

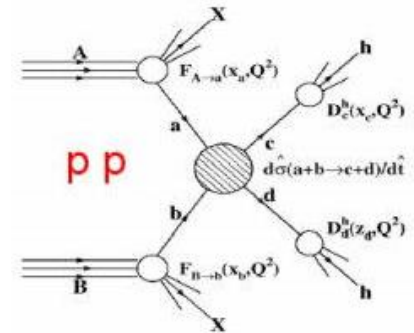


# The Muon Collider

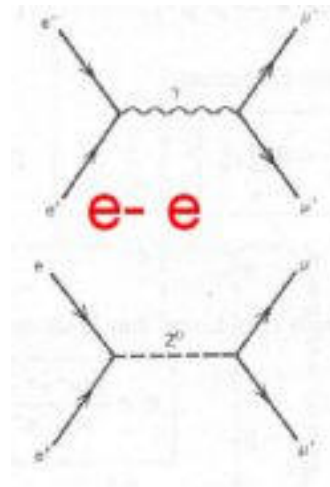
Daniel Schulte

# Collider Choices

- Hadron collisions: compound particles
  - Protons or ions
  - Mix of quarks, anti-quarks and gluons: variety of processes
  - Parton energy spread
  - QCD processes large background sources
  - **Hadron collisions**  $\Rightarrow$  can typically achieve higher collision energies



- Lepton collisions: elementary particles
  - Electrons, positrons and probably muons
  - Collision process known
  - Well defined energy
  - Less background
  - **Lepton collisions**  $\Rightarrow$  precision measurements



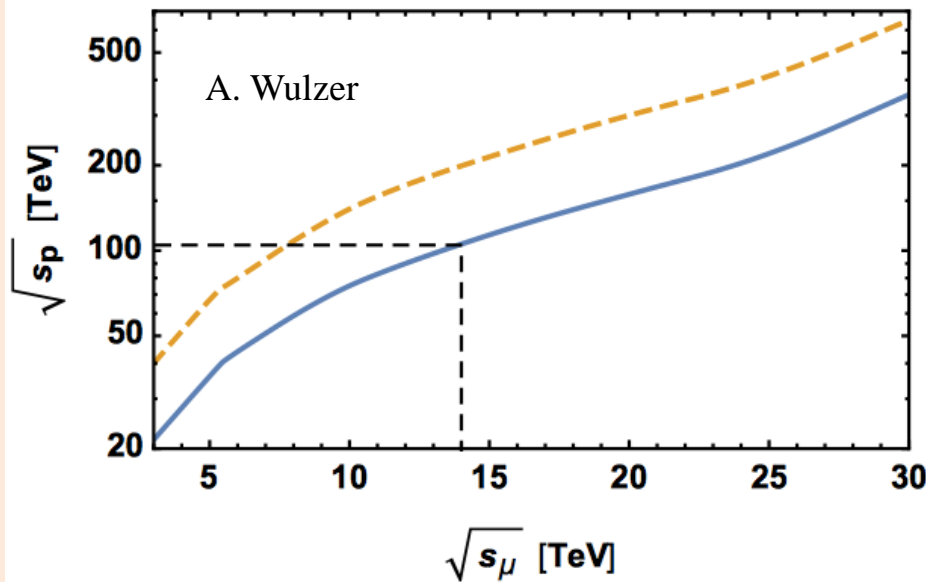
- Photons also possible

# Lepton Physics at High Energy

High energy lepton colliders are precision and discovery machines

$$V = \frac{1}{2} m_h^2 h^2 + (1 + k_3) \lambda_{hhh}^{SM} v h^3 + (1 + k_4) \lambda_{hhhh}^{SM} h^4$$

Chiesa, Maltoni, Mantani,  
Mele, Piccinini, Zhao  
[Muon Collider -  
Preparatory Meeting](#)



## Precision potential

Measure  $k_4$  to some 10%  
With 14 TeV, 20  $\text{ab}^{-1}$

## Discovery reach

14 TeV lepton collisions are comparable to  
100 TeV proton collisions for production of  
heavy particle pairs

## Luminosity goal

(Factor O(3) less than CLIC at 3 TeV)  
 $4 \times 10^{35} \text{ cm}^{-2} \text{ s}^{-1}$  at 14 TeV

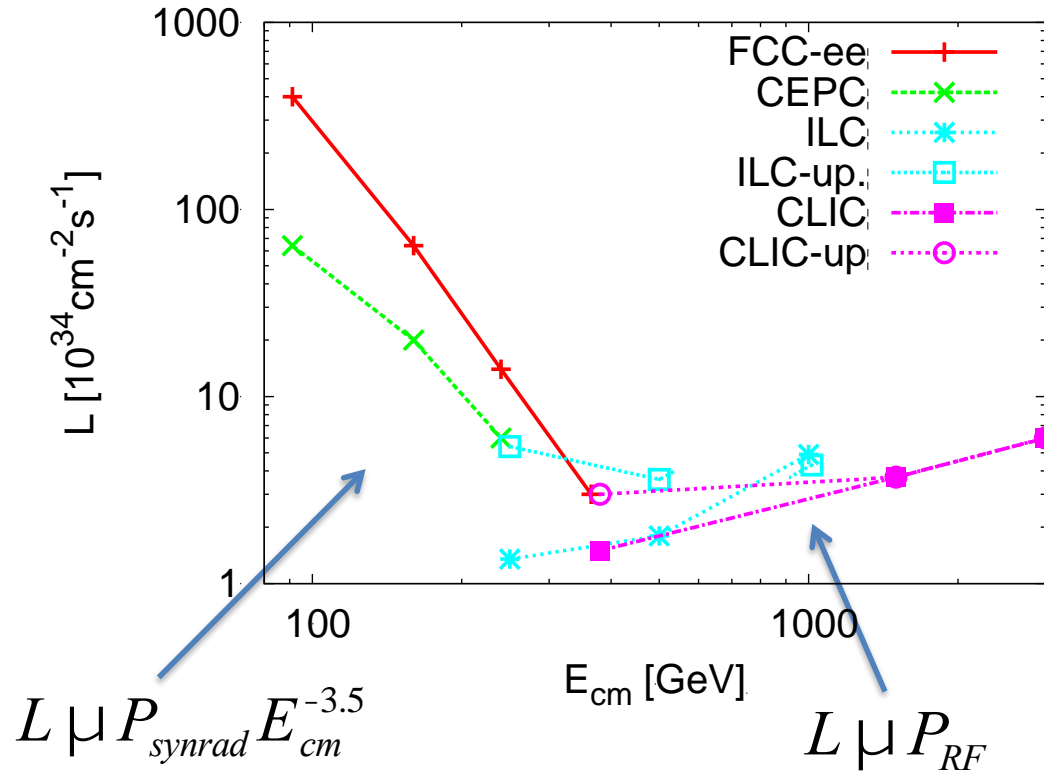
$$L \gtrsim \frac{5 \text{ years}}{\text{time}} \left( \frac{\sqrt{s_\mu}}{10 \text{ TeV}} \right)^2 2 \cdot 10^{35} \text{ cm}^{-2} \text{ s}^{-1}$$

# Comparisons of Projects with CDR

Project	Type	Energy [TeV]	Int. Lumi. [ $\text{a}^{-1}$ ]	Oper. Time [y]	Power [MW]	Cost
ILC	ee	0.25	2	11	129 (upgr. 150-200)	4.8-5.3 GILCU + upgrade
		0.5	4	10	163 (204)	7.8 GILCU
		1.0			300	?
CLIC	ee	0.38	1	8	168	5.9 GCHF
		1.5	2.5	7	(370)	+5.1 GCHF
		3	5	8	(590)	+7.3 GCHF
CEPC	ee	0.091+0.16	16+2.6		149	5 G\$
		0.24	5.6	7	266	
FCC-ee	ee	0.091+0.16	150+10	4+1	259	10.5 GCHF
		0.24	5	3	282	
		0.365 (+0.35)	1.5 (+0.2)	4 (+1)	340	
LHeC	ep	60 / 7000	1	12	(+100)	1.75 GCHF
FCC-hh	pp	100	30	25	580 (550)	17 GCHF (+7 GCHF)
HE-LHC	pp	27	20	20		7.2 GCHF

# Proposed Lepton Colliders (ESU)

## Luminosity per facility



Maximum proposed energy CLIC 3 TeV

- Cost estimate total of 18 GCHF
  - In three stages
  - Largely main linac, i.e. energy
- Power 590 MW
  - Part in luminosity, a part in energy
- Similar to FCC-hh (24 GCHF, 580 MW)

Technically possible to go higher in energy

But is it affordable?

$$L \gtrsim \frac{5 \text{ years}}{\text{time}} \left( \frac{\sqrt{s_\mu}}{10 \text{ TeV}} \right)^2 2 \cdot 10^{35} \text{ cm}^{-2} \text{ s}^{-1}$$

Cost roughly is linear with energy

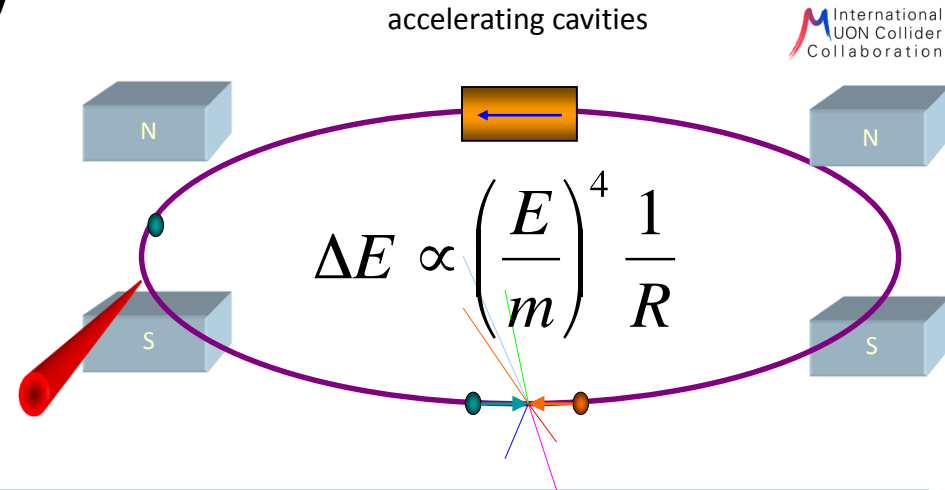
Power consumption roughly goes with the square of energy

# Energy Limit

**Electron-positron rings are multi-pass colliders limited by synchrotron radiation**

Strong dependence on particle mass

Hence **proton rings** are energy frontier



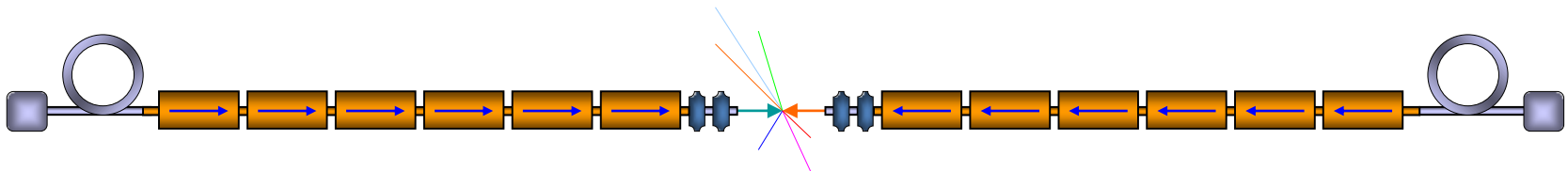
**Electron-positron linear colliders** avoid synchrotron radiation, but **single pass**

## Energy challenge

Need full voltage in main linac which is costly

## Luminosity challenge

Need very small beam size at collision is required, leads to strong beam-beam effects, requires extremely tight tolerances



# Comparing Luminosity in MAP vs. CLIC

CLIC is at the limit of what one can do (decades of R&D)

- No obvious way to improve

Luminosity per beam power increases with energy in muon collider

- **power efficient**

Site is **compact**

- 10 TeV comparable to 3 TeV CLIC

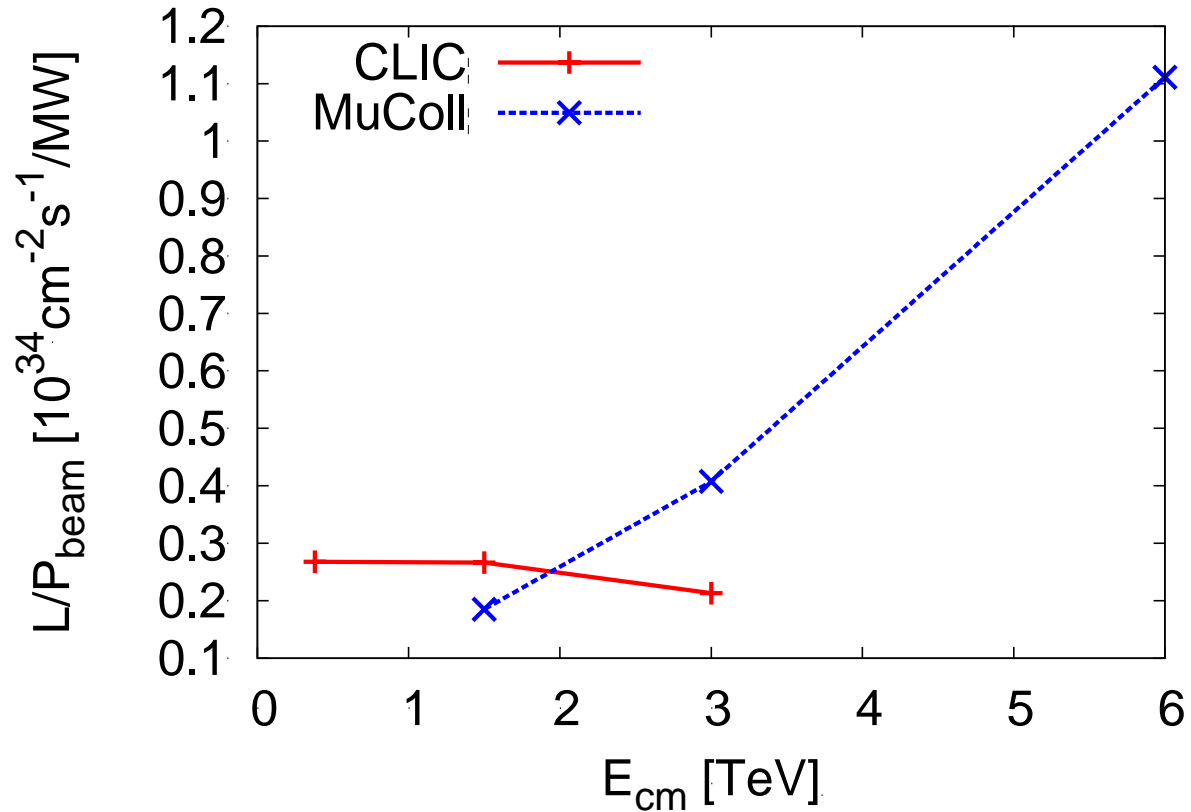
**Staging** is natural

- acceleration by a factor of a few is done in rings

Appears to promise **cost effectiveness**

- but need detailed study

Other **synergies** exist (neutrino/higgs)



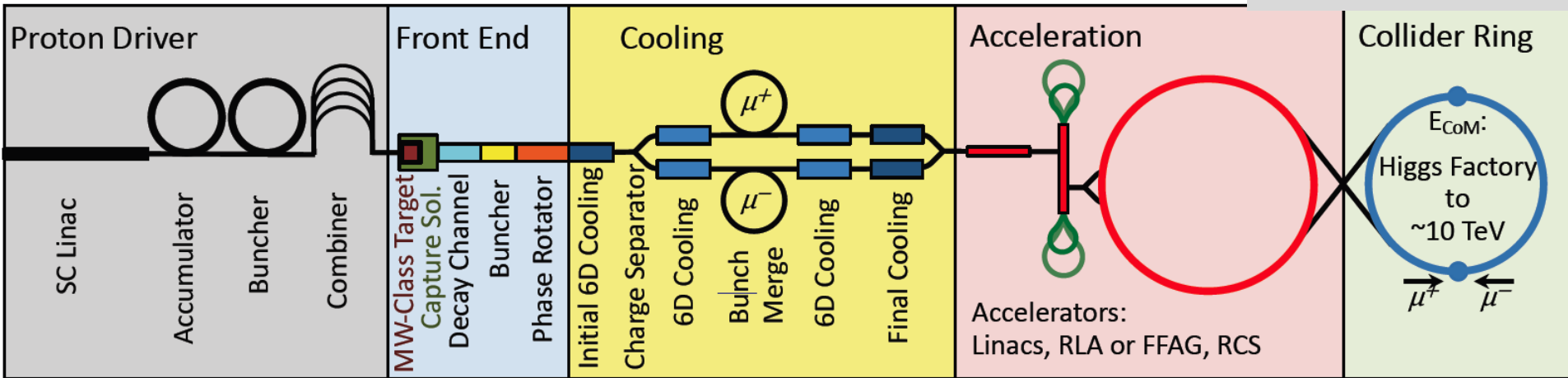
Muon collider promises unique opportunity for a **high-energy, high-luminosity lepton collider**

Main challenge **muons are not stable** (2.2  $\mu\text{s}$  at rest)

# Proton-driven Muon Collider Concept

The muon collider has been developed by the MAP collaboration mainly in the US  
 Muon cooling demonstration by MICE in the UK, some effort on alternative mainly at INFN

MAP collaboration



Short, intense proton bunches to produce hadronic showers

Protons produce pions  
 Pions decay to muons

Muon are captured, bunched and then cooled by ionisation cooling in matter

Acceleration to collision energy

Collision

Muon collider promises unique opportunity for a **high-energy, high-luminosity lepton collider**

Main challenge: **muons are not stable** (2.2  $\mu$ s at rest)



# Luminosity Goals

## Target integrated luminosities

$\sqrt{s}$	$\int \mathcal{L} dt$
3 TeV	1 ab <sup>-1</sup>
10 TeV	10 ab <sup>-1</sup>
14 TeV	20 ab <sup>-1</sup>

**Note: currently no staging**  
**Would only do 10 or 14 TeV**


- Tentative parameters achieve goal in 5 years
- FCC-hh to operate for 25 years
- Might integrate some margins
- Aim to have two detectors

**Now study if these parameters lead to realistic design with acceptable cost and power**

## Tentative target parameters Scaled from MAP parameters

Parameter	Unit	3 TeV	10 TeV	14 TeV
L	10 <sup>34</sup> cm <sup>-2</sup> s <sup>-1</sup>	1.8	20	40
N	10 <sup>12</sup>	2.2	1.8	1.8
f <sub>r</sub>	Hz	5	5	5
P <sub>beam</sub>	MW	5.3	14.4	20
C	km	4.5	10	14
<B>	T	7	10.5	10.5
ε <sub>L</sub>	MeV m	7.5	7.5	7.5
σ <sub>E</sub> / E	%	0.1	0.1	0.1
σ <sub>z</sub>	mm	5	1.5	1.07
β	mm	5	1.5	1.07
ε	μm	25	25	25
σ <sub>x,y</sub>	μm	3.0	0.9	0.63

Comparison:  
CLIC at 3 TeV: 28 MW



# Physics Potential



A. Wulzer et al.

Collaboration

The muon collider physics potential emerges from a **variety** of measurements and searches that offer **opportunities** for new physics **discoveries** that are **comparable** or **superior** to “standard” future colliders.

Our studies must be illustrative of the MC potential for new physics exploration in **multiple directions**.

Our plans for Snowmass21:

[https://indico.cern.ch/event/944012/contributions/3989516/attachments/2091456/3518021/Physics\\_SnowMass\\_Lol.pdf](https://indico.cern.ch/event/944012/contributions/3989516/attachments/2091456/3518021/Physics_SnowMass_Lol.pdf)

## Letter of Interest: Muon Collider Physics Potential

D. BUTTAZZO, R. CAPEDEVILLA, M. CHIESA, A. COSTANTINI, D. CURTIN, R. FRANCESCHINI,  
T. HAN, B. HEINEMANN, C. HELSENS, Y. KAHN, G. KRnjaIC, I. LOW, Z. LIU,  
F. MALTONI, B. MELE, F. MELONI, M. MORETTI, G. ORTONA, F. PICCININI, M. PIERINI,  
R. RATTAZZI, M. SELVAGGI, M. VOS, L.T. WANG, A. **WULZER**, M. ZANETTI, J. ZURITA

On behalf of the forming muon collider international collaboration [1]

We describe the plan for muon collider physics studies in order to provide inputs to the Snowmass process. The goal is a first assessment of the muon collider physics potential. The target accelerator design center of mass energies are 3 and 10 TeV or more [2]. Our study will consider energies  $E_{\text{CM}} = 3, 10, 14$ , and the more speculative  $E_{\text{CM}} = 30$  TeV, with reference integrated luminosities  $\mathcal{L} = (E_{\text{CM}}/10 \text{ TeV})^2 \times 10 \text{ ab}^{-1}$  [3]. Variations around the reference values are encouraged, aiming at an assessment of the required luminosity of the project based on physics performances. Recently, the physics potentials of several future collider options have been studied systematically [4], which provide reference points for comparison for our studies.

# Physics Potential



A. Wulzer et al.

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The muon collider physics potential emerges from a **variety** of measurements and searches that offer **opportunities** for new physics **discoveries** that are **comparable** or **superior** to “standard” future colliders.

Our studies must be illustrative of the MC potential for new physics exploration in **multiple directions**.

**And we are not alone**

Electroweak multiplets at the Muon Collider

R. Capdevilla, D. Curtin, Y. Kahn, G. Krnjaic, F. Meloni, J. Zurita

August 2020

## MUON COLLIDER: A WINDOW TO NEW PHYSICS

Douglas Berry<sup>1</sup>, Kevin Black<sup>2</sup>, Anadi Canepa<sup>1</sup>, Swapan Chattopadhyay<sup>1,3</sup>, Matteo Cremonesi<sup>1</sup>, Sridhara Dasu<sup>2</sup>, Dmitri Denisov<sup>4</sup>, Karri Di Petrillo<sup>1</sup>, Melissa Franklin<sup>5</sup>, Zoltan Gece<sup>1</sup>, Allison Hall<sup>1</sup>, Ulrich Heintz<sup>6</sup>, Christian Herwig<sup>1</sup>, James Hirschauer<sup>1</sup>, Tova Holmes<sup>7</sup>, Andrew Ivanov<sup>8</sup>, Bodhitha Jayatilaka<sup>1</sup>, Sergio Jindariani<sup>1</sup>, Young-Kee Kim<sup>9</sup>, Jacobo Konigsberg<sup>10</sup>, Lawrence Lee<sup>5</sup>, Miaoyuan Liu<sup>11</sup>, Zhen Liu<sup>12</sup>, Chang-Seong Moon<sup>13</sup>, Meenakshi Narain<sup>6</sup>, Scarlet Norberg<sup>14</sup>, Isobel Ojalvo<sup>15</sup>, Katherine Pachal<sup>16</sup>, Simone Pagan Griso<sup>17</sup>, Kevin Pedro<sup>1</sup>, Alex Perloff<sup>18</sup>, Elodie Resseguie<sup>17</sup>, Stefan Spanier<sup>7</sup>, Maximilian Swiatlowski<sup>19</sup>, Ann Miao Wang<sup>5</sup>, Lian-Tao Wang<sup>9</sup>, Xing Wang<sup>20</sup>, Hannsjörg Weber<sup>1\*</sup>, David Yu<sup>6</sup>

<sup>1</sup>Fermi National Accelerator Laboratory, <sup>2</sup>University of Wisconsin, Madison, <sup>3</sup>Northern Illinois University, <sup>4</sup>Brookhaven National Laboratory, <sup>5</sup>Harvard University, <sup>6</sup>Brown University, <sup>7</sup>University of Tennessee, Knoxville, <sup>8</sup>Kansas State University, <sup>9</sup>University of Chicago, <sup>10</sup>University of Florida, <sup>11</sup>Purdue University, <sup>12</sup>University of Maryland, <sup>13</sup>Kyungpook National University, <sup>14</sup>University of Puerto Rico, Mayagüez, <sup>15</sup>Princeton University, <sup>16</sup>Duke University, <sup>17</sup>Lawrence Berkeley National Laboratory, <sup>18</sup>University of Colorado, Boulder, <sup>19</sup>TRIUMF <sup>20</sup>University of California, San Diego

**Letter of Interest: EW effects in very high-energy phenomena**

C. ARINA, G. CUOMO, T. HAN, Y. MA, F. MALTONI, A. MANOHAR, S. PRESTEL, R. RUIZ, L. VECCHI, R. VERHEYEN, B. WEBBER, W. WAALEWIJN, A. WULZER, K. XIE to be submitted to the Theory Frontier (TF07) and Energy Frontier (EF04)

## HIGGS AND ELECTROWEAK PHYSICS AT THE MUON COLLIDER: AIMING FOR PRECISION AT THE HIGHEST ENERGIES

Aram Apyan<sup>1</sup>, Jeff Berryhill<sup>1</sup>, Pushpa Bhat<sup>1</sup>, Kevin Black<sup>2</sup>, Elizabeth Brost<sup>3</sup>, Anadi Canepa<sup>1</sup>, Sridhara Dasu<sup>2</sup>, Dmitri Denisov<sup>3</sup>, Karri Di Petrillo<sup>1</sup>, Zoltan Gece<sup>1</sup>, Tao Hann<sup>4</sup>, Ulrich Heintz<sup>5</sup>, Rachel Hyneman<sup>6</sup>, Young-Kee Kim<sup>7</sup>, Da Liu<sup>8</sup>, Mia Liu<sup>9</sup>, Zhen Liu<sup>10</sup>, Ian Low<sup>11,12</sup>, Sergio Jindariani<sup>1</sup>, Chang-Seong Moon<sup>13</sup>, Isobel Ojalvo<sup>14</sup>, Meenakshi Narain<sup>5</sup>, Maximilian Swiatlowski<sup>15\*</sup>, Marco Valente<sup>15</sup>, Lian-Tao Wang<sup>7</sup>, Xing Wang<sup>16</sup>, Hannsjörg Weber<sup>1</sup>, David Yu<sup>5</sup>

**Muon Collider: Study of Higgs couplings and self-couplings precision**

C. Aimè<sup>a</sup>, F. Balli<sup>b</sup>, N. Bartosik<sup>c</sup>, L. Buonincontri<sup>d</sup>, M. Casarsa<sup>e</sup>, M. Chiesa<sup>f</sup>, F. Collamati<sup>g</sup>, C. Curatolo<sup>d</sup>, D. Lucchesi<sup>d</sup>, B. Mele<sup>g</sup>, F. Maltoni<sup>h</sup>, B. Mansoulié<sup>b</sup>, A. Nisati<sup>g</sup>, N. Pastrone<sup>e</sup>, F. Piccinini<sup>i</sup>, C. Riccardi<sup>a</sup>, P. Sala<sup>l</sup>, P. Salvini<sup>i</sup>, L. Sestini<sup>m</sup>, I. Vai<sup>a</sup>, D. Zuliani<sup>d</sup>

**Beyond the Standard Model with High-Energy Lepton Colliders**

Hind Al Ali<sup>1</sup>, Nima Arkani-Hamed<sup>2</sup>, Ian Banta<sup>1</sup>, Sean Benevides<sup>1</sup>, Tianji Cai<sup>1</sup>, Junyi Cheng<sup>1</sup>, Tim Cohen<sup>3</sup>, Nathaniel Craig<sup>1</sup>, JiJi Fan<sup>4</sup>, Isabel Garcia Garcia<sup>5</sup>, Seth Koren<sup>6,1</sup>, Giacomo Koszegi<sup>1</sup>, Zhen Liu<sup>7</sup>, Kunfeng Lyu<sup>8</sup>, Amara McCune<sup>1</sup>, Patrick Meade<sup>9</sup>, Isobel Ojalvo<sup>10</sup>, Umut Oktem<sup>1</sup>, Matthew Reece<sup>11</sup>, Raman Sundrum<sup>7</sup>, Dave Sutherland<sup>12</sup>, Timothy Trott<sup>1</sup>, Chris Tully<sup>10</sup>, Ken Van Tilburg<sup>5</sup>, Lian-Tao Wang<sup>6</sup>, and Menghang Wang<sup>1</sup>

# Muon Collider Collaboration and Panel



Following the European Strategy, Laboratory Directors Group (LDG) initiated an **International Muon Collider Collaboration**  
A **Memorandum of Cooperation** can be signed to join

**Council** charged LDG to deliver a **European Accelerator R&D Roadmap**

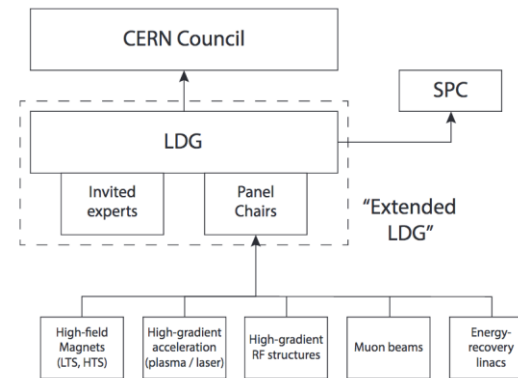
The extended LDG will deliver to **council** a report with a **prioritised workplan**

Five panels:

Magnets: P. Vedin, Plasma: R. Assmann, RF: S. Bousson, Muons: D. Schulte, ERL: M. Klein

Muon Panel Members: Daniel Schulte (CERN, chair), Mark Palmer (BNL, co-chair, Tabea Arndt (KIT), Antoine Chance (CEA/IRFU), Jean-Pierre Delahaye (retired), Angeles Faus-Golfe (IN2P3/IJClab), Simone Gilardoni (CERN), Philippe Lebrun (European Scientific Institute), Ken Long (Imperial College London), Elias Metral (CERN), Nadia Pastrone (INFN-Torino), Lionel Quettier (CEA/IRFU), Magnet Panel link, Tor Raubenheimer (SLAC), Chris Rogers (STFC-RAL), Mike Seidel (EPFL and PSI), Diktys Stratakis (FNAL), Akira Yamamoto (KEK and CERN)

Contributors: Alexej Grudiev (CERN), RF panel link, Roberto Losito (CERN), Test Facility link, Donatella Lucchesi (INFN) MDI link



# International Muon Collider Collaboration



## Objective:

In time for the next European Strategy for Particle Physics Update, the study aims to **establish whether the investment into a full CDR and a demonstrator is scientifically justified.**

It will provide a baseline concept, well-supported performance expectations and assess the associated key risks as well as cost and power consumption drivers. It will also identify an R&D path to demonstrate the feasibility of the collider.

## Scope:

- Focus on two energy ranges:
  - **3 TeV**, if possible with technology ready for **construction in 15-20 years**
  - **10+ TeV**, with more advanced technology, **the reason to do muon colliders**
- Explore synergy with other options (neutrino facility/higgs factory at resonance)
- Define **R&D path**

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If no higgs factory not in Europe, might be candidate for next large project after LHC, i.e. operation mid 2040s would need massive ramp-up after next ESPPU

If interest after other higgs factory or if no other higgs factory

# Community Meeting Convener



Conveners list (to be updated)

**Radio-Frequency (RF):** Alexej Grudiev (CERN), Jean-Pierre Delahaye (CERN retiree), Derun Li (LBNL), Akira Yamamoto (KEK).

**Magnets:** Lionel Quettier (CEA), Toru Ogitsu (KEK), Soren Prestemon (LBNL), Sasha Zlobin (FNAL), Emanuela Barzi (FNAL).

**High-Energy Complex (HEC):** Antoine Chance (CEA), J. Scott Berg (BNL), Alex Bogacz (JLAB), Christian Carli (CERN), Angeles Faus-Golfe (IJCLab), Eliana Gianfelice-Wendt (FNAL), Shinji Machida (RAL).

**Muon Production and Cooling (MPC):** Chris Rogers (RAL), Marco Calviani (CERN), Chris Densham (RAL), Diktys Stratakis (FNAL), Akira Sato (Osaka University), Katsuya Yonehara (FNAL).

**Proton Complex (PC):** Simone Gilardoni (CERN), Hannes Bartosik (CERN), Frank Gerigk (CERN), Natalia Milas (ESS).

**Beam Dynamics (BD):** Elias Metral (CERN), Tor Raubenheimer (SLAC and Stanford University), Rob Ryne (LBNL).

**Radiation Protection (RP):** Claudia Ahdida (CERN).

**Parameters, Power and Cost (PPC):** Daniel Schulte (CERN), Mark Palmer (BNL), Jean-Pierre Delahaye (CERN retiree), Philippe Lebrun (CERN retiree and ESI), Mike Seidel (PSI), Vladimir Shiltsev (FNAL), Jingyu Tang (IHEP), Akira Yamamoto (KEK).

**Machine Detector Interface (MDI):** Donatella Lucchesi (University of Padova), Christian Carli (CERN), Anton Lechner (CERN), Nicolai Mokhov (FNAL), Nadia Pastrone (INFN), Sergo R Jindariani (FNAL).

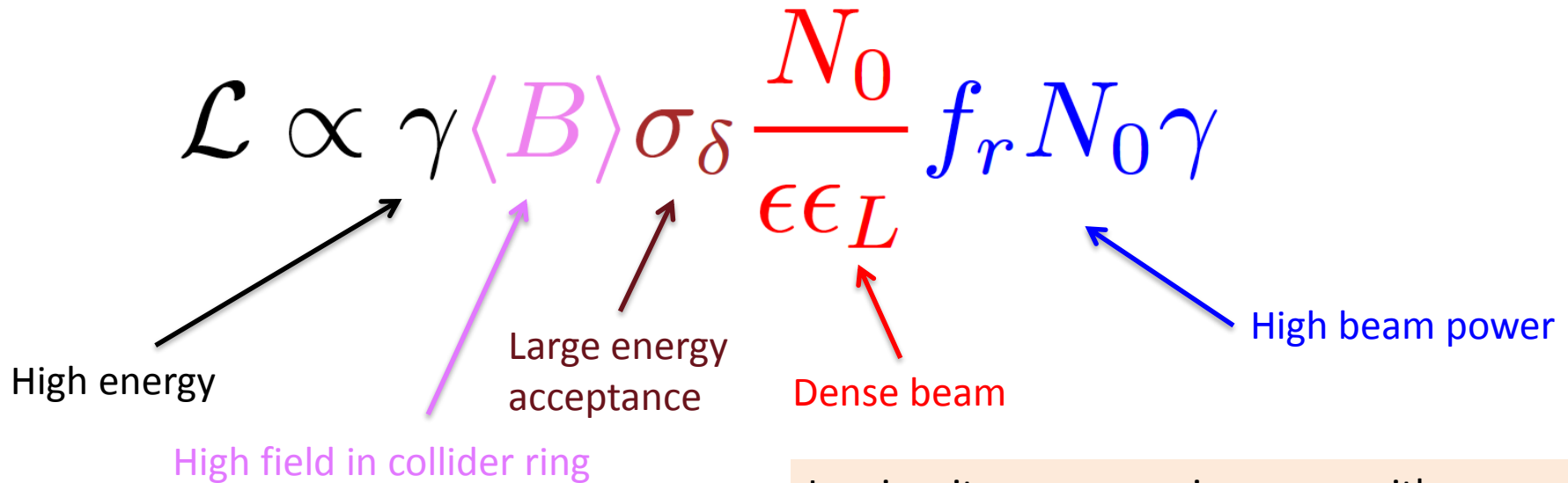
**Synergy:** Kenneth Long (Imperial College), Roger Ruber (Uppsala University), Koichiro Shimomura (KEK).

**Test Facility (TF):** Roberto Losito (CERN), Alan Bross (FNAL), Tord Ekelof (ESS, Uppsala University).

# Muon Collider Luminosity Drivers

Fundamental limitation

Requires emittance preservation and advanced lattice design



$$L \gtrsim \frac{5 \text{ years}}{\text{time}} \left( \frac{\sqrt{s}_\mu}{10 \text{ TeV}} \right)^2 2 \cdot 10^{35} \text{ cm}^{-2} \text{ s}^{-1}$$

Luminosity per power increases with energy  
 Provided all technical limits can be solved

Constant current for required luminosity

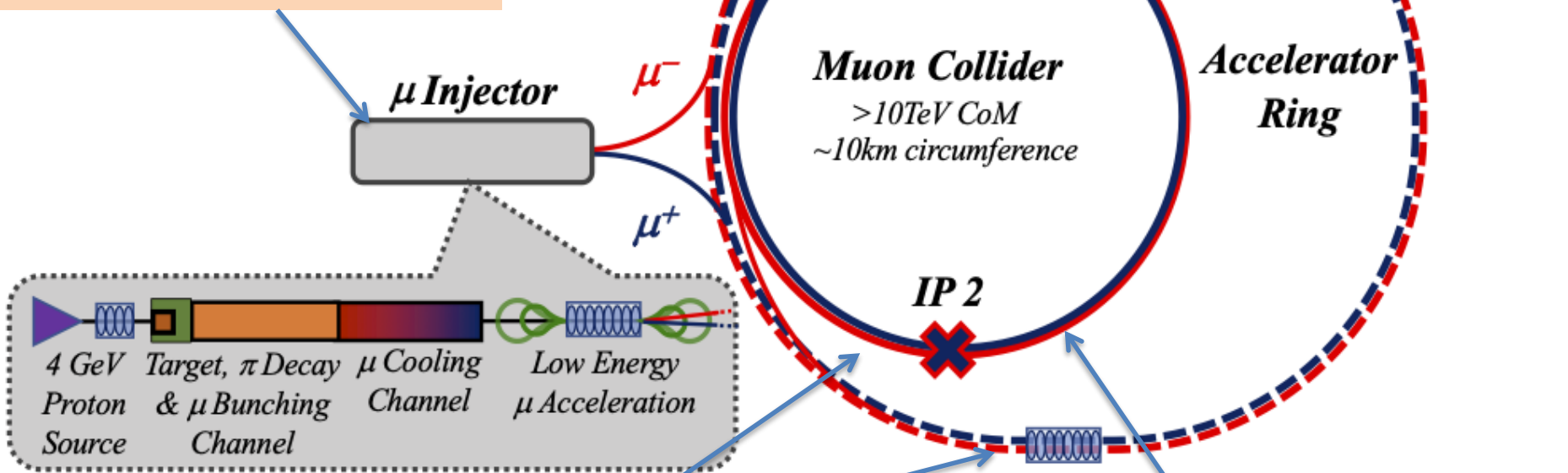
Better scaling than linear colliders



# Key Challenges

Drives the **beam quality**  
quite detailed MAP design  
still challenging design with  
challenging components  
*optimise as much as possible*

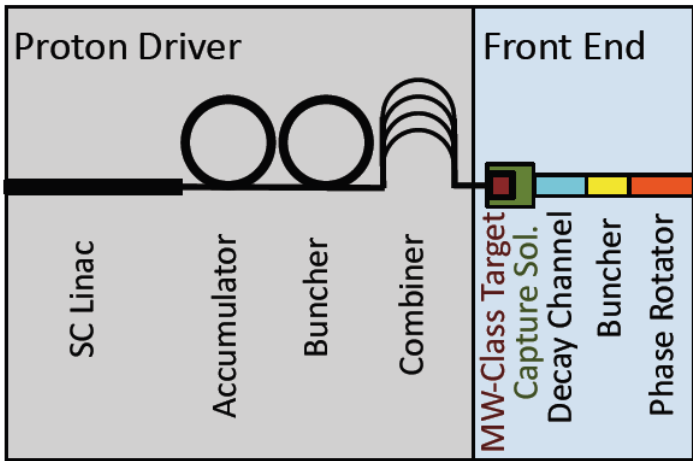
**Beam induced  
background**



**Cost and power** consumption drivers, limit energy reach  
e.g. 30 km accelerator for 10/14 TeV, 10/14 km collider ring  
Also impacts **beam quality**

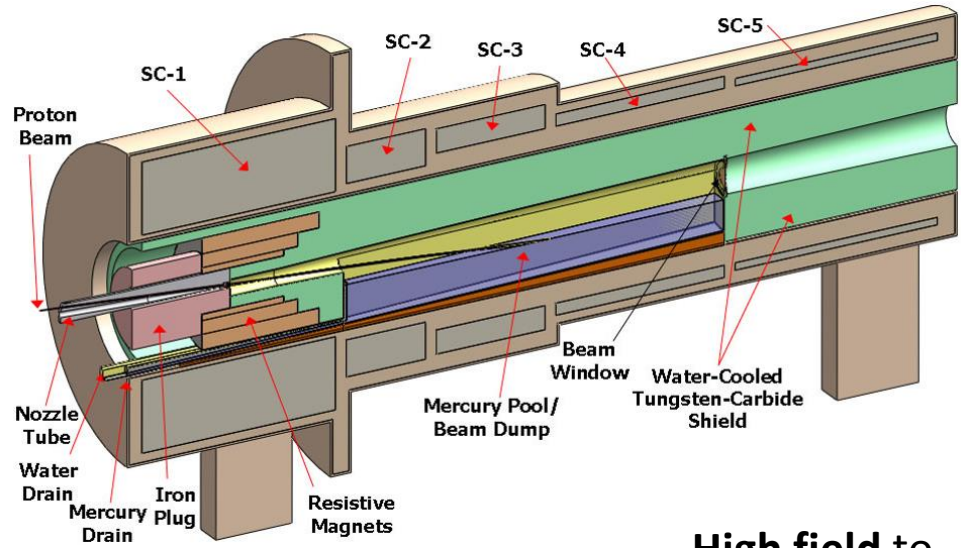
**Dense neutrino flux**  
mitigated by mover system  
and site selection

# Proton Complex and Target Area



Proton beam power is no issue, some look required at  
**H- source and accumulator complex**

**2 MW proton beam** requires radiation protection mainly of the solenoid

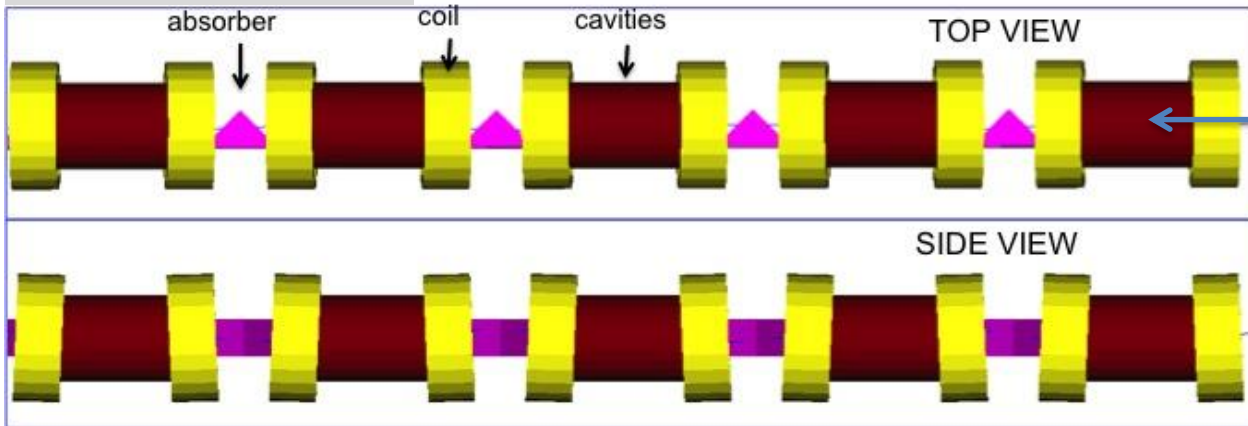


**Large aperture**  $O(1m)$  to allow shielding

**High field** to efficiently collect pions/muons: 20 T then tapering

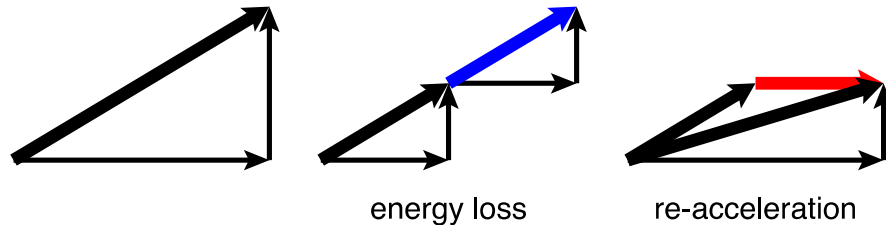
# Cooling Concept

MAP collaboration



Limit muon decay, cavities with **high gradient in a magnetic field** tests much better than design values but need to develop

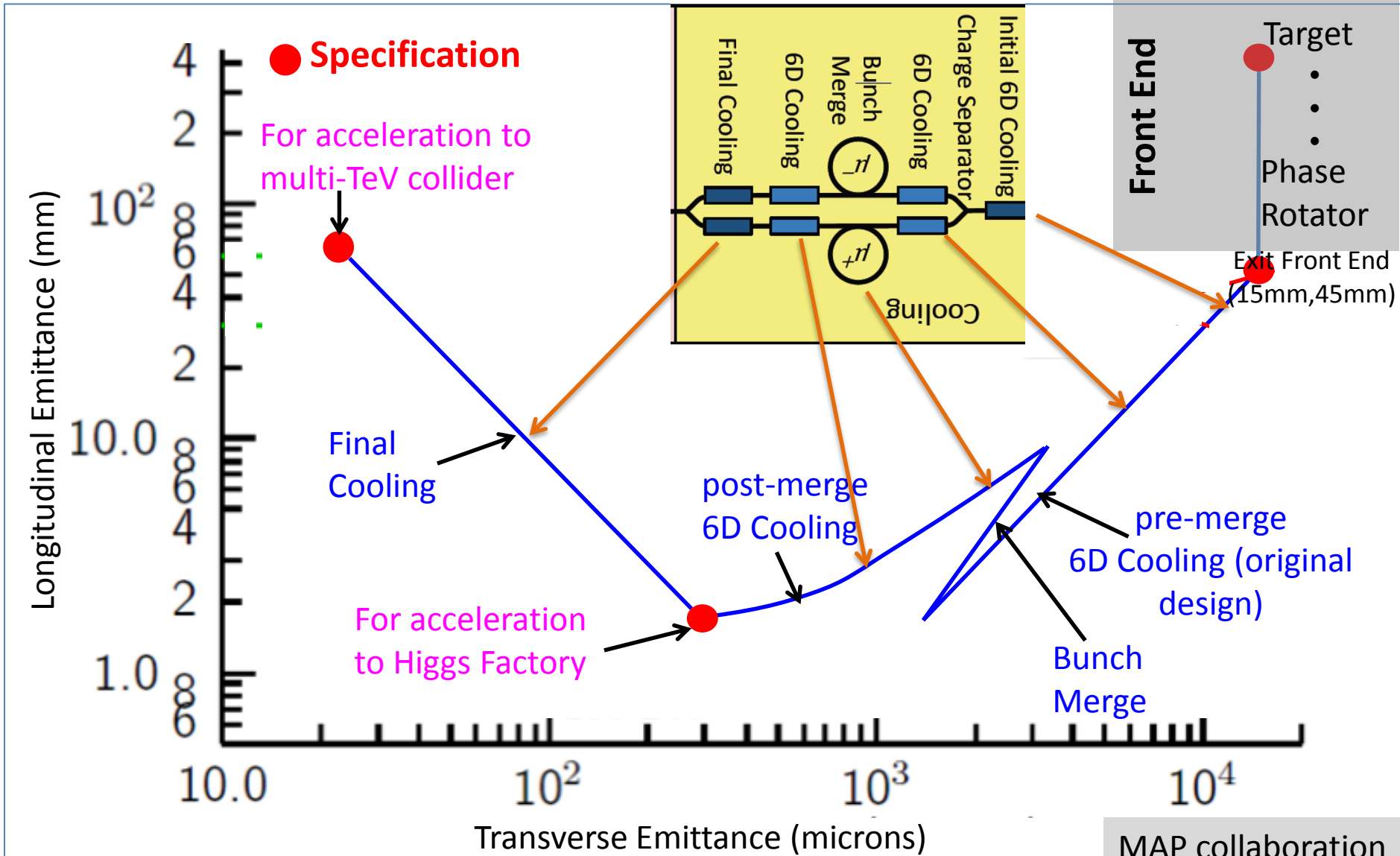
**Compact integration** to minimise muon loss



Minimise betafunctor with **strongest solenoids (40+ T)**  
32 T achieved, 40+ T planned

$$\frac{d\epsilon_{\perp}}{ds} = -\frac{1}{(v/c)^2} \frac{dE}{ds} \frac{\epsilon_{\perp}}{E} + \frac{1}{2} \frac{1}{(v/c)^3} \left( \frac{14 \text{ MeV}}{E} \right)^2 \frac{\beta\gamma}{L_R}$$

# Cooling: The Emittance Path



# Cooling Challenges and Status

Cavities with very high accelerating gradient in strong magnetic field

**Very strong solenoids (> 30 T)** for the final cooling

- simplified: Luminosity is proportional to the field

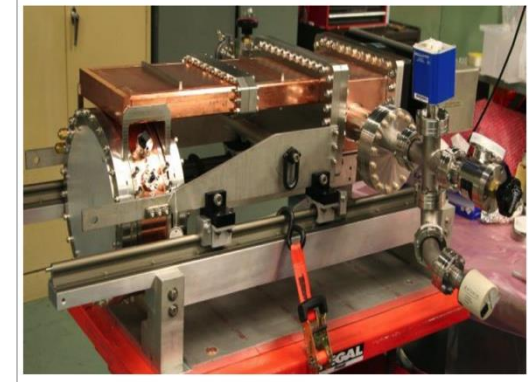
**Integrated system test**

**MuCool: >50**

**MV/m in 5 T field**

Two solutions

- Copper cavities filled with hydrogen
- Be end caps



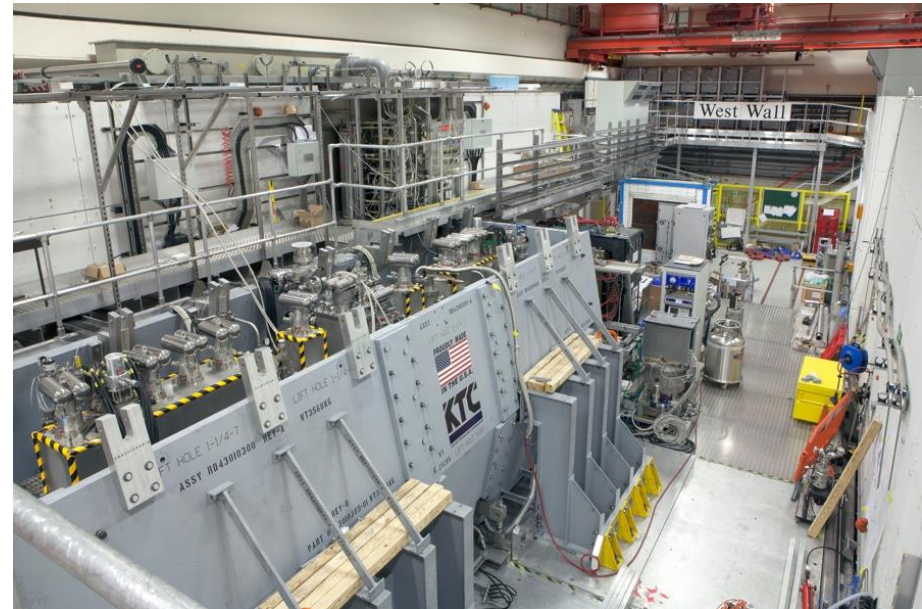
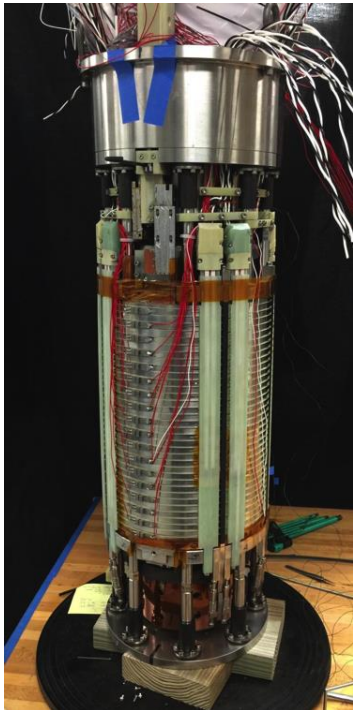
**NHFML**

32 T solenoid with low-temperature HTS

We would like to push even further

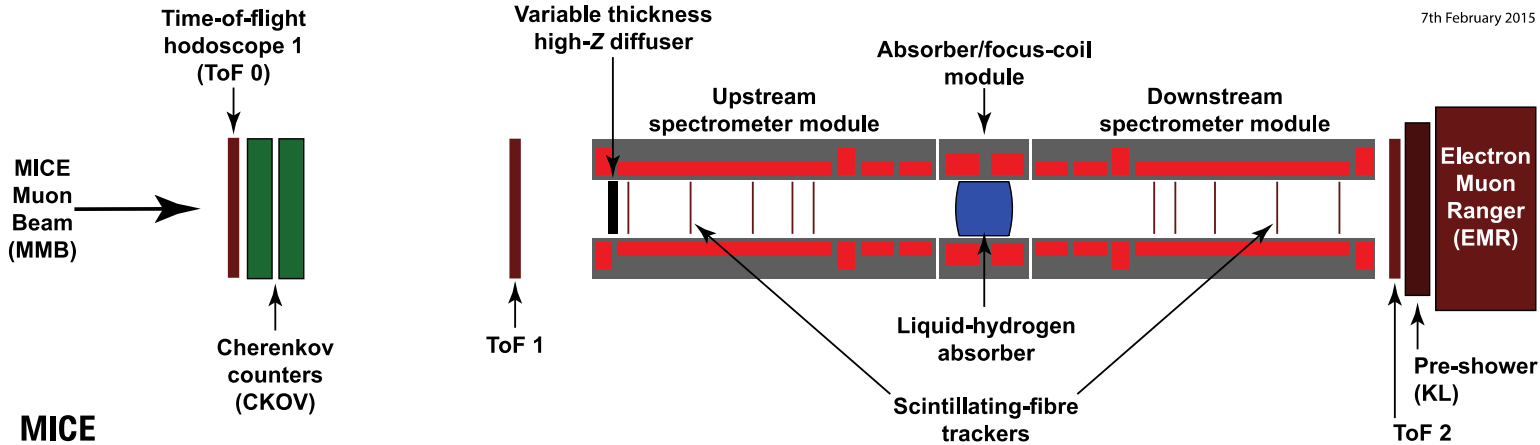
Plans for 40+ T exist

**MICE  
(UK)**



# MICE (in the UK)

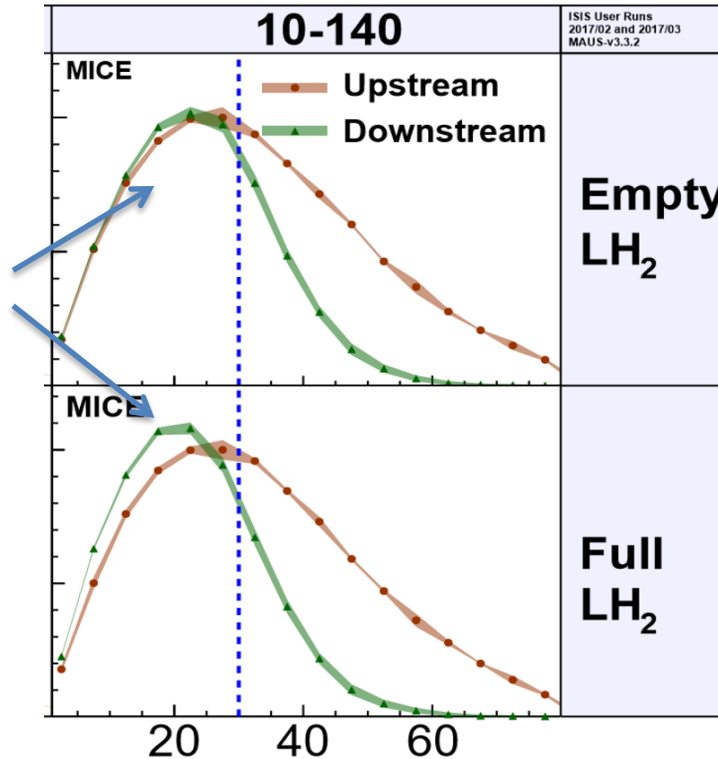
7th February 2015



MICE

More particles at smaller amplitude after absorber is put in place

Principle of ionisation cooling has been demonstrated



Nature volume 578,  
pages 53-59 (2020)

More complete experiment with higher statistics, more than one stage required

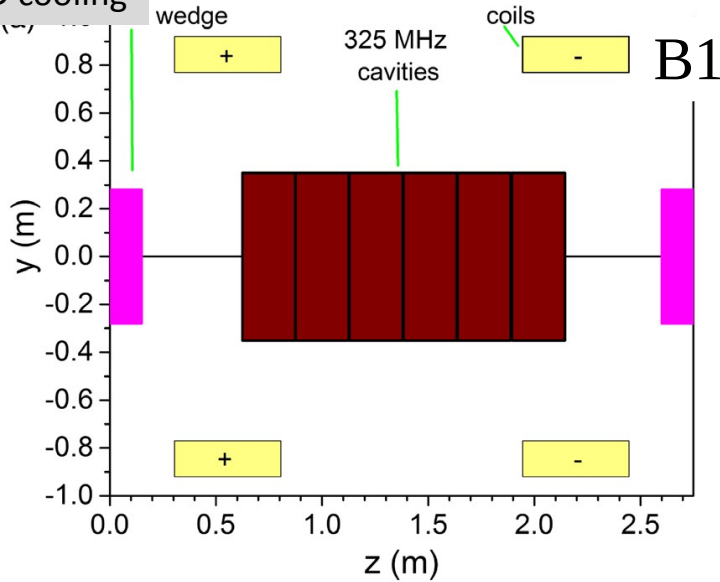
Integration of magnets, RF, absorbers, vacuum is engineering challenge

# Example Cell Designs

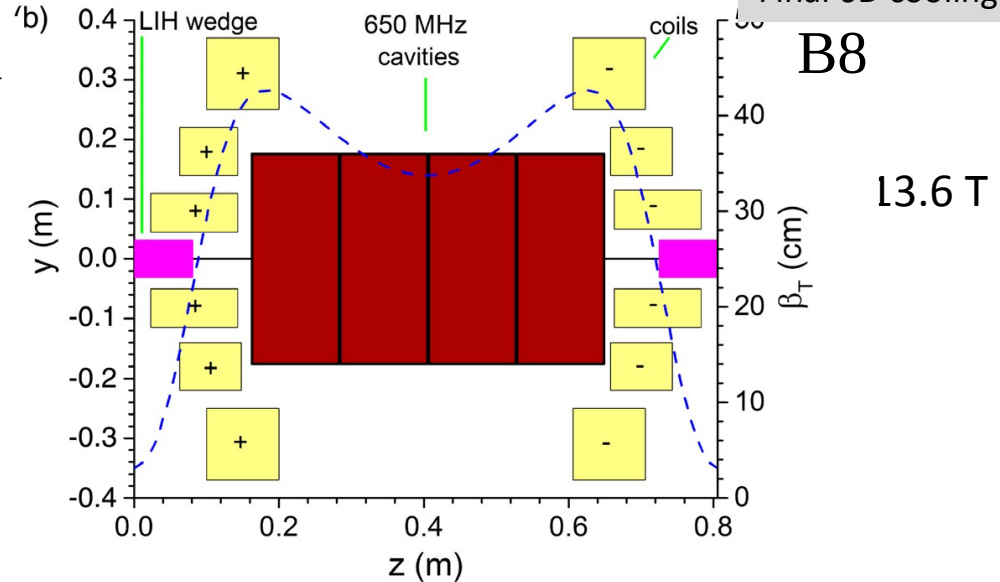
Main 6D-cooling has many magnets and needs **tight integration** with RF and absorbers

Initial 6D cooling

2.2 T



Final 6D cooling



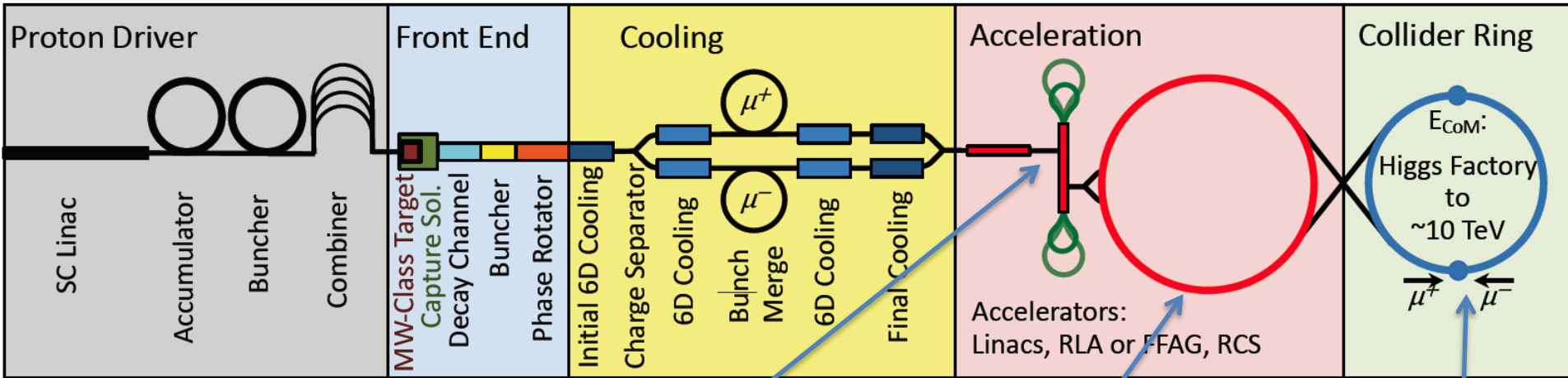
Are already aware of slightly violated space constraints

- maybe cool copper can help both gradient, space and peak power

Alignment has to be integrated (e.g. additional bellows)

Beam operation is important, e.g. beam position on absorber wedge, diagnostics integration, ...

# High-energy Complex



## Initial acceleration

Linacs/recirculating linacs

**Detailed designs from MAP**

Alex Bogacz

## Final acceleration

- FFAG (static superconducting magnets)
- or RCS (rapid cycling synchrotron)

**High-energy designs required**

Start-to-end simulations

**To be started**

Collider ring

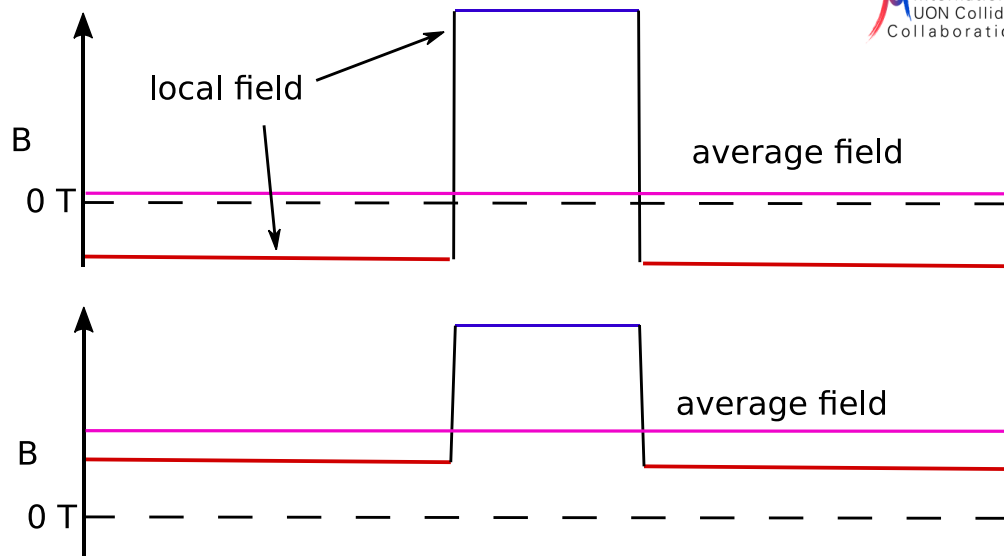
**High-energy designs required**



# High-energy Acceleration

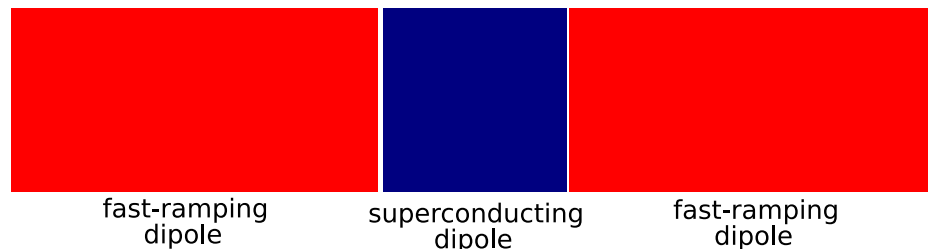
## Rapid cycling synchrotron (RCS)

- Ramp magnets to follow beam energy
- Combine static and ramping magnets
- Possible circumference
  - 14-26.7 km at 3 TeV
  - O(30 km) for 10 and 14 TeV
- Power consumption of fast-ramping systems is important



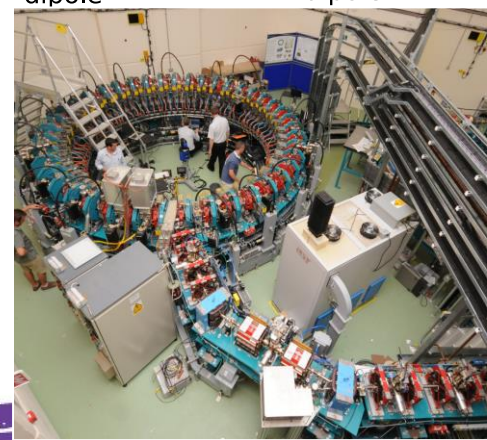
## FFAG

- Fixed (high-field) magnets but large energy acceptance
- Challenging lattice design for large bandwidth and limited cost
- Complex high-field magnets
- Challenging beam dynamics



**EMMA** proof of FFA principle

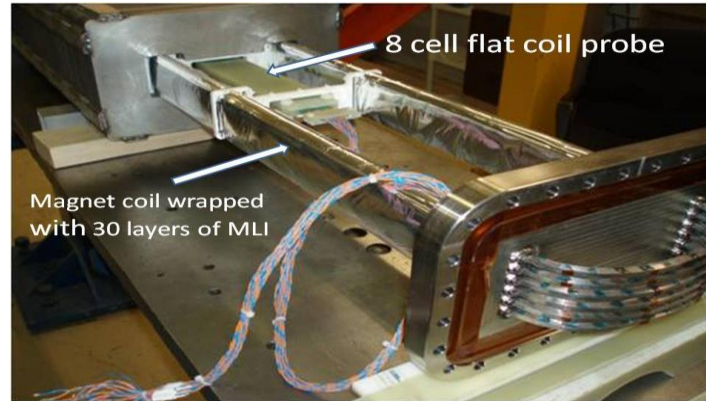
Nature Physics 8,  
243–247 (2012)



# Key RCS Components

⇒ S. Prestemon, D.14.00003

Test of **fast-ramping normal-conducting magnet** design



**Fast-ramping, normal-conducting magnets**  
(5 km of 2 T of per TeV beam energy in hybrid design)  
Design optimisation needed

**Fast, high-field HTS ramping magnets** could benefit 10+ TeV design  
Need O(100) improvement in speed and O(few) in amplitude

Acceleration 0.3 to 1.5 TeV				
Length	km	13.8	26.7	26.7
8 T dipole	km	2.36	2.36	-
$L_{\text{ramp}}$	km	6.3	15.8	18.2
$B_{\text{ramp}}$	T	-2 / 2	-1 / 1	0.34 / 1.7



**FNAL**  
12 T/s HTS  
0.6 T max

Need to push in field and speed

**Power converters** (recovery of energy in ramping magnets, O(200 MJ) at 14 TeV) *Design started*

**RF** (also for FFA):  
Single-bunch beam, high charge (10 x HL-LHC), maintain small longitudinal emittance, high efficiency  
*Design started*

# Final Focus

Pushing beta-functions down at higher energy is key to luminosity

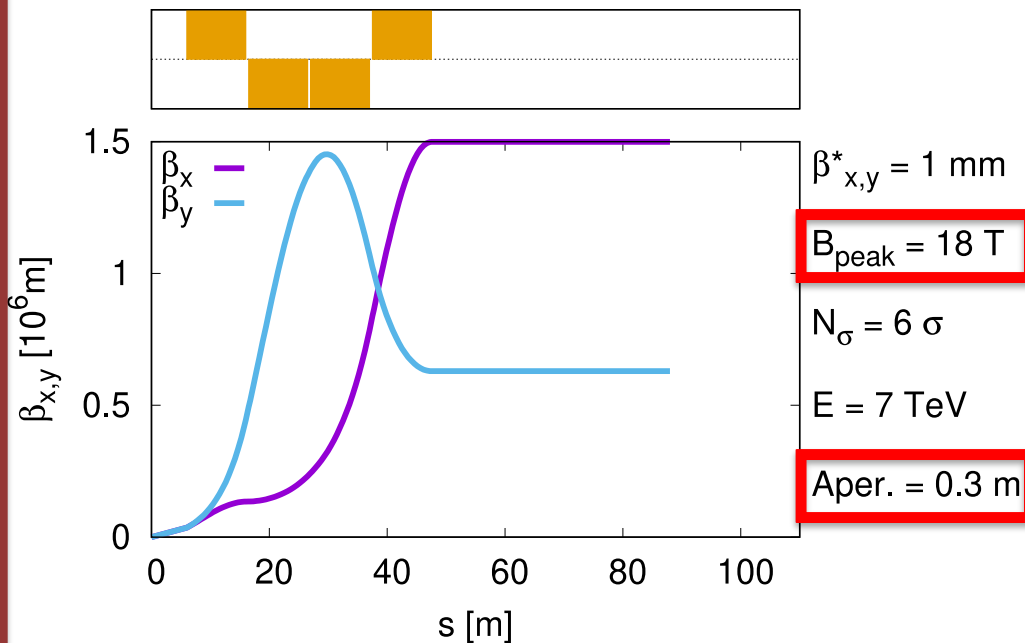
$$b^* \propto \frac{1}{E}$$

Focusing of higher energy beam is more difficult

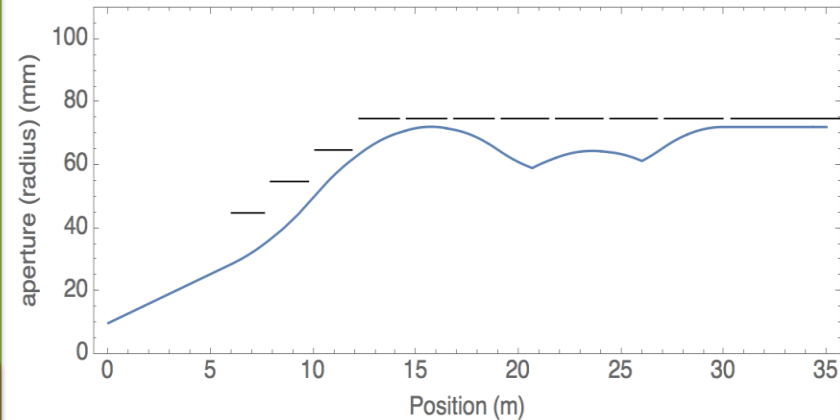
Want to keep triplet short

- rule of thumb: shorter triplet is better for the beam

## First considerations on 14 TeV Design (R: Tomas)



## 3 TeV Design (MAP)



Parameter	Q1	Q1	Q3	Q4
Aperture (mm)	90	110	130	150
Gradients (T/m)	267	218	-154	-133.5
Peak field (T)	12	12	10+	10+
Dipole field (T)	0	0	2.00	2.00

High gradient important at high energy

- HL-LHC level at 3 TeV
- HTS at 14 TeV

What can we hope for from Nb<sub>3</sub>Sn and HTS in 30 years?

# Collider Ring Arcs

## MAP 3 TeV example:

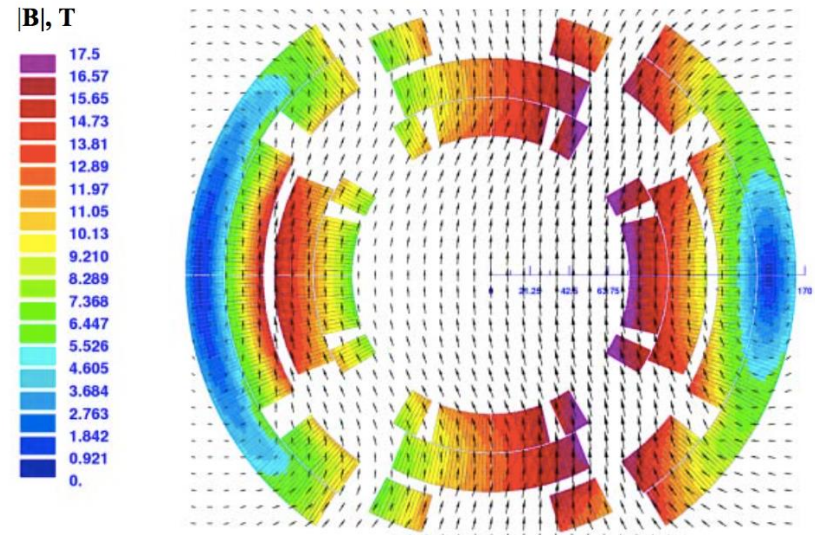
- **10.4 T** 6 m-long dipoles, 150 mm aperture
- and combined function magnets
- 500 W/m losses
- **50/30 mm shielding**
- In cold mass 1.5 mW/g but 10 W/m

## 10 TeV:

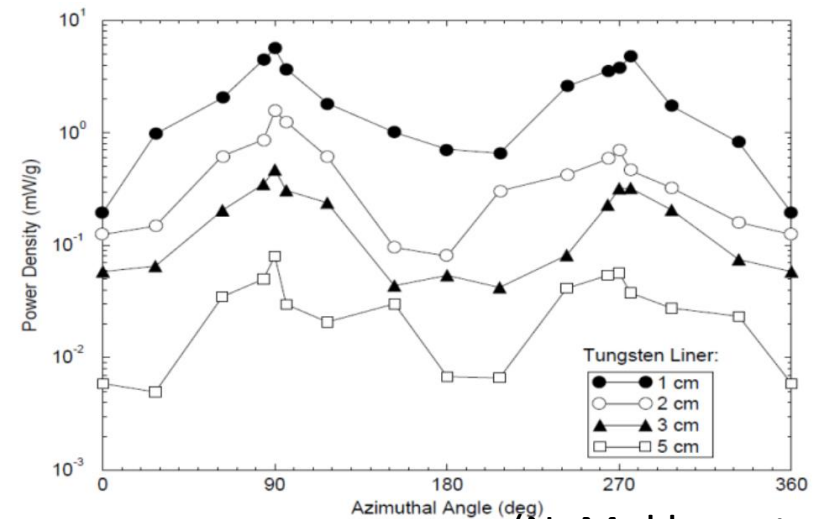
- Expect modest increase of shielding/aperture with beam energy
- Currently no 10 TeV design is developed
- Would go for highest reasonable field (16 T?)

At 3 TeV: Is NbTi worth considering for cost effectiveness?

What can we expect for each technology?



(V.V. Kashikhin et al.)



(N. Mokhov et al.)

# Technology Progress

Important progress on high-field magnets for many projects, HL-LHC, FCC, ...

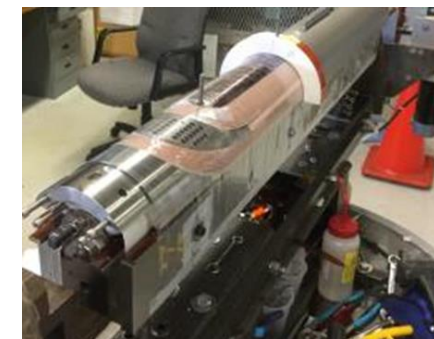
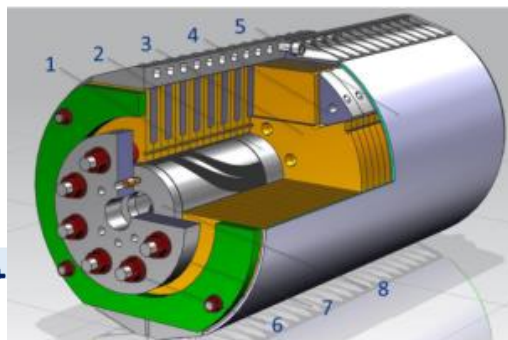
General development of magnets (Nb<sub>3</sub>Sn and HTS) in all regions

Consider more conventional for first stage, more advanced technology for second stage

Development of conductors (FCC)



15 T dipole demonstrator, 60-mm aperture, 4-layer graded coil



Switzerland



Japan



Russia



China



Austria



Korea



Germany



Italy



Finland/USA



Italy



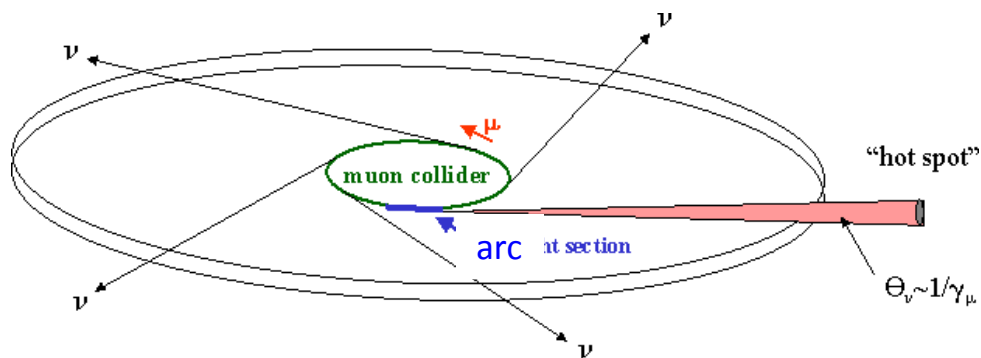
Switzerland



7 companies, two universities and two national research institutes

**Magnet progress is important  
Need to share magnet work for muon collider**

# Neutrino Radiation



Typical legal limit 1 mSv/year

MAP goal < 0.1 mSv/year

No legal procedure < 10  $\mu$ Sv/year

LHC achieved < 5  $\mu$ Sv/year

## Important luminosity limitation

Particularly high in direction of the straights

⇒ buy land in direction of straights

Have to still cover arcs

No mitigation, 500 m deep tunnel:

3 TeV: close to LHC

14 TeV: around legal limit

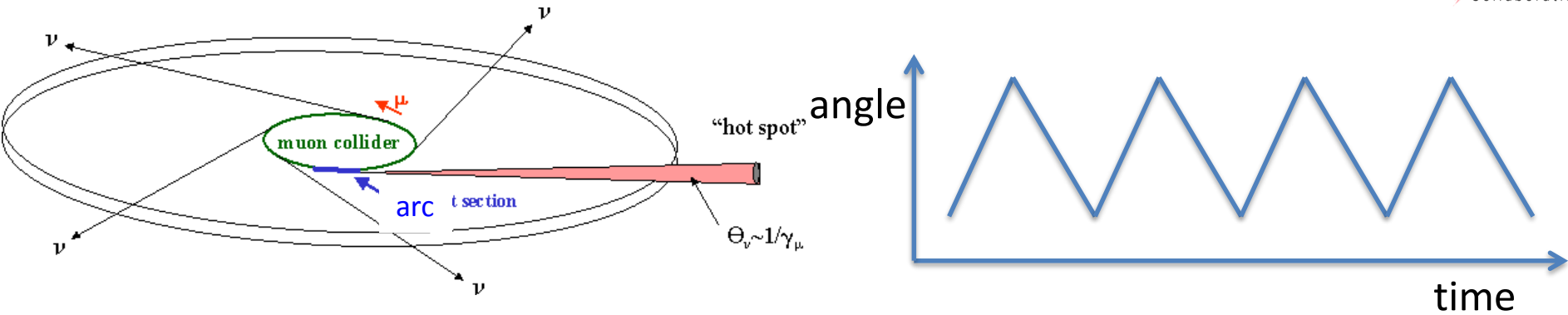
**Needed to find a solution**

Work with **Radiation Protection, Civil Engineering, Geometers** and **Lattice Design** started to find solutions

Mitigate radiation to a level as low as reasonably achievable

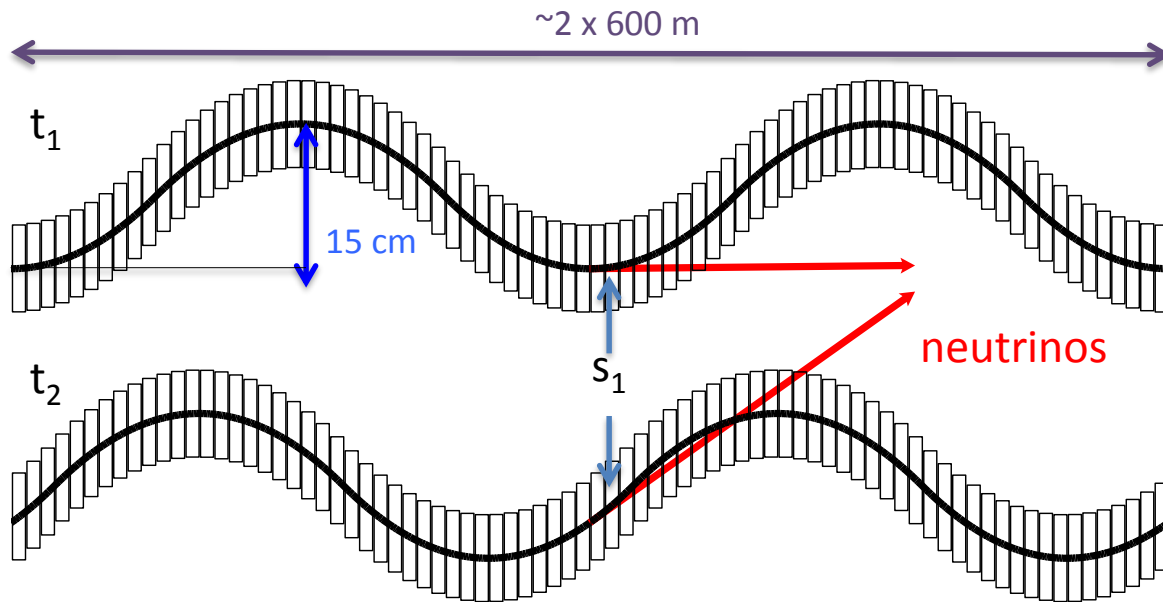
Similar to LHC

# Neutrino Radiation Mitigation Proposal



Mokhov, Ginneken: move beam in collider aperture

Investigating: move collider ring components, e.g. vertical bending with 1% of main field



Opening angle  $\pm 1$  mradian

Even at 14 TeV  
200 m deep tunnel would be  
comparable to LHC case

**Need to study impact on beam  
operation, e.g. dispersion  
control, and components**

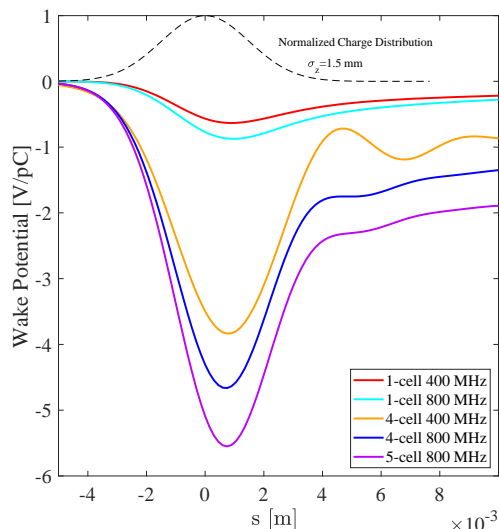
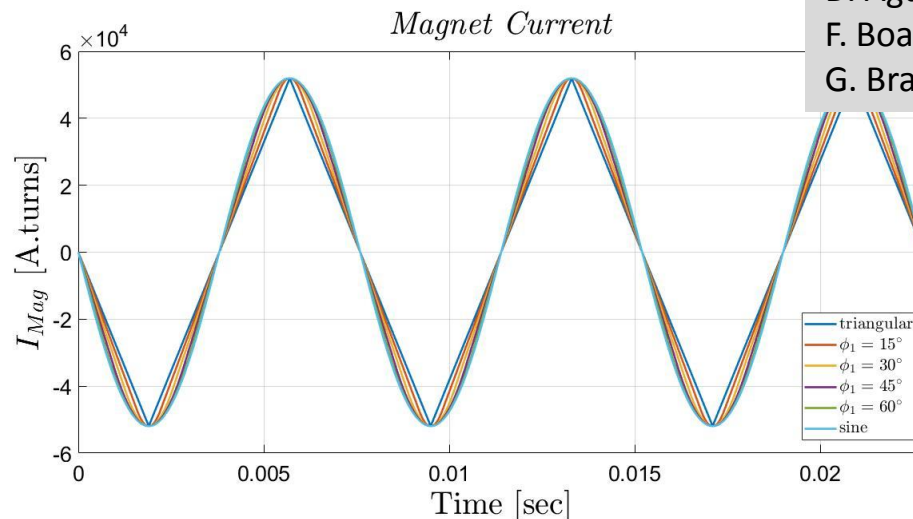
# Selected Recent Progress



D. Aguglia  
F. Boattini  
G. Brauchli

## Ramping magnet challenge

At 14 TeV, energy in field is  $O(200 \text{ MJ})$   
Need to recover it pulse to pulse  
Started to develop **powering scheme**  
with energy recovery



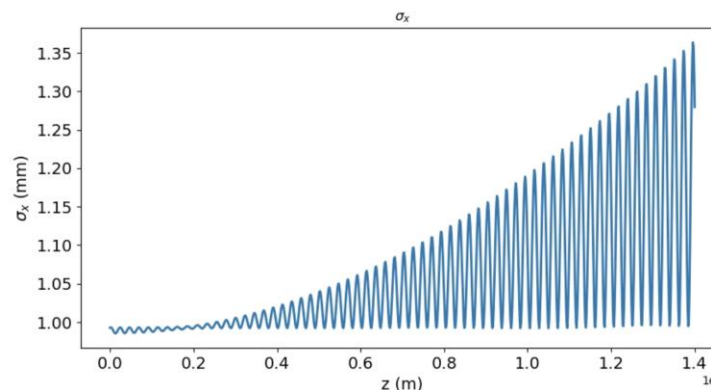
S. Zadeh  
U. van Rienen

## RF challenge (also for FFA):

High efficiency for power consumption  
High-charge ( $10 \times \text{HL-LHC}$ ), short, single-bunch beam  
Maintain small longitudinal emittance  
Studies on cavity wakefields and longitudinal dynamics started

## Collective effects might be a bottleneck

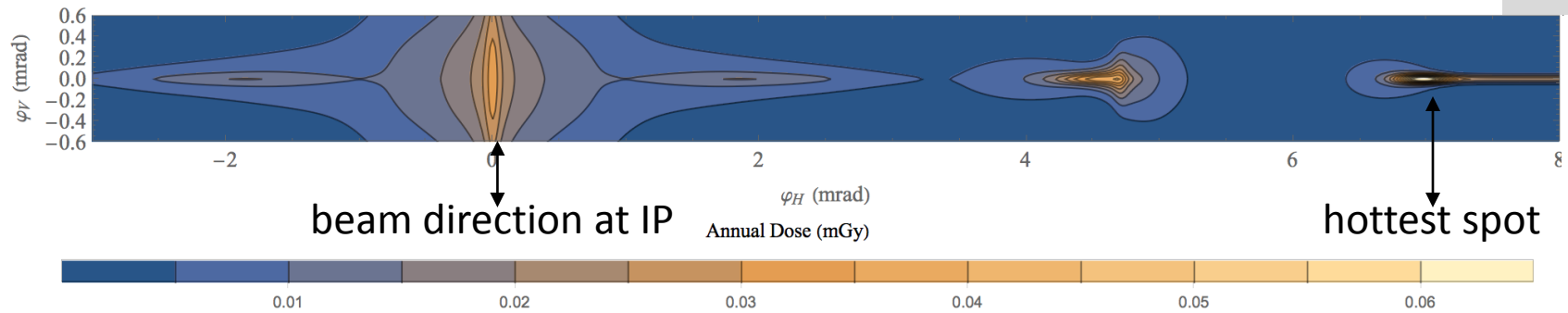
Revisiting for higher energies  
Need to develop tools for collective effects in matter



M. Magliorati  
E. Metral,  
T. Raubenheimer  
D.S.



# Selected Recent Progress, cont.

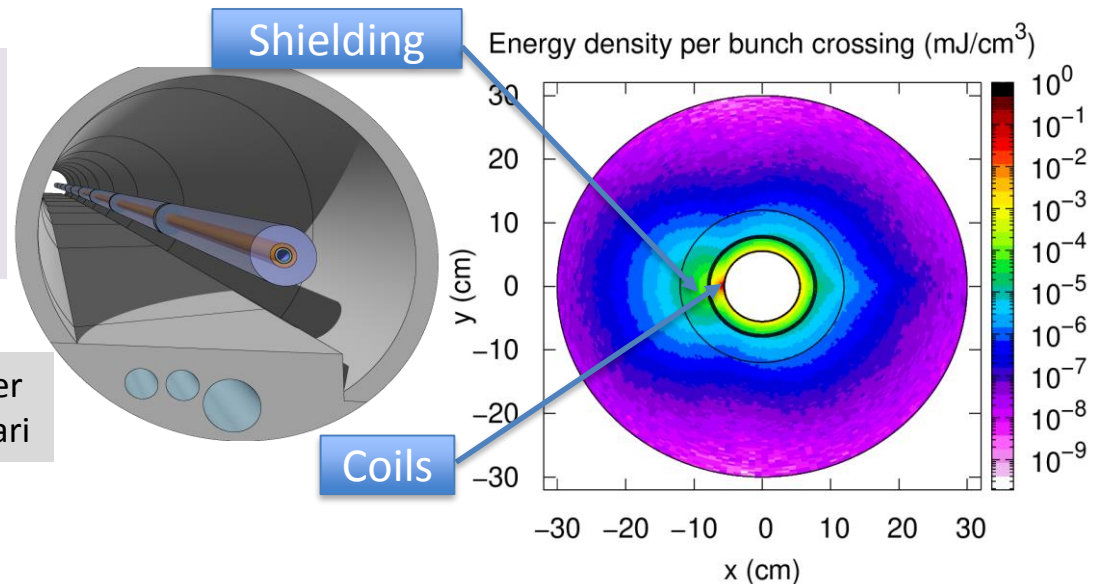


## Collider Ring Lattice Design:

Based on MAP design, lattice design for high energy is starting  
 Started production of **radiation maps** and identified hot spots around IP and in arcs  
 Need to include radiation considerations in lattice design

**Loss challenge** in collider ring:  
 Loss per unit length is constant  
 fewer, but higher energy particles  
 Simulations of shielding started

A. Lechner  
 D. Calzolari



# Tentative Detector Performance Specification

10+ TeV collider enters uncharted territory

Need to establish physics case and detector feasibility

**Established tentative detector performance specifications** in form of DELPHES card (thanks to M. Selvaggi, Werner Riegler, Ulrike Schnoor, A. Sailer, D. Lucchesi, N. Pastrone M. Pierini, F. Maltoni, A. Wulzer et al.), based on FCC-hh and CLIC performances, including masks against beam induced background (BIB)

- For use by physics potential studies
  - Are the performances sufficient or too good?
- For detector studies to work towards
  - make sure technologies are reasonable
  - ensure background is OK
- Please find the card here: <https://muoncollider.web.cern.ch/node/14>

Detector simulation studies/design will now have to verify/ensure that this is realistic considering background and technologies

# Physics Potential, Detector and MDI

## Main background sources

- Muon decay products (40,000 muons/m/crossing at 14 TeV)
  - tertiary muons produced far from collision point
  - showers products produced in final triplets
- Beam-beam background

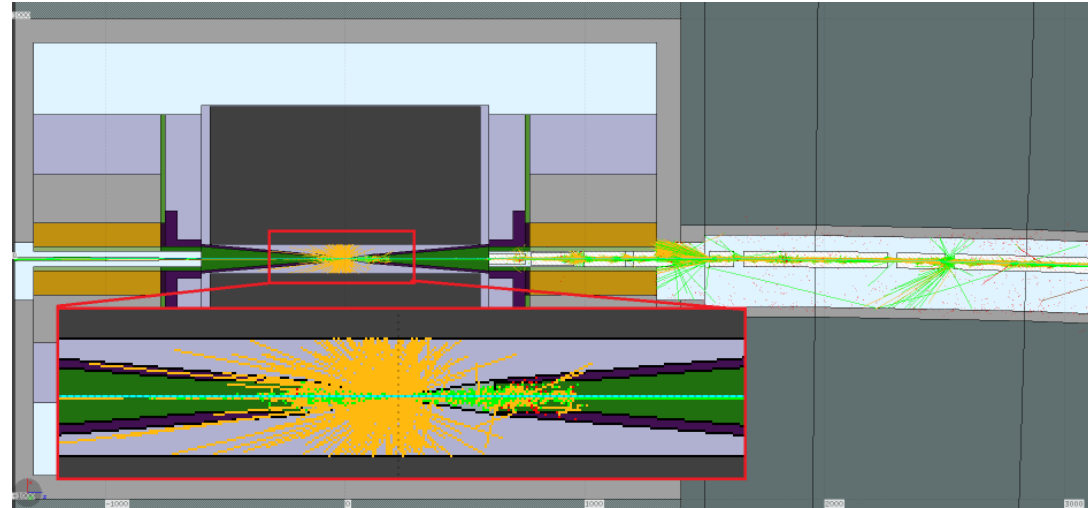
## Background reduces while beam decays

- part luminosity delivered with lower background
- consider worst condition

## Mitigation methods

- masks
- detector granularity
- detector timing
- solenoid field
- event reconstruction strategies
- ...

Need to ensure they do not compromise physics



Simulation tools exist

First studies at lower energies (125 GeV and 1.5 TeV are encouraging)

Will develop systems for higher energies

# Test Programme

## **High-energy complex mostly consists of known components with pushed performances**

- Can be tested as individual prototypes
- Synergies with other developments exist
- Some beam experiments might be useful but could be considered at other accelerators, e.g. control of longitudinal phase space

## **Muon cooling complex is novel and unique to the muon collider**

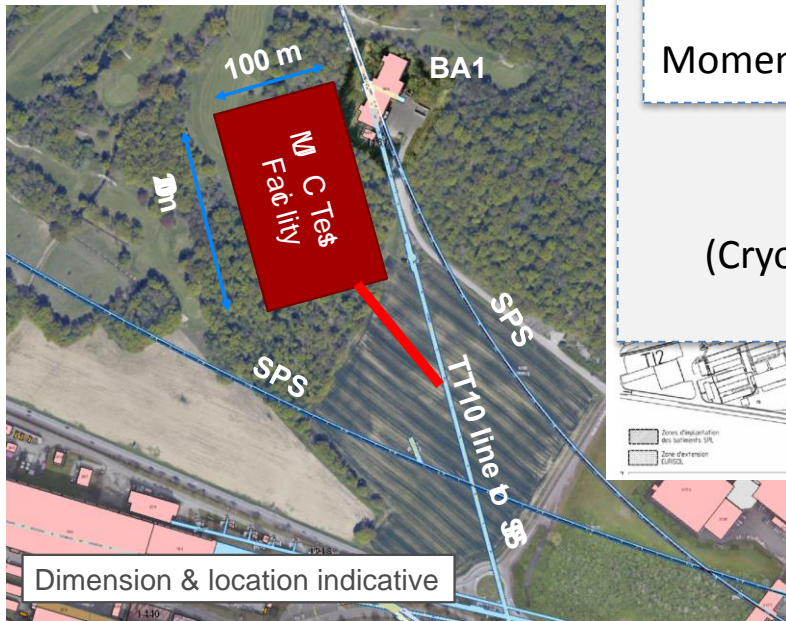
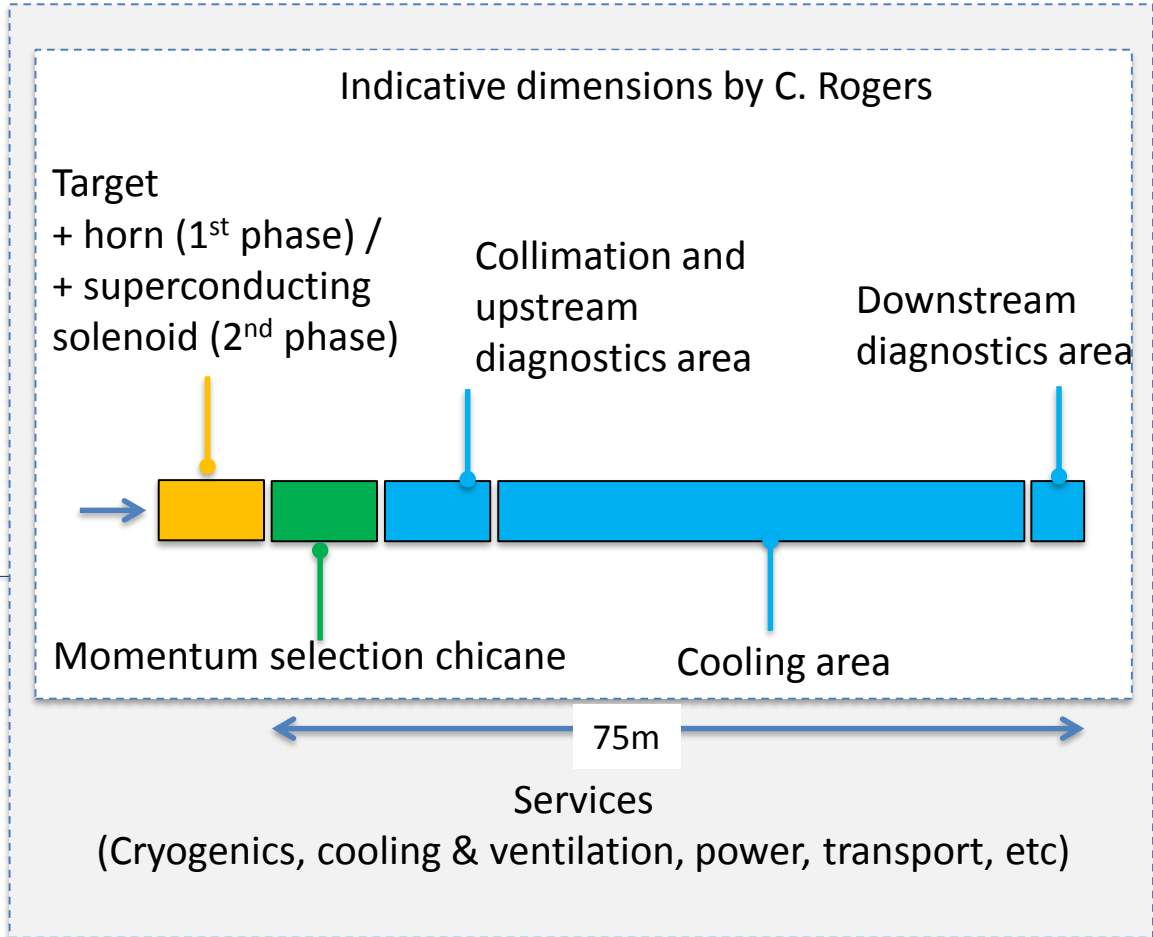
- Many components are unconventional
  - e.g. high-gradient cavities in magnetic field with Be windows or filled with gas
  - massive use of absorbers in the beam path
- Novel technologies beyond MAP design can be considered
  - e.g. very short RF pulse to reduce breakdown probability
  - e.g. use of cooled copper
- Also compact integration is required to maximise muon survival
  - strong superconducting solenoids next to RF at room temperature
  - complex lattice design optimisation
- Almost no experience with beam in these components
  - MICE has been a limited model (no RF, single muons, ...)

⇒ **Need to test a string of cooling cells, ultimately with beam**

# Test Facility Considerations

Test **cooling cell string**, ultimately with beam

Option:  
 CERN land, extract PS beam from TT10 ( $10^{13}$  26 GeV protons in 5 ns, O(10%) of collider, with O(Hz))  
 In molasse (no ground water)



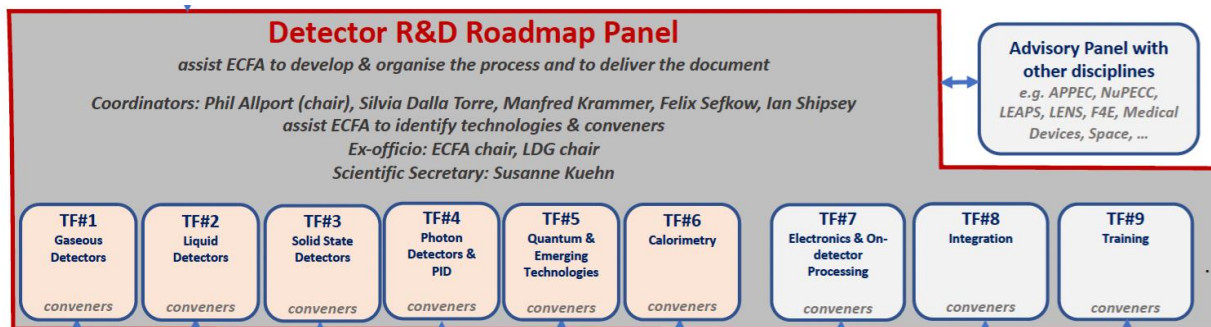
Other options to be explored

Accumulator test at ESS?

# Detector Technologies

Will rely largely on European Detector R&D Roadmap (ECFA)

- Will provide link persons to relevant working groups



Currently consider the following most important (N. Pastrone)

- solid state tracking
- calorimetry
- emerging technologies
- electronics and in detector processing

Will also include other regions

Physics potential studies and machine background studies will verify if performances similar to CLIC and FCC-hh are sufficient

# US Snowmass/P5



Submitted a number of proposals for white papers

- physics potential
- detector
- accelerator

Growing interest in the community

Aiming to coordinate the regional efforts

International Muon Collider Collaboration (corresponding author: D. Schulte)

Muon Collider Facility (c.a.: D. Schulte)

Muon Collider Physics Potential (c.a.: A. Wulzer)

Machine Detector Interface Studies at a Muon Collider (c.a.: D. Lucchesi)

Muon Collider experiment: requirements for new detector R&D and reconstruction tools (c.a.: N. Pastrone)

A Proton-Based Muon Source for a Collider at CERN (c.a.: Chr. Rogers)

Issues and Mitigations for Advanced Muon Ionization Cooling (c.a.: Chr. Rogers)

LEMMA: a positron driven muon source for a muon collider (c.a.: M.E. Biagini)

Applications of Vertical Excursion FFAs(vFFA)and Novel Optics (c.a.: Sh. Machida)

# Timeline

## Initial Design Phase 2021-2025

*Establish whether investment into full CDR and demonstrator is scientifically justified.*

Provide a baseline concept, well-supported performance expectations and assess the associated key risks as well as cost and power consumption drivers.

Identify an R&D path toward the collider, considering High-field Magnet and RF Roadmap results.

## Conceptual Design Phase 2026-

*Develop concept and technology to be ready to commit*

Verify performance of all key components. In particular, build cooling cell string and test with beam. Build and test magnet models and RF components. Start building industrial base for production. Develop site and infrastructure. Determine cost, power, construction schedule. Optimise design.

## Technical Design Phase

*Prepare approval and project implementation*

Prepare industrial production of components, e.g. build magnet prototypes and preseries with industry. Prepare site for construction. Refine cost, power and construction schedule.

## Strategy decision

*Define performance goals and timeline for muon collider*  
*Potentially ramp up of muon collider effort*

## Decision to move to technical design

*Pre-commitment to project*

## Project Approval



# Timeline Discussion

Muon collider is a **long-term direction** toward high-energy, high-luminosity lepton collider

But should also prudently have **back-up scenario** in case Europe does not build higgs factory

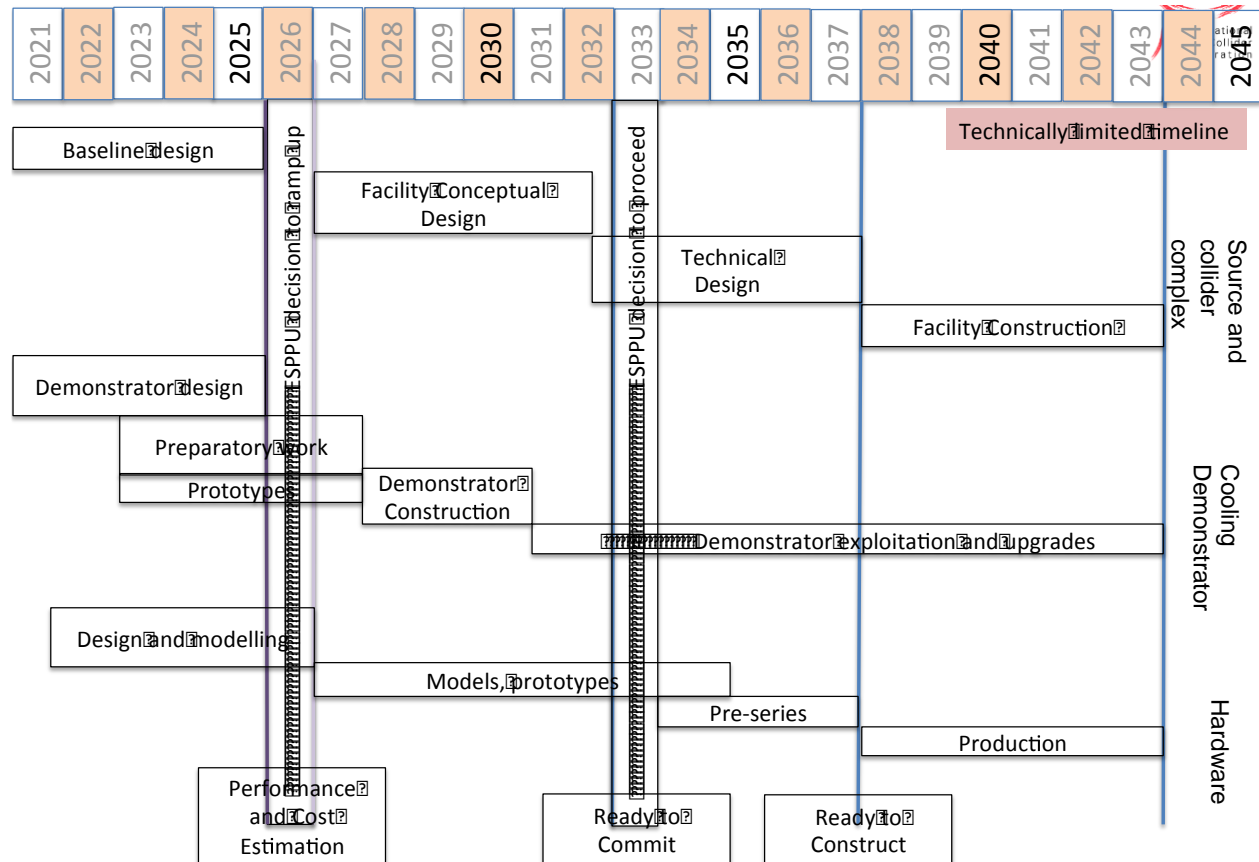
- next large project after HL-LHC

Exploring shortest possible aggressive timeline

- Initial 3 TeV stage on the way to 10+ TeV

Will have more solid evaluation for next ESPPU

Tentative Target for Aggressive Timeline  
to assess when 3 TeV could be realised, assuming massive ramp-up in 2026



# My Impression of Panel Discussions



## **Muon collider has a high potential**

- The muon collider presents enormous potential for fundamental physics research at the energy frontier.
- Not as mature as some other lepton collider options such as ILC and CLIC; but promises attractive cost, power consumption and time scale for the energy frontier, reaching beyond linear colliders.

## **Challenges but no showstoppers**

- The panel identified the key R&D challenges.
- At this stage the panel did not identify any showstopper in the concept.
- Strong support of feasibility from previous studies.
- The panel considers baseline parameter set viable starting point.

## **Panel sees way forward**

- The panel will propose the R&D effort that it considers essential to address these challenges during the next five years to a level that allows estimation of the performance and cost with greater certainty.
- Ongoing developments in underlying technologies will be exploited as they arise in order to ensure the best possible performance.
- This R&D effort will allow the next ESPPU to make fully informed decisions. It will also benefit equivalent strategy processes in other regions.

## **and potential ramp-up**

- Based on these decisions a significant ramp-up of resources could be envisaged, in particular if a fast implementation is deemed essential.

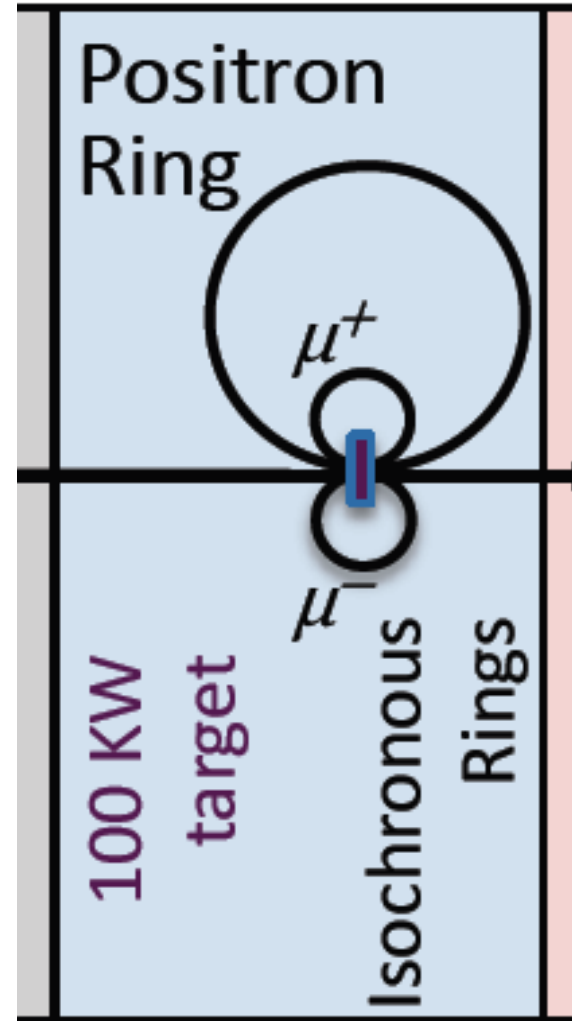
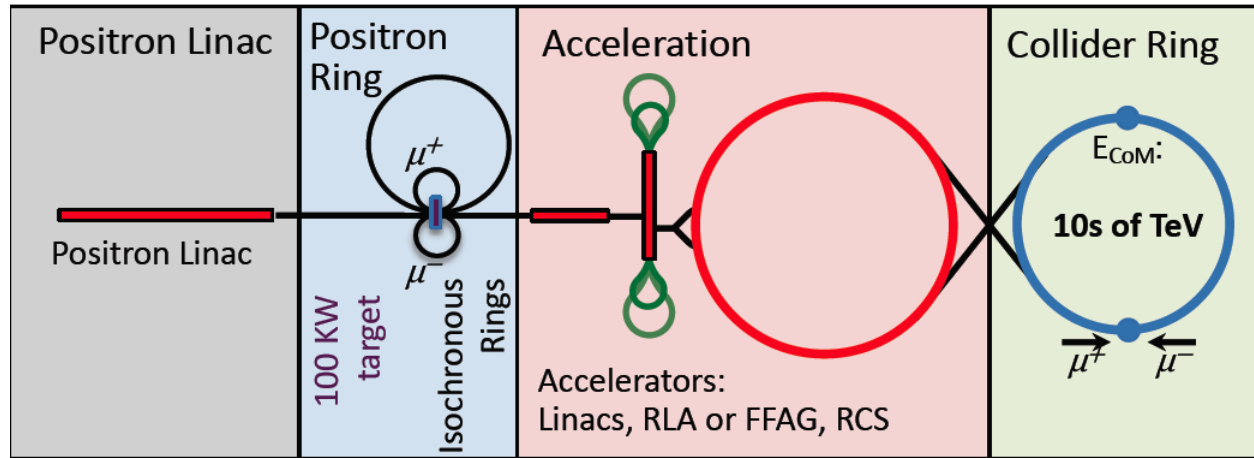
# Conclusion

- Muon colliders are a unique opportunity for a high-energy, high-luminosity lepton collider
  - high luminosity to beam power ratio
  - cost efficiency to be assessed
- Two different options considered
  - 3 TeV collider that can start construction in less than 20 years
  - 10 TeV collider that uses advanced technologies
- Not as mature as ILC or CLIC
  - have to address important R&D items
  - but no showstopper identified
- Need to develop concept to a maturity level that allows to make informed choices by the next ESPPU and other strategy processes
  - Baseline design
  - R&D and demonstration programme
- An important opportunity that we should not miss
- Web page: <http://muoncollider.web.cern.ch>

Many thanks to the Muon Beam Panel, the collaboration, the MAP study, the MICE collaboration, Mark Palmer, Chris Rogers and many others

# Reserve

# Alternative: The LEMMA Scheme



45 GeV positrons to produce muon pairs  
Accumulate muons from several passages

Low-emittance muon beam can reduce radiation

Less mature than proton-driven scheme  
Large positron current required  
Target is challenging  
Large positron production rate [ $O(10^{17}/s)$ ]  
Currently do not reach luminosity goal

# Target Parameter Scaling

Parameter	Unit	3 TeV	10 TeV	14 TeV
L	$10^{34} \text{ cm}^{-2}\text{s}^{-1}$	1.8	20	40
N	$10^{12}$	2.2	1.8	1.8
$f_r$	Hz	5	5	5
$P_{\text{beam}}$	MW	5.3	14.4	20
C	km	4.5	10	14
$\langle B \rangle$	T	7	10.5	10.5
$\epsilon_L$	MeV m	7.5	7.5	7.5
$\sigma_E / E$	%	0.1	0.1	0.1
$\sigma_z$	mm	5	1.5	1.07
$\beta$	mm	5	1.5	1.07
$\epsilon$	$\mu\text{m}$	25	25	25
$\sigma_{x,y}$	$\mu\text{m}$	3.0	0.9	0.63

Scaled from MAP parameters

Emittance is constant

$$\sigma_E \sigma_z = \text{const}$$

Collider ring acceptance is constant

$$\frac{\sigma_E}{E} = \text{const}$$

Bunch length decreases

$$\sigma_z \propto \frac{1}{\gamma}$$

Betafunction decreases

$$\mathcal{L} \propto \gamma \langle B \rangle \sigma_\delta \frac{N_0}{\epsilon \epsilon_L} f_r N_0 \gamma$$