

# Particle Optics – part I

N.Biancacci

With many thanks to J.De Conto, A.Latina, P.Lebrun and E.Métral for all the input in the preparation!



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# Objectives of this course

## Part I: Light optics

- Understand the basic laws of light propagation
- Being able to apply the principle of geometrical optics to trace rays of light across common optical elements (boundaries, thin and thick lenses, complex optical systems, ...)



Strong analogy

## Part II: Particle optics

- Understand the basic laws of particle transport
- Being able to apply the principles of light optics to trace particles across common accelerators elements (dipoles, quadrupoles, ...)
- Understand the effect of dispersion on spectrometry

# Outline of part I

- Recall of Maxwell equations
  - Integral form
  - Local form
  - Boundary conditions
  - The plane wave
  - Law of reflection & refraction
- Ray optics
  - Eikonal equation
  - ABCD matrix for simple elements (mirror, drifts, planar and curved interfaces)
  - Thick and thin lenses
  - Image formation
  - Complex systems: principal planes
  - Light transport in periodic channels: stability condition
  - Limit of the treatment: aberrations

# Recall of Maxwell's equations

Short recap, more details in H.Henke's lectures!

**juas...**

**JUAS - TIMETABLE 2020 - WEEK 1**

Schedule 2020	Monday Jan 13	Tuesday Jan 14	Wednesday Jan 15	Thursday Jan 16	Friday Jan 17
09:00		Relativity <i>H. Henke</i>	Electro-magnetism <i>H. Henke</i>	Intro. to Accelerator Design <i>Ph. Bryant</i>	Intro. to the Mini-Workshop <i>Ph. Bryant</i>
10:00		Coffee Break	Coffee Break	Coffee Break	Coffee Break
10:15		Relativity <i>H. Henke</i>	Electro-magnetism <i>H. Henke</i>	Intro. to Accelerator Design <i>Ph. Bryant</i>	Intro. to the Mini-Workshop <i>Ph. Bryant</i>
11:15		Relativity <i>H. Henke</i>	Electro-magnetism <i>H. Henke</i>	Intro. to Accelerator Design <i>Ph. Bryant</i>	Bus leaves at 11:15 from JUAS (Lunch at CERN, R3, offered by ESI)
12:15	12:00 OFFICIAL OPENING (welcome & building visit)	BREAK	BREAK	BREAK	13:30 Visit of LHC Magnets Test Hall <i>M. Bajko</i>
14:00	13:00 WELCOME LUNCH	Relativity <i>H. Henke</i>	Electro-magnetism <i>H. Henke</i>	Intro. to Accelerator Design <i>Ph. Bryant</i>	15:00 Introduction to CERN & its Accelerator Network Seminar - <i>R. Alemany</i>
15:00	14:00 Presentation of JUAS & Introduction of students <i>P. Lebrun</i>	Particle optics <i>N. Biancacci</i>	Particle optics <i>N. Biancacci</i>	Intro. to Accelerator Design <i>Ph. Bryant</i>	16:30 Visit of CERN Control Center <i>R. Alemany</i>
16:00	History of particle accelerators Seminar <i>V. Vaccaro</i>	Coffee Break	Coffee Break	Coffee Break	Bus leaves at 17:30 from CERN
16:15	CHECK-IN AT THE RESIDENCE & SHOPPING FOR GROCERIES	Particle optics <i>N. Biancacci</i>	Particle optics <i>N. Biancacci</i>	Intro. to Accelerator Design <i>Ph. Bryant</i>	
17:15		Particle optics <i>N. Biancacci</i>			
18:15		AFTER WORK AT ESI			

# Maxwell equations: integral form

In a homogeneous, linear, isotropic medium:

Gauss' law 
$$\int_{\partial S} \bar{\mathcal{E}} \cdot dS = \int_V \frac{\rho}{\varepsilon} dV$$

Gauss' law for magnetism 
$$\int_{\partial S} \bar{\mathcal{H}} \cdot dS = 0$$

Faraday's law 
$$\oint_{\partial S} \bar{\mathcal{E}} \cdot d\mathbf{l} = -\frac{d}{dt} \int_{\partial S} \mu \bar{\mathcal{H}} \cdot dS$$

Ampère-Maxwell law 
$$\oint_{\partial S} \bar{\mathcal{H}} \cdot d\mathbf{l} = \int_S \bar{\mathcal{J}} \cdot dS + \frac{d}{dt} \int_S \varepsilon \bar{\mathcal{E}} \cdot dS$$

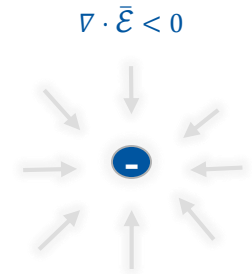
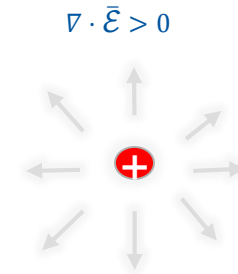
$\varepsilon$ : medium permittivity

$\mu$ : medium permeability

# Local form

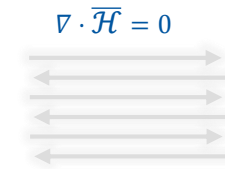
Gauss' law

$$\nabla \cdot \vec{\mathcal{E}} = \frac{\rho}{\varepsilon}$$



Gauss' law for magnetism

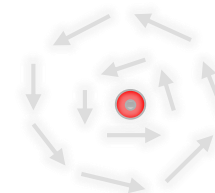
$$\nabla \cdot \vec{\mathcal{H}} = 0$$



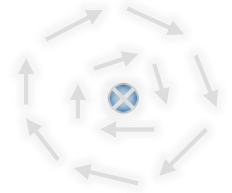
Faraday's law

$$\nabla \times \vec{\mathcal{E}} = -\frac{d}{dt}\mu\vec{\mathcal{H}}$$

$\nabla \times \vec{\mathcal{H}} > 0$



$\nabla \times \vec{\mathcal{H}} < 0$



Ampère-Maxwell law

$$\nabla \times \vec{\mathcal{H}} = \vec{\mathcal{J}} + \frac{d}{dt}\varepsilon\vec{\mathcal{E}}$$

# Local form in frequency domain

Suppose the field is stationary, we can apply the **Fourier transform**:

$$\bar{E}(\omega) = \int_{-\infty}^{+\infty} \bar{E}(t) e^{j\omega t} dt$$

With this,  $\frac{d}{dt} \rightarrow j\omega$  and dropping the time dependence we have:

Gauss' law  $\nabla \cdot \bar{E} = \rho/\varepsilon$

Gauss' law for magnetism  $\nabla \cdot \bar{H} = 0$

Faraday's law  $\nabla \times \bar{E} = -j\omega\mu\bar{H}$

Ampère-Maxwell law  $\nabla \times \bar{H} = j\omega\varepsilon\bar{E} + \bar{J}$

# Boundary conditions

Let's now recap the behavior of the fields at the boundaries:

Electric field

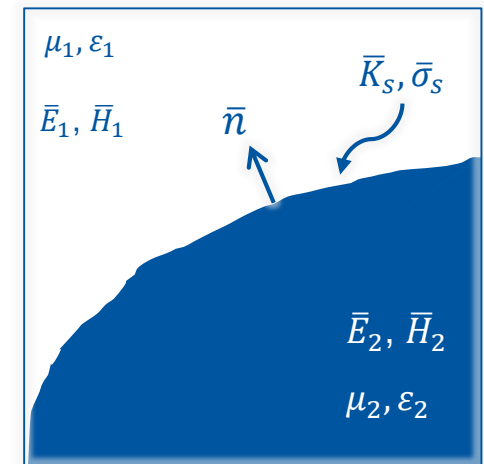
$$\bar{n} \times (\bar{E}_1 - \bar{E}_2) = 0$$

$$\bar{n} \cdot (\epsilon_1 \bar{E}_1 - \epsilon_2 \bar{E}_2) = \bar{\sigma}_s$$

Magnetic field

$$\bar{n} \times (\bar{H}_1 - \bar{H}_2) = \bar{K}_s$$

$$\bar{n} \cdot (\mu_1 \bar{H}_1 - \mu_2 \bar{H}_2) = 0$$



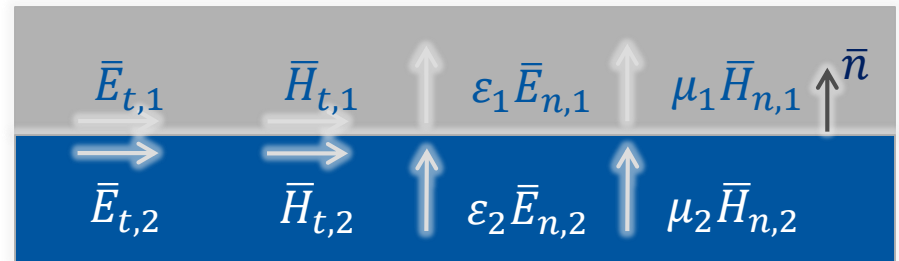
where  $\bar{K}_s$  and  $\bar{\sigma}_s$  are free surface electric current and charges.



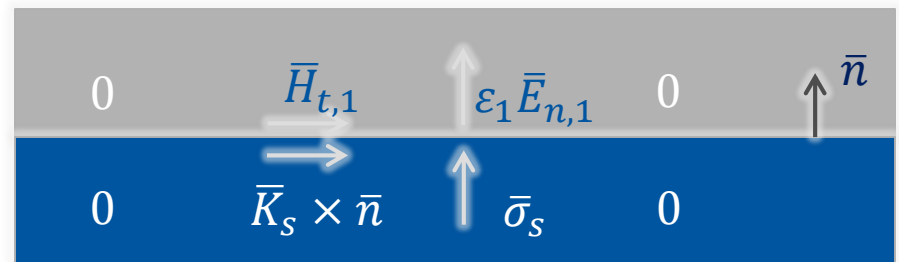
# Boundary conditions

Example for **dielectrics** and **perfect electric conductor (PEC)**

For a **dielectric**  $\rightarrow \bar{\sigma}_s, \bar{K}_s$  are null



For a **PEC**  $\rightarrow \bar{E}_2, \bar{H}_2$  are null but free current and charges can be present on the surface



# Plane wave

For an homogeneous, isotropic medium (e.g. vacuum, dielectric, magnetic material):

$$\nabla \times \bar{H} = j\omega\varepsilon\bar{E} + \bar{J}$$

$$\nabla \times \bar{E} = -j\omega\mu\bar{H}$$

$$\nabla \times \left( \frac{1}{-j\omega\mu} \nabla \times \bar{E} \right) = j\omega\varepsilon\bar{E} + \bar{J}$$

$$\nabla \times (\nabla \times \bar{E}) = \omega^2\mu\varepsilon\bar{E} - j\omega\mu\bar{J}$$

$$\nabla \times (\nabla \times \bar{E}) = \nabla(\nabla \cdot \bar{E}) - \nabla^2 \bar{E}$$

$$\nabla(\nabla \cdot \bar{E}) - \nabla^2 \bar{E} = \omega^2\mu\varepsilon\bar{E} - j\omega\mu\bar{J}$$

$$\nabla \left( \frac{\rho}{\varepsilon} \right) - \nabla^2 \bar{E} = \omega^2\mu\varepsilon\bar{E} - j\omega\mu\bar{J}$$

$$\nabla \cdot \varepsilon\bar{E} = \rho$$

In vacuum ( $\rho, \bar{J}$  null):

$$\nabla^2 \bar{E} + \omega^2\mu\varepsilon\bar{E} = 0$$

$$k^2 = \omega^2\mu\varepsilon \rightarrow k = \frac{\omega}{c} \mathbf{n} \text{ with } \mathbf{n} = \sqrt{\varepsilon_r} \text{ refraction index}$$

(real in dielectrics with negligible absorption)

# Plane wave

Maxwell equations in **free space** condense into Helmholtz equation:

$$\nabla^2 \bar{E} + k^2 \bar{E} = 0 \quad \text{with } k = \frac{\omega}{c} n$$

The solution is a plane wave in free space with propagation vector  $\bar{k} \cdot \bar{k}_0 = k_x \bar{x}_0 + k_y \bar{y}_0 + k_z \bar{z}_0$ .

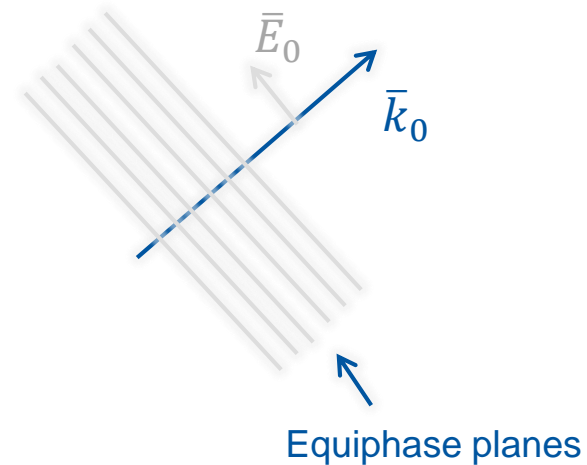
$$\bar{E} = \bar{E}_0 e^{-j(k_x x + k_y y + k_z z)} = \bar{E}_0 e^{-j(\bar{k} \cdot \bar{r})}$$

Back in time domain this is:

$$\bar{\mathcal{E}} = \bar{E}_0 \cos(\bar{k} \cdot \bar{r} - \omega t + \phi) = \bar{E}_0 \cos\left(\frac{\omega}{c} n \cdot \bar{k}_0 \cdot \bar{r} - \omega t + \phi\right)$$

where the **equiphase planes** are given by:  $\frac{\omega}{c} n \cdot \bar{k}_0 \cdot \bar{r} - \omega t = \text{const}$

and travel in the **direction of  $\bar{k}_0$**  with phase velocity  $\frac{dr}{dt} \bar{k}_0 \cdot \bar{r}_0 = \frac{c}{n} = v_p$



# Laws of reflection & refraction

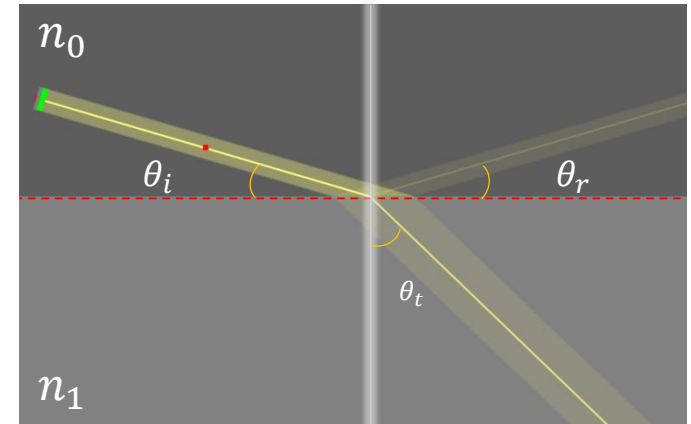
Consider two media with different index  $n_0$  and  $n_1$ . In order to satisfy the boundary conditions a **reflected** and a **refracted** wave are produced.

- The angle of **reflection** is equal to the angle of incidence (*law of reflection*):

$$\theta_i = \theta_r$$

- The angle of **refraction** is given by (*Snell law*):

$$n_0 \sin(\theta_i) = n_1 \sin(\theta_t)$$



E.g. going from a medium with lower  $n$  to a larger one, the angle of propagation of the refracted wave gets closer to the normal to the interfaces.

P.S.: Have fun with this nice optics simulator at <https://ricktu288.github.io/ray-optics/> 😊

# Ray optics

Now let's consider  $n(\vec{r})$ , i.e. a **refraction index varying in space**.

$$\nabla^2 \bar{E} + \frac{\omega^2}{c^2} n(\vec{r})^2 \bar{E} = 0 \quad \longrightarrow \quad \nabla^2 \bar{E} + k_0^2 n(\vec{r})^2 \bar{E} = 0$$

We look for a solution as for a plane wave:

$$\bar{E} = \bar{E}_0 e^{-j(k_0 R(\vec{r}))}$$

$$\frac{1}{k_0^2} \nabla^2 (\bar{E}_0 e^{-j(k_0 R(\vec{r}))}) + n(\vec{r})^2 \bar{E}_0 e^{-j(k_0 R(\vec{r}))} = 0$$

The computation of  $\nabla^2 (\bar{E}_0 e^{-j(k_0 R(\vec{r}))})$  is rather lengthy and it is left in appendix for curiosity.

# Ray optics

The result is this:

$$\left[ -\frac{j}{k_0} \nabla^2 R(\vec{r}) \vec{E}_0 - \right.$$



$$- \hat{\mathbf{z}} \nabla E_{0,z} \cdot \nabla R(\vec{r}) \Big)$$

# Light optics

$$\left[ -\frac{j}{k_0} \nabla^2 R(\vec{r}) \bar{E}_0 - \frac{2j}{k_0} \left( \hat{x} \nabla E_{0,x} \cdot \nabla R(\vec{r}) + \hat{y} \nabla E_{0,y} \cdot \nabla R(\vec{r}) + \hat{z} \nabla E_{0,z} \cdot \nabla R(\vec{r}) \right) \right]$$

For **very short wavelengths** (light  $\rightarrow$  400 – 700 nm scale),  $1/k_0 \rightarrow 0$  and the solution simplifies to:

$$|\nabla R(\vec{r})|^2 = n(\vec{r})^2 \quad \text{Eikonal equation (\textit{εἰκὼν}, image)}$$

Or equivalently:

$$\nabla R(\vec{r}) = n(\vec{r}) \hat{s}(r)$$

- The **energy flows** in the direction of  $\nabla R(\vec{r})$ , i.e. **the gradient of the wavefronts**.
- For a **homogeneous medium**  $n = \text{const}$  and  $R(\vec{r}) = n/k_0(k_x x + k_y y + k_z z)$  represents the **plane wavefronts** (verify that  $|\nabla R(\vec{r})|^2 = n^2$ ).
- By this **simplification** we pay the price of not being able to describe phenomena as **diffraction, reflection**.

# Daylife example

Consider a **very hot day** in summer, asphalt gets hot and the **air close to it expands lowering the refraction index**.

$$n(y) = n_0 \sqrt{1 + \frac{y^2}{h^2}}$$

We can compute the ray direction:

$$|\nabla R(\vec{r})|^2 = n(\vec{r})^2$$

$$\left(\frac{\partial R}{\partial x}\right)^2 + \left(\frac{\partial R}{\partial y}\right)^2 = n(\vec{r})^2 = n_0^2 \left(1 + \frac{y^2}{h^2}\right)$$

$$R(x, y) = (n_0 x) \hat{x}_0 + \left(1 + \frac{y^2}{2h^2}\right) \hat{y}_0$$

The direction of light rays is then:

$$\hat{s}(x, y) = \frac{\nabla R(x, y)}{|\nabla R(x, y)|} = \frac{\left(\frac{y}{h} \hat{y}_0 + \hat{x}_0\right)}{\sqrt{1 + \frac{y^2}{h^2}}}$$



Source: <https://www.scienceabc.com>





# Daylife example

The ray covers a distance  $dx$  with direction given by  $\hat{s}(x, y)$ . The gained altitude  $dy$  can be then approximated by:

$$dy = \frac{s_y}{s_x} dx$$

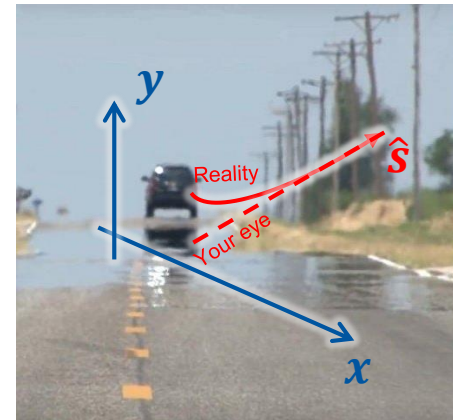
From this we have

$$\frac{dy}{dx} = \frac{s_y}{s_x} = \frac{y}{h}$$

With solution

$$y = e^{\pm(x-x_0)/h}$$

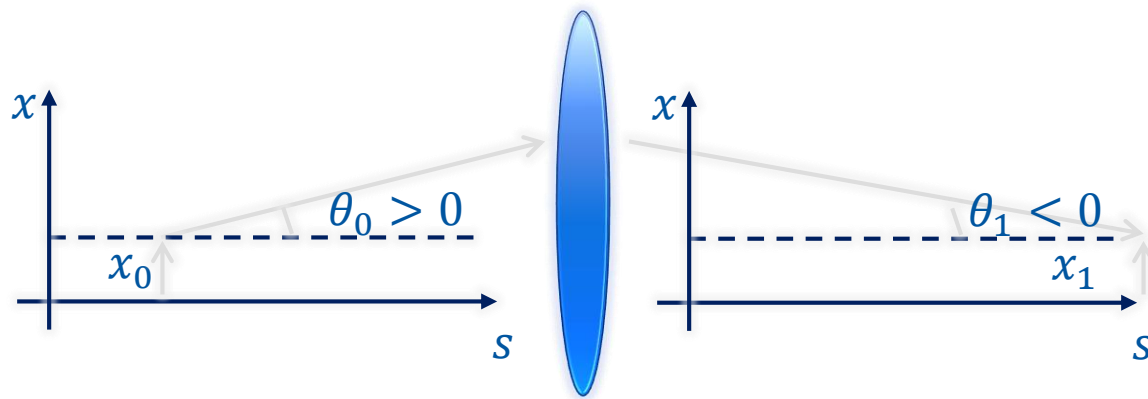
Light bends up (solution with +) emulating a reflection from a wet surface.



Source: <https://www.scienceabc.com>

# Matrix treatment

The interaction of a wave can be reduced to the study of the **interaction** of the ray direction **at an object boundary**.



Element's matrix:

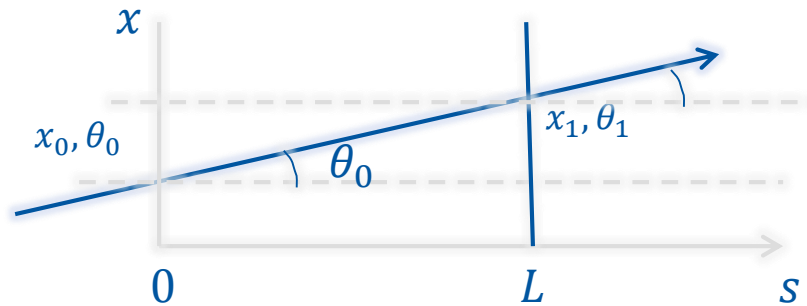
$$\begin{pmatrix} x_1 \\ \theta_1 \end{pmatrix} = \begin{bmatrix} A & B \\ C & D \end{bmatrix} \begin{pmatrix} x_0 \\ \theta_0 \end{pmatrix}$$

$$M \equiv \frac{x_1}{x_0} = A \text{ magnification}$$

$$M_\theta \equiv \frac{\theta_1}{\theta_0} = D \text{ angular magnification}$$

# Drift space

Given a point on a wavefront of a ray, if we are in a homogeneous medium, it will just **drift in space**:



$$\begin{cases} \theta_1 = \theta_0 \\ x_1 = x_0 + L \sin \theta_0 = x_0 + L\theta_0 \end{cases}$$

(paraxial approximation  $\sin \theta \simeq \theta$ )

Matrix form

$$\begin{pmatrix} x_1 \\ \theta_1 \end{pmatrix} = M \begin{pmatrix} x_0 \\ \theta_0 \end{pmatrix} = \begin{bmatrix} 1 & L \\ 0 & 1 \end{bmatrix} \begin{pmatrix} x_0 \\ \theta_0 \end{pmatrix}$$

N.B.:  $\det(M) = 1$

# Spherical mirror

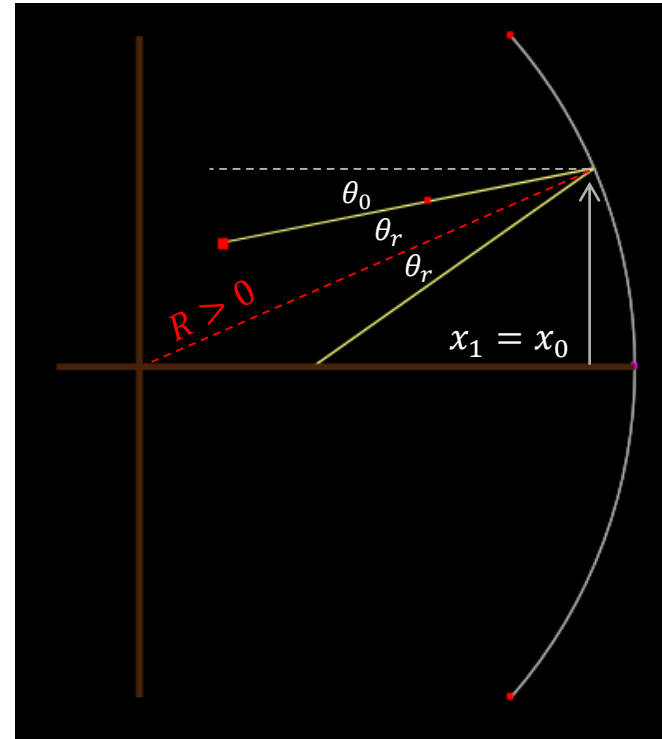
Let us consider a **spherical mirror** (concave  $R>0$ , convex  $R<0$ ).

$$\begin{cases} x_1 = x_0 \\ \theta_1 = -(\theta_0 + 2\theta_r) \\ \theta_r + \theta_0 = \frac{x_1}{R} \end{cases}$$

$$\downarrow$$
$$\theta_1 = -\frac{2x_1}{R} + \theta_0$$

$$\begin{pmatrix} x_1 \\ \theta_1 \end{pmatrix} = \begin{bmatrix} 1 & 0 \\ -2/R & 1 \end{bmatrix} \begin{pmatrix} x_0 \\ \theta_0 \end{pmatrix}$$

$$\text{N.B.: } \det(M) = 1$$



# Curved interface

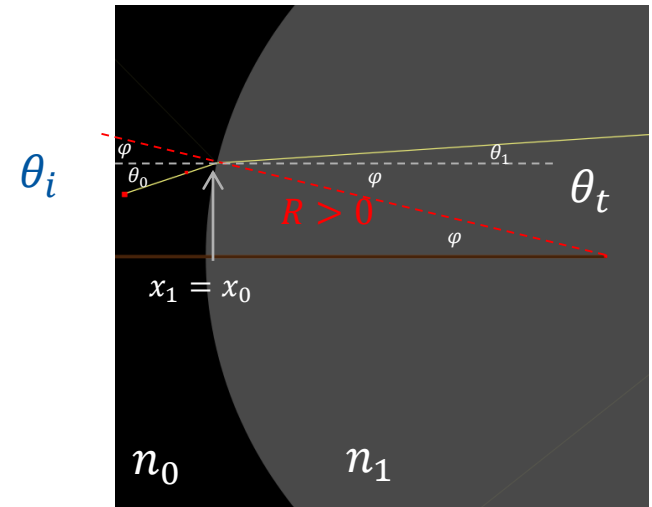
Let us consider a curved interface (convex  $R>0$ , concave  $R<0$ )

$$\left\{ \begin{array}{l} x_1 = x_0 \\ n_0 \theta_i = n_1 \theta_t \text{ (Snell's law)} \\ \theta_i = \varphi + \theta_0 \\ \theta_t = \varphi + \theta_1 \\ \varphi = x_1/R \end{array} \right.$$



$$\begin{aligned} n_0(x_1/R + \theta_0) &= n_1(x_1/R + \theta_1) \\ \frac{n_0}{n_1} \left( \frac{x_1}{R} + \theta_0 \right) - x_1/R &= \theta_1 \end{aligned}$$

$$\begin{pmatrix} x_1 \\ \theta_1 \end{pmatrix} = \begin{bmatrix} 1 & 0 \\ \left( \frac{n_0}{n_1} - 1 \right) / R & n_0/n_1 \end{bmatrix} \begin{pmatrix} x_0 \\ \theta_0 \end{pmatrix}$$



N.B.:  $\det(M) = 1$  if  $n_0 = n_1$ :  
this is a result applicable to any  
optical system with same  
start/end refractive index.

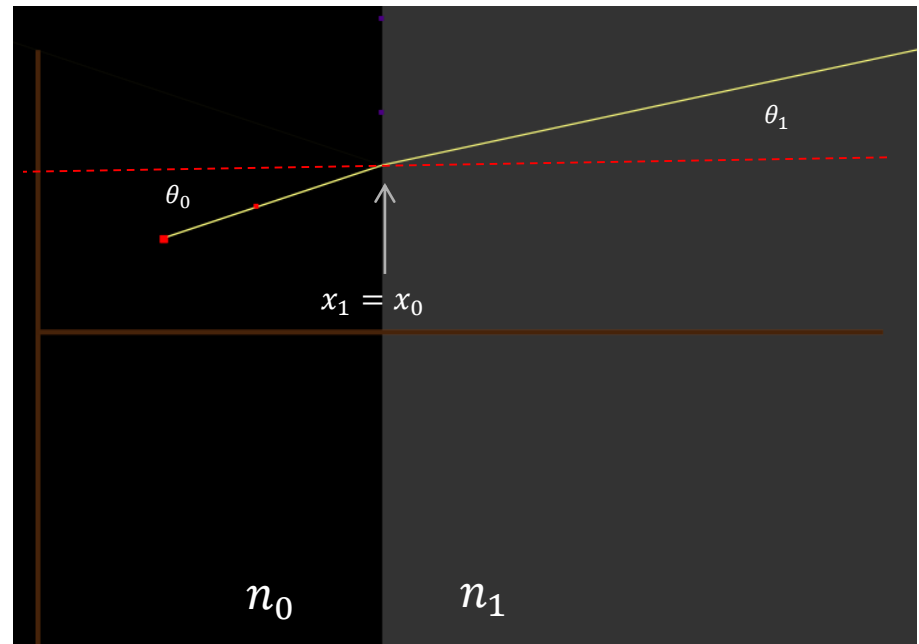
# Curved interface: planar interface

For a planar interface we simply take the limit of large radius.

$$\begin{pmatrix} x_1 \\ \theta_1 \end{pmatrix} = \begin{bmatrix} 1 & 0 \\ \left(\frac{n_0}{n_1} - 1\right)/R & n_0/n_1 \end{bmatrix} \begin{pmatrix} x_0 \\ \theta_0 \end{pmatrix}$$

$\downarrow R \rightarrow \infty$

$$\begin{pmatrix} x_1 \\ \theta_1 \end{pmatrix} = \begin{bmatrix} 1 & 0 \\ 0 & n_0/n_1 \end{bmatrix} \begin{pmatrix} x_0 \\ \theta_0 \end{pmatrix}$$

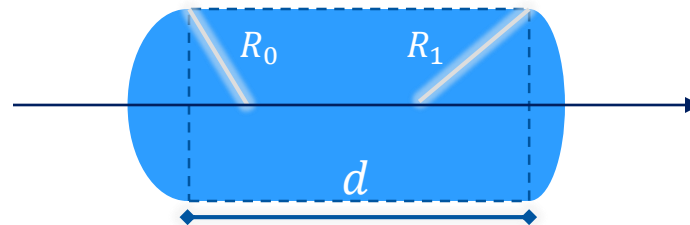


# Exercise 1: Thick lens

We have derived the **basic elements** that allow us to derive the matrices for more **complex optical systems** by simple matrix multiplication:

- 1) propagation through a region of uniform index,
- 2) reflection from a curved mirror
- 3) transmission through a curved interface of regions with different indices.

**Classwork:** Derive the ABCD matrix for a thick lens made of material  $n_1 = n$  surrounded by air ( $n_0 = 1$ ). Let the lens have curvatures  $R_0$  and  $R_1$  and thickness  $d$ .



What happens for  $R_0, R_1 \rightarrow \infty$ ? Make a sketch.

# Exercise 1: Thick lens

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$$\begin{pmatrix} x_1 \\ \theta_1 \end{pmatrix} = \begin{bmatrix} 1 - \frac{d}{R_0} \left( 1 - \frac{1}{n} \right) & \frac{d}{n} \\ -(n-1) \left( \frac{1}{R_0} - \frac{1}{R_1} \right) + \frac{d}{R_0 R_1} \left( 2 - n - \frac{1}{n} \right) & 1 + \frac{d}{R_1} \left( 1 - \frac{1}{n} \right) \end{bmatrix} \begin{pmatrix} x_0 \\ \theta_0 \end{pmatrix}$$



# Thin lens

Considering a **vanishing length**  $d$  between the two lens surfaces we have:

$$\begin{pmatrix} x_1 \\ \theta_1 \end{pmatrix} = \begin{bmatrix} 1 & 0 \\ -(n-1)\left(\frac{1}{R_0} - \frac{1}{R_1}\right) & 1 \end{bmatrix} \begin{pmatrix} x_0 \\ \theta_0 \end{pmatrix}$$

Which can be written as

$$\begin{pmatrix} x_1 \\ \theta_1 \end{pmatrix} = \begin{bmatrix} 1 & 0 \\ -1/f & 1 \end{bmatrix} \begin{pmatrix} x_0 \\ \theta_0 \end{pmatrix}$$

Where  $f$  is the **focal length** given by:

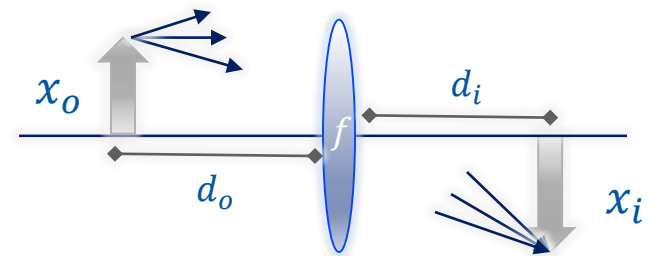
$$1/f = (n-1) \left( \frac{1}{R_0} - \frac{1}{R_1} \right) \quad \text{Lens maker's equation}$$

When the **radii are the same** (in modulo):  $\frac{1}{f} = (n-1) \frac{2}{R}$

# Image formation

Given a **transfer matrix**, we would like to know if, placing an object in front of the optical system ( $d_o$ ), an **image is formed**, and **where** ( $d_i$ ).

$$\begin{pmatrix} x_i \\ \theta_i \end{pmatrix} = \begin{bmatrix} A & B \\ C & D \end{bmatrix} \begin{pmatrix} x_o \\ \theta_o \end{pmatrix}$$



To have an image, all the rays should go to the same point regardless of the angle they start. A point  $x_o$  is mapped into a point  $x_i$  regardless of the angles  $\theta_o$  the rays have which means:

**$B = 0$**  (condition of image formation).

# Image formation

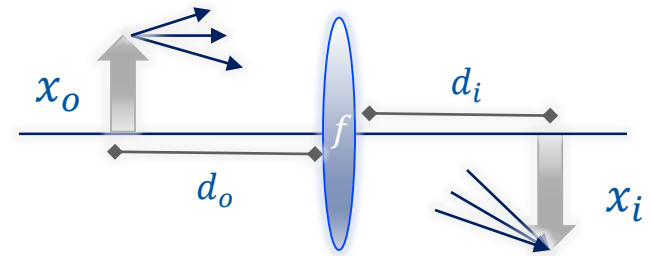
If the object is at distance  $d_o$ , for a **thin lens** we have:

$$\begin{pmatrix} x_i \\ \theta_i \end{pmatrix} = \begin{bmatrix} 1 & d_i \\ 0 & 1 \end{bmatrix} \begin{bmatrix} 1 & 0 \\ -1/f & 1 \end{bmatrix} \begin{bmatrix} 1 & d_o \\ 0 & 1 \end{bmatrix} \begin{pmatrix} x_o \\ \theta_o \end{pmatrix}$$

$$\begin{pmatrix} x_i \\ \theta_i \end{pmatrix} = \begin{bmatrix} 1 - \frac{d_i}{f} & d_o + d_i \left( -\frac{1}{f} d_o + 1 \right) \\ -\frac{1}{f} & -\frac{1}{f} d_o + 1 \end{bmatrix} \begin{pmatrix} x_o \\ \theta_o \end{pmatrix}$$

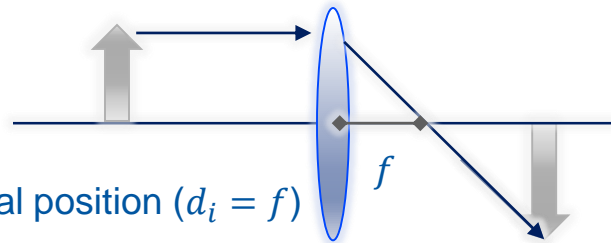
$$\mathbf{B} = \mathbf{0} \rightarrow \begin{cases} \frac{1}{f} = \frac{1}{d_o} + \frac{1}{d_i} \\ M = A = -\frac{d_i}{d_o} \\ M_\theta = D = -\frac{d_o}{d_i} \end{cases} \longrightarrow$$

$$\begin{pmatrix} x_i \\ \theta_i \end{pmatrix} = \begin{bmatrix} -\frac{d_i}{d_o} & 0 \\ -\frac{1}{f} & -\frac{d_o}{d_i} \end{bmatrix} \begin{pmatrix} x_o \\ \theta_o \end{pmatrix}$$



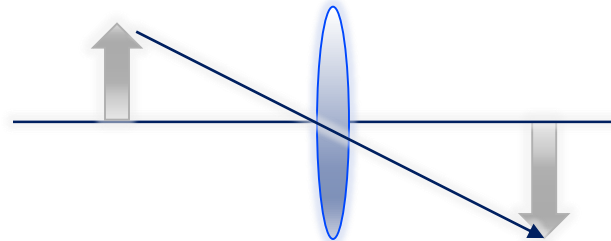
# Practical drawing

1) A ray coming **parallel** to the lens is cross the axis at the **focal point** after the lens

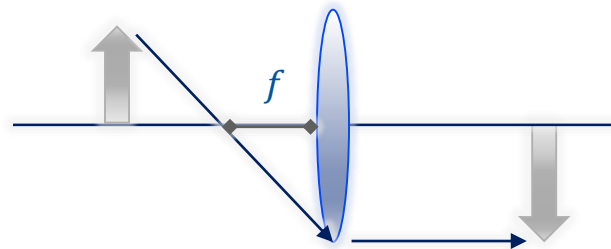


→ an image from far away is produced in the focal position ( $d_i = f$ )

2) A ray passing through the **lens center** is **un-deflected**



3) A ray passing through the **focal point** before the lens **exits parallel**.

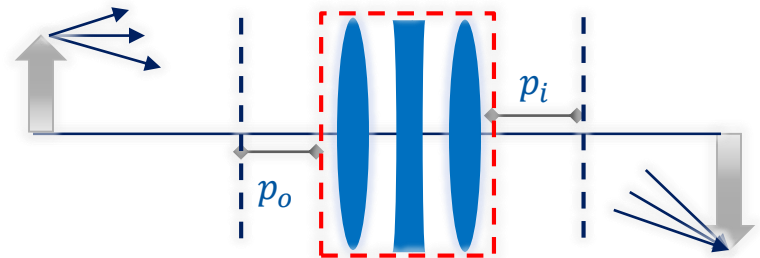


# Principal planes

For a general arrangement of optical elements starting and ending to the same refractive index there exist two planes such that the system allows image formation.

$$\begin{pmatrix} x_i \\ \theta_i \end{pmatrix} = \begin{bmatrix} 1 & p_i \\ 0 & 1 \end{bmatrix} \begin{bmatrix} A & B \\ C & D \end{bmatrix} \begin{bmatrix} 1 & p_o \\ 0 & 1 \end{bmatrix} \begin{pmatrix} x_o \\ \theta_o \end{pmatrix}$$

$$M = \begin{bmatrix} A + p_i C & A p_o + B + p_i p_o C + p_i D \\ C & C p_o + D \end{bmatrix}$$



$$\begin{cases} A + p_i C = 1 \rightarrow p_i = \frac{1 - A}{C} \\ D + p_o C = 1 \rightarrow p_o = \frac{1 - D}{C} \\ C \equiv -\frac{1}{f_{eff}} \end{cases}$$

$$\rightarrow A p_o + B + p_i p_o C + p_i D = 1 - \det(M) = 0$$

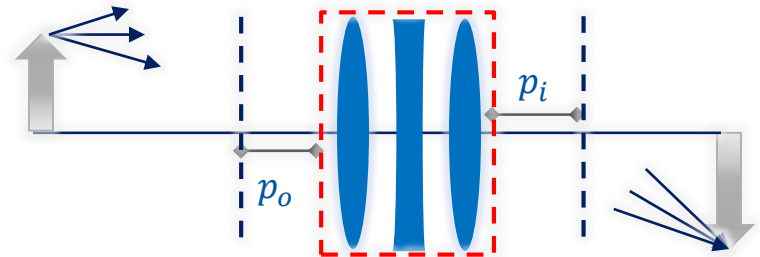
$\stackrel{=1}{\text{(same } n \text{ at start/end)}}$

# Principal planes

For a general arrangement of optical elements **starting and ending to the same refraction index** there exist **two planes** such that the system allows **image formation**.

$$\begin{pmatrix} x_i \\ \theta_i \end{pmatrix} = \begin{bmatrix} 1 & p_i \\ 0 & 1 \end{bmatrix} \begin{bmatrix} A & B \\ C & D \end{bmatrix} \begin{bmatrix} 1 & p_o \\ 0 & 1 \end{bmatrix} \begin{pmatrix} x_o \\ \theta_o \end{pmatrix}$$

$$M = \begin{bmatrix} A + p_i C & Ap_o + B + p_i p_o C + p_i D \\ C & Cp_o + D \end{bmatrix}$$



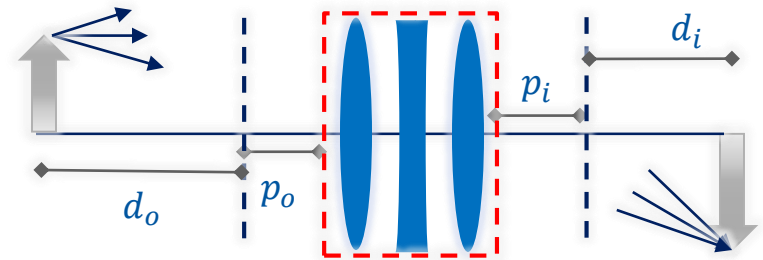
$$M = \begin{bmatrix} 1 & 0 \\ -\frac{1}{f_{eff}} & 1 \end{bmatrix}$$

The planes passing by  $p_o$  and  $p_i$  are called *principal planes*: from there any system will look like a **thin lens** system.

# Principal planes

For a generic system with equal start/end refractive index, we can rewrite the **condition for image formation**:

$$\left\{ \begin{array}{l} \frac{1}{d_o} + \frac{1}{d_i} = \frac{1}{f} \\ M = -\frac{d_i}{d_o} \\ M_\theta = -\frac{d_o}{d_i} \end{array} \right. \quad \begin{array}{l} \text{image condition} \\ \text{magnification} \\ \text{angular magnification} \end{array}$$



These are the same result as of the thin lens case, with exception that now the length is taken **with respect to the principal planes**.

# Practical experiment

Find the focal length of the given lens.

You can help yourself with a light ☺



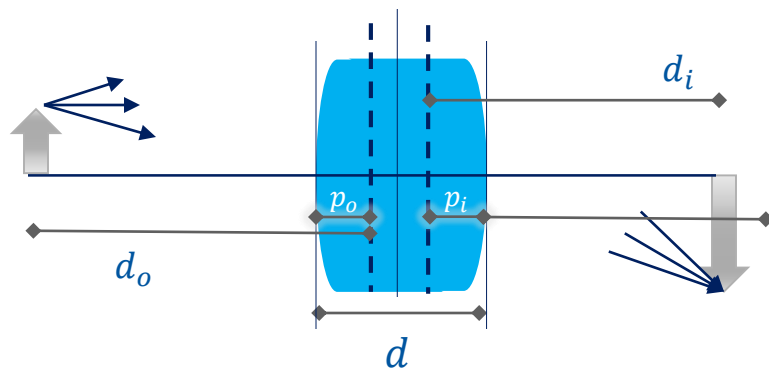
The refraction index of glass is  $n \sim 1.5$ : what is the **curvature radius** of the lens in the **thin lens** assumption?



# Exercise 2

Find the **image position and magnification** of an object placed 30 cm apart from a thick lens of  $d = 1$  cm thickness, in/out radii of 20 cm, made of glass.

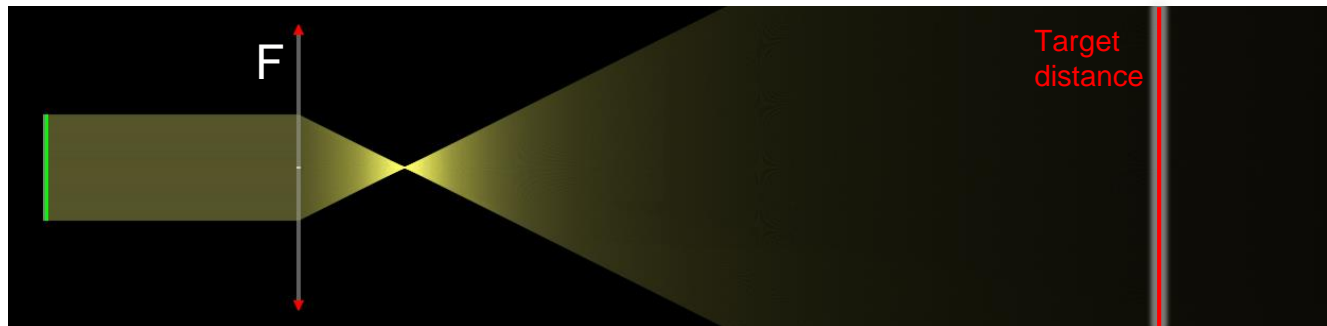
*Reminder: a **thick lens is a complex system**, distances need to be computed from the **principal planes**.*



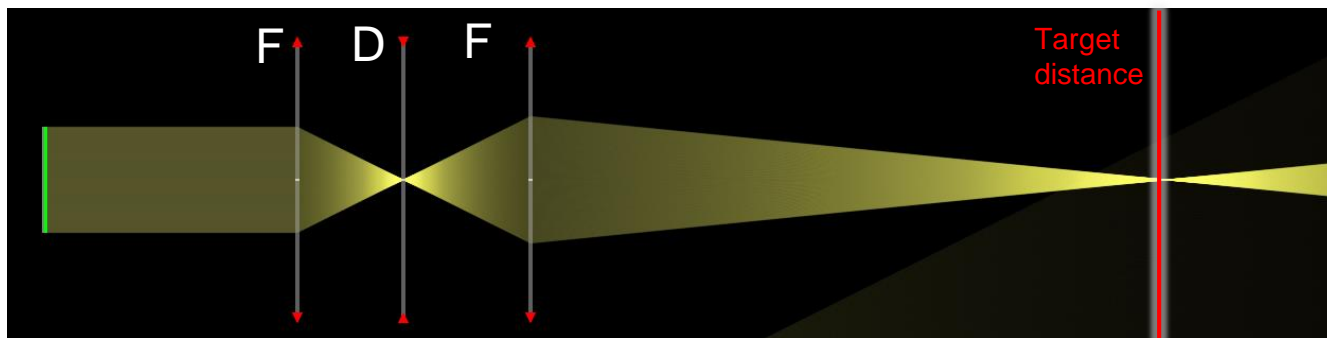
$$M_{thick} = \begin{bmatrix} 1 - \frac{d}{R_0} \left( 1 - \frac{1}{n} \right) & \frac{d}{n} \\ -(n-1) \left( \frac{1}{R_0} - \frac{1}{R_1} \right) + \frac{d}{R_0 R_1} \left( 2 - n - \frac{1}{n} \right) & 1 + \frac{d}{R_1} \left( 1 - \frac{1}{n} \right) \end{bmatrix}$$

# Light transport

Suppose we want to **focus light** at a **large distance**.  
A single lens would not do a great job..



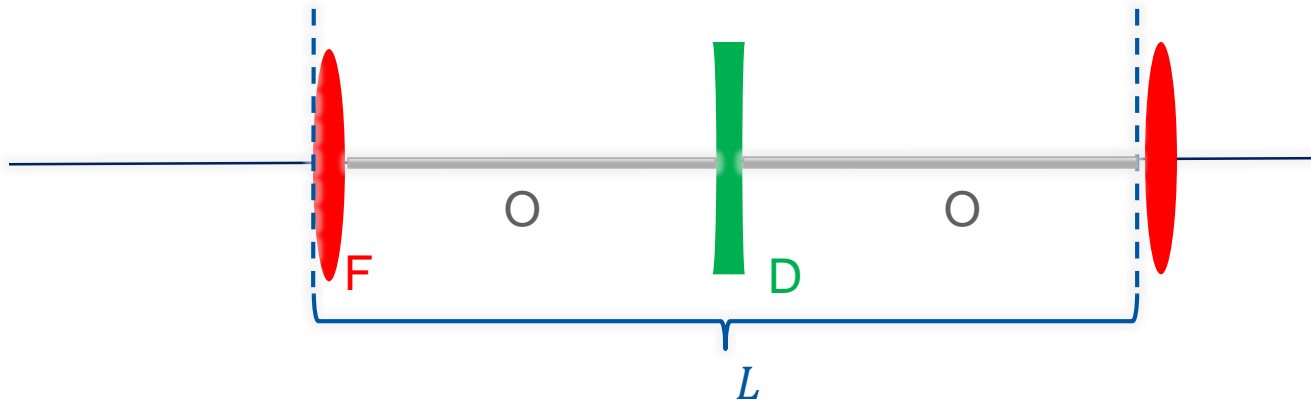
But a **sequence of focusing-defocusing (F-D) elements** could do it!



# Light transport

Sequence of F-D lenses spaced by  $L/2$  drift:

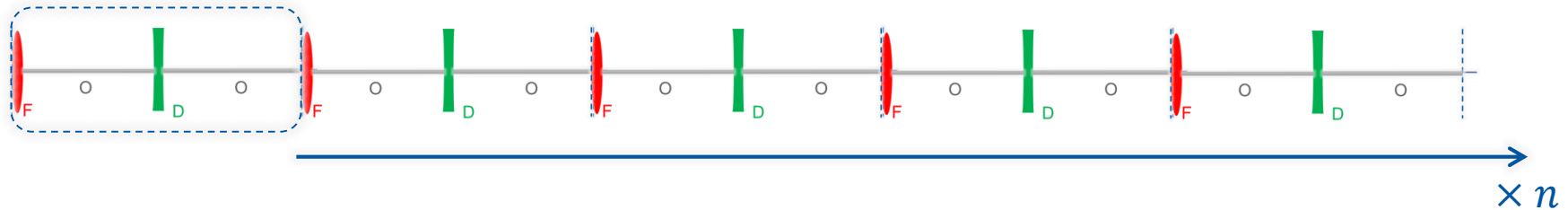
$$FODO = \begin{bmatrix} 1 & 0 \\ -1/f & 1 \end{bmatrix} \begin{bmatrix} 1 & L/2 \\ 0 & 1 \end{bmatrix} \begin{bmatrix} 1 & 0 \\ 1/f & 1 \end{bmatrix} \begin{bmatrix} 1 & L/2 \\ 0 & 1 \end{bmatrix}$$



$$FODO = \begin{bmatrix} 1 + \frac{L}{2f} & L + \frac{L^2}{4f} \\ -\frac{L}{2f^2} & 1 - \frac{L}{2f} - \frac{L^2}{4f^2} \end{bmatrix}$$

# Light transport

Repeating the FODO cell we can focus light from one point to another.



For  $n$  cells:

$$M = FODO^n = \begin{bmatrix} A & B \\ C & D \end{bmatrix}^n$$

Starting and ending to the same refractive index,  $\det(M) = 1$ .

We can apply **Sylvester's theorem** which states:

$$FODO^n = \frac{1}{\sin \theta} \begin{bmatrix} A \sin N \theta - \sin(N-1) \theta & B \sin N \theta \\ C \sin N \theta & D \sin N \theta - \sin(N-1) \theta \end{bmatrix}$$

with  $\cos \theta = \frac{1}{2}(A + D)$   $\xrightarrow{\theta \text{ is real} \leftrightarrow \text{stability}}$   $\left| \frac{1}{2}(A + D) \right| \leq 1$

# Light transport

In a FODO cell:

$$FODO = \begin{bmatrix} 1 + \frac{L}{2f} & L + \frac{L^2}{4f} \\ -\frac{L}{2f^2} & 1 - \frac{L}{2f} - \frac{L^2}{4f^2} \end{bmatrix}$$

The terms  $A$  and  $D$  are:

$$\begin{cases} A = 1 + \frac{L}{2f} \\ D = 1 - \frac{L}{2f} - \frac{L^2}{4f^2} \end{cases}$$

The **stability** of the FODO transport system is ensured if:

$$\left| \frac{1}{2}(A + D) \right| \leq 1 \quad \longrightarrow \quad \left| \left( 1 - \frac{L^2}{8f^2} \right) \right| \leq 1 \quad \longrightarrow \quad \boxed{L \leq 4f}$$

Stability for a  
FODO transport

# Light transport

In a FODO cell:

$$FODO = \begin{bmatrix} 1 + \frac{L}{2f} & L + \frac{L^2}{2f} \\ -\frac{L}{2f^2} & 1 - \frac{L}{2f} \end{bmatrix}$$

The terms  $A$  and  $D$  are:

$$\begin{cases} A = 1 + \frac{L}{2f} \\ D = 1 - \frac{L}{2f} - \frac{L^2}{4f^2} \end{cases}$$

The stability of the FODO

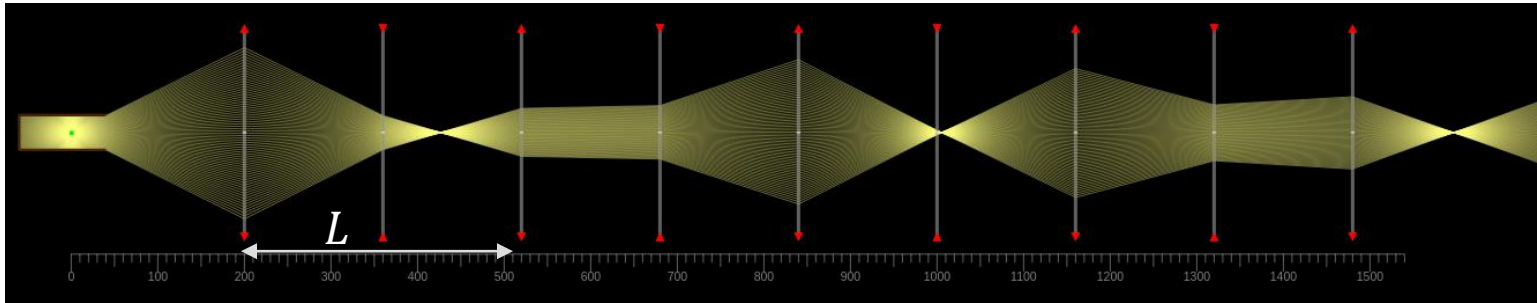
$$\left| \frac{1}{2} (A + D) \right| \leq 1 \quad \rightarrow$$

You will see the same stability condition  
for transverse particle optics in  
A.Latina's lecture next week!

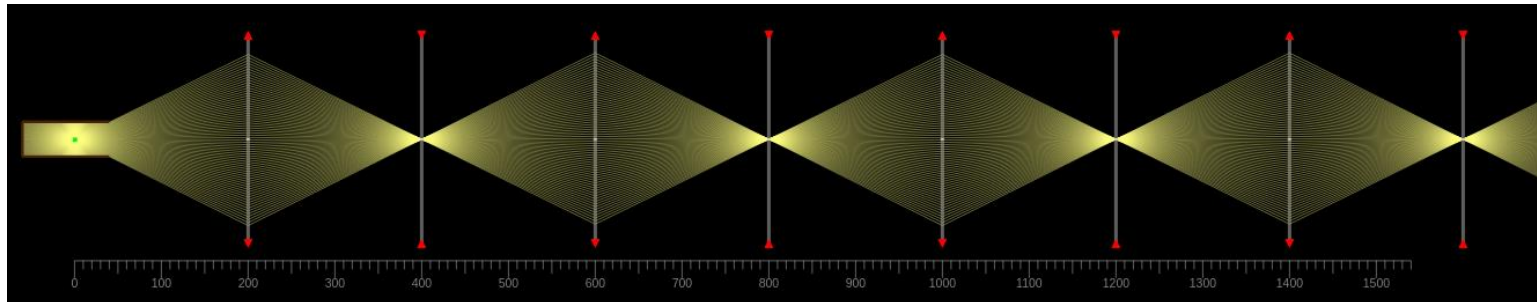
**JUAS - TIMETABLE 2020 - WEEK 2**

Schedule 2020	Monday Jan 20	Tuesday Jan 21	Wednesday Jan 22	Thursday Jan 23	Friday Jan 24
09:00	Transverse Dynamics A. Latina	Transverse Dynamics A. Latina	Transverse Dynamics A. Latina	Cyclotrons B. Jacquot	Linacs D. Alesini
10:00	Coffee Break	Coffee Break	Coffee Break	Coffee Break	Coffee Break
10:15	Transverse Dynamics A. Latina	Transverse Dynamics A. Latina	Transverse Dynamics A. Latina	Cyclotrons B. Jacquot	Linacs D. Alesini
11:15	Transverse Dynamics A. Latina	Transverse Dynamics A. Latina	Transverse Dynamics A. Latina	Cyclotrons B. Jacquot	Linacs D. Alesini
12:15	WORKING LUNCH	BREAK	BREAK	BREAK	BREAK
14:00	Intro. to MAD-X G. Sterbini	Transverse Dynamics A. Latina	Cyclotrons B. Jacquot	Linacs D. Alesini	Transverse Dynamics A. Latina
15:00	MADX N. Fuster Martinez / H. Garcia Morales / A. Latina / G. Sterbini	MADX N. Fuster Martinez / H. Garcia Morales / A. Latina / G. Sterbini	Cyclotrons B. Jacquot	Linacs D. Alesini	Transverse Dynamics A. Latina
16:00	Coffee Break	Coffee Break	Coffee Break	Coffee Break	Coffee Break
16:15	MADX N. Fuster Martinez / H. Garcia Morales / A. Latina / G. Sterbini	MADX N. Fuster Martinez / H. Garcia Morales / A. Latina / G. Sterbini	Cyclotrons B. Jacquot	Linacs D. Alesini	MADX N. Fuster Martinez / H. Garcia Morales / A. Latina / G. Sterbini
17:15	Hybrid collisions in the LHC J. Jowett		European Projects for Collaborative Accelerator R&D Seminar M. Vretenar		MADX N. Fuster Martinez / H. Garcia Morales / A. Latina / G. Sterbini
18:15			AFTER WORK AT ESI		

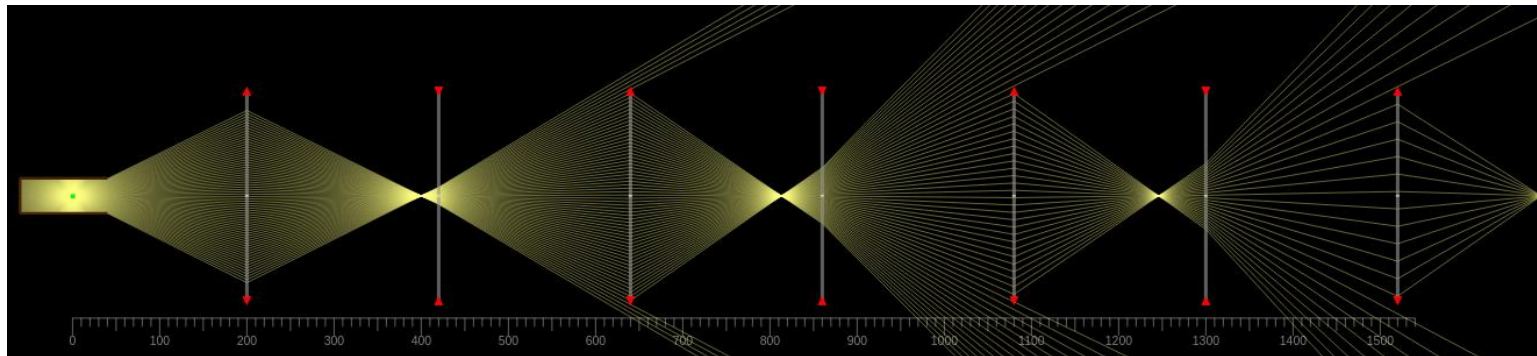
Example of light transport **between focusing/defocusing lenses** with  $f = 100$ .



$$L < 4f$$



$$L = 4f$$

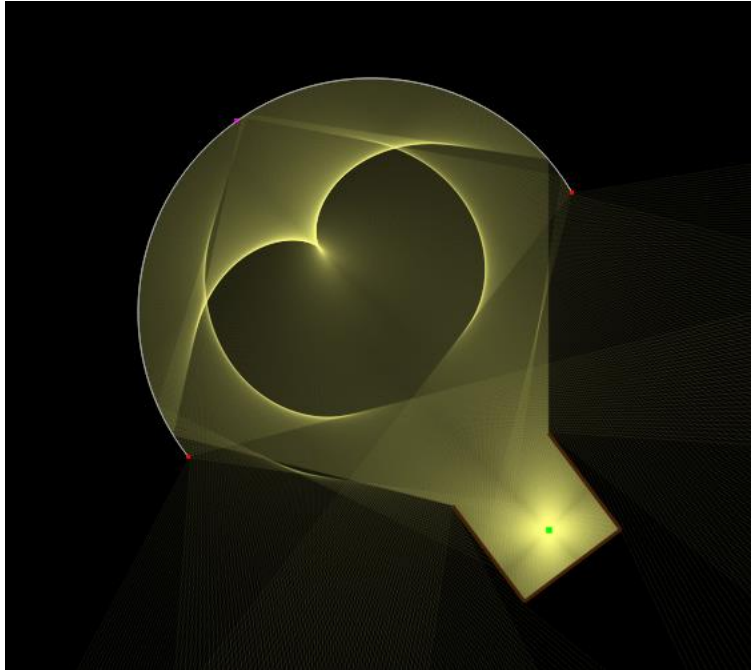


$$L > 4f$$

# The not so ideal world...

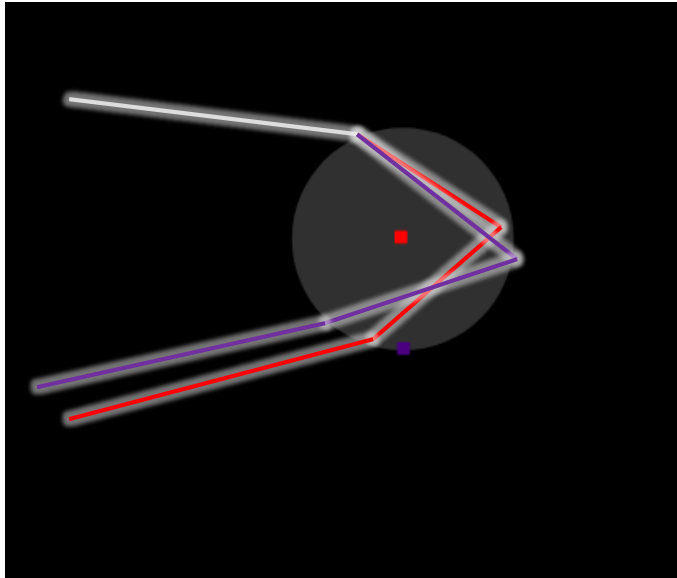


# Spherical aberrations



- In circular mirrors different focal length for large angles smears the focus in a so-called *caustic* line.
- We can use parabolic mirror or additional lens corrections to prevent this.

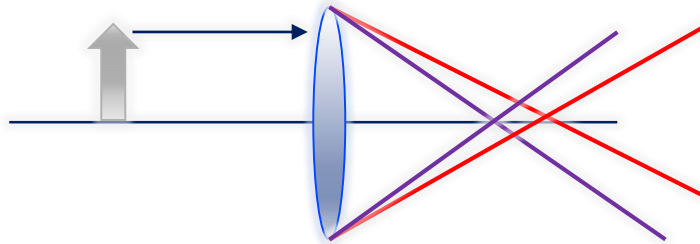
# Chromatic aberrations I



Refractive index depends on wavelength → nicely decomposed in rainbows (primary and secondary depending on number of reflection in rain drops).

# Chromatic aberrations II

Also a thin lens exhibit different focusing depending on wavelength (“chromaticity” of the lens).



Keep “chromaticity” word in mind → key parameter in accelerator design and control.

# Appendix

# Recall of some vector relations

Nabla operator in cartesian coordinates:

$$\nabla = \left[ \frac{d}{dx} \hat{x} + \frac{d}{dy} \hat{y} + \frac{d}{dz} \hat{z} \right]$$

If  $S$  is a vector field and  $\phi$  a scalar function we have:

$$\nabla \cdot (\phi S) = \phi \nabla \cdot S + S^T \nabla \phi$$

The gradient of a vector field is a dyadic:

$$\nabla S = [\nabla S_x \quad \nabla S_y \quad \nabla S_z]$$

Taking the divergence we get back a vector:

$$\nabla \cdot (\nabla f) = \left[ \frac{d}{dx} \hat{x} + \frac{d}{dy} \hat{y} + \frac{d}{dz} \hat{z} \right] \cdot [\nabla S_x \quad \nabla S_y \quad \nabla S_z] = \begin{pmatrix} \nabla \cdot \nabla S_x \\ \text{same with } \nabla S_y \\ \text{same with } \nabla S_z \end{pmatrix}$$

# Eikonal equation derivation

$$\frac{1}{k_0^2} \nabla^2 (\bar{E}_0 e^{-j(k_0 R(\vec{r}))}) + n(\vec{r})^2 \bar{E}_0 e^{-j(k_0 R(\vec{r}))} = 0$$

We need to compute  $\nabla^2 \bar{f} = \nabla \cdot (\nabla \bar{f})$  with  $\bar{f} = \bar{E}_0 e^{-j(k_0 R(\vec{r}))}$

The  $\nabla \bar{f}$  is a dyadic:

$$\begin{aligned} \nabla (\bar{E}_0 e^{-j(k_0 R(\vec{r}))}) &= (-jk_0 \nabla R(\vec{r}) \bar{E}_0 + \nabla \bar{E}_0) e^{-j(k_0 R(\vec{r}))} \\ &= \begin{pmatrix} -jk_0 \nabla R \cdot E_{0,x} + \nabla E_{0,x} \\ \text{same with } \nabla E_{0,y} \\ \text{same with } \nabla E_{0,z} \end{pmatrix} e^{-j(k_0 R(\vec{r}))} \end{aligned}$$

$\nabla \cdot (\nabla \bar{f})$  is then given by

$$\begin{pmatrix} \nabla \cdot ((-jk_0 \nabla R \cdot E_{0,x} + \nabla E_{0,x}) e^{-j(k_0 R(\vec{r}))}) \\ \text{same with } \nabla E_{0,y} \\ \text{same with } \nabla E_{0,z} \end{pmatrix}$$

# Eikonal equation derivation

For the  $x$  component we have:

$$\begin{aligned} \nabla \cdot \left( -jk_0 \nabla R(\vec{r}) E_{0,x} e^{-j(k_0 R(\vec{r}))} + \nabla E_{0,x} e^{-j(k_0 R(\vec{r}))} \right) &= e^{-j(k_0 R(\vec{r}))} \cdot \\ &\cdot \left( -jk_0 \nabla^2 R(\vec{r}) E_{0,x} - jk_0 \nabla E_{0,x} \cdot \nabla R(\vec{r}) - k_0^2 \nabla R(\vec{r}) E_{0,x} \nabla R(\vec{r}) + \nabla^2 E_{0,x} - jk_0 \nabla E_{0,x} \cdot \nabla R(\vec{r}) \right) \end{aligned}$$

And for all three components:

$$\begin{aligned} &\left[ -jk_0 \nabla^2 R(\vec{r}) E_{0,x} - 2jk_0 \nabla E_{0,x} \cdot \nabla R(\vec{r}) + \nabla^2 E_{0,x} - k_0^2 E_{0,x} \nabla R(\vec{r}) \cdot \nabla R(\vec{r}) \right] e^{-j(k_0 R(\vec{r}))} \hat{x} \\ &\left[ -jk_0 \nabla^2 R(\vec{r}) E_{0,y} - 2jk_0 \nabla E_{0,y} \cdot \nabla R(\vec{r}) + \nabla^2 E_{0,y} - k_0^2 E_{0,y} \nabla R(\vec{r}) \cdot \nabla R(\vec{r}) \right] e^{-j(k_0 R(\vec{r}))} \hat{y} \\ &\left[ -jk_0 \nabla^2 R(\vec{r}) E_{0,z} - 2jk_0 \nabla E_{0,z} \cdot \nabla R(\vec{r}) + \nabla^2 E_{0,z} - k_0^2 E_{0,z} \nabla R(\vec{r}) \cdot \nabla R(\vec{r}) \right] e^{-j(k_0 R(\vec{r}))} \hat{z} \end{aligned}$$