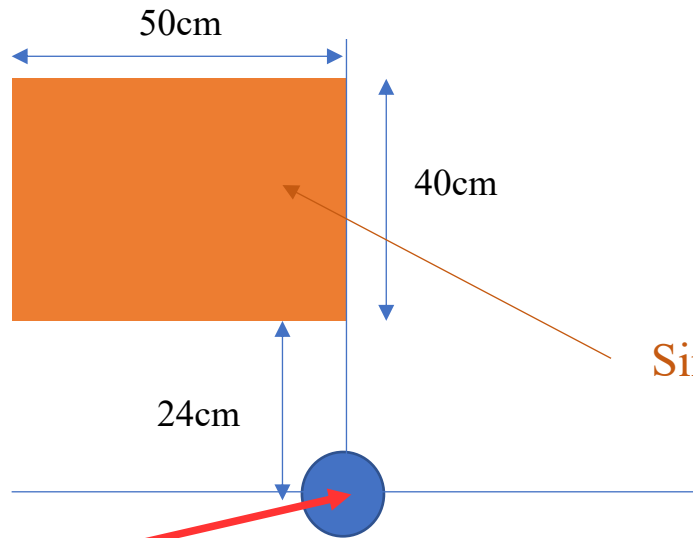


Scattering and Neutrino Detector in TI18

Giovanni De Lellis

Detector configuration



$$7.1 < \eta < 8.3$$

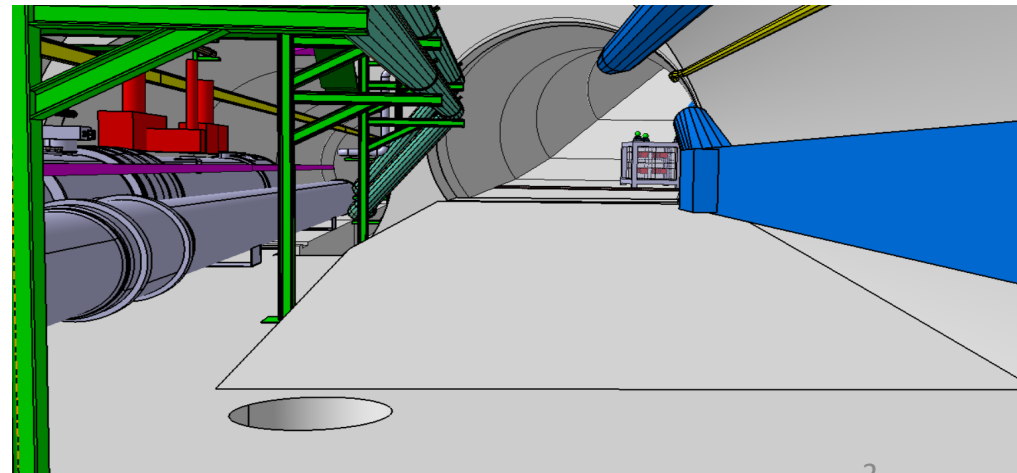
Simulation refers to this area

η	Distance (m)
9.86	0.05
7.38	0.6
6.5	1.4
6	2.4

Beam axis
prolongation

5.8 cm above the l.o.s. due to crossing angle

TI18, Pablo Santos Diaz, EN



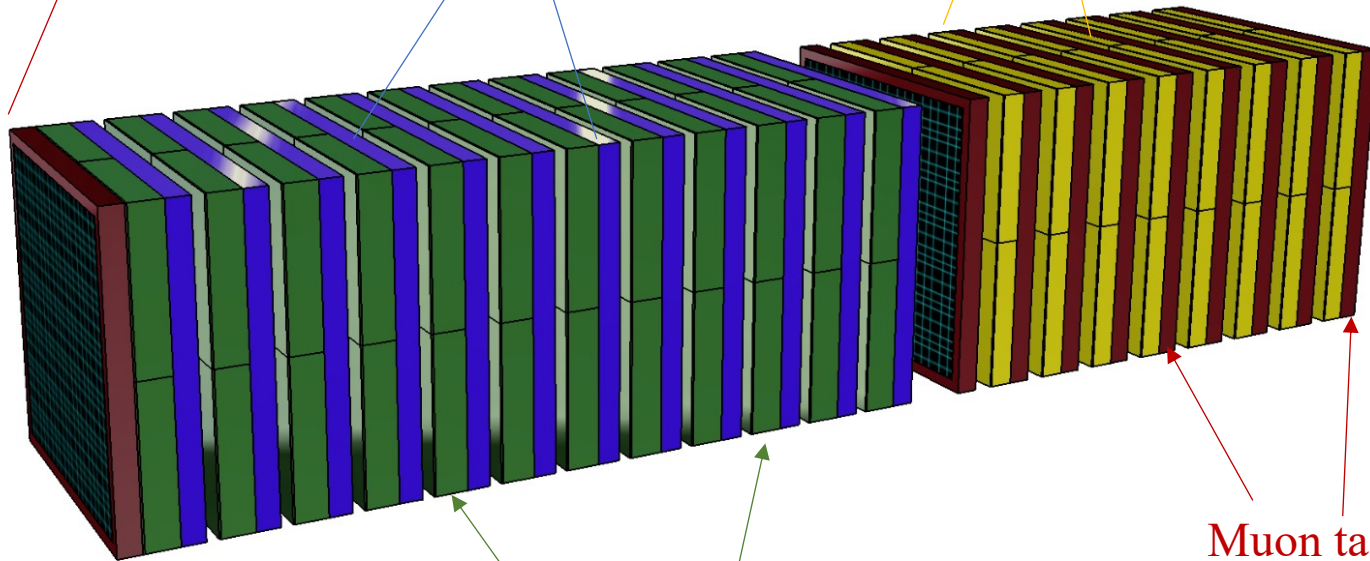
SND detector setup

Plastic scintillator read out
by SiPM arrays
Veto plane

SciFi read out
by SiPM
TT-emulsion matching
Energy measurement

ECC+SciFi = Wall
20 walls in this first attempt
12 active + 8 passive (muon)

Passive material
for muon filter



Emulsion Cloud Chamber
57 emulsion films interleaved
by 56 1mm lead plates (10 X⁰)
VTX + tau id

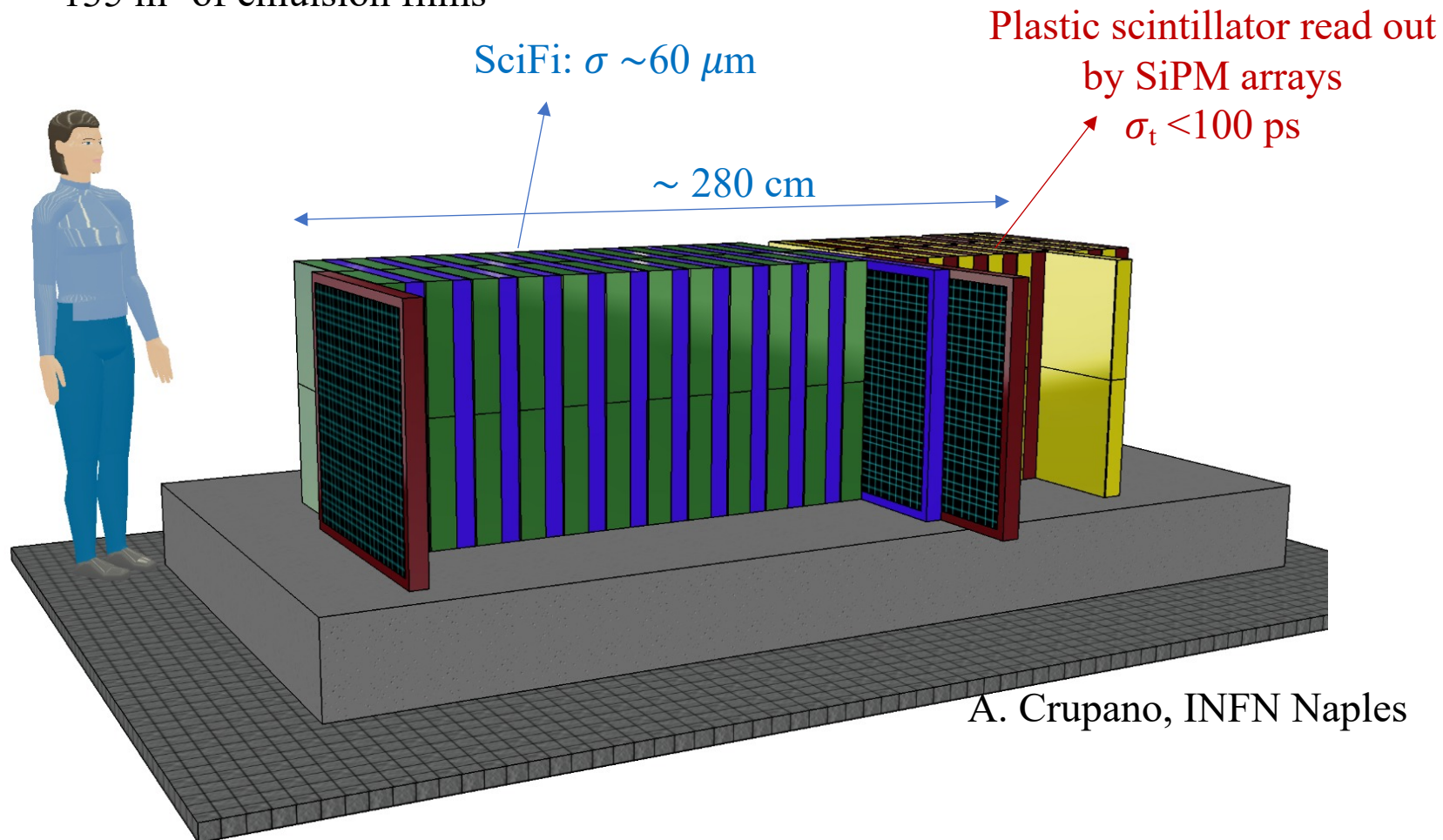
Muon tagging
Fast timing

Plastic scintillator read out
by SiPM arrays

SND detector setup

ν target: 1.5 tons

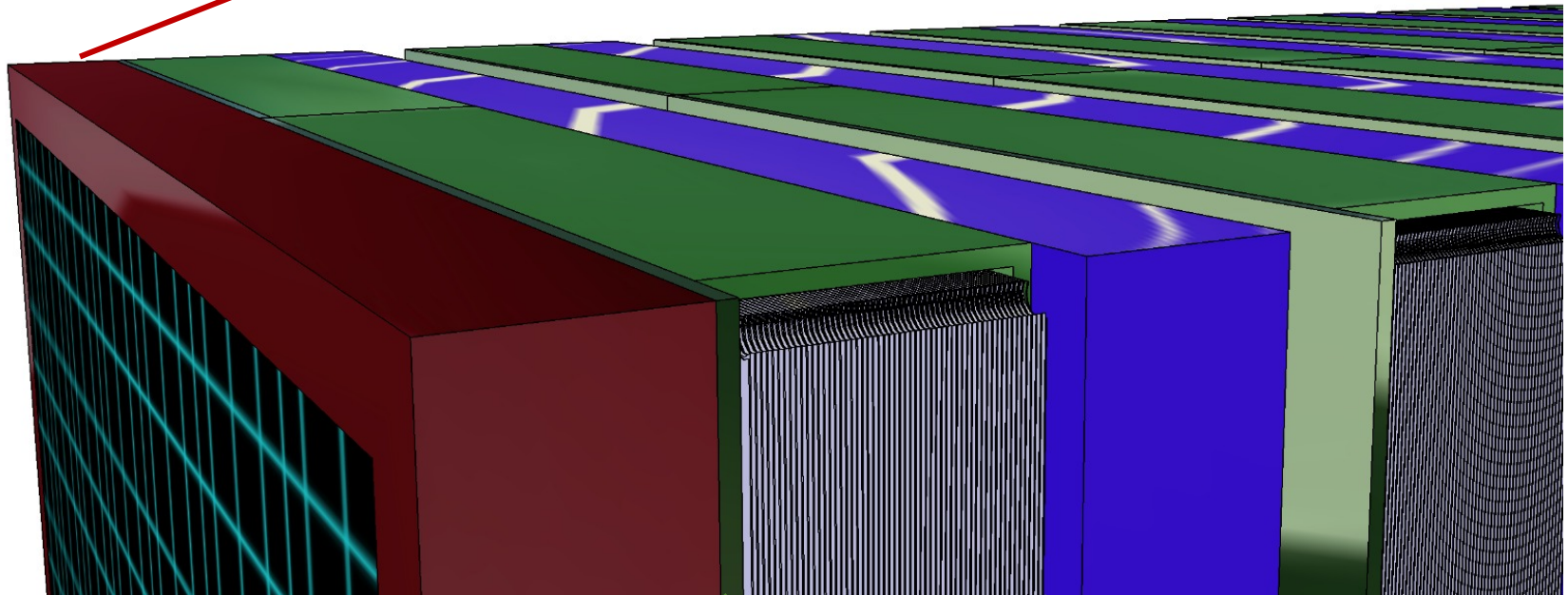
135 m² of emulsion films



Veto plane to discriminate between neutrino-induced and penetrating muons
Fast timing required to distinguish between neutrinos and heavier particles

SND detector setup

Bar dimension: $50 \times 1 \times 1 \text{ cm}^3$ (X and Y arrangement)
In SHIP $168 \times 6 \text{ cm}^2 \rightarrow$ better timing in this configuration



Emulsion Cloud Chamber
57 emulsion films interleaved
by 56 1mm lead plates (10 X⁰)

Detector considerations

- Emulsion as a vertex detector, with optimal e/π^0 separation
- Hybrid detector for all 3 neutrino flavours
- Muon identification with a muon filter (scintillator bars): muon neutrinos and $\tau \rightarrow \mu$ decays
- Energy measurement with SciFi (and scintillator bars) in combination with emulsions
- SciFi is a good option for several reasons:
 - The high intensity flux of muons is similar to the SHiP conditions ($10^5/\text{cm}^2$)
 - Hence, matching with emulsion tracks requires high position accuracy ($< 100 \mu\text{m}$)
 - Very light infrastructure required (low radiation rate, $< 1\text{Hz}/\text{cm}^2 \rightarrow$ no cooling needed)
 - It complies with the difficulty to make sizeable infrastructure work for Run3 in T118
- A scintillator station upstream as a veto. Same technology used for the muon filter
- Time stamp with good timing resolution ($< 100 \text{ps}$) required to separate neutrinos from DM at 480m distance

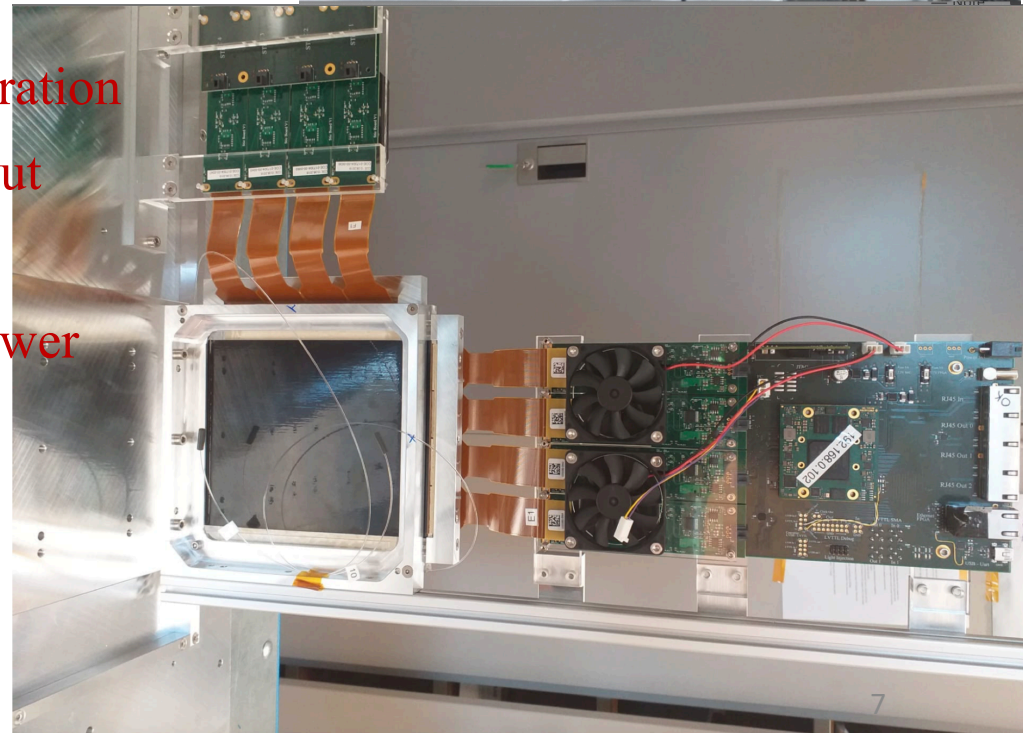
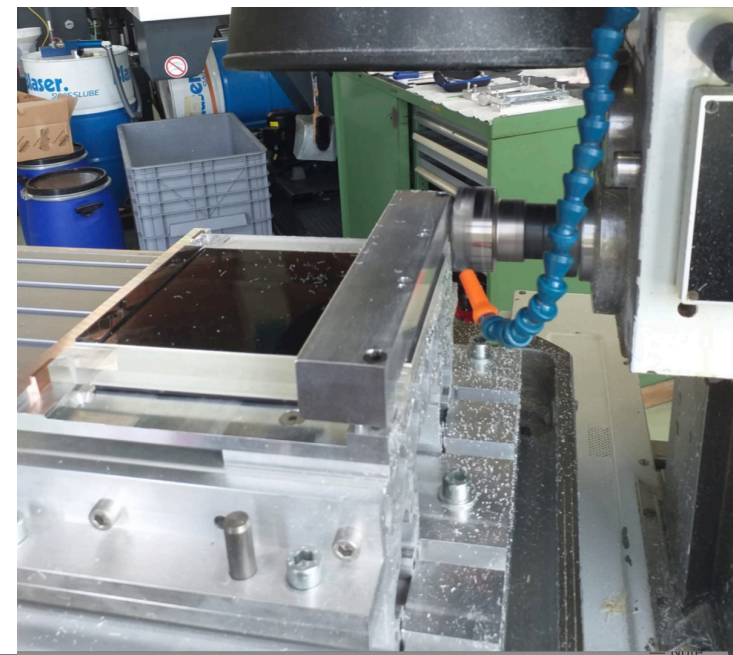
SciFi in TI18 (EPFL)

Current status:

- 8 planes (or 4xy) available with the size of $(400\text{mm})^2$
- possible to double this number by 2021
- power consumption ~ 125 W/plane
- cooling power required ~ 125 W/plane, \rightarrow 250 W/plane effectively
- one socket of 220 V is enough for operation
- Ethernet connection for the data readout

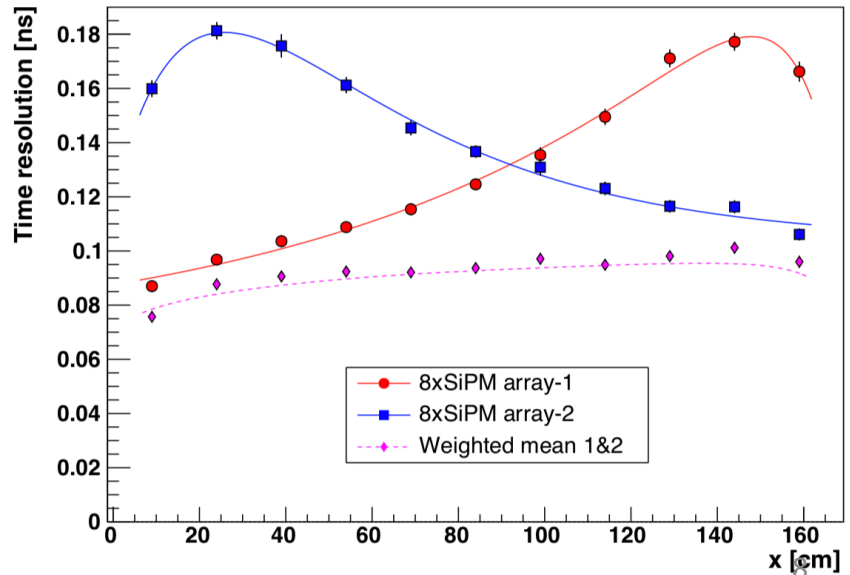
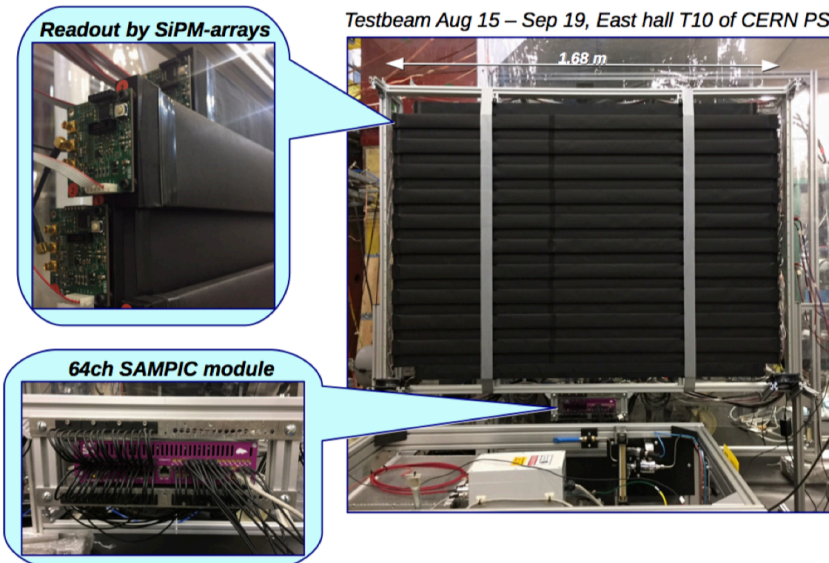
Possible developments:

- conceivable to have new chips with lower power consumption, ~ 15 W/plane

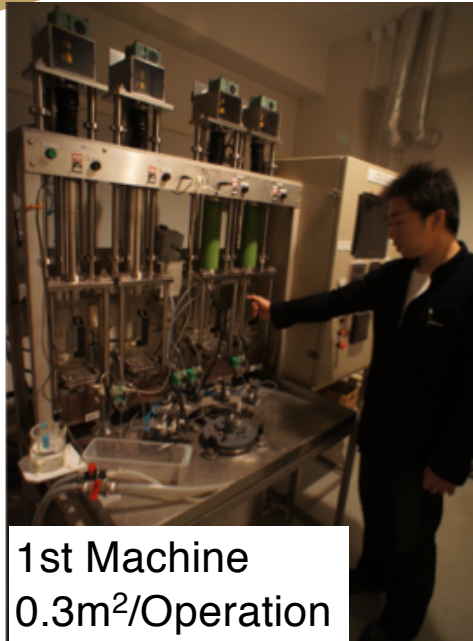
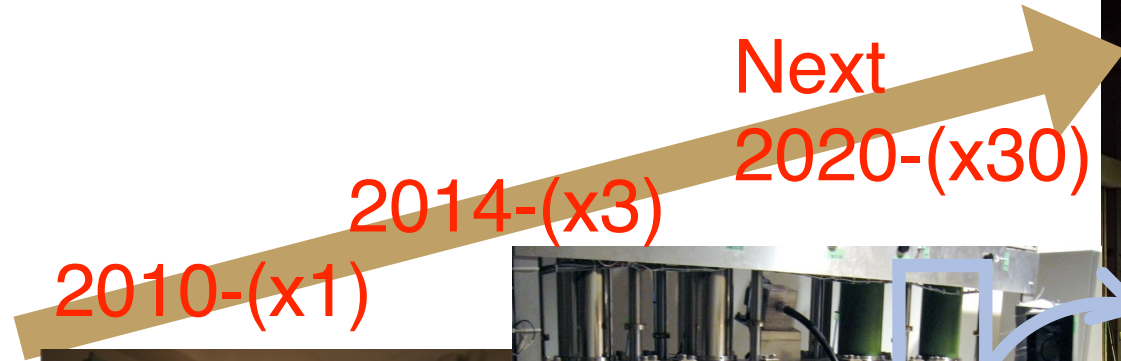


Scintillating bars with high timing performance (Zurich)

- Length to be shortened (50-60 cm)
- Smaller width (\sim cm)



Gel Production Facility at Nagoya Univ. for all emulsion projects



1st Machine
0.3m²/Operation



2nd Machine
1 m²/Operation

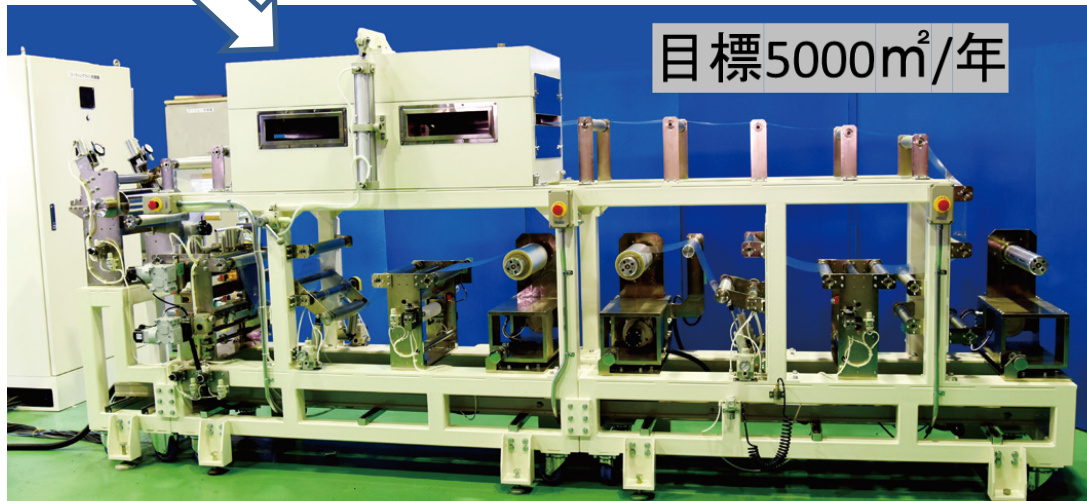


3rd Machine
(under construction)
10 m²/Operation
→2000-3000m²/year

Automation of film production

Production by hand limited to 600m²/year

Aiming at 5000m²/year by a roll to roll pouring system



135 m² per 1.5 ton detector
Assume to replace emulsions
every $\sim 20 \text{ fb}^{-1}$, 7-8 times in Run3
 $\sim 1000 \text{ m}^2$

Emulsion production at Slavich



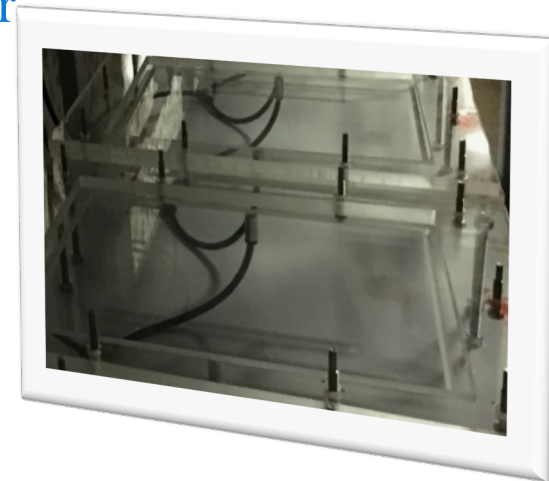
- Emulsion production by SLAVICH company started in 2009, since 2012 pilot batches manufactured
- Production scale today is about 190kg/year
- Emulsion production could be as large as ~600 kg/year



Emulsion production at Slavich



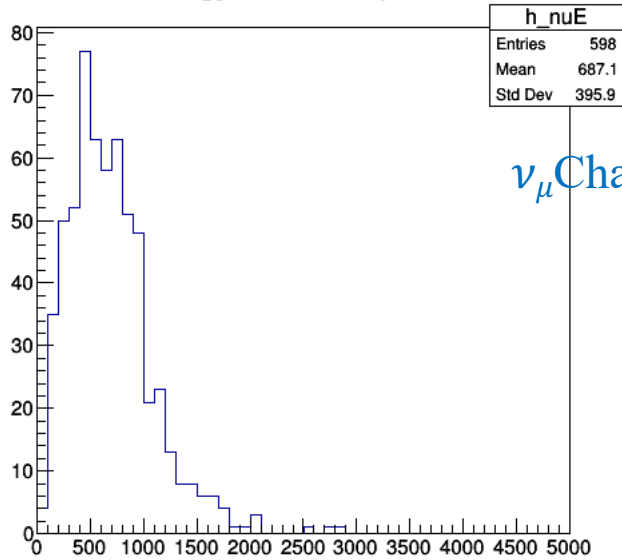
- Currently, films are produced by hand pouring with 18 vacuum tables
- Speed production by hand pouring now is 80 m² per year



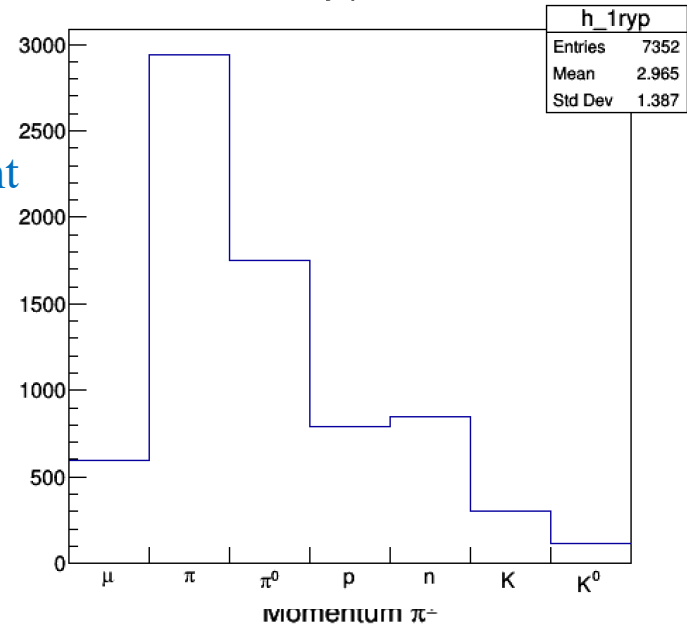
**For a large production, an automated chain could be used
Capacity of pilot automated pouring machine is 480 m² per year**

Energy spectra of produced particles: ν flux Pythia (Luca) + ν int Genie (Annarita)

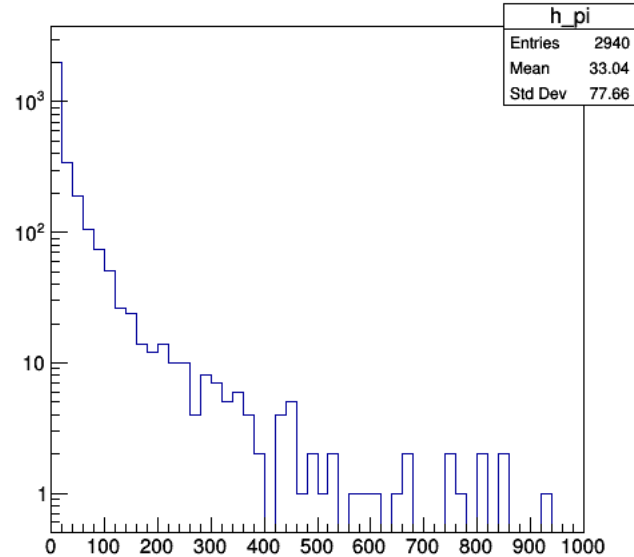
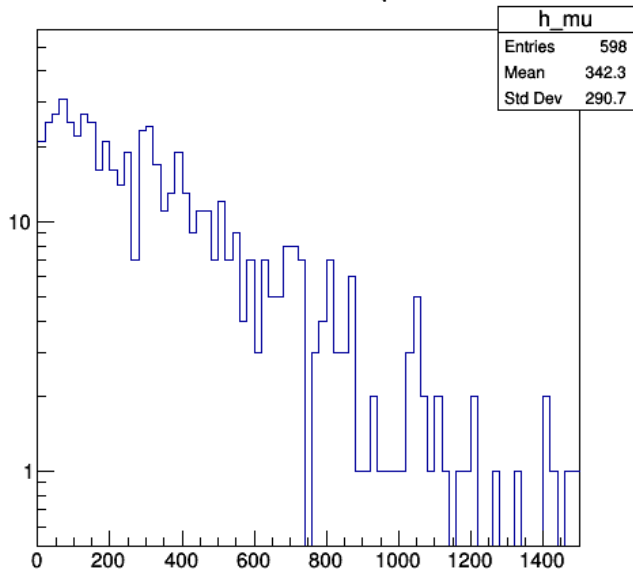
Energy nu in acceptance



Primary particles



Momentum μ



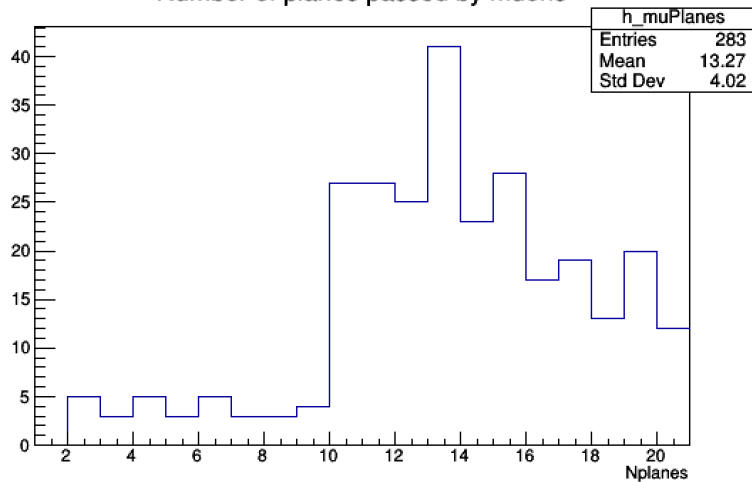
Penetrating power in terms of walls

Fairship simulation by Annarita

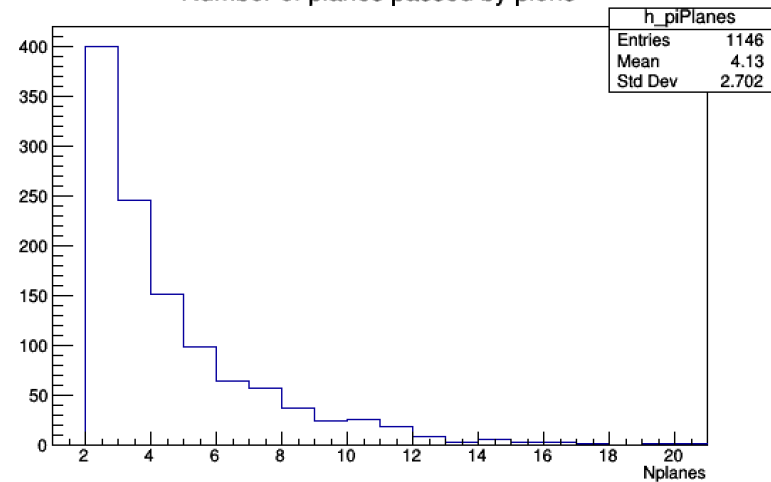
15000 events simulated, 288 in acceptance

Interacting in the first 12 walls

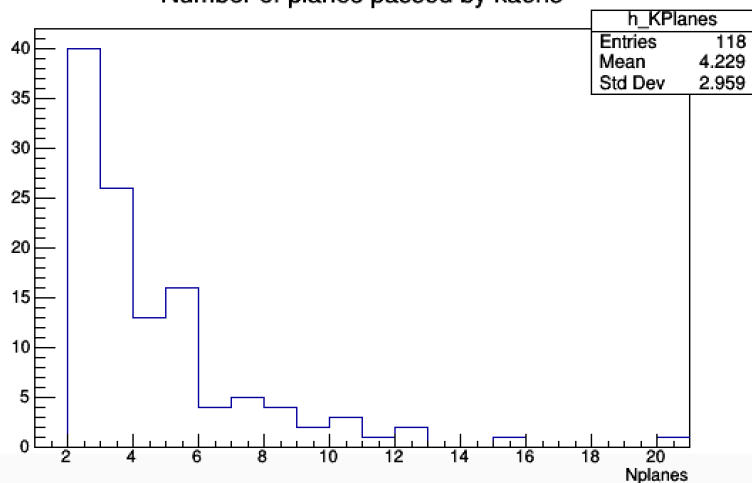
Number of planes passed by muons



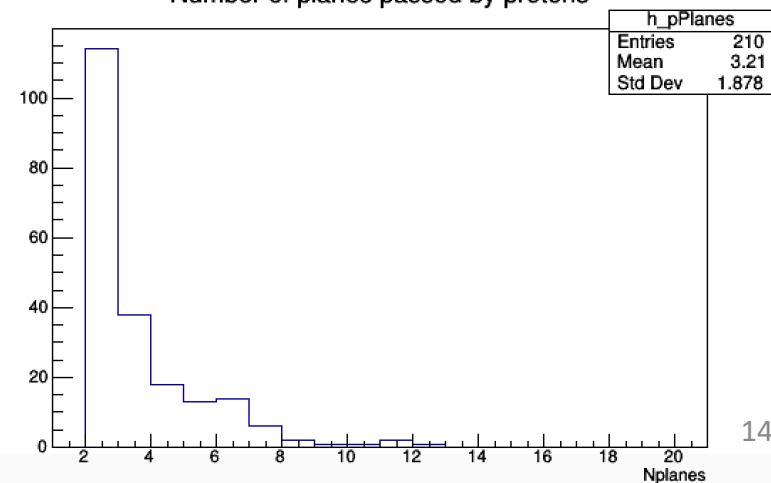
Number of planes passed by pions



Number of planes passed by kaons

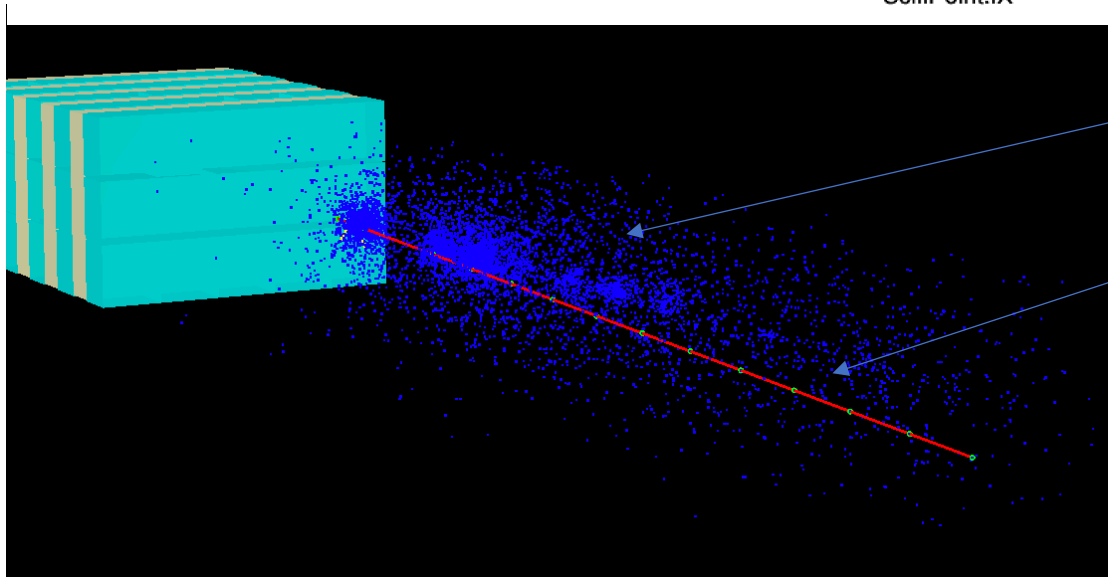
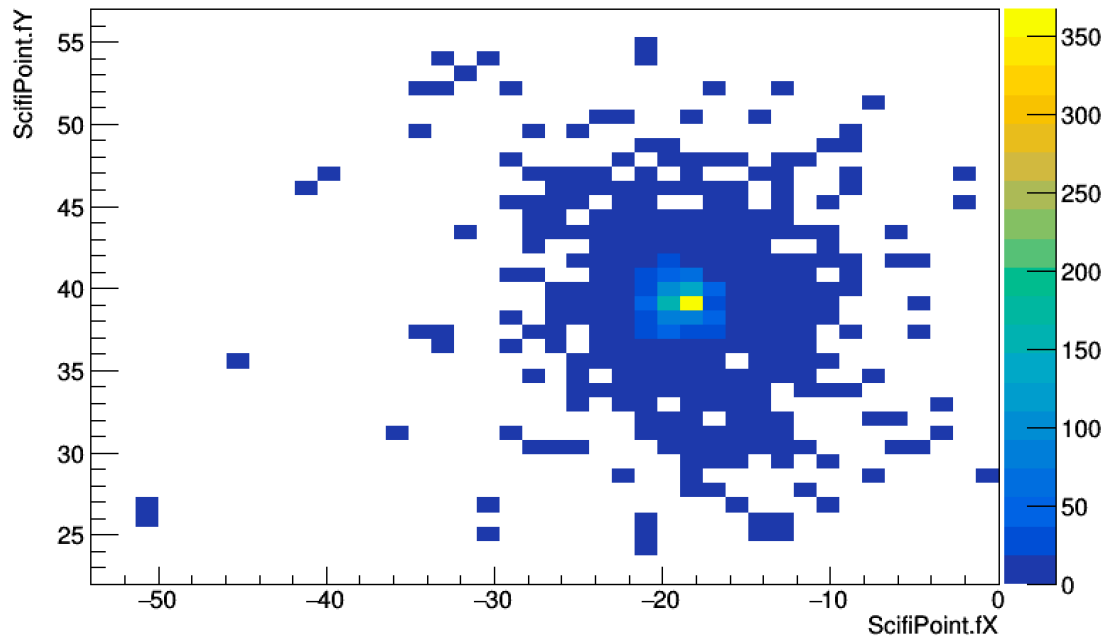


Number of planes passed by protons



One event as an example

Hits in the first electronic detector

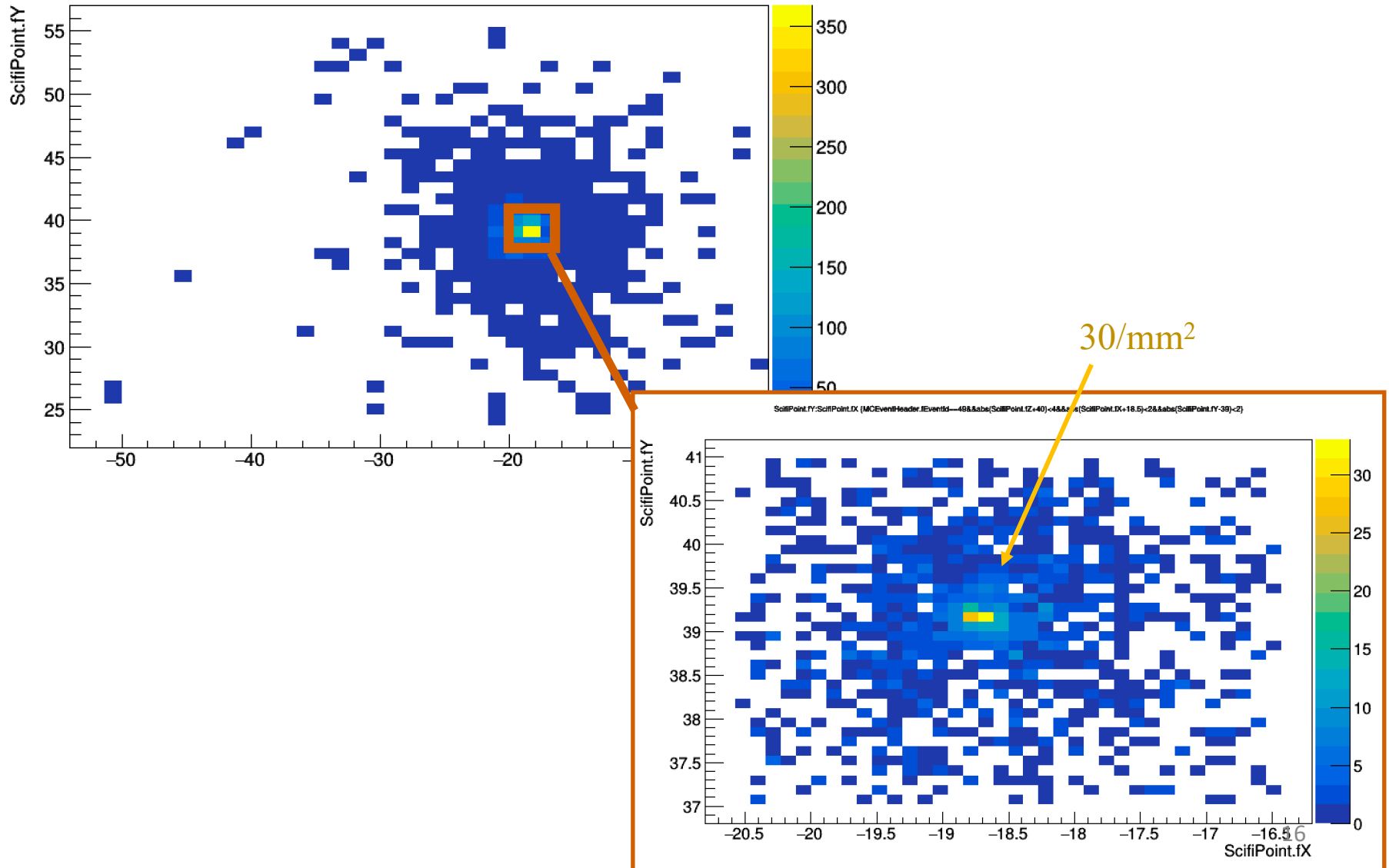


Particle hits

Muon track

One event as an example

Hits in the first electronic detector



Event classification

Npl cut	N 0 μ	N 1 μ	N 2 μ	N 3 μ	N > 3 μ
> 12	109	168	9	2	0

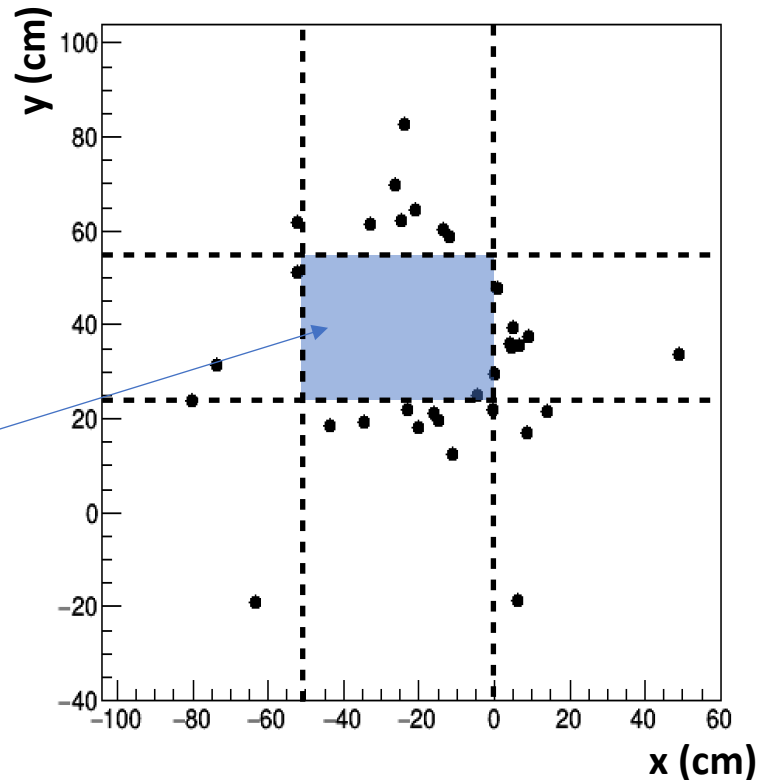
Purity = $(N_{\mu \rightarrow \mu}) / N_{1\mu} = 162 / 168 = 0.96$

$\epsilon \sim 60\%$

Main reason is insufficient thickness
 Currently 8 walls (~ 50 cm Fe equivalent)
 To be increased \rightarrow 80 cm Fe

Muon identification efficiency
 Can be largely optimised
 Also by increasing the size (~ 10 cm)

Interactions sufficiently upstream
 With unidentified muon
 Projected position muon last wall zy projection



Remarks

- Optimise the geometry (more passive material at the end)
- Fix the number of muon stations in the downstream part (3 sufficient)
- Consider the infrastructure required (Pablo)