

# **Black holes in holographic quantum gravity**

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# Black hole thermodynamics & AdS

- “Thermodynamic” behavior of black holes is one of the mysteries of GR.
- One can ask various interesting questions:
  - precise statistical origin: entropy, Hawking radiation, phase transitions, ...
  - information loss & recovery, scrambling, chaos, complexity, ...
- However, general thermodynamics of quantum gravity couldn’t be studied well:
  - Does not know how to define quantum gravity at general (especially high) temperature.
  - Other conceptual subtleties to study full-fledged gravitational thermodynamics
- Some of these subtleties led Hawking & Page to study BH’s in AdS spacetime.
  - Global AdS is a gravitationally confining box.

$$ds_{D+1}^2 = d\rho^2 - \cosh^2 \frac{\rho}{\ell} dt^2 + \ell^2 \sinh^2 \frac{\rho}{\ell} ds^2(S^{D-1})$$

- Microscopic definition of quantum gravity, via CFT at the boundary  $S^{D-1} \times R$ : “AdS/CFT”
- Full-fledged study of BH thermodynamics is in principle possible.

# High temperature limit of CFT

- This leads us to study the thermodynamics of CFT on  $S^{D-1} \times R$ .
- It is easier to extract the physics at high temperature: This will be today's topic.
- Euclidean CFT on  $S_r^{D-1} \times S_\beta^1$ , at  $\beta \equiv T^{-1} \ll r^{-1}$ .

- Kaluza-Klein (KK) modes on small  $S^1$ :

- Divergent contributions to  $\log Z$  come from these KK modes at  $\beta \rightarrow 0$ .
- Encoded in the **effective action of background fields** (chemical potentials) on  $S^{D-1}$ .
- Universal background fields: metric, gravi-photon & dilaton:

$$ds_D^2 = ds_{D-1}^2 + e^{-2\Phi} (d\tau + a)^2 \quad \tau \sim \tau + \beta \quad , \quad \beta e^{-\Phi} \sim \text{radius of } S^1$$

- E.g. on  $S^3$  from temperature  $\beta^{-1}$  and angular velocities  $\omega_{1,2}/\beta$  :

$$ds^2 = r^2 \left[ d\theta^2 + \sum_{i=1}^2 n_i^2 \left( d\phi_i - \frac{i\omega_i}{\beta} d\tau \right)^2 \right] + d\tau^2 = r^2 \left[ d\theta^2 + \sum_i n_i^2 d\phi_i^2 + \frac{r^2 (\sum_i \omega_i n_i^2 d\phi_i)^2}{\beta^2 (1 - r^2 \sum_i \frac{n_i^2 \omega_i^2}{\beta^2})} \right] + e^{-2\Phi} (d\tau + a)^2$$

$$e^{-2\Phi} = 1 - r^2 \sum_i \frac{n_i^2 \omega_i^2}{\beta^2} \quad , \quad a = -i \frac{r^2 \sum_i \omega_i n_i^2 d\phi_i}{\beta (1 - r^2 \sum_i \frac{n_i^2 \omega_i^2}{\beta^2})} \quad (n_1, n_2) = (\cos \theta, \sin \theta)$$

# Effective action & derivative expansion

- In CFT, **small  $\beta$  expansion =  $r^{-1}$  expansion = derivative expansion on  $S^{D-1}$ .**
- $\infty$ -tower of derivative expansion. Most coefficients depend on coupling constants.

$$\frac{A_1}{\beta^3} \int d^3x e^{3\Phi} \sqrt{g} + \frac{A_2}{\beta} \int d^3x \sqrt{g} e^{\Phi} R + A_3 \int d^3x \sqrt{g} (\beta e^{-\Phi}) (\beta^{-1} da)^2 + \dots$$

cosmological constant
Einstein-Hilbert
gravi-photon kinetic

- However, Chern-Simons terms on  $S^{D-1}$  (for even  $D$ ) are coupling independent.
- 't Hooft anomalies of global symmetries determine them. (More later)
- They are sub-leading terms in the normal thermal free energy.

$$k\beta^{-2} \int a \wedge da + k_I \beta^{-1} \int \boxed{\mathcal{A}^I} \wedge da + k_{IJ} \int \mathcal{A}^I \wedge d\mathcal{A}^J + \dots$$

background gauge fields ( $\sim$  chemical potential) for global symmetries

- CFT w/ AdS dual: The expansion constrains AdS black hole's thermodynamics.

# Kerr black holes in AdS<sub>5</sub>

- Plug in our  $S^3$  background fields to:

$$\frac{A_1}{\beta^3} \int d^3x e^{3\Phi} \sqrt{g} + \frac{A_2}{\beta} \int d^3x \sqrt{g} e^{\Phi} R + A_3 \int d^3x \sqrt{g} (\beta e^{-\Phi}) (\beta^{-1} da)^2 + \dots$$

- Leading terms in derivative expansion. (Take  $\Omega \equiv \omega_1 = \omega_2$  for simplicity.)

$$S = \log Z + \beta E + 2\Omega J = A_1 \frac{\beta}{(\beta^2 - \Omega^2)^2} + A_2 \frac{3\beta^3 - 4\beta\Omega^2}{(\beta^2 - \Omega^2)^2} + A_3 \frac{\beta\Omega^2}{(\beta^2 - \Omega^2)^2} + \beta E + 2\Omega J + \mathcal{O}(\beta^0)$$

- Extremize in  $\beta, \Omega$  & solve for  $E, J, S$ . 2 more eqns.
- Expressions for  $S, E, J$  in terms of  $\beta, \Omega$ , and the unknowns  $A_i$ .
- Fitting three  $A_i$ 's, one reproduces all divergent parts of  $E, J, S$  carried by Kerr BH's. (Here,  $\Omega = -a\beta + \dots$ .) Although 3 parameters are fitted by hand, still quite nontrivial constraints.

$$E = \frac{\pi l^2}{G} \left( \frac{3 + a^2}{8(1 - a^2)^3} \frac{l^4 \pi^4}{\beta^4} - \frac{3 + 3a^2 - 2a^4}{8(1 - a^2)^3} \frac{l^2 \pi^2}{\beta^2} - \frac{3 - 11a^2 + 6a^4 - 2a^6}{16(1 - a^2)^3} + \mathcal{O}(\beta^2) \right)$$

$$J = \frac{\pi l^2}{G} \left( \frac{a}{4(1 - a^2)^3} \frac{l^4 \pi^4}{\beta^4} - \frac{2a - a^3}{4(1 - a^2)^3} \frac{l^2 \pi^2}{\beta^2} + \frac{a}{8(1 - a^2)^3} + \mathcal{O}(\beta^2) \right)$$

$$S = \frac{\pi^2 l^3}{G} \left( \frac{1}{2(1 - a^2)} \frac{l^3 \pi^3}{\beta^3} - \frac{3 - 2a^2}{4(1 - a^2)^2} \frac{l\pi}{\beta} + \mathcal{O}(\beta) \right)$$

# High T expansion of indices

- AdS/CFT models w/ SUSY: “BPS states” w/ coupling-independent spectrum.
- “Witten index”: coupling independent partition function. For 4d SCFTs on  $S^3 \times S^1$ ,

$$Z(\omega_1, \omega_2) \equiv \text{Tr} \left[ e^{-\beta(E - \frac{3}{2}R - J_1 - J_2)} e^{-\frac{1}{2}\Delta R} e^{-\omega_1 J_1 - \omega_2 J_2} \right]$$

$$\Delta = \omega_1 + \omega_2 + 2\pi i \quad E - \frac{3}{2}R - J_1 - J_2 \sim \{\mathcal{Q}, \mathcal{Q}^\dagger\}$$

- On  $S^3$ , coupling-independent CS terms determine the “high T” or large charge asymptotics.
- $\beta$  is merely a regulator of the index. Easy to compute at  $\beta \rightarrow 0^+$ .
- Since the leading term is  $O(\beta^0)$ ,  $\beta \rightarrow 0$  is a fake thermal circle parameter.
- True derivative expansion parameters are  $|\omega_i| \ll 1$ .

$$e^{-2\Phi} = 1 - \boxed{r^2 \sum_i \frac{n_i^2 \omega_i^2}{\beta^2}}, \quad a = -i \frac{r^2 \sum_i \omega_i n_i^2 d\phi_i}{\beta(1 - r^2 \sum_i \frac{n_i^2 \omega_i^2}{\beta^2})}$$

- Background fields: same metric, dilaton, gravi-photon (up to small shifts of  $\omega_i$  by  $\beta$ )
- Background  $U(1)_R$  & other flavor symmetries’ gauge fields:

$$A^I = -\frac{i\Delta_I}{\beta} d\tau \xrightarrow{\text{rearrange to 3d fields}} A^I = A_4^I(d\tau + a) + \mathcal{A}^I \quad A_4^I = -\frac{i\Delta^I}{\beta} \quad \mathcal{A}^I = -A_4^I a$$

# CS coefficients from anomalies

- Knowing these CS coefficients, one can determine the leading free energy.
- There turn out to be two types of CS terms.

- Gauge non-invariant CS terms

- 4d effective action  $S_{\text{eff}} = -\log Z$  respects 't Hooft anomaly:  $\delta_\epsilon S_{\text{eff}} \sim \epsilon F \wedge F + \dots$
- This should be reflected in the 3d background fields' effective action.
- It demands the existence of certain gauge non-invariant CS terms.

[Banerjee, Bhattacharya, Bhattacharyya, Jain, Minwalla, Sharma] (2012)

$$S_{\text{non-inv}} = -i \frac{\beta(5a - 3c)}{8\pi^2} \int_{S^3} C_{IJK} \left( A_4^I \mathcal{A}^J \wedge d\mathcal{A}^K + A_4^I A_4^J \mathcal{A}^K \wedge da + \frac{1}{3} A_4^I A_4^J A_4^K a \wedge da \right)$$

- Gauge invariant CS terms: More elaborate arguments by [Jensen, Loganayagam, Yarom] (2013) determine them all from anomalies. (See also [Di Pietro, Komargodski] (2014).)

$$S_{\text{inv}} \sim -i(a - c) \int_{S^3} k_I \mathcal{A}^I \wedge da$$

[ $C_{IJK}$ ,  $k_I$  are coefficients of the cubic anomaly polynomial in a suitable normalization.]

# Free energy & BH's

- Plugging in our background fields, one obtains the leading free energy.

$$\log Z \sim (5a - 3c) \frac{C_{IJK} \Delta^I \Delta^J \Delta^K}{6\omega_1 \omega_2} + \frac{4\pi^2 (a - c) k_I \Delta^I}{\omega_1 \omega_2}$$

- If one only turns on the chemical potential for the U(1) superconformal R-symmetry, one obtains a universal formula in terms of two central charges.

$$\log Z \sim -\frac{16i}{27} \frac{3c - 2a}{\omega_1 \omega_2} \quad C_{111} = 6, \quad k_1 = 1, \quad \Delta^1 = \frac{2\pi i}{3}$$

- It is known that  $3c - 2a > 0$  is always met by interacting SCFTs. [Hofman, Maldacena]
- Its Legendre transformation is subtle. The proper interpretations established only recently.
- As a result, at large  $J_1, J_2 \gg a, c$ , one obtains a macroscopic entropy.

$$S = \sqrt{3} \left[ 2(3c - 2a) \left( J_1 + \frac{R}{2} \right) \left( J_2 + \frac{R}{2} \right) \right]^{\frac{1}{3}}$$

- This entropy precisely agrees with the Bekenstein-Hawking entropy of BPS black holes in  $AdS_5$ , accounting for their microstates from CFT.



# Conclusion & remarks

- Implications to CFT:

- What I explained generalize the Cardy formula of 2d CFT, to all even dimensional SCFTs.

$$Z(\tau) \sim \text{Tr} [e^{2\pi i\tau L_0}] \sim \exp \left[ \frac{\pi i c}{12\tau} \right] \quad \text{at } \tau \rightarrow i0^+$$

- 2d conformal symmetry is very strong. Universal high T asymptotics.
- For SCFTs, continues to find universal asymptotics in terms of  $3c - 2a$ .

- Implications to black hole physics:

- Black holes in AdS from dual CFT: deconfined quark-gluon plasma [Witten] (1998)
  - However, only recently we have been able to confirm and study this quantitatively.
  - Today, I explained to you the simplest route to understand the recent developments.
- Today's talk is only a small part of the recent advances in BPS AdS black holes.
  - We counted the microstates of BPS  $AdS_{D+1}$  black holes from  $SCFT_D$  Cardy indices for all  $D = 3,4,5,6$ . [Choi, J. Kim, SK, Nahmgoong] [Choi, SK] [J. Kim, SK, Song] [Nahmgoong] [Choi, Hwang, SK] [Choi, Hwang] .....