



The CMB-S4 Experiment

Science and Instruments

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for the CMB-S4 Collaboration
29 August 2023



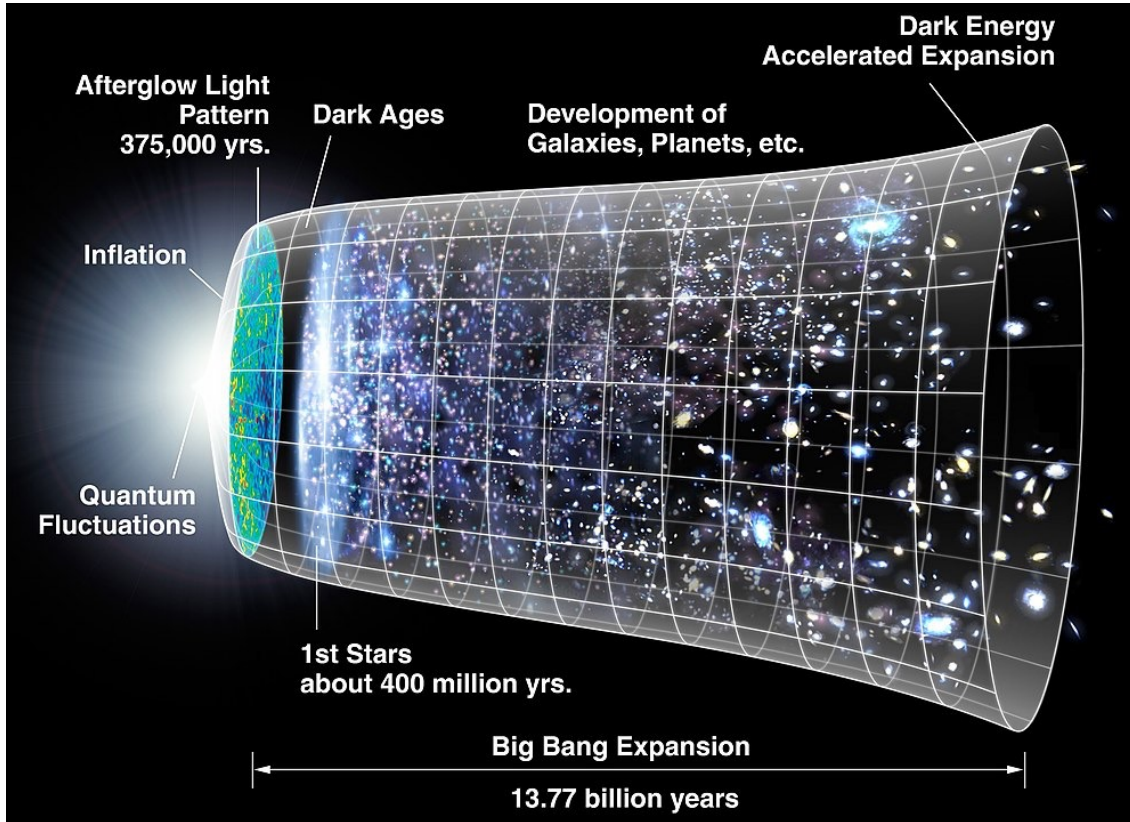
CMB-S4



U.S. DEPARTMENT OF
ENERGY

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Science

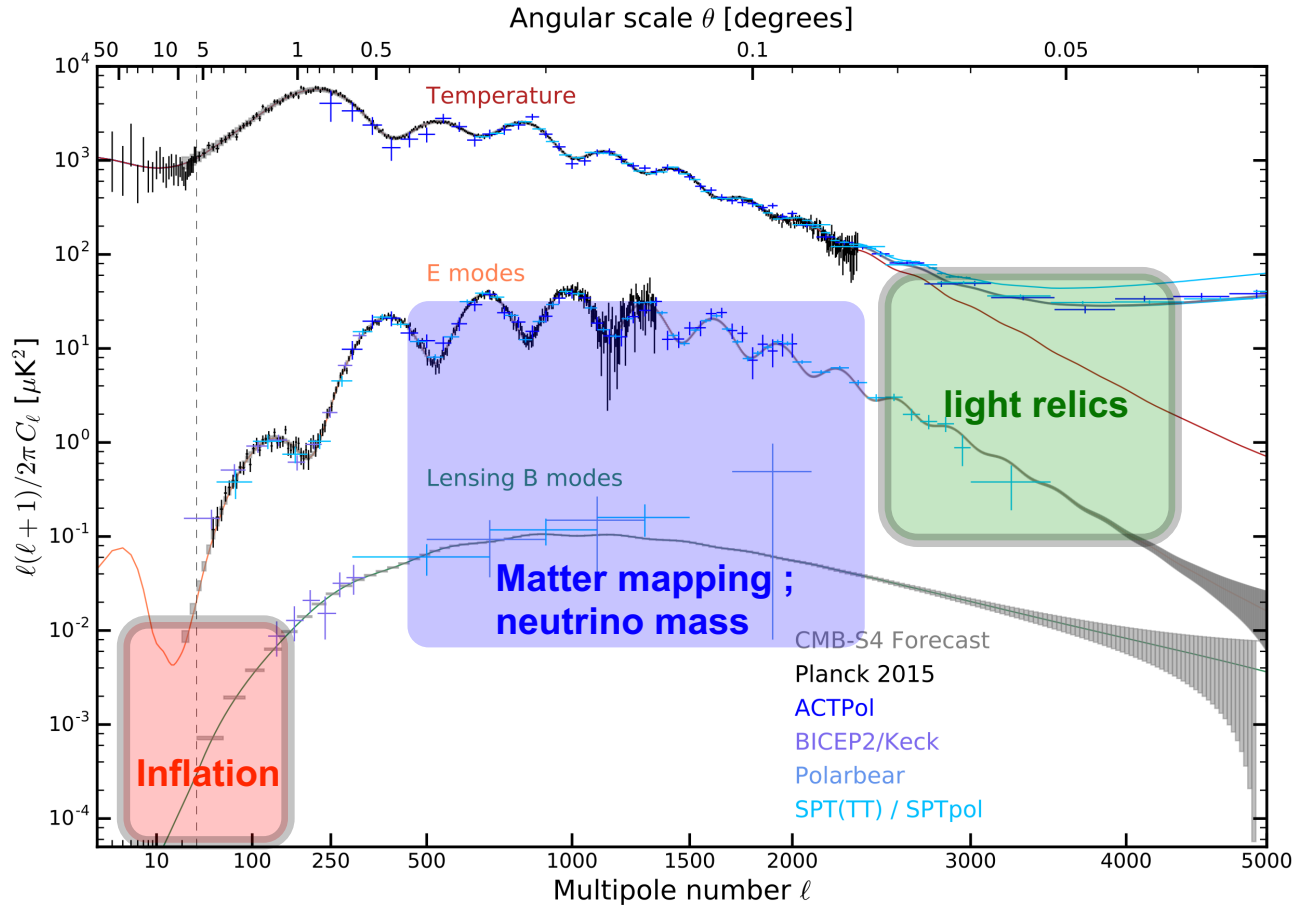
A Unique Probe of Fundamental Physics



The cosmic microwave background (CMB) probes the fundamental physics of the universe in two ways:

- Snapshot of the high energy physics of the early universe
- Backlight through the full formation and evolution of structure

CMB-S4 Science Themes

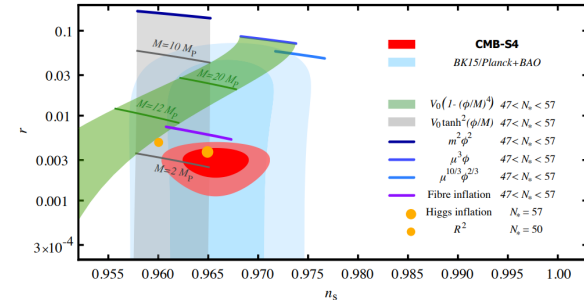
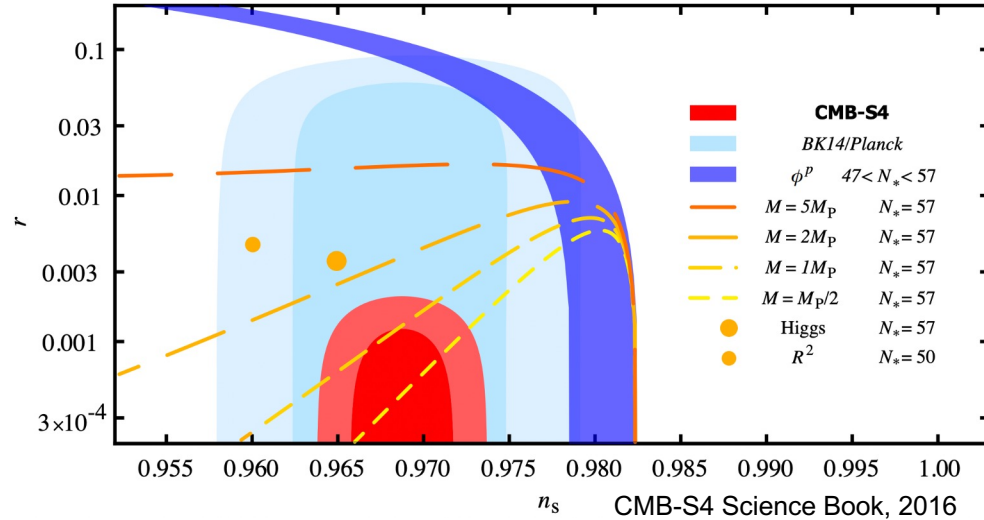


Time-variable millimeter-wave sky

Deepest mm map of our Galaxy – with polarization

Critical Thresholds on Inflation

- Inflation would have left a unique imprint in the polarization of the CMB (**B-modes**)
 - Tensor-to-scalar ratio r gives energy scale of inflation
 - Probe quantum gravity and fundamental physics $\sim 10^{-36}$ s after the universe began, at grand unification theory energy scales (10^{16} GeV)
- CMB-S4 will reach $\sigma(r) < 0.0005$ → Two orders of magnitude lower than current constraints!
- Discover or rule out the most simple and compelling models of inflation



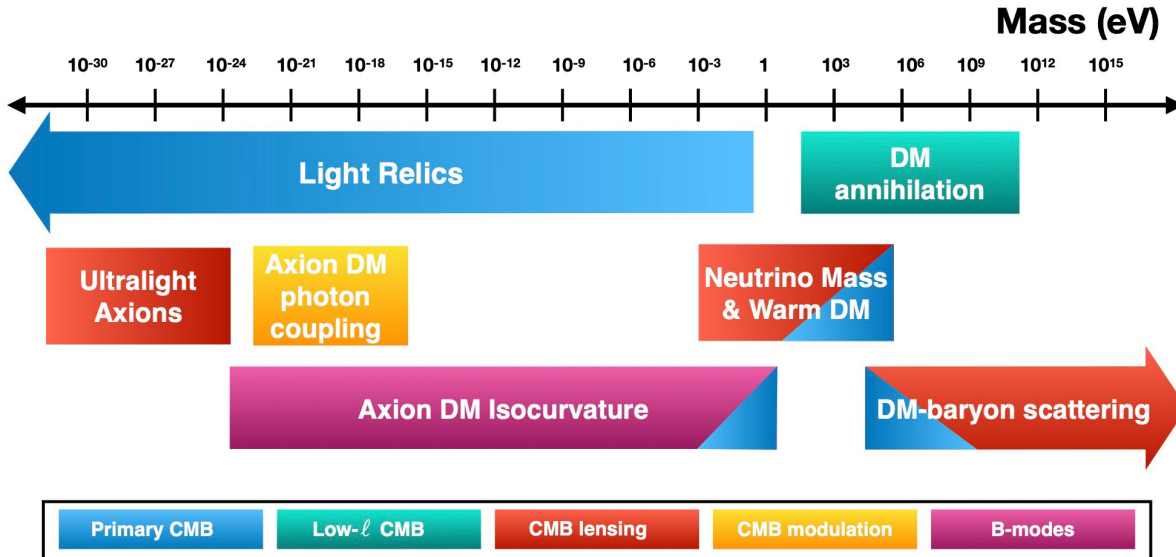
Probing Light Relics

- Any light relics would modify the radiation density \rightarrow CMB power spectra
- CMB gives insights to the dark sector across the mass spectrum

$$\rho_{rad} = \left[1 + \frac{7}{8} \left(\frac{4}{11} \right)^{4/3} N_{eff} \right] \rho_{\gamma}$$

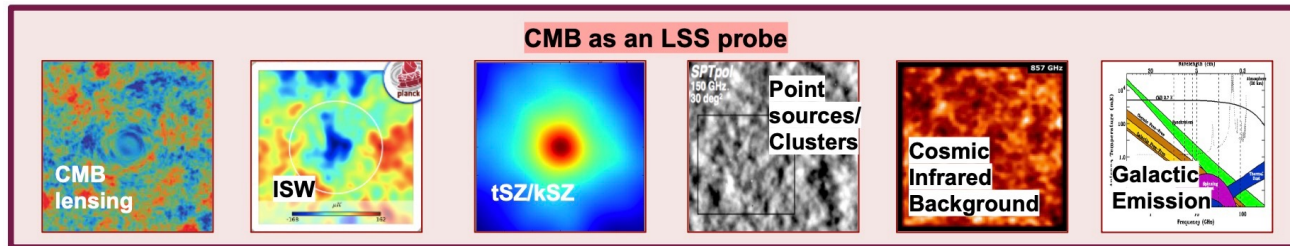
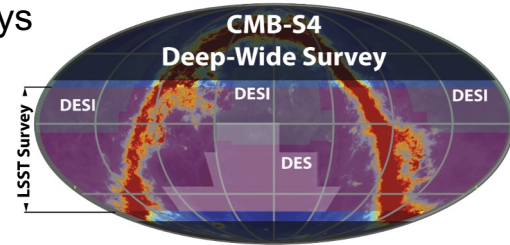
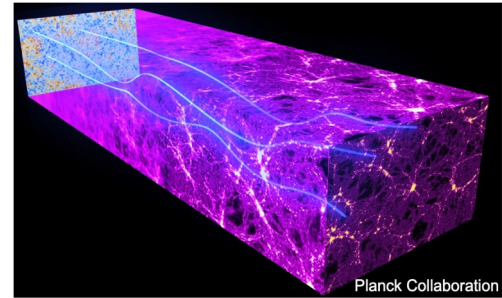
Planck (current): $\sigma(N_{eff}) \sim 0.2$

CMB-S4: $\sigma(N_{eff}) \sim 0.027$



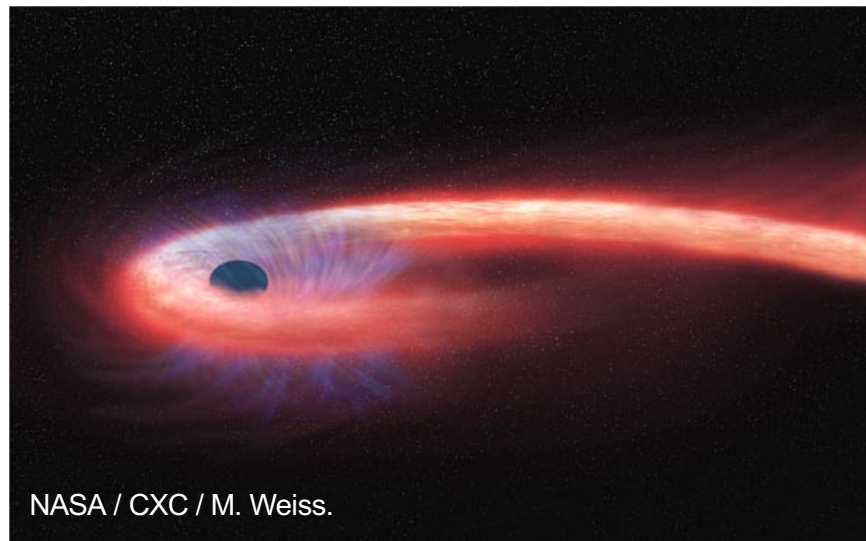
Matter Mapping

- Map matter through **gravitational lensing**, map galaxy clusters via the **Sunyaev-Zel'dovich effect**
- Information about the growth of structure → dark matter, dark energy, sum of the neutrino masses
 - Complementary to neutrino oscillation experiments Δm_ν
 - Highly complementary to supernovae and large-scale structure surveys
 - **Cross-correlations** remove systematics, give crosscheck
 - Baryon pressure, velocity, feedback
- Galaxy clusters to higher z , lower M



The Time-VARIABLE Millimeter-Wave Sky

- First millimeter-wave survey expected to routinely detect transients → population statistics + long-term high cadence monitoring
- Many transient events evolve from low frequency to high frequency in time → CMB-S4 will issue maps of ~hourly observations on few hour timescales
 - Early detection
 - Follow up at other wavelengths → moving into age of multi-messenger astronomy
- Potential for new discoveries!



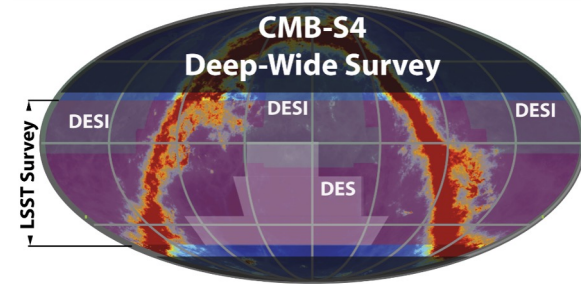
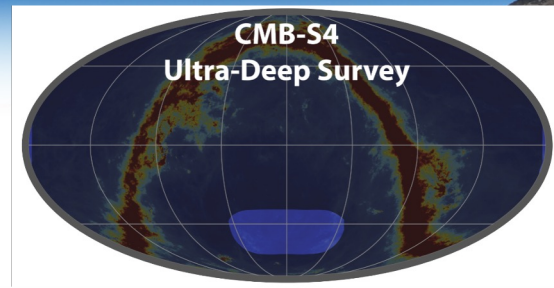
CMB-S4 Design

Ultra-Deep Survey from the South Pole and **Deep-Wide Survey** from Chile with arcmin angular resolution

7 frequency bands (20-300 GHz) for foreground subtraction

Uses mature technology successfully demonstrated in current experiments (ACT, PolarBear, SPT, BK Array, etc.)

Exceeds current generation by **~10x in channel count**



South Pole



Chile

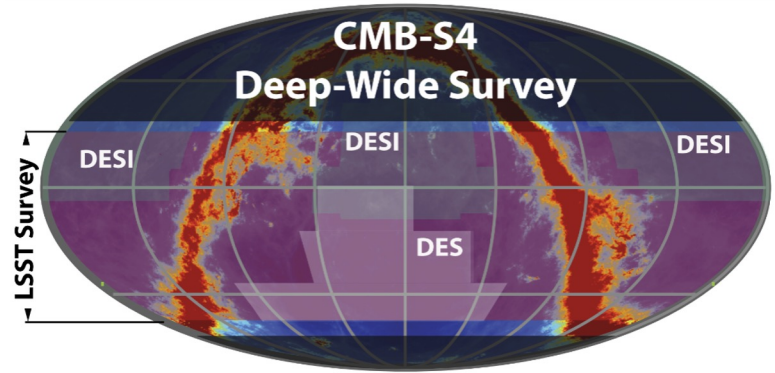


SAT=Small Aperture Telescope ; LAT=Large Aperture Telescope

Two Large Aperture Telescopes in Chile



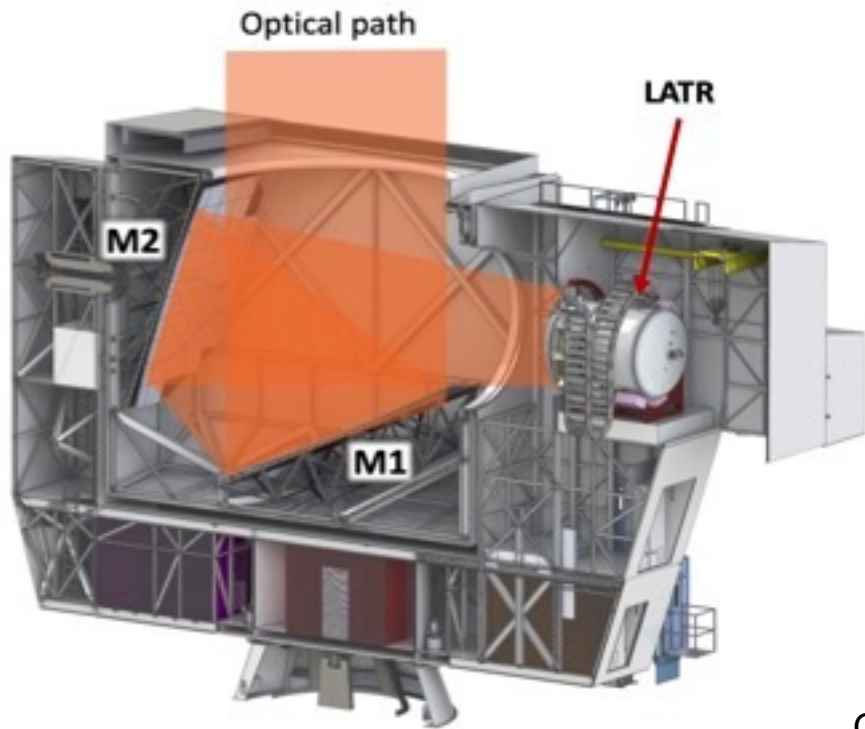
Two 6-m Large Aperture Telescopes (LAT)



Deep and Wide field survey optimized for matter mapping via CMB lensing, neutrino mass, light relics, and transient mm-wave phenomenon.

Overlaps with other optical surveys (DES, DESI, LSST) for cross-correlation

Two Large Aperture Telescopes in Chile



6-meter primary diameter
Crossed-Dragone Design

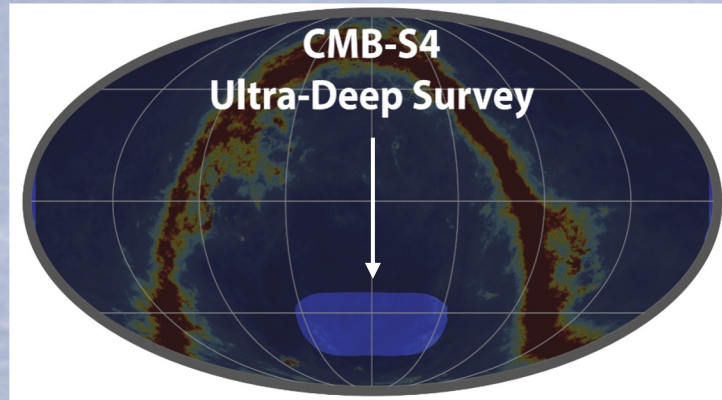
1.4' angular resolution at 150 GHz

Design based on experience from
CCAT-Prime (FYST) and the
Simons Observatory

Gallardo et. al.
<https://arxiv.org/abs/2207.10012>

South Pole Large and Small Area Telescopes

South Pole

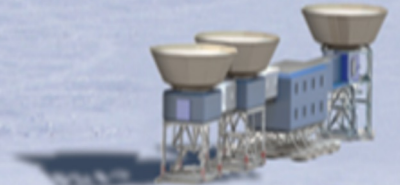


Large Area
Telescope
for B-mode
Delensing

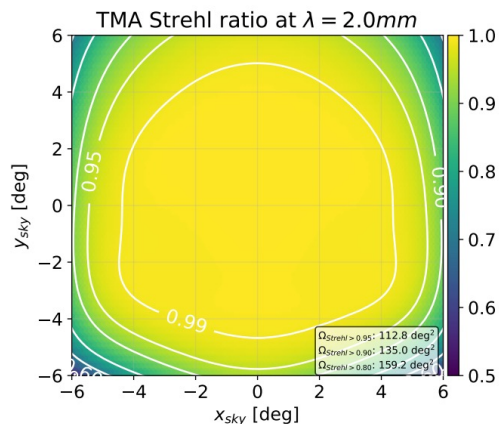
3 Small Area
Telescopes for
B-mode
survey



SPT and Bicep/Keck Array

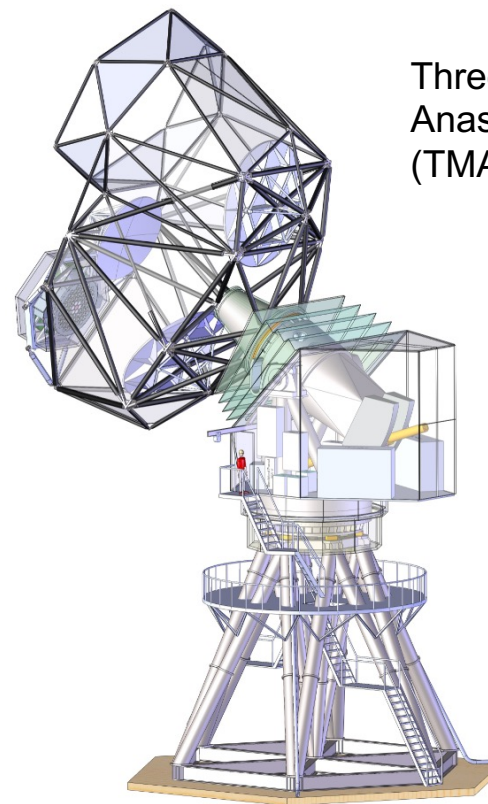


South Pole Large Aperture Telescope

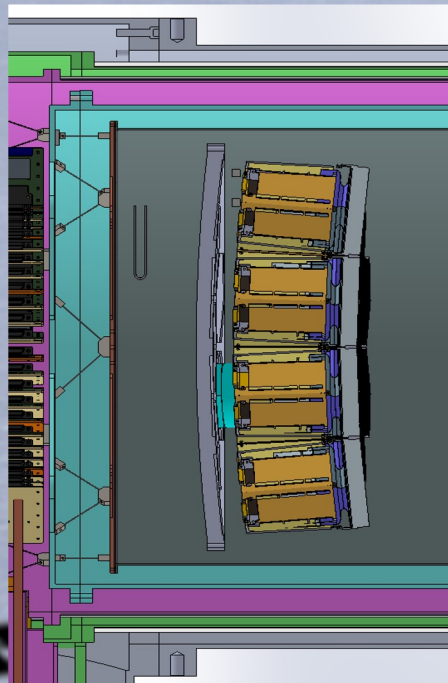


Optimized for the B-mode
Delensing Survey

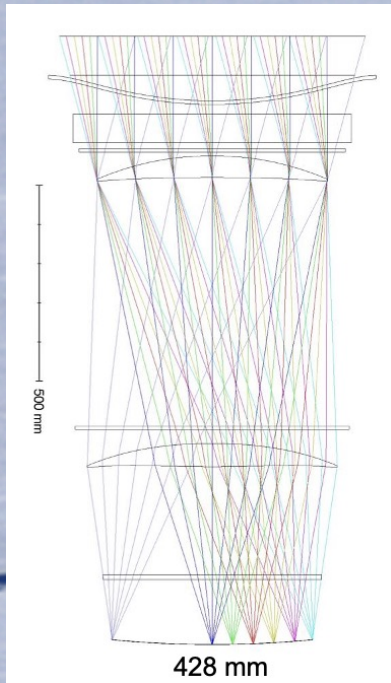
- Optimal beam over FOV
- Gapless mirrors to prevent B-mode contamination
- Bore-sight rotation to cancel polarization systematics



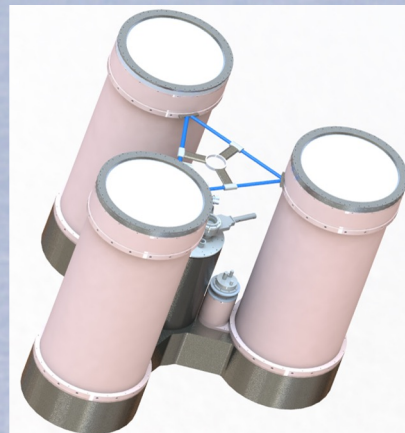
South Pole Small Area Telescopes



Focal Plane



Cold Refractor Optics



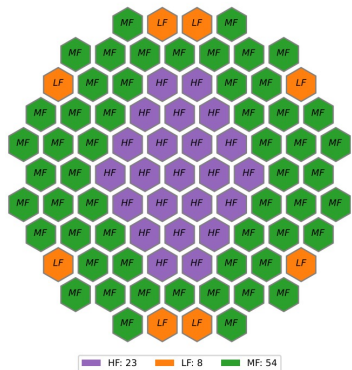
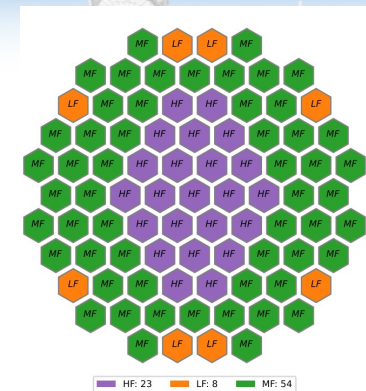
Cryostat



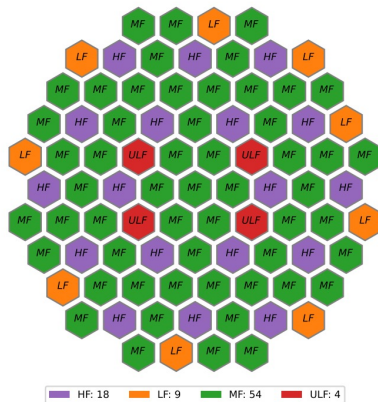
Support Tower and Ground shield

Builds upon proven Bicep/Keck Array Design

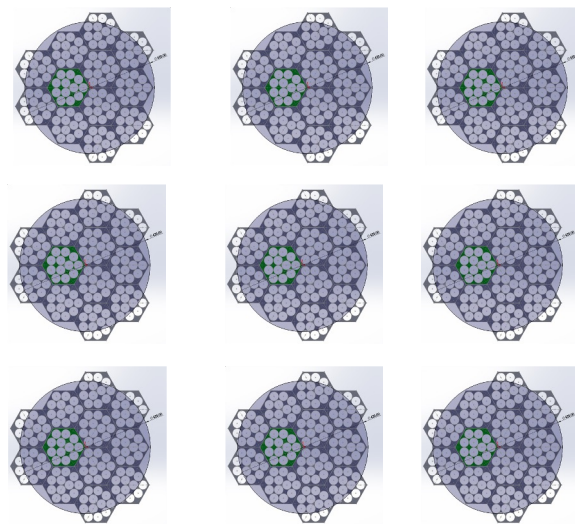
Telescope Focal Plane



Chilean LAT Focal Plane
~ 270,000 detectors



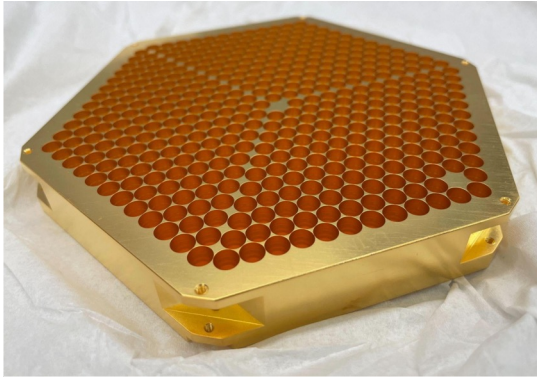
South Pole LAT Focal Plane
~ 130,000 detectors



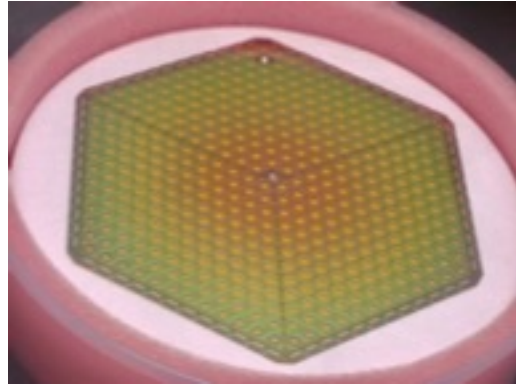
South Pole SAT Focal Plane
~ 100,000 detectors
(low frequency detectors shown for illustration)

7 frequency bands (20-300 GHz)

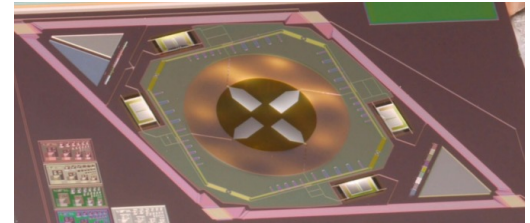
CMB-S4 will be the most sensitive CMB experiment to date



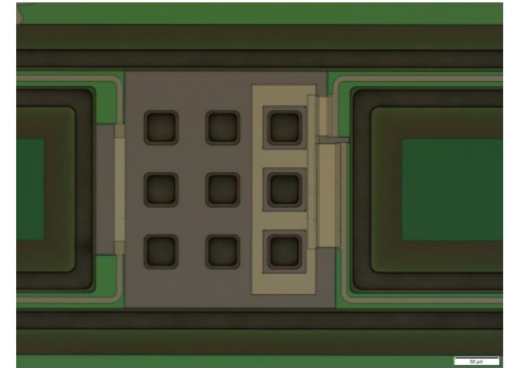
Horn antenna array for best beam optics
(90 GHz design shown for illustration)



100 mK Super conducting
Circuits on 6" Nb Si wafers



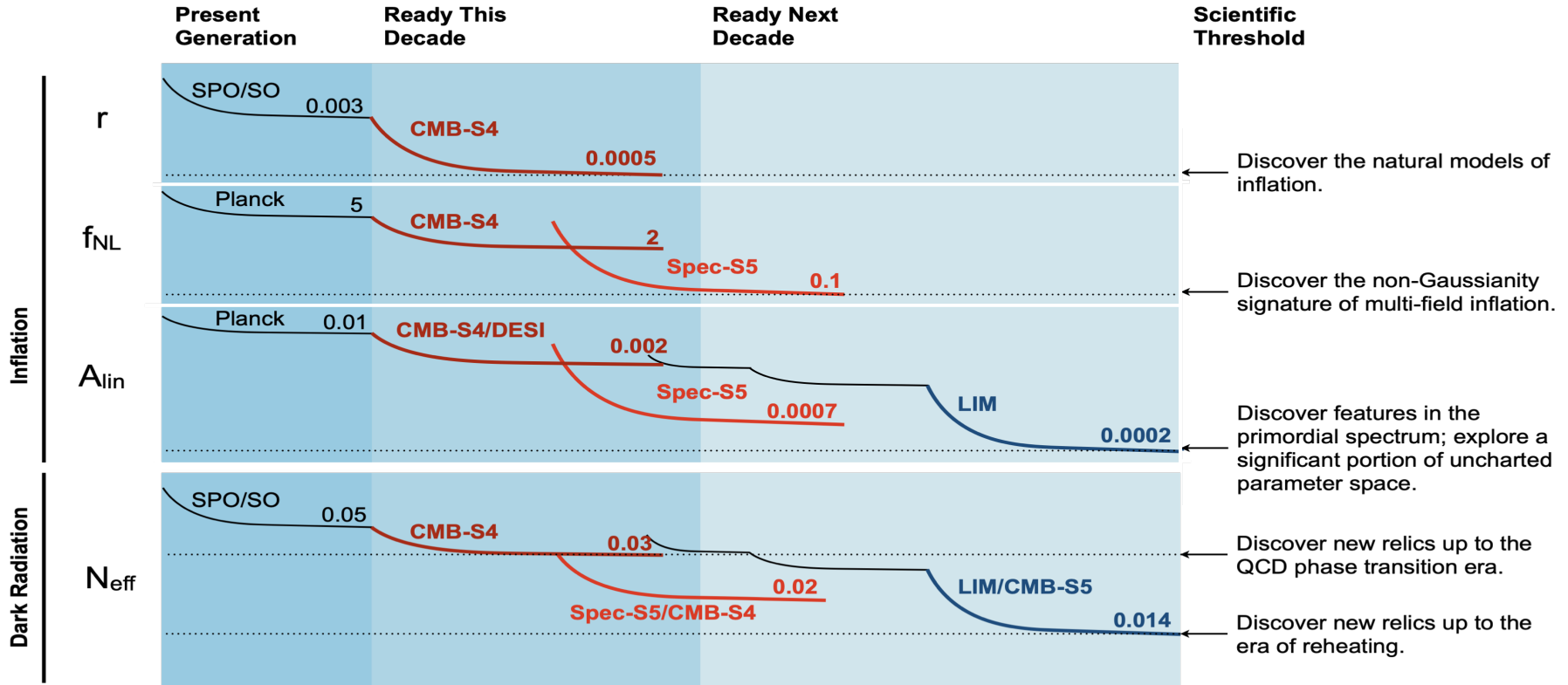
Polarization-sensitive
orthomode transducers



Transition Edge Sensors read out with 80-to-1 multiplexing in the time domain (i.e. sequential readout).

- **Detector Noise practically limited only by photon statistics**
- **Mature technology used by ~all ground-based experiments**

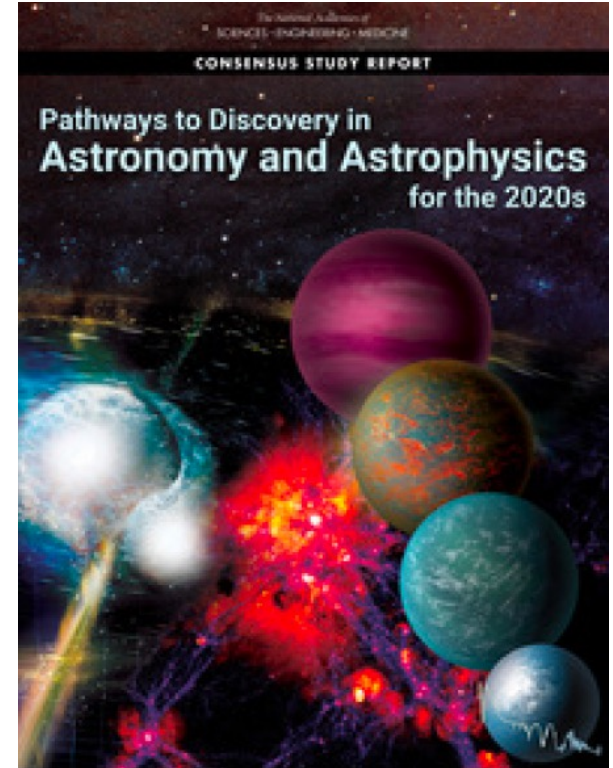
CMB-S4 Science Goals in 7-9 years



US Snowmass Cosmic Frontier 5 report

CMB-S4 Concept Endorsement by US Community

- CMB community has united around “Stage 4” experiment concept, **CMB-S4**:
 - Designed to make transformational discoveries in fundamental physics, cosmology, astrophysics & astronomy
 - Recommended by 2014 P5 under all budget scenarios
 - Recommended by 2015 NAS report A Strategic Vision for NSF Investments in Antarctic and Southern Ocean Research
 - **Strongly recommended by Astro2020: “NSF & DOE should jointly pursue the design and implementation of CMB-S4”**
 - Prominently featured in recent Snowmass process / ongoing 2023 P5: Snowmass Cosmic Frontier report “**Our top project priority is to complete construction of CMB-S4**”.





CMB-S4 has members from 26 US institutions and 19 countries



enthusiasm from
the international
community!

Summary

- The next generation of CMB observations are poised to make tremendous discoveries
 - r : Observe gravity operating on quantum scales
 - N_{eff} : Probe for light particles beyond the standard model
 - $\sum m_\nu$: Constrain the masses of neutrinos
 - New insights into dark energy, dark matter, structure formation
 - New discovery space for transients in the millimeter sky
- CMB-S4 offers a unique opportunity to study fundamental physics of the universe **scaling up mature technology**
- CMB-S4 will impact many fields: HEP, Cosmology, and Traditional Astronomy



CMB-S4 Collaboration meeting August 2023

Thank you!



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