

Initial Stages 2021

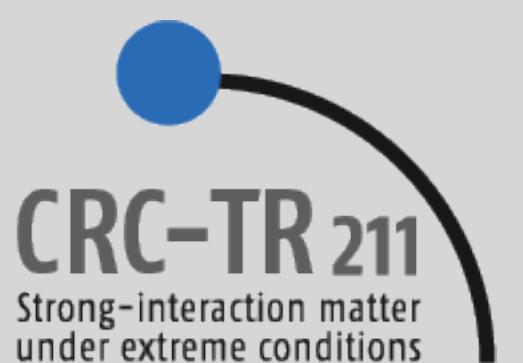
On the way to collectivity in rarely interacting systems

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In collaboration with
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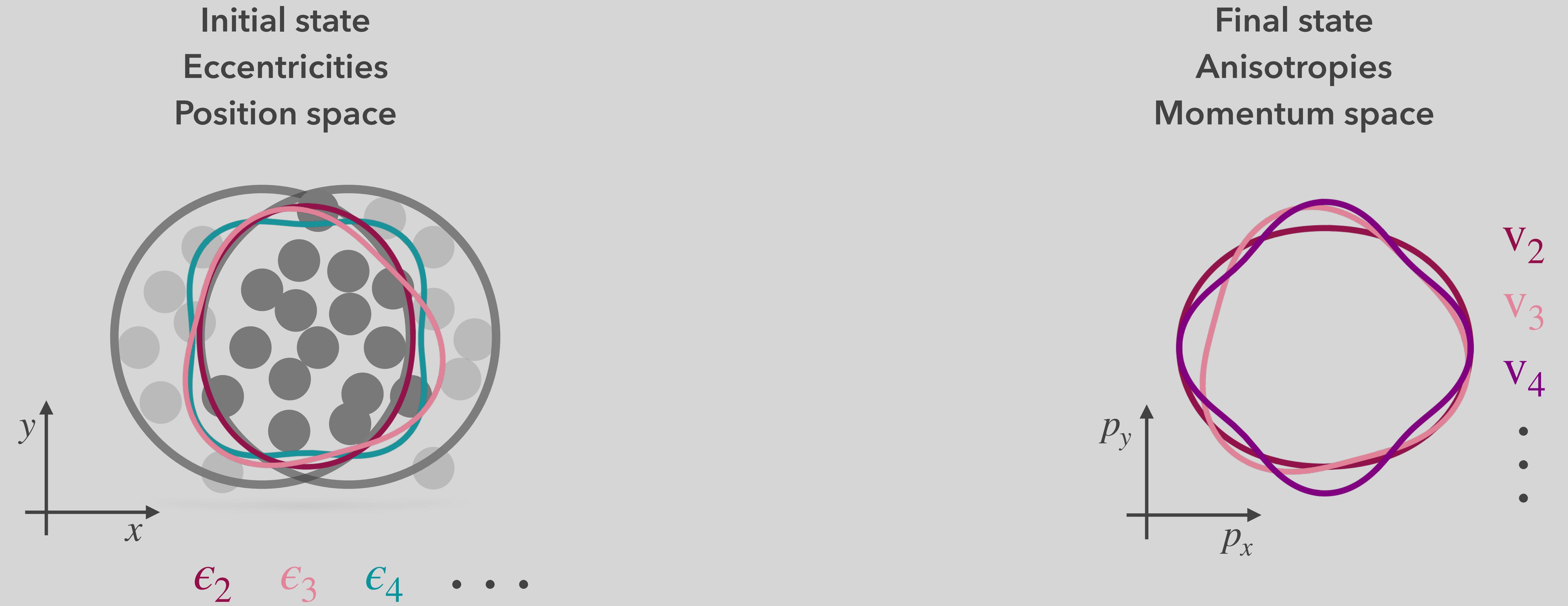


DFG

Motivation

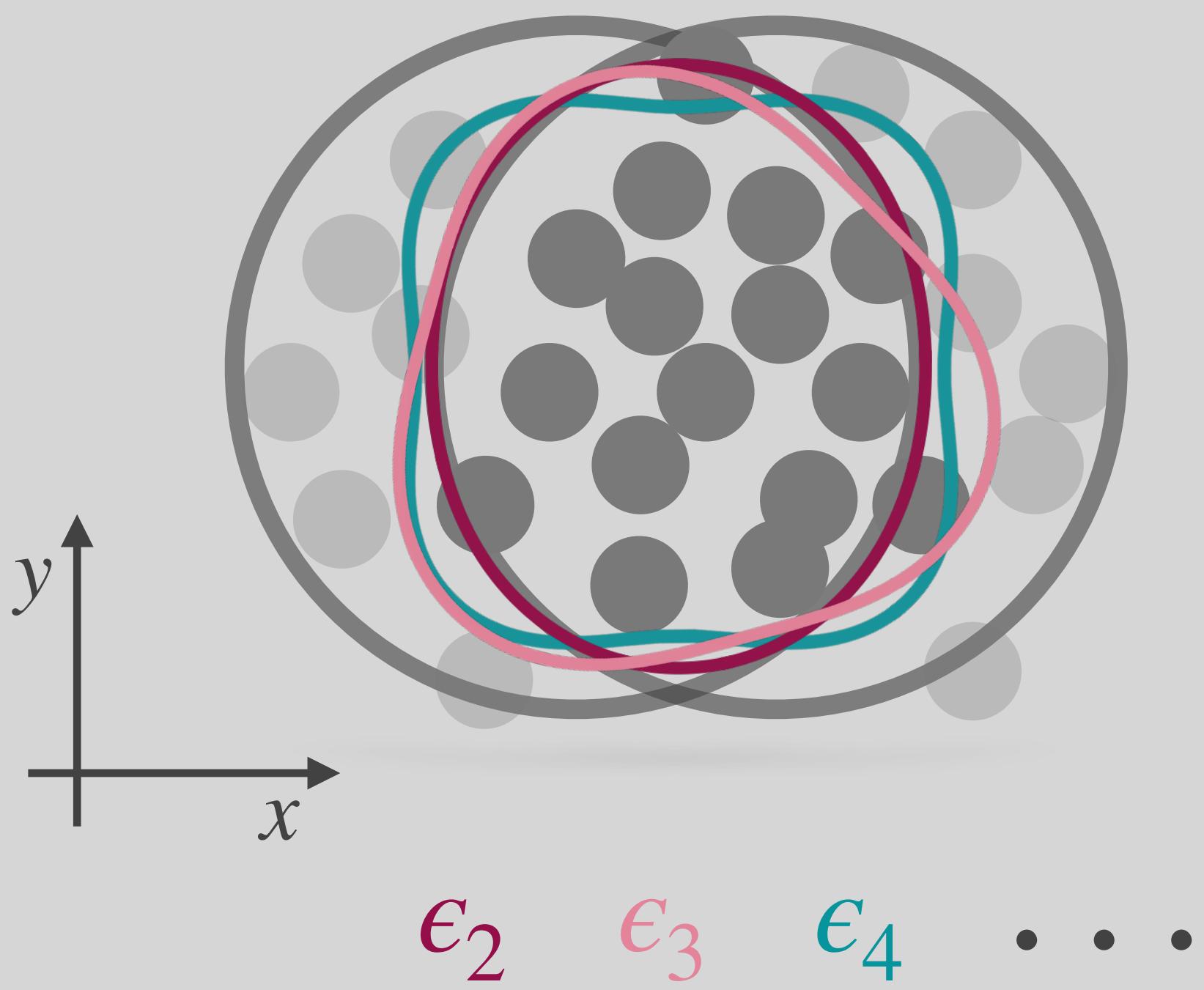


Link between initial eccentricities and flow harmonics



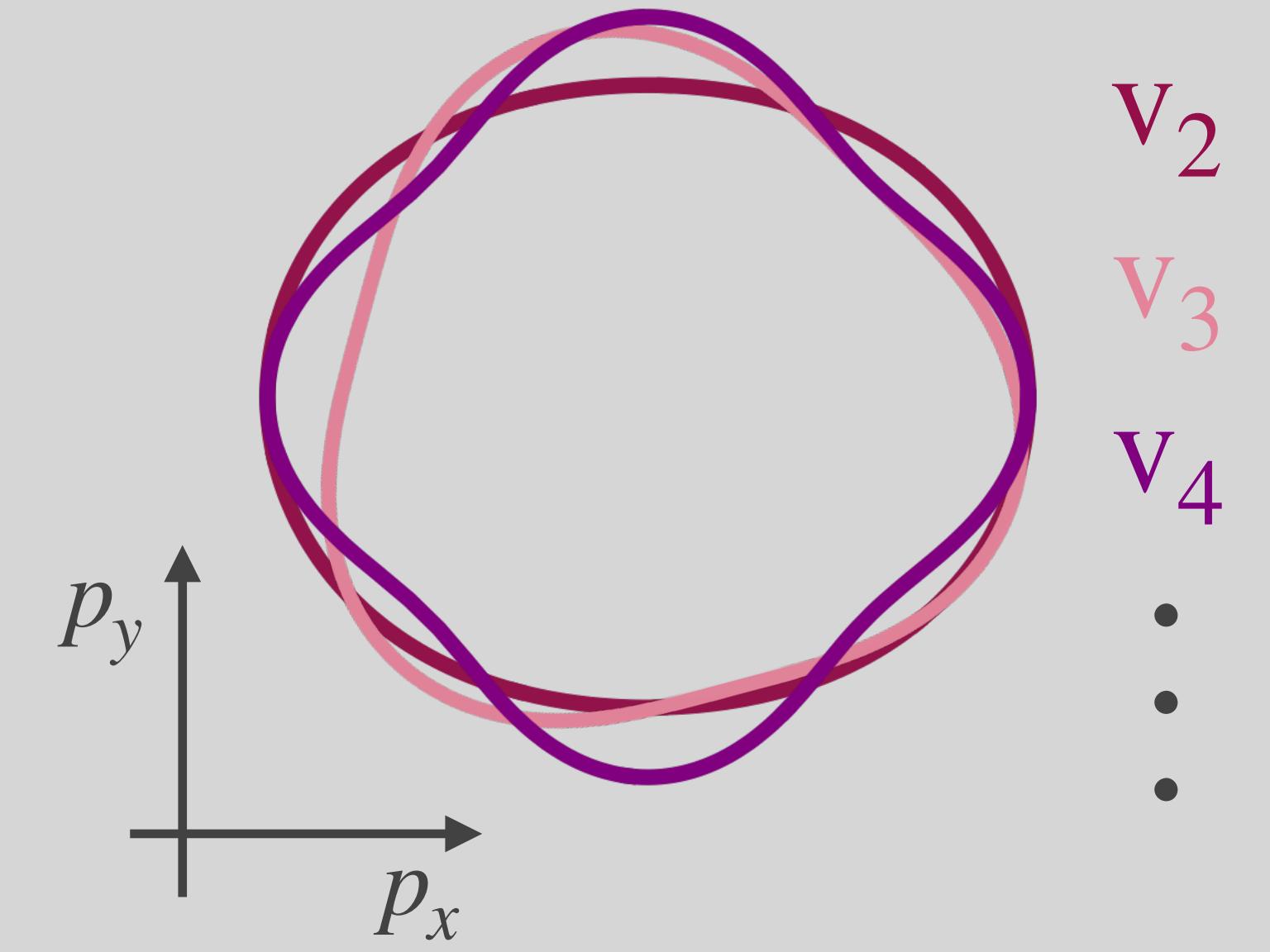
Link between initial eccentricities and flow harmonics

Initial state
Eccentricities
Position space



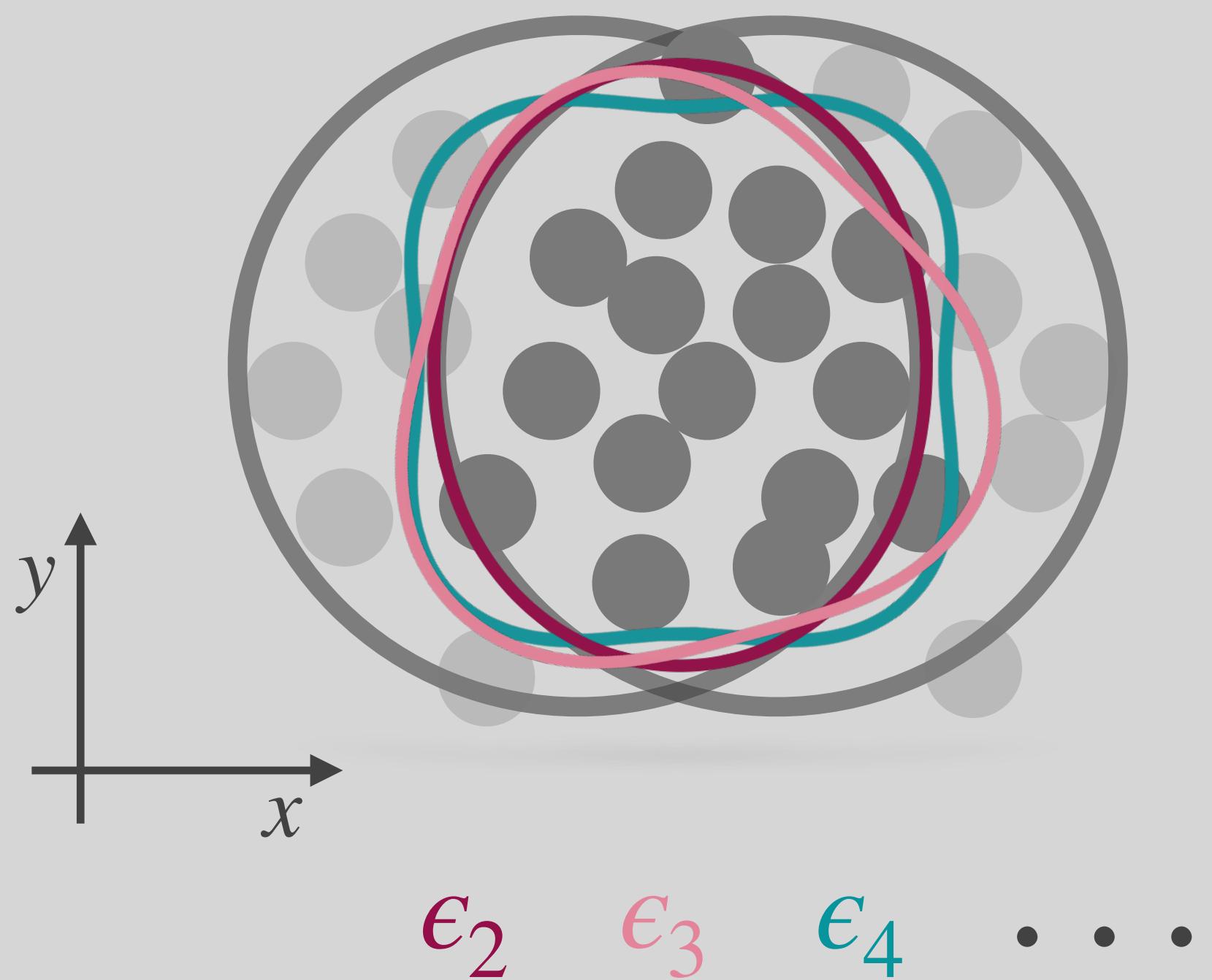
Theoretical
description

Final state
Anisotropies
Momentum space



Link between initial eccentricities and flow harmonics

Initial state
Eccentricities
Position space



Theoretical
description

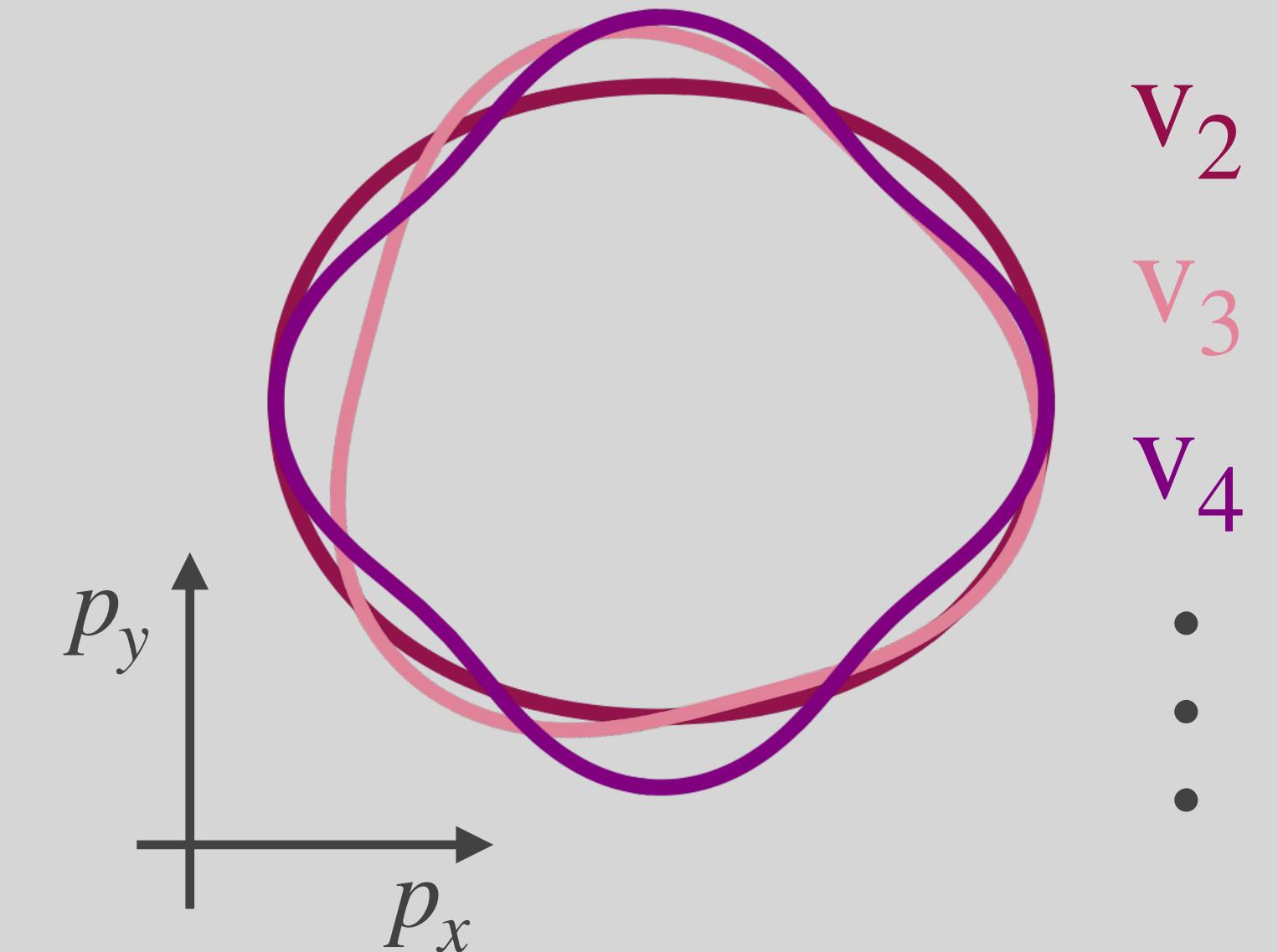
Hydrodynamics
Collective behaviour

$$N_{\text{resc}} \gg 1$$

Kinetic theory
Few collision regime

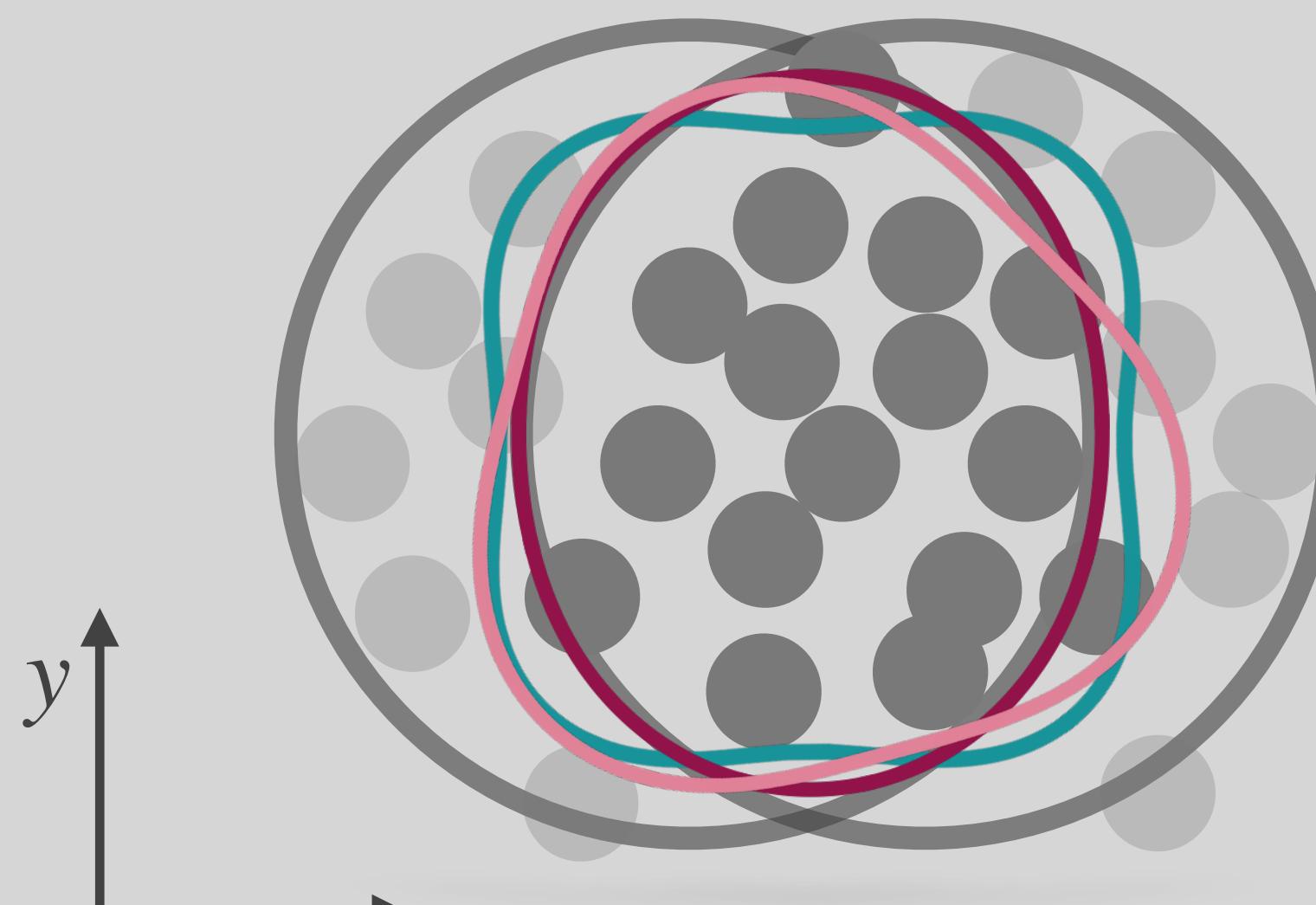
$$N_{\text{resc}} < 1$$

Final state
Anisotropies
Momentum space



Link between initial eccentricities and flow harmonics

Initial state
Eccentricities
Position space



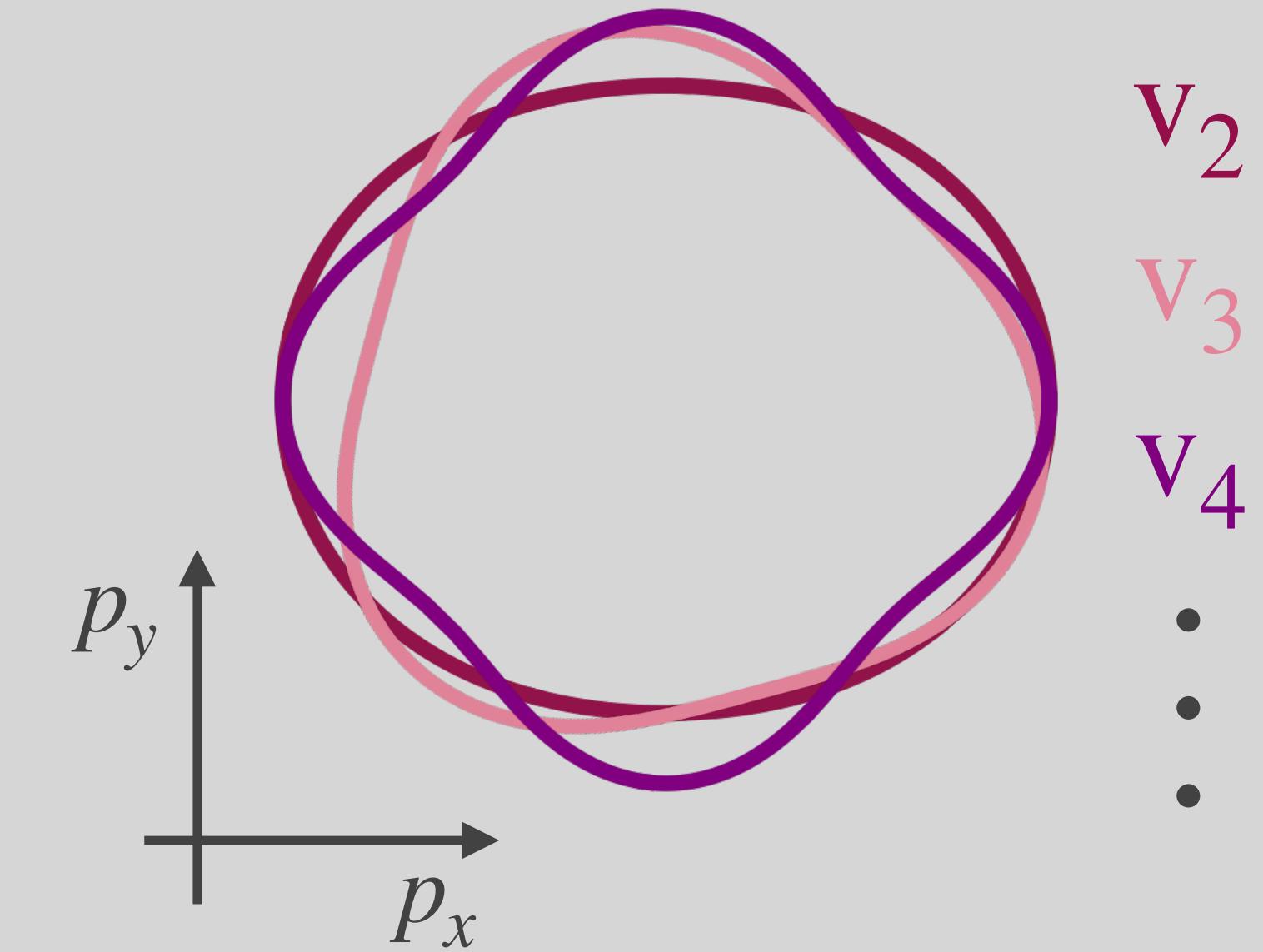
ϵ_2 ϵ_3 ϵ_4 \dots



Hydrodynamics
Collective behaviour
 $N_{\text{resc}} \gg 1$

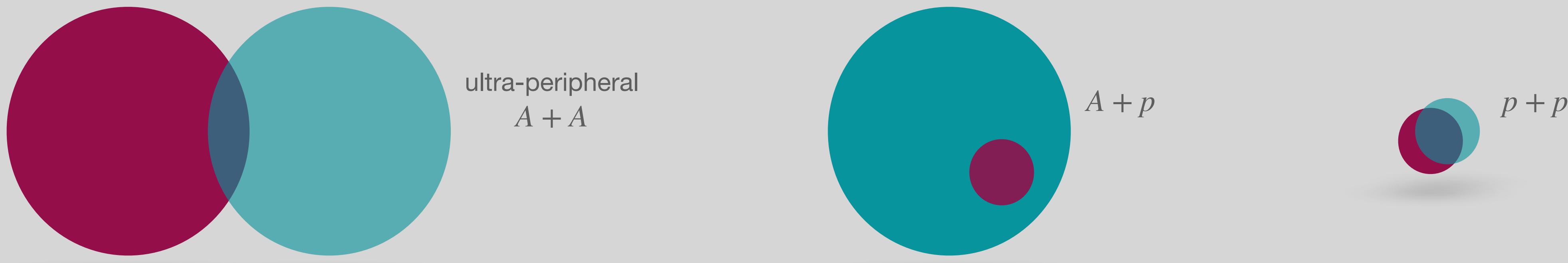
Kinetic theory
Few collision regime
 $N_{\text{resc}} < 1$

Final state
Anisotropies
Momentum space



Reasons for kinetic theory approach

Small and dilute systems



Question to be answered

for our approaches

How do flow harmonics v_n evolve in time with dependence on
the various initial eccentricities ϵ_m ?

Methods

Analytical approach

to calculate time-dependent v_n

Analytical
model

Calculate time-dependent anisotropic flow coefficients

$$v_n(t, p_\perp) = \frac{\iint \cos(n(\phi - \Psi_n)) f(t, \mathbf{x}, p_\perp, \phi) d^2\mathbf{x} d\phi}{\iint f(t, \mathbf{x}, p_\perp, \phi) d^2\mathbf{x} d\phi}$$

Analytical approach

to calculate time-dependent v_n

- Calculate time-dependent AFC

$$v_n(t, p_\perp) = \frac{\iint \cos\left(n(\phi - \Psi_n)\right) f(t, \mathbf{x}, p_\perp, \phi) d^2\mathbf{x} d\phi}{\iint f(t, \mathbf{x}, p_\perp, \phi) d^2\mathbf{x} d\phi}$$

Analytical
model

**Loss term
in Boltzmann
equation**

Use classical relativistic Boltzmann equation as equation of motion

$$[\partial_t + \mathbf{v} \cdot \nabla] f(t, \mathbf{x}, \mathbf{p}) = C_{coll} [f(t, \mathbf{x}, \mathbf{p})]$$

with loss term of 2-to-2 elastic collision kernel

Analytical approach

to calculate time-dependent v_n

- Calculate time-dependent AFC

$$v_n(t, p_\perp) = \frac{\iint \cos\left(n(\phi - \Psi_n)\right) f(t, \mathbf{x}, p_\perp, \phi) d^2\mathbf{x} d\phi}{\iint f(t, \mathbf{x}, p_\perp, \phi) d^2\mathbf{x} d\phi}$$

- Use

$$[\partial_t + \mathbf{v} \cdot \nabla_{\mathbf{x}}] f = 0$$

Few collision limit

Dependence on free-streaming distribution function

$$f^{(0)}(t, \mathbf{x}, \mathbf{p}) = f^{(0)}(0, \mathbf{x} - t\mathbf{v}, \mathbf{p})$$

Analytical model

Loss term

in Boltzmann equation

$$\mathcal{O}(N_{\text{resc}}) = 1$$

Analytical approach

to calculate time-dependent v_n

- Calculate time-dependent AFC

$$v_n(t, p_\perp) = \frac{\iint \cos(n(\phi - \Psi_n)) f(t, \mathbf{x}, p_\perp, \phi) d^2\mathbf{x} d\phi}{\iint f(t, \mathbf{x}, p_\perp, \phi) d^2\mathbf{x} d\phi}$$

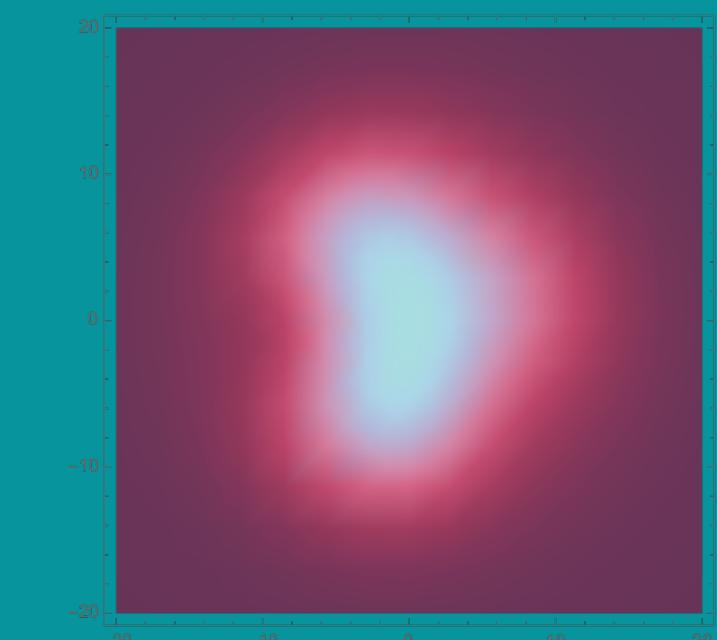
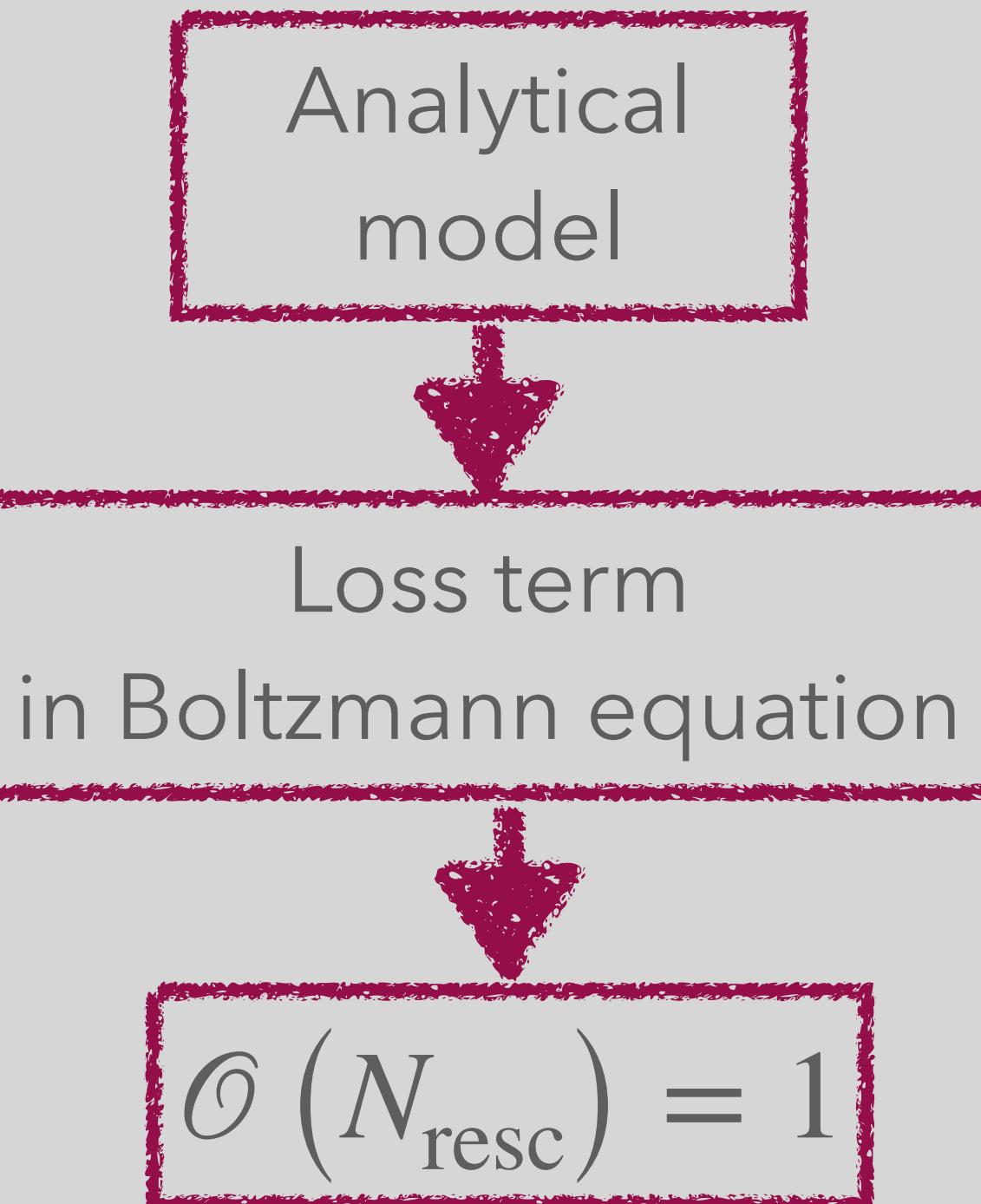
- Use classical relativistic Boltzmann equation as equation of motion

$$[\partial_t + \mathbf{v} \cdot \nabla] f(t, \mathbf{x}, \mathbf{p}) = C_{coll} [f(t, \mathbf{x}, \mathbf{p})]$$

with loss

Initial distribution function (including initial eccentricities) as input
for my calculation

- Few constraints
- Dependence



Analytical approach

to calculate time-dependent v_n

- Calculate time-dependent AFC

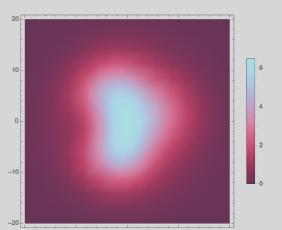
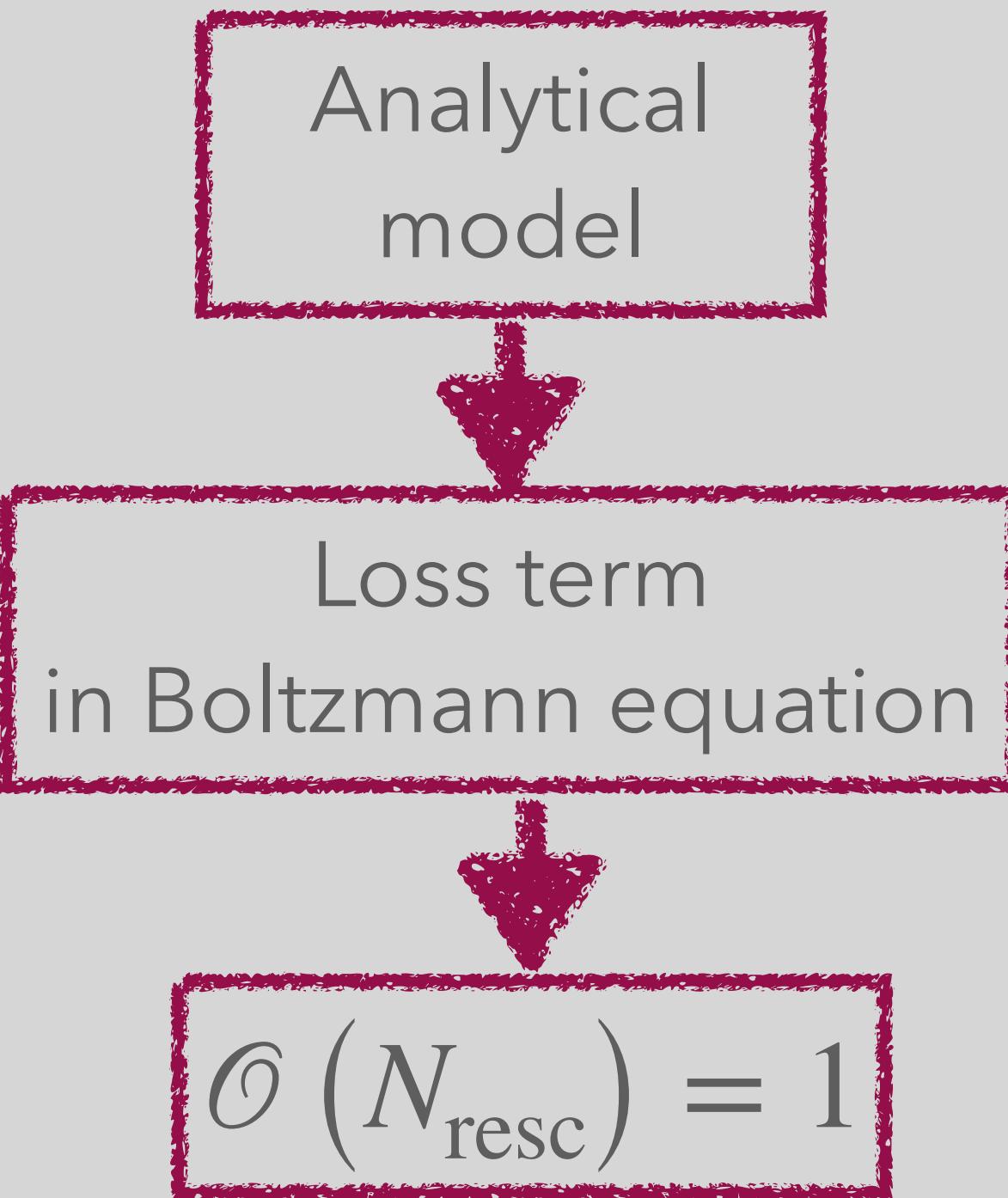
$$v_n(t, p_\perp) = \frac{\iint \cos(n(\phi - \Psi_n)) f(t, \mathbf{x}, p_\perp, \phi) d^2\mathbf{x} d\phi}{\iint f(t, \mathbf{x}, p_\perp, \phi) d^2\mathbf{x} d\phi}$$

- Use classical relativistic Boltzmann equation as equation of motion

$$[\partial_t + \mathbf{v} \cdot \nabla] f(t, \mathbf{x}, \mathbf{p}) = C_{coll} [f(t, \mathbf{x}, \mathbf{p})]$$

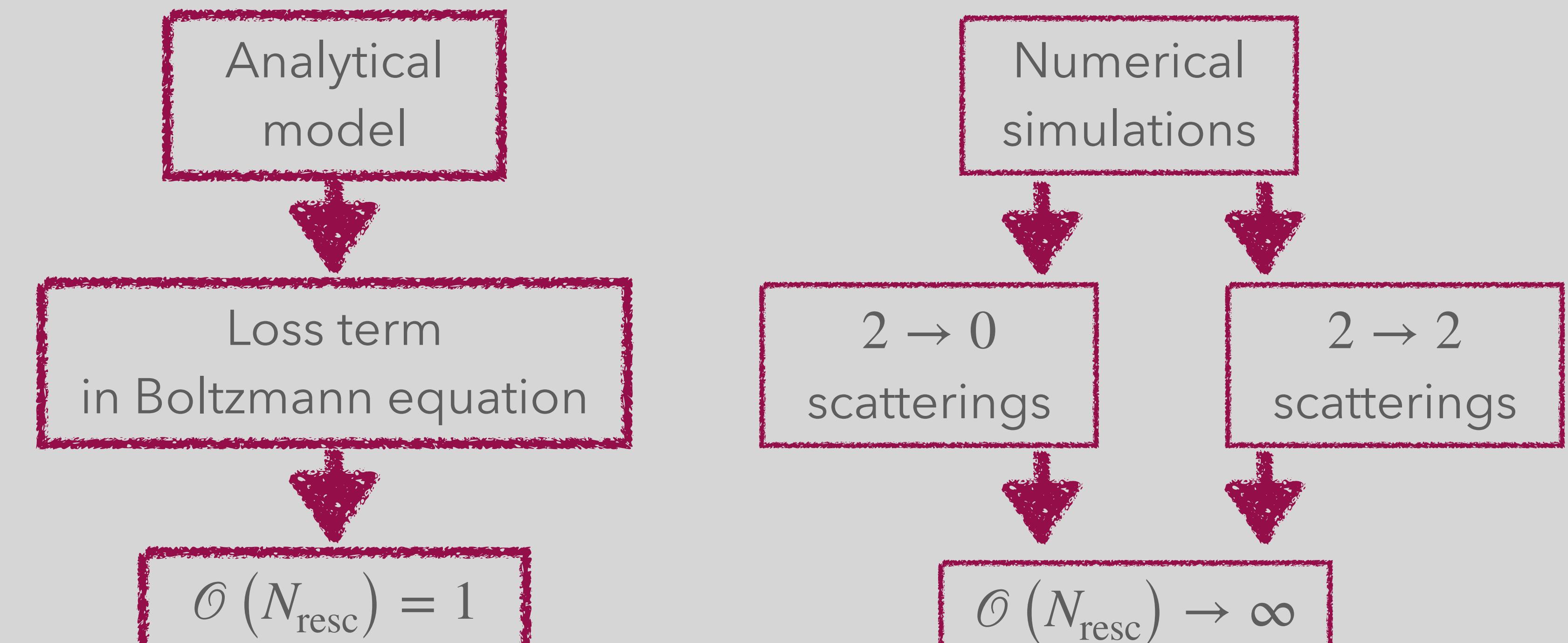
with loss term of 2-to-2 elastic collision kernel

- Few collision limit
- Dependence on free-streaming distribution function $f^{(0)}(t, \mathbf{x}, \mathbf{p}) = f^{(0)}(0, \mathbf{x} - t\mathbf{v}, \mathbf{p})$
- Initial distribution function (including initial eccentricities) as input for my calculation



Our approaches to calculate v_n

- Kinetic theory
- 2 dimensional
- Massless particles

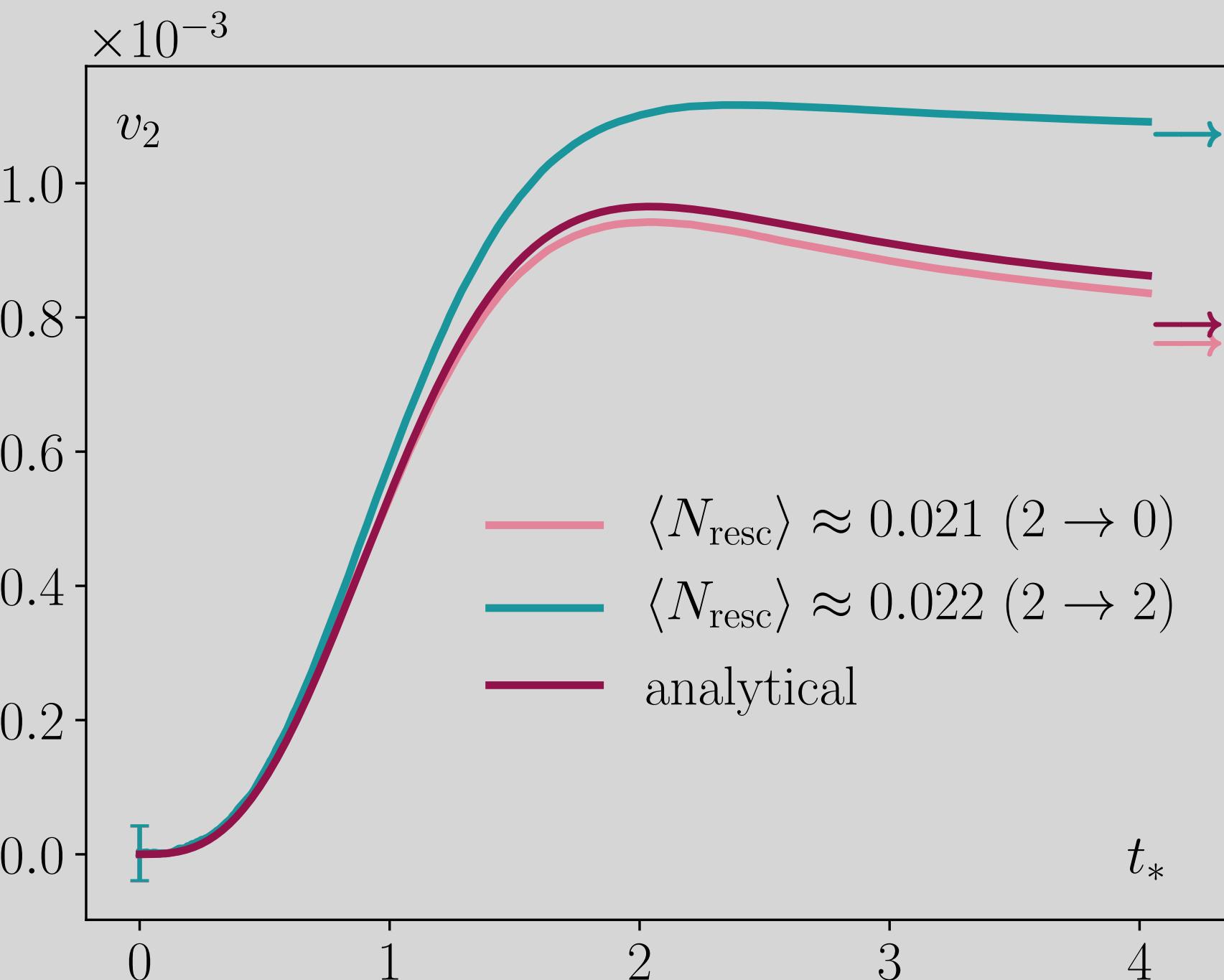


Hendrik Roch
Poster in session T

Results



v_2 from our approaches



Loss term dominates signal for $N_{\text{reesc}} \lesssim 0.35$

Higher orders in N_{reesc} can be neglected in few collision regime

Deviation increases with growing N_{reesc}

Good agreement for small N_{reesc}

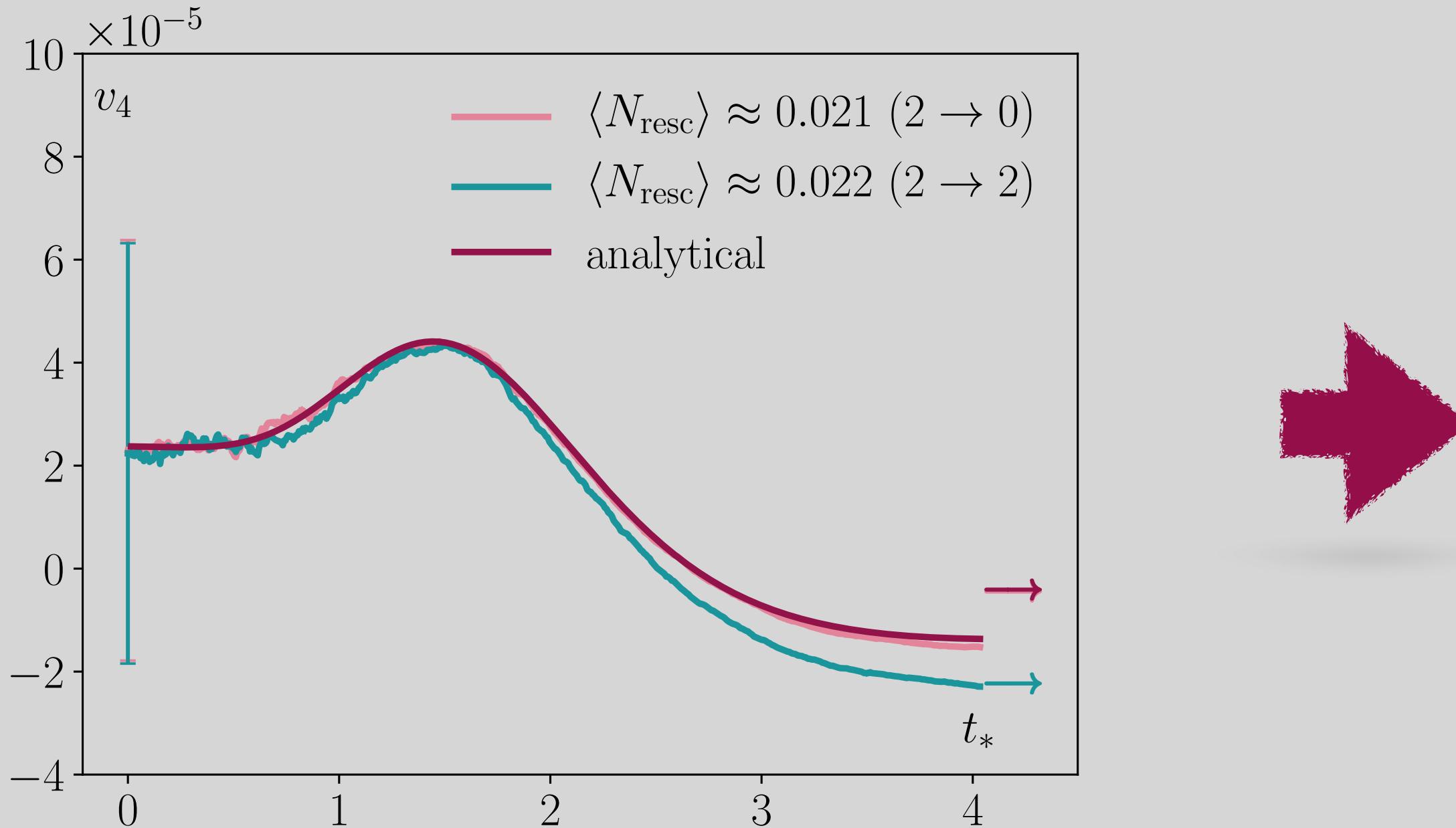
Expansion around small times $t \Rightarrow v_2 \propto \epsilon_2 t^3$

- Analytical formula (with $t_* \equiv t/R$)

$$v_2(t_*) = \frac{64 \sqrt{\pi} N_{\text{reesc}} \epsilon_2}{27 (8 + 3 \sqrt{2} \epsilon_2^2)} e^{-\frac{2 t_*^2}{3}} \left[-t_* I_0 \left(\frac{2 t_*^2}{3} \right) + \left(2 t_* + \frac{3}{t_*} \right) I_1 \left(\frac{2 t_*^2}{3} \right) \right]$$

v_4 from our approaches

with dependence on ϵ_4



Good agreement for small N_{resc}

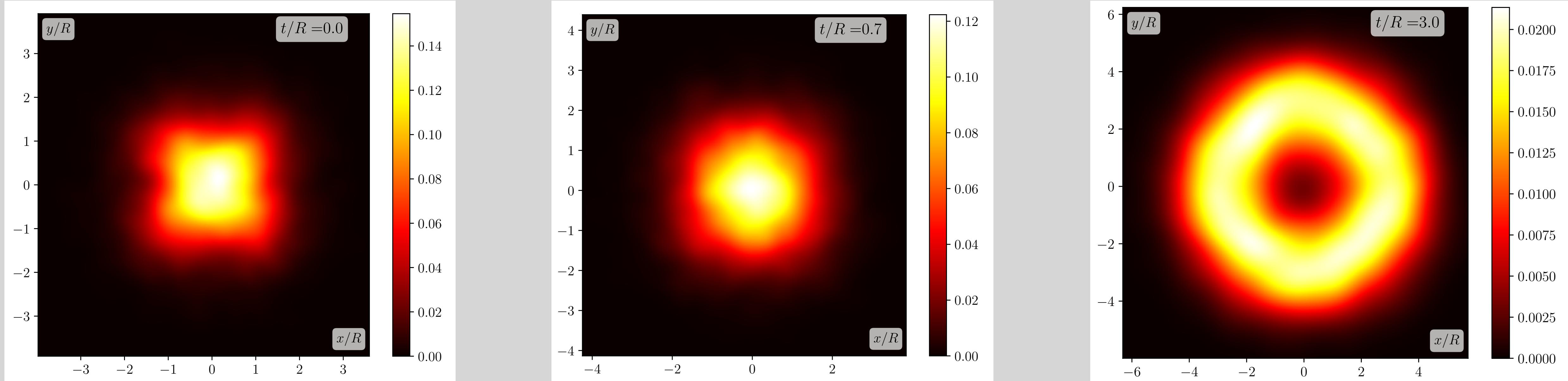
Expansion around small times t

$$\Rightarrow v_4 \propto \epsilon_4 t^5$$

- Analytical formula (with $t_* \equiv t/R$)

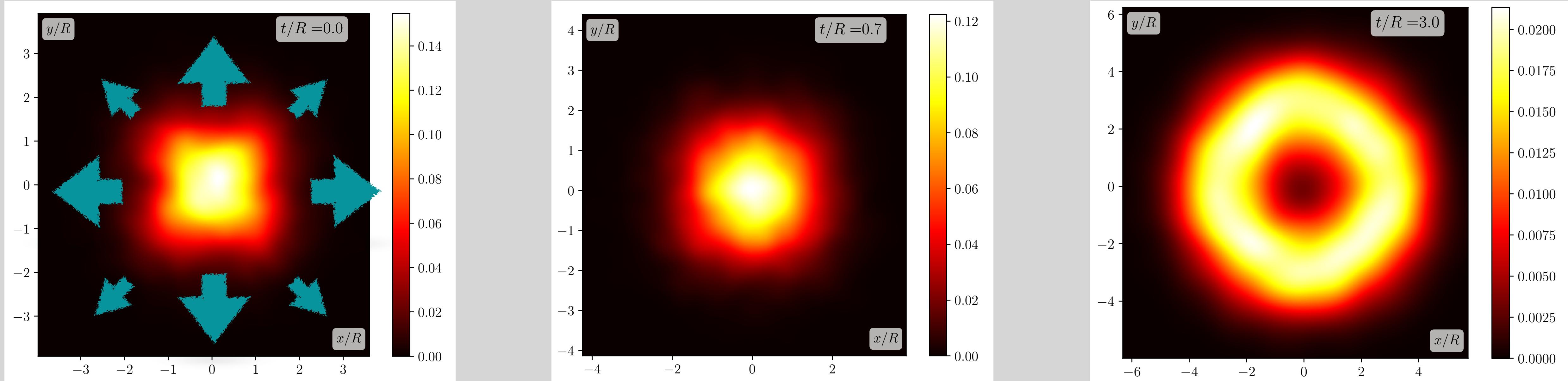
$$v_4(t_*) = \frac{2048 \sqrt{\pi} N_{\text{resc}} \epsilon_4}{405 (384 + 35 \sqrt{2} \epsilon_4^2)} e^{-\frac{2 t_*^2}{3}} \left[-\left(5 t_*^3 + 21 t_* + \frac{54}{t_*} \right) I_0\left(\frac{2 t_*^2}{3}\right) + \left(5 t_*^3 + 24 t_* + \frac{63}{t_*} + \frac{162}{t_*^3} \right) I_1\left(\frac{2 t_*^2}{3}\right) \right]$$

Density plots from numerical simulations



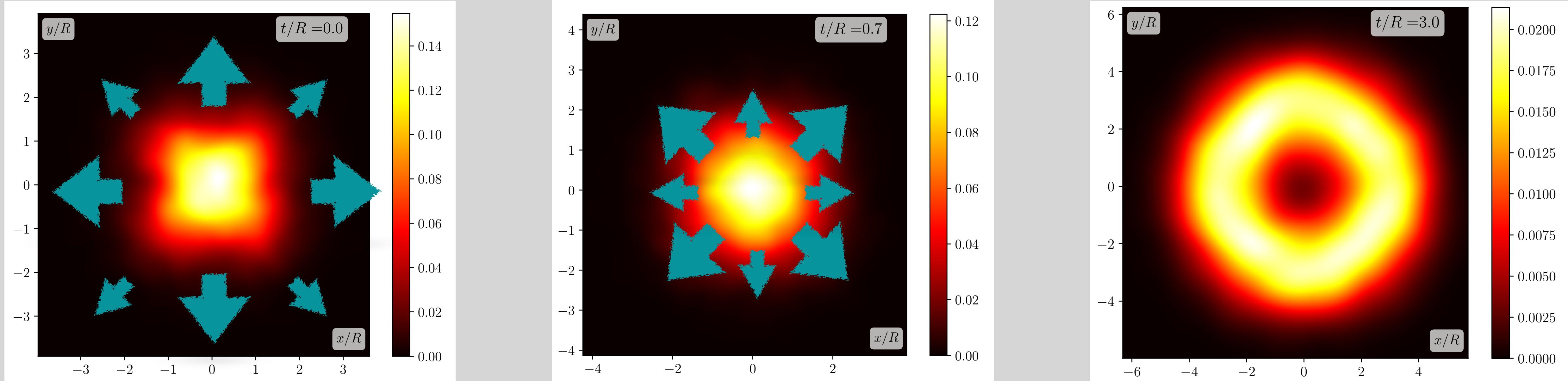
- Initial eccentricity ϵ_4 only
 - $\pi/4$ -rotated high density region \Rightarrow locally $\epsilon_4 < 0$
 - Dilute regions \Rightarrow small contribution to v_4
- More diluted system
- Denser regions still with $\epsilon_4 < 0$

Density plots from numerical simulations



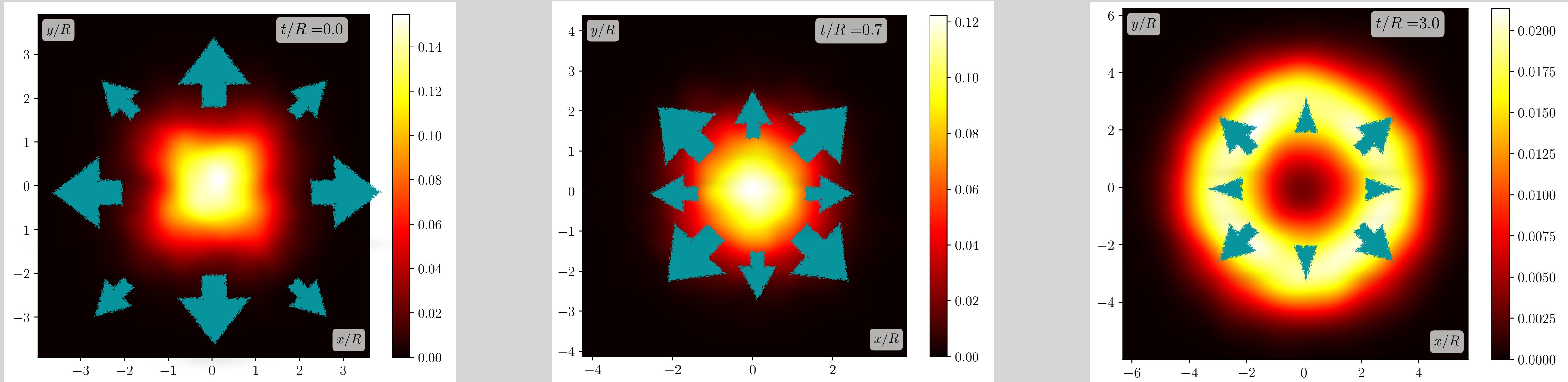
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Density plots from numerical simulations



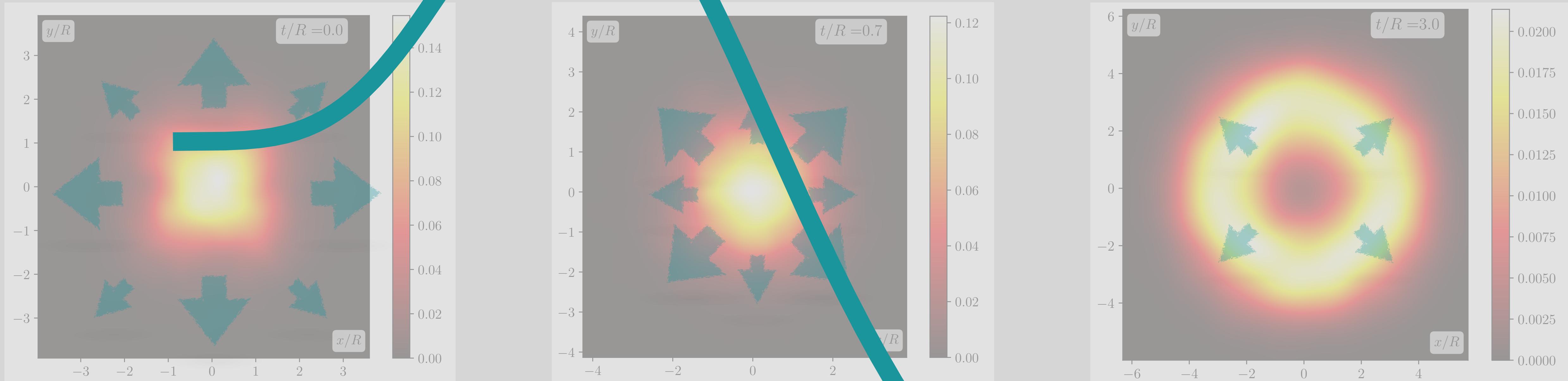
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Density plots from numerical simulations



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Density plots from numerical simulations



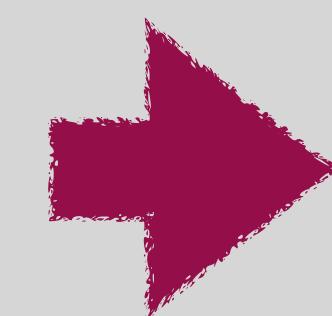
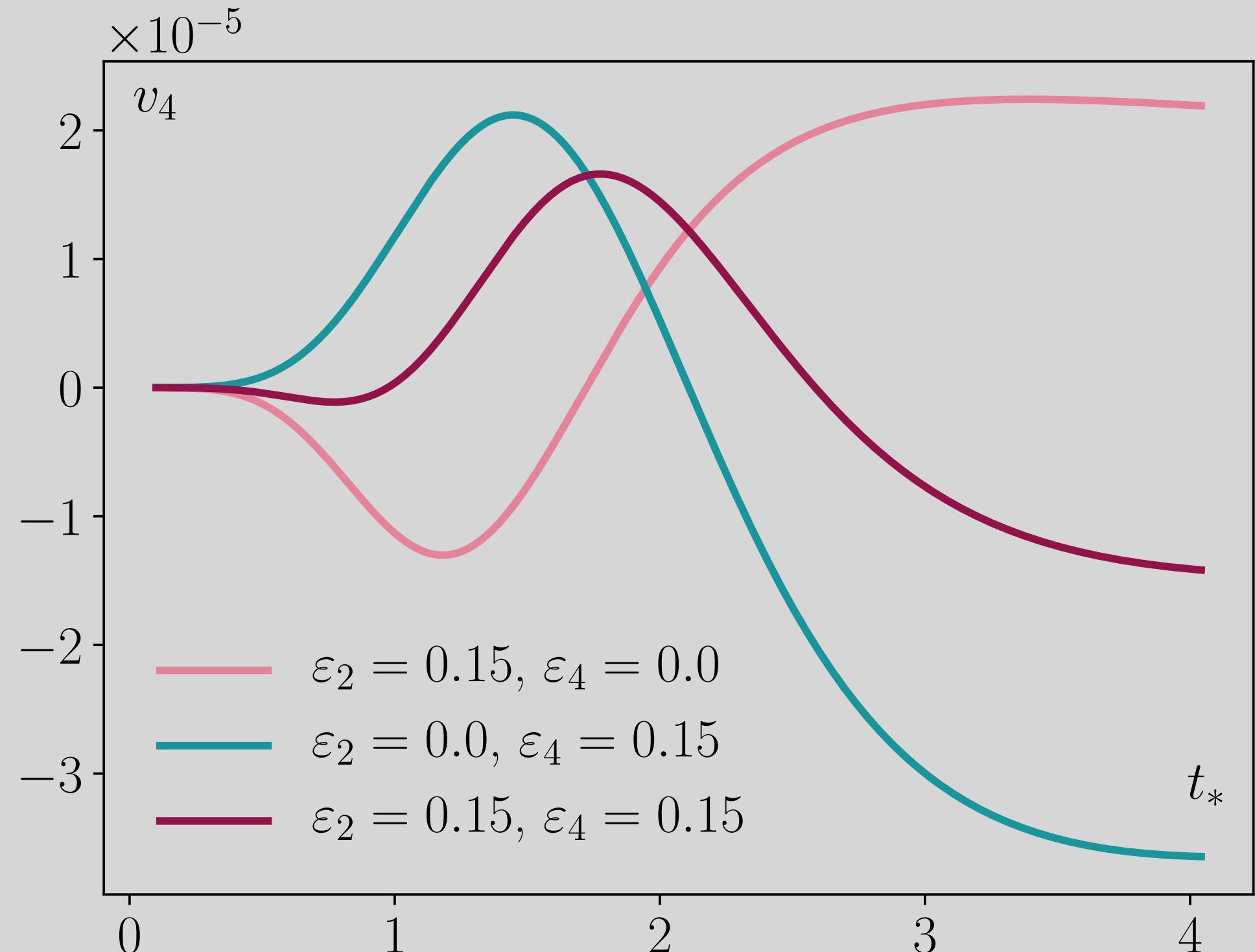
- Initial eccentricity ϵ_4 only

- $\pi/4$ -rotated high density region \Rightarrow locally $\epsilon_4 < 0$
- Dilute regions \Rightarrow small contribution to v_4

- More diluted system
- Denser regions still with $\epsilon_4 < 0$

$v_4(t_*)$

v_4 from our analytical approach with dependence on ϵ_2 and ϵ_4

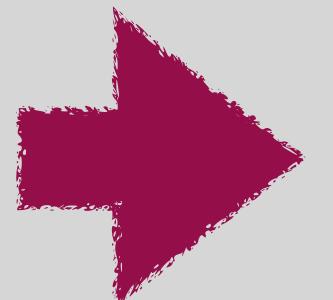
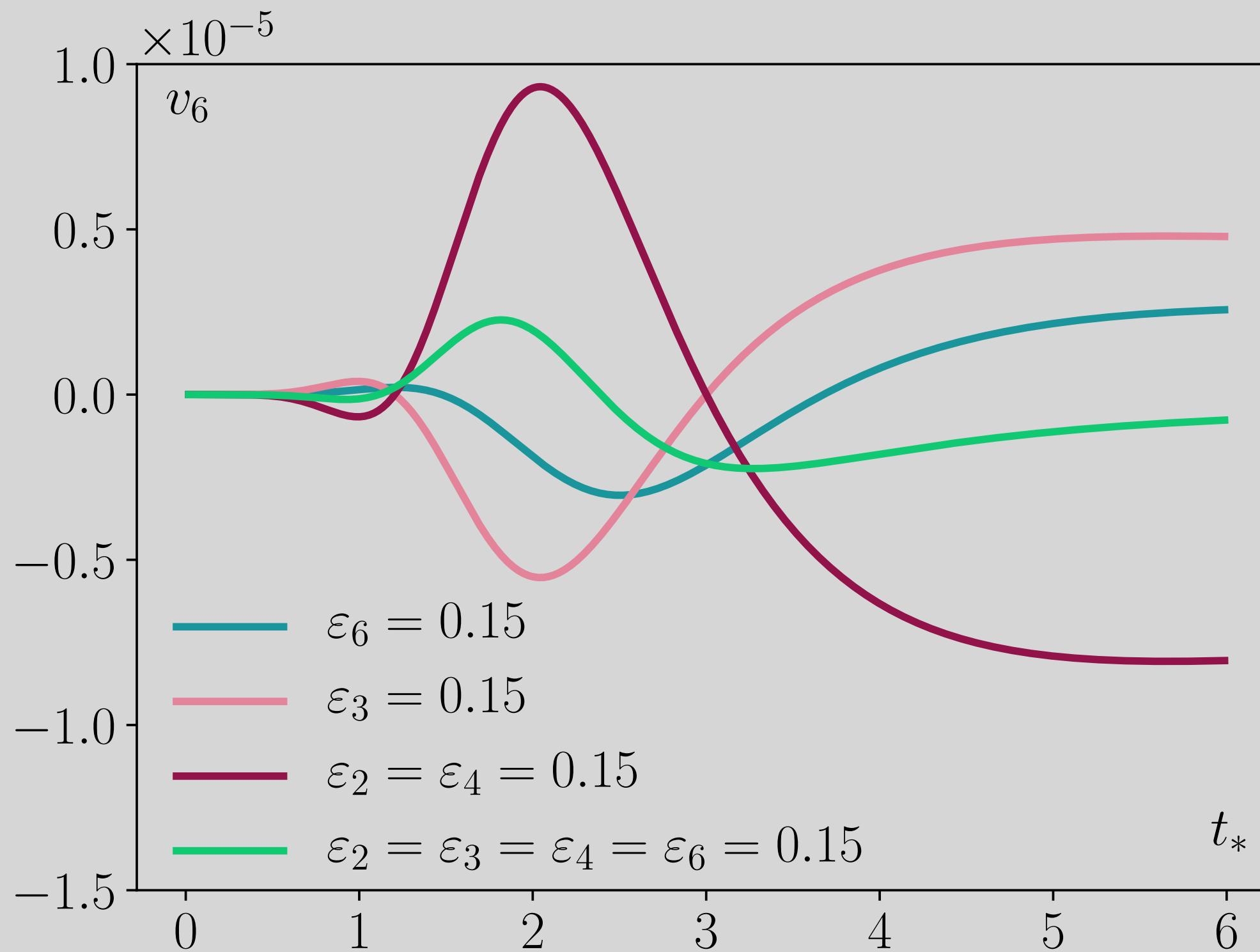


Two contributions from ϵ_4 and ϵ_2^2 to v_4

Oscillating v_4 with dependence on time t

Contributions compensate each other

v_6 from our analytical approach with dependence on $\epsilon_2, \epsilon_3, \epsilon_4$ and ϵ_6



Three contributions from ϵ_6, ϵ_3^2 and $\epsilon_2\epsilon_4$ to v_6

Oscillating v_6 with dependence on time t

Additional zero compared to v_4 signal

Contributions compensate each other

Summary

of our results on time-dependent anisotropic flow coefficients v_{2n}
in few collision regime

- Loss term dominates the signal of anisotropic flow coefficients v_{2n}
- Rediscovery of $v_{2n} \propto t^{2n+1}$ for kinetic theory (expansion around small times t)
- Oscillation of v_{2n} with dependence on time
- Compensation of linear and non-linear eccentricity-contributions to higher flow harmonics v_{2n}

Outlook

- Further comparison with RTA-results obtained by our colleagues
- Extension to 3-dimensions and description of massiv particles

Outlook

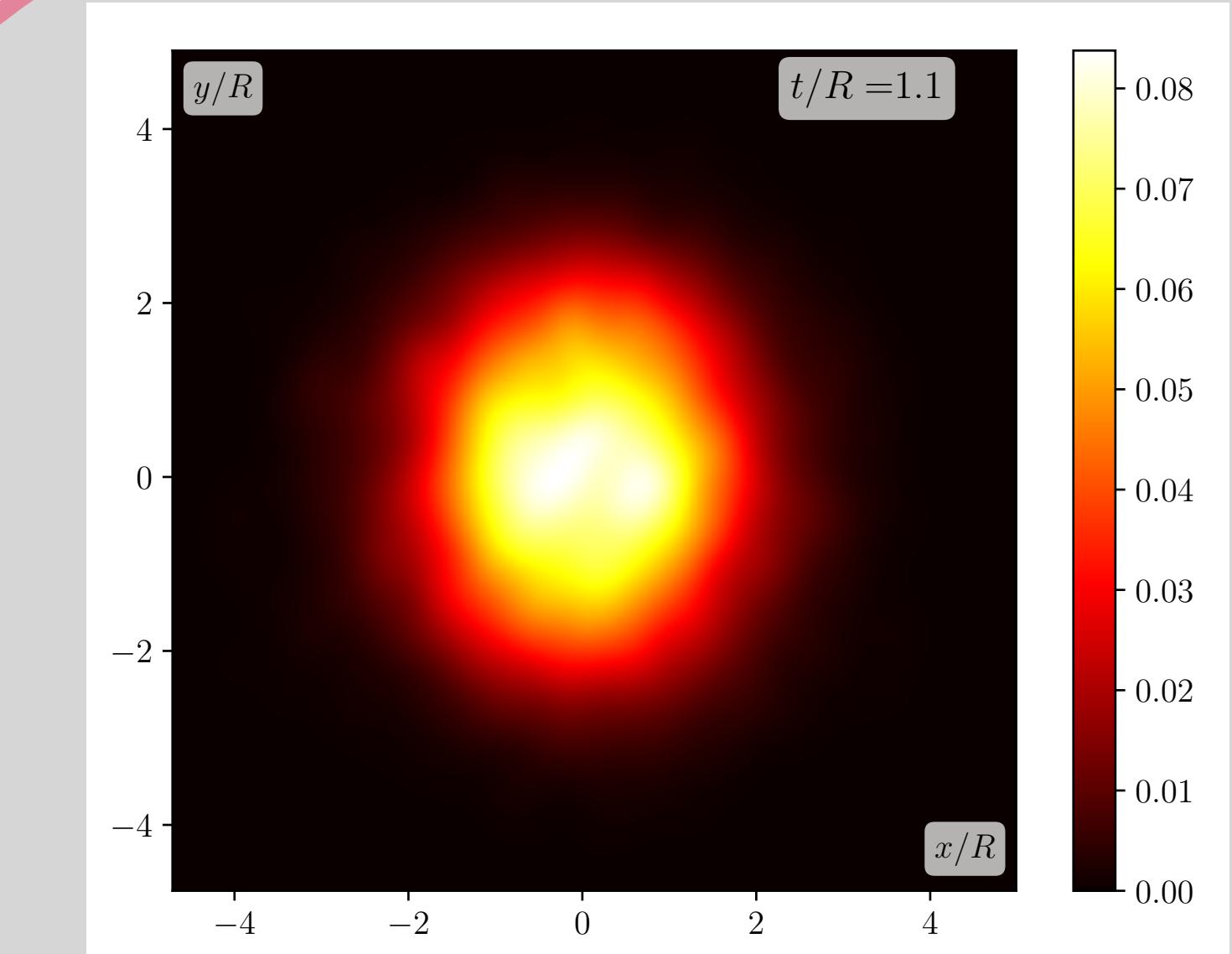
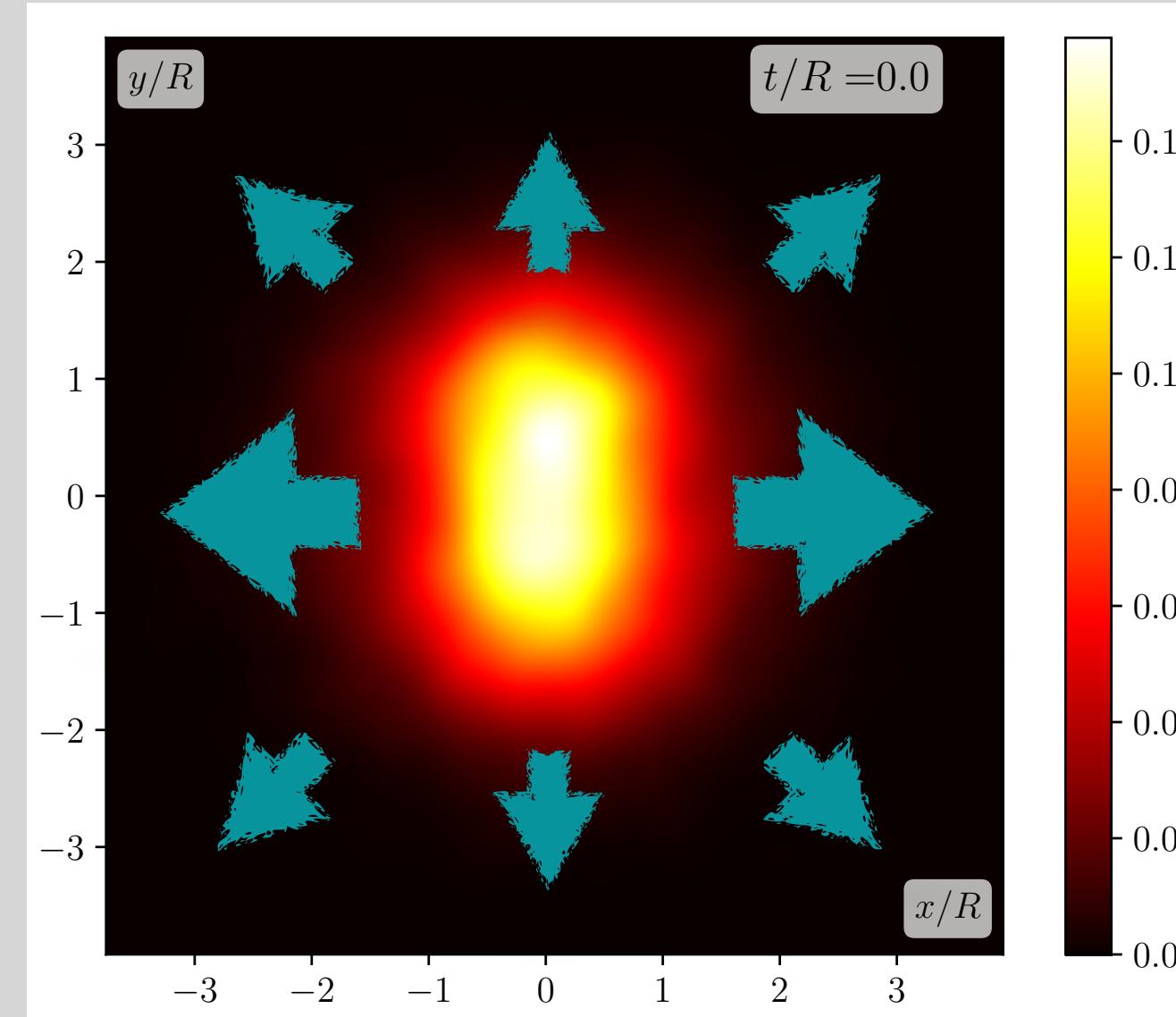
- Further comparison with RTA-results obtained by our colleagues
- Extension to 3-dimensions and description of massiv particles

Thanks for your attention

Backup



Density plots from numerical simulations



$v_4(t_*)$

- Initial eccentricity ϵ_2 only
- Negative v_4 created

- Eccentricity $\epsilon_4 > 0$ created
- Leads to increase of v_4

Note: The v_2 signal is much stronger than the v_4 signal.
It is not easy to draw conclusions from this density plot.

AFC v_4 for our approaches

with dependence on ϵ_2 and ϵ_4

- Analytical formula

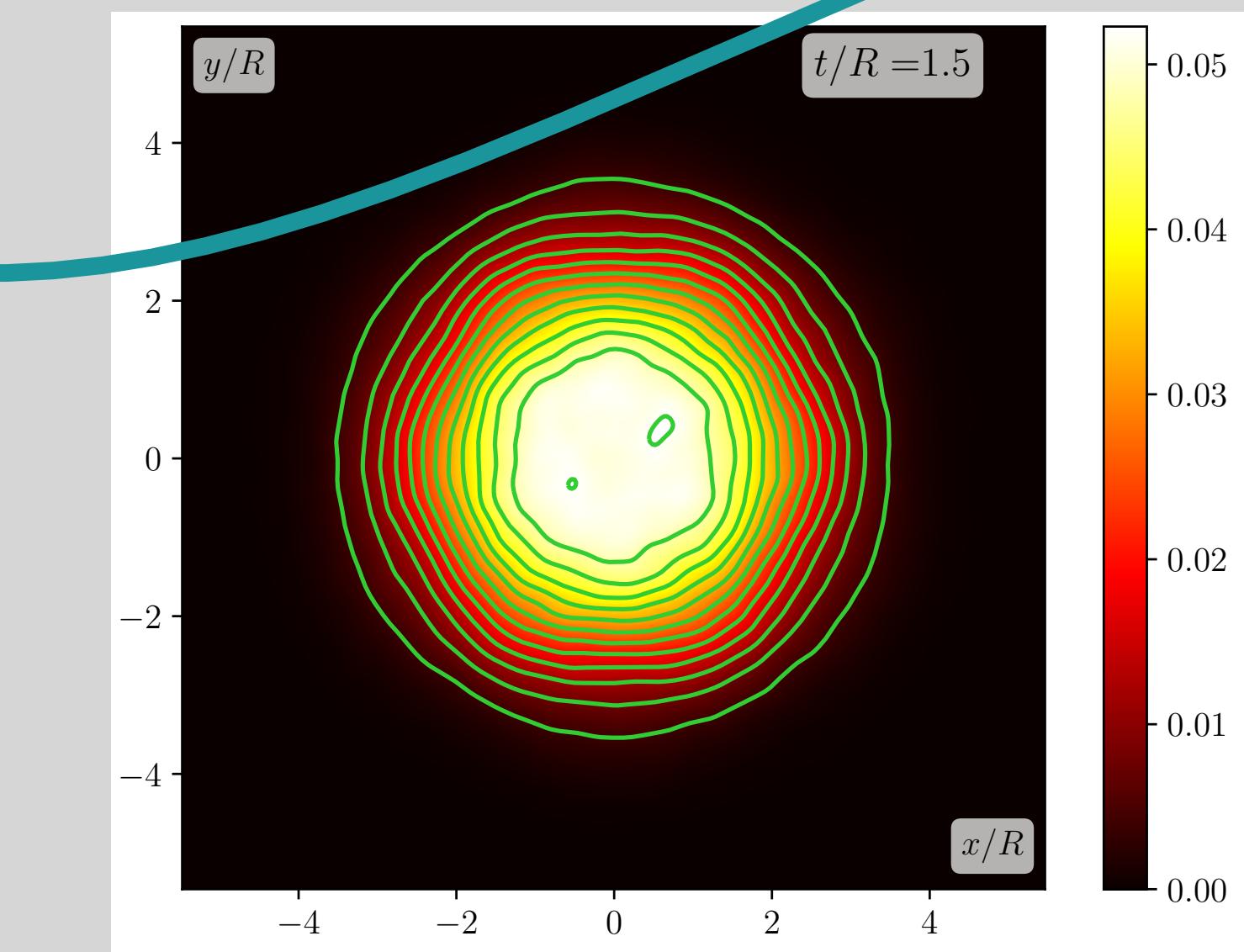
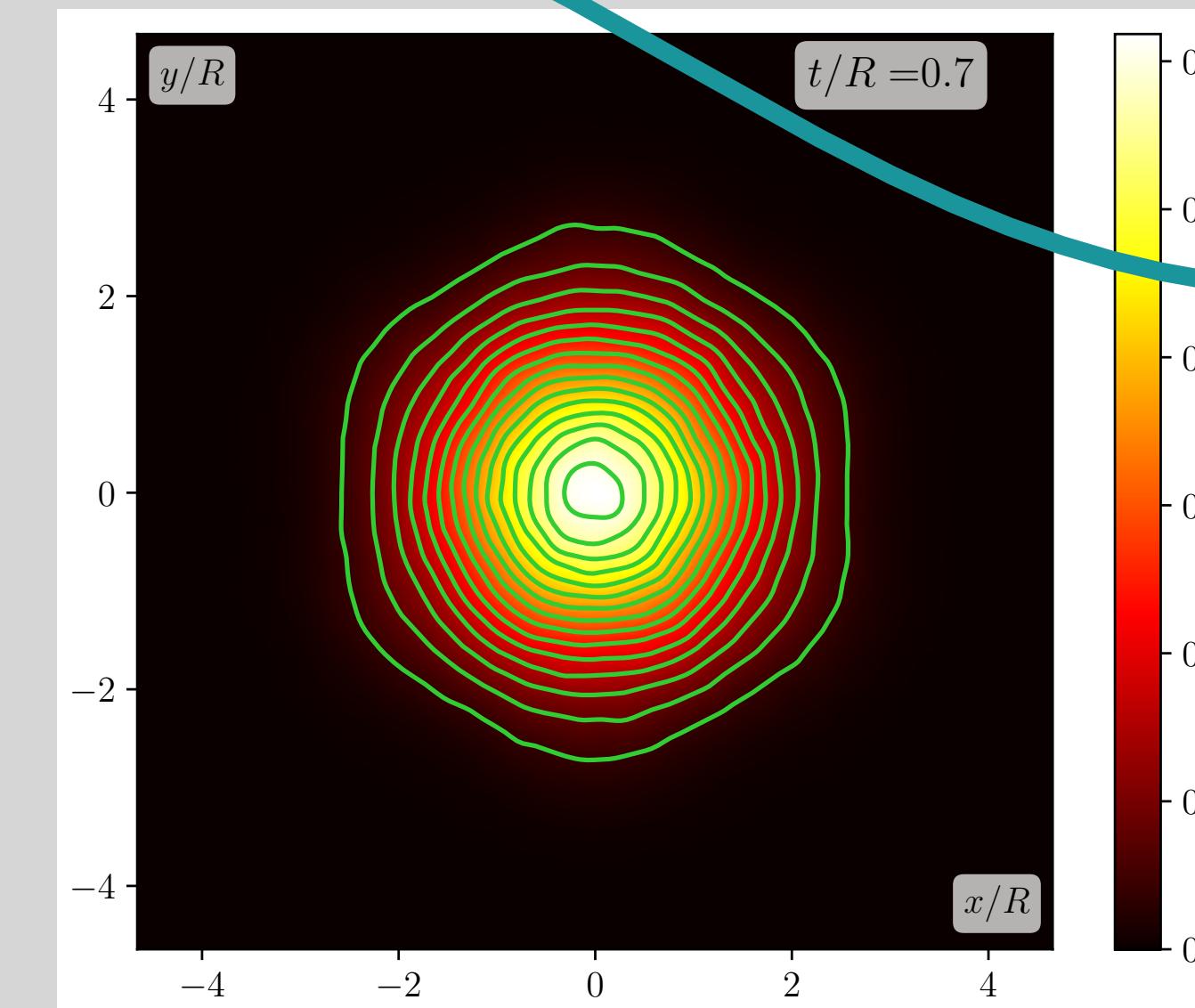
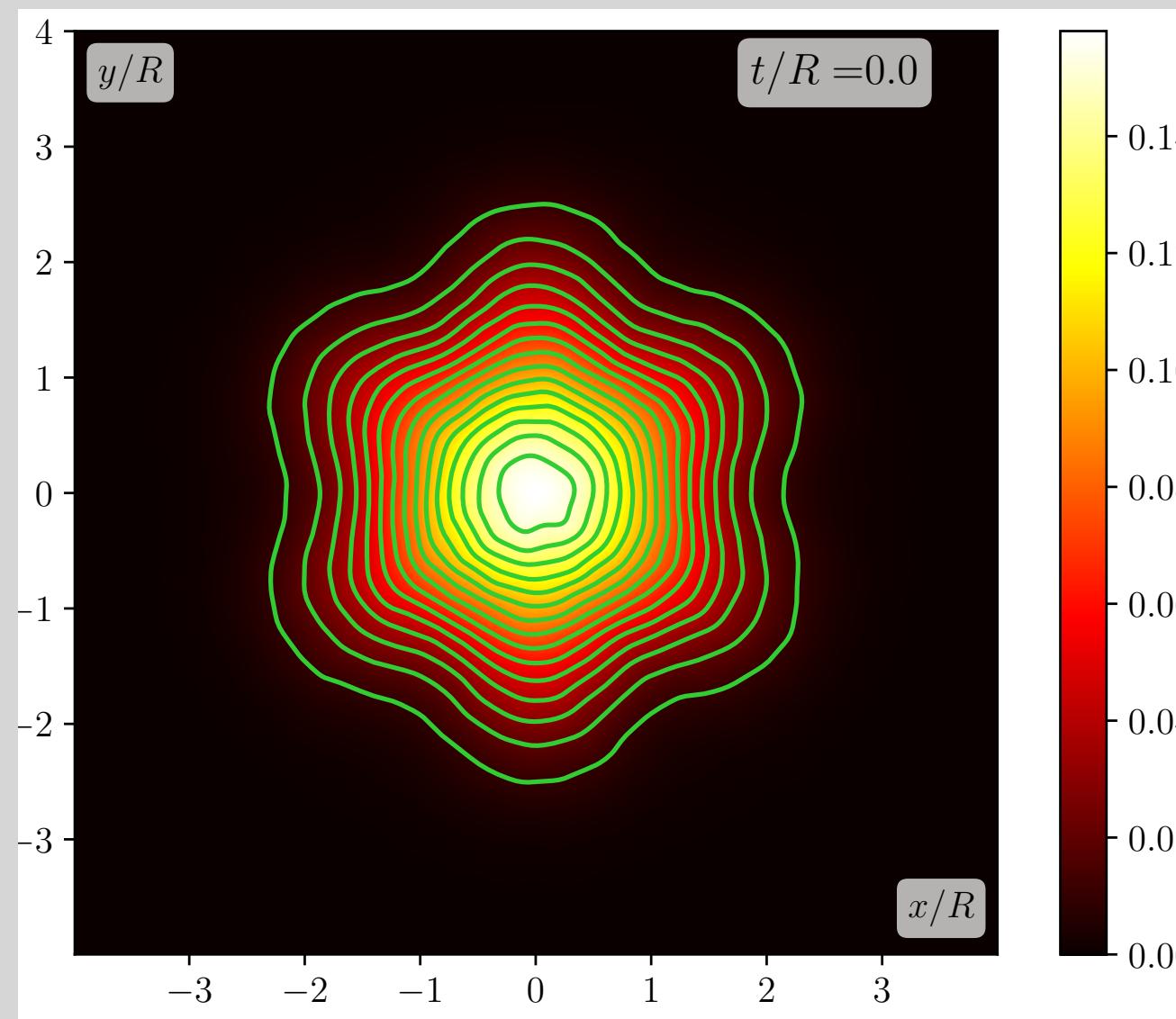
$$v_4(t_*) = \frac{64 \sqrt{\pi} N_{\text{resc}}}{405 \left(384 + 144 \sqrt{2} \epsilon_2^2 + 35 \sqrt{2} \epsilon_4^2 \right)} e^{-t_*^2}$$
$$\left[243 \epsilon_2^2 \left(\left(5t_*^3 + 14t_* + \frac{24}{t_*} \right) I_0(t_*^2) - \left(5t_*^3 + 16t_* + \frac{28}{t_*} + \frac{48}{t_*^3} \right) I_1(t_*^2) \right) \right.$$
$$\left. + 32\epsilon_4 e^{\frac{t_*^2}{3}} \left(- \left(5t_*^3 + 21t_* + \frac{54}{t_*} \right) I_0\left(\frac{2 t_*^2}{3}\right) + \left(5t_*^3 + 24t_* + \frac{63}{t_*} + \frac{162}{t_*^3} \right) I_1\left(\frac{2 t_*^2}{3}\right) \right) \right]$$

AFC v_4 for our approaches with dependence on ϵ_2 and ϵ_4

- Analytical formula

$$v_4(t_*) = \frac{64 \sqrt{\pi} N_{\text{resc}}}{405 (384 + 144 \sqrt{2} \epsilon_2^2 + 35 \sqrt{2} \epsilon_4^2)} e^{-t_*^2}$$
$$\left[2^{2/3} \epsilon_4^{1/3} \left[v_4 \propto N_{\text{resc}} [\kappa_{4,4} \epsilon_4 + \kappa_{4,22} \epsilon_2^2] t_*^5 \right. \right.$$
$$\left. \left. + 32 \epsilon_4 e^{\frac{t_*^2}{3}} \left(- \left(5t_*^3 + 21t_* + \frac{54}{t_*} \right) I_0 \left(\frac{2t_*^2}{3} \right) + \left(5t_*^3 + 24t_* + \frac{63}{t_*} + \frac{162}{t_*^3} \right) I_1 \left(\frac{2t_*^2}{3} \right) \right) \right]$$

Density plots from numerical simulations



- Initial eccentricity ϵ_6 only

- $\pi/6$ -rotated high density region \Rightarrow locally $\epsilon_6 < 0$
- Dilute regions \Rightarrow small contribution to v_6

- Original orientation in denser region regained
- Dilute regions \Rightarrow small contribution to v_6

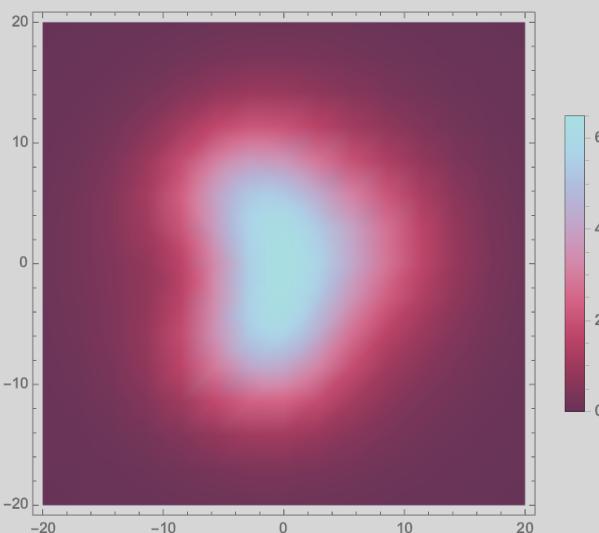
Initial distribution function

- Initially factorized in position and momentum space

$$f^{(0)}(0, \mathbf{r}, \mathbf{p}) = G(\mathbf{r}) F(\mathbf{p})$$

- Initial distribution function in position space

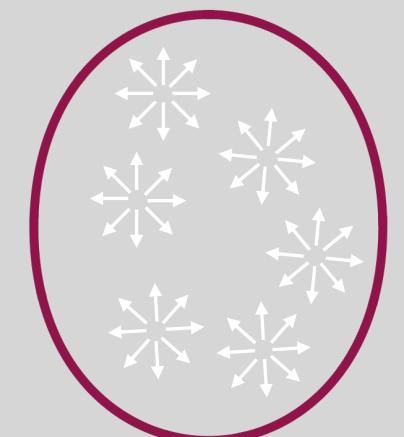
$$G(\mathbf{r}) \propto \exp\left(-\frac{r^2}{2R^2}\right) \left[1 + \sum_{n=2}^{\infty} \tilde{\epsilon}_n \left(\frac{r}{R}\right)^n \cos(n(\theta - \Psi_n)) \exp\left(-\frac{r^2}{2R^2}\right) \right]$$



- Initially isotropy in momentum space

$$\Rightarrow F(\mathbf{p}) = F(p_T)$$

OR



- Anisotropic initial momentum distribution

$$F(\mathbf{p}) = \tilde{F}(p_{\perp}) \left[1 + 2 \sum_{k=2}^{\infty} \left(w_{k,c} \cos(k\phi) + w_{k,s} \sin(k\phi) \right) \right]$$

Eccentricity

$$\bullet \text{ Formula } \epsilon_n = \frac{\int G(r, \theta) \cos(n(\theta - \Psi_n)) r^{n+1} d\theta dr}{\int G(r, \theta) r^{n+1} d\theta dr}$$

weighted with r^n

⇒ Outer regions contribute stronger to ϵ_n