# $B \rightarrow X_{c} \gamma$ with Neural Networks

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<sup>†</sup>Progress report of work done in collaboration with P. Gambino, M. Misiak, S. Schacht

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### Motivations

- $B \to X_{s}\gamma$ 
  - OPE at NNLO for  $E_{\gamma}^{\text{th}} > 1.6 \text{ GeV}$

  - Spectrum at NNLO: depends on a the (mostly) unknown B-meson Shape Function • Experimental data requires  $E_{\nu}^{exp} > [1.6-1.9]$  GeV (most precise at large  $E_{\nu}$ )
- Fits to the spectrum allow to
  - already have information from  $B \to X_c \ell \nu$ )
  - improve the extrapolation down to 1.6 GeV extract information on m<sub>h</sub> and various hadronic matrix elements (on which we
  - extract the Shape Function itself
- Knowledge of the Shape function is essential to  $B \to X_u \ell \nu$  and to the  $m_{X_c}$  extrapolation in  $B \to X_s \ell \ell$

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- In this talk the CLN parameterization will not be mentioned

- We work in the kinetic Scheme for the b mass and all hadronic parameters:  $m_b$ ,  $\mu_{\pi}^2$ ,  $\mu_G^2$ ,  $\rho_D^3$ ,  $\rho_{IS}^3$  and the Shape Function
- Introduce a Wilsonian cutoff  $\mu \sim 1 \text{ GeV}$ 

  - mentioned non-perturbative parameters and in the Shape Functions
- This approach is rigorous and has been used to the
  - construction of the widely used kinetic scheme for  $m_b$  and other hadronic parameters from various small-velocity  $B \rightarrow X_{u,c} \ell \nu$  sum-rules
  - calculation of moments of the  $B \rightarrow X_{c} \gamma$  spectrum
  - extraction of  $V_{cb}$ ,  $m_b$  and other hadronic parameters from  $B \to X_c \ell \nu$
  - extraction of  $V_{\mu b}$  from  $B \to X_{\mu} \ell \nu$

## $B \rightarrow X_{c} \gamma$ with a Wilsonian cutoff

• Virtual and real gluons with energy larger than  $\mu$  are calculated in perturbation theory • Gluons with  $E_g < \mu$  are intrinsically non-perturbative and are included via the above

### References

- Wilsonian cutoff in B physics and kinetic scheme: Bigi, Shifman, Uraltsev, Vainshtein, hep-h/9312359 Bigi, Shifman, Uraltsev, Vainshtein, hep-ph/9402360 Bigi, Shifman, Uraltsev, Vainshtein, hep-ph/9405410 Bigi, Shifman, Uraltsev, Vainshtein, hep-ph/9312359 Uraltsev, hep-ph/9610425 Bigi, Shifman, Uraltsev, hep-ph/9703290 Uraltsev, hep-ph/0010328
- $B \rightarrow X_{c}\gamma$  moments: Bigi, Uraltsev, hep-ph/0308165 Benson, Bigi, Uraltsev, hep-ph/041080
- $B \to X_c \ell \nu$  fits: Benson, Bigi, Mannel, Uraltsev, hep-ph/0302262
- $B \to X_{\mu} \ell \nu$  fits: Gambino, Giordano, Ossola, Uraltsev, 0707.2493

• We write the spectrum as ( $\mu \simeq 1 \text{ GeV}$  is the Wilsonian cutoff):

$$\frac{d\Gamma}{dE_{\gamma}} = \int dk_{+} f(k_{+}, \mu) \frac{d\Gamma^{pert}}{dE_{\gamma}} \left( E_{\gamma} - \frac{k_{+}}{2}, \mu \right)$$
$$= \Gamma_{0} \sum_{i \le j=1}^{8} C_{i}^{\text{eff}*}(\mu_{b}) C_{j}^{\text{eff}}(\mu_{b}) \int_{-\infty}^{\lambda} d\kappa F(\kappa, \mu) \nabla_{j}^{k} d\kappa F(\kappa, \mu) \nabla_{j}^{k} d\kappa F(\kappa, \mu)$$

where 
$$F(\kappa, \mu) = m_b f(m_b \kappa, \mu)$$
  
 $m_b = m_b^{\text{kin}}(\mu)$   
 $\xi = 2E_{\gamma}/m_b$   
 $\lambda = (m_B - m_b)/m_b$   
 $\Gamma_0 = \frac{G_F^2 \alpha m_b^2 m_b^{\overline{\text{MS}}}(\mu_b)^2}{16\pi^4} |V_{tb}V_{ts}^*|^2$ 

 $W_{ij}^{pert}(\xi - \kappa, \mu, \mu_b)$ 

Shape Function in the kinetic scheme

• For instance, the  $|C_7|^2$  hard scattering kernel has the following structure:

$$W_{77}^{pert}(\xi,\mu,\mu_b) = \left[1 + \frac{\alpha_s(\mu_b)}{\pi} C_F\left(V^{(1)} + \beta_0 \frac{\alpha_s}{\pi} V^{(2,\text{BLM})} + \frac{\alpha_s(\mu_b)}{\pi} C_F\left[B^{(1)} + \beta_0 \frac{\alpha_s}{\pi} B^{(2,\text{BLM})} + \frac{1}{2} \frac{\alpha_s(\mu_b)}{\pi} C_F\left[C^{(1)} + \beta_0 \frac{\alpha_s}{\pi} C^{(2,\text{BLM})} + \frac{\alpha_s(\mu_b)}{\pi} C_F\left[H^{(1)}(\xi,\mu) + \beta_0 \frac{\alpha_s}{\pi} D_F^{(2,\text{BLM})}\right]\right]$$

moment matches the OPE result

virtual corrections ( $E_g > \mu$ ): 0th moment (total rate) 
$$\begin{split} & H^{(2,\text{BLM})}(\xi,\mu) + \frac{\alpha_s}{\pi} V^{(2)} \end{pmatrix} \bigg] \delta(1-\xi) \\ & \text{1st moment} \\ & \text{1st moment} \\ & \text{1st moment} \\ & \text{2nd moment} \\ & \text{gluon Bremsstrahlung} (E_g > \mu) \\ & H^{(2,\text{BLM})}(\xi,\mu) + \frac{\alpha_s}{\pi} H^{(2)}(\xi) \bigg] \theta(1-\xi) \theta(\xi) \end{split}$$

• Virtual contributions can be either calculated directly or extracted by requiring that each









- Soft-gluon contributions (with  $E_g < 1 \text{ GeV}$ ) have been absorbed into the SF
- For instance:

$$\int d\xi \int d\kappa F(\kappa) W_{77}^{pert}(\xi - \kappa) = \int d\kappa F(\kappa) \int d\xi W_{77}^{pert}(\xi - \kappa) = \int d\kappa F(\kappa) \int d\xi' W_{77}^{pert}(\xi')$$
  
total rate at NNLO 
$$= 1 + O(1/m_b^2)$$

conditions for the moments of the Shape Function. For instance, we find:

$$\int d\kappa F(\kappa,\mu) = 1 - \frac{\mu_{\pi}^2 + 3\mu_G^2}{2m_b^2} - \frac{11\rho_D^3 - 9\rho_{LS}^3}{6m_b^3} + \frac{\alpha_s}{\pi} \left[ C_F \left(\frac{4\eta}{3} + \frac{3\eta^2}{4} + \frac{11\eta^3}{18}\right) \frac{\mu_{\pi}^2}{m_b^2} + \left(\frac{67}{36} - \frac{26\pi^2}{27} - \frac{9L_b}{4} + \frac{16\eta}{3} + 3\eta^2 + \frac{22\eta^3}{9}\right) \frac{\mu_G^2}{m_b^2} \right]$$

• The moments of the photon spectrum depend only on corresponding moments of the SF.

• Power corrections in the OPE for the moments are reproduced by appropriate matching

- We include effects of all other operators in a similar way.
- For instance the 78 kernel is:

$$\begin{split} W_{78}^{pert}(\xi,\mu,\mu_b) &= C_F \frac{\alpha_s(\mu_b)}{\pi} \Big[ \Big( V_{78}^{(1)} + \frac{\alpha_s}{\pi} V_{78}^{(2)} \Big) \delta(1-\xi) + \frac{\alpha_s}{\pi} B_{78}^{(2)} \delta'(1-\xi) + \frac{1}{2} \frac{\alpha_s}{\pi} C_{78}^{(2)} \delta''(1-\xi) \\ &+ \Big( H_{78}^{(1)}(\xi,\mu) + \frac{\alpha_s}{\pi} H_{78}^{(2)}(\xi,\mu) \Big) \theta(1-\xi) \theta(\xi) \Big] \end{split}$$

The spectrum is obtained by convoluting this kernel with same Shape Function introduced for 77 because soft and collinear end-point singularities factorize

• For all other kernels we have only BLM contributions at NNLO:

$$W_{ij}^{pert}(\xi,\mu,\mu_b) = C_F \frac{\alpha_s(\mu_b)}{\pi} \Big[ \Big( V_{ij}^{(1)} + \beta_0 \frac{\alpha_s}{\pi} V_{ij}^{(2BLM)} \Big) \delta(1-\xi) + \left( H_{ij}^{(1)}(\xi,\mu) + \beta_0 \frac{\alpha_s}{\pi} H_{ij}^{(2BLM)}(\xi,\mu) \right) \theta(1-\xi) + \left( H_{ij}^{(1)}(\xi,\mu) + \beta_0 \frac{\alpha_s}{\pi} H_{ij}^{(2BLM)}(\xi,\mu) \right) \theta(1-\xi) + \left( H_{ij}^{(1)}(\xi,\mu) + \beta_0 \frac{\alpha_s}{\pi} H_{ij}^{(2BLM)}(\xi,\mu) \right) \theta(1-\xi) + \left( H_{ij}^{(1)}(\xi,\mu) + \beta_0 \frac{\alpha_s}{\pi} H_{ij}^{(2BLM)}(\xi,\mu) \right) \theta(1-\xi) + \left( H_{ij}^{(1)}(\xi,\mu) + \beta_0 \frac{\alpha_s}{\pi} H_{ij}^{(2BLM)}(\xi,\mu) \right) \theta(1-\xi) + \left( H_{ij}^{(1)}(\xi,\mu) + \beta_0 \frac{\alpha_s}{\pi} H_{ij}^{(2BLM)}(\xi,\mu) \right) \theta(1-\xi) + \left( H_{ij}^{(1)}(\xi,\mu) + \beta_0 \frac{\alpha_s}{\pi} H_{ij}^{(2BLM)}(\xi,\mu) \right) \theta(1-\xi) + \left( H_{ij}^{(1)}(\xi,\mu) + \beta_0 \frac{\alpha_s}{\pi} H_{ij}^{(2BLM)}(\xi,\mu) \right) \theta(1-\xi) + \left( H_{ij}^{(1)}(\xi,\mu) + \beta_0 \frac{\alpha_s}{\pi} H_{ij}^{(2BLM)}(\xi,\mu) \right) \theta(1-\xi) + \left( H_{ij}^{(1)}(\xi,\mu) + \beta_0 \frac{\alpha_s}{\pi} H_{ij}^{(2BLM)}(\xi,\mu) \right) \theta(1-\xi) + \left( H_{ij}^{(1)}(\xi,\mu) + \beta_0 \frac{\alpha_s}{\pi} H_{ij}^{(2BLM)}(\xi,\mu) \right) \theta(1-\xi) + \left( H_{ij}^{(1)}(\xi,\mu) + \beta_0 \frac{\alpha_s}{\pi} H_{ij}^{(2BLM)}(\xi,\mu) \right) \theta(1-\xi) + \left( H_{ij}^{(1)}(\xi,\mu) + \beta_0 \frac{\alpha_s}{\pi} H_{ij}^{(2BLM)}(\xi,\mu) \right) \theta(1-\xi) + \left( H_{ij}^{(1)}(\xi,\mu) + \beta_0 \frac{\alpha_s}{\pi} H_{ij}^{(2BLM)}(\xi,\mu) \right) \theta(1-\xi) + \left( H_{ij}^{(1)}(\xi,\mu) + \beta_0 \frac{\alpha_s}{\pi} H_{ij}^{(2BLM)}(\xi,\mu) \right) \theta(1-\xi) + \left( H_{ij}^{(1)}(\xi,\mu) + \beta_0 \frac{\alpha_s}{\pi} H_{ij}^{(2BLM)}(\xi,\mu) \right) \theta(1-\xi) + \left( H_{ij}^{(1)}(\xi,\mu) + \beta_0 \frac{\alpha_s}{\pi} H_{ij}^{(2BLM)}(\xi,\mu) \right) \theta(1-\xi) + \left( H_{ij}^{(1)}(\xi,\mu) + \beta_0 \frac{\alpha_s}{\pi} H_{ij}^{(2BLM)}(\xi,\mu) \right) \theta(1-\xi) + \left( H_{ij}^{(1)}(\xi,\mu) + \beta_0 \frac{\alpha_s}{\pi} H_{ij}^{(2BLM)}(\xi,\mu) \right) \theta(1-\xi) + \left( H_{ij}^{(1)}(\xi,\mu) + \beta_0 \frac{\alpha_s}{\pi} H_{ij}^{(2BLM)}(\xi,\mu) \right) \theta(1-\xi) + \left( H_{ij}^{(1)}(\xi,\mu) + \beta_0 \frac{\alpha_s}{\pi} H_{ij}^{(2BLM)}(\xi,\mu) \right) \theta(1-\xi) + \left( H_{ij}^{(1)}(\xi,\mu) + \beta_0 \frac{\alpha_s}{\pi} H_{ij}^{(2BLM)}(\xi,\mu) \right) \theta(1-\xi) + \left( H_{ij}^{(1)}(\xi,\mu) + \beta_0 \frac{\alpha_s}{\pi} H_{ij}^{(2BLM)}(\xi,\mu) \right) \theta(1-\xi) + \left( H_{ij}^{(1)}(\xi,\mu) + \beta_0 \frac{\alpha_s}{\pi} H_{ij}^{(2BLM)}(\xi,\mu) \right) \theta(1-\xi) + \left( H_{ij}^{(1)}(\xi,\mu) + \beta_0 \frac{\alpha_s}{\pi} H_{ij}^{(2BLM)}(\xi,\mu) \right) \theta(1-\xi) + \left( H_{ij}^{(1)}(\xi,\mu) + \beta_0 \frac{\alpha_s}{\pi} H_{ij}^{(2BLM)}(\xi,\mu) \right) \theta(1-\xi) + \left( H_{ij}^{(1)}(\xi,\mu) + \beta_0 \frac{\alpha_s}{\pi} H_{ij}^{(2BLM)}(\xi,\mu) \right) \theta(1-\xi) + \left( H_{ij}^{(1)}(\xi,\mu) + \beta_0 \frac{\alpha_s}{\pi} H_{ij}^{(2$$

 Resolved contributions are not present in this formalism and need to be taken into account separately [Lee, Neubert, Pax, hep-ph/0609224]

Korchemski and Sterman, hep-ph/9902341 Akhoury and Rothstein, hep-ph/9512303



• Shape Function vs Hard Scattering spectra:



# Shape Function with a Neural Network

- We consider a NN with one input ( $\xi^{(0)}$ ), one layer with N nodes ( $\xi_i^{(1)}$ ) and a single output ( $\xi^{(2)}$ )



• g is a non-linear activation function. We adopt a sigmoid:  $g(x) = -\frac{1}{2}$ 

[Gambino, Healey, Mondino, 1604.07598]

Neural Networks can be used to provide unbiased parameterizations of continuous functions





# Shape Function with a Neural Network

- help speeding the training without (possibly) introducing a strong bias
- explicit cut-off ( $\kappa < \lambda$ )
- For instance we consider:

 $F(\kappa) = \bar{w}_1 e^{\bar{w}_2 \kappa} (\lambda - \kappa)^{\bar{w}_3} \quad N(\{w, \theta\}, \kappa) \quad \theta(\lambda - \kappa)$ 

• The actual Shape Function we consider contains an *underlying function* whose purpose is to

Finally we impose few further conditions: positivity of the overall Shape Function and an

• The extra parameters  $\bar{w}_i$  (with  $\bar{w}_1 > 0$ ) are treated on equal footing with the other weights

• In the preliminary tests we have run we considered N=6, implying 22 adjustable weights





# Training

- We adopt a genetic algorithm
- Each full iteration begins with a "parent" SF with random weights and ends with a "child" SF for which a  $\chi^2$  comprised of experimental (BaBar,Belle) and theoretical (SF moments) information is below a given threshold.
- Each successful iteration yields a replica: a possible Shape Function which is fully compatible with both experimental and theoretical constraints.
- There are many (important) details related to • **overtraining** (e.g. divide  $\chi^2$  into training and validation)

  - avoiding local minima
- Distribution of replicas is not expected to be strongly affected/biased by minor changes in the chosen training scheme



- Training a Neural Network requires an enormous number of  $\chi^2$  evaluations
- Binned experimental results are presented in the B and  $\Upsilon(4S)$  rest frames and require multiple numerical integrations
- An excellent solution is approximating the NN Shape Function with a cubic spline.
- We found that dividing the  $\kappa \in [-1,\lambda]$  into 50 intervals (with  $4 \times 50 = 200$  spline coefficients) yields 0.1~% approximations to realistic Shape Functions and to their corresponding spectra and integrated observables
- All observables can be easily calculated for each of the 200 basic spline building blocks. The  $\chi^2$  can then be easily obtained by contracting the latter with the 200 spline coefficients

# $\chi^2$ calculation



- An alternative approach we considered is projecting the NN Shape Function onto a (hopefully) small set of basis functions

$$F(\kappa) = \sum_{n=0}^{\infty} c_n f_n (1 - \kappa/\lambda) \qquad Y(x) = \frac{128}{3} x^3 e^{-4x} \qquad y(x) = -1 + 2 \int_0^x Y(w) \, dw$$
  
$$f_n(x) = \sqrt{Y(x)} \sqrt{2n+1} P_n[y(x)] \qquad Shape \text{ Functions with this "typical" space approximated by few basis elements}$$

- Need many basis functions to reasonably approximate SFs with widths different from the "built-in" one
- Basis elements with large n yield highly oscillating spectra and usually partake into delicate cancellations

## $\chi^2$ calculation

• As an example we tried a basis inspired by the one adopted by the SIMBA collaboration:





## Experimental data

 We considered data from 2012 BaBar fully inclusive and sum over exclusive analyses (in the B rest frame) and 2016 Belle results (in the  $\Upsilon(4S)$  rest frame):



### Some very preliminary results Babar fully incl. Babar sum excl.



Belle fully incl.

### Comparison with SCET approach [SIMBA collaboration: Bernlochner, Lacker, Ligeti, Stewart, F. Tackmann, K. Tackmann]

- the (leading) SF
- on different combinations of leading and subleading SFs.
- can only be used judiciously (i.e. with added uncertainties) in other inclusive B decays
- Within SCET it is possible to resum hard-collinear logs. We have  $Q_h \sim m_h = 4.57$  GeV, therefore quite small:  $\log[Q_h/Q_{hc}] = \log[Q_{hc}/Q_s] = 0.93$
- I conclude that these two approaches are effectively equivalent and offer a complementary  $\bullet$ approach to a simultaneous analysis of inclusive radiative and (rare-)semileptonic B decays

• Within SCET the separation of soft and hard-collinear scales is achieved within an effective theory framework (based on the method of regions and dim-reg): this corresponds to a different scheme for

• At leading power the SCET SF is universal for all inclusive B decays. At subleading power several additional universal SCET SFs appear. Unfortunately  $B \to X_s \gamma$ ,  $B \to X_u \ell \nu$  and  $B \to X_s \ell \ell$  depend

• Effectively both the kinetic and SCET approach extract from  $B \to X_s \gamma$  one effective SF which

 $Q_{hc} \sim \sqrt{m_b(m_B - m_b)} = 1.8 \text{ GeV}$  and  $Q_s \sim (m_B - m_b) = 0.71 \text{ GeV}$ . The relevant logs are



# Implications for $B \to X_{c}\ell\ell$

- SF needed for extrapolation in  $m_{X_c}$  and to improve the EvtGen Monte Carlo event generator which is the heart of Belle, BaBar and Belle II analyses.
- Hadronic spectrum is based on the Fermi motion implementation presented in Ali, Hiller, Handoko, Morozumi hep-ph/9609449:

$$\frac{d\Gamma_B}{ds \, du \, dp} = \int du' \frac{m_b(p)^2}{m_B} p \left[ \frac{4}{\sqrt{\pi p_F^3}} \exp(-p^2/p_F^2) \right] \left( u'^2 + 4m_b(p)^2 s \right)^{-1/2} \left[ \frac{d\Gamma_b}{ds \, du} \right]_{m_b \to m_b(p)}^{\star} \text{momentum dependent b mass}$$

- We need to urgently update the code!
- Work in progress on the complete triple differential rate at  $O(\alpha_{\rm s})$ [T. Huber, T. Hurth, J. Jenkins, EL, in preparation]

[EvtGen: Ryd, Lange, Kuznetsova, Versille, Rotondo, Kirkby, Wuerthwein, Ishikawa; Maintained by J. Back, M. Kreps and T. Latham at University of Warwick]

parton level with

pb = \_calcprob->FermiMomentum(\_pf); effective b-quark mass  $mb = mB*mB + _mq*_mq - 2.0*mB*sqrt(pb*pb + _mq*_mq);$ ( mb>0. && sqrt(mb)-\_ms < 2.0\*ml ) mb= -10.; mb = sqrt(mb);



