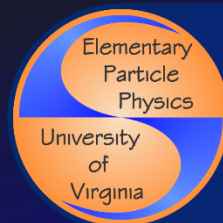
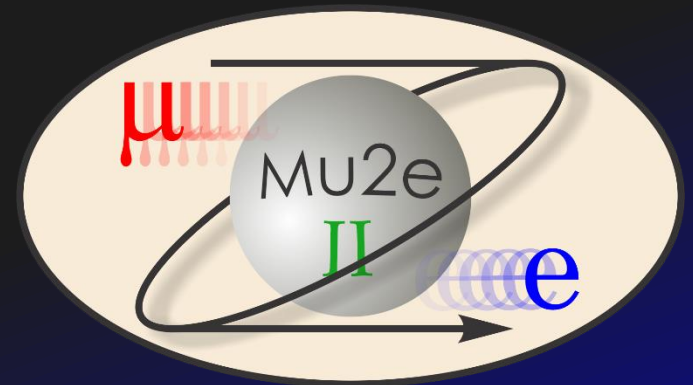


# Mu2e-II: The Mu2e Experiment in the PIP-II Era

E. Craig Dukes  
(for the Mu2e-II Collaboration)  
University of Virginia

NuFact 2021  
September 8, 2021



Frontier Physics Group  
University of Virginia



# Why We Think the Standard Model is Incomplete

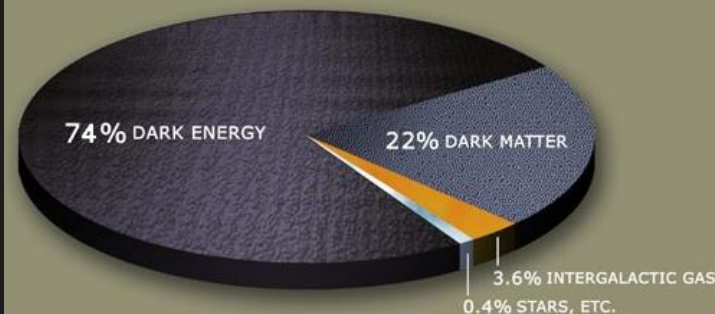
## Theory

- Quantum theory of gravity
- Origin of neutrino mass hierarchy
- Solution to hierarchy problem  $\Rightarrow$  supersymmetry, something else?



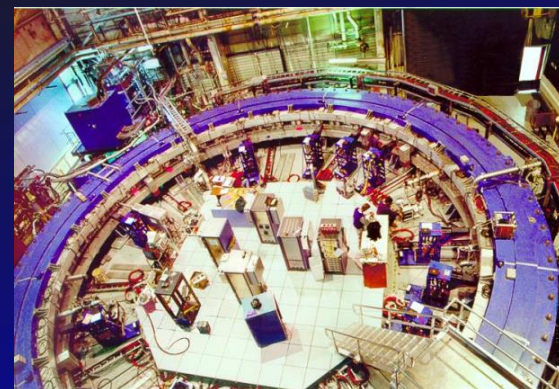
## Cosmology

- Matter-antimatter asymmetry in the universe
- Dark matter
- Dark energy

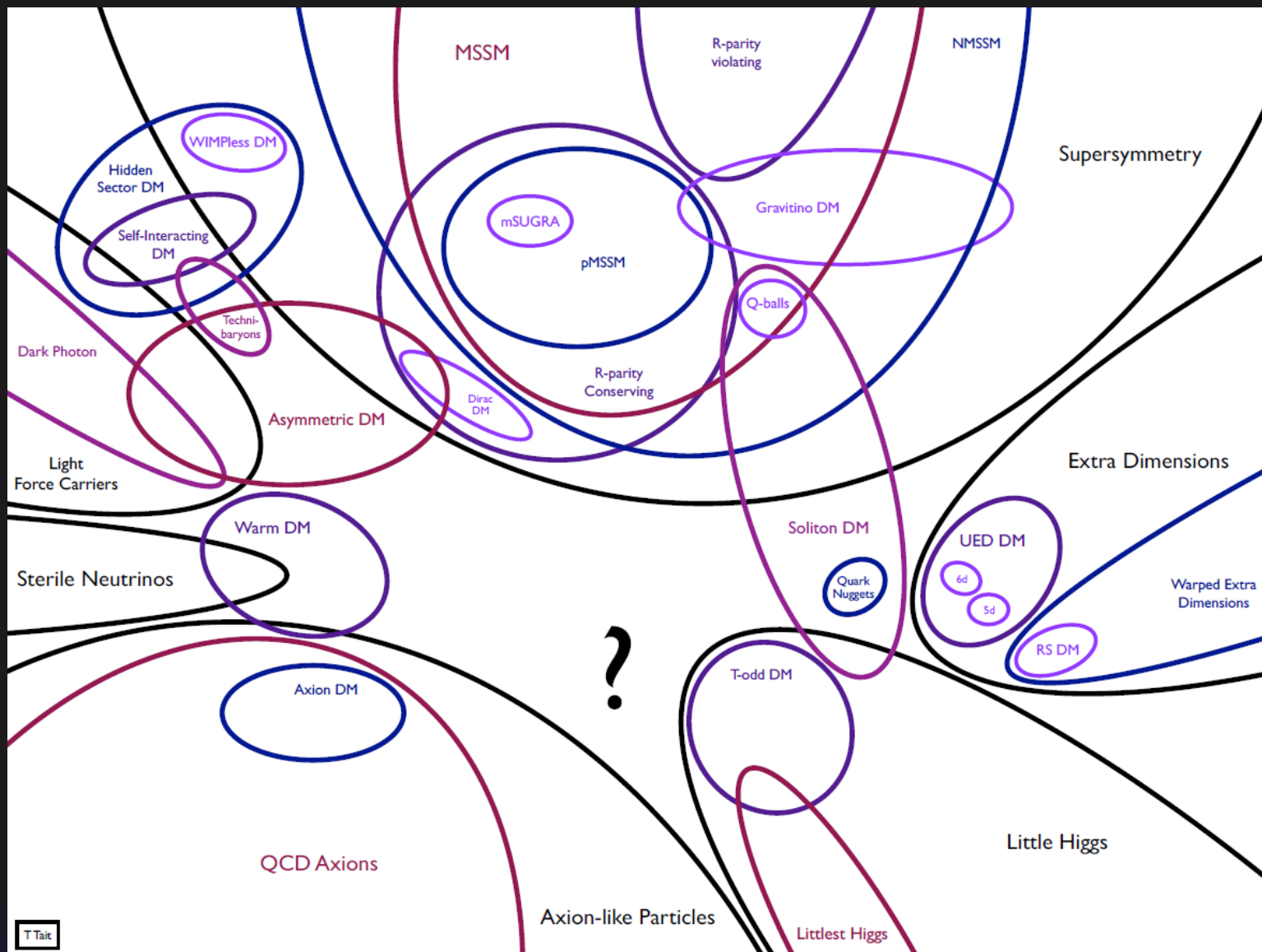


## Experiment

- Neutrino mass  $\Rightarrow$  first evidence of physics beyond the standard model
- Occasional hints appear, and often disappear: muon  $g-2$ ,  $B^+ \rightarrow l^+ l^-$ , NuTeV, CP phases in  $B_s$  mixing,  $D_s$  decay rates, W+jets, Top AFB



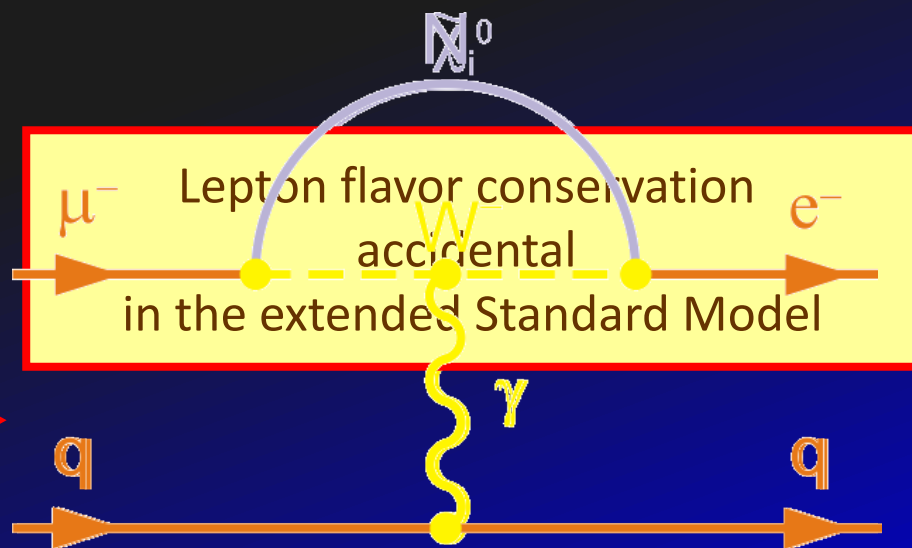
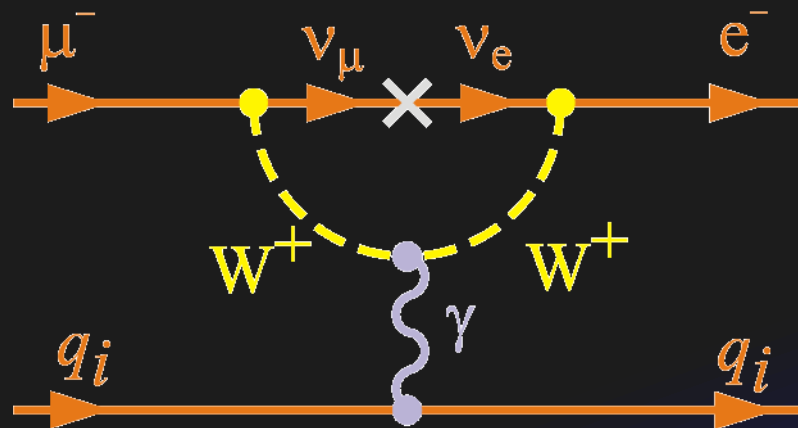
# No Lack of Theoretical Ideas, but Little Guidance



# Why Search for Charged Lepton Flavor Violation?

$$\mu^- N \rightarrow e^- N$$

- In Standard Model not there  $\Rightarrow$  neutrino mass discovery implies an unobservable  $10^{-52}$  rate
- Hence, any signal unambiguous evidence of new physics
- Exquisite sensitivities can be obtained experimentally
  - $\Rightarrow$  sensitivities that allow favored beyond-the-standard-model theories to be tested



**New heavy neutrino**



# Why Muon-to-Electron Conversion?

Probes of different  
SUSY and non-SUSY  
BSM models

★★★ Large effects

★★ Visible, but  
small

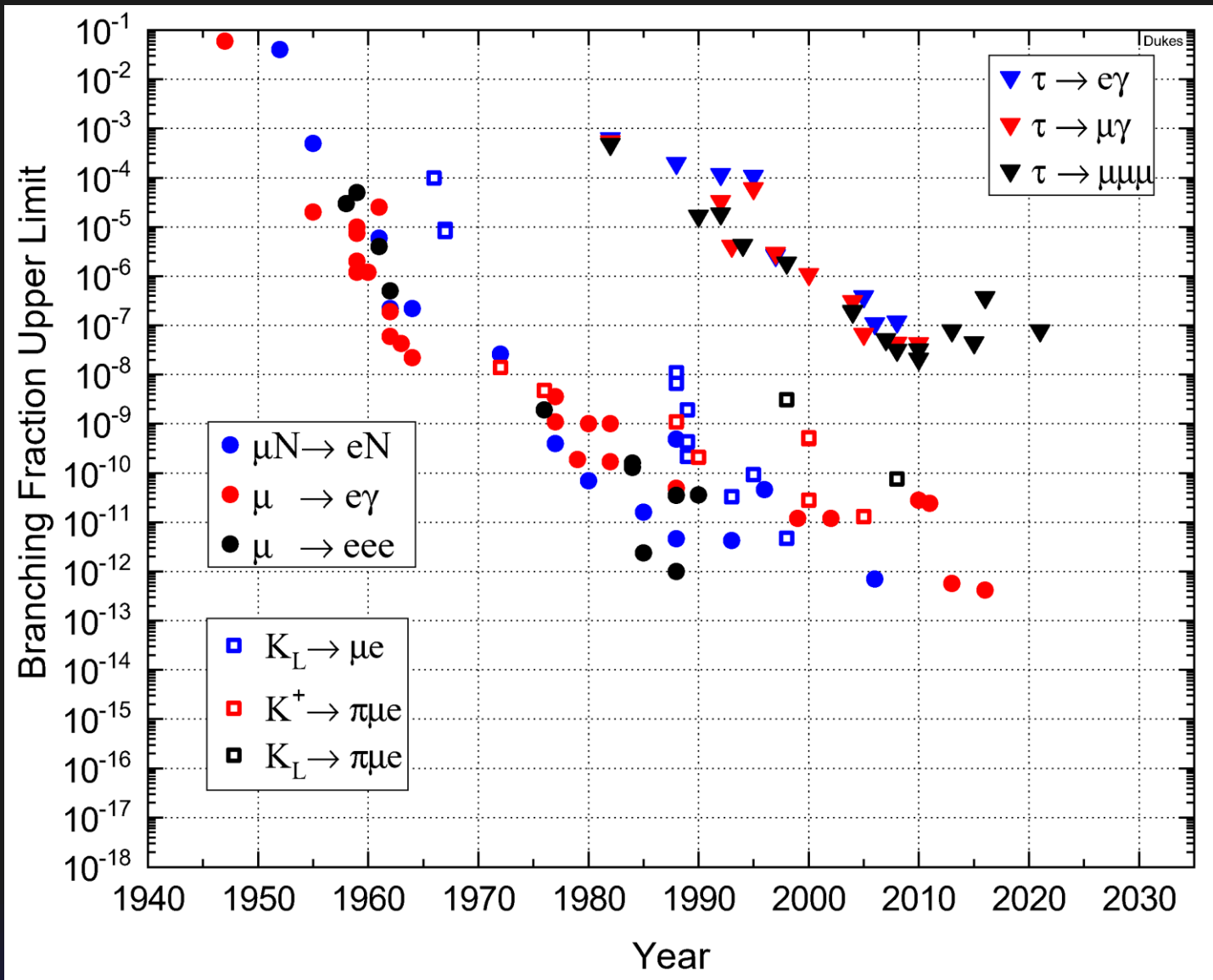
★ No sizable  
effect

Altmannshofer, Buras, et  
al., NPB 830, 17 (2010)

	AC	RVV2	AKM	$\delta$ LL	FBMSSM	LHT	RS
$D^0 - \bar{D}^0$	★★★	★	★	★	★	★★★	?
$\epsilon_K$	★	★★★	★★★	★	★	★★	★★★
$S_{\psi\phi}$	★★★	★★★	★★★	★	★	★★★	★★★
$S_{\phi K_S}$	★★★	★★	★	★★★	★★★	★	?
$A_{CP}(B \rightarrow X_s \gamma)$	★	★	★	★★★	★★★	★	?
$A_{7,8}(B \rightarrow K^* \mu^+ \mu^-)$	★	★	★	★★★	★★★	★★	?
$A_9(B \rightarrow K^* \mu^+ \mu^-)$	★	★	★	★	★	★	?
$B \rightarrow K^{(*)} \nu \bar{\nu}$	★	★	★	★	★	★	★
$B_s \rightarrow \mu^+ \mu^-$	★★★	★★★	★★★	★★★	★★★	★	★
$K^+ \rightarrow \pi^+ \nu \bar{\nu}$	★	★	★	★	★	★★★	★★★
$K_L \rightarrow \pi^0 \nu \bar{\nu}$	★	★	★	★	★	★★★	★★★
$\mu \rightarrow e \gamma$	★★★	★★★	★★★	★★★	★★★	★★★	★★★
$\tau \rightarrow \mu \gamma$	★★★	★★★	★	★★★	★★★	★★★	★★★
$\mu + N \rightarrow e + N$	★★★	★★★	★★★	★★★	★★★	★★★	★★★
$d_n$	★★★	★★★	★★★	★★	★★★	★	★★★
$d_e$	★★★	★★★	★★	★	★★★	★	★★★
$(g-2)_\mu$	★★★	★★★	★★	★★★	★★★	★	?

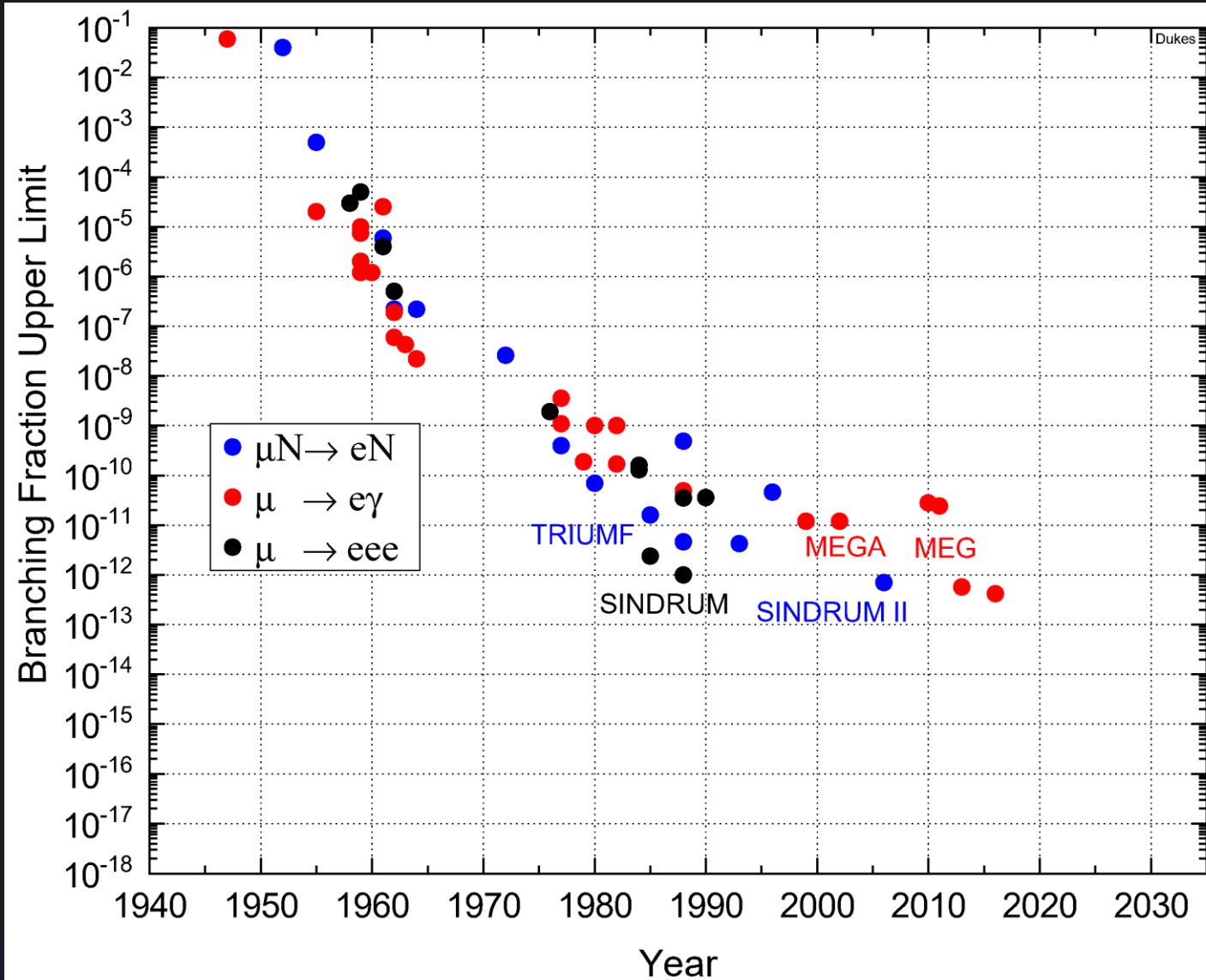
Table 8: “DNA” of flavour physics effects for the most interesting observables in a selection of SUSY and non-SUSY models ★★★ signals large effects, ★★ visible but small effects and ★ implies that the given model does not predict sizable effects in that observable.

# (Incomplete) History of CLFV Searches

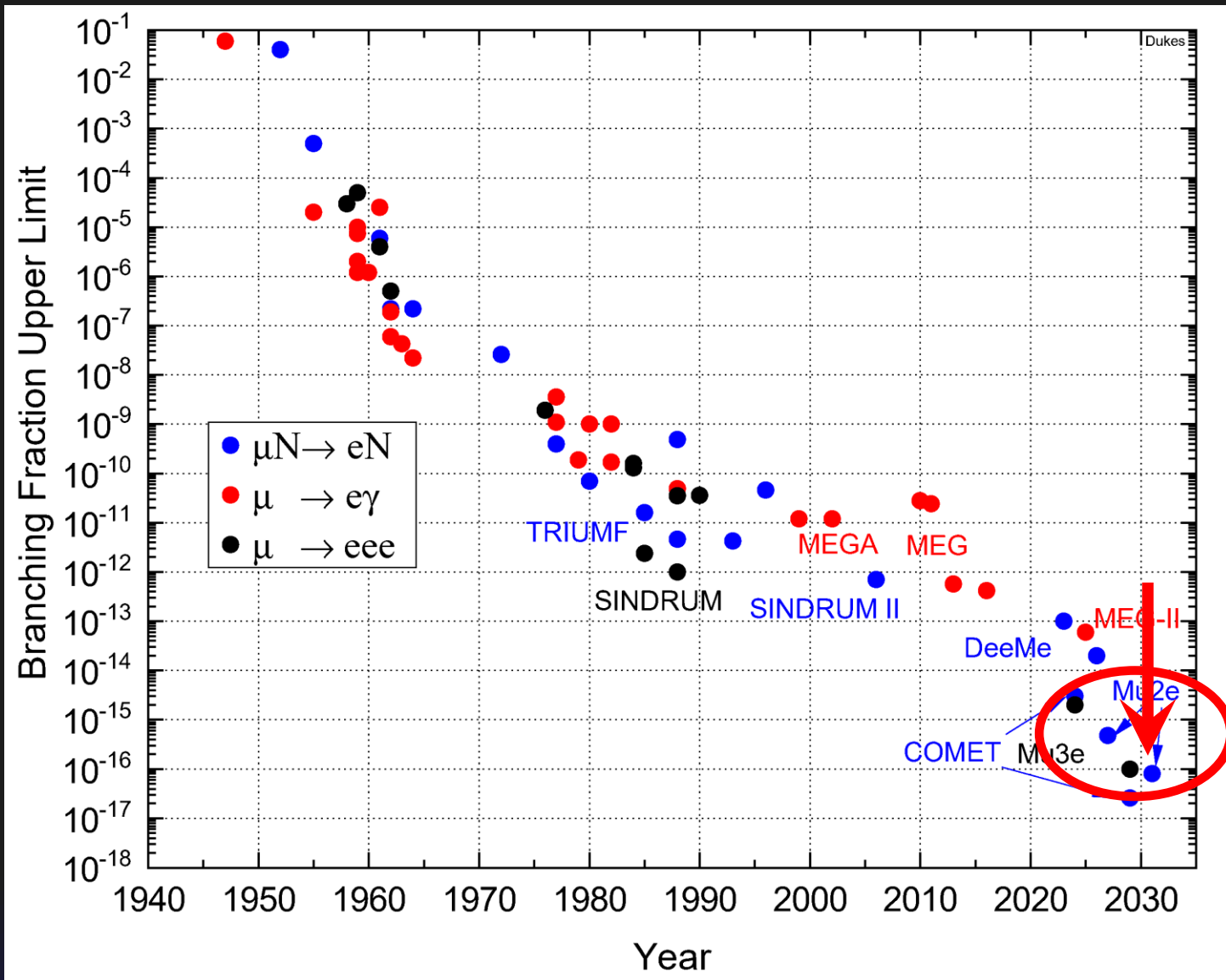




# History of Muon CLFV Searches

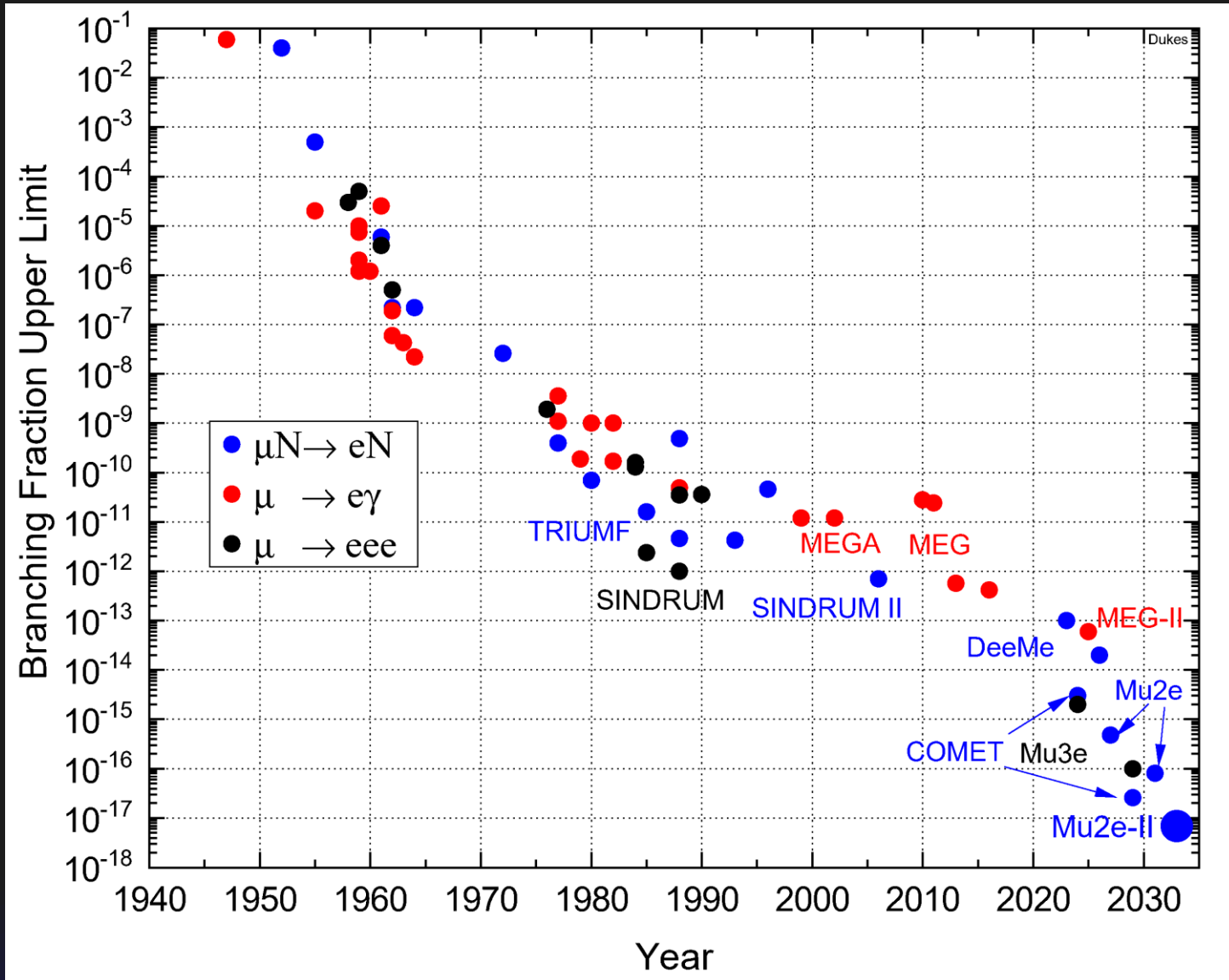


# Future Searches of Muon CLFV





# Mu2e-II Goal: $O(10^{-18})$ Single Event Sensitivity

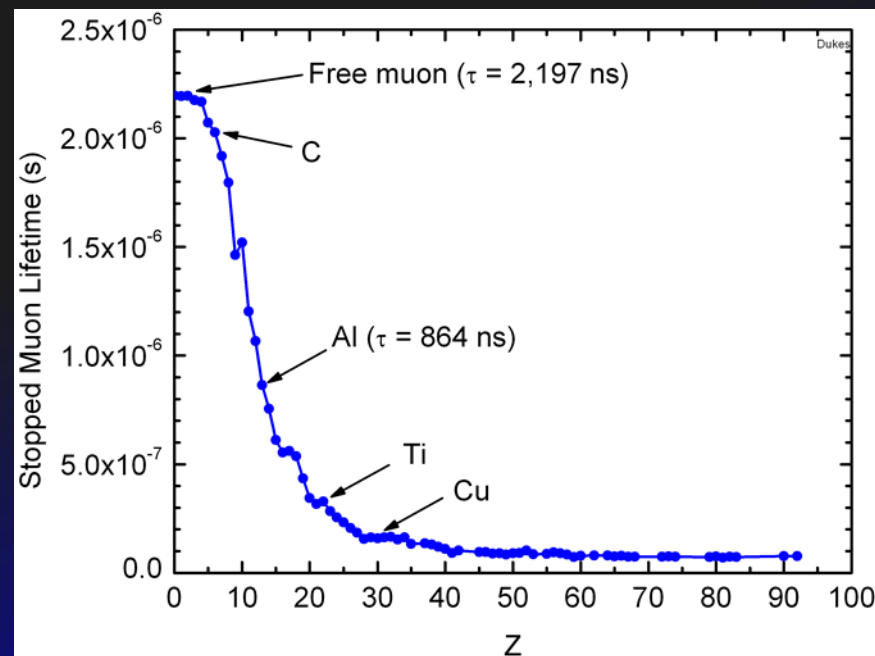
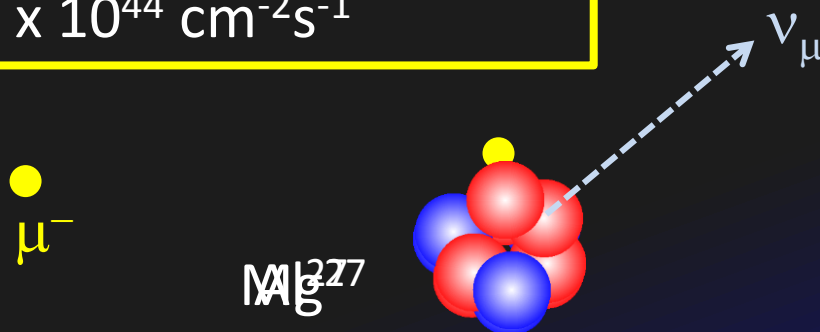


# How to Search for $\mu^-N \rightarrow e^-N$

- Stop muon in atom
- Muon rapidly ( $10^{-13}s$ ) cascades to 1S state
- Circles the nucleus for up to  $\sim 2 \mu s$  (in Al  $\tau = 864$  ns)
- Two things most likely happen:

1. muon is captured by the nucleus:  $\mu^- N_{A,Z} \rightarrow \nu_\mu N_{A,Z-}$

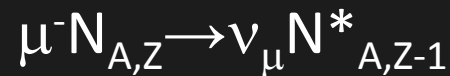
Single atom luminosity:  
 $\sim 1 \times 10^{44} \text{ cm}^{-2} \text{ s}^{-1}$



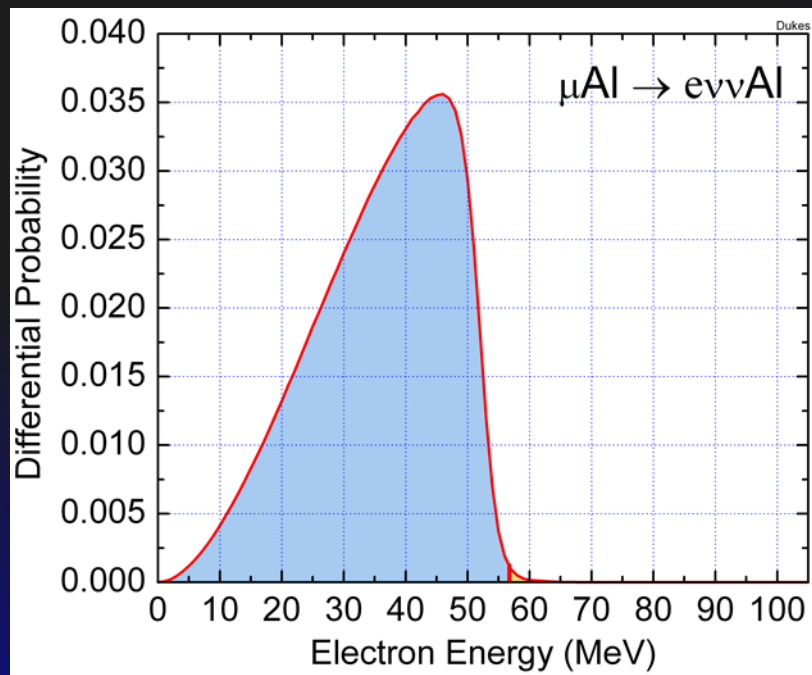
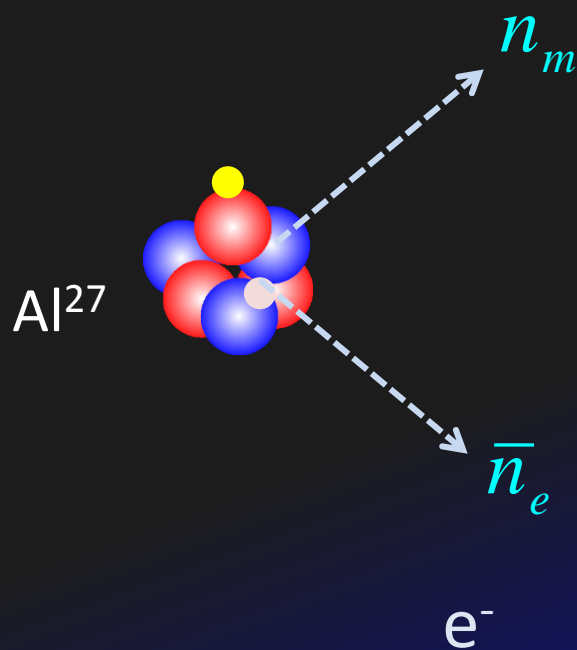
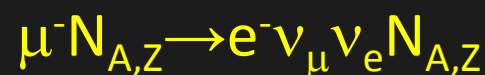
# How to Search for $\mu^-N \rightarrow e^-N$

- Stop muon in atom
- Muon rapidly ( $10^{-16}$ s) cascades to 1S state
- Circles the nucleus for up to  $\sim 2 \mu\text{s}$
- Two things most likely happen:

1. muon is captured by the nucleus:



2. muon decays in orbit:



# Mu2e Searching for a Third Process: $\mu^-N \rightarrow e^-N$

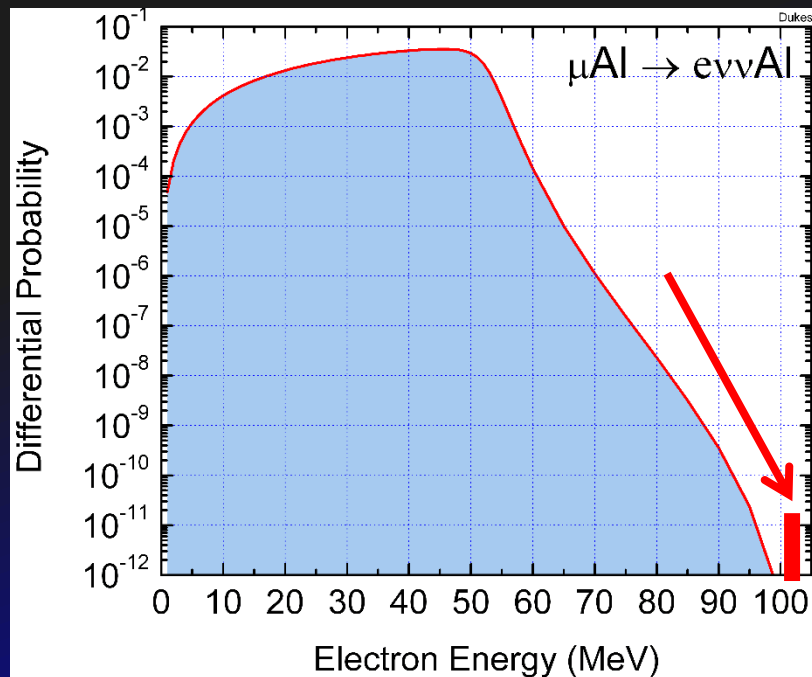
The muon turns into an electron  $\mu^-N \rightarrow e^-N$  leaving the nucleus in ground state

- signature single delayed ( $\tau = 864$  ns in Al) isolated electron
- Electron energy given by the rest mass of the muon minus the nucleus recoil energy and the binding energy:

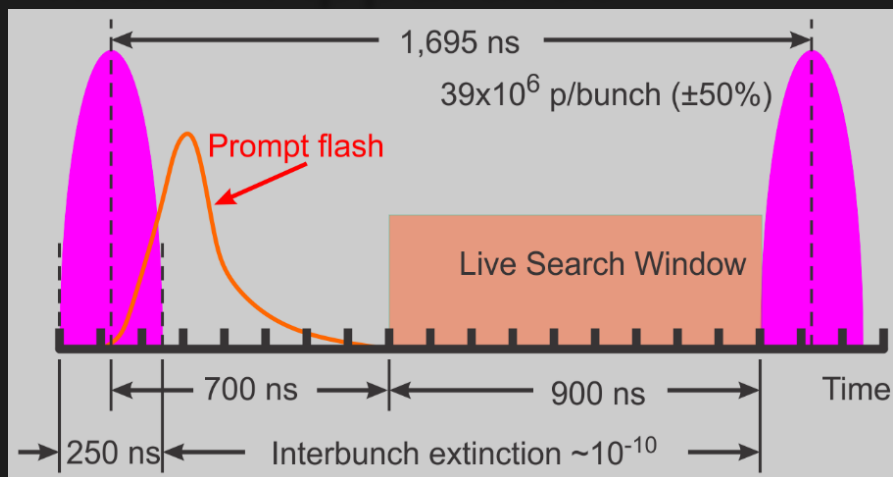
$$E_e = m_\mu - E_{NR} - E_b \sim 104.97 \text{ MeV (Al)}$$



e<sup>-</sup>



# Mu2e Apparatus

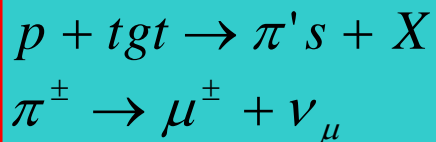
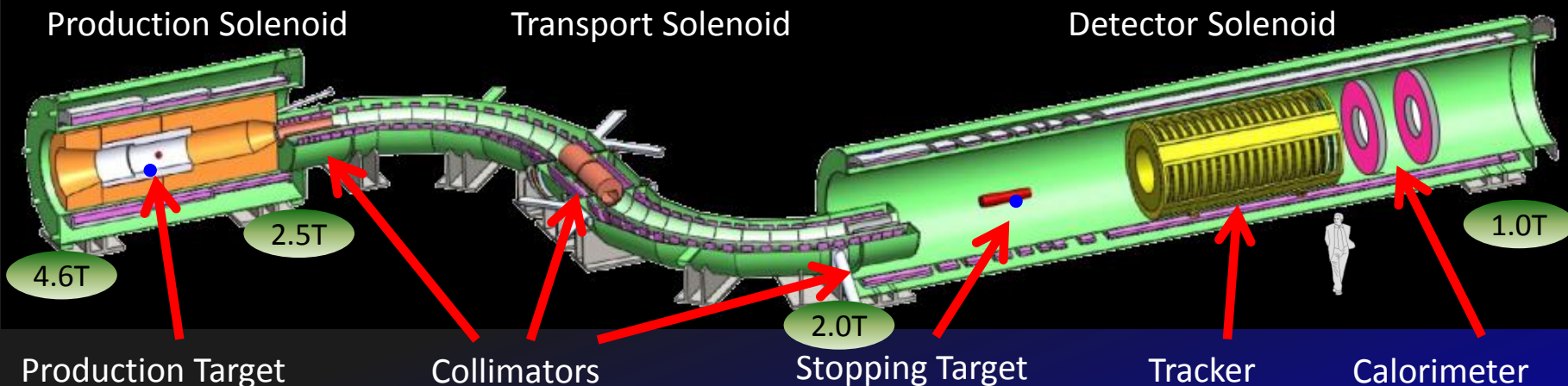


$39 \times 10^6$  protons every  $1.7 \mu\text{s}$   
 $74 \times 10^3 \mu^-$  stops every  $1.7 \mu\text{s}$   
 43 billion  $\mu^-$  stops/spill-second

← 25 m →

Muon Beam

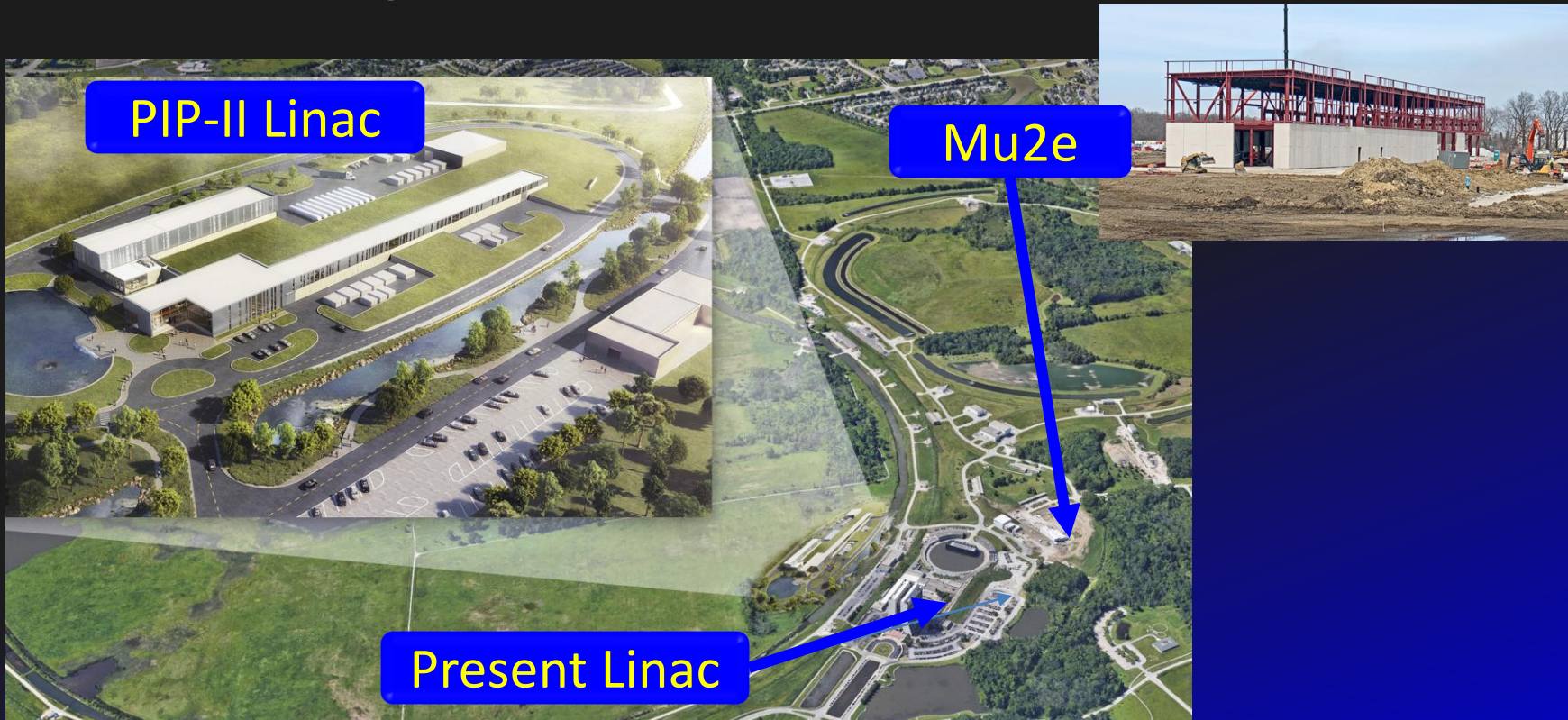
Spectrometer





# PIP-II: Fermilab Proton Improvement Plan II

- Present Fermilab Linac replaced with 800 MeV Continuous Wave SRF Linac
- High intensity  $H^-$  beam: up to  $4 \times 10^8$  p/bunch, 162.5 MHz bunch frequency
- Design driven by needs of the Fermilab neutrino program
  - However: LBNF/DUNE only needs  $\sim 1\%$  of the available beam a fraction of the time, limited by (the increased) Booster rep rate of 20 Hz
- Construction has begun; scheduled to end in 2027





# Mu2e-II: Goal of Mu2e in the PIP-II Era

We wish to seize the opportunity provided by upgrades of the Fermilab accelerator complex being built for DUNE to:

- **Increase sensitivity over Mu2e by 10X while keeping backgrounds < 1 event**
- This is to be done by:  **$O(10^{-18})$  SES**
  - ~3X increase in muon beam intensity (through ~30X p beam intensity)
  - ~3X increase in live time (through a better duty factor)

## Advantages:

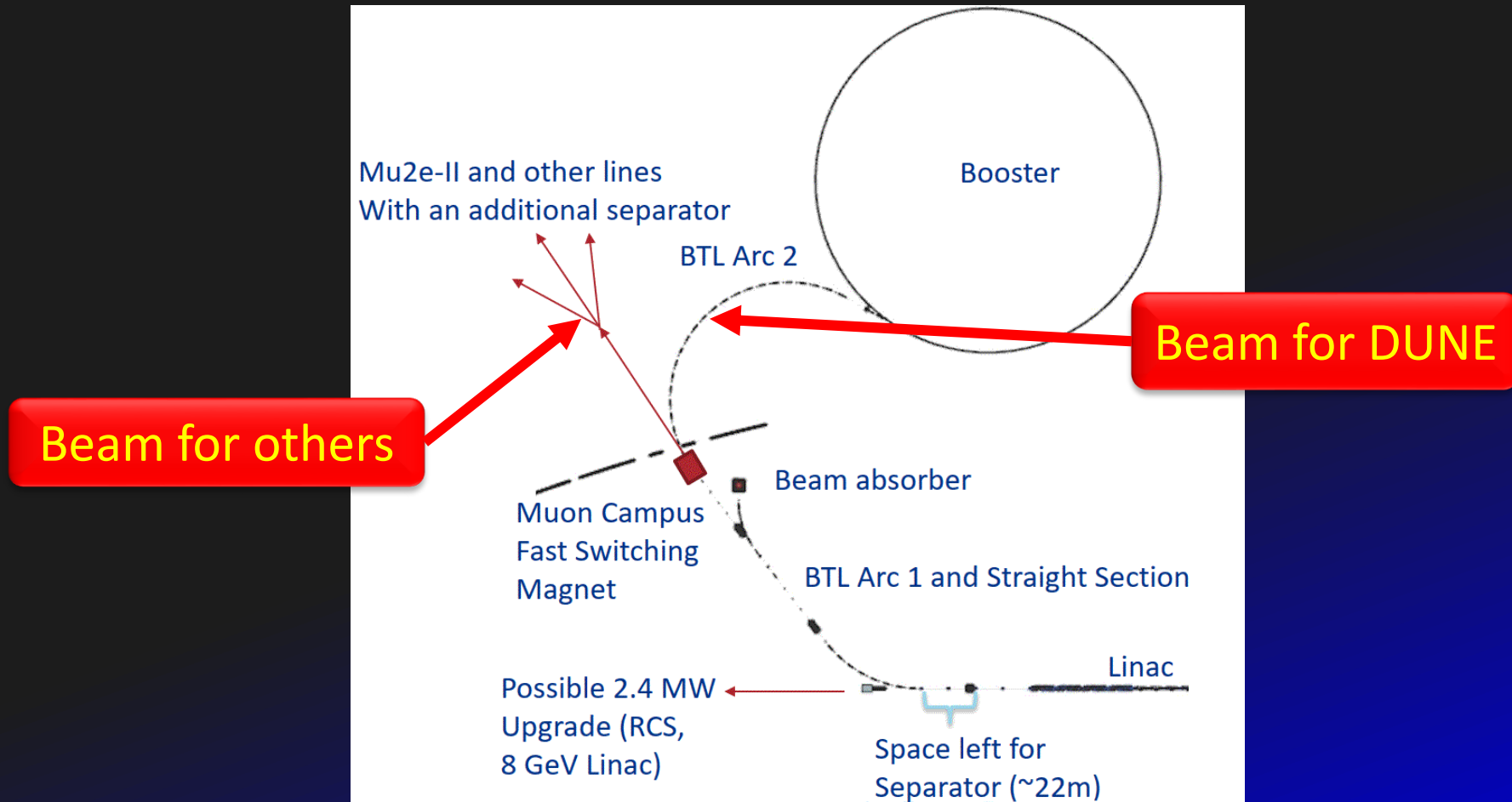
- Higher duty factor (~3X)
- More intense muon beam (~3X)
- Beam structure can be tuned to the needs of the Mu2e target choice
- Narrower beam pulse
- Lower energy eliminates anti-proton background

## Challenges:

- Getting the lower-energy beam on the production target
- Dealing with higher rates
- Dealing with higher radiation levels

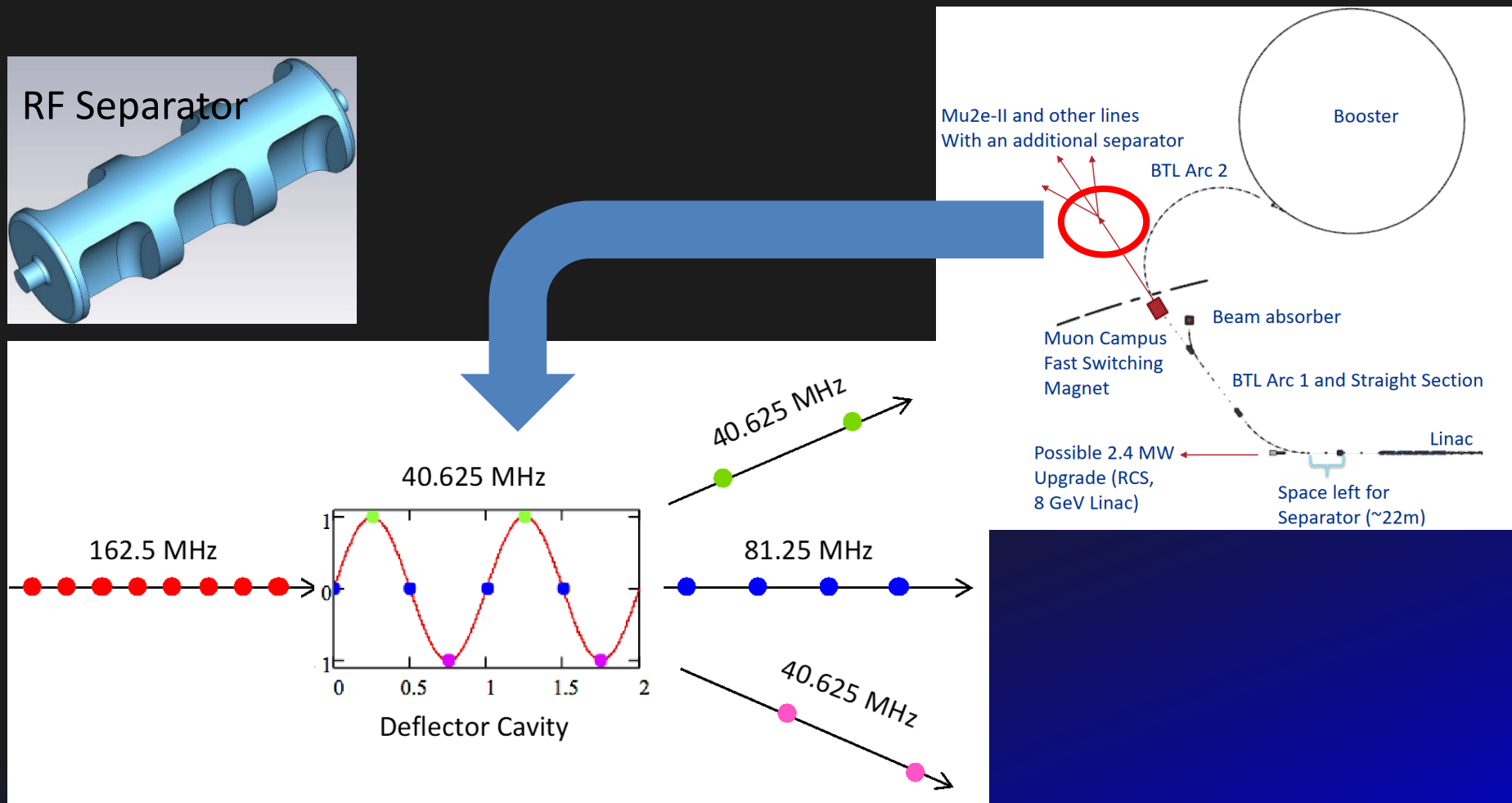
# PIP-II: Getting the Proton Beam to Mu2e

- Fast dipole magnet (20  $\mu$ s) switches beam between Booster and Mu2e-II + other potential experiments/beamlines
- Not in PIP-II baseline: an ongoing dialog with PIP-II designers heading off potential show-stoppers



# Mu2e-II: Proton Beam Splitting

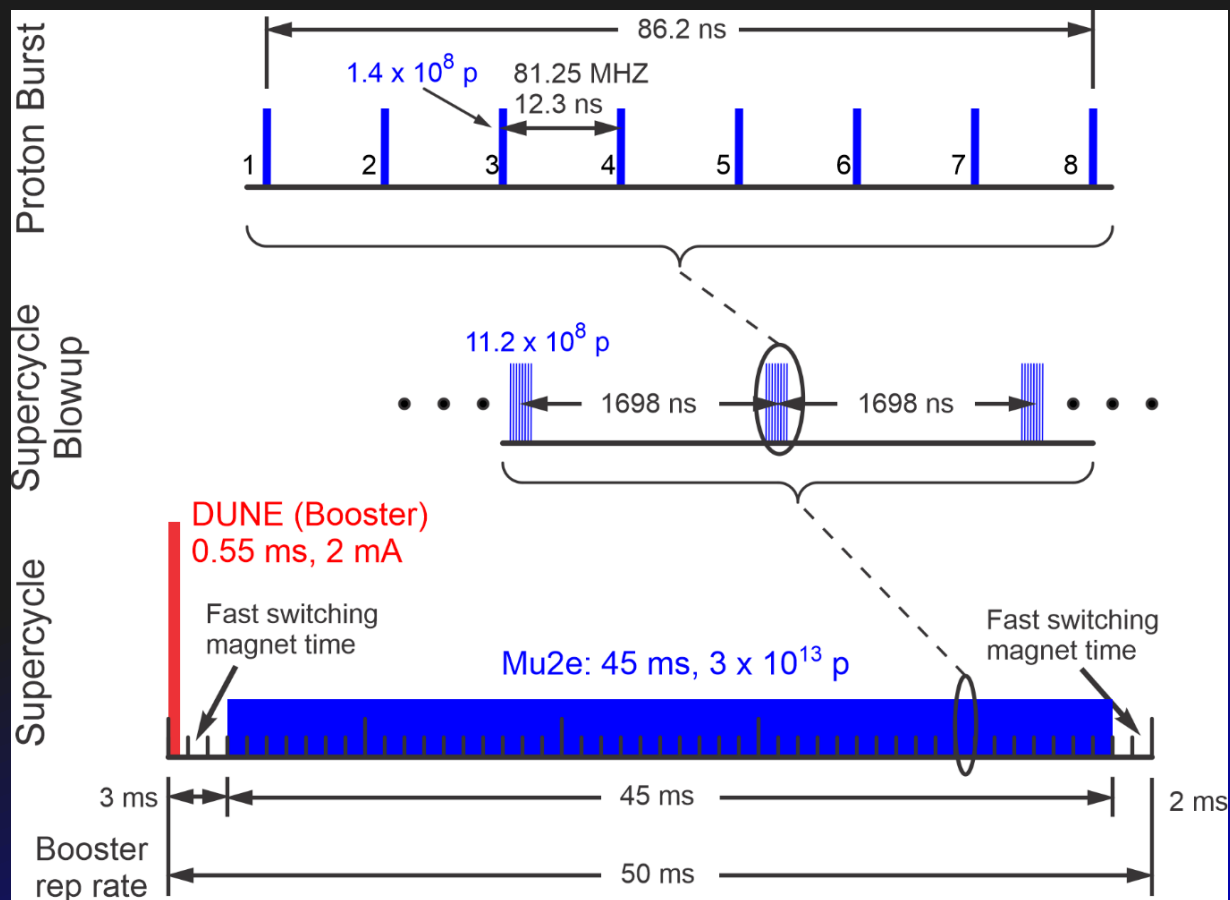
- Mu2e would take  $\frac{1}{2}$  of full bunch rate: 162.5 MHz
- Two other beam lines could be selected with an RF beam separator



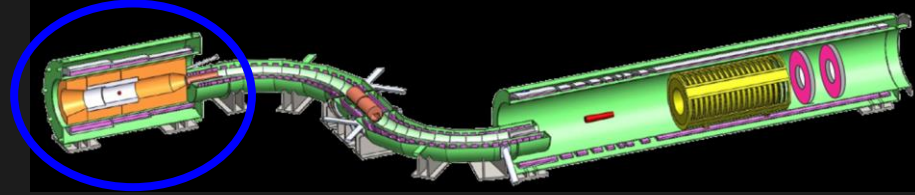
# Mu2e-II: Beam Structure

- PIP-II allows programmable proton pulse patterns
- Need a burst separation of  $\sim 1700$  ns (with Al target)
- Burst: 8 pulses ( $1.4 \times 10^8$  p each), each separated by 12.3 ns (81.25 MHz)
- Burst train: Burst followed by gap of 1698 ns, then another burst of 8 pulses, this repeated for  $\sim 45$  ms of beam every 50 ms
- 90% duty factor

76 kW average power  
(compared to 7.3 kW  
for Mu2e)



# Production Solenoid: Challenges



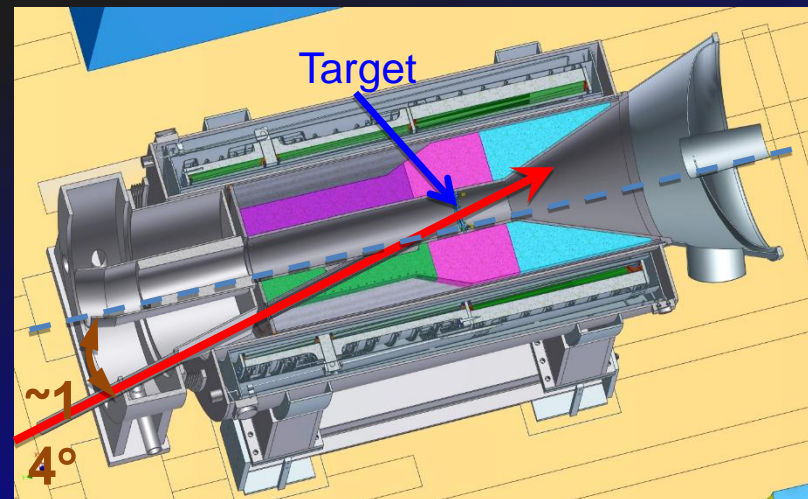
## Radiation and heat load:

- Some coils will have been subjected to  $\sim 7$  MGy and become activated
- Insulation damage (conventional epoxy limit  $\sim 10$  MGy)
- Degradation of Al stabilizer (RRR)
- Large heat load: power density increases by 10X
  - present magnet already pushed to limit
  - $\Delta T$  in coil goes from 0.25 K to 2.5 K: quench temperature is  $6.6 \text{ K} - 2.5 \text{ K} = 4.1 \text{ K}$ . With a thermal margin of 1.5 K the magnet temperature of 2.6 K is close to the lambda point (2.17 K)

## Beam transport with lower momentum beam

- 0.800 GeV rather than 8 GeV
- How do we steer it onto the target, dump, and extinction monitor?

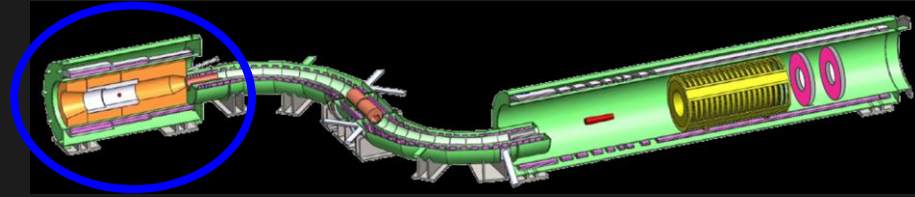
## Finding a target to handle 76 kW beam power



# Production Solenoid: Solutions

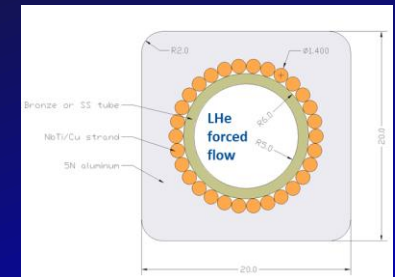
## Use the present Production Solenoid

- Upgrade the cryo-system
- Replace bronze Heat/Radiation shield (HRS) with W to reduce power density by 2.5X → note, may not be viable
- Will have to operate it at a lower temperature and/or with lower margin



## Replace or rebuild much of the Production Solenoid

- Some parts – vacuum vessel, thermo-shield, cold-mass supports reused
- New cable and coils would have to be made:
  - Cable-in-Conduit Conductor
    - Direct cooling increases heat load capability
    - Technically challenging, but being used (ITER)
  - Internally-cooled Al-stabilized cable
  - Non superconducting magnet:
    - Room temp resistive coil (replace HRS): ~ 5 MW
    - Cyro-cooled resistive coil (replace HRS): ~ 1 MW

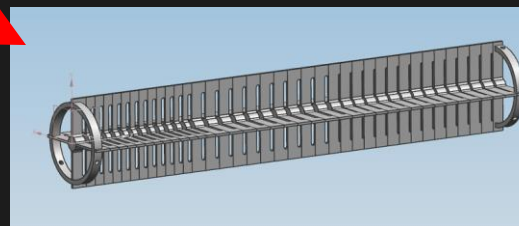
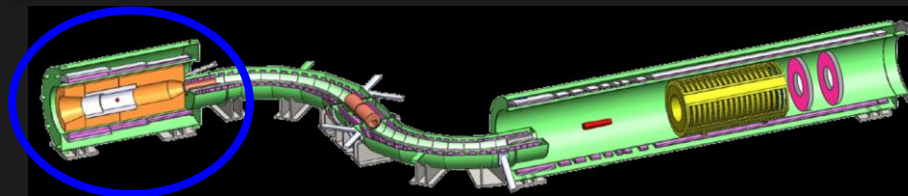




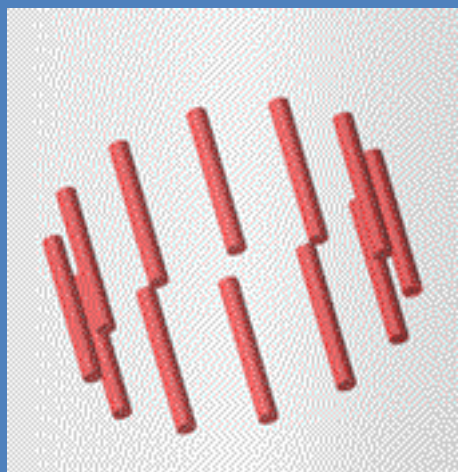
# Production Solenoid: Target Solutions

- Mu2e: ~1 kW in passively cooled W target
- Mu2e-II: ~15 kW in target
  - DPA  $\gg$  1

New target design needed

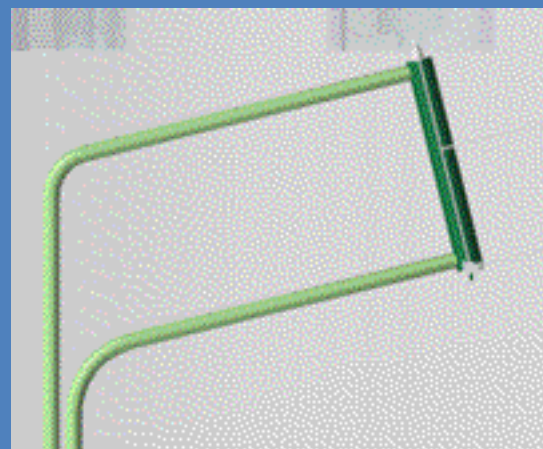


### Rotating Elements



- Pro:** survives radiation
- Con:** large profile

### Fixed Granular w Gas Cooling



- Pro:** small profile
- Con:** peak DPA  $>$  300/yr

### Conveyor Tube w Balls



Favored Target

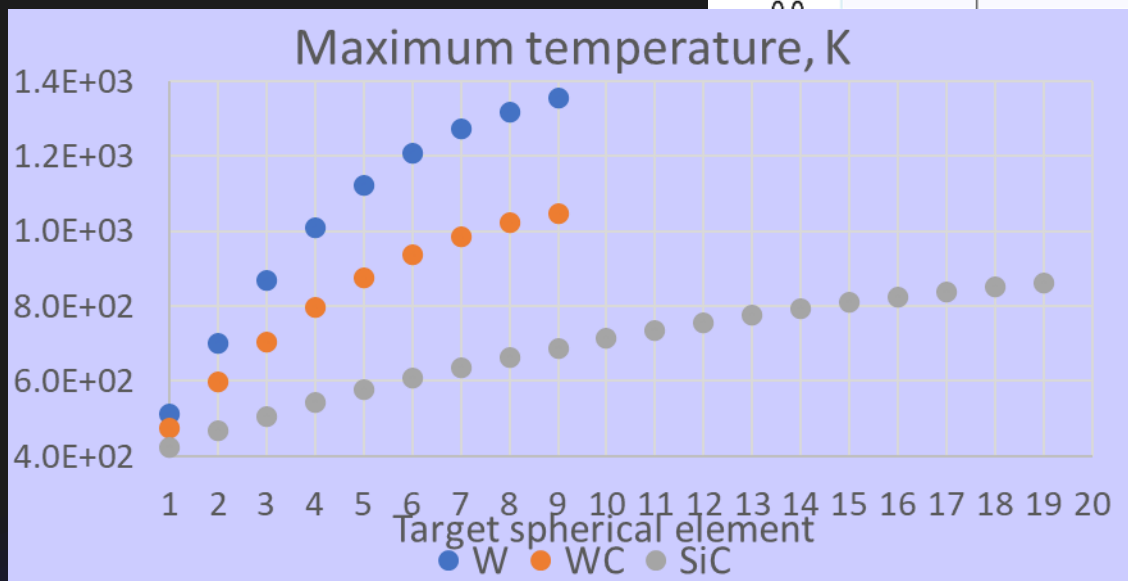
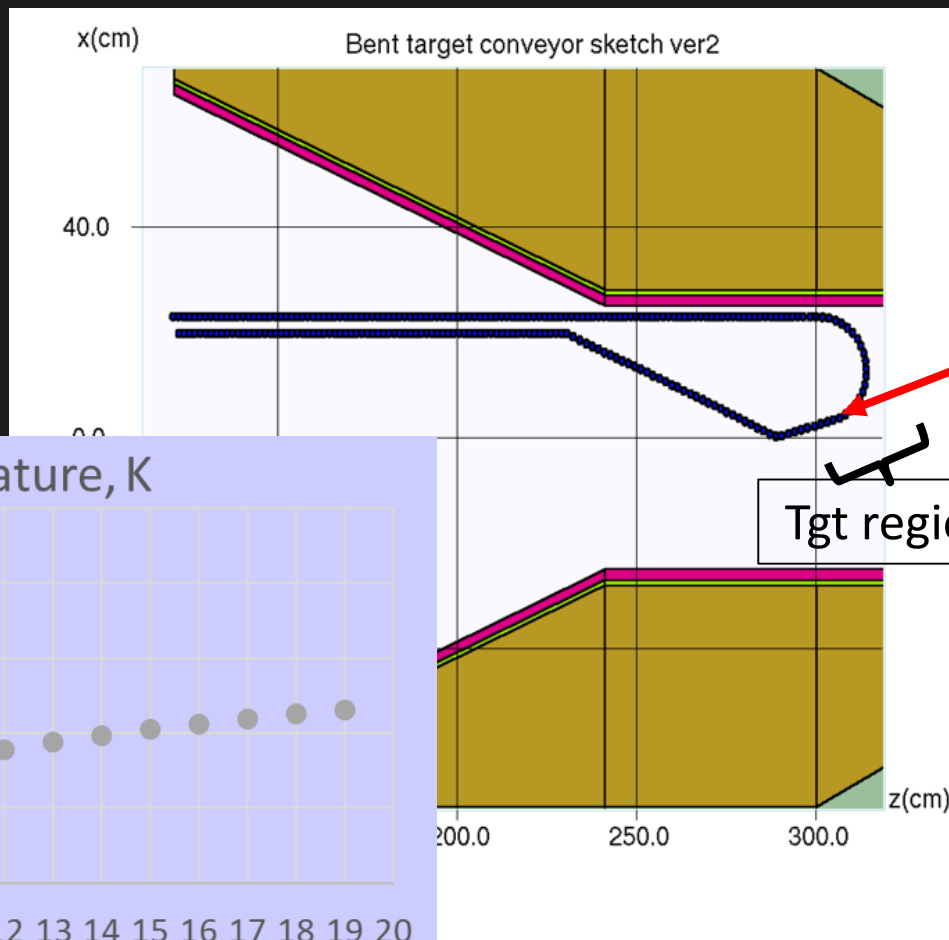
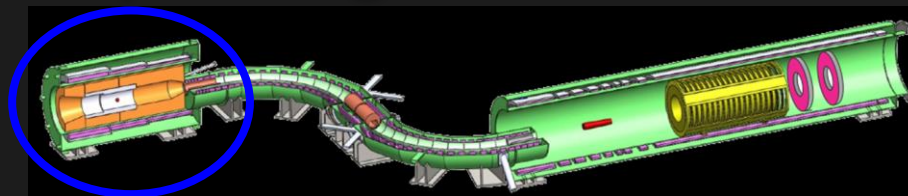
- Pro:** modest profile
- Con:** technically challenging

# Production Solenoid: Conveyor Target

Exploring different target materials

- W/WC balls: 9
- SiC balls: 19
- C balls: 28

Max temperatures well below melting points

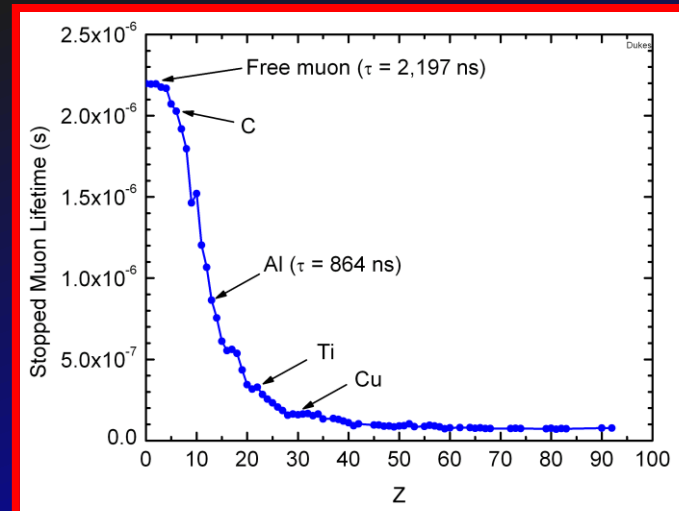
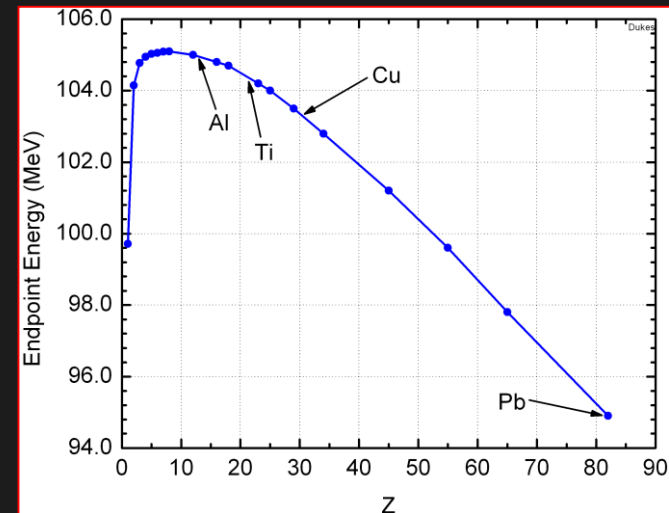
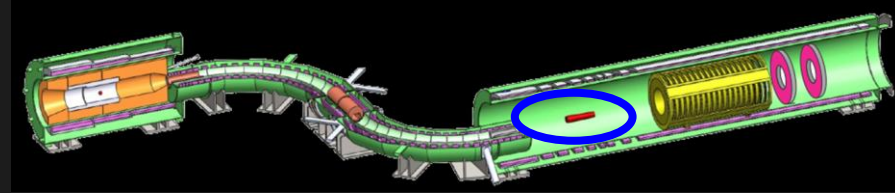


# Mu2e-II: Stopping Target

## Al is ideal first target

- High endpoint energy
  - muons captured on other (higher Z) detector material not a background
- Capture daughter more massive  $m_{Z-1} > m_Z$ 
  - keeps max. energy of radiative capture muons below signal electrons
- Long lifetime
  - keeps proton blast separated from live window

Would continue with this target if Mu2e sees nothing



# Mu2e-II: Other Stopping Target Choices

If Mu2e observes conversion, a different target would be ideal in order to narrow down the physics process

The opportunities and challenges of the following targets are being explored:

Lithium

- Low discrimination

Sulphur

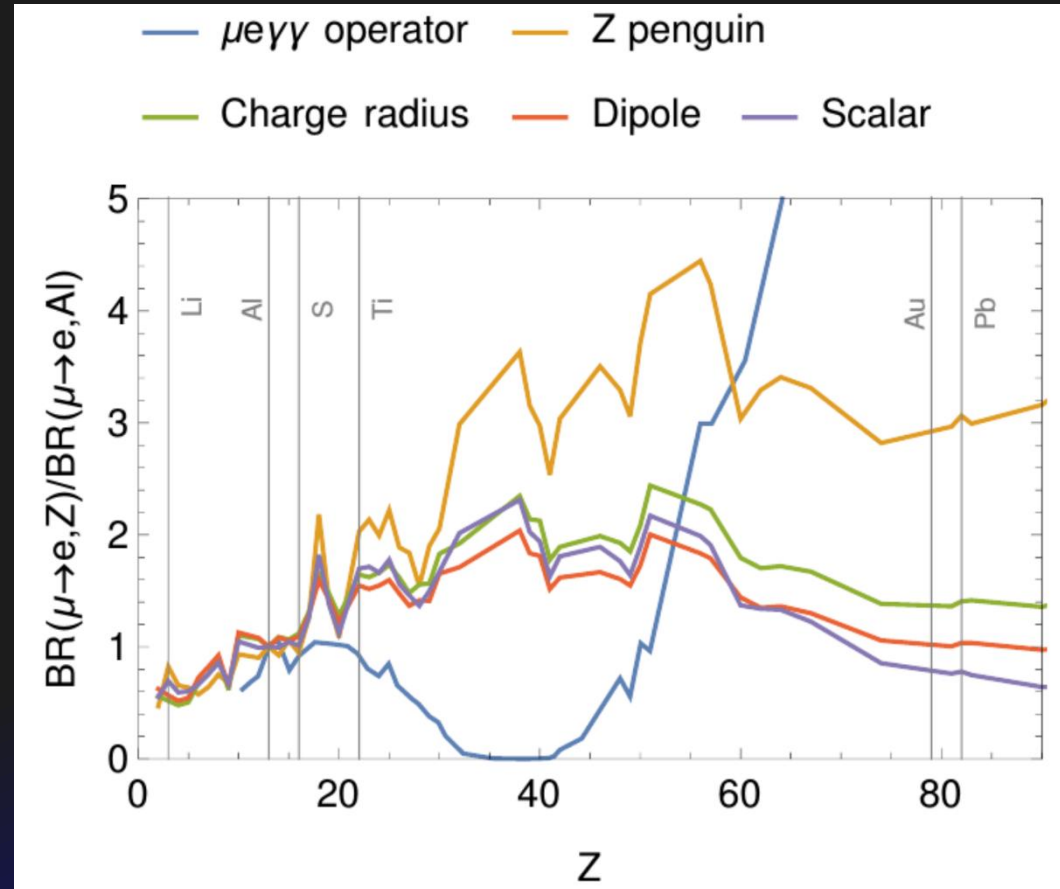
- Advantages for  $e^+$  channel

Titanium

- Multiple isotopes

Au/Pb

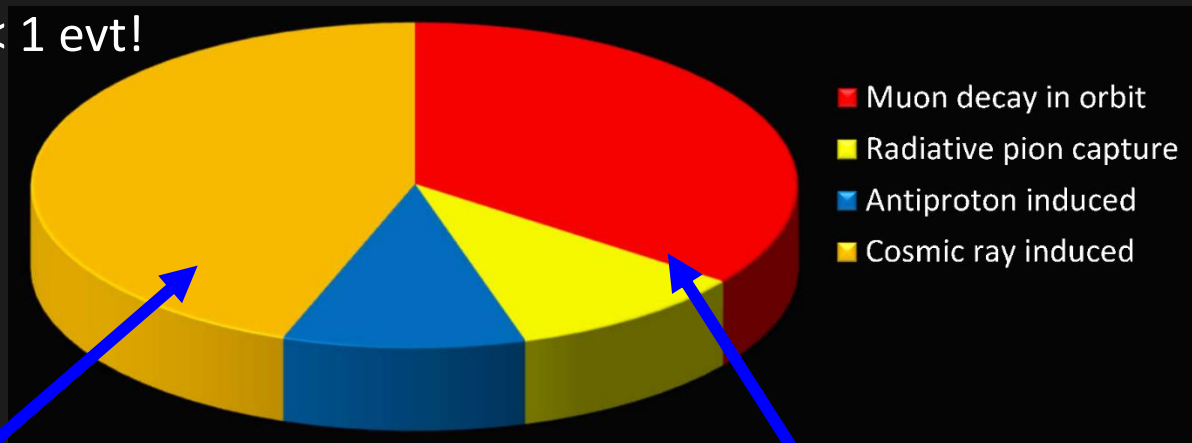
- Good discrimination
- Short lifetime  $\rightarrow$  low rate



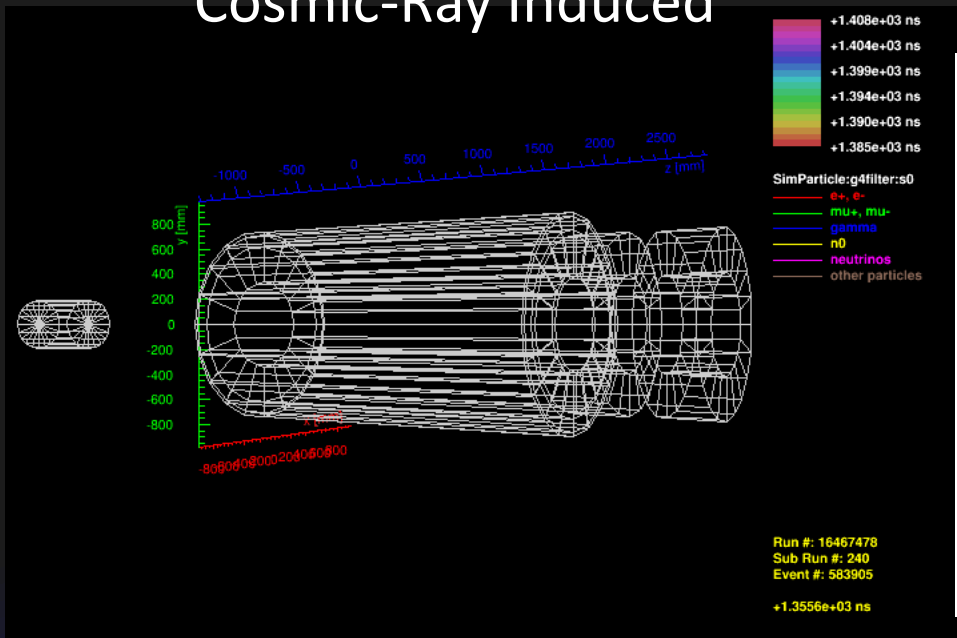
# Keeping Backgrounds in Check

Need to keep backgrounds to  $< 1$  evt!

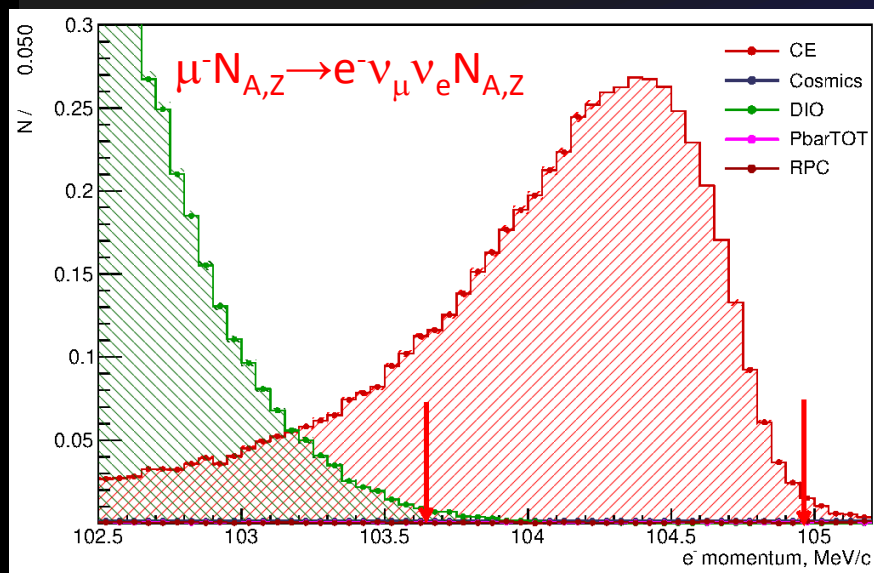
Two biggest backgrounds are Muon decay-in-orbit and cosmic-ray induced electrons



## Cosmic-Ray Induced



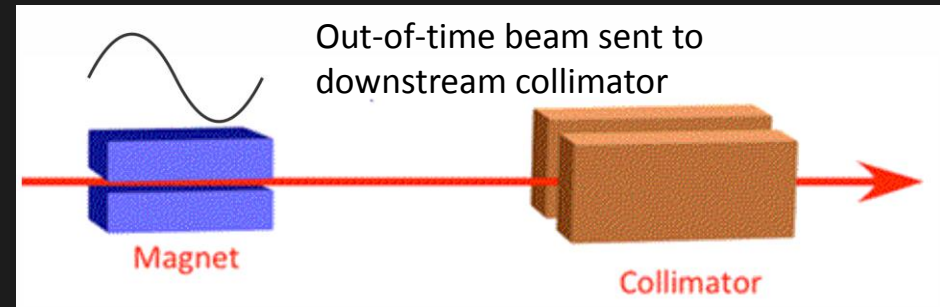
## Muon Decay in Orbit





# Extinction

Needed to remove out of time protons that can produce conversion-like electrons from captured pions and from electrons that scatter in the stopping target



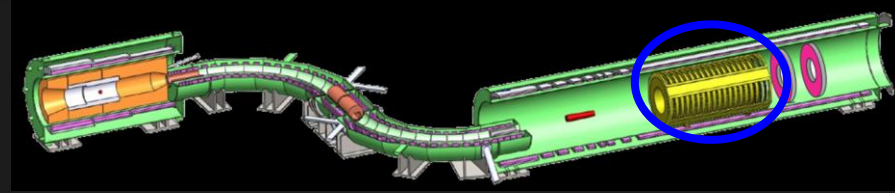
- Mu2e:
  - Need  $10^{-10}$  reduction (out of bunch/in bunch)
  - $10^{-5}$  from bunch formation in Recycler/Delivery Ring
  - $10^{-7}$  from AC Dipole extinction system
  - 100X safety factor
- Mu2e-II:
  - Need  $10^{-11}$  reduction
  - $10^{-4}$  (at least) from chopper
  - Need additional  $10^{-9}$  from extinction system to get same safety factor
  - Should be easy: lower energy, smaller emittance beam, etc

**This looks very feasible**

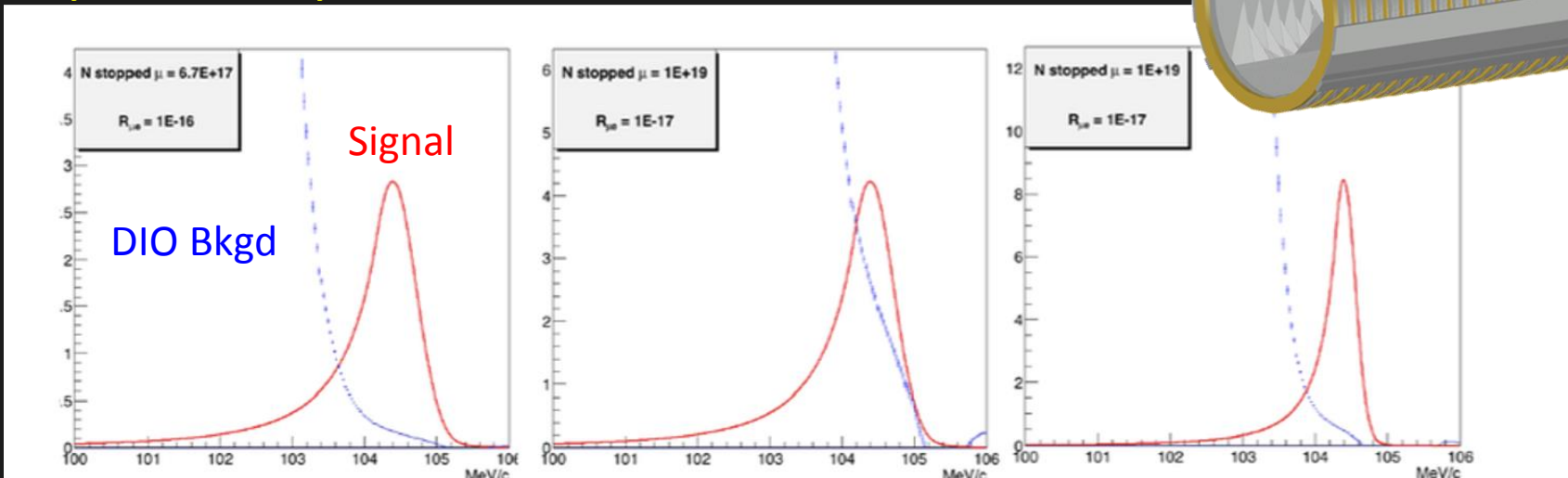


# Tracker: Would the Mu2e-II Tracker Work?

- 21,000: 5 mm diam straws Ar:CO<sub>2</sub>
- 15  $\mu\text{m}$  Mylar thick
- metalized inside/outside



## Toy MC Study:



Mu2e Rate and Background

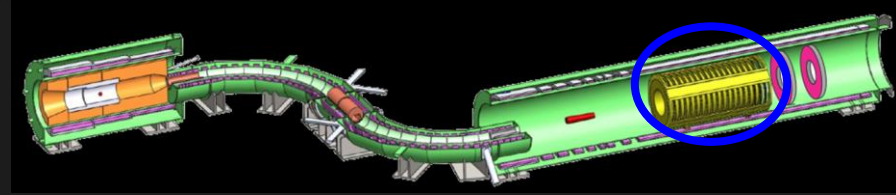
Mu2e-II Rate and Background  
Mu2e Tracker: 15  $\mu\text{m}$  straws

Mu2e-II Rate and Background  
Tracker w  $\frac{1}{2}$  resolution

**No! It must be replaced as background would exceed 1 event**

- DIO background moves to the right (into signal region)
- Signal region needs to be narrowed, DIO moved to left, by reducing material

# Tracker: Challenges & Solutions



## Challenge: increased radiation load

- Radiation-hard front-end electronics: ASICs, DC-DC converters, optical readout

## Challenge: ~4X Increase in bunch intensity

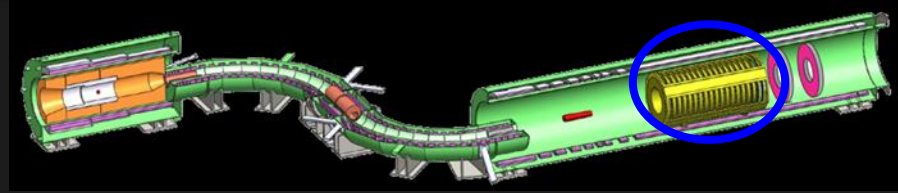
- Only produces a 5% reduction in momentum resolution and reconstruction efficiency
- Current design and software is capable of this

## Challenge: Lower mass to meet momentum resolution goal

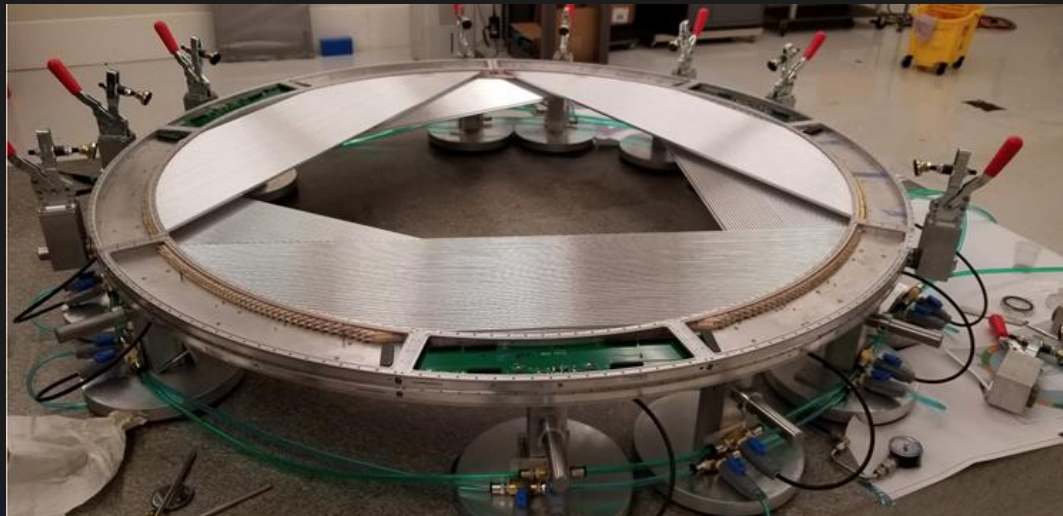
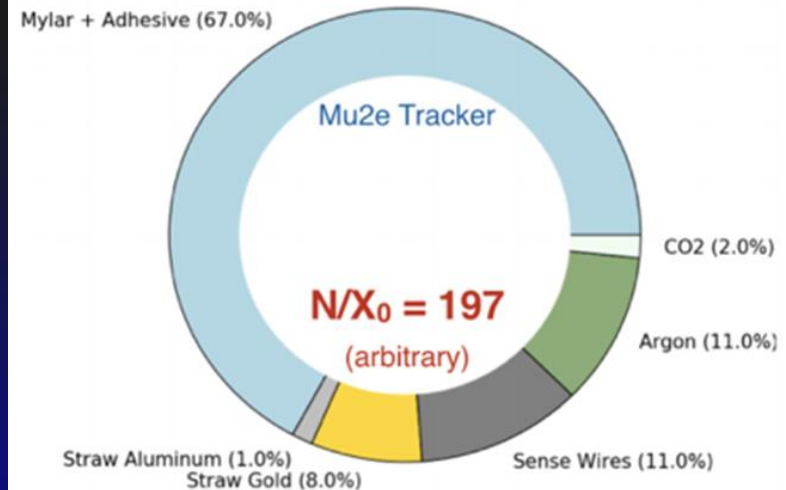
- Solutions:
  - Thinner straws
  - Lower mass gas/sense wires
  - Completely different technology

# Tracker: Straw Challenges

- Straw thickness: 15  $\mu\text{m}$  must be reduced to at least  $\sim 8 \mu\text{m}$
- Fabricating issues: butt vs overlap seams, winding schemes, etc
- Sustaining 1 atm pressure difference a challenge (significantly higher Hoop Stress)
- Increased leak rate (Mu2e has a  $\sim 15\%$  straw failure rate)
- Mechanical properties: less strength to keep straw straight
- Aging: large additional charge of  $\sim 10 \text{ C/cm}$  (Mu2e:  $\sim 1 \text{ C/cm}$ )

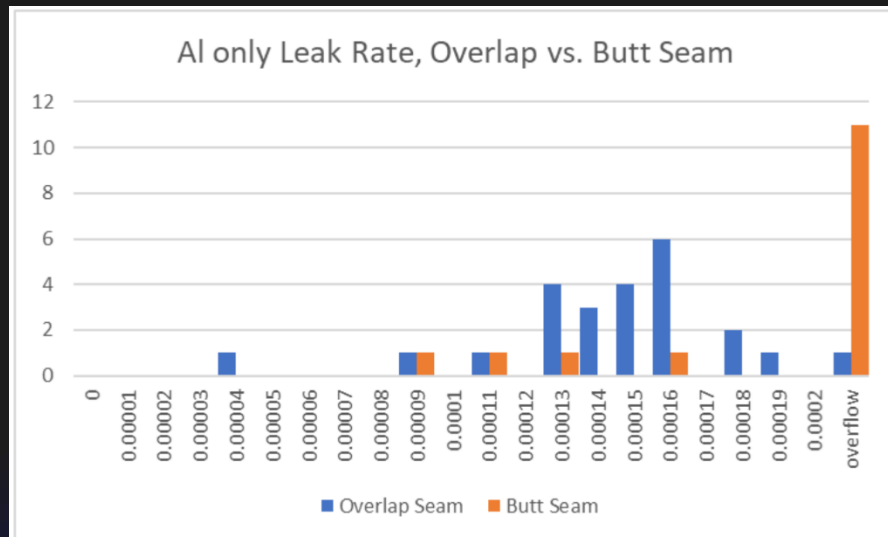


## Material Budget for Mu2e Straws



# Tracker: Straw Solutions

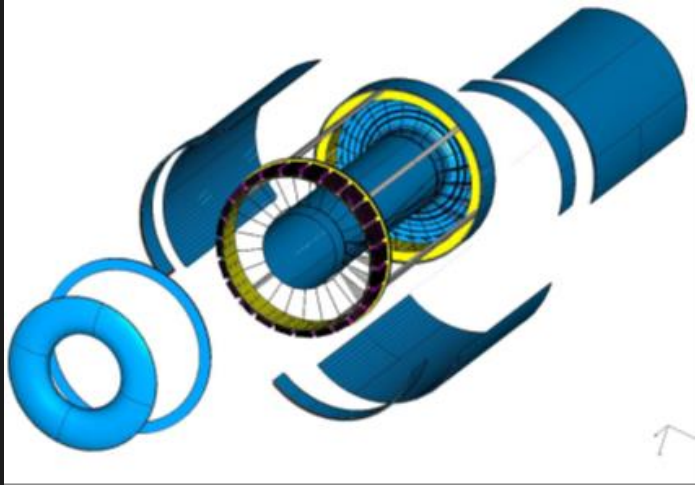
- Fermilab has LDRD team (Brendan Casey et al) researching thinner straws
- Fabricating straws with 3-8  $\mu\text{m}$  wall thickness
- Gases:
  - Can the pressure be less than 1 atm?
    - Reduces leak rate ( $\text{CO}_2$ ) through straws
    - Reduces Hoop Stress
    - Slight reduction in mass
- Note: thinner walls reduce charge load (largely caused by photon conversions in walls)



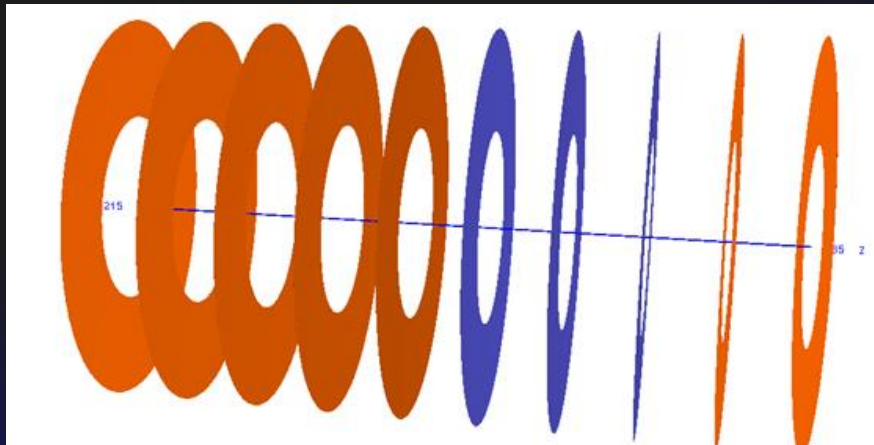


# Tracker Solutions: Straw Alternatives

Drift chamber similar to MEG-II

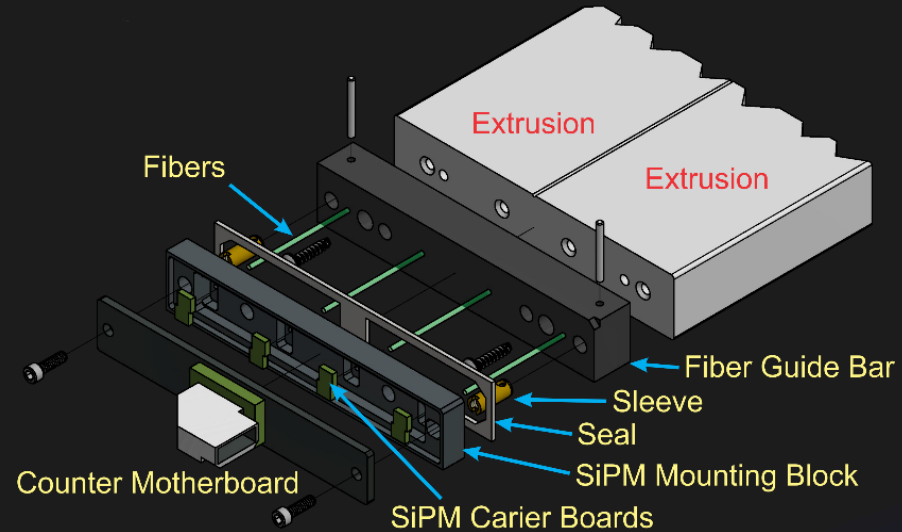


Other alternatives: Light Si, Micro Pattern Gas Detector, radial TPC



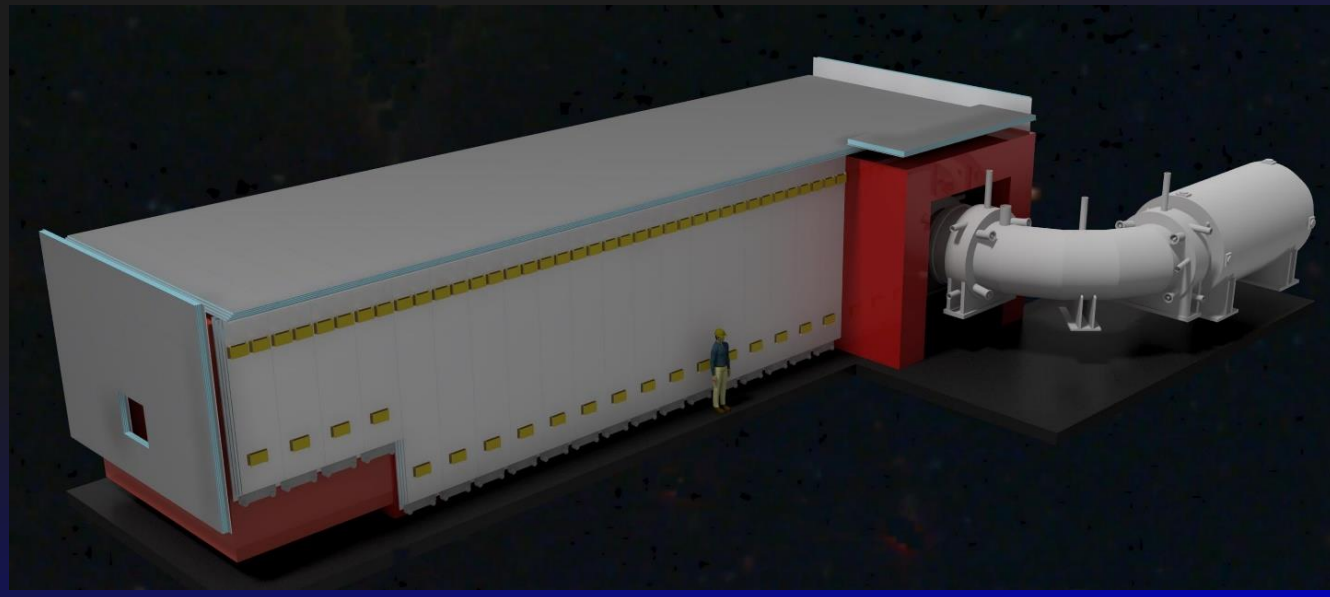
# Cosmic Ray Veto (CRV): Mu2e

- About 1 fake event/day from cosmic-ray muons
- Hence need  $\sim 99.99\%$  efficiency in an intense radiation environment
- Surround Detector Solenoid by 4 layers of scintillator read out by waveshifting fibers, and silicon photomultipliers



## Details:

- 5,344 counters
- Area: 335 m<sup>2</sup>
- 10,688 fibers
- 19,392 SiPMs





# Cosmic Ray Veto: Challenges



**Live time: ~3X higher (due to larger duty factor)**

- Directly increases cosmic-ray induced background by same amount → Cosmic-ray induced background scales as live time, not beam intensity!

**Light yield degradation: expected scintillator light yield will be significantly less**

- Scintillator aging directly impacts efficiency in detecting cosmic-ray muons

With present CRV expect ~4 background events in a 3-year run of  $4.56E7$  s beginning in 2030 due to the live time increase and effects of light yield decline

**Noise rates: expect ~3X higher noise rates from neutrons, gammas, etc. coming from the production target, stopping target, collimators, and muon beam stop**

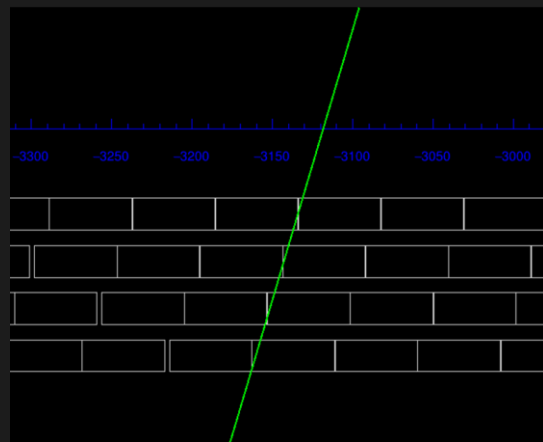
- Rates in some sectors of the CRV are already at the limit of what the electronics can handle
- Radiation damage to photodetectors and front-end electronics close to becoming an issue
  - Mu2e max non-ionizing:  $1 \times 10^{10}$  n/cm<sup>2</sup> for Front-end electronics,  $1 \times 10^{11}$  n/cm<sup>2</sup> for SiPMs

# CRV: Improving Muon Veto Efficiency

Overall efficiency needs to be improved by  $\sim 3X$  do to longer lifetime Note: 20% of CRV vetoes 80% of background-creating muons



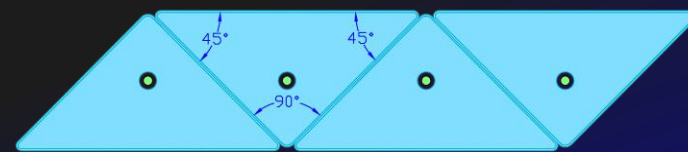
Most non-vetoed CR muons come at nearly vertical angles, and traverse gaps between counters



Replace rectangular counter with triangular design to avoid vertical gaps



Mu2e "di-counter"



Mu2e-II "quad-counter"

- Increase light yield: New counters reset light-yield decline; use new, higher efficiency SiPMs; pot fibers in their channels
- Other solutions: employ high-rate gas detectors in critical sectors

# Cosmic Ray Veto: Handling 3X Muon Intensity

## Problems:

- Front-end Electronics throughput at its maximum limit
- Radiation damage becoming an issue, particularly for SiPMs
- Deadtime, presently about 5-10%, will become uncomfortably large
- Larger duty factor reduces 'off-spill' time used for data transmission, calibration, etc

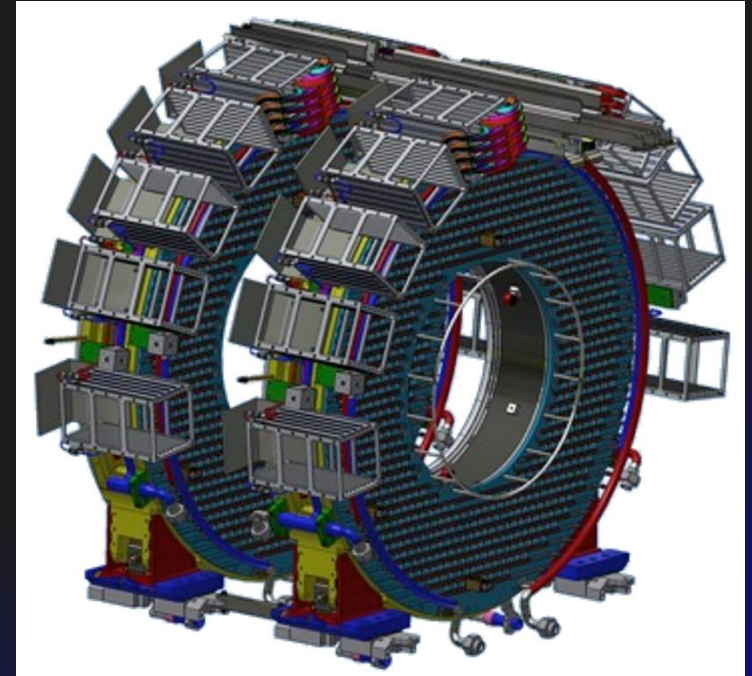
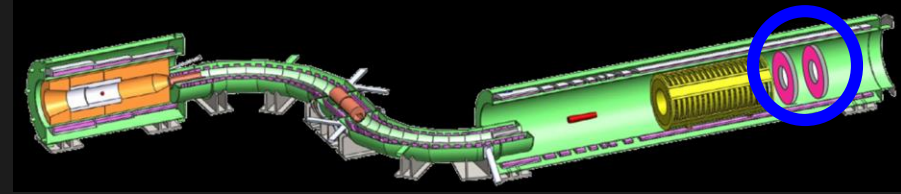
## Solutions:

- Note: only about 10% of the CRV is adversely affected: increased rates not a problem for the remainder of the CRV
- Better shielding: Boron and barite loaded concrete
- Smaller counters: Triangular counters
- Non-scintillator designs: employ high-rate gas detectors in critical sectors

# Calorimeter

Needed for:

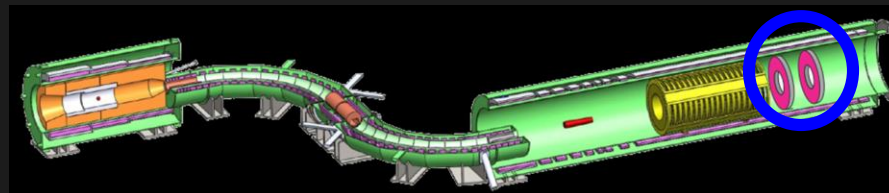
- Seeding track finder
- Particle ID: e vs  $\mu$
- Trigger
- Two disks separated by  $\frac{1}{2} \lambda$  of helix
- 2 x 674 un-doped CsI crystals: 34x34x200 mm<sup>3</sup> ( $10X_0$ )
- Dual UV-extended SiPM readout
- Conversion electron resolution:  $\sigma_E/E \sim O(5\%)$
- Timing:  $\sigma_t < 0.50$  ns



# Calorimeter: Problems and Solutions

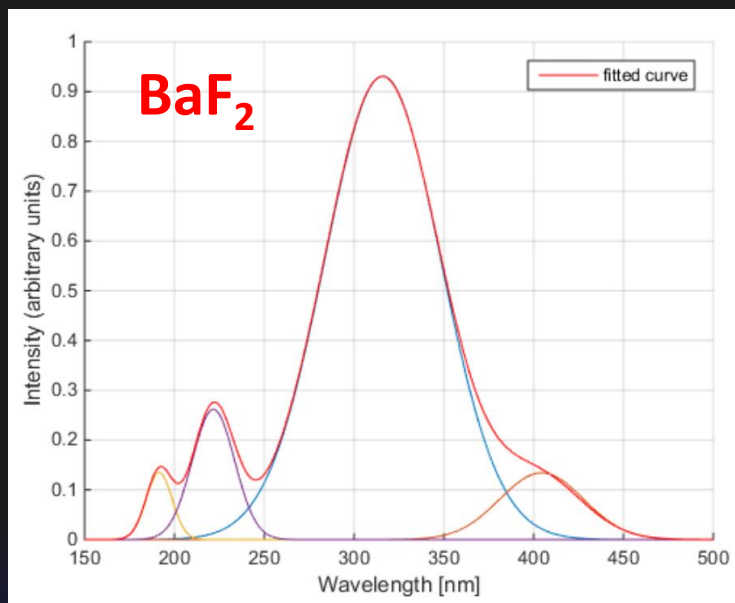
## Problems:

- Not sufficiently rad hard for Mu2e-II:  $\sim 10$  kGy/yr IR;  $\sim 10^{13}$  n/cm<sup>2</sup>
- Not fast enough: 30 ns



## Solutions:

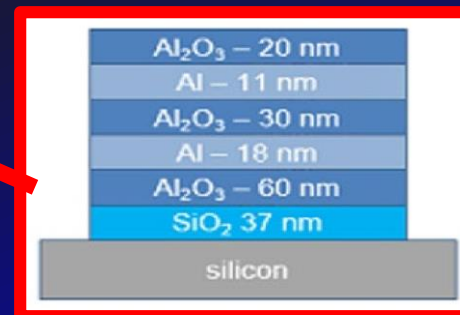
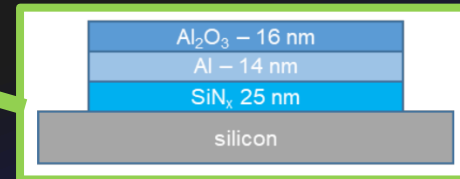
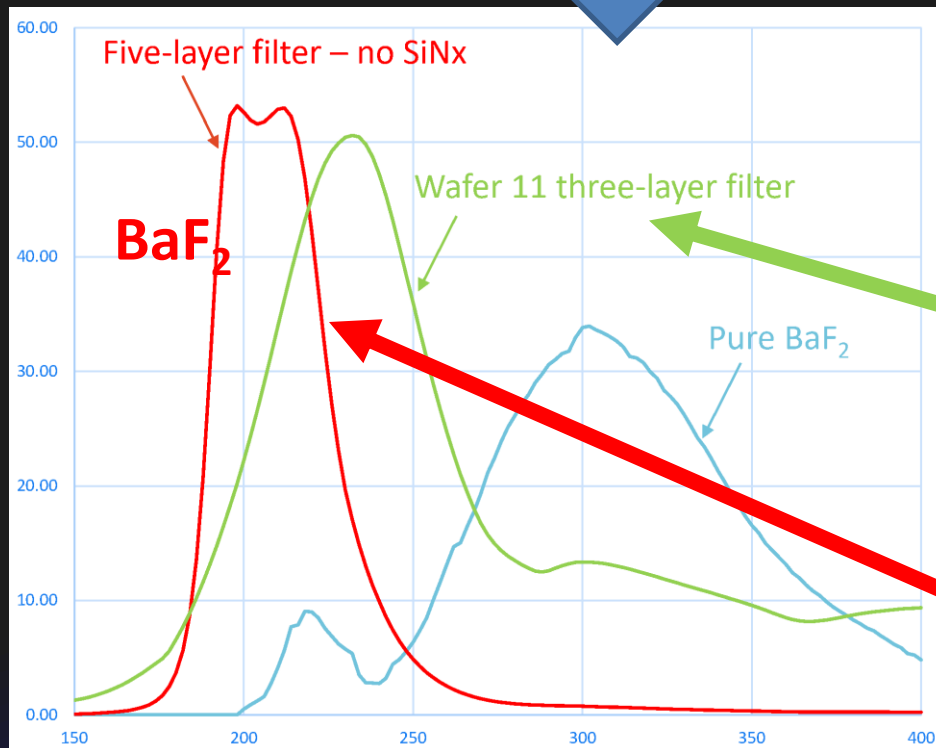
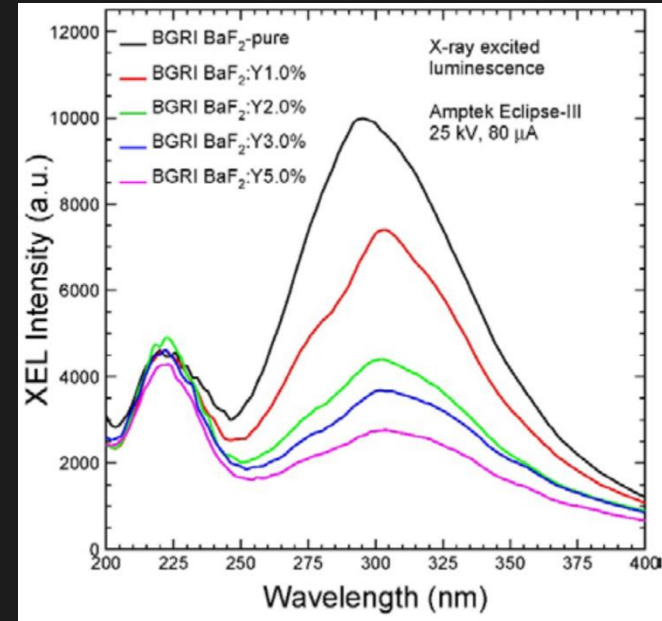
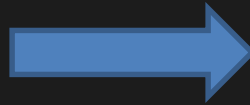
- Fast, rad hard crystals:
  - BaF<sub>2</sub> with suppressed slow component readout
  - BaF<sub>2</sub> doped with Y to suppress slow component



Crystal	CsI	BaF <sub>2</sub>	BaF <sub>2</sub> (Y)
Density (g/cm <sup>3</sup> )	4.51	4.89	4.89
Hygroscopicity	Slight	None	None
$\lambda_{\text{peak}}$ (nm)	420 310	300 220	300 220
Light Yield (% NaI(Tl))	3.6 1.1	42 4.8	1.7 4.8
Decay Time (ns)	30 6	600 0.5	600 0.5

# Calorimeter: Fast Readout of BaF<sub>2</sub>

- Caltech working on Y-doped BaF<sub>2</sub>
- Caltech, JPL, FBK working on UV-only sensitive SiPM





# Summary

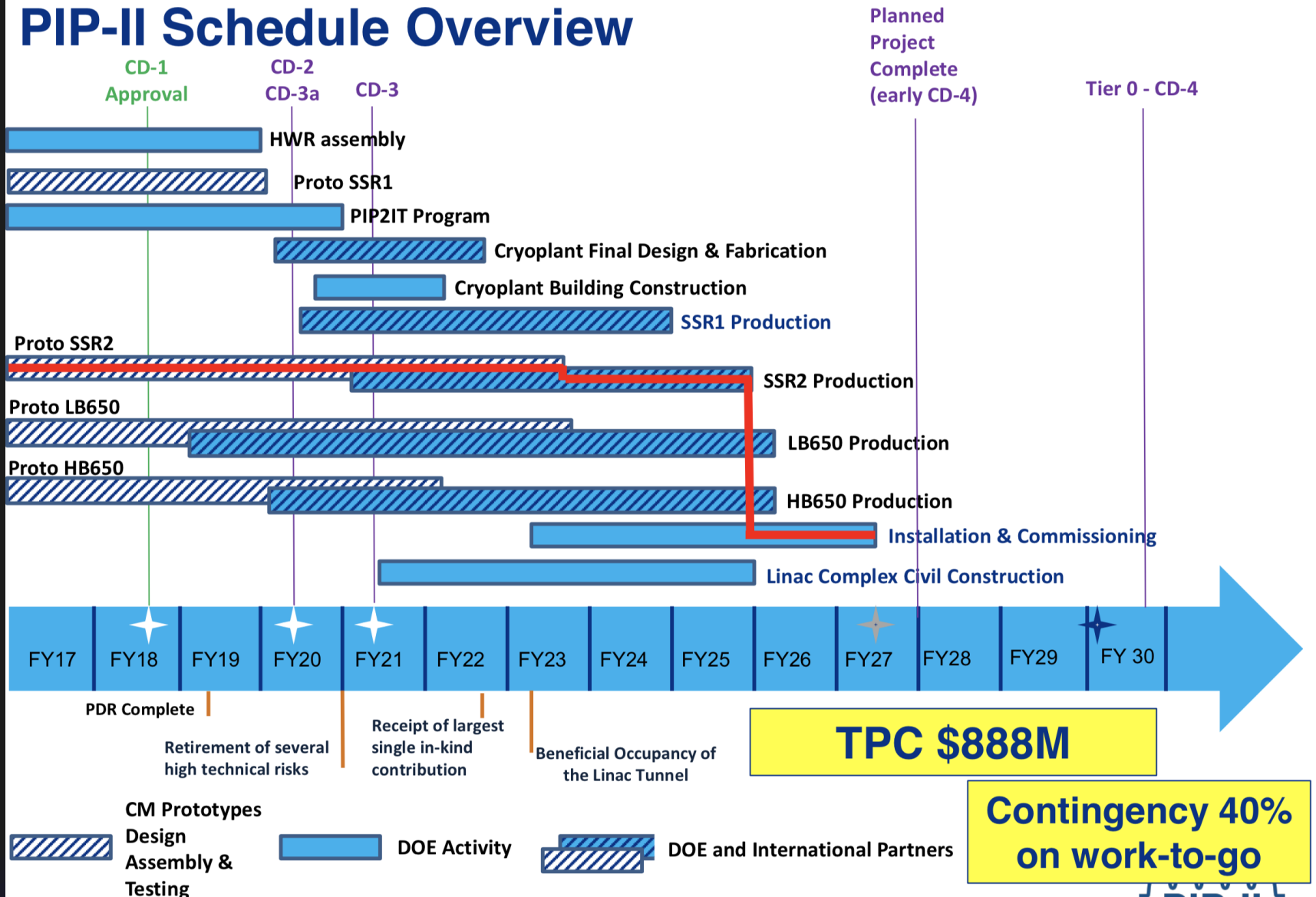
- Muon-to-electron conversion provides one of the most sensitive probes of new physics
- PIP-II has the potential to allow an improvement of 10X over the expected Mu2e sensitivity through a more intense muon beam and better duty factor
- The PIP-II beam provides several other advantages over the present booster beam
- However, there are significant challenges:
  - Handling the lower energy proton beam, 0.800 GeV vs 8 GeV
  - Handling the higher rates
- Significant parts of the apparatus will have to be replaced
- **Opportunities of achieving a SES  $O(10^{-18})$  far outweigh the challenges**
- A small, enthusiastic team is working on Mu2e-II, the immediate goal is to produce a strong conceptual design for the US Snowmass-2022 process, which is mapping out a USA plan for the next generation of experiments and facilities
- If you are interested in joining our team, there are many opportunities

**Thanks to my many colleagues on Mu2e-II and Mu2e!**

# Backup Slides

# PIP-II Schedule

## PIP-II Schedule Overview



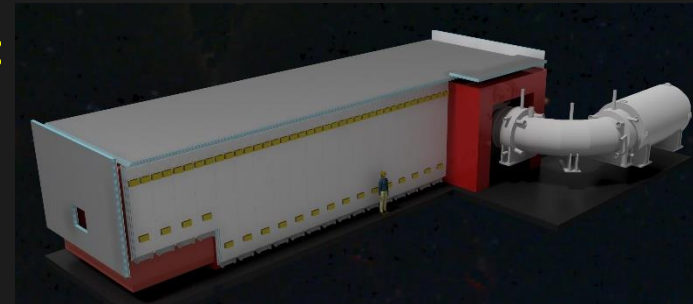
# General requirements for Mu2e PS

- Magnetic:
  - Nominal peak field on the axis 4.6 T;
  - Maximum peak field on axis 5.0 T;
  - Axial gradient -1 T/m;
  - Gradient uniformity  $\pm 5$  %.
- Electrical:
  - Operating margins:  $\geq 30$  % in  $I_c$ ,  $\geq 1.5$  K in  $T_c$ ;
  - Operating current 9÷10 kA;
  - Peak quench temperature  $\leq 130$  K;
  - Voltage across terminals  $\leq 600$  V.
- Structural:
  - Withstand forces at all conditions while part of the system or stand-alone;
  - Cryostated magnet weight  $\leq 60$  tons;
  - Compliance with applicable structural codes.
- Cryogenic:
  - Cooling agent: LHe at 4.7 K;
  - Total heat flow to LHe  $\leq 100$  W;
  - Cryostat ID 1.5 m;
  - Conduction cooling.
- Radiation:
  - Absorbed dose  $\leq 7$  MGy total;
  - Minimum RRR of Al stabilizer in the operating cycle  $\geq 100$ .

# CRV: Reducing Neutron Induced Background

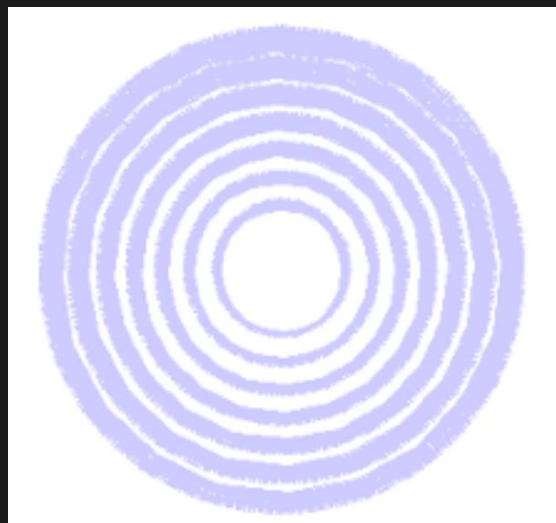
Estimate 0.007 background events per 1E6 seconds:  
0.175 events in a run of 2.5E7 live seconds

- Increase shielding above detector pit to reduce number of cosmic-ray neutrons
  - This is feasible and sufficient
- Find a way to place a veto around the stopping target region
  - Extremely challenging – high rates, low mass (including cable plant), operation in vacuum
  - Could replace much of the present Mu2e CRV



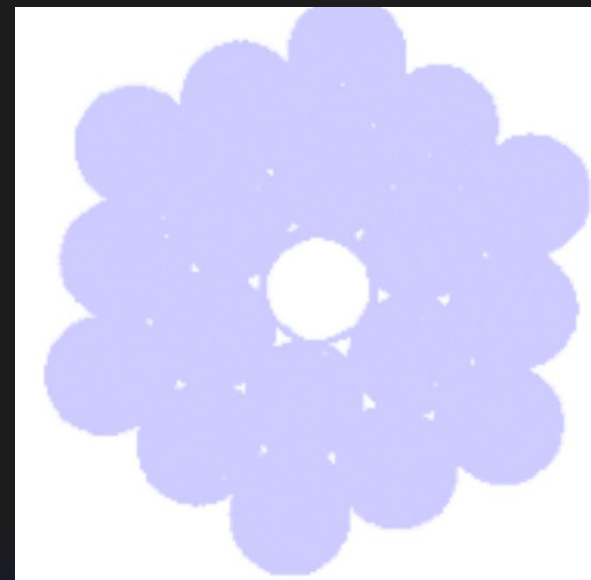
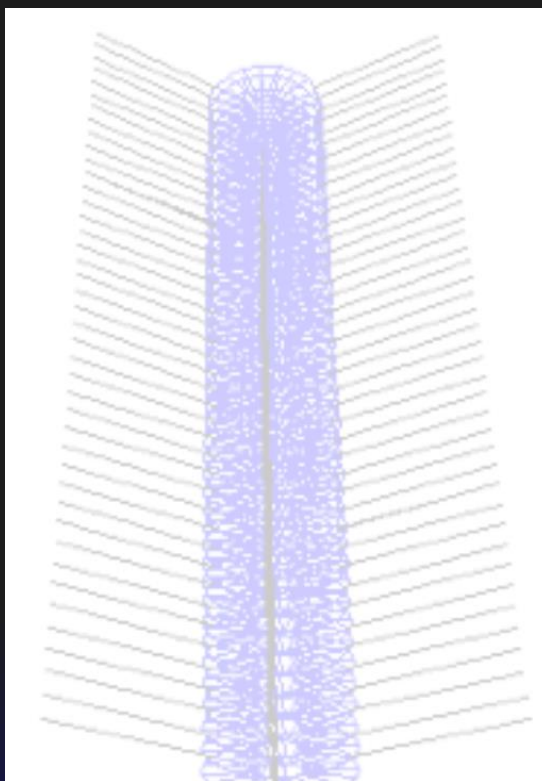
# Mu2e-II: Other Stopping Target Designs

Can we increase the stopping fraction; lower the electron escape mass?



Concentric Cylinders

Foils & Frame



Hexagonal Cylinders

No significant improvement in stopping fraction: Present design close to optimal