

Probing axion-like particles with reactor neutrino experiments

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in collab. with D. Aristizabal, V. De Romeri & L.J. Flores



Operational Programme
Human Resources Development,
Education and Lifelong Learning

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Outline

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ALP motivations

Axions: *Nambu-Goldstone bosons from the breaking of a color anomalous global chiral $U(1)$ symmetry which is spontaneously broken in the vacuum.*

[Peccei & Quinn, PRL 38 (1977) 1440], [Weinberg, PRL 40 (1978) 223], [Wilczek, PRL 40 (1978) 279]

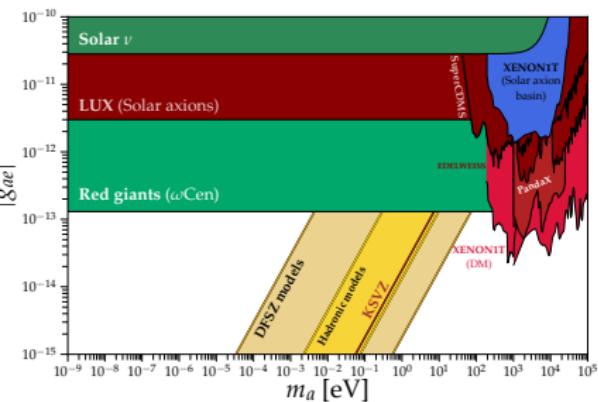
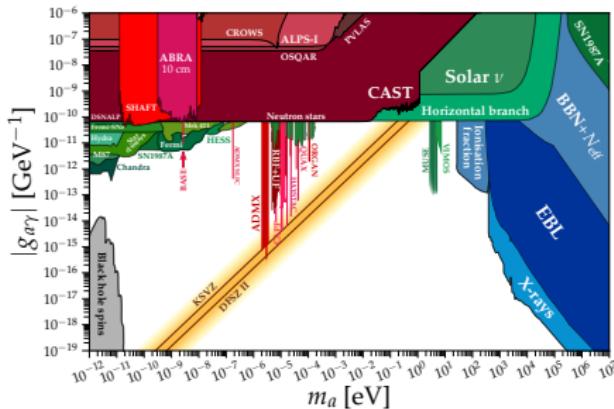
- solution to the strong CP problem
- dark matter candidate
- m_a and f_a are related $\rightarrow m_a = 5.7 \left(10^{12} \text{GeV}/f_a\right) \mu\text{eV}$

Axion-like particles (ALPs): *Pseudo Nambu-Goldstone bosons of spontaneously broken global symmetries.*

- Lepton symmetry: Majoron [Chikashige et. al, PLB 98 (1981) 2651981]
- Family symmetry: Familon [Wilczek, PRL 49 (1982) 1549]
- Flavor symmetry: Flavon
- m_a and f_a are not related \rightarrow mass does not arise from QCD effects

Current status of ALP-related experimental searches

- **helioscopes** (CAST) & **haloscopes** (Abracadabra, ADMX, CASPER...)
- **interferometry** (ADBC, DANCE) & **polarization exps** (PVLAS)
- **beam dump & fixed target** experiments (FASER, LDMX, NA62, NA64..)
- **colliders & dark matter DD** experiments (XENON, LUX, CDMS..)
- **astrophysical observations** (Stellar energy-losses)



plots from: C. O'Hare (<https://github.com/cajohare/AxionLimits>) 10.5281/zenodo.3932430
and references therein

Our goal: Explore ALPs in view of reactor neutrino experiments via nuclear and electron recoil measurements [Dent et al. PRL. 124 (2020) 21, 211804]

ALP production @ nuclear reactors

Nuclear reactors utilized as a high intensity photon flux

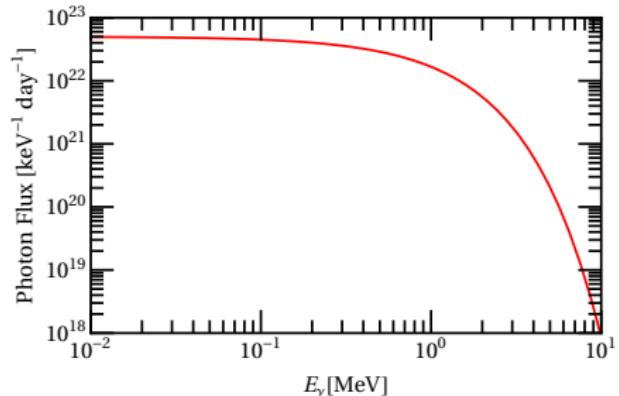
Photon flux

$$\frac{d\Phi_{\gamma'}}{dE_{\gamma'}} = \frac{5.8 \times 10^{17}}{\text{MeV} \cdot \text{sec}} \left(\frac{P}{\text{MW}} \right) e^{-1.1 E_{\gamma'} / \text{MeV}}$$

- P: reactor power in MW

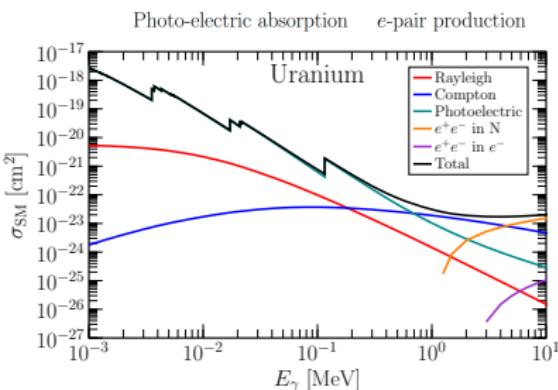
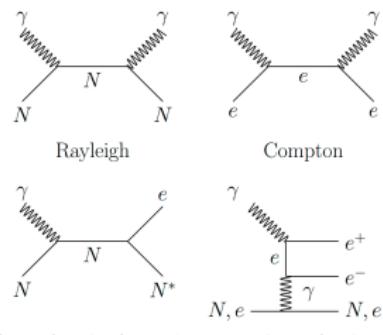
[Bechteler et al., Technical Report, Inst. fuer Kernphysik (1984)]

Reactor Power P=1 GW



SM processes

[<https://physics.nist.gov/PhysRefData/Xcom/html/xcom1.html>]

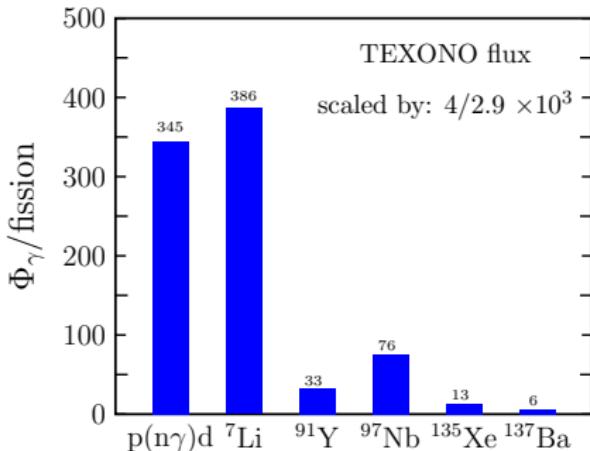
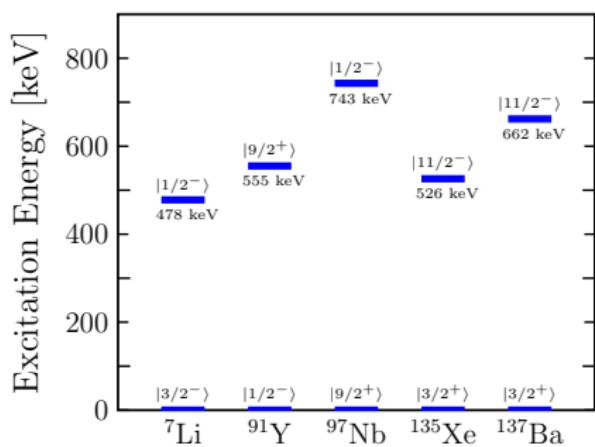


a fraction of these photons can be converted into ALPs

Nuclear reactors utilized as a high intensity photon flux

Photon flux from nuclear transitions

TEXONO Collab., Phys. Rev. D 75 (2007) 052004



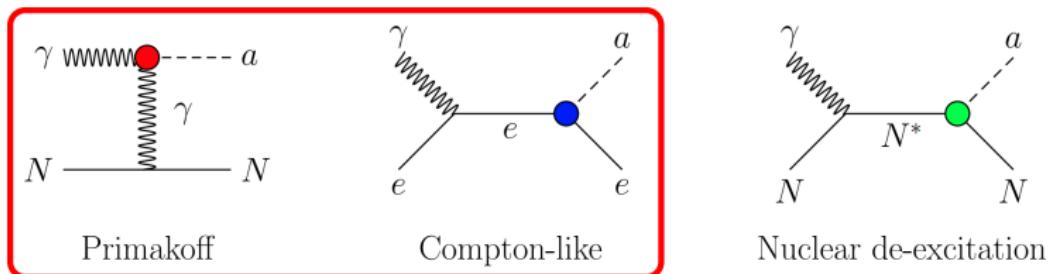
- (mainly) M1 transition from the excited state to the ground state of ^{7}Li
- M4 transitions from the excited to the ground state of ^{91}Y , ^{97}Nb , ^{135}Xe and ^{137}Ba
- thermal neutron capture on proton in the cooling water, $p + n \rightarrow d + \gamma$. The deuteron ground state has magnetic dipole and electric quadrupole moments, the emitted γ is therefore mainly M1

a fraction of these photons can be converted into ALPs

ALP production mechanisms

Phenomenological parametrization via dim-5 effective operators

$$\mathcal{L} = -\frac{1}{4} \textcolor{red}{g_{a\gamma\gamma}} a F_{\mu\nu} \tilde{F}^{\mu\nu} - i \textcolor{blue}{g_{ae\bar{e}}} a \bar{e} \gamma_5 e - i a \bar{n} \gamma_5 \left(\textcolor{green}{g_{ann}^{(0)}} + \tau_3 \textcolor{green}{g_{ann}^{(1)}} \right) n$$



Continuous ALP flux

$$\frac{d\Phi_a^P}{dE_a} = \mathcal{P}_{\text{surv}} \int_{E_{\gamma',\text{min}}}^{E_{\gamma',\text{max}}} \frac{1}{\sigma_{\text{Tot}}} \frac{d\sigma_{\text{ALP}}^{\text{prod}}}{dE_a}(E_{\gamma'}, E_a) \frac{d\Phi_{\gamma'}}{dE_{\gamma'}} dE_{\gamma'} , \quad \text{with } \sigma_{\text{Tot}} = \sigma_{\text{SM}} + \sigma_{\text{ALP}}^{\text{prod}}$$

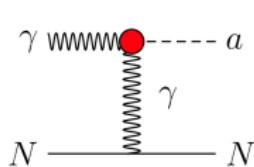
[Dent et al. PRL 124 (2020) 21, 211804]

- Survival probability, assuring that the ALP flux reaches the detector: $\mathcal{P}_{\text{surv}} = e^{-LE_a/|\vec{p}_a|\tau}$
- L : distance between the reactor and detector
- τ : ALP lifetime in the fixed target frame

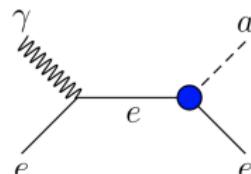
ALP production mechanisms

Phenomenological parametrization via dim-5 effective operators

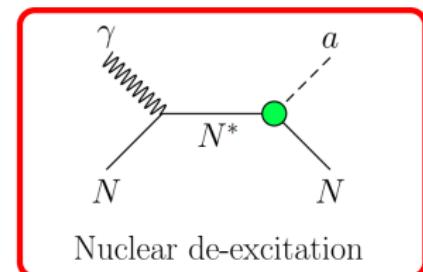
$$\mathcal{L} = -\frac{1}{4} \textcolor{red}{g_{a\gamma\gamma}} a F_{\mu\nu} \tilde{F}^{\mu\nu} - i \textcolor{blue}{g_{ae\bar{e}}} a \bar{e} \gamma_5 e - i a \bar{n} \gamma_5 \left(\textcolor{green}{g_{ann}^{(0)}} + \tau_3 \textcolor{green}{g_{ann}^{(1)}} \right) n$$



Primakoff



Compton-like



Nuclear de-excitation

Monochromatic ALP flux for the i -th transition

$$\left(\frac{d\Phi_a^{MT}}{dE_a} \right)_i = \phi_a^i \delta(E_{\gamma'} - E_a) = R_f \Phi_{\gamma}^i \left(\frac{\Gamma_a}{\Gamma_{\gamma}} \right)_i \mathcal{P}_{\text{surv}} \delta(E_{\gamma'} - E_a) \quad (i = p(n, \gamma)d, \text{ MJ}) ,$$

- fission rate: R_f
- photon flux per fission: Φ_{γ}^i
- branching ratio of ALP to photon e mission in the nuclear transitions: $\left(\frac{\Gamma_a}{\Gamma_{\gamma}} \right)_i$

ALP production cross sections

$g_{a\gamma\gamma}$ coupling

- Primakoff scattering: $\gamma + N \rightarrow a + N$ [Aloni et al. PRL 123 (2019) 7, 071801]

$$\frac{d\sigma_{\text{Prim}}^{\text{prod}}}{dt} = 2\alpha Z^2 F^2(t) g_{a\gamma\gamma}^2 \frac{M_N^4}{t^2(M_N^2 - s)^2(t - 4M_N^2)^2} \left\{ m_a^2 t(M_N^2 + s) - m_a^4 M_N^2 \right. \\ \left. - t [(M_N^2 - s)^2 + st] \right\}$$

Primakoff scattering: $E_\gamma \simeq E_a$ photon energy is coherently converted into ALP energy

g_{aee} coupling

- Compton-like scattering: $\gamma + e^- \rightarrow a + e^-$ [Brodsky et al. PRL 56 (1986) 1763]

$$\frac{d\sigma_{\text{Compt}}^{\text{prod}}}{dE_a} = \frac{Z\pi g_{aee}^2 \alpha x}{4\pi(s - m_e^2)(1 - x)E_{\gamma'}} \left[x - \frac{2m_a^2 s}{(s - m_e^2)^2} + \frac{2m_a^2}{(s - m_e^2)^2} \left(\frac{m_e^2}{1 - x} + \frac{m_a^2}{x} \right) \right],$$

$$\text{where } x = 1 - \frac{E_a}{E_{\gamma'}} + \frac{m_a^2}{2E_{\gamma'} m_e}.$$

ALP production from MJ transitions

g_{ann} coupling

- neutron capture isovector $M1$ transitions ($\text{pn} \rightarrow \text{d}\gamma$) depend only on kinematics

$$\left(\frac{\Gamma_a}{\Gamma_\gamma} \right)_{\text{pn}} = \frac{1}{2\pi\alpha} \left(\frac{|\vec{p}_a|}{|\vec{p}_\gamma|} \right)^3 \left(\frac{g_{ann}^{(1)}}{\mu_1} \right)^2 ,$$

[Barroso, Mukhopadhyay, PRC C24 (1981) 2382]

- MJ transitions are nuclear structure dependent

$$\left(\frac{\Gamma_a}{\Gamma_\gamma} \right)_{MJ} = \frac{1}{\pi\alpha} \left(\frac{1}{1+\delta^2} \right) \left(\frac{J}{J+1} \right) \left(\frac{|\vec{p}_a|}{|\vec{p}_\gamma|} \right)^{2J+1} \left(\frac{g_{ann}^{(0)}\kappa + g_{ann}^{(1)}}{(\mu_0 - 1/2)\kappa + (\mu_1 + \eta)} \right)^2 .$$

- Isovector magnetic moment: $\mu_1 = \mu_p - \mu_n = 4.71 \mu_N$
- Isosinglet magnetic moment: $\mu_0 = \mu_p + \mu_n = 0.88 \mu_N$
- δ, η, κ are nuclear structure dependent

[TEXONO collab., PRD 75 (2007) 052004]

[Avignone III et al., PRD 35 (1987) 2752]

Typical ALP fluxes from a nuclear reactor

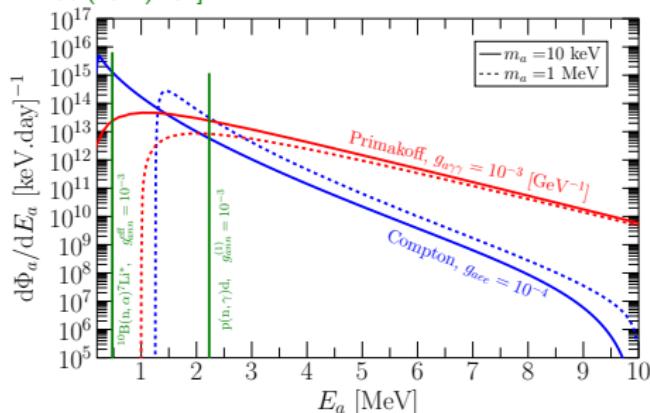
Experiment	Nuclear Reactor	Power [GW]
TEXONO [41]	Kuo-Sheng Nuclear Power Station	2.9
CONUS [37]	Brokdorf	3.9
ν GeN [72]	Kalinin Nuclear Power Plant	~ 1
MINER [36]	TRIGA 1	10^{-3}
ν CLEUS [38]	FRM2	4
Ricochet [39]	Chooz Nuclear Power Plant	8.54
RED-100 [40]	Kalinin Nuclear Power Plant	~ 1
SBC [73]	ININ (or Laguna Verde)	10^{-3} (2)
CONNIE [74]	Angra 2	3.8
vIOLETA [75]	Atucha II	2
SoLid [76]	BR2	$(0.4, 1) \times 10^{-1}$
NEON [77]	Hanbit Nuclear Power Plant	2.8

[Aristizabal, De Romeri, Flores, DKP, JHEP 03 (2021) 294]

Detector	Experiment	Material	$m_{\text{det}} [\text{kg}]$	$L [\text{m}]$
Semiconductor detectors (ionization)	TEXONO [41]	Ge	1.06	28
	CONUS [37]	Ge	1	17.1
	ν GeN [72]	Ge	1.6-5	10-12
Low temperature bolometers	MINER [36]	Ge, Si	4	1-2.5
	ν CLEUS [38]	CaWO ₄ , Al ₂ O ₃	10^{-2}	15-100
	Ricochet [39]	Ge, Zn	10	355/469
Liquid noble-gas detectors (TPC)	RED-100 [40]	Xe	100	19
	SBC [73]	LaBr, Xe	10	3/30
CCD	CONNIE [74]	Si	~ 0.05	30
	vIOLETA [75]	Si	1	12
Scintillators	SoLid [76]	⁶ LiF : ZnS(Ag)	1600	~ 7.6
	NEON [77]	NaI[Tl]	3.3-10	24

Typical reactor

- $P = 4 \text{ GW}$
- $L = 10 \text{ m}$



ALP detection

ALP detection cross sections

$g_{a\gamma\gamma}$ coupling

- inverse Primakoff scattering: $a + N \rightarrow \gamma + N$
same as the production cross section but a factor 2 larger due to spin

g_{aee} coupling

- inverse Compton-like scattering: $a + e^- \rightarrow \gamma + e^-$ [Avignone et al. PRD 37 (1988) 618-630]

$$\frac{d\sigma_{\text{Compt}}^{\det}}{dE_\gamma} = \frac{Z g_{aee}^2 \alpha E_\gamma}{4m_e^2 |\vec{p}_a|} \left| \frac{2(E_a + m_e - |\vec{p}_a| \cos\theta)^2}{|\vec{p}_a| y} \right| \times \left(1 + \frac{4m_e^2 E_\gamma^2}{y^2} - \frac{4m_e E_\gamma}{y} - \frac{4m_a^2 |\vec{p}_a|^2 m_e E_\gamma (1 - \cos^2\theta)}{y^3} \right), \quad y = 2m_e E_a + m_a^2$$

- axio-electric cross section: $a + e^- + Z \rightarrow e^- + Z$ [Derevianko et al. PRD 82 (2010) 065006]

$$\sigma_{\text{axioel}}^{\det} = \frac{g_{aee}^2}{\beta} \frac{3E_a^2}{16\pi\alpha m_e^2} \left(1 - \frac{\beta^{2/3}}{3} \right) \sigma_{\text{PE}}, \quad \beta = |\vec{p}_a|/E_a$$

ALP decays

$g_{a\gamma\gamma}$ coupling

- ALP diphoton decay:

$$\Gamma_{a \rightarrow 2\gamma} \equiv \Gamma(a \rightarrow \gamma\gamma) = \frac{g_{a\gamma\gamma}^2 m_a^3}{64\pi}$$

g_{aee} coupling

- ALP decay to electron pair:

$$\Gamma_{a \rightarrow e^+ e^-} = \frac{g_{aee}^2 m_a}{8\pi} \sqrt{1 - 4 \frac{m_e^2}{m_a^2}}$$

Number of events

scattering processes

$$\frac{d\mathcal{N}_X^{\text{scatt}}}{dE_a} = m_{\text{det}} \frac{N_T \Delta t}{4\pi L^2} \int \frac{d\Phi_a}{dE_a} \frac{d\sigma_X^{\text{det}}}{dE_\gamma} dE_\gamma, \quad X = \{\text{Prim., Compt.}\}$$

$$\frac{d\mathcal{N}_{\text{axioel}}}{dE_a} = m_{\text{det}} \frac{N_T \Delta t}{4\pi L^2} \frac{d\Phi_a}{dE_a} \sigma_{\text{axioel}}^{\text{det}}(E_\gamma, E_a)$$

decay processes

$$\frac{d\mathcal{N}_X^{\text{decay}}}{dE_a} = \frac{\mathcal{A} \Delta t}{4\pi L^2} \frac{d\Phi_a}{dE_a} \mathcal{P}_{\text{decay}}^X, \quad X = \{\text{Prim., Compt.}\}$$

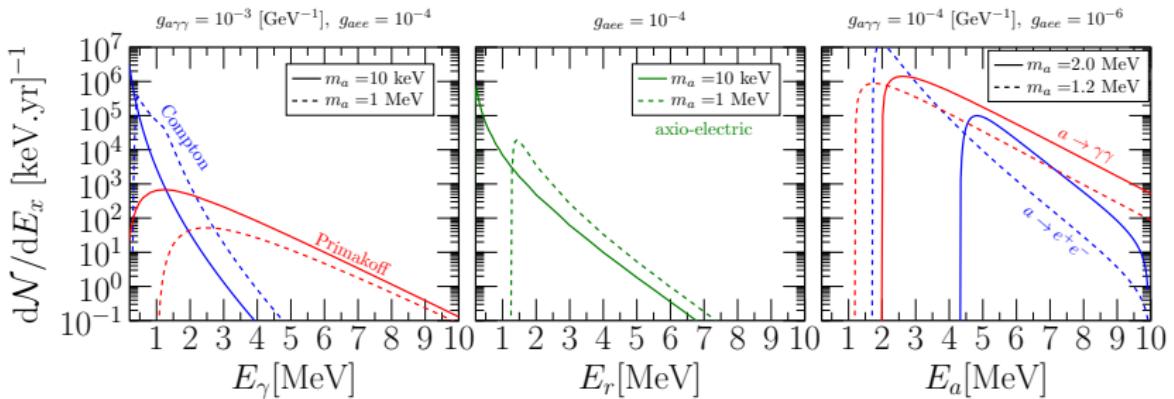
- P_{decay} : probability that the decay occurs within the detector

$$\mathcal{P}_{\text{decay}}^X = 1 - e^{-L_{\text{det}} E_a / |\vec{p}_a| \tau_X}$$

- $\mathcal{A} = L_{\text{det}}^2$ denotes the detector transverse area.

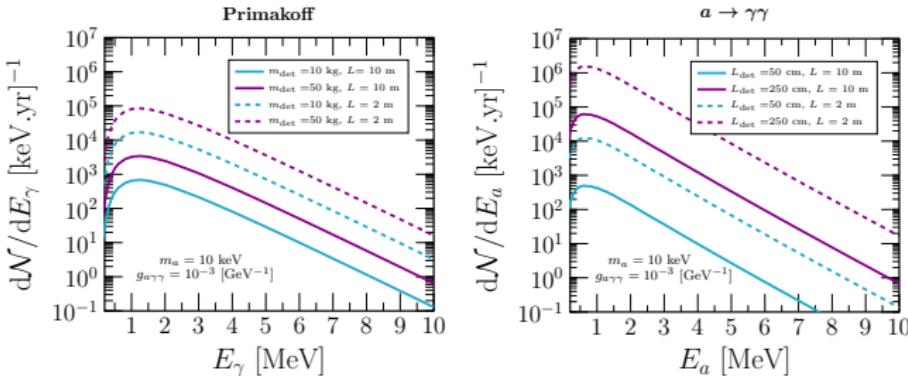
Number of events

varying the ALP mass



[Aristizabal, De Romeri, Flores, DKP, JHEP 03 (2021) 294]

varying the detector specifications

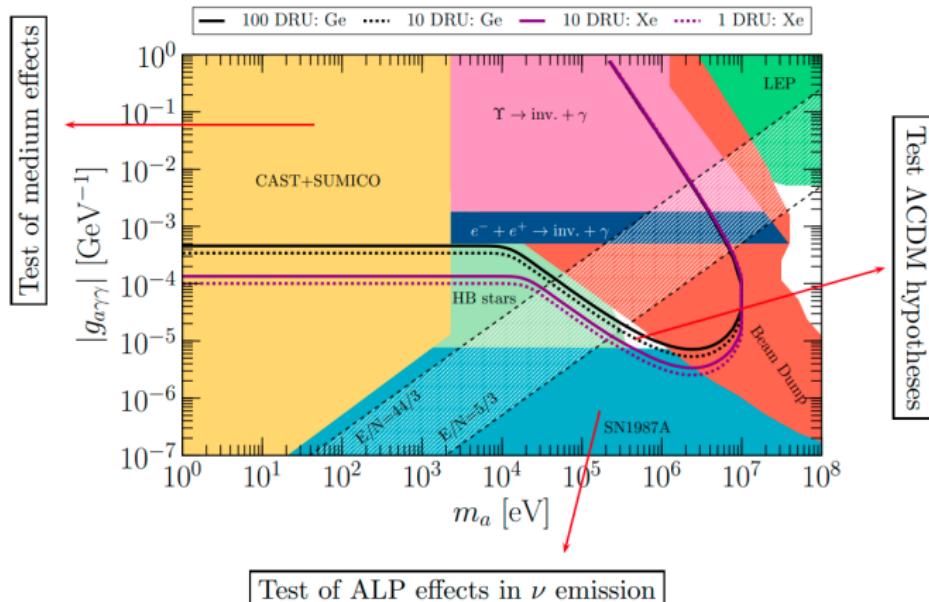


Sensitivities

Probing the $g_{a\gamma\gamma}$ coupling

Assumed detector specifications: current vs. future

P[GW]	PM	TM	$m_{\text{det}}[\text{kg}]$	$L[\text{m}]$	$L_{\text{det}}[\text{cm}]$	bkg [1/keV/day/kg]
4	^{235}U	Ge	10	10	50	10–100
8	^{235}U	Xe	10^3	10	140	1–10

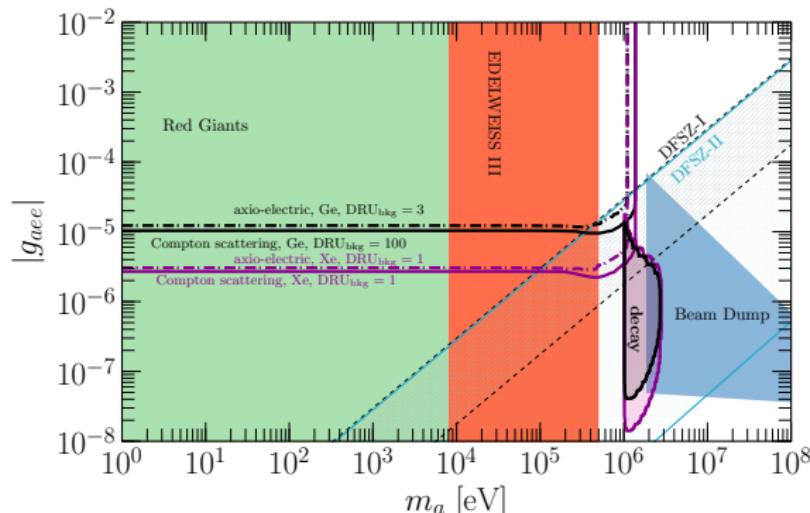


Probing the g_{aee} coupling

Assumed detector specifications: current vs. future

P[GW]	PM	TM	$m_{\text{det}}[\text{kg}]$	$L[\text{m}]$	$L_{\text{det}}[\text{cm}]$	bkg [1/keV/day/kg]
4	^{235}U	Ge	10	10	50	10–100
8	^{235}U	Xe	10^3	10	140	1–10

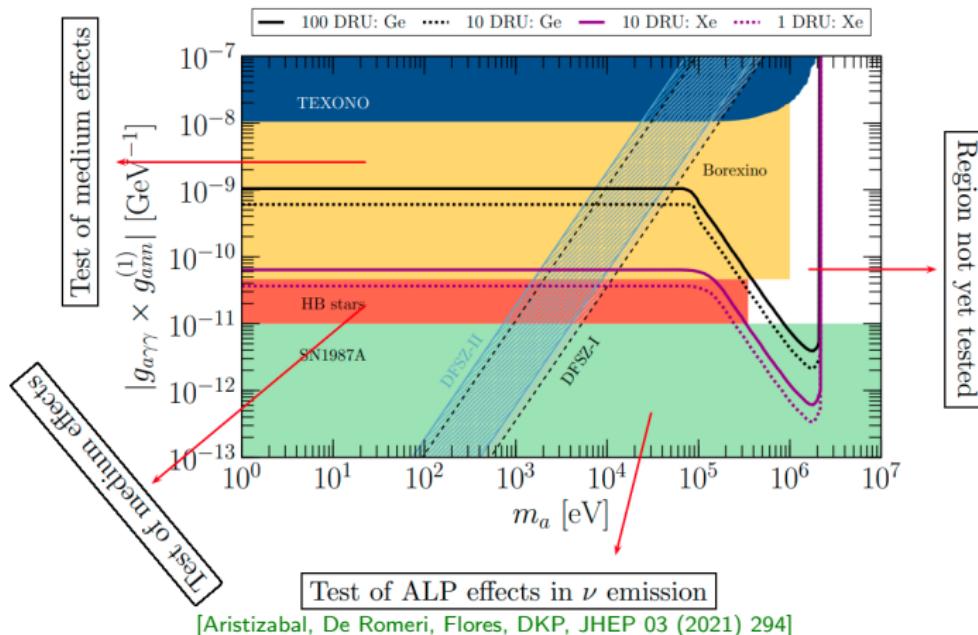
Testing: environmental effects in Red Giants, unexplored regions



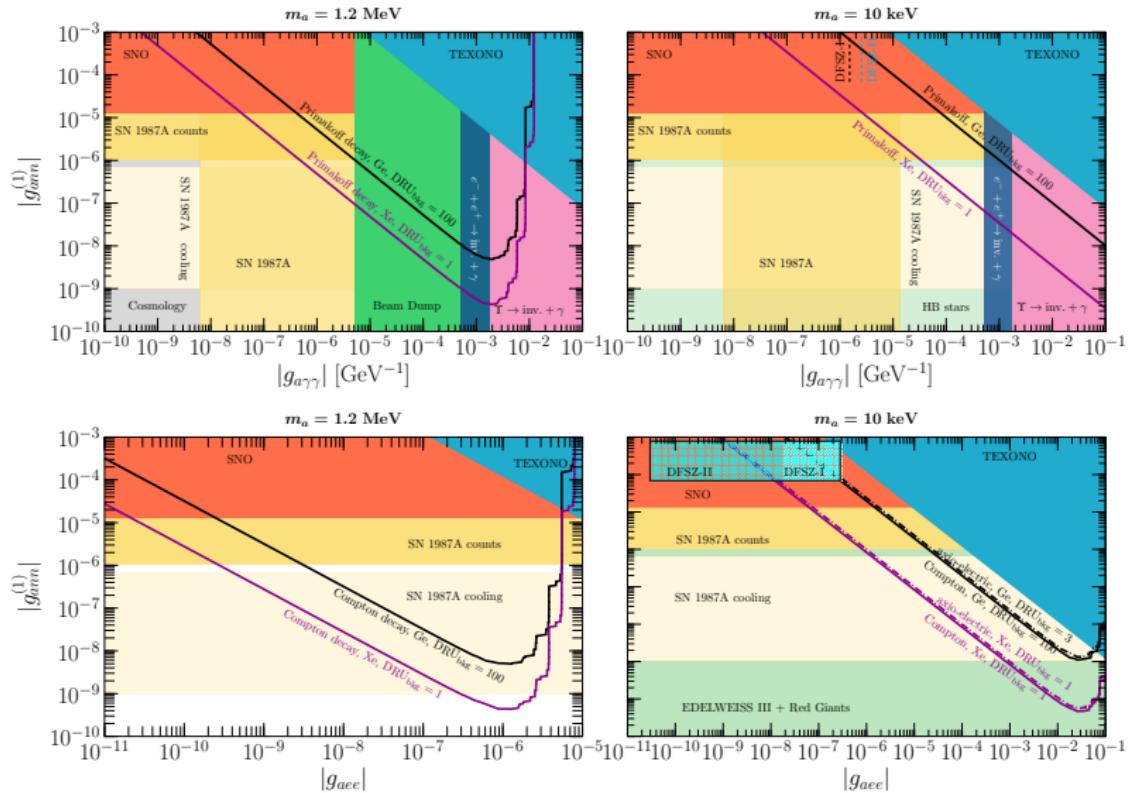
Probing the $g_{a\gamma\gamma}$ and g_{ann} couplings

Assumed detector specifications: current vs. future

P[GW]	PM	TM	$m_{\text{det}}[\text{kg}]$	$L[\text{m}]$	$L_{\text{det}}[\text{cm}]$	bkg [1/keV/day/kg]
4	^{235}U	Ge	10	10	50	10–100
8	^{235}U	Xe	10^3	10	140	1–10



Combined analyses



[Aristizabal, De Romeri, Flores, DKP, JHEP 03 (2021) 294]

complementary laboratory probes especially in the regions currently tested only by astrophysics

Summary

Summary ALP production and detection mechanisms considered

Scattering processes				
Process	Coupling	Prod	Det	
Primakoff	$\gamma + N \leftrightarrow a + N$	$g_{a\gamma\gamma}$	✓	✓
Compton-like	$\gamma + e^- \leftrightarrow a + e^-$	g_{aee}	✓	✓
Nuclear de-excitation	$\gamma + N \leftrightarrow N^* \rightarrow a + N$	g_{ann}	✓	✓
Axio-electric	$a + e^- + Z \rightarrow e^- + Z$	g_{aee}	✗	✓
e -pair production in N	$a + N \rightarrow e^- + e^- + N$	g_{aee}	✗	✓
e -pair production in e	$a + e^- \rightarrow e^- + e^+ + e^-$	g_{aee}	✗	✓
Decay processes				
Process	Coupling	Prod	Det	
γ -pair final state	$a \rightarrow \gamma + \gamma$	$g_{a\gamma\gamma}$	✗	✓
e -pair final state	$a \rightarrow e^- + e^+$	g_{aee}	✗	✓
n -pair final state	$a \rightarrow n + n$	g_{ann}	✗	✗

Reactor experiments can be used as ALP factories:

- extend the physics reach of reactor neutrino programmes
- probe ALPs with $m_a \leq 10$ MeV utilizing their intense photon flux
- complementary information on ALPs in the low-energy frontier

Thank you for your attention !