



# Signatures of ULTRALIGHT SCALARS in neutrino oscillation experiments

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in collaboration with **Abhish Dev** and **Pedro A. N. Machado** 

First EuCAPT Annual Symposium Based on JHEP 01 (2021) 094

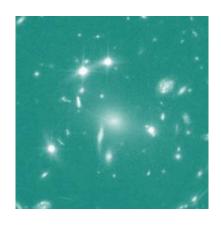




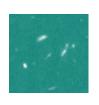








Neutrino oscillation experiments could dectect ultralight scalars via modulations of the signal with time and distorted neutrino oscillations.



## ULTRALIGHT SCALARS and NEUTRINOS

An ultralight scalar field,

$$\phi(x,t) \simeq \frac{\sqrt{2\rho_{\phi}}}{m_{\phi}} \sin\left[m_{\phi}(t - \vec{v} \cdot \vec{x})\right]$$

A. Berlin.

PRL 117 (2016) 23, 231801

G.Krnjaic et al.

PRD 97 (2018) 7, 075017

V. Brdar et al.

PRL117 (2016) 23, 231801

which couples effectively to neutrinos,

$$\mathcal{L}_{\text{eff}} = -m_{\nu} \left( 1 + y \frac{\phi}{\Lambda} \right) \bar{\nu} \nu + \text{h.c.}$$

induces a time dependence on the mass splittings (also on the mixing angles...)

$$\Delta m_{ij}^2(t) \equiv m_i^2(t) - m_j^2(t) \simeq \Delta m_{ij}^2 \left[ 1 + 2\eta \cos(m_\phi t) \right]$$

# Time between two events (inverse of the event rate)

#### **Running time**

$$au_{\phi} \sim au_{\nu} \sim 10^{\text{-}2}$$
 -  $10^{\text{-}3} \; \mathrm{s}$ 

$$\tau_{\phi} \sim \tau_{\rm evt} \sim 1 \, \, {\rm day}$$

$$\tau_{\phi} \sim \tau_{\rm exp} \sim 10 \text{ years}$$

$$m_{\phi}$$
 ~10<sup>-13</sup> –10<sup>-14</sup> eV

$$m_\phi$$
 ~10<sup>-20</sup> eV

$$m_\phi$$
 ~10<sup>-23</sup> eV

The **mass** of the ultralight scalar is related to the **modulation period** 

$$\tau_{\phi} \equiv \frac{2\pi\hbar}{m_{\phi}} = 0.41 \left(\frac{10^{-14} \text{ eV}}{m_{\phi}}\right) \text{ seconds.}$$

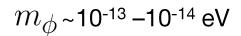
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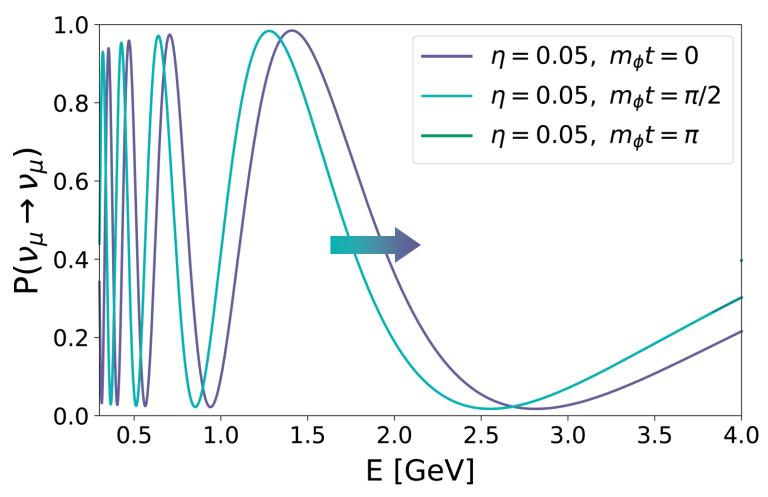
$$m_\phi$$
 ~10<sup>-23</sup> eV

#### A SIGNAL TIME MODULATION

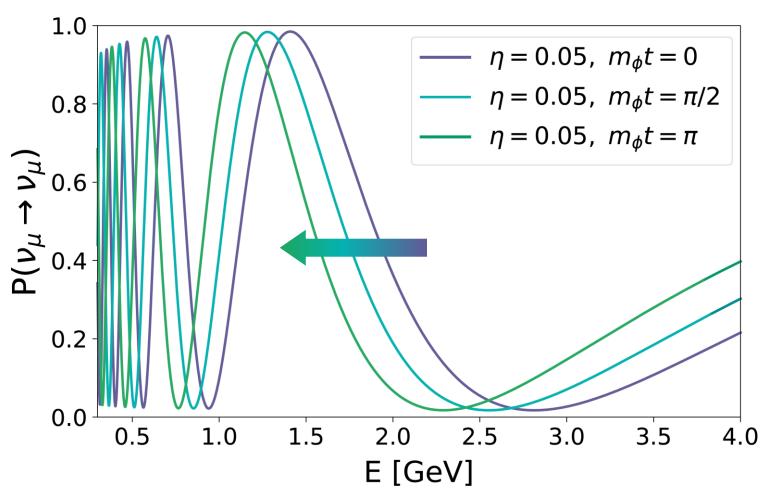
could be measured in experiments with large statistics

 $au_{\phi}$ 

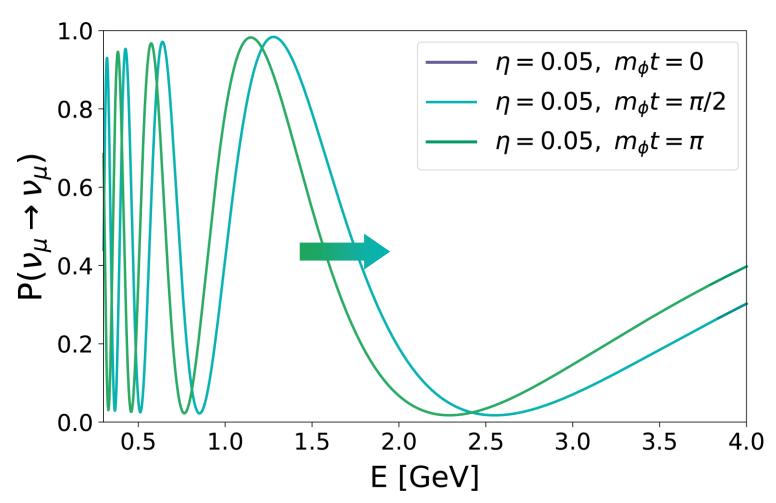
In **DUNE** for a modulating mass splitting only,  $\Delta m^2_{31}$  .



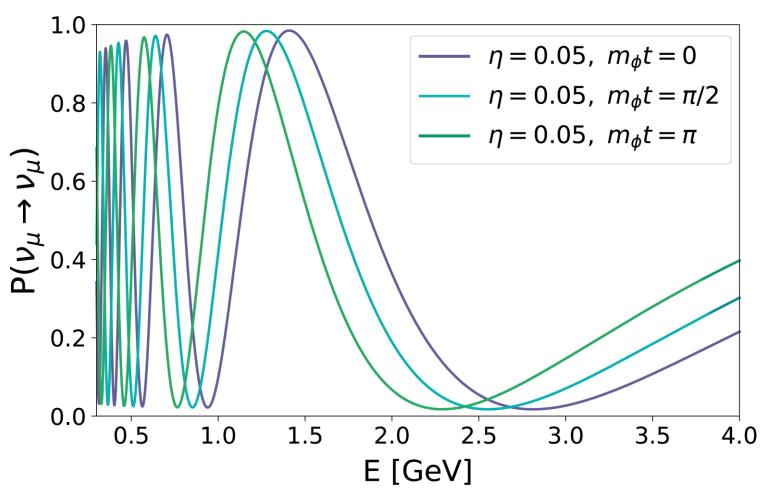
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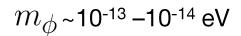
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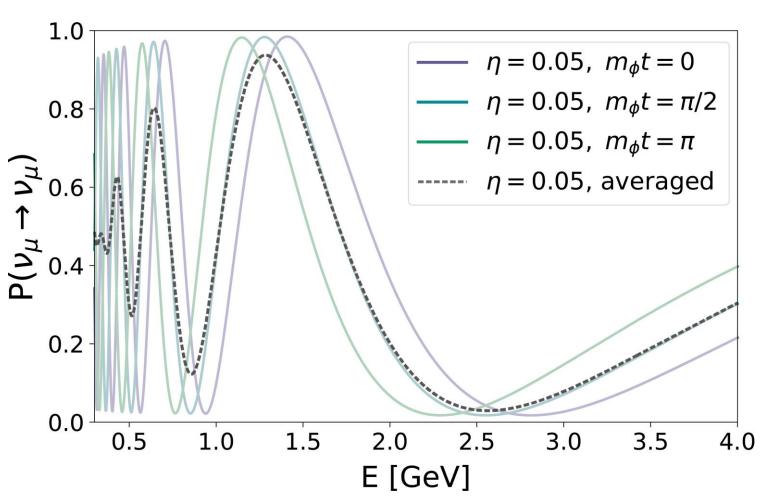
$$m_\phi$$
 ~10<sup>-20</sup> eV

$$m_\phi$$
 ~10<sup>-23</sup> eV

 $au_{\phi}$ 

Modulation gets averaged leading to **DISTRORTED NEUTRINO OSCILLATIONS** 

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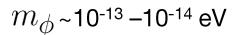
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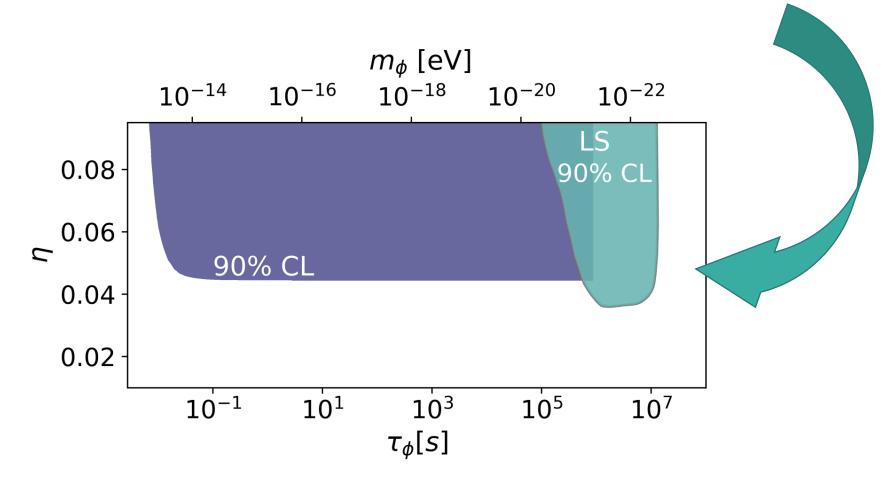


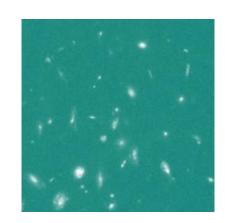
Modulation becomes too fast and manifests as a **fast-varying matter potential** 



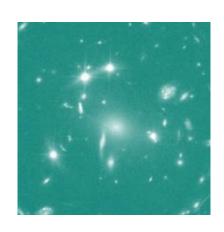
and eventually one can not distinguish it from standard oscillations.

For a more detailed discussion of the signatures and a particular example (DUNE), see <u>JHEP 01 (2021) 094</u>.





### TAKE HOME MESSAGE



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