

# Fluctuations of Energy Density in Ultra-Central Collisions

Liner Santos and Matthew Luzum

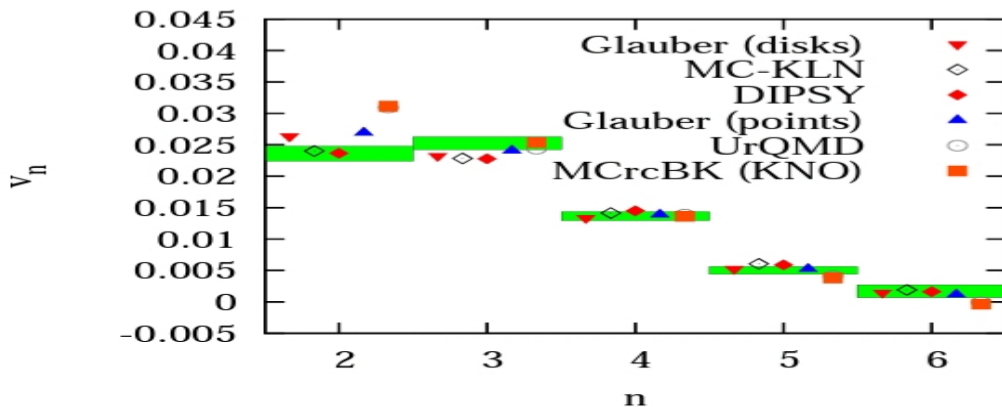
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- Problem - In ultra-central regime, we have  $v_3 \sim v_2$  and there is no one existing model that can calculate  $v_n\{2\}$  with a good accuracy.
- Goal - Find what must be the initial conditions in order to get the experimental data in this regime.
- Method - Relate  $v_n$  to n-point functions of initial energy density.
- Low density regions - Fluctuations are larger than monte carlo existing models.

# Motivation

The existing initial condition models that calculate the harmonic flows through a hydrodynamic evolution aren't able to generate results with a good accuracy related to experimental data in ultra-central regime.



Harmonic flows from ATLAS collaboration in 0 – 1% cent. bin. We can see that the results got from existing models are not able to fit with the experimental data. Image from (*Luzum, M. et. al. Nuclear Physics A 904-905 (2013) pp. 377-380*)

# N-point functions and Harmonic flows

- Starting point  $\rho(z) = \langle \rho(z) \rangle + \delta\rho(z)$ ,  
where  $\rho(z)$  is the energy density in one event and  $\delta\rho(z)$  is the fluctuation of energy density. We have used  $z = x + iy$ .

## Assumptions

- **Local Fluctuations** - Neglect correlations.  
 $\Rightarrow \langle \delta\rho(z_1)\delta\rho(z_2) \rangle = \delta(z_1 - z_2)\kappa_2(z_1)$   
 $\Rightarrow \langle \delta\rho(z_1)\delta\rho(z_2)\delta\rho(z_3) \rangle = \delta(z_1 - z_2)\delta(z_1 - z_3)\kappa_3(z_1)$ .
- **Small Fluctuations** - The RMS values of  $\epsilon_n$  are given by  $\langle \rho(z) \rangle$  and by 1 and 2-point functions  $\rightarrow$  power series in  $\delta\rho$ .

In these equations,  $\kappa_1(z) = \langle \rho(z) \rangle$ ,  $\kappa_2(z) \propto \text{var}(\rho(z))$  are the cumulants per surface unit. These results are from *ArXiv 1604.07230*.

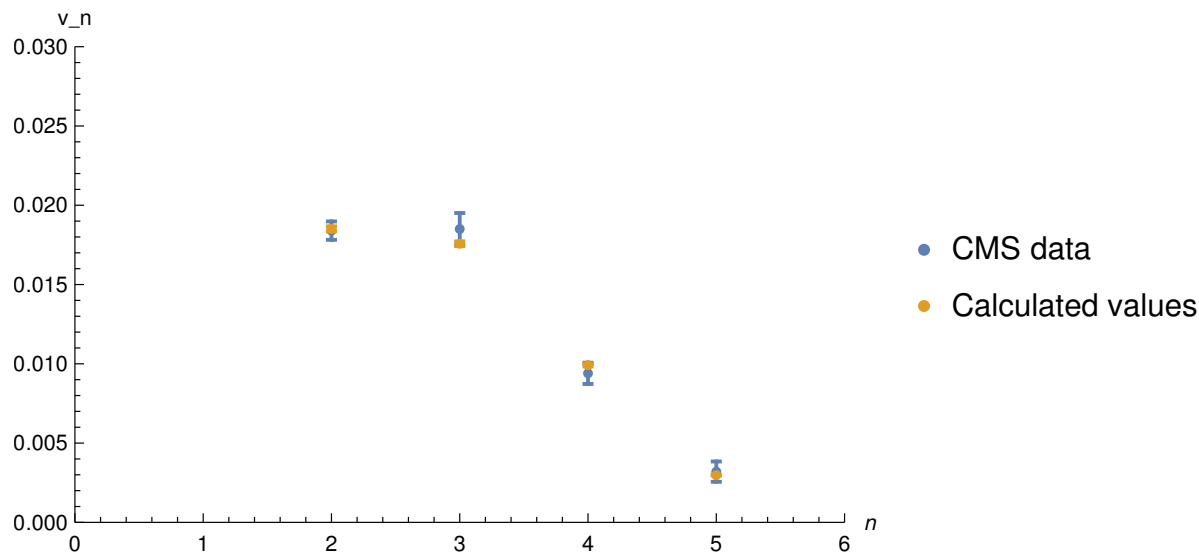
# N-point functions and Harmonic flows

- $\kappa_1$  It's the mean of energy density for each point over several events  $\rightarrow$  CGC, TRENTO, etc.
- $\kappa_2$  Means the fluctuations of each spacepoint. We have parameterized the fluctuations in terms of the mean:

$$\kappa_2 = C^2 \cdot \kappa_1^{2p} \quad (1)$$

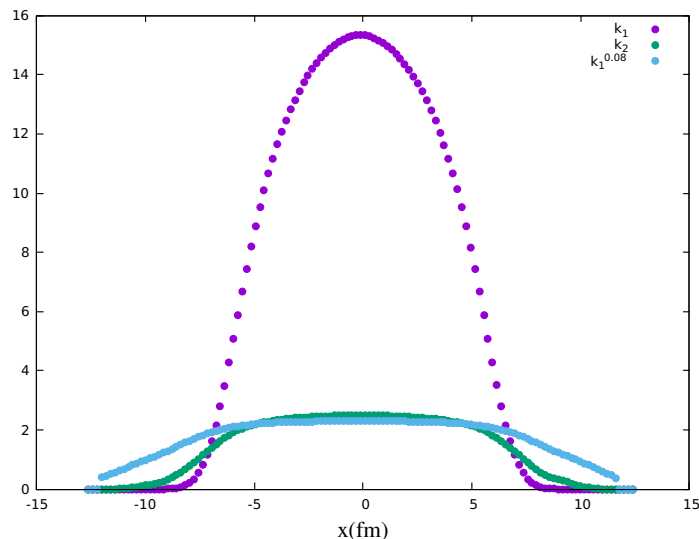
- **Response coefficient**  $K_n \equiv \frac{v_n}{\epsilon_n}$  - MUSIC simulations, assuming  $K_n \equiv K_n(\frac{\eta}{s})$ , where  $\frac{\eta}{s}$  is the shear viscosity.

# Results



This plot compares the harmonic flows from CMS 0 – 0.02% collaboration and the values got from the power parameterization of the mean. One can see that this parameterization generates harmonic flows with a reasonable accuracy.

# Results - Initial conditions



Green dots are the variance per unit surface from TRENTO results and blue dots shows  $\kappa_2(z)$  got from exponential law with  $p = 0.08$ . Purple points are the mean per surface unit from TRENTO initial conditions. We can see that the fluctuations in peripheral zone are larger with the power parameterization than we consider TRENTO results.

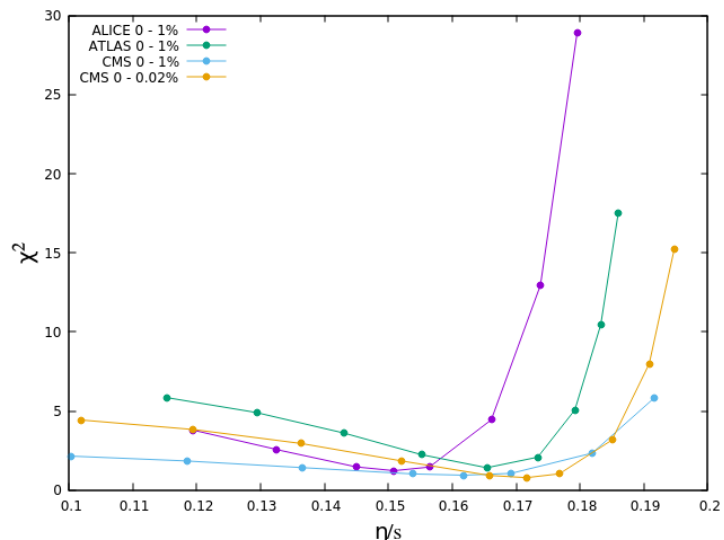
Thank you!



Using the power parameterization, we have the results below

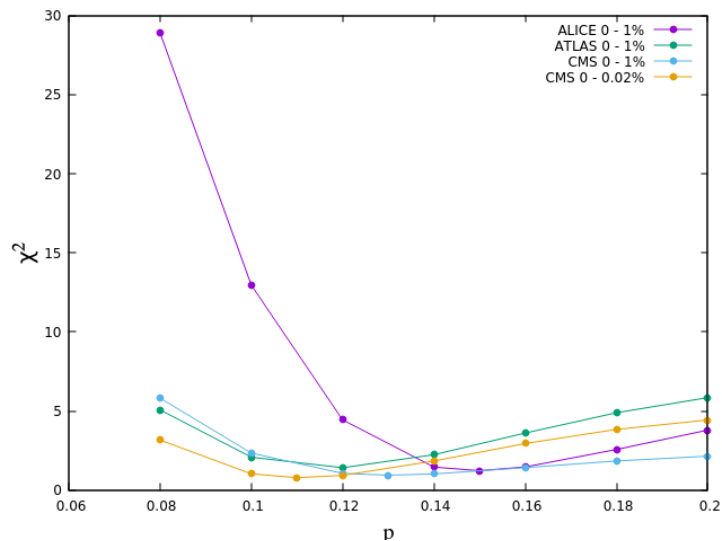
$\kappa_1$	Collaboration	$p$	$C$	$\eta/s$	$\chi^2$
CGC	ALICE 0 – 1%	0.15	12.64	0.15	1.23
CGC	ATLAS 0 – 1%	0.12	10.58	0.16	1.42
CGC	CMS 0 – 1%	0.13	10.70	0.16	0.94
CGC	CMS 0 – 0.02%	0.11	09.18	0.17	0.80
TRENTO	ALICE 0 – 1%	0.12	3.38	0.14	0.78
TRENTO	ATLAS 0 – 1%	0.05	2.20	0.16	0.72
TRENTO	CMS 0 – 1%	0.08	2.55	0.16	0.58
TRENTO	CMS 0 – 0.02%	0.04	1.85	0.16	0.37

# Extra material



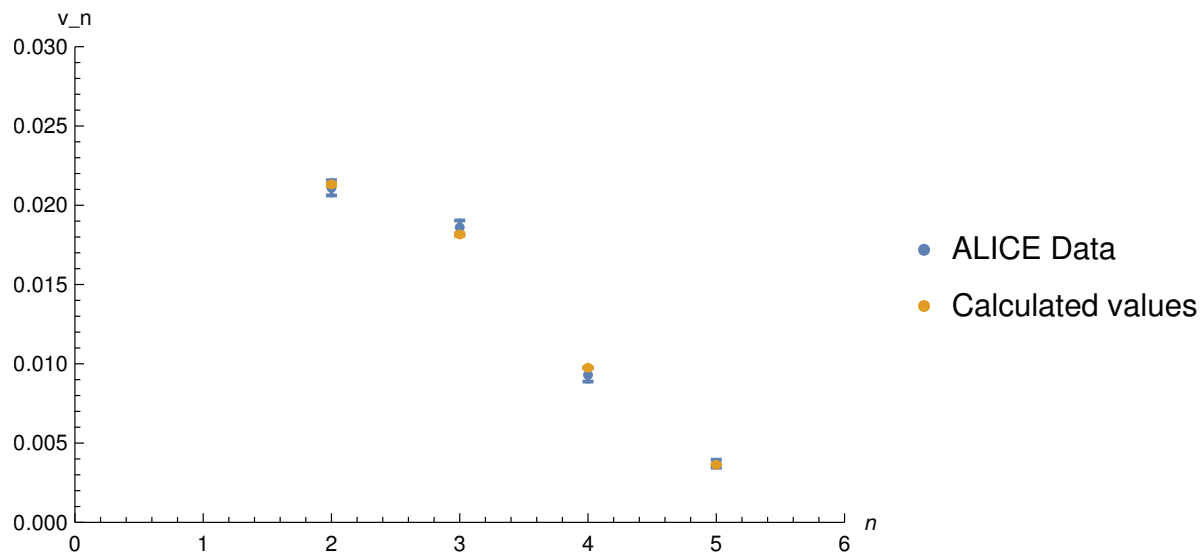
The shear viscosity values and the corresponding  $\chi^2$  got from power parameterization and the least square.

# Extra material



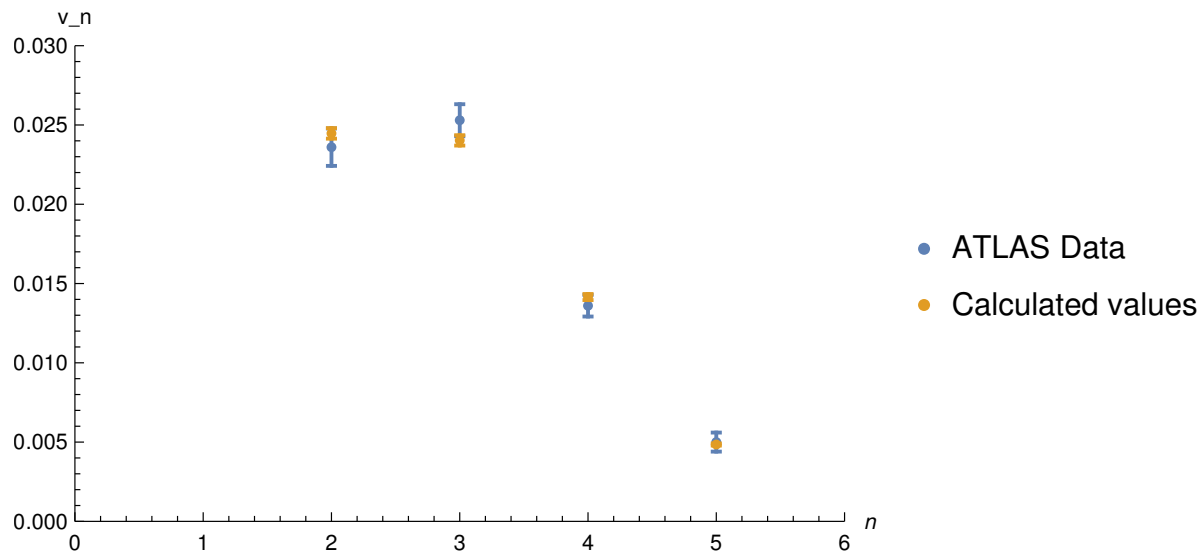
The free parameter  $p$  values and the corresponding  $\chi^2$  got from power parameterization and the least square.

# Extra material



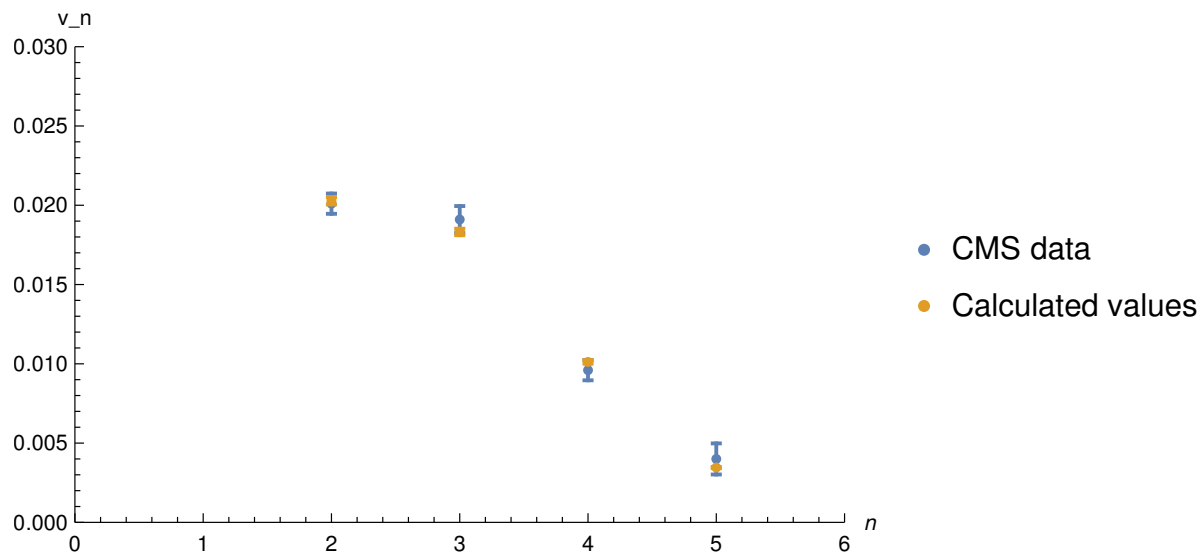
This plot compares the harmonic flows from ALICE 0 – 1% collaboration and the values got from the power parameterization of the mean. One can see that this parameterization generates harmonic flows with a reasonable accuracy.

# Extra material



This plot compares the harmonic flows from ATLAS 0 – 1% collaboration and the values got from the power parameterization of the mean. One can see that this parameterization generates harmonic flows with a reasonable accuracy.

# Extra material



This plot compares the harmonic flows from CMS 0 – 1% collaboration and the values got from the power parameterization of the mean. One can see that this parameterization generates harmonic flows with a reasonable accuracy.

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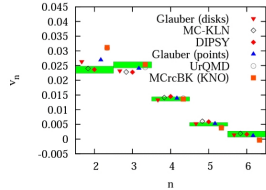


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## Introduction

The existing initial conditions models for heavy ion collisions give the flow coefficients  $v_n$  with a good accuracy for the most of cases. However, in ultra-central events, i.e. centrality bin  $0-1\%$  or more central, there is no model for initial conditions that provide all the  $v_n$  in a satisfactory way, as you can see on fig.(1). Thus, it is *sine qua non* to understand what properties the system must have under these centralities in order to the flow coefficients are compatible with the experimental data, remembering that the fluctuations play an important role in these cases, since the participant is almost symmetric.

In this work, we have made a parameterization of energy density and its fluctuations via 1 and 2 point functions, assuming that ones are small and local. Besides, we constrain them by using experimental data for  $v_n$ (2) in conjunction with hydrodynamic simulations performed on MUSIC, which is a C++ code that simulates all system evolution. We believe that it is possible to describe the experimental data with the requirement that the fluctuations are larger than mean in regions of low density. We comment on the implications of this finding, as well as the limitations of this initial analysis and how it can be improved in the future



Harmonic flows from ATLAS collaboration in centrality bin  $0-1\%$ . We can see that no initial condition model used could calculated all  $v_n$  with reasonable accuracy. Image from [4].

## N-Point Functions and Harmonic flows

The starting point is to write energy density as  $\rho(z) = \langle \rho(z) \rangle + \delta\rho(z)$ , where  $\delta\rho(z)$  is the fluctuation of energy density and  $z$  is a shorthand notation to  $(x, y)$ . Besides this, if we consider that the fluctuations are small and the correlations are local, we find a constrain between the fluctuations and the correlations:

$$\langle \delta\rho(z_1)\delta\rho(z_2) \rangle = \kappa_2(z_1)\delta(z_1 - z_2) \quad (1)$$

where  $\kappa_2(z)$  is  $\kappa_2$  per unit surface. Combining this equation with the definition of eccentricities:

$$e_n = \frac{\int d^2r r^n \rho(z)}{\int d^2r |\rho(z)|^2} \quad (2)$$

We get a expression for rms value of eccentricities  $e_n$ (2)[1]:

$$e_n(2) \equiv \langle |e_n(2)|^2 \rangle^{1/2} \approx a = \sqrt{\frac{\int d^2r r^{2n} \kappa_2(z)}{\int d^2r \rho^2(z)}} \quad (3)$$

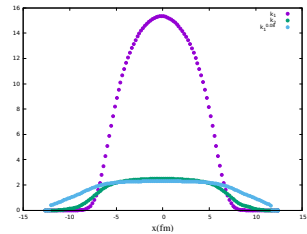
$$\Rightarrow v_n(2) = K_n * \sqrt{\frac{\int d^2r r^{2n} \kappa_2(z)}{\int d^2r \rho^2(z)}} \quad (4)$$

The coefficients  $K_n$  were calculated by MUSIC, which is a C++ code that calculates all hydro evolution of the system, from initial conditions to kinetic freeze-out e generates the spectra of final stage. This coefficient is defined by the ratio between the harmonic flow  $v_n$  and the eccentricity  $e_n$ . According to hydro description, the harmonic flows are strongly related to the shear viscosity to entropy ratio  $\eta/s$ . Thus, we perform simulations changing the shear viscosity to get the expressions for  $K_n$  via linear fit.

Also, we use the fact that the eccentricities are small in ultra-central collisions. In other centralities bins, the relation between  $v_n$  and  $e_n$  has terms as  $\kappa_n * e_n$ . This assumption proved to be essential to the model to work. In the fig.(2), we show the values of fluctuations obtained by TRENTO, which is a code that generates fluctuating initial conditions, and the values of fluctuations using an exponential relation between the mean  $\kappa_1(z)$  and the variance  $\kappa_2(z)$ . TRENTO model doesn't work in ultra-central events whereas, the exponential relation

$$\kappa_2(z) = C^2 * \kappa_1(z)^2 \quad (5)$$

where  $p$  is a parameter put *ad hoc*. works. Thus, we have thought that the fluctuations must be small and larger than one calculated by Monte-Carlo existing models in regions of low density.



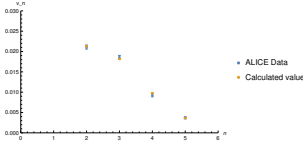
Comparison between the mean and the fluctuations obtained by TRENTO (green dots) and by exponential law (blue dots). X-axis means the position related to CM.

## Results

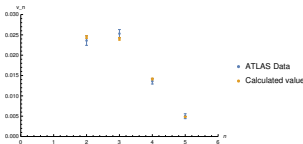
In order to calculate the harmonic flows, we found a relation between  $\kappa_1$  and  $\kappa_2$  so that we can apply the eq.(4) to calculate the harmonic flows. Then, we have applied the least square method using the experimental data. Thus, we have found the "fitted values" of harmonic flows, and the value of the shear viscosity. In the table below, we see the values that that the calculated values of  $v_n$  and experimental data from ALICE, ATLAS and CMS collaborations[2] in centrality bin  $0-1\%$  have a reasonable accuracy.

Collaboration	$p$	$C$	$\eta/s$	$\chi^2_{dof}$	Cent. bin
ALICE	0.15	12.64	0.15	1.23	$0-1\%$
ATLAS	0.12	10.58	0.16	1.42	$0-1\%$
CMS	0.13	10.70	0.16	0.94	$0-1\%$
CMS	0.11	9.18	0.17	0.80	$0-0.02\%$

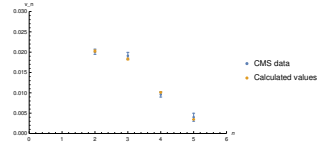
$p, C, \eta/s \propto \kappa^2$  for each collaboration with centrality bin  $0-1\%$ .



Experimental  $v_n$ (2) in  $0-1\%$  from ALICE collaboration and calculated values with this mapping.



Experimental  $v_n$ (2) in  $0-1\%$  from ATLAS collaboration and calculated values with this mapping.



Experimental  $v_n$ (2) in  $0-1\%$  from CMS collaboration and calculated values with the mapping.

The table below shows the extracted values of shear viscosity for each collaboration and the corresponding  $\kappa^2$

## Final Considerations

In ultra-central events, the triangular flow  $v_3$ (2) is very close, and bigger than in some cases, to elliptic flow  $v_2$ (2) and this one doesn't occurs in less central events. In order to describe these results, we parameterize the energy density and its fluctuations via 1 and 2 point functions and we have assumed that the fluctuations are bigger than the mean in low density regions. We have combined this ansatz with hydro simulations and we have compared to experimental data through the least square method and we have got results with a good accuracy. Besides, the shear viscosity obtained is close to phenomenological results [4]. For the future works, we think that to improve these results adding other physical parameters, as bulk viscosity to s ratio ( $\zeta/s$ ) to response coefficients, as well to think in asymetry on fluctuations.

## References

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