

$ar{B} ightarrow X_u \, I \, ar{ u}$ theory

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Outline

- Theoretical framework
- Recent progress
- Critical review of current approaches
- Open problems and future progress

Theoretical framework

• For *B* decays: $5 \text{ GeV} \sim m_b \gg \Lambda_{\text{QCD}} \sim 0.5 \text{ GeV}$ Observables expandable in $\Lambda_{\text{QCD}}/m_b \sim 0.1$

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$$M_X^2 \sim m_b^2$$
local OPE("OPE region") $M_X^2 \sim m_b \Lambda_{QCD}$ Non local OPE("end point region") $M_X^2 \sim \Lambda_{QCD}^2$ No inclusive description("resonance region")Glass (Wave State University) $R \rightarrow X$ (5 theory)

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- The same shape function appears at leading power for $ar{B} o X_s \, \gamma$

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c perturbative, *f* non-perturbative parton distribution function (PDF)

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- At subleading power in Λ_{QCD}/m_b :
- Several subleading shape functions (SSF) appear
- Different linear combinations for $\bar B \to X_u\, /\, \bar \nu$ and $\bar B \to X_s\, \gamma$
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- For $\bar{B} \rightarrow X_s \gamma$ need Effective Hamiltonian

$$\mathcal{H}_{\text{eff}} = \frac{G_F}{\sqrt{2}} \sum_{p=u,c} \lambda_p \left(C_1 Q_1^p + C_2 Q_2^p + \sum_{i=3,...,10} C_i Q_i + C_{7\gamma} Q_{7\gamma} + C_{8g} Q_{8g} \right) + \text{h.c.}$$

- At leading power only $\mathit{Q}_{7\gamma} \mathit{Q}_{7\gamma}$ contribute
- At higher orders need other $Q_i Q_j$ contributions
- Most important: $Q_{7\gamma}, Q_{8g}$, and Q_1

$$Q_{7\gamma} = \frac{-e}{8\pi^2} m_b \bar{s} \sigma_{\mu\nu} (1+\gamma_5) F^{\mu\nu} b$$

$$Q_{8g} = \frac{-g_s}{8\pi^2} m_b \bar{s} \sigma_{\mu\nu} (1+\gamma_5) G^{\mu\nu} b$$

$$Q_1^q = (\bar{q}b)_{V-A} (\bar{s}q)_{V-A} (q=u,c)$$

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• Q_1-Q_{8g} and Q_1-Q_1 give $\Lambda^2_{
m QCD}/m_b^2$ effects

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Top line: Bottom left: Bottom right: $egin{aligned} Q_{7\gamma} - Q_{8g} &\Rightarrow g_{78} \ Q_{8g} - Q_{8g} &\Rightarrow g_{88} \ Q_1 - Q_{7\gamma} &\Rightarrow g_{17} \end{aligned}$

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- Moments of $g_{17} \leftrightarrow \mathsf{HQET}$ parameters. Can we improve those?

Recent progress

Power corrections

• Dimension 7 and 8 HQET operators contribution to $\bar{B} \rightarrow X_c \ell \bar{\nu}_\ell$ [Mannel, Turczyk, Uraltsev JHEP **1011**, 109 (2010)]

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Dimension 7 and 8 HQET operators contribution to B
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Dimension 7 and 8 HQET operators extracted from inclusive B decays
[Gambino, Healey, Turczyk PLB 763, 60 (2016)]

Table 2

Default fit results: the second and third columns give the central values and standard deviations.

m_{b}^{kin}	4.546	0.021	r_1	0.032	0.024
$\overline{m}_{c}(3 \text{ GeV})$	0.987	0.013	r_2	-0.063	0.037
μ_{π}^2	0.432	0.068	r_3	-0.017	0.025
μ_G^2	0.355	0.060	r_4	-0.002	0.025
ρ_D^3	0.145	0.061	r_5	0.001	0.025
$\rho_{LS}^{\bar{3}}$	-0.169	0.097	r_6	0.016	0.025
\overline{m}_1	0.084	0.059	r_7	0.002	0.025
\overline{m}_2	-0.019	0.036	r_8	-0.026	0.025
\overline{m}_3	-0.011	0.045	r_9	0.072	0.044
\overline{m}_4	0.048	0.043	r ₁₀	0.043	0.030
\overline{m}_5	0.072	0.045	<i>r</i> ₁₁	0.003	0.025
\overline{m}_6	0.015	0.041	r ₁₂	0.018	0.025
\overline{m}_7	-0.059	0.043	r ₁₃	-0.052	0.031
\overline{m}_8	-0.178	0.073	r ₁₄	0.003	0.025
\overline{m}_9	-0.035	0.044	r ₁₅	0.001	0.025
χ^2/dof	0.46		r_{16}	0.001	0.025
BR(%)	10.652	0.156	r ₁₇	-0.028	0.025
10 ³ V _{cb}	42.11	0.74	r_{18}	-0.001	0.025

Moments of shape functions and HQET parameters

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- Answer given in [Gunawardana, GP, JHEP 1707 137 (2017)] (See also appendix A of [Heinonen, Mannel, arXiv:1609.01334])
- Method of [Gunawardana, GP, JHEP 1707 137 (2017)] allows to
- 1) Find such relations
- 2) List HQET parameters, in principle, to arbitrary dimension
- 3) Construct NRQED and NRQCD bilinear operators, in principle, to *arbitrary* dimension

• Example: Spin-dependent dimension 8 HQET operators

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$$\begin{split} &\frac{1}{2M_{H}} \langle H|\bar{h}iD^{\mu}iD^{\mu}iD^{\mu}iD^{\mu}iD^{\mu}iD^{\mu}iD^{\mu}iD^{\mu}i\delta^{\mu}i\rho^{\mu}s^{\lambda}h|H\rangle = i\bar{c}^{(8)}v^{\mu}v^{\mu}v^{\mu}v^{\mu}e^{\rho\mu_{1}\mu_{5}\lambda}v_{\rho} \\ &i\bar{s}^{(2)}_{15} \left(v^{\mu}3\Pi^{\mu_{1}\mu_{2}}e^{\rho\mu_{4}\mu_{5}\lambda}v_{\rho} - v^{\mu}3\Pi^{\mu_{4}\mu_{5}}e^{\rho\mu_{2}\mu_{1}\lambda}v_{\rho}\right) + i\bar{s}^{(8)}_{14} \left(v^{\mu}3\Pi^{\mu_{1}\mu_{4}}e^{\rho\mu_{2}\mu_{5}\lambda}v_{\rho} - v^{\mu}3\Pi^{\mu_{5}\mu_{2}}e^{\rho\mu_{4}\mu_{1}\lambda}v_{\rho}\right) + \\ &+i\bar{s}^{(8)}_{15} v^{\mu}3\Pi^{\mu_{1}\mu_{5}}e^{\rho\mu_{2}\mu_{4}\lambda}v_{\rho} + i\bar{s}^{(8)}_{24}v^{\mu}3\Pi^{\mu_{2}\mu_{4}}e^{\rho\mu_{1}\mu_{5}\lambda}v_{\rho} + \\ &+i\bar{b}^{(8)}_{13} \left(v^{\mu}2\Pi^{\mu_{1}\mu_{3}}e^{\rho\mu_{4}\mu_{5}\lambda}v_{\rho} - v^{\mu}4\Pi^{\mu_{5}\mu_{3}}e^{\rho\mu_{2}\mu_{1}\lambda}v_{\rho}\right) + i\bar{b}^{(8)}_{14} \left(v^{\mu}2\Pi^{\mu_{1}\mu_{4}}e^{\rho\mu_{3}\mu_{5}\lambda}v_{\rho} - v^{\mu}4\Pi^{\mu_{5}\mu_{2}}e^{\rho\mu_{3}\mu_{1}\lambda}v_{\rho}\right) + \\ &+i\bar{b}^{(8)}_{15} \left(v^{\mu}2\Pi^{\mu_{1}\mu_{5}}e^{\rho\mu_{3}\mu_{4}\lambda}v_{\rho} - v^{\mu}4\Pi^{\mu_{1}\mu_{5}}e^{\rho\mu_{3}\mu_{2}\lambda}v_{\rho}\right) + i\bar{b}^{(8)}_{34} \left(v^{\mu}2\Pi^{\mu_{3}\mu_{4}}e^{\rho\mu_{1}\mu_{5}\lambda}v_{\rho} - v^{\mu}4\Pi^{\mu_{3}\mu_{2}}e^{\rho\mu_{5}\mu_{1}\lambda}v_{\rho}\right) + \\ &+i\bar{b}^{(8)}_{35} \left(v^{\mu}2\Pi^{\mu_{3}\mu_{5}}e^{\rho\mu_{1}\mu_{4}\lambda}v_{\rho} - v^{\mu}4\Pi^{\mu_{3}\mu_{1}}e^{\rho\mu_{5}\mu_{2}\lambda}v_{\rho}\right) + i\bar{b}^{(8)}_{45} \left(v^{\mu}2\Pi^{\mu_{4}\mu_{5}}e^{\rho\mu_{1}\mu_{3}\lambda}v_{\rho} - v^{\mu}4\Pi^{\mu_{2}\mu_{5}}e^{\rho\mu_{5}\mu_{3}\lambda}v_{\rho}\right) . \end{split}$$

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• Relation to $r_8 - r_{18}$

$$\begin{split} r_8 &= 6 \tilde{c}^{(8)}, \ r_9 &= -6 \left[\tilde{b}^{(8)}_{14} + \tilde{b}^{(8)}_{15} - \tilde{b}^{(8)}_{34} - \tilde{b}^{(8)}_{35} - 3 \tilde{b}^{(8)}_{45} \right], \ r_{10} &= 6 \left[3 \tilde{b}^{(8)}_{13} + \tilde{b}^{(8)}_{14} - \tilde{b}^{(8)}_{15} + \tilde{b}^{(8)}_{34} - \tilde{b}^{(8)}_{35} \right], \\ r_{11} &= 6 \left[\tilde{b}^{(8)}_{13} + 3 \tilde{b}^{(8)}_{14} + \tilde{b}^{(8)}_{15} + \tilde{b}^{(8)}_{34} - \tilde{b}^{(8)}_{45} \right], \ r_{12} &= 6 \left[- \tilde{b}^{(8)}_{13} + \tilde{b}^{(8)}_{15} + \tilde{b}^{(8)}_{34} + 3 \tilde{b}^{(8)}_{35} + \tilde{b}^{(8)}_{45} \right], \\ r_{13} &= -6 \left[\tilde{b}^{(8)}_{13} - \tilde{b}^{(8)}_{14} - 3 \tilde{b}^{(8)}_{15} - \tilde{b}^{(8)}_{35} + \tilde{b}^{(8)}_{45} \right], \ r_{14} &= 6 \left[\tilde{b}^{(8)}_{13} + \tilde{b}^{(8)}_{14} + 3 \tilde{b}^{(8)}_{34} + \tilde{b}^{(8)}_{35} + \tilde{b}^{(8)}_{45} \right], \ r_{15} &= 6 \left[3 \tilde{a}^{(8)}_{12} - \tilde{a}^{(8)}_{15} + 3 \tilde{a}^{(8)}_{24} \right], \\ r_{16} &= 6 \left[-2 \tilde{a}^{(8)}_{12} + 2 \tilde{a}^{(8)}_{14} + 3 \tilde{a}^{(8)}_{15} \right], \ r_{17} &= 6 \left[2 \tilde{a}^{(8)}_{12} + 2 \tilde{a}^{(8)}_{14} + 3 \tilde{a}^{(8)}_{24} \right], \ r_{18} &= 6 \left[3 \tilde{a}^{(8)}_{14} + \tilde{a}^{(8)}_{15} + \tilde{a}^{(8)}_{24} \right]. \end{split}$$
• Example: Spin-dependent dimension 8 HQET operators

$$\begin{split} &\frac{1}{2M_{H}} \langle H|\bar{h}\,iD^{\mu_{1}}\,iD^{\mu_{2}}\,iD^{\mu_{3}}\,iD^{\mu_{4}}\,iD^{\mu_{5}}\,s^{\lambda}h|H\rangle = i\bar{c}^{(8)}v^{\mu_{2}}v^{\mu_{3}}v^{\mu_{4}}\epsilon^{\rho\mu_{1}\mu_{5}\lambda}v_{\rho} \\ &i\bar{s}_{12}^{(8)}\left(v^{\mu_{3}}\Pi^{\mu_{1}\mu_{2}}\epsilon^{\rho\mu_{4}\mu_{5}\lambda}v_{\rho} - v^{\mu_{3}}\Pi^{\mu_{4}\mu_{5}}\epsilon^{\rho\mu_{2}\mu_{1}\lambda}v_{\rho}\right) + i\bar{s}_{14}^{(8)}\left(v^{\mu_{3}}\Pi^{\mu_{1}\mu_{4}}\epsilon^{\rho\mu_{2}\mu_{5}\lambda}v_{\rho} - v^{\mu_{3}}\Pi^{\mu_{5}\mu_{2}}\epsilon^{\rho\mu_{4}\mu_{1}\lambda}v_{\rho}\right) + \\ &+i\bar{s}_{15}^{(8)}v^{\mu_{3}}\Pi^{\mu_{1}\mu_{5}}\epsilon^{\rho\mu_{2}\mu_{4}\lambda}v_{\rho} + i\bar{s}_{24}^{(8)}v^{\mu_{3}}\Pi^{\mu_{2}\mu_{4}}\epsilon^{\rho\mu_{1}\mu_{5}\lambda}v_{\rho} + \\ &+i\bar{b}_{15}^{(8)}\left(v^{\mu_{2}}\Pi^{\mu_{1}\mu_{3}}\epsilon^{\rho\mu_{4}\mu_{5}\lambda}v_{\rho} - v^{\mu_{4}}\Pi^{\mu_{5}\mu_{3}}\epsilon^{\rho\mu_{2}\mu_{1}\lambda}v_{\rho}\right) + i\bar{b}_{14}^{(8)}\left(v^{\mu_{2}}\Pi^{\mu_{1}\mu_{4}}\epsilon^{\rho\mu_{3}\mu_{5}\lambda}v_{\rho} - v^{\mu_{4}}\Pi^{\mu_{5}\mu_{2}}\epsilon^{\rho\mu_{3}\mu_{1}\lambda}v_{\rho}\right) + \\ &+i\bar{b}_{15}^{(8)}\left(v^{\mu_{2}}\Pi^{\mu_{1}\mu_{5}}\epsilon^{\rho\mu_{3}\mu_{4}\lambda}v_{\rho} - v^{\mu_{4}}\Pi^{\mu_{3}\mu_{1}}\epsilon^{\rho\mu_{5}\mu_{2}\lambda}v_{\rho}\right) + i\bar{b}_{45}^{(8)}\left(v^{\mu_{2}}\Pi^{\mu_{4}\mu_{5}}\epsilon^{\rho\mu_{1}\mu_{3}\lambda}v_{\rho} - v^{\mu_{4}}\Pi^{\mu_{2}\mu_{1}}\epsilon^{\rho\mu_{5}\mu_{3}\lambda}v_{\rho}\right) + \\ &-i\bar{b}_{35}^{(8)}\left(v^{\mu_{2}}\Pi^{\mu_{3}\mu_{5}}\epsilon^{\rho\mu_{1}\mu_{4}\lambda}v_{\rho} - v^{\mu_{4}}\Pi^{\mu_{3}\mu_{1}}\epsilon^{\rho\mu_{5}\mu_{2}\lambda}v_{\rho}\right) + i\bar{b}_{45}^{(8)}\left(v^{\mu_{2}}\Pi^{\mu_{4}\mu_{5}}\epsilon^{\rho\mu_{1}\mu_{3}\lambda}v_{\rho} - v^{\mu_{4}}\Pi^{\mu_{2}\mu_{1}}\epsilon^{\rho\mu_{5}\mu_{3}\lambda}v_{\rho}\right) + \\ &-i\bar{b}_{35}^{(8)}\left(v^{\mu_{2}}\Pi^{\mu_{3}\mu_{5}}\epsilon^{\rho\mu_{1}\mu_{4}\lambda}v_{\rho} - v^{\mu_{4}}\Pi^{\mu_{3}\mu_{1}}\epsilon^{\rho\mu_{5}\mu_{2}\lambda}v_{\rho}\right) + i\bar{b}_{45}^{(8)}\left(v^{\mu_{2}}\Pi^{\mu_{4}\mu_{5}\lambda}v_{\rho} - v^{\mu_{4}}\Pi^{\mu_{2}\mu_{1}}\epsilon^{\rho\mu_{5}\mu_{3}\lambda}v_{\rho}\right) + i\bar{b}_{45}^{(8)}\left(v^{\mu_{2}}\Pi^{\mu_{4}\mu_{5}\lambda}v_{\rho} - v^{\mu_{4}}\Pi^{\mu_{2}\mu_{1}}\epsilon^{\rho\mu_{5}\mu_{3}\lambda}v_{\rho}\right) + i\bar{b}_{45}^{(8)}\left(v^{\mu_{2}}\Pi^{\mu_{4}\mu_{5}\lambda}v_{\rho} - v^{\mu_{4}}\Pi^{\mu_{2}\mu_{1}}v_{\rho}^{\mu_{5}\mu_{3}\lambda}v_{\rho}\right) + i\bar{b}_{45}^{(8)}\left(v^{\mu_{2}}\Pi^{\mu_{4}\mu_{5}\lambda}v_{\rho} - v^{\mu_{4}}\Pi^{\mu_{2}\mu_{1}}v_{\rho}^{\mu_{5}\mu_{3}\lambda}v_{\rho}\right) + i\bar{b}_{45}^{(8)}\left(v^{\mu_{2}}\Pi^{\mu_{4}\mu_{5}\lambda}v_{\rho} - v^{\mu_{4}}\Pi^{\mu_{2}\mu_{2}\lambda}v_{\rho}^{\mu_{5}\mu_{5}\lambda}v_{\rho}\right) + i\bar{b}_{45}^{(8)}\left(v^{\mu_{2}}\Pi^{\mu_{4}\mu_{5}\lambda}v_{\rho} - v^{\mu_{4}}\Pi^{\mu_{2}\mu_{5}\lambda}v_{\rho}\right) + i\bar{b}_{45}^{(8)}\left(v^{\mu_{2}}\Pi^{\mu_{4}\mu_{5}\lambda}v_{\rho} - v^{\mu_{4}}\Pi^{\mu_{5}\mu_{5}\lambda}v_{\rho}\right) + i\bar{b}_{45}^{(8)}\left(v^{\mu_{4}\Pi^{\mu_{5}\mu_{5}\lambda}v_{$$

• Relation to $r_8 - r_{18}$

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• Moments in ω and ω_1 are related to HQET parameters

$$\begin{split} \langle \omega^{l} \, \omega_{1}^{k} \, g_{17} \rangle &\equiv \int_{-\infty}^{\bar{h}} d\omega \, \omega^{l} \int_{-\infty}^{\infty} d\omega_{1} \, \omega^{k} \, g_{17}(\omega, \omega_{1}, \mu) = \left(iv^{\rho} \, \epsilon_{\rho \mu \alpha_{\perp} \lambda} \bar{n}^{\mu} - g_{\alpha_{\perp} \lambda} \right) (-1)^{k} \\ &\times \qquad \frac{1}{2M_{B}} \langle \bar{B} | \bar{h} \, (in \cdot D)^{l} \underbrace{\left[i\bar{n} \cdot D, \left[i\bar{n} \cdot D, \cdots \left[i\bar{n} \cdot D \right], \cdots \left[i\bar{n} \cdot D \right] \cdots \right] \right] s^{\lambda} h | \bar{B} \rangle. \\ & \qquad k \text{ times} \end{split}$$

Gil Paz (Wayne State University)

 $\bar{B} \to X_{\mu} I \bar{\nu}$ theory

Numerical Results: moments of g_{17}

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- Using different models, including some Λ_{QCD}^2/m_b^2 corrections and larger m_c range, a smaller reduction was found in [Benzke, Hurth PRD **102** 114024 (2020)]

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What about V_{ub} ?

- Above information not applied yet to the photon spectrum but it can be done
- Moments of g_{78} and g_{88} are *not* related to to these HQET parameters. Very little is known about them, in particular, their E_{γ} dependence
- It is not clear if such modeling of g_{78} and g_{88} is better than assuming a $\Lambda_{\rm QCD}/m_b$ uncertainty on the extraction of the leading shape function from $\bar{B} \to X_s \gamma$

Critical review of current approaches

• We need to distinguish between incorrect statements and differences in what is "reasonable"

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- I'll try to distinguish the two in the following

BLNP

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- Based on $d\Gamma \sim H \cdot J \otimes S + rac{1}{m_b} \sum_{\cdot} H \cdot J \otimes s_i + ...$
- Leading power $H \cdot J \otimes S$ at $\mathcal{O}(\alpha_s)$
- Subleading shape functions: $H \cdot J \otimes s_i$ at $\mathcal{O}(\alpha_s^0)$
- $\alpha_s/m_b^{\{1,2\}}$ and $1/m_b^2$ incorporated as naive convolution with LO SF
- *S* extracted from $\bar{B} \rightarrow X_s \gamma$, s_i modeled (\sim 700 models)
- Smoothly merges to local OPE when integrated over phase space
- Hard, Jet, and Soft scale separated with NLO resummation

$$d\Gamma \sim H \cdot J \otimes S + \frac{1}{m_b} \sum_i H \cdot J \otimes s_i + \dots$$

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 Resummation: 2009 Study of implication of O(α²_s) on |V_{ub}| [Greub, Neubert, Pecjak, EPJC **65** 501 (2010)]
 "factorization ... perturbative coefficient...into jet and hard functions is not strictly necessary: using ... fixed-order... does not lead to large scale uncertainties ... nor to a poor convergence ..."

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3) S has a negative radiative tail for large ω that is "glued": not very elegant

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- S has a negative radiative tail for large ω that is "glued": not very elegant
- Uses the shape function mass scheme, not easy to switch to other schemes. e.g. kinetic

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- A major difference from BLNP is the treatment of the leading shape function using

$$S(\omega,\mu_{\Lambda})=\int dk \ C_0(\omega-k,\mu_{\Lambda})F(k)$$

"where $C_0(\omega, \mu_{\Lambda})$ is the *b* quark matrix element of the shape function operator calculated in perturbation theory, and F(k) is a nonperturbative function that can be extracted from data." [Ligeti, Stewart, Tackmann PRD **78** 114014 (2008)]

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- "The advantage of our construction [...] is that the tail automatically turns on in a smooth manner when it dominates over the nonperturbative function *F*(*k*) and provides the proper μ dependence for *S*(ω, μ) at any ω."
 [Ligeti, Stewart, Tackmann PRD **78** 114014 (2008)]
- The same paper also suggested to use orthonormal basis functions Among other things it allows to better fit $\bar{B} \to X_s \gamma$ photon spectrum

SIMBA: Possible issues: Positivity

1) The same paper also introduced

$$S(\omega) = \int dk \ C_0(\omega - k)F(k) = \int dk \ \widehat{C}_0(\omega - k) \ \widehat{F}(k)$$

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- From [Gunawardana, GP '17] and [Gambino, Healey, Turczyk '16)] $\int d\omega S(\omega) = 1, \int d\omega \omega S(\omega) = 0, \int d\omega \omega^2 S(\omega) = \mu_{\pi}^2/3 = 0.144 \pm 0.023 \text{ GeV}^2,$ $\int d\omega \omega^3 S(\omega) = -\rho_D^3/3 = -0.048 \pm 0.020 \text{ GeV}^3,$ $\int d\omega \omega^4 S(\omega) = m_1/5 - m_2/3 = 0.023 \pm 0.017 \text{ GeV}^4,$ $\int d\omega \omega^5 S(\omega) = (-8r_1 + 2r_2 + 2r_3 + 2r_4 + r_5 + r_6 + r_7)/15 = -0.027 \pm 0.015 \text{ GeV}^5$
 - Starting with the third, the moments are not positive

2) "We expect on physical grounds that $\widehat{F}(k)$ is positive, so we can expand its square root.."

$$\widehat{F}^{\mathrm{mod}}(k) = \frac{1}{\lambda} \left[\sum_{n=0}^{2} c_n f_n\left(\frac{k}{\lambda}\right) \right]^2$$

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Figure 1: The variation of F against k; red curve represent combination $\{{\rm c0}\to 0.949031, {\rm c1}\to -0.308648, {\rm c2}\to 0.0638516\}$

[Image by A. Gunawardana, private communication]

Gil Paz (Wayne State University)

 $\bar{B} \to X_{\mu} \, l \, \bar{\nu}$ theory

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Figure 2: The variation of F against k; blue curve represent combination {c0 \rightarrow 0.988539, c1 \rightarrow -0.115211, c2 \rightarrow -0.0975604} [Image by A. Gunawardana, private communication] Gil Paz (Wayne State University) $\bar{B} \rightarrow X_{\mu} l \bar{\nu}$ theory

SIMBA: Possible issues: Small momentum behavior

3) "Due to the short distance subtractions ... to ensure that $S(\omega, \mu)$ goes to zero at $\omega = 0$, we need $\widehat{F}(k)$ to go to zero at least as k^3 for $k \to 0$ "

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- 3) "Due to the short distance subtractions ... to ensure that $S(\omega, \mu)$ goes to zero at $\omega = 0$, we need $\widehat{F}(k)$ to go to zero at least as k^3 for $k \to 0$ "
 - If we want to include higher moments, we might need to increase the power of k [B. Lange, private communication]

GGOU

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$$W_i \sim F_i \otimes W_i^{pert}$$

- W_i structure functions that give $d\Gamma$
- W_i^{pert} known perturbative quantities
- $F_i(k_+, q^2, \mu)$ OPE-constrained non-perturbative distribution functions
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- *F_i* moments are constrained by OPE About 100 forms considered in GGOU Each parameterized by simple 2-parameter functional forms [Gambino, CKM 2016 talk]

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- I am not familiar with factorization formula that shows such a symmetry for all power corrections

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The future looks promising for $\bar{B} \rightarrow X_u \, I \, \bar{\nu}$ and inclusive $|V_{ub}|!$