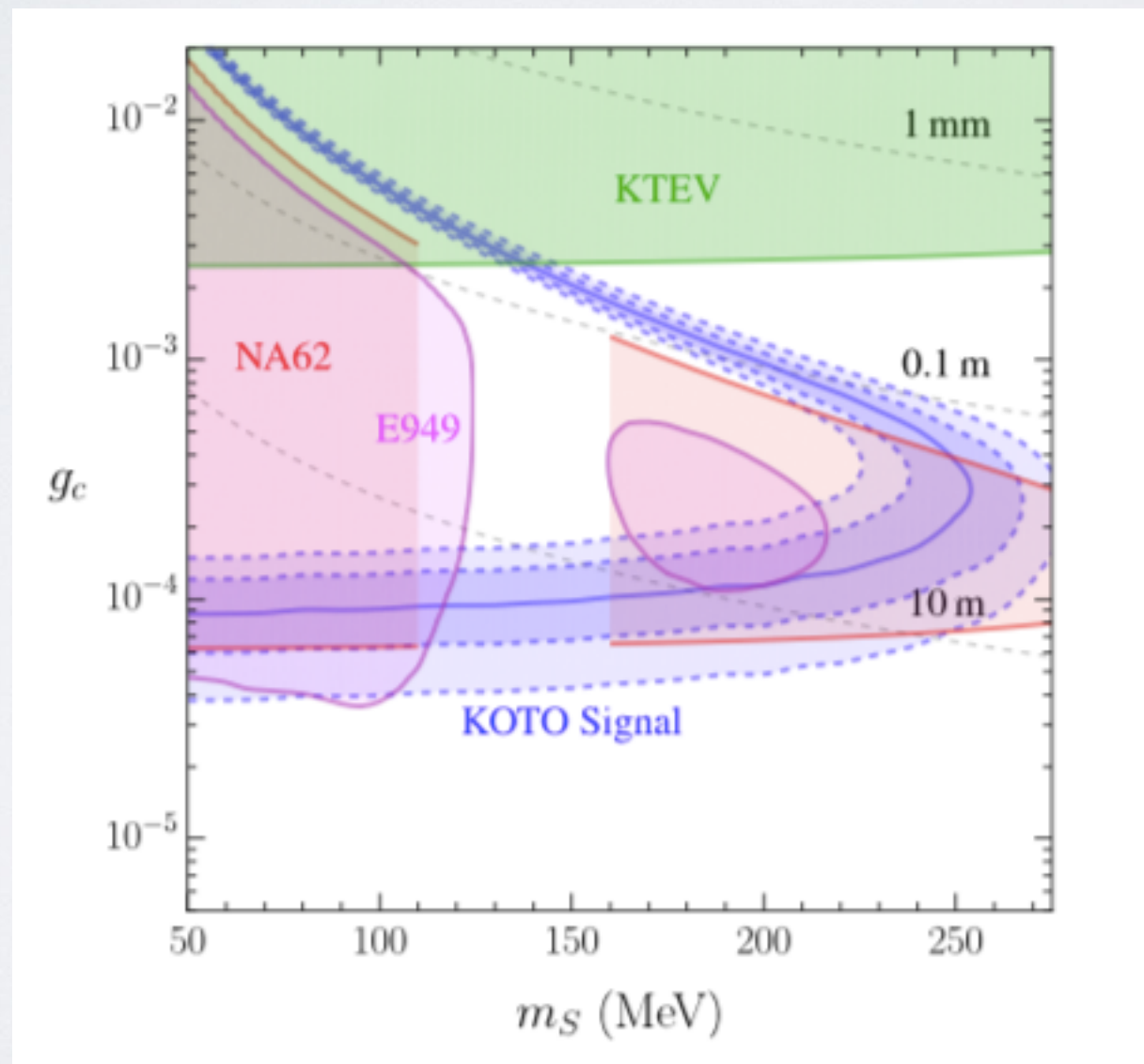
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# WHAT COULD BE EVEN CRAZIER...

KOTO  $K \rightarrow \pi^0 S$



# CONCLUSIONS

- Flavorful BSM physics is typically the third rail of theory... but avoiding it biases us to flavor universal or third generation searches
- Flavorful physics of the 1st and 2nd generation (and 3rd) CAN exist at LHC energies with Spontaneous Flavor Violation (SFV) AND have big cross sections while being compatible with current LHC constraints
- SFV can also **the** Higgs -
  - Light Yukawa couplings need measured!
  - Di-Higgs confusion with triple Higgs!
- Flavor dependence can also occur at LOW energies, so instead of just thinking about dark photons proportional to charge/mass and universal scalars, you need to think about individual quark flavor couplings if you want to cover the space!



FIN

# CAN WE GENERALIZE CONCEPT OF ALIGNMENT WITHOUT A FULL THEORY?

YES... Aligned Flavor Violation (AFV)

Introduce new spurions  $\kappa^u, \kappa^d$  transforming like  $y^u, y^d$

We want alignment, but without a “basis” dependence”

The mass eigenbasis is defined only up to a  $U(1)^6$  reparametrization symmetry

This symmetry is really what forbids tree level FCNCs...



# FLAVOR + REPARAMETRIZATION

$$y^u = U_{Q_u} Y^u U_{\bar{u}}^\dagger \equiv U_{Q_u} \text{diag}(y_u^{\text{SM}}, y_c^{\text{SM}}, y_t^{\text{SM}}) U_{\bar{u}}^\dagger$$
$$y^{d\dagger} = U_{Q_d} Y^d U_{\bar{d}}^\dagger \equiv U_{Q_d} \text{diag}(y_d^{\text{SM}}, y_s^{\text{SM}}, y_b^{\text{SM}}) U_{\bar{d}}^\dagger$$

$V$  is charged under reparametrization so we have to expand in it

$$\kappa^u = U_{Q_u} \left[ K^u + K^{u'} V^* K^{u'} V^T K^{u'''} + \mathcal{O}(V^4) \right] U_{\bar{u}}^\dagger$$
$$\kappa^{d\dagger} = U_{Q_d} \left[ K^d + K^{d'} V^T K^{d''} V^* K^{d'''} + \mathcal{O}(V^4) \right] U_{\bar{d}}^\dagger$$

$K$ 's are diagonal because of Reparametrization symmetry

# ALIGNED FLAVOR VIOLATION

- Interesting extension of MFV but...
- Relies on a “fictitious” symmetry
- And only is aligned up to CKM which isn't enough to get flavor physics down to the TeV scale



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Is there a way to get this, that isn't basis dependent and also doesn't require a full theory of flavor?



Operator	$\Lambda_{\text{NP}}^{\text{anarchic}}$ [TeV]	$\Lambda_{\text{NP}}^{\text{SFV}}$ [TeV]	$\Lambda_{\text{NP}}^{\text{MFV}}$ [TeV]
$(Q_1^\dagger \bar{\sigma}^\mu Q_2)^2$	$1.5 \times 10^4_{(\text{Im})}$	$262.7  \kappa_d^2 - \kappa_s^2 $	5.1
$(Q_1 \bar{d}_3)(Q_3^\dagger \bar{d}_1^\dagger)$	$2.1 \times 10^3_{(\text{Abs})}$	$19.3 \sqrt{ \kappa_d \kappa_b }$	—
$(Q_1 \bar{d}_2)(Q_2^\dagger \bar{d}_1^\dagger)$	$2.4 \times 10^5_{(\text{Im})}$	$72.7 \sqrt{ \kappa_d \kappa_s }$	—
$2eH\sigma^{\mu\nu} Q_2 \bar{d}_3 F_{\mu\nu}$	$276.3_{(\text{Re})}$	$54.3 \sqrt{ \kappa_b }$	7.0
$2eH\sigma^{\mu\nu} Q_3 \bar{d}_2 F_{\mu\nu}$	$276.3_{(\text{Re})}$	$54.3 \sqrt{ \kappa_s }$	7.0
$2eH\sigma^{\mu\nu} Q_3 \bar{d}_1 F_{\mu\nu}$	$140.5_{(\text{Abs})}$	$13.2 \sqrt{ \kappa_d }$	7.0

TABLE II. 95% CL bounds on the new physics scale  $\Lambda_{\text{NP}}$ , for anarchic, SFV and MFV operator coefficients (from [1, 31–33]). Subscripts on the anarchic operator limits indicates that the limit is on the real, imaginary or absolute value of the operator coefficient.

$$\begin{aligned}
V(H_1, H_2) &= m_1^2 H_1^\dagger H_1 + m_2^2 H_2^\dagger H_2 + \left( m_{12}^2 H_1^\dagger H_2 + \text{h.c.} \right) \\
&+ \frac{1}{2} \lambda_1 (H_1^\dagger H_1)^2 + \frac{1}{2} \lambda_2 (H_2^\dagger H_2)^2 + \lambda_3 (H_2^\dagger H_2) (H_1^\dagger H_1) + \lambda_4 (H_2^\dagger H_1) (H_1^\dagger H_2) \\
&+ \left[ \frac{1}{2} \lambda_5 (H_1^\dagger H_2)^2 + \lambda_6 H_1^\dagger H_1 H_1^\dagger H_2 + \lambda_7 (H_2^\dagger H_2) (H_1^\dagger H_2) + \text{h.c.} \right].
\end{aligned}$$

$$H_1 = \begin{pmatrix} 0 \\ H_1^0 \end{pmatrix} = \frac{1}{\sqrt{2}} \begin{pmatrix} 0 \\ v + h_1 \end{pmatrix}$$

$$H_2 = \begin{pmatrix} H^+ \\ H_2^0 \end{pmatrix} = \frac{1}{\sqrt{2}} \begin{pmatrix} \sqrt{2} H^+ \\ h_2 + i h_3 \end{pmatrix}$$

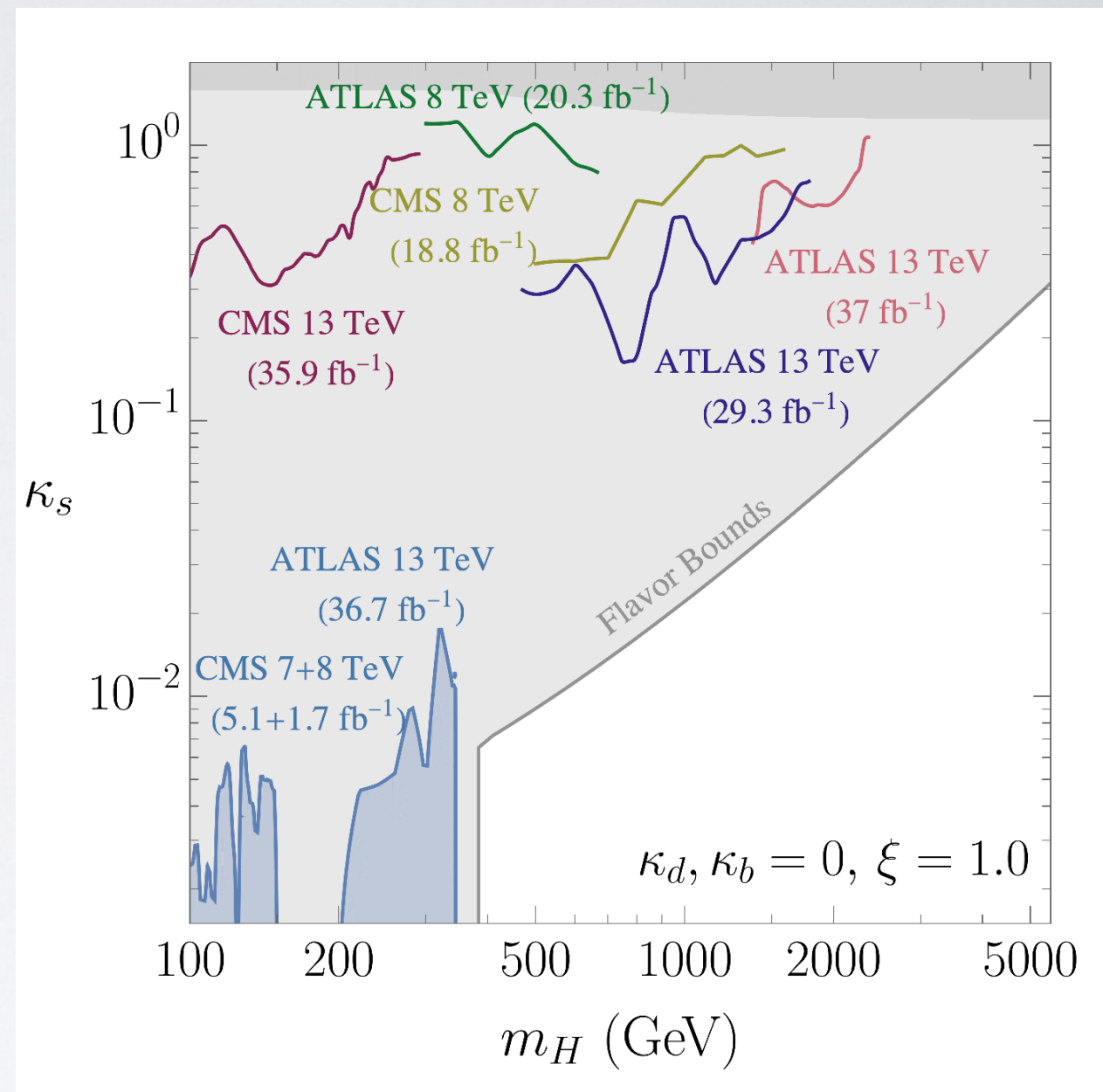
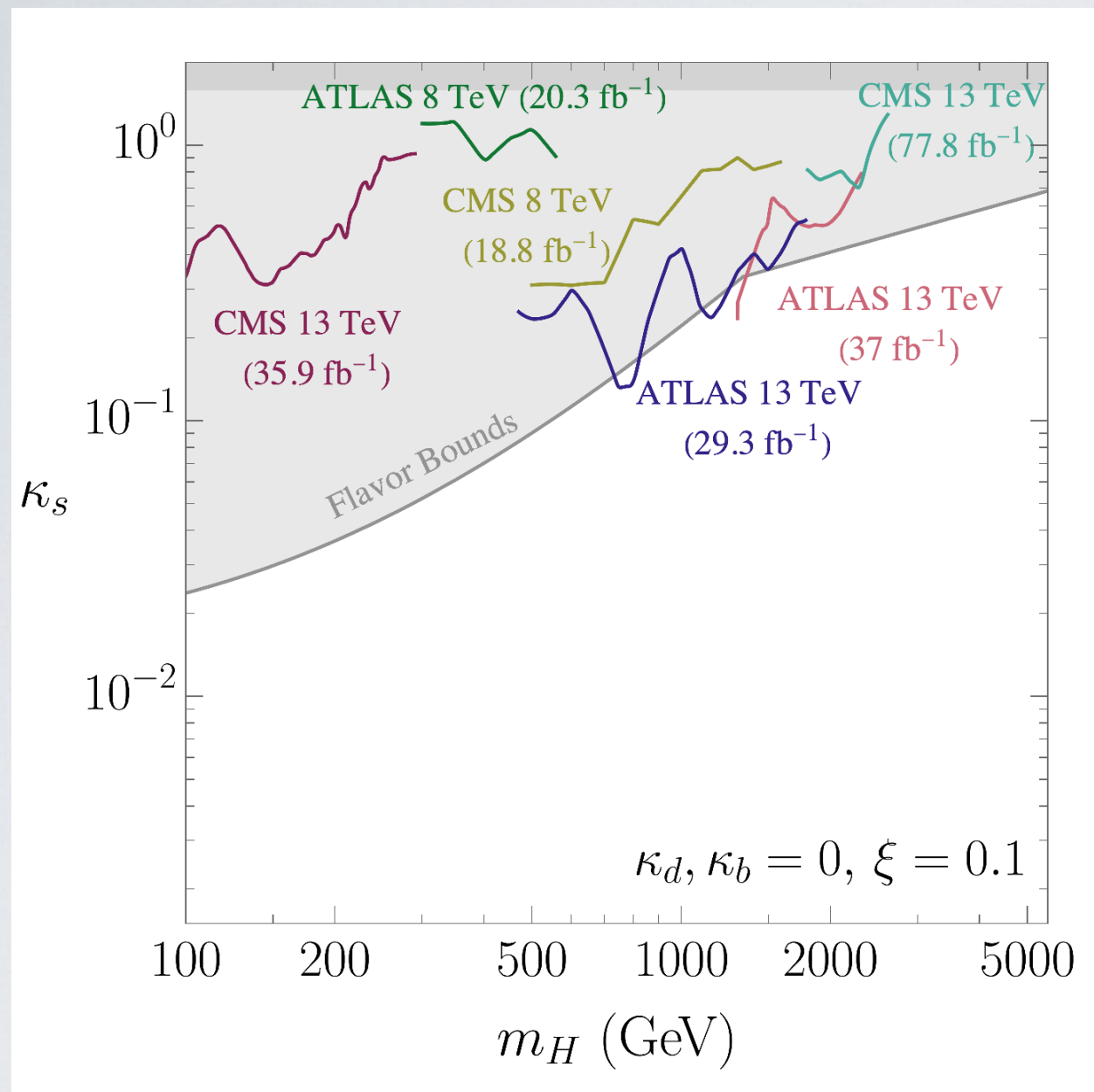
$$\begin{aligned}
\tan [2(\beta - \alpha)] &= \frac{-2\mathcal{M}_{12}^2}{\mathcal{M}_{22}^2 - \mathcal{M}_{11}^2} \\
&= \frac{2\lambda_6 v^2}{\lambda_1 v^2 - (m_2^2 + \frac{1}{2}(\lambda_3 + \lambda_4 + \lambda_5)v^2)}
\end{aligned}$$

$$h \equiv \sin(\beta - \alpha) h_1 + \cos(\beta - \alpha) h_2$$

$$H \equiv -\cos(\beta - \alpha) h_1 + \sin(\beta - \alpha) h_2$$

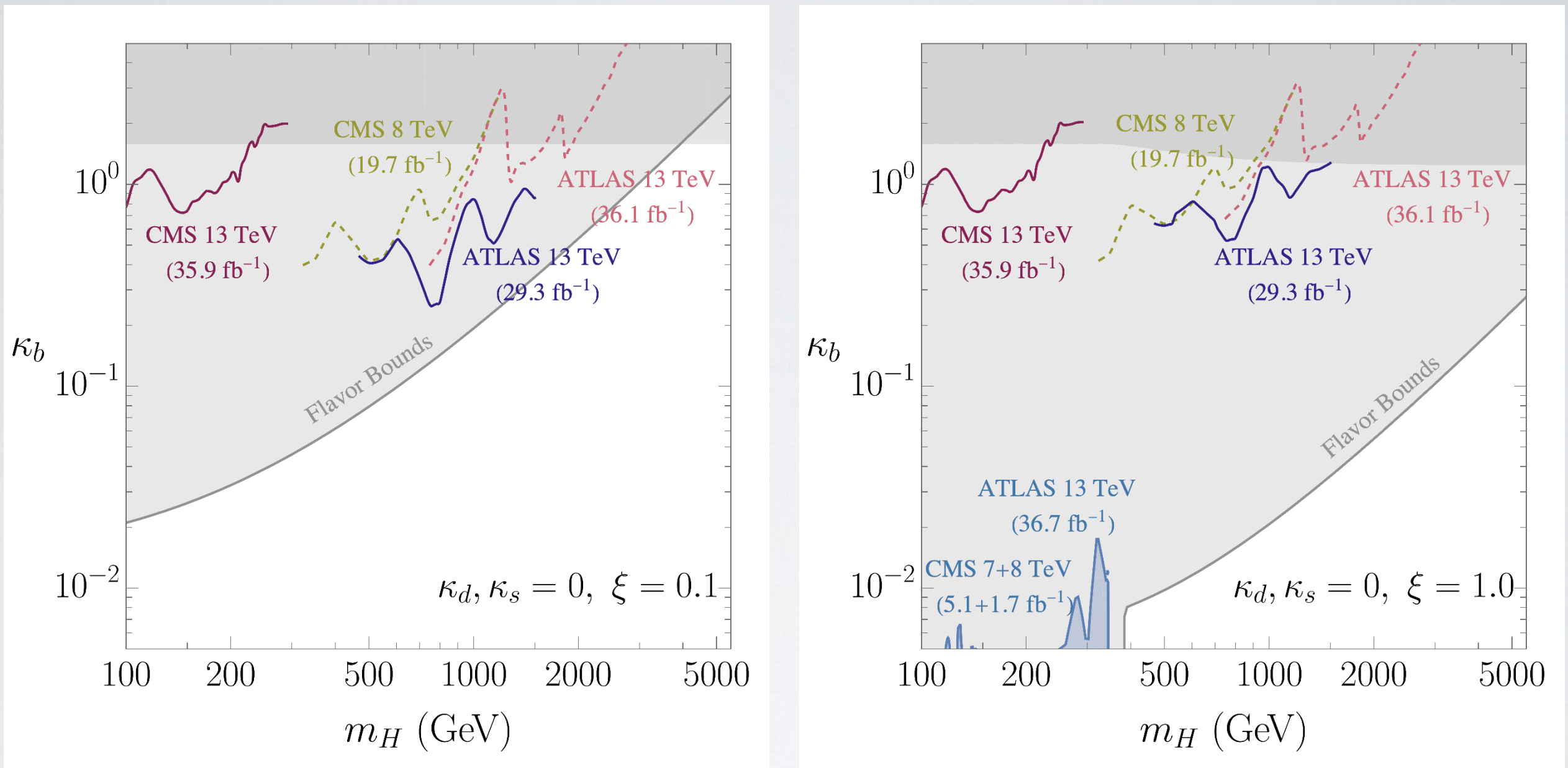


$\lambda_{hu_i\bar{u}_j}$	$\delta_{ij} Y_i^u [\sin(\beta - \alpha) + \xi \cos(\beta - \alpha)]$	$\lambda_{Hu_i\bar{u}_j}$	$\delta_{ij} Y_i^u [-\cos(\beta - \alpha) + \xi \sin(\beta - \alpha)] ,$
$\lambda_{hd_i\bar{d}_j}$	$\delta_{ij} [Y_i^d \sin(\beta - \alpha) + K_i^d \cos(\beta - \alpha)]$	$\lambda_{Hd_i\bar{d}_j}$	$\delta_{ij} [-Y_i^d \cos(\beta - \alpha) + K_i^d \sin(\beta - \alpha)]$
$\lambda_{h\ell_i\bar{\ell}_j}$	$\delta_{ij} Y_i^\ell [\sin(\beta - \alpha) + \xi^\ell \cos(\beta - \alpha)]$	$\lambda_{H\ell_i\bar{\ell}_j}$	$\delta_{ij} Y_i^\ell [-\cos(\beta - \alpha) + \xi^\ell \sin(\beta - \alpha)]$
$\lambda_{Au_i\bar{u}_j}$	$i\xi\delta_{ij} Y_i^u$	$\lambda_{H^+d_i\bar{u}_j}$	$-\left[\xi V^T Y^u\right]_{ij}$
$\lambda_{Ad_i\bar{d}_j}$	$-i\delta_{ij} K_i^d$	$\lambda_{H^-u_i\bar{d}_j}$	$\left[V^* K^d\right]_{ij}$
$\lambda_{A\ell_i\bar{\ell}_j}$	$-i\xi^\ell\delta_{ij} Y_i^\ell$	$\lambda_{H^- \ell_i\bar{\ell}_j}$	$\left[\xi^\ell Y^\ell\right]_{ij}$



**Figure 11:** The same as Fig. 10, but for  $\kappa_s$ , with  $\kappa_d = \kappa_b = 0$ .





**Figure 12:** The same as Fig. 10 but for  $\kappa_b$ , with  $\kappa_d = \kappa_s = 0$ . Solid lines indicate limits from ordinary dijet searches while the dashed lines indicate searches using  $b$ -tagging information (see text for details.)