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# Overview of Intensity-Frontier Efforts

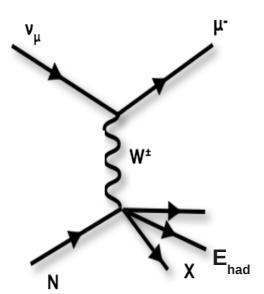
Minerba Betancourt, Fermilab 15 November 2019

#### Introduction

- Oscillation probability depends on neutrino energy  $E_{\nu}$
- We need to reconstruct the neutrino energy precisely

$$P(\nu_{\alpha} \to \nu_{\beta}) \approx 1 - \sin^2 2\theta \sin^2(\frac{\Delta m^2 L}{E_{\nu}})$$

- Neutrino energy reconstruction is obtained using the final state particles of neutrino-nucleus interactions
  - Fully active experiments reconstruct the energy using:  $E_v = E_{lepton} + hadron$

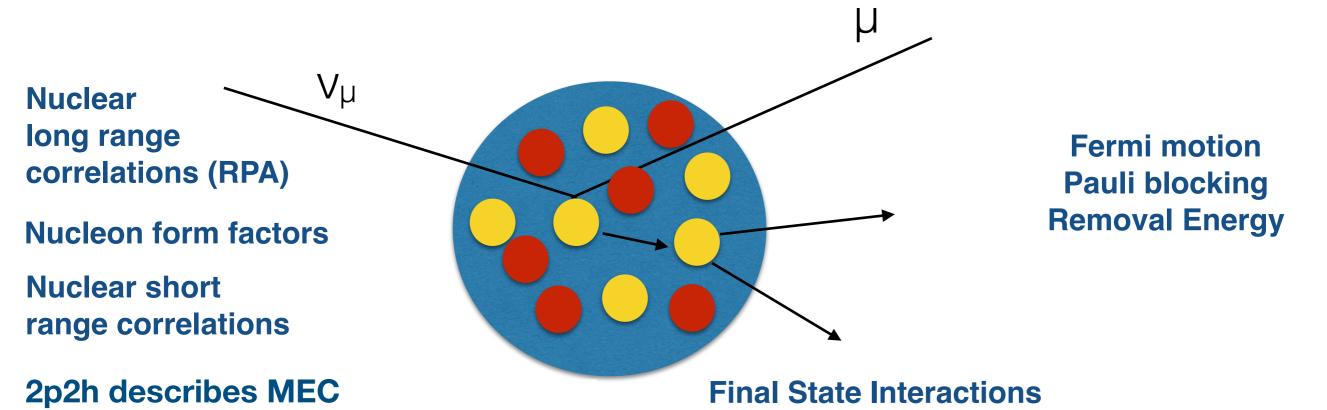


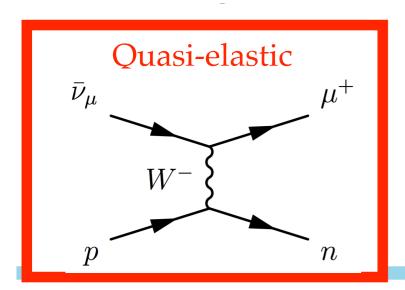
 Nuclear effects modify the kinematics of the particles and the reconstruction of the neutrino energy

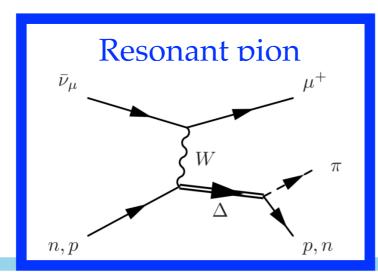


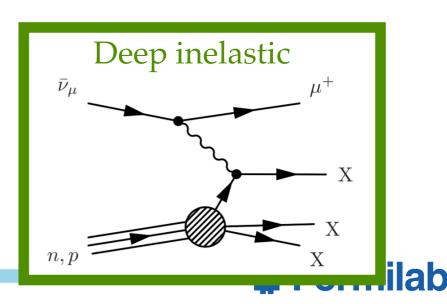
## **Neutrino Scattering**

- Understanding neutrino interactions is challenging
- Modeling the interactions and measuring them present different types of challenges



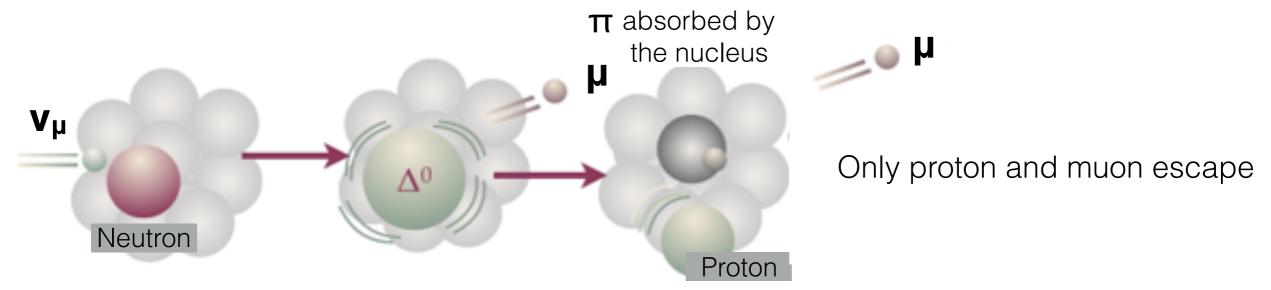






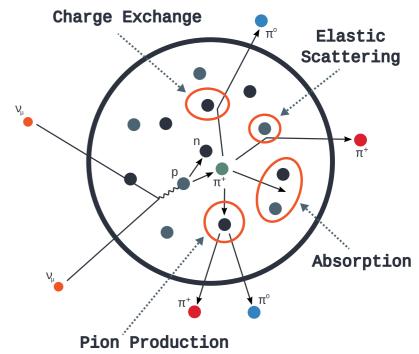
## **Example of Nuclear Effects (Final State Interactions)**

- Final state interactions (FSI):
  - Due to final state interactions, particles can interact with nucleons, pions can be absorbed before exiting the nucleus, and other nucleons get knocked out



Start as a RES interaction, the pion is absorbed and the interaction looks QE like in our detector

• Nuclear effects modify the true/reco neutrino energy relationship and final-state particle kinematics!

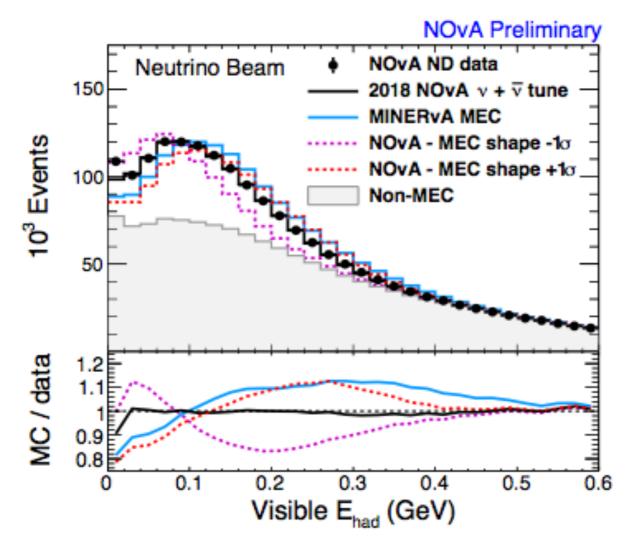


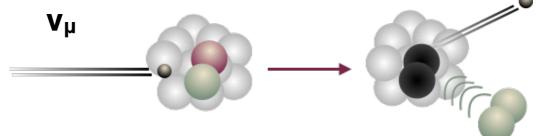


## Detailed Understanding of Neutrino Interactions is Critical for Oscillation Experiments

•Example of different models for multi-nucleon effects compared to NOvA Near

Detector data





#### nucleons

#### The most important systematics

- Neutrino cross sections
  - Particularly nuclear effects (RPA and MEC)
- Neutron uncertainty

Fermilab Wine and Cheese Seminar, June 15 2018, Alex Himmel

MINERVA MEC: Theoretical prediction from Valencia Model PRC 70, 055503 (2004) tuned with MINERVA data NOVA MEC: GENIE Empirical MEC tuned with NOvA ND data



# Source of uncertainty for oscillation experiment

NOvA T2K

Source of uncertainty	$\nu_e$ signal (%)	Total beam background (%)
Cross sections and FSI	7.7	8.6
Normalization	3.5	3.4
Calibration	3.2	4.3
Detector response	0.67	2.8
Neutrino flux	0.63	0.43
$\nu_e$ extrapolation	0.36	1.2
Total systematic uncertainty	9.2	11
Statistical uncertainty	15	22
Total uncertainty	18	25

Phys. Rev. D 91.072010

Source of uncertainty	$\nu_{\mu}$ CC	$\nu_e$ CC
Flux and common cross sections		
(w/o ND280 constraint)	21.7%	26.0%
(w ND280 constraint)	2.7%	3.2%
Independent cross sections	5.0%	4.7%
SK	4.0%	2.7%
FSI+SI(+PN)	3.0%	2.5%
Total		
(w/o ND280 constraint)	23.5%	26.8%
(w ND280 constraint)	7.7%	6.8%

Phys. Rev. D 98.032012



#### **Present and Future**

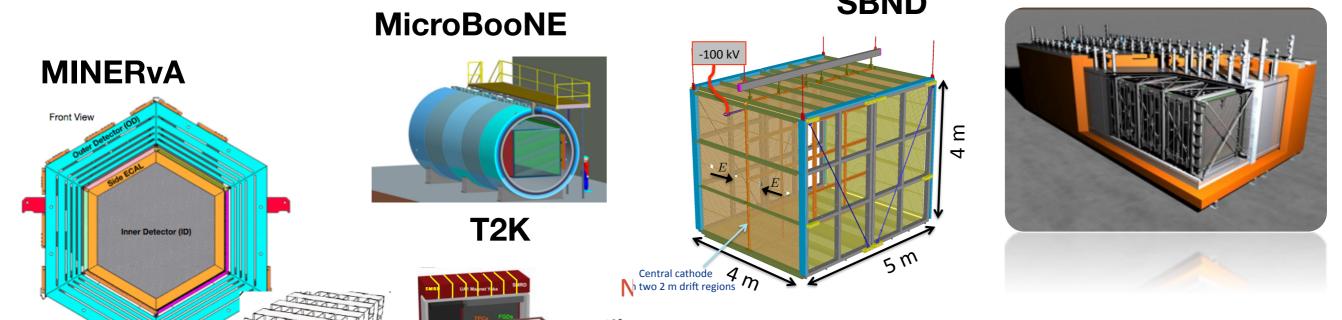
· We are using heavy targets for oscillation experiments, such as carbon and liquid argon

• The field of oscillation experiments are getting more statistics: accelerators are doing

better than ever, and detectors are getting bigger

SBND

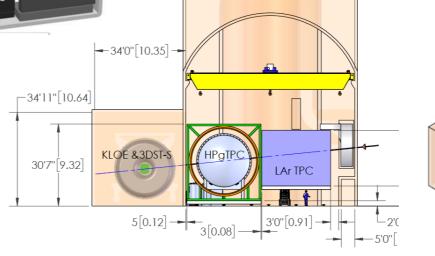
ICARUS

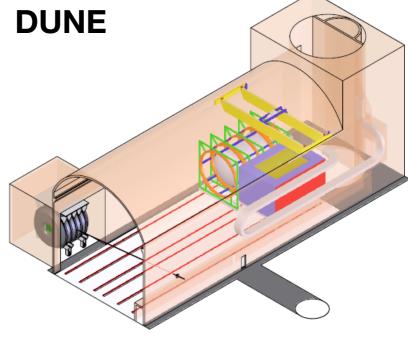


DUNE Near Detector
Complex detector, main components:
Liquid argon+downstream tracker
with gaseous argon target

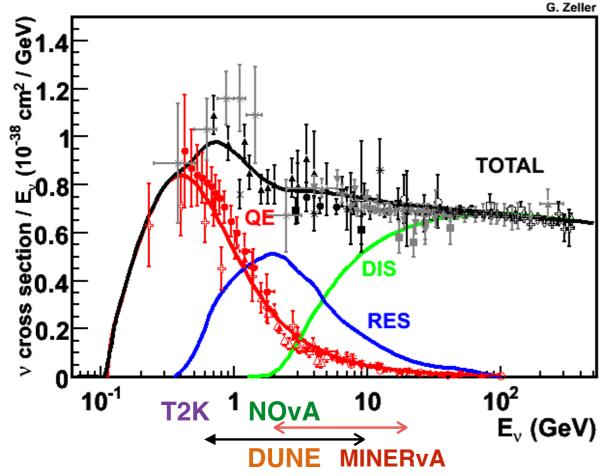
**NOvA** 

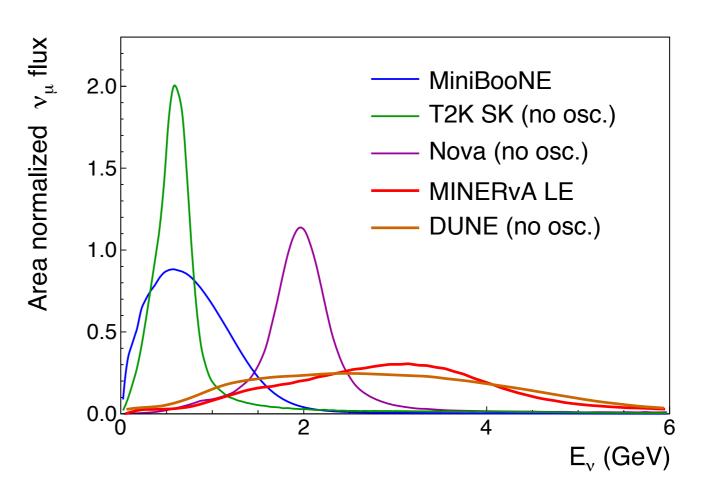
Detectors can move to off-axis fluxes





## **Neutrino Energies for Different Experiments**





J.A. Formaggio, G. Zeller, Reviews of Modern Physics, 84 (2012)

**Plot courtesy of Phil Rodrigues** 

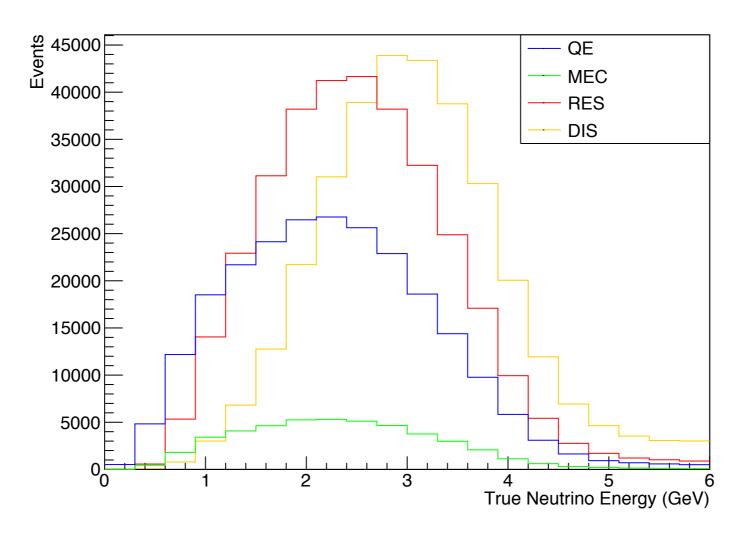
- We have a lot of progress these past years for the simplest interactions CCQE  $(0\pi)$
- Starting to focus on the pion production, mainly because NOvA has a good contribution from pion events and T2K is including pions events in their signal to get more statistics
- Next crucial channel is the DIS scattering, the largest contribution for DUNE
- Showing results from the last few years



#### Muon Neutrino at the Near Detector in DUNE

#### **GENIE** predictions for argon

DUNE Near Detector V<sub>µ</sub>

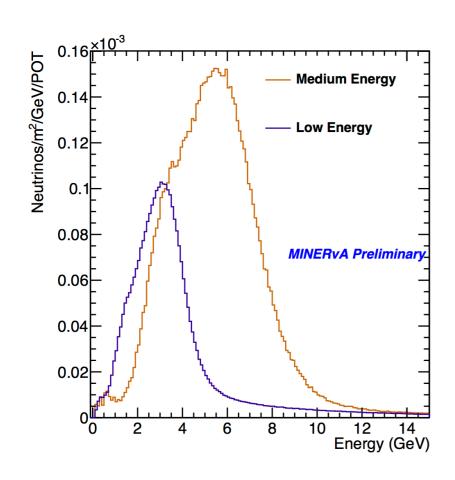


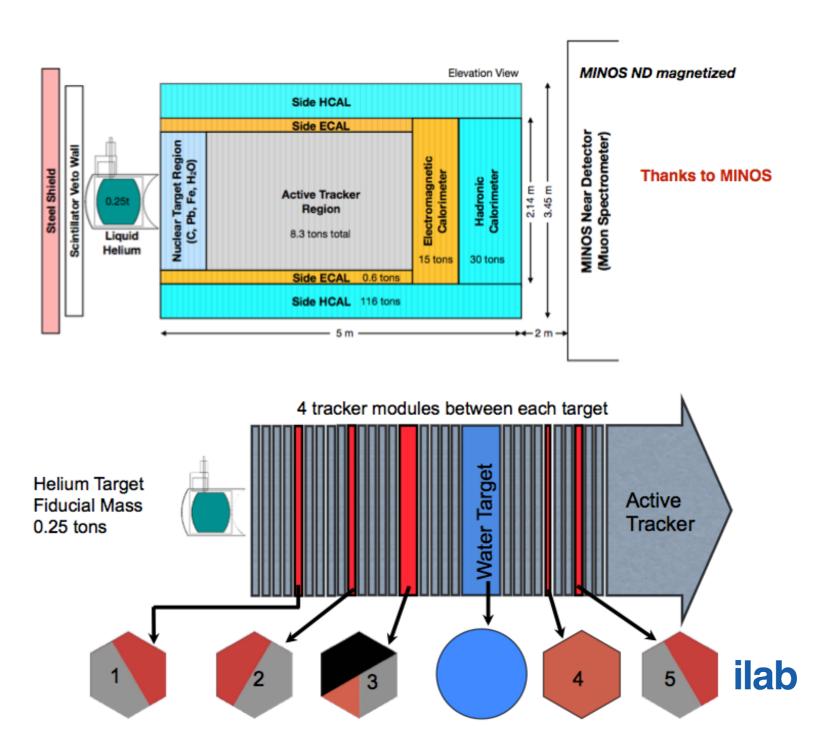
Significant contributions from RES and DIS interactions



## **MINERvA** Experiment

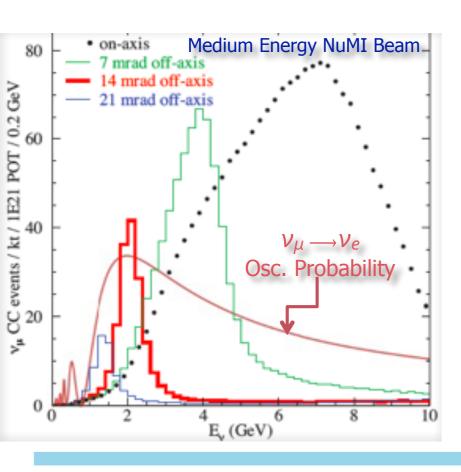
- •Fine-grained scintillator tracker surrounded by calorimeters
- •We have different nuclear targets which allows us to study nuclear effects and see how the different processes are affected by the nucleus

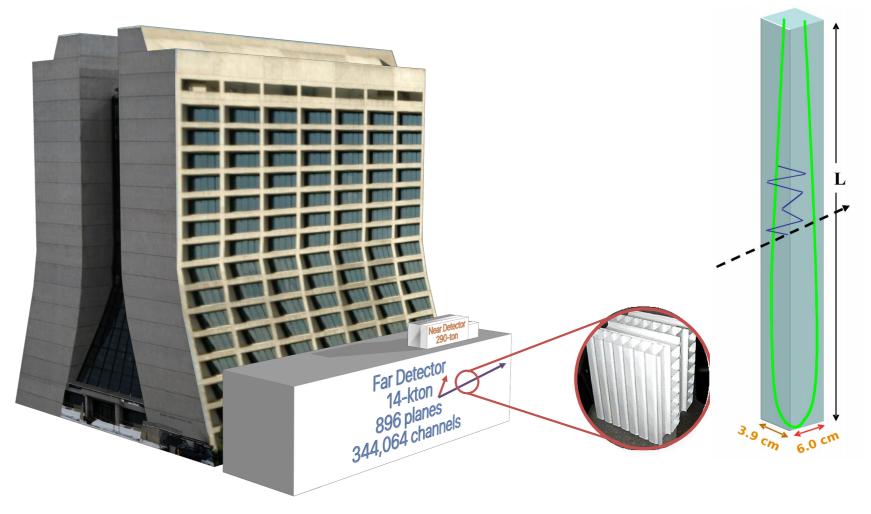




## **NOvA Experiment**

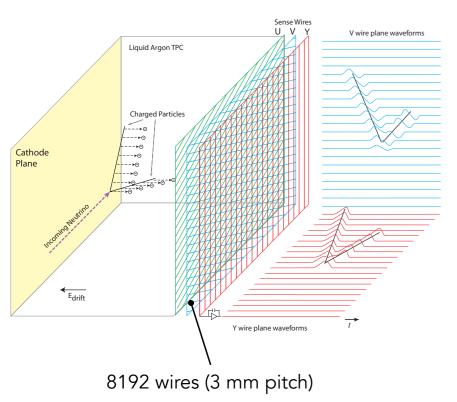
- Near Detector is 290 tons placed 300 ft underground at Fermilab
- Consists of plastic cells filled with liquid scintillator
- Off axis beam neutrinos

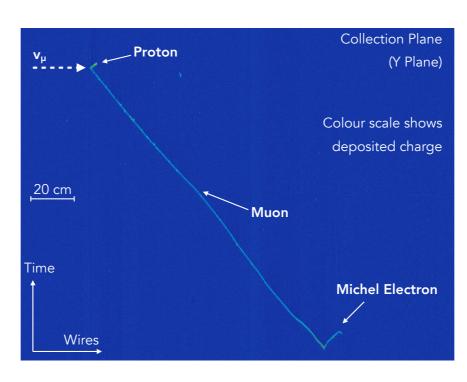


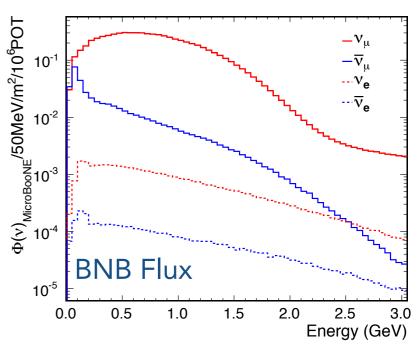


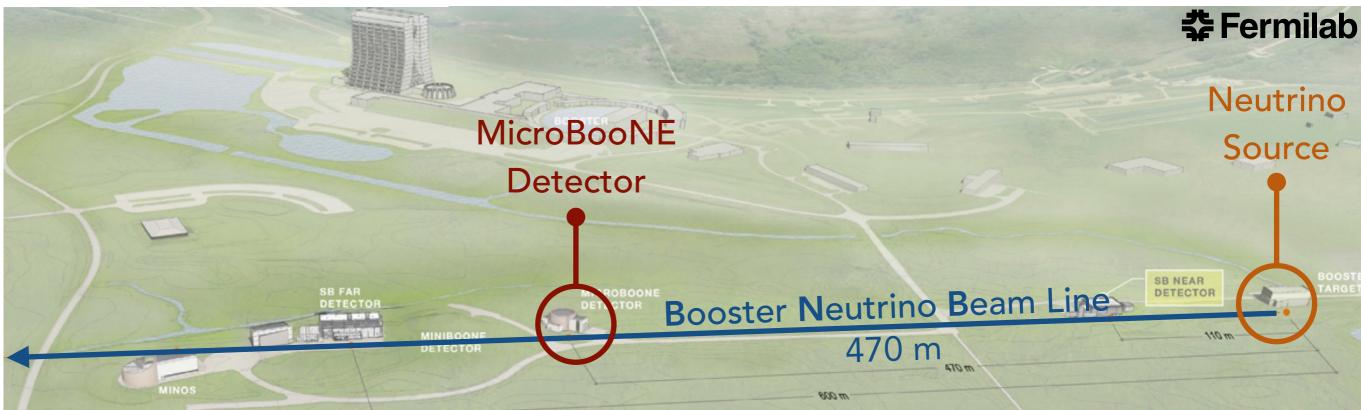


## **MicroBooNE Experiment**



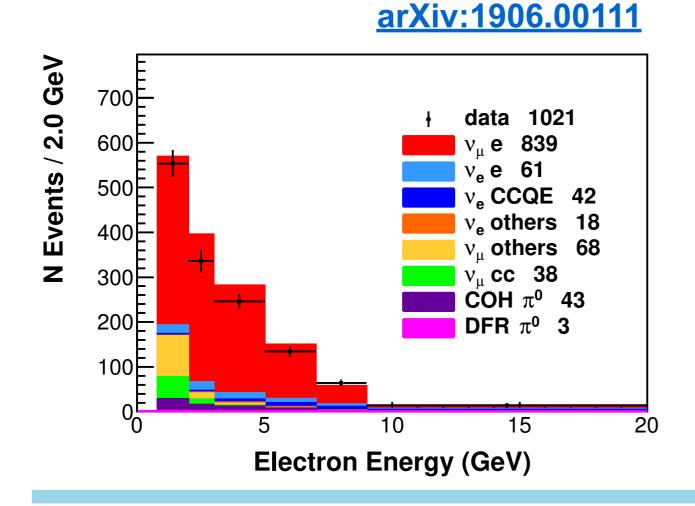


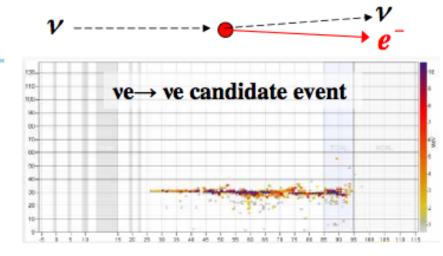


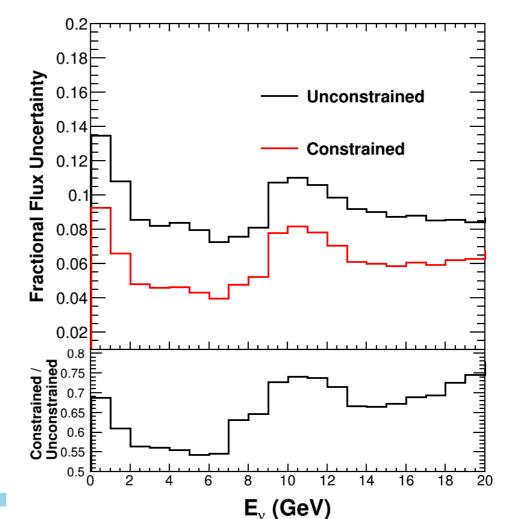


#### Flux Prediction

- MINERvA constrain flux with in-situ measurements:
   V+e scattering
- Using a total sample of 810 signal events
- The measurement reduces the normalization uncertainty on the  $\nu_{\mu}$  NuMI flux from 7.5% to 3.9%
- The most precise flux constraint ever







#### **Neutrino Event Generators**

We have several event generators: GENIE, NEUT, NuWro and GiBUU

#### **NuWro**



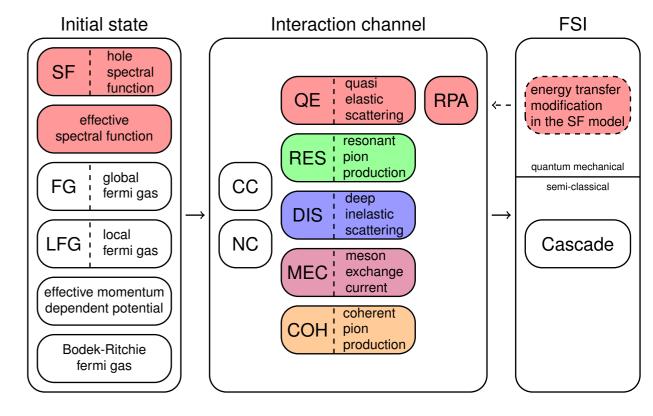
#### **GENIE**

- CCQE models
  - Llewellyn Smith
  - Nieves, Amaro and Valverde
- MEC models
  - Empirical
  - Nieves Simo Vacas
- Nuclear Models
  - Relativistic Fermi Gas
  - Local Fermi Gas
  - Effective Spectral Functions

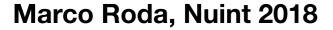


- Single Kaon
- Λ production

- RES
  - Rein-Sehgal
  - Berger-Sehgal
  - Kuzmin-Lyubushkin-Naumov
- COH
  - Rein-Sehgal
  - Berger-Sehgal
  - Alvarez Ruso
- FSI Intranuke
  - Full Intra-Nuclear cascade
  - Schematic based on Hadron-nucleus data



Jan Sobczyk, Nuint 2018



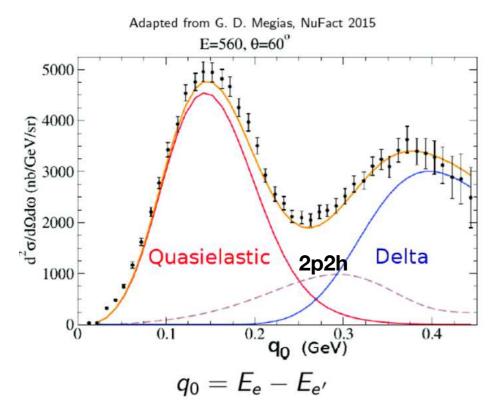


### Inclusive CC Double Differential Cross Section from MINERvA

- Cross section in q<sub>0</sub> and q<sub>3</sub>
  - q<sub>0</sub> is calorimetric hadronic energy and q<sub>3</sub> is the three momentum transfer

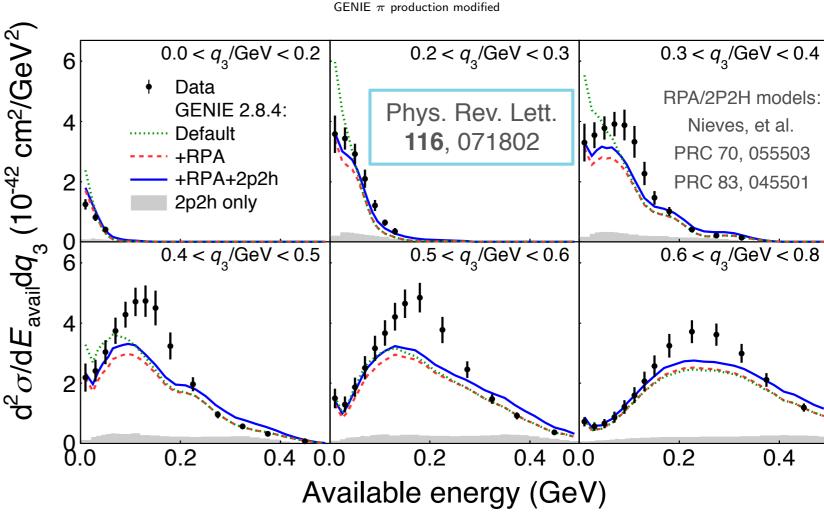
$$E_{
u} = E_{\mu} + q_0$$
  $q_3 \equiv |\mathbf{q}| = \sqrt{Q^2 + q_0^2}$ 

#### From electron scattering



New model predictions for 2p2h and RPA from the Valencia group (PRC 70, 055503 PRC 83, 045501)

# Similar measurement for neutrinos using the hadronic system and the lepton

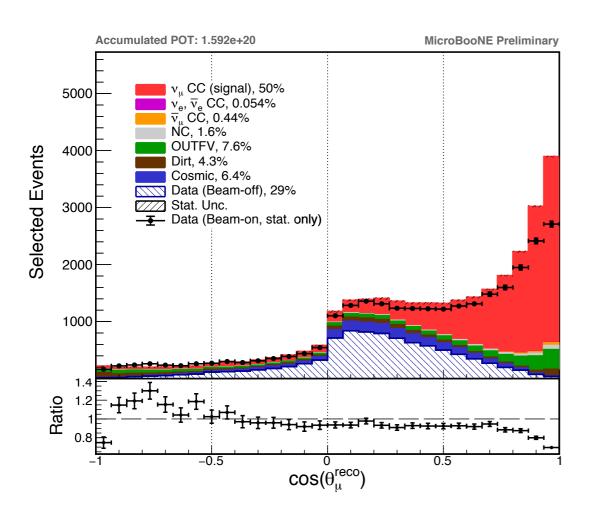


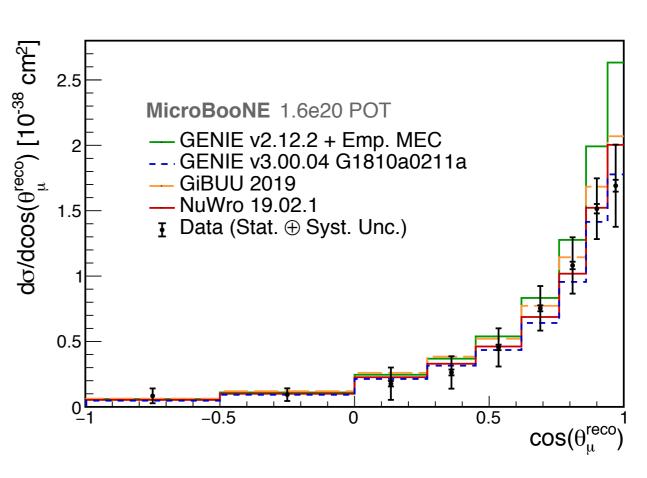
Model does not provide enough strength in some regions

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## First v<sub>µ</sub> CC Inclusive Cross Section from MicroBooNE

- MicroBooNE is on Earth's surface
- Cosmic rays are the main background
- Rate 5.5 kHz —> 25 cosmic rays per recorded event
- GENIE over predicts the cross section in the very high momentum and forward region





https://arxiv.org/abs/1905.09694



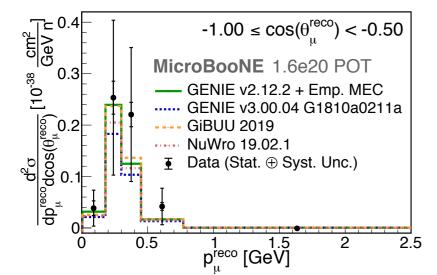
uBooNE

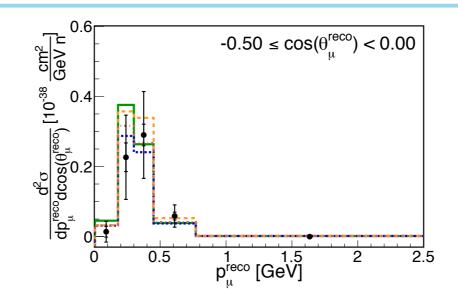
## **MicroBooNE Double Differential Cross Section**

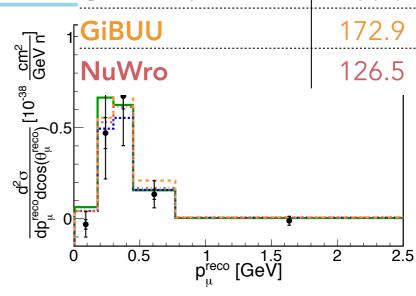
 Model
 χ²

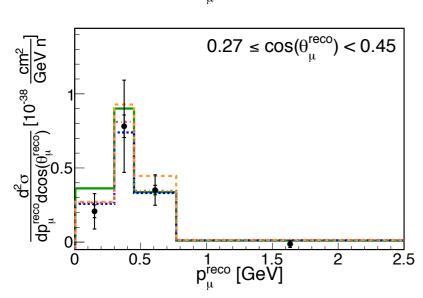
 GENIE v2 + MEC
 245.9

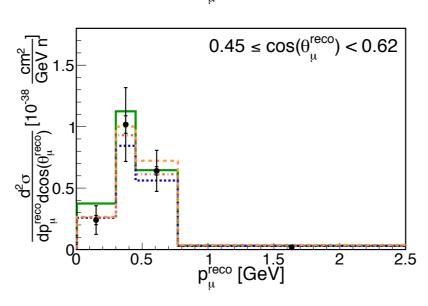
 GENIE v3
 108.8

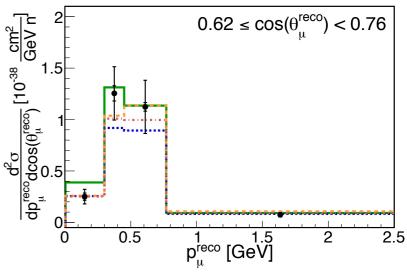


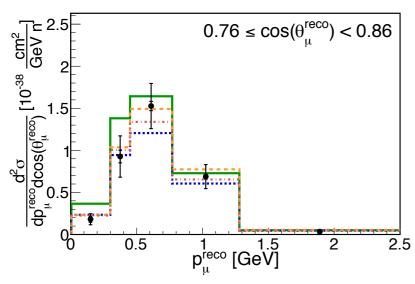


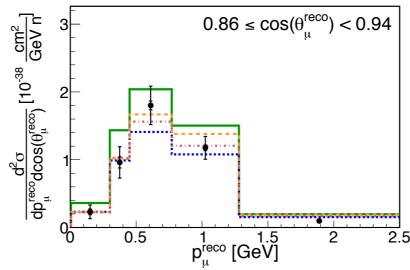


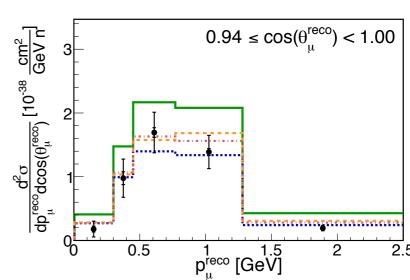








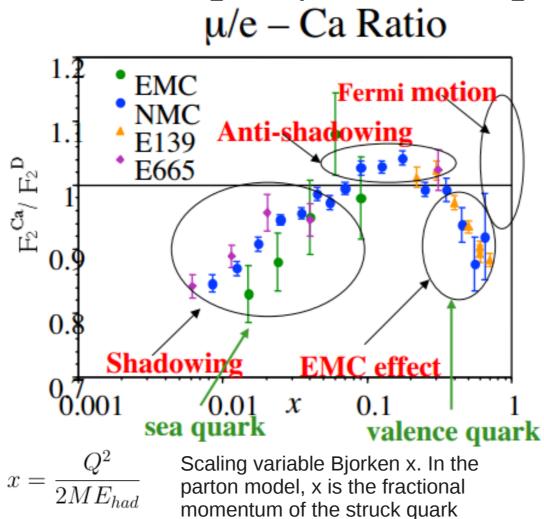




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## **Deep Inelastic Scattering**

- One of the most important channels for the next generation of oscillation experiment (DUNE)
- Different effects observed in charged lepton scattering



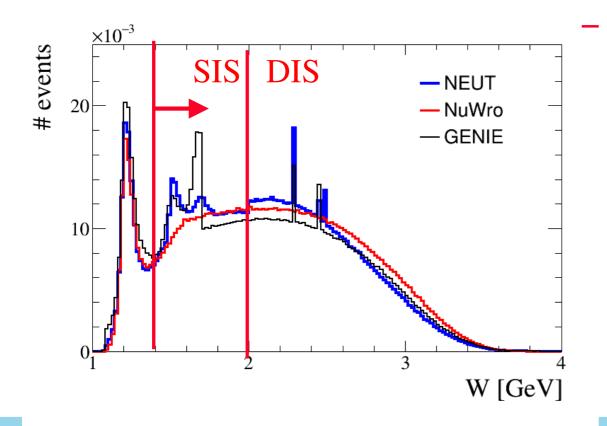
• We use x-dependent nuclear effects for Fe and Pb based on charged lepton scattering data, need to test with neutrinos

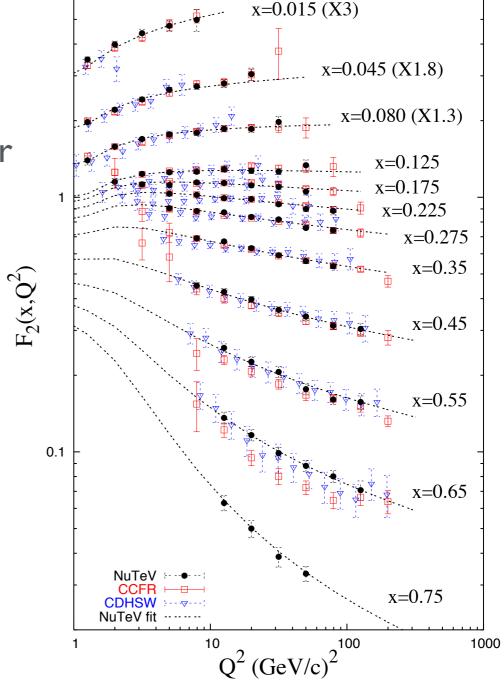


#### **Old DIS Data**

- NuTeV accumulated over 3 million neutrino / antineutrino events with 20<E<400</li>
   GeV, NuTeV operated for 15 months in 1996
- NuTeV measured F<sub>2</sub> on Fe
- At low x very different Q<sup>2</sup> dependence
- At high x (x>0.5) NuTeV is systematically higher

# What do the event generators say?

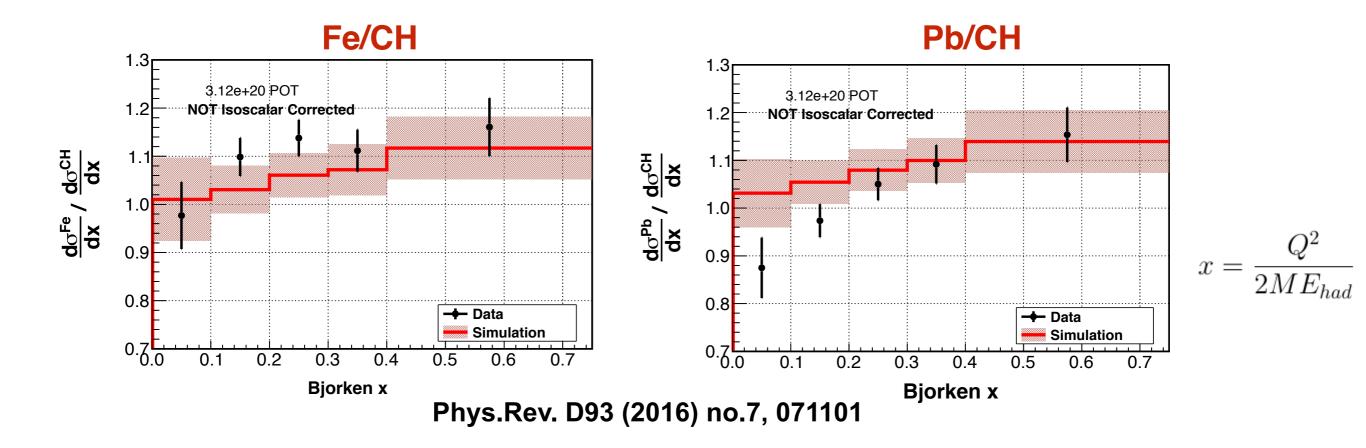






#### MINERvA DIS Cross Sections On Different Nuclei

 MINERvA measured deep inelastic ratios from nuclear targets to study x dependence nuclear effects, new analysis with more statistic is underway



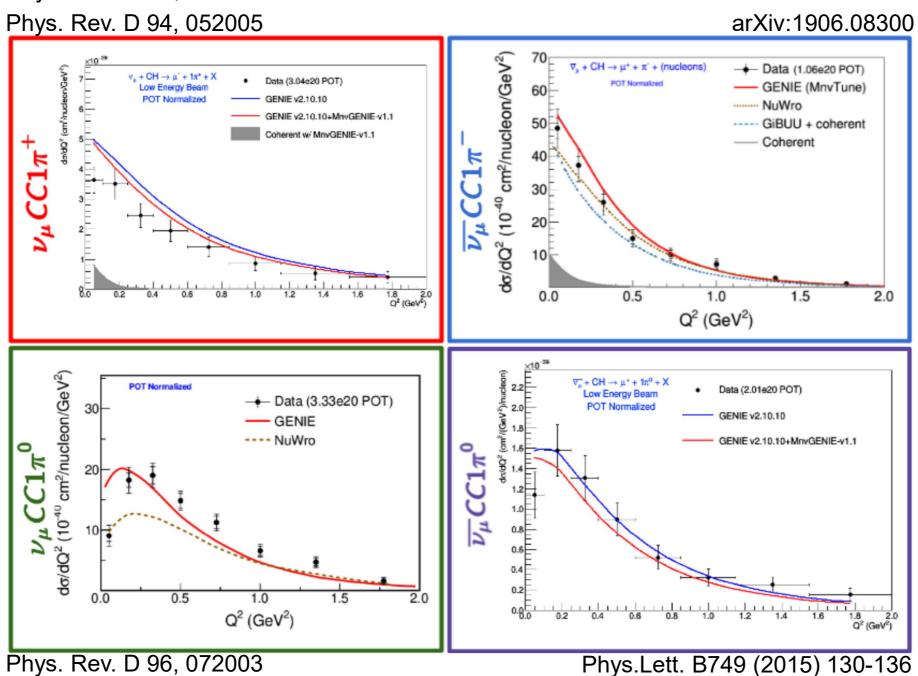
- The data suggest additional nuclear effects in the lowest x bin for Pb/CH
- In the region (0.3<x<0.75) we found good agreement between data and simulation
- New analysis in preparation with unprecedented statistics in Fe and Pb for neutrinos and antineutrinos



#### **MINERvA Pion Production Measurements**

• Several pion cross section measurements, MINERvA sees deficit of pion production at low Q<sup>2</sup>

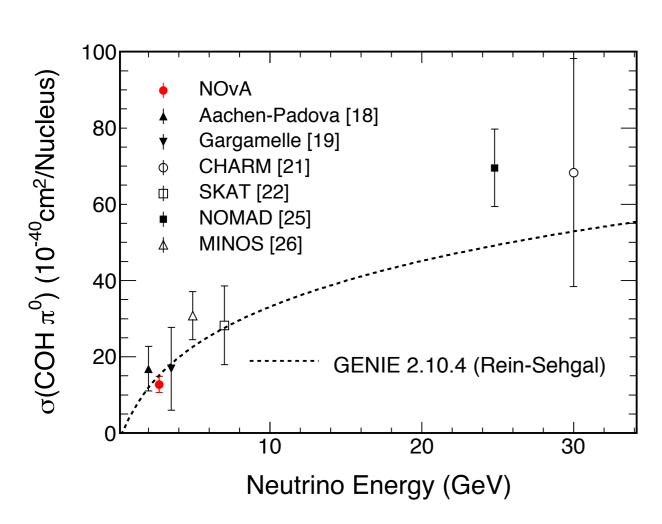
Phys. Rev. D 92, 092008



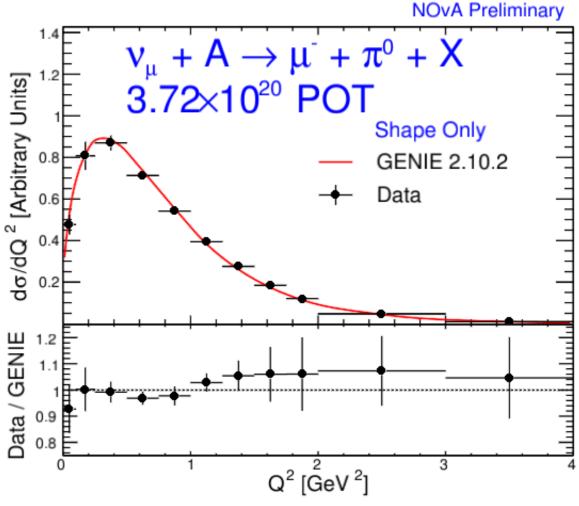
Phys. Rev. D 94, 052005

## **NOvA** Neutral Current Coherent π<sup>0</sup> and Charged Current semi-inclusive π<sup>0</sup>

- NOvA Near Detector measured neutral-current coherent  $\pi^0$  production on a carbon-dominated target
- The flux-averaged cross section is consistent with the model prediction
- Generally consistent with GENIE predictions





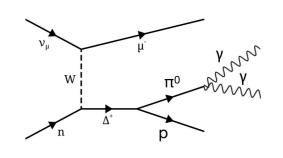


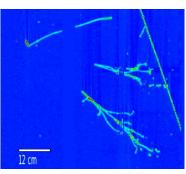
Fermilab Wine and Cheese Seminar, Dec 2017, Dan Pershey



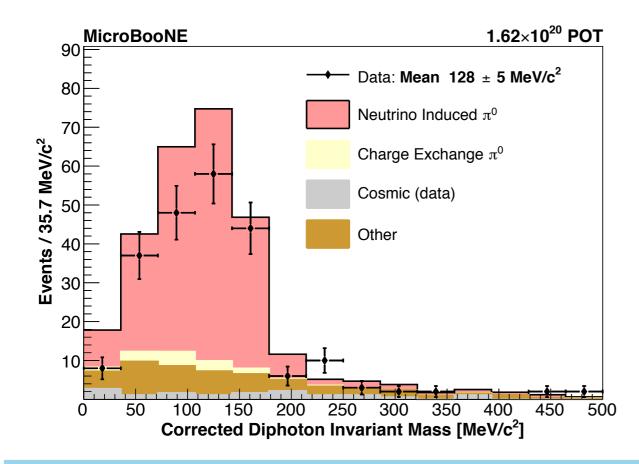
## First Charged Current Π<sub>0</sub> Production on Argon

- First generation MicroBooNE analysis is a great demonstration of reconstruction in LArTPCs
- Invariant mass expected to be m=135 MeV

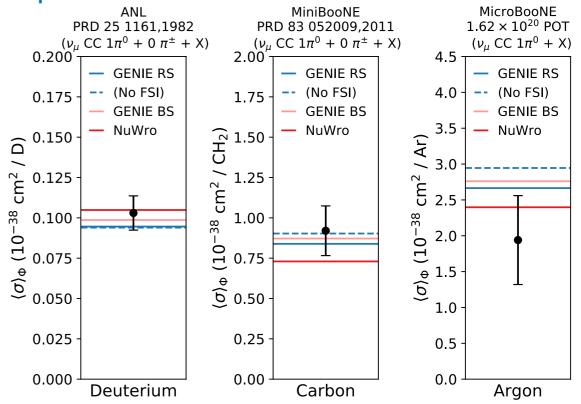




Reconstructed invariant mass of the two photon candidates associated to the neutrino interaction vertex



Results are compared to past measurements on lighter nuclei to compare the models predictions



Phys.Rev. D99 (2019) no.9, 091102



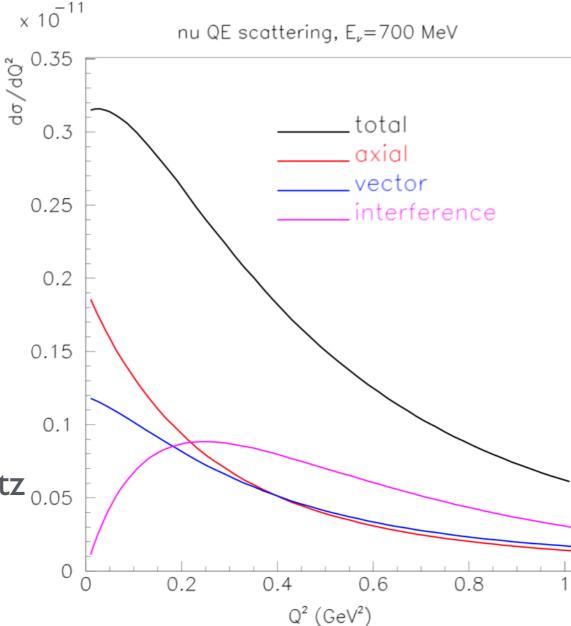
## **Quasi-Elastic Scattering (CCQE)**

- The Quasi-elastic process gives the largest contribution for the signal in many
  - oscillation experiments
- We use a free nucleon CCQE formalism to determine the cross section

$$\frac{d\sigma}{dq^2} \propto (F_1, F_2, F_A)$$

- Depend on the form factors  $F_1$ ,  $F_2$  and the axial form factor FA
- The vector form factors  $F_1$ ,  $F_2$  are known from electron-nucleon scattering
- ullet The axial form factor is described using an ansatz  $_{0.05}$

$$F_A(Q^2) = \frac{F_A(0)}{(1 - \frac{q^2}{M_A^2})^2}$$



• FA(0) is constrained from neutron beta decay and  $M_A$  is the axial mass



#### **Axial Form Factor**

The dipole axial form factor ansatz:

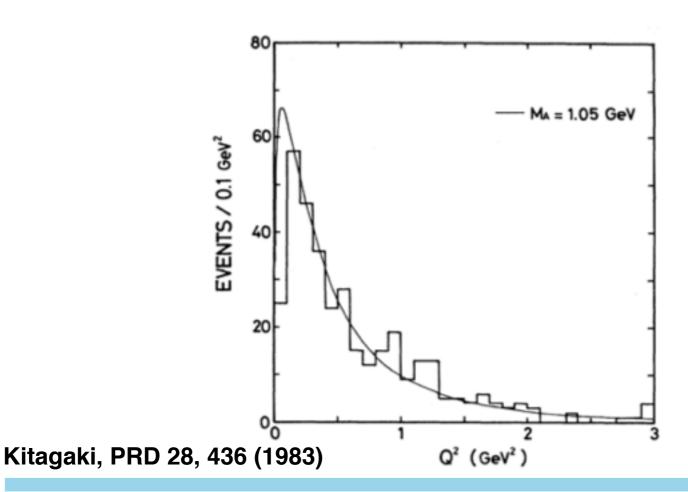
#### **Bubble chamber experiment at Fermilab**

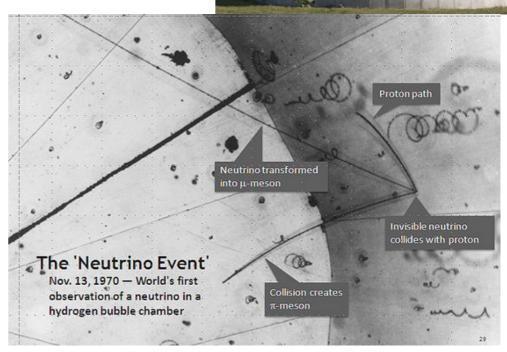
$$F_A(Q^2) = \frac{F_A(0)}{(1 - \frac{q^2}{M_A^2})^2}$$

 $\bullet$  Experiments with deuterium targets have employed this anzatz, obtaining a world average  $M_A$ 

$$M_A = 1.014 \pm 0.014 GeV$$

Eur. Phys. J. C 53, 349 (2008)





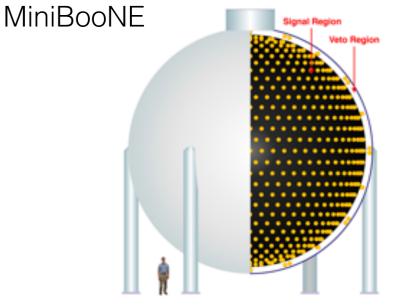


#### **Axial Form Factor**

The dipole axial form factor ansatz:

$$F_A(Q^2) = \frac{F_A(0)}{(1 - \frac{q^2}{M_A^2})^2}$$

 Modern experiments using heavy targets, like carbon from MiniBooNE reported a higher axial mass



- Other experiments such as K2K, SciBar and MINOS find similar higher axial mass compared with the world average
- This high mA is an effective parameter that we expect represents multi-nucleons effects, and not directly the form factor itself



#### **Axial Form Factor**

- A model independent description of the axial form factor called z-expansion is derived in Phys. Rev. D84 (2011)
- The form factor can be expressed as a power series of a new variable z

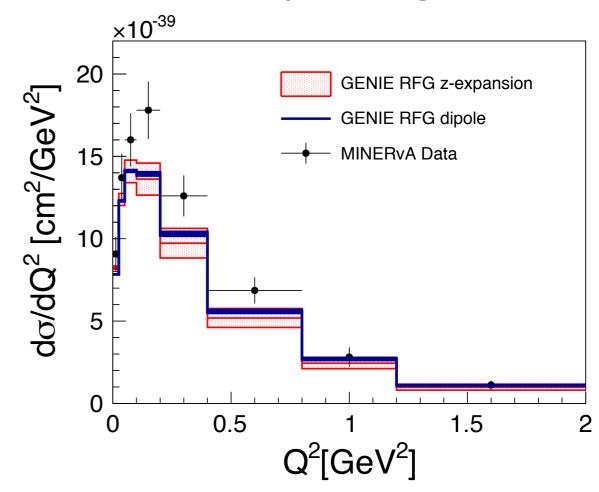
$$F_A(q^2) = \sum_{k=0}^{k_{\text{max}}} a_k z(q^2)^k$$

• where the expansion coefficients ak are dimensionless numbers representing nucleon

structure information

Derived from first principles of QCD

Extensively used in meson decay

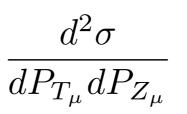


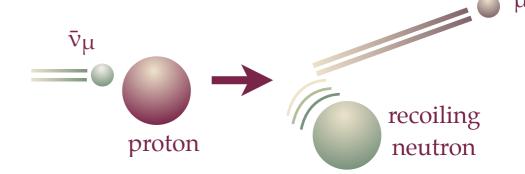
Aaron Meyer, MB, Richard Gran and Richard Hill, Phys. Rev D93(2016)

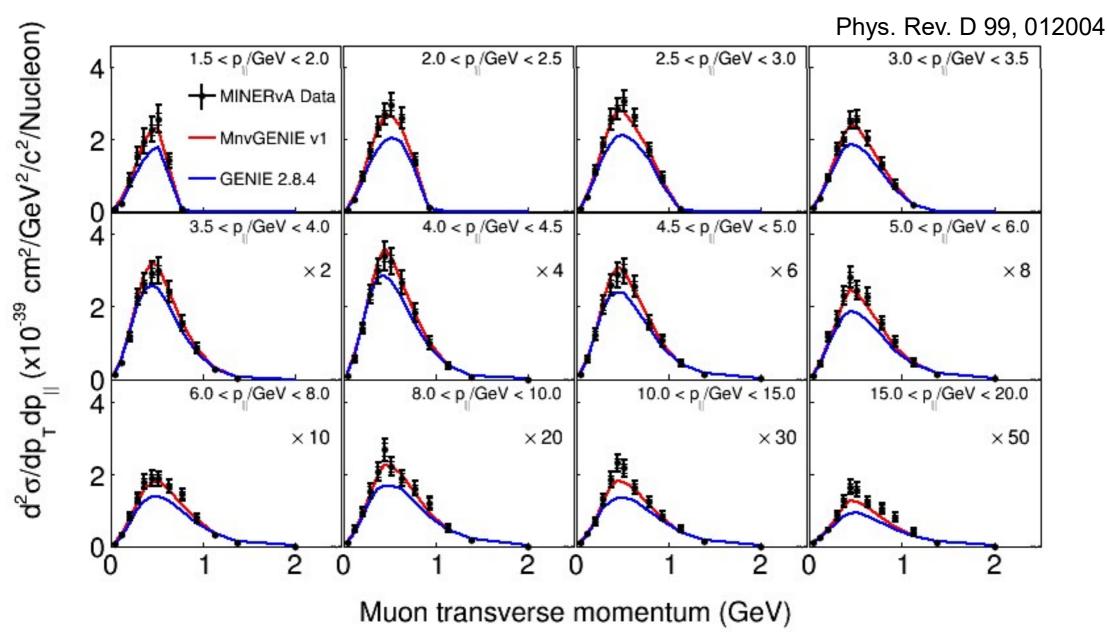


## Double Differential Cross Sections (Antineutrinos and Neutrinos)

- Using the kinematics of the muon
- Data agrees with the simulation that includes nuclear effects (RPA and 2p2h)









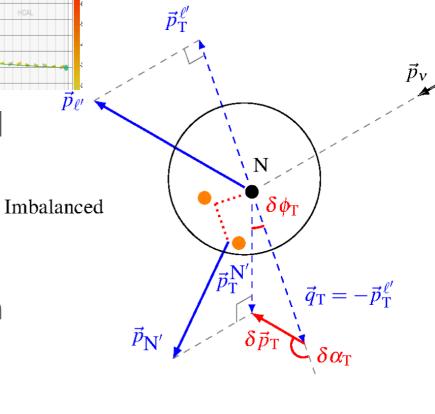
**Fermilab** 

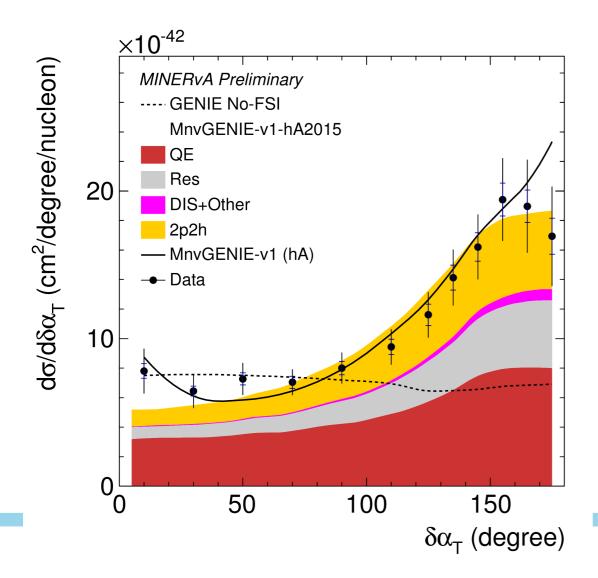
### **MINERvA New Observables**

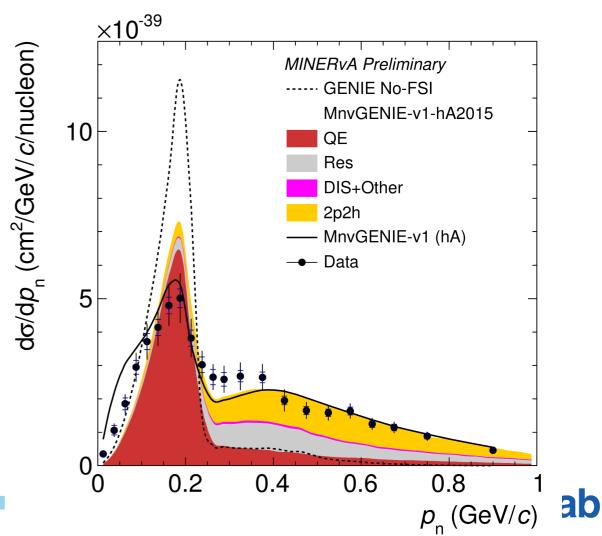
• Probing final state nuclear effects using the transverse imbal of the leading proton and the lepton

• First differential cross section in initial struck neutron momentum

- Useful to constrain initial nuclear effects, deficiencies in the at low and mid region of the neutron momentum







Phys.Rev.Lett. 121 (2018) no.2, 022504

## Joint Group Between Theorists and Experimentalists at Fermilab

#### What we have

- Rich neutrino program at Fermilab
- Remaining questions of neutrino oscillation
- Neutrino phenomenology
- HEP theory
- Experimentalists
- Software expertise (collider, lattice, V)

#### What we need

- Communication between theorists and experimentalists
- Accurate models
- New models to be developed and incorporated in the simulations
- Detailed understanding of each of the component of theory

- Walter Giele, Adi Ashkenazi and Minerba Betancourt are leading the effort
- Members of the group are Fermilab scientists, distinguish Fellows, intensity Frontiers and neutrino physics center fellows



## Joint group between theorists and Experimentalists

- We have been meeting once a month since 2017
  - Details at Indico https://indico.fnal.gov/category/724/
- We identified specific topics for the working groups and started to work
- Interfacing theory and GENIE:
  - Nuclear ab initio
  - DIS
  - Radiative corrections and nue/numu cross section differences
  - Lattice QCD
- Phenomenology
- We welcome new collaborators!



## **Summary**

- Several challenges from the theoretical model side and experimental side to understanding neutrino interactions
- Many theoretical developments to model meson exchange current and nuclear long range correlations
- More development needed for pion production and deep inelastic scattering
- Rich set of new cross section measurements from MINERvA, MicroBooNE and NOvA
  - Data -model disagreement across many experiments
- We are learning a lot from neutrino-nucleus interactions and building a rich set of cross section results for the oscillation experiments
- Oscillation experiments depend on modeling nuclear effects correctly and knowledge of cross sections to a few percent for precision oscillation measurements

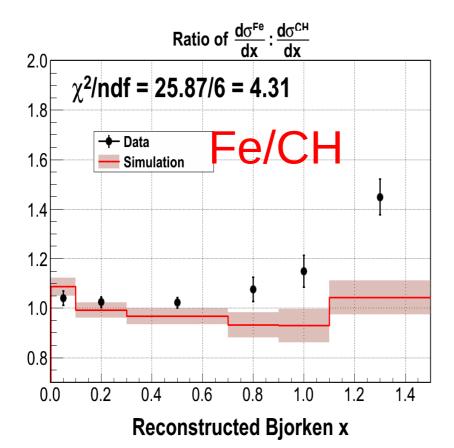


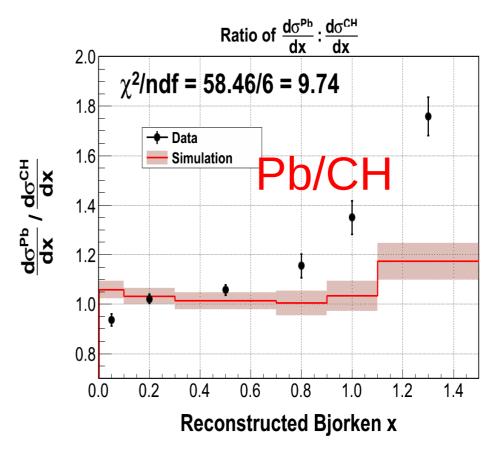
# **BackUp Slides**



#### MINERvA Inclusive Cross Sections On Different Nuclei

- MINERvA measured inclusive charged current cross section ratios as a function of Bjorken x  $x = \frac{Q^2}{2ME_{had}}$
- x corresponds to the fraction of the initial nucleon's momentum that is carried by the struck quark
- Model of nuclear effects is wrong for heavier nuclei
- The high X region greater than one is dominated by short range correlation



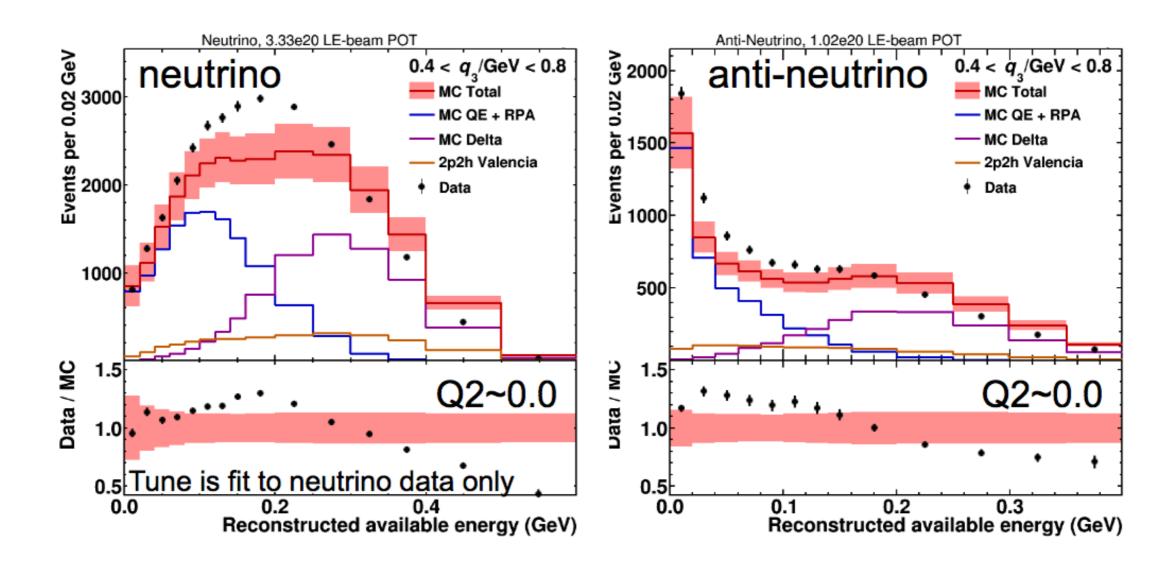


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#### The MINERvA Tune

• Adding in models of multi nucleon and RPA improves agreement in some regions



- Performed an empirical fit to scale up the multi nucleon contribution
- The fit is performed using neutrino data, then applied in antineutrino analysis



#### The MINERvA Tune

• The tune from neutrino data also agrees with antineutrino data

