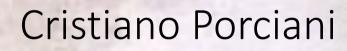
The finger of the observer effect

exploiting the impact of our peculiar velocity on galaxy redshift surveys

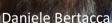


Borzyszkowski et al. 1703.03407 Elkhashab et al. 2108.13424 Elkhashab et al. 2023 in prep

An In fü As

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Mohamed Yousry Elkhashab





PECULIAR VELOCITY OF THE SUN AND THE COSMIC MICROWAVE BACKGROUND

D. W. Sciama*[†] Columbia University, New York, New York, and New York University, New York, New York (Received 17 April 1967) $\begin{bmatrix} 3361.90 \pm 0.04 \text{ (stat.)} \pm 0.36 \text{ (syst.)} \end{bmatrix} \mu \text{K} \implies v_{obs} \approx 370 \text{ km/s}$ $\ell = 263^{\circ}.959 \pm 0^{\circ}.003 \text{ (stat.)} \pm 0^{\circ}.017 \text{ (syst.)} \text{ Delouis et al. (2021)}$ $b = 48^{\circ}.260 \pm 0^{\circ}.001 \text{ (stat.)} \pm 0^{\circ}.007 \text{ (syst.)}$

Fixsen et al. (1994)

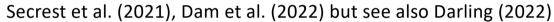
Mon. Not. R. astr. Soc. (1984) 206, 377-381

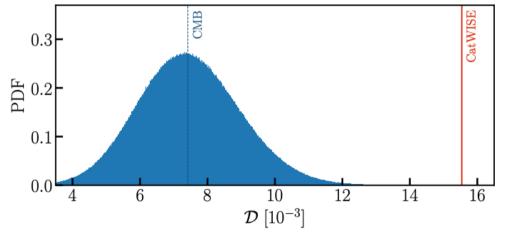
On the expected anisotropy of radio source counts

G. F. R. Ellis* and J. E. Baldwin[†] Orthodox Academy of Crete, Kolymbari, Crete

Systematic errors? Low-redshift contamination? ...

... Bulk flows? Tilted universe? Superhorizon isocurvature perturbations?

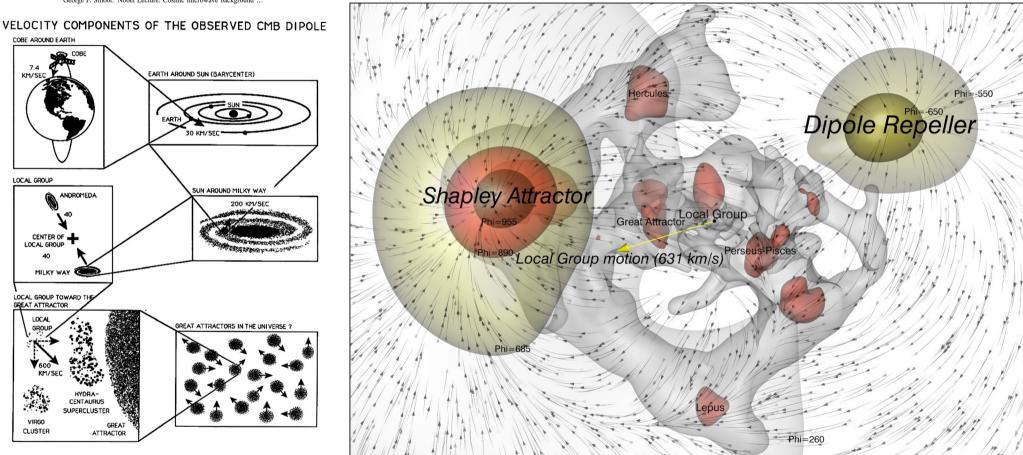




Mass inhomogenities are pulling and pushing us

George F. Smoot: Nobel Lecture: Cosmic microwave background

Hoffman et al. (2017)



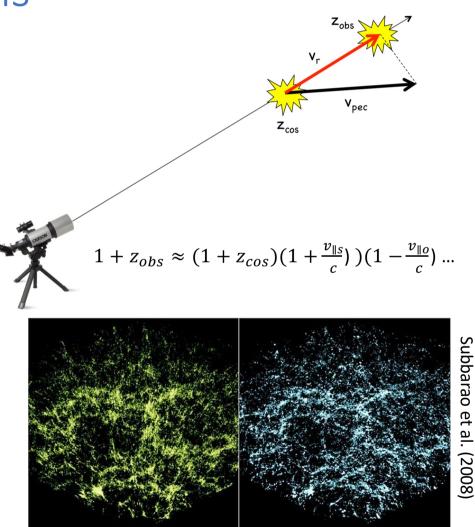
The peculiar velocity generated within 200 Mpc/h points 10° away from the CMB dipole and an external contribution of 160 km/s should arise from sources lying beyond the this volume (Carrick et al. 2015)

Motivation

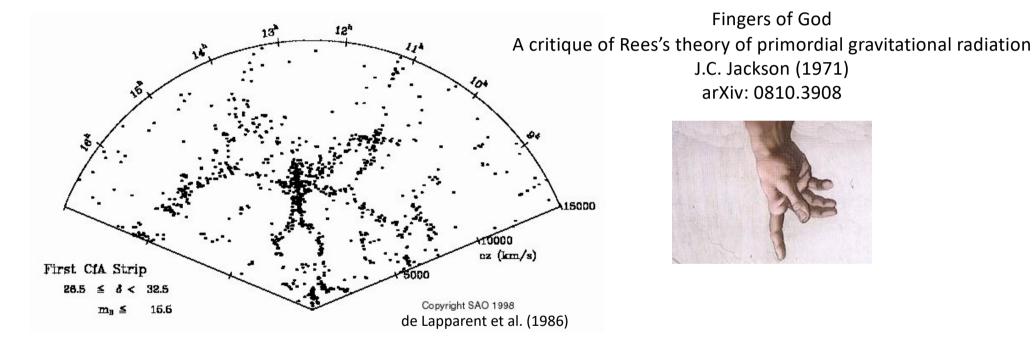
- The current photometric quasar and radio-galaxy samples have limitations
- The quest for dark-energy is driving unprecedentedly large galaxy surveys with accurate redshift measurements $(\frac{\Delta z}{1+z} \ll 10^{-3})$
- Can we measure \mathbf{v}_{obs} from summary statistics of galaxy clustering?
- And, in general, what is the effect of \mathbf{v}_{obs} on our observables?

Redshift-space distortions

- Redshift catalogs list angular positions and redshifts for a selected sample of galaxies (often flux limited)
- We use the observed redshift as a distance indicator for galaxies
- The conversion from redshift to comoving radial distance is done using an unperturbed FLRW model
- However, z_{obs} differs from z_{cos} due to perturbations
- The reconstructed position of a galaxy is shifted with respect to the actual one

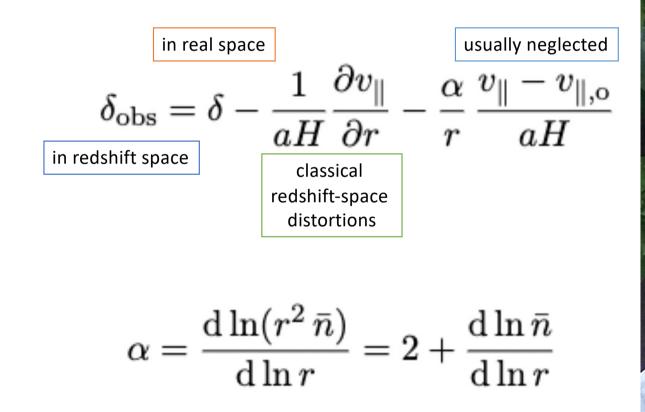


Finger-of-God effect



"The galaxies appear to fall into long chains or cigar-shaped configurations, all pointing at the Earth. Unless one is prepared to assign to the Earth a very special place in the Universe, one must conclude that D is not a good distance indicator, and that in reality the galaxies exist in roughly spherical configurations whose internal velocity dispersions are several times that which would be observed if these systems were expanding with the Universe"

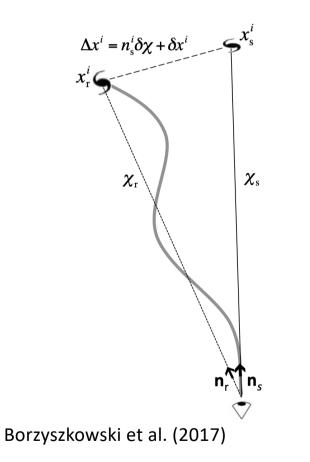
Nick Kaiser's 1987 seminal paper





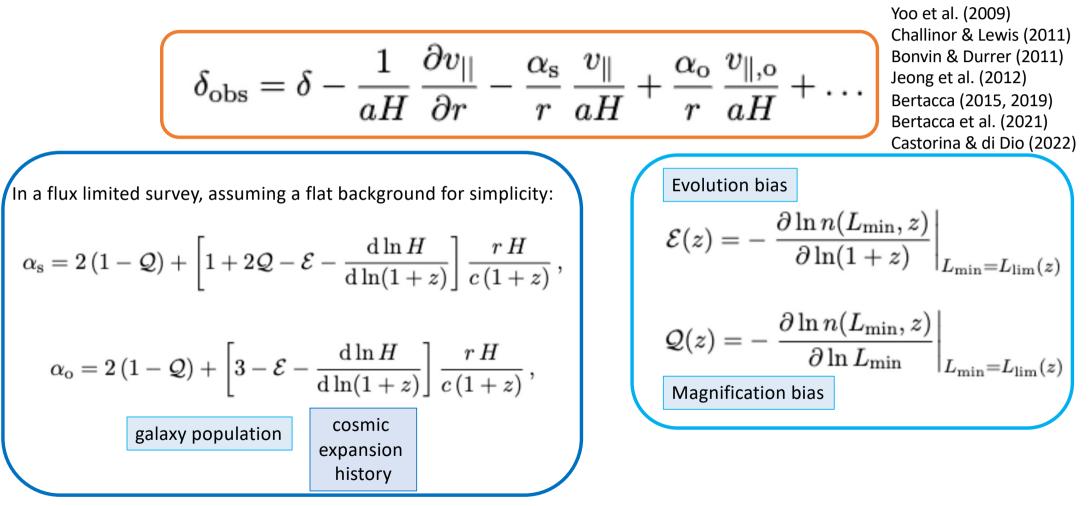
Nick Kaiser (1954-2023)

General relativistic (or light-cone projection) effects



- Bending of light rays due to intervening density fluctuations (gravitational-lensing deflection and magnification)
- Gravitational redshift, integrated and notintegrated Sachs-Wolfe efffects
- Difference between the rest frames of sources and observer (relativistic aberration)
- Shapiro time delay

General relativistic effects (linear order)



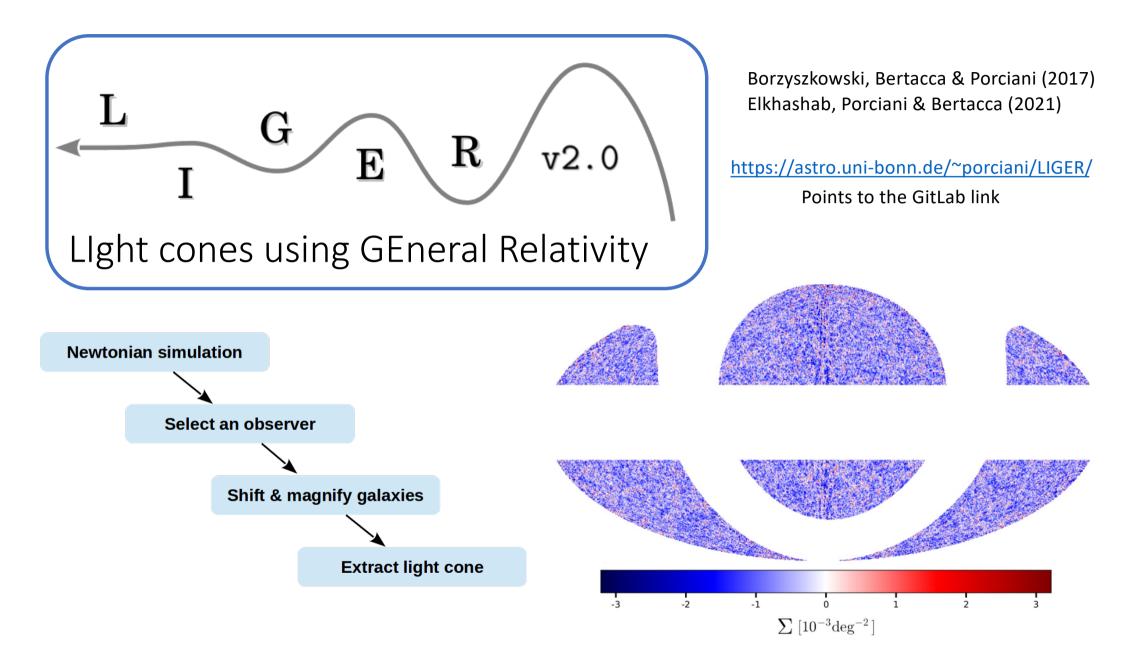
The Finger of the Observer (FOTO) Effect



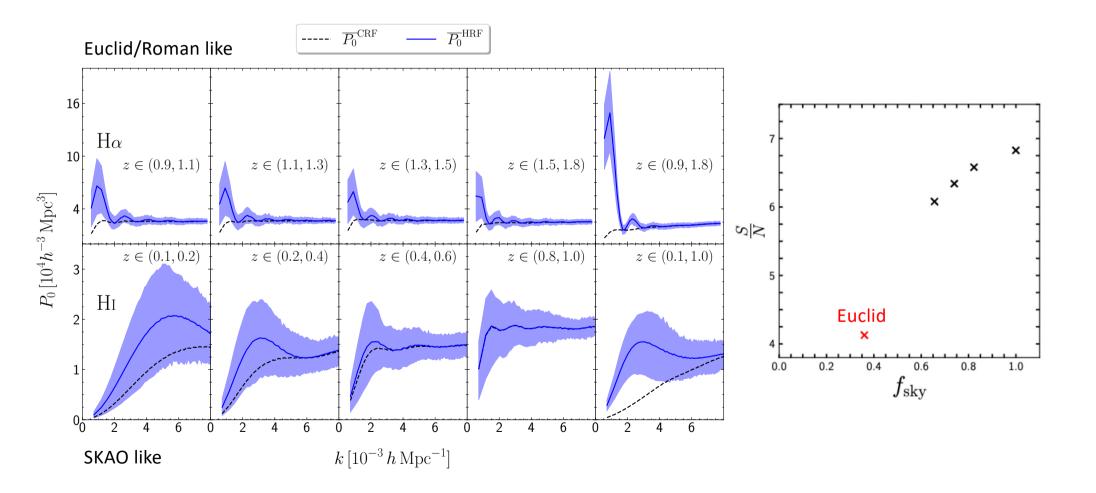
The FOTO effect superimposes oscillations to the monopole of the power spectrum on large scales (also adds other signatures to higher multipoles). The correction can be computed analytically, in general. For a full-sky thin radial shell:

$$P_{0, ext{dip}}(k) \simeq rac{4\pi}{3} \, lpha_{ ext{o}}^2 \, \left[\left(rac{v_{ ext{o}}}{aH}
ight)^2 \, \Delta r
ight] \, j_1^2(kr)$$

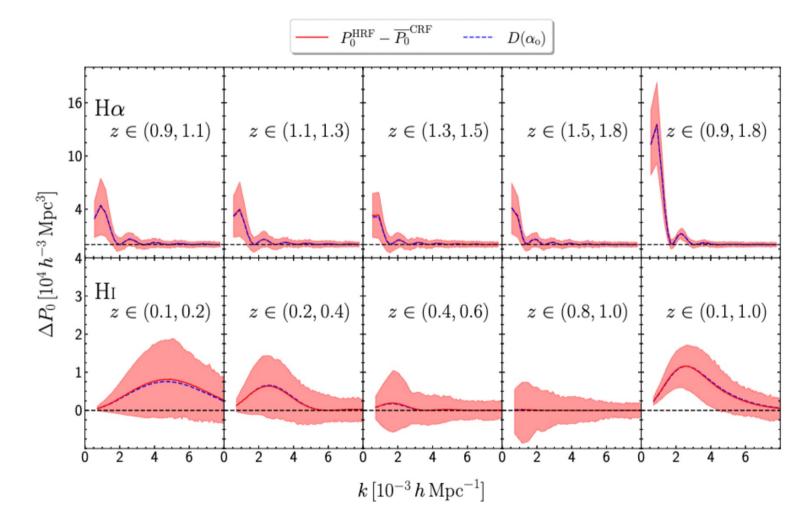
Elkhashab et al. 2021



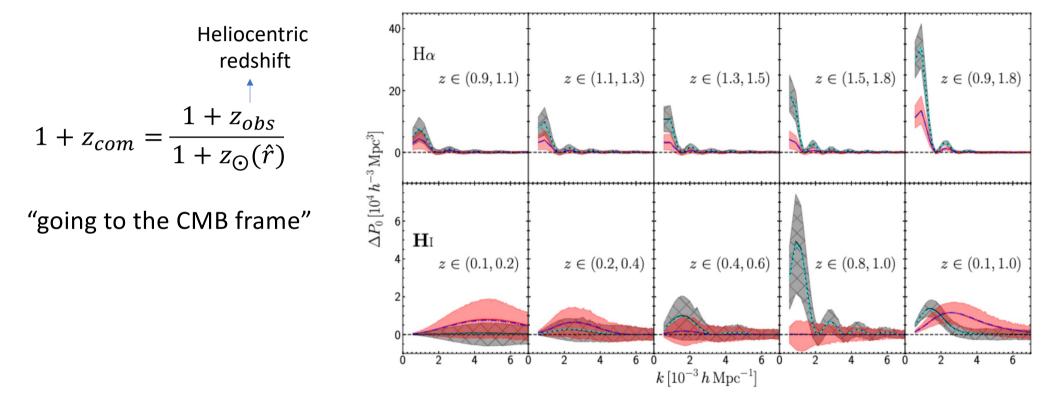
Measurements from 140 mock catalogs



Is the model accurate?



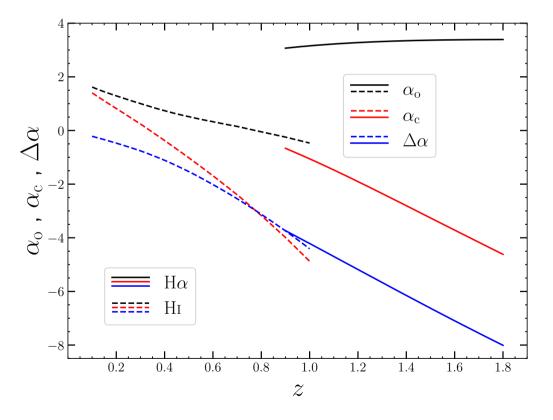
Does a redshift boost cancel the FOTO effect?



- The FOTO effect is not erased because the correction only changes the redshifts of the sources and not their angular positions, sizes and luminosities. In many cases the effect is actually enhanced!
- The net effect of the redshift correction is to modify the α_o parameter in a calculable way.

Enhancing the signal with artificial redshift boosts

$$\delta_{\rm cor} = \delta_{\rm com} + \frac{\alpha_{\rm o}}{a \, H \, r} [(\boldsymbol{v}_{\rm cor} + \boldsymbol{v}_{\odot}) \cdot \hat{\mathbf{r}}] + \frac{\Delta \alpha}{a \, H \, r} (\boldsymbol{v}_{\rm cor} \cdot \hat{\mathbf{r}})$$

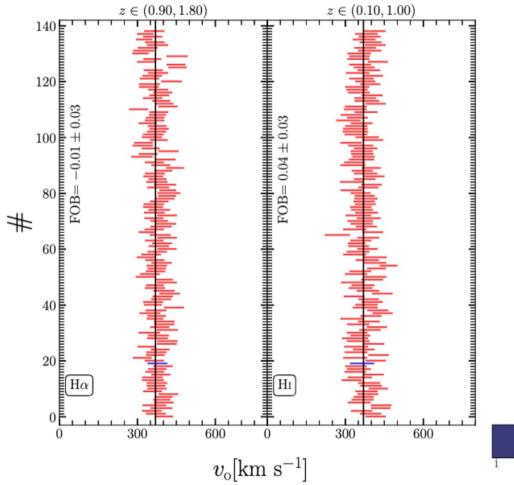


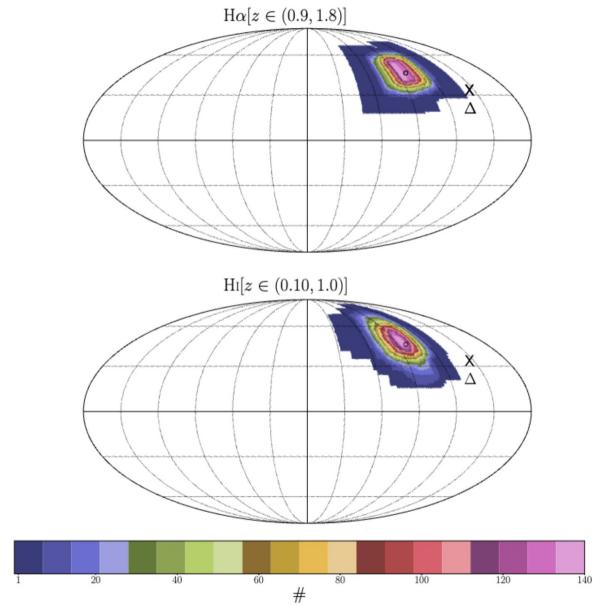
- We simultaneously analyse spectra taken by using up to 5 redshift boosts with different v_{cor} and accounting for their covariances
- This gives us a handle on the direction of the peculiar velocity of the Sun

Possible objectives

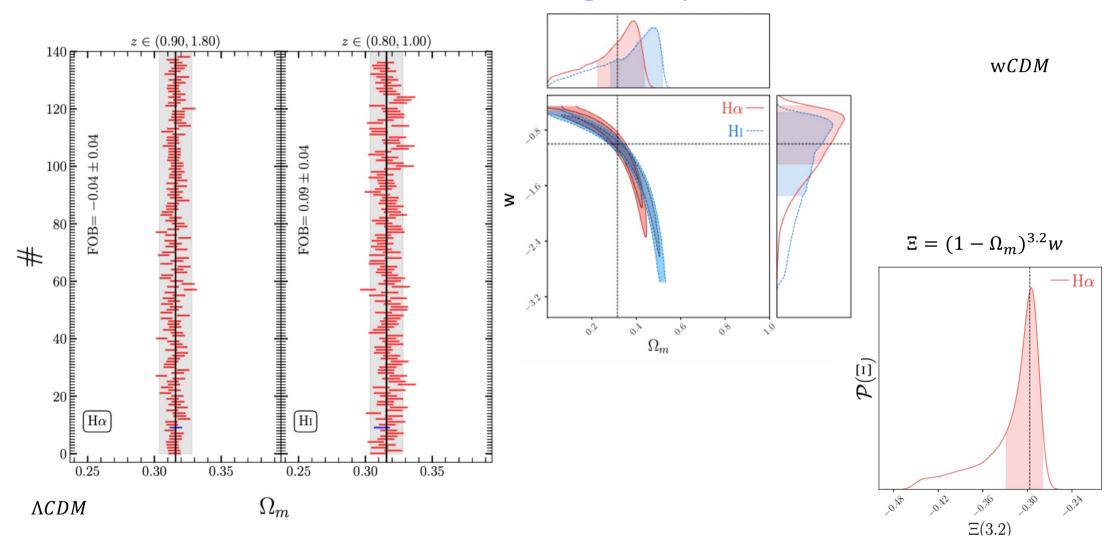
- Low ambition: use priors on cosmology and v_{obs} from CMB studies and measure the evolution and magnification bias of the sources
- Medium ambition: use priors on cosmology and the measurements of the luminosity function to set constraints on v_{obs}
- High ambition (madness?): use priors on v_{obs} from CMB and measurements of the luminosity function to the FOTO effect as a cosmological probe







Could we constrain cosmological parameters?



Caveat

- Measuring the power spectrum on very large scales is challenging
- Variations of the flux limit between areas observed at different times and other systematic effects (e.g. dust corrections) could create spurious clustering
- On the other hand, the signal we are after has a very characteristic signature

Conclusions

- Our peculiar velocity modifies the redshift, size, and luminosity of cosmological sources
- The observed galaxy overdensity contains a dipolar deterministic term proportional to $v_{\|obs}$ (the FOTO effect)
- This generates characteristic oscillatory patterns in the monopole moment of the power spectrum on large scales
- This signal cannot be erased with a simple redshift transformation. Actually, we can take advantage of it to enhance the effect.
- If clustering statistics can be robustly measured on such large scales, the FOTO effect gives a handle to measure v_{obs} and constrain the expansion history of the Universe