



Reactor antineutrino anomaly in light of recent flux model refinements

Z. Xin (IHEP, Beijing)

Sept. 2021

On the basis of the work with *C. Giunti, Y.F. Li, C.A. Ternes*

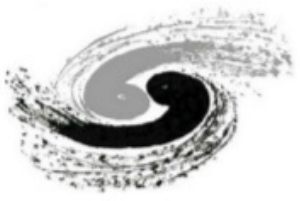
NuFact 2021: The 22nd International Workshop on Neutrinos from Accelerators

Online Event

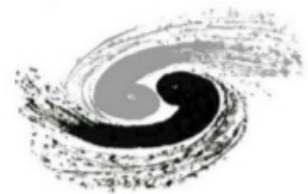
Outline



- **Motivation**
- **New calculation of IBD yield**
- **Method of analysis**
- **Fit of reactor rates**
- **Fit of reactor fuel evolution data**
- **Best-fit model**
- **Summary**

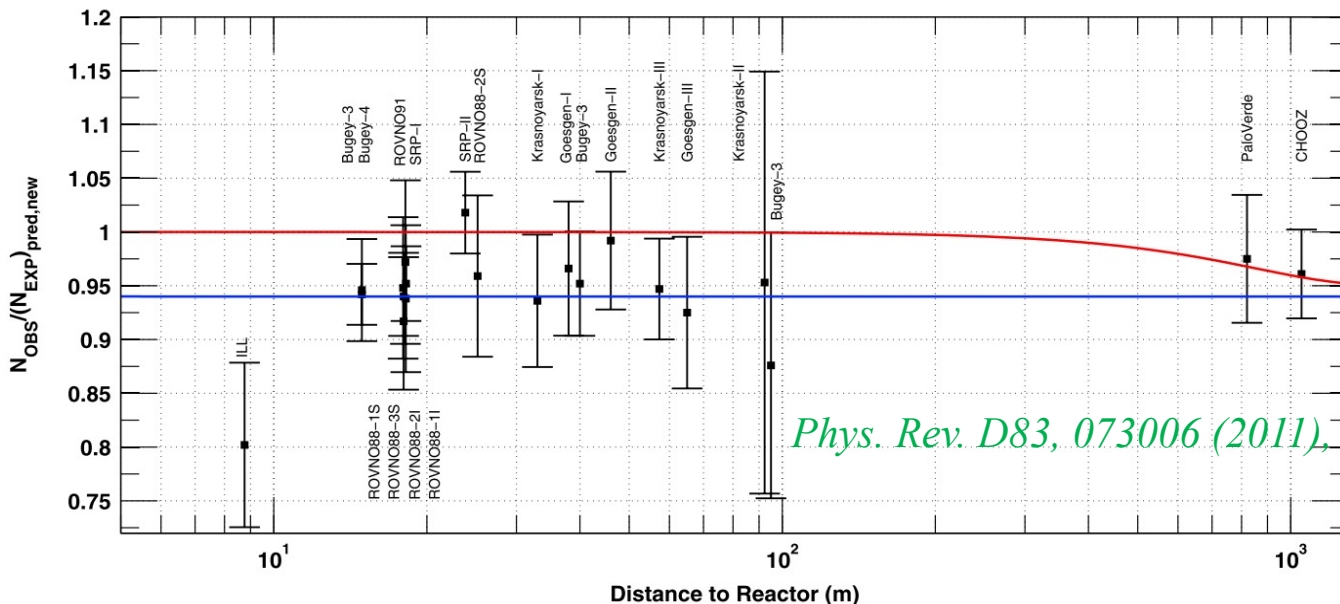


Part 1: Motivation



Motivation

- Reactor antineutrino anomaly



mean averaged ratio:
 $\bar{R} = 0.943 \pm 0.024$

- Reactor experiments data → test RAA for different models.

Reactor data

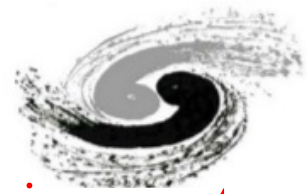
Models

- Huber-Mueller model *Phys. Rev. C 83, 054615 (2011)*
- Estienne-Fallot summation model *Phys. Rev. C 85, 029901 (2012)*
- Hayen-Kostensalo-Severijns-Suhonen model *Phys. Rev. C 100, no.5, 054323 (2019)*
- Recent Kurchatov Institute measurements *arXiv:2103.01684*
 - HM → KI model
 - HKSS → HKSS-KI model

- Reactor rates data (27)
 - 80s-90s, 2000s
 - Recent Prospect & STEREO
- Fuel evolution data (8+8)
 - Daya Bay
 - RENO



Part 2: New calculation of IBD yield



Reactor flux models

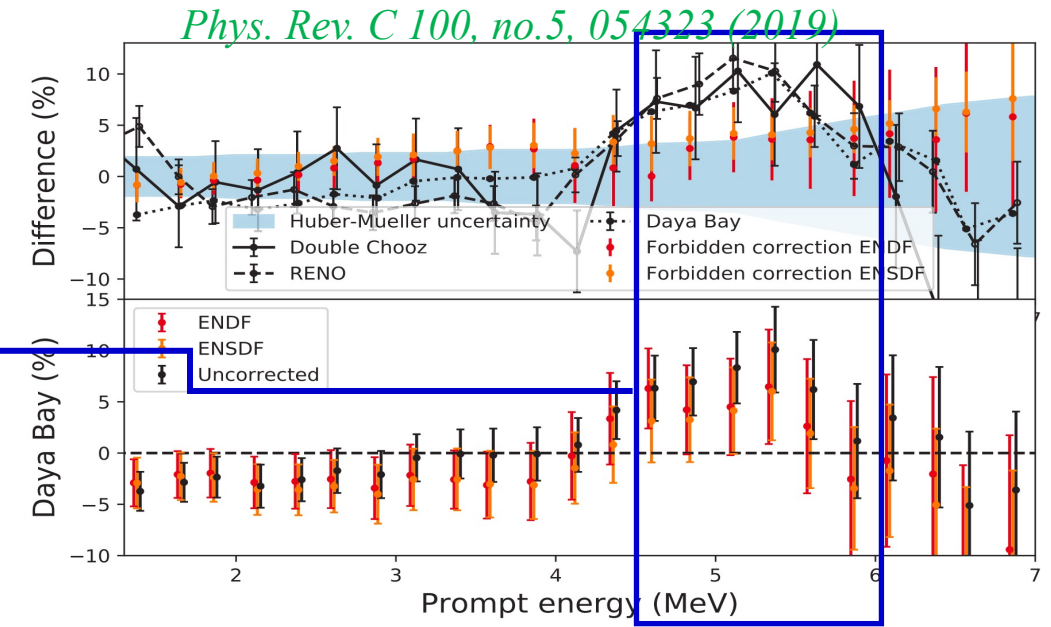
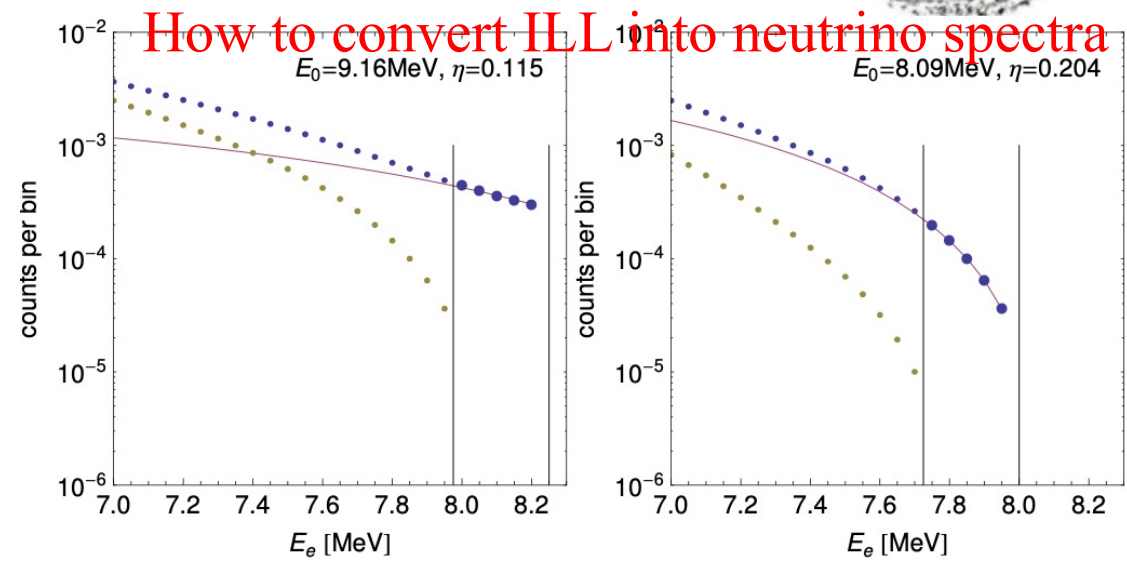
- Theoretical reactor antineutrino spectra
 - Conversion method
Measured β spectra \rightarrow neutrino spectra
 - Summation method
Sum all the decay branches [database](#)

- Huber-Mueller model** *Phys. Rev. C 85, 029901 (2012)*
 - ^{235}U , ^{239}Pu , ^{241}Pu : **ILL** β spectrum \rightarrow neutrino spectrum *Phys. Rev. C 83, 054615 (2011)*
 - ^{238}U : sum all β decay branches
 - Allowed approximation**

forbidden transition

Partially explain "5 MeV bump"!

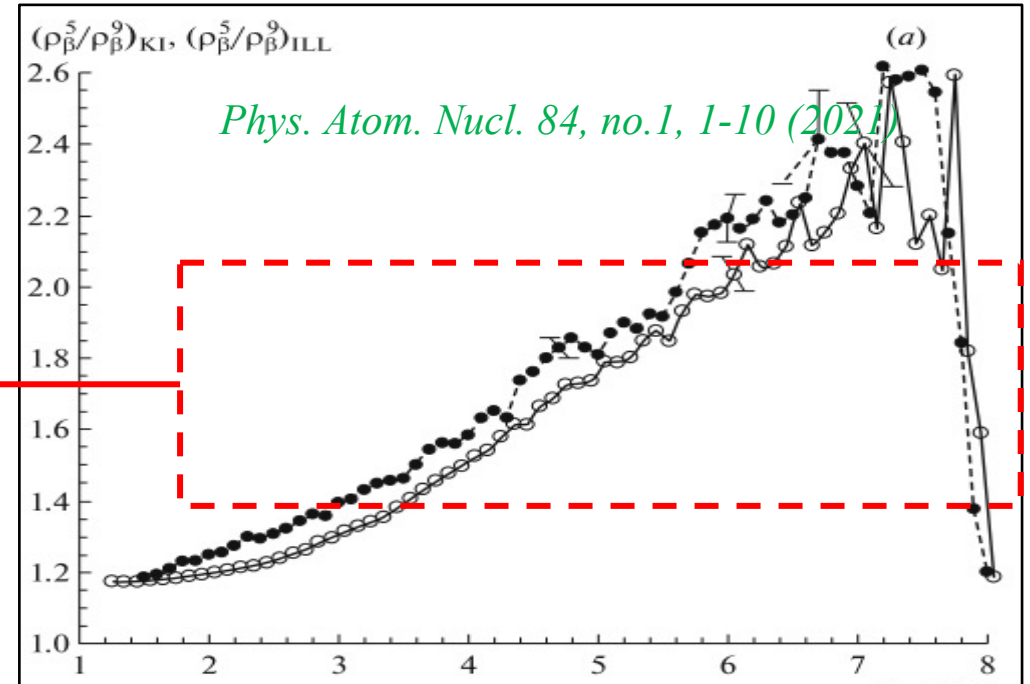
- HKSS model** *Phys. Rev. C 100, no.5, 054323 (2019)*
 - Forbidden transition contribution





Reactor flux models

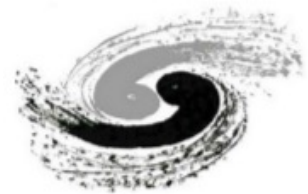
- **Kurchatov Institute model:** [arXiv:2103.01684](https://arxiv.org/abs/2103.01684)
 - ^{235}U HM model + KI measurement
 - ^{238}U **conversion spectrum** + KI measurement
 - Pu spectra: same with HM model
- **HKSS-KI model:**
 - ^{235}U HKSS model + KI measurement
 - ^{238}U and Pu: same with HKSS model



With the assumption of the unchanged ^{239}Pu comparing with ILL

KI measurement: Reduction of ^{235}U !

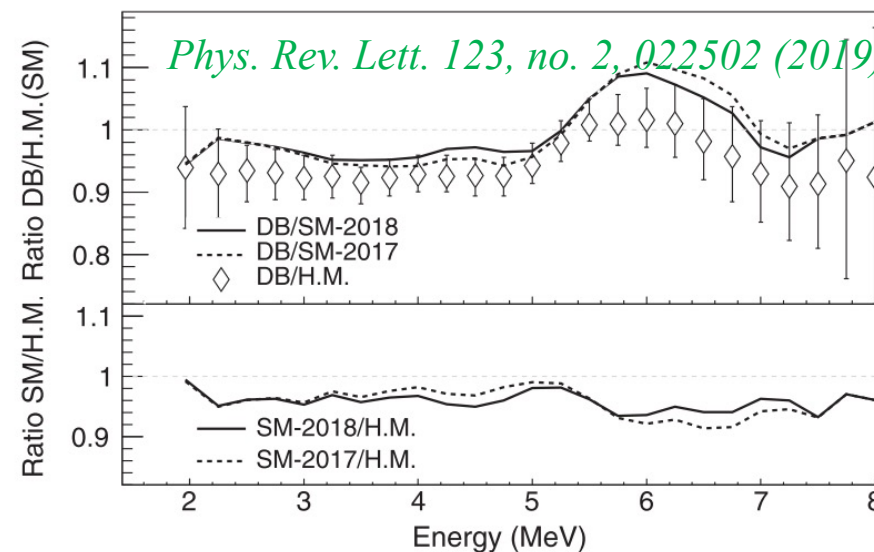
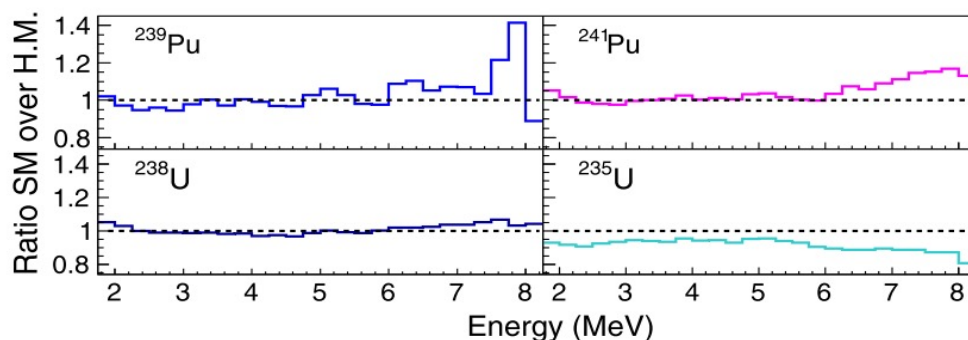
The **Kurchatov Institute measurement** (open circles) directly measured the ratio of ^{235}U beta spectrum and ^{239}Pu beta spectrum, which is lower than **HM model** (closed circles) in most region.



Reactor flux models

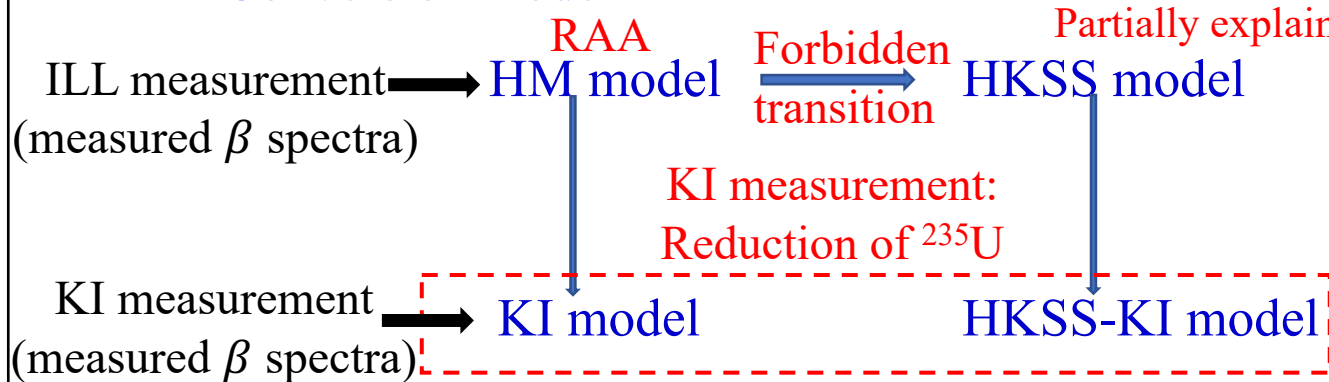
Phys. Rev. Lett. 123, no. 2, 022502 (2019)

- Estienne-Fallot summation model
 - Summation method
 - Nuclear database + Pandemonium-free data



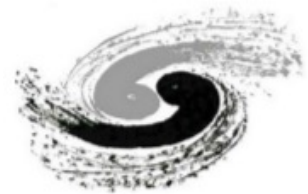
Model considered in this work

Conversion model



Summation model

EF model
The event rate is only 1.9% deviation from Daya Bay.



Updated IBD yields

- IBD yield σ_f $\sigma_{f,a} = \sum_i f_i^a \sigma_i$,

$i = 235, 238, 239$, and 241 for ^{235}U , ^{238}U , ^{239}Pu and ^{241}Pu .

- The individual IBD yield σ_i

$$\sigma_i = \int_{E_{\min}}^{E_{\max}} dE \Phi_i(E) \sigma_{\text{IBD}}(E),$$

1. IBD cross section: *Phys. Rev. D60, 053003 (1999)*

1st-order Vogel-Beacom IBD cross section w/ PDG 2020

0th-order cross section $\sigma_{\text{tot}}^{(0)} = \frac{2\pi^2/m_e^5}{f_{p.s.}^R \tau_n} E_e^{(0)} p_e^{(0)}$

Neutron lifetime $\tau_n = 879.4\text{s}$

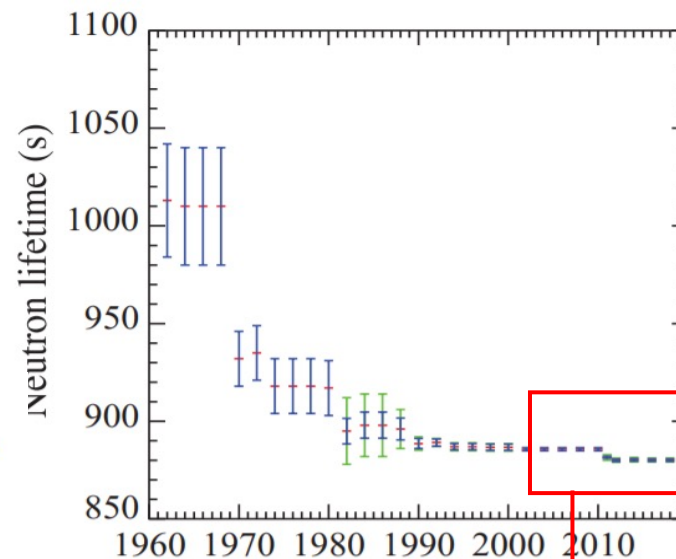
Phase space factor $f_{p.s.}^R = 1.7152$

0th-order $E_e^{(0)} = E_\nu - \Delta,$
 $\Delta = M_n - M_p.$

1st-order

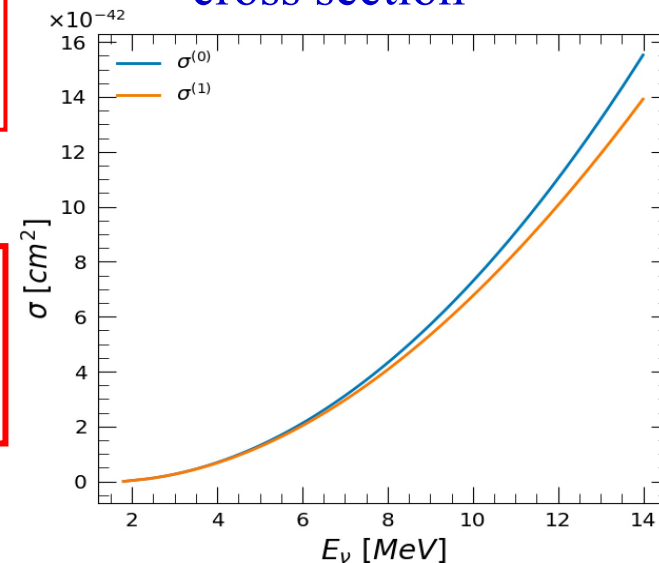
$$E_e^{(1)} = E_e^{(0)} \left[1 - \frac{E_\nu}{M} (1 - v_e^{(0)} \cos \theta) \right] - \frac{y^2}{M}$$

$$y^2 = (\Delta^2 - m_e^2)/2.$$

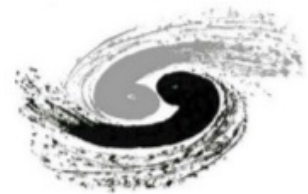


A historical perspective of values of neutron lifetime τ_n

0th and 1st order IBD cross section



2. Integral energy regions



Updated IBD yields

• The **individual IBD yield** σ_i

1. **IBD cross section** *Phys. Rev. D60, 053003 (1999)*

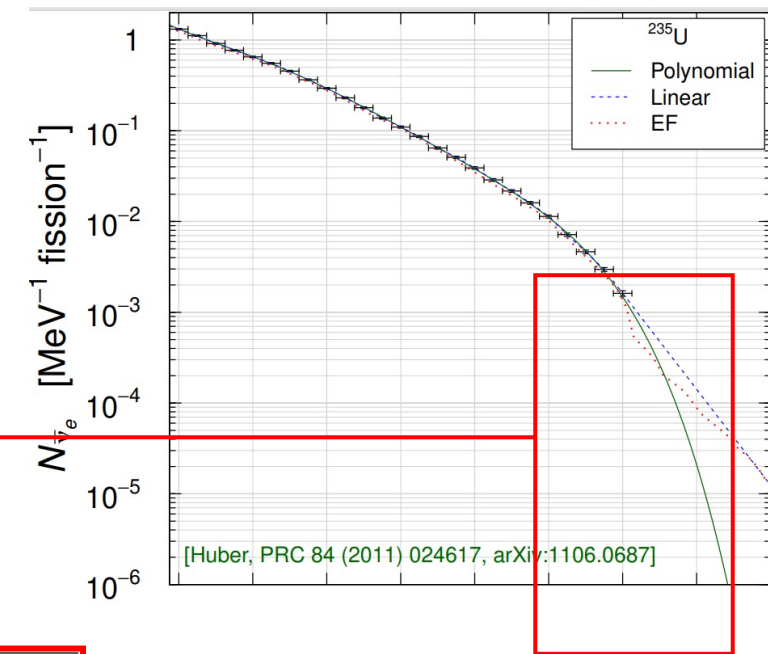
2. **Integral energy regions (1.8→10.0 MeV)**

- Low energy region (1.8 → 8.0 MeV)

extrapolate and interpolate with the original spectra.

- High energy region approximation (8.0 → 10.0 MeV)

EF summation model spectra with a very conservative **100% uncertainty**.



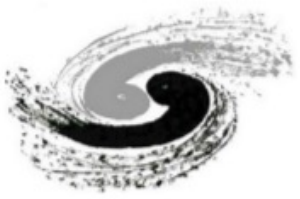
original IBD yields

Model	σ_{235}	σ_{238}	σ_{239}	σ_{241}
HM	6.69 ± 0.14	10.10 ± 0.82	4.40 ± 0.11	6.03 ± 0.13
EF	6.28 ± 0.31	10.14 ± 1.01	4.42 ± 0.22	6.23 ± 0.31
HKSS	6.74 ± 0.17	10.33 ± 0.85	4.43 ± 0.13	6.07 ± 0.16
KI	6.27 ± 0.13	9.34 ± 0.47	4.33 ± 0.11	6.01 ± 0.13

our selected IBD yields input

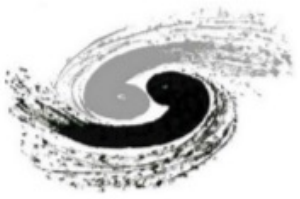
Model	σ_{235}	σ_{238}	σ_{239}	σ_{241}
HM	6.62 ± 0.16	10.09 ± 0.82	4.34 ± 0.13	6.02 ± 0.16
EF	6.23 ± 0.31	10.07 ± 1.00	4.37 ± 0.22	6.17 ± 0.31
HKSS	6.70 ± 0.17	10.19 ± 0.84	4.39 ± 0.13	6.09 ± 0.16
KI	6.29 ± 0.13	9.44 ± 0.48	4.34 ± 0.13	6.02 ± 0.16
HKSS-KI	6.36 ± 0.13	10.19 ± 0.84	4.39 ± 0.13	6.09 ± 0.16

Small contribution **above 8 MeV**:
 0.3% for ^{235}U , 0.9% for ^{238}U ,
 0.2% for ^{239}Pu , 0.3% for ^{241}Pu .



Part 3: Method of analysis

LSM with Wilks' theorem



How to treat the **systematic theoretical uncertainties** in the least-squares function.

Method A

Phys. Rev. D83, 073006 (2011)

JHEP 1706, 135 (2017)

A covariance matrix with experimental and theoretical uncertainties added in quadrature.

$$\chi^2 = \sum_{a,b} \left(\sigma_{f,a}^{\text{exp}} - R_{\text{NP}}^a \sigma_{f,a}^{\text{th}} \right) (V^{\text{tot}})^{-1}_{ab} \left(\sigma_{f,b}^{\text{exp}} - R_{\text{NP}}^b \sigma_{f,b}^{\text{th}} \right)$$

$$V^{\text{tot}} = V^{\text{exp}} + \underline{V^{\text{th}}} \quad \sigma_{f,a}^{\text{th}} = \sum_i f_i^a \sigma_i^{\text{mod.}}$$

A **strongly-correlated theoretical** matrix derived from the covariance matrix V_{ij}^{mod} among ^{235}U , ^{238}U , ^{239}Pu , and ^{241}Pu



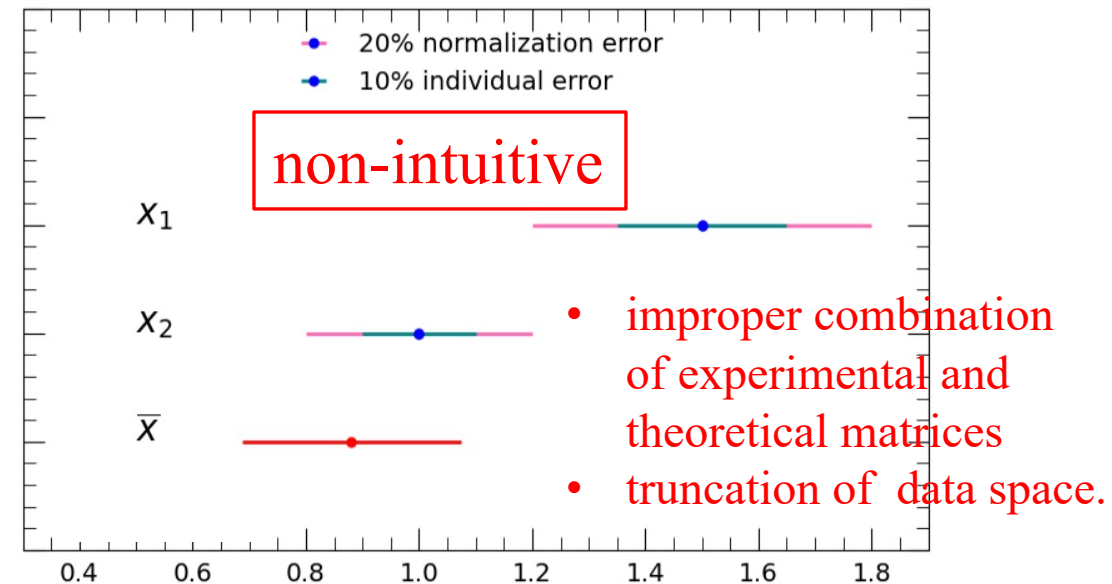
The method A will suffer the PPP!

Journal of Nuclear Science and Technology 31, 770 (1994).

Peelle's Pertinent Puzzle

strongly correlated data

the best-fit average can be lower than most of the data



LSM with Wilks' theorem



Method B *Phys. Rev. D87, 073018 (2013)*

Calculate the fit results considering only the experimental uncertainties and add by hand a global theoretical uncertainty to the final result.

$$\chi^2 = \sum_{a,b} \left(\sigma_{f,a}^{\text{exp}} - R_{\text{NP}}^a \sigma_{f,a}^{\text{th}} \right) (V^{\text{exp}})_{ab}^{-1} \left(\sigma_{f,b}^{\text{exp}} - R_{\text{NP}}^b \sigma_{f,b}^{\text{th}} \right)$$

hard to calculate

Method C **Method C is adopted in this work!**

Phys.Rev.Lett. 120, 022503 (2018),

Phys.Rev. D99, 073005 (2019)

Consider the theoretical uncertainties with appropriate **pull terms**

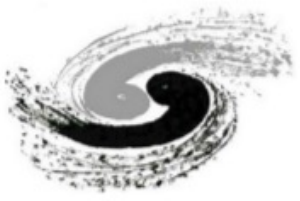
$$\chi^2 = \sum_{a,b} \left(\sigma_{f,a}^{\text{exp}} - R_{\text{NP}}^a \sigma_{f,a}^{\text{th}} \right) (V^{\text{exp}})_{ab}^{-1} \left(\sigma_{f,b}^{\text{exp}} - R_{\text{NP}}^b \sigma_{f,b}^{\text{th}} \right)$$

$$+ \sum (r_i - 1) \left(\tilde{V}^{\text{mod}} \right)_{ij}^{-1} (r_j - 1),$$

$$\sigma_{f,a}^{\text{th}} = \sum_i r_i f_i^a \sigma_i^{\text{mod}}. \quad \tilde{V}_{ij}^{\text{mod}} = V_{ij}^{\text{mod}} / (\sigma_i^{\text{mod}} \sigma_j^{\text{mod}})$$

V_{ij}^{mod} covariance matrix for these four isotopes

PPP is avoided by decoupling the minimization of **physical parameters** from the minimization of **pull coefficients!**

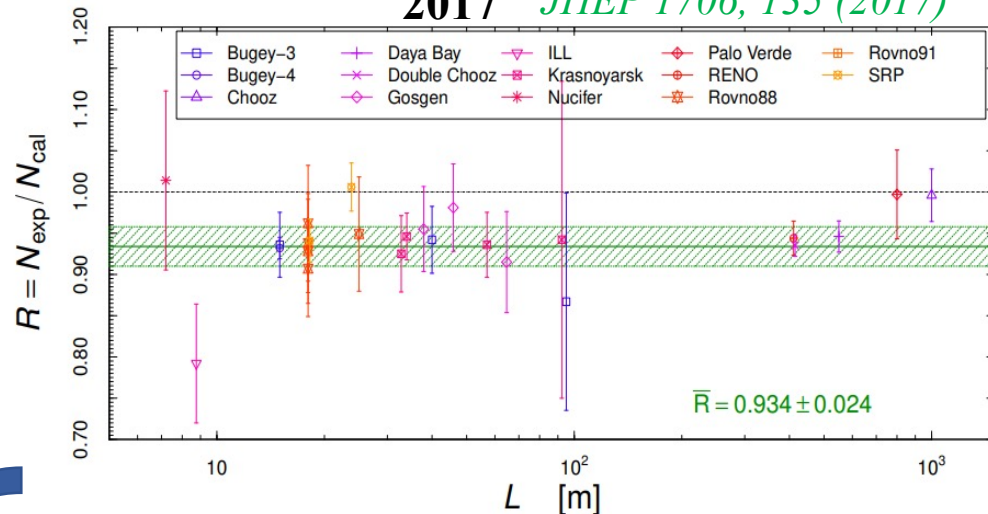


Part 4: Fit of reactor rates & evolution data



Fit of reactor rates

2017 *JHEP 1706, 135 (2017)*

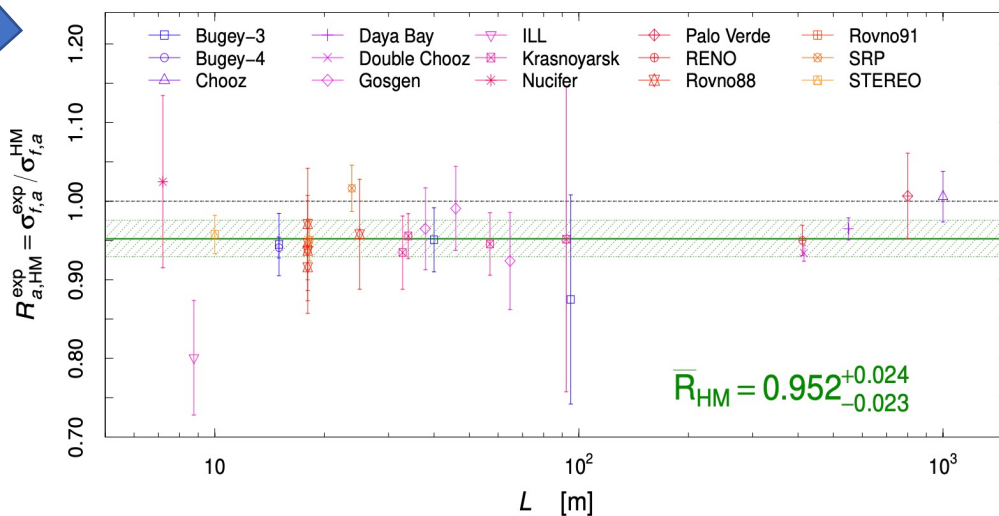


HM model

Original HM IBD yield	Updated HM IBD yield
6.69 ± 0.14	6.62 ± 0.16
10.10 ± 0.82	10.09 ± 0.82
4.40 ± 0.11	4.34 ± 0.13
6.03 ± 0.13	6.02 ± 0.13



Our work



Work in 2017

RAA: 2.8σ w/ new IBD yields
 0.934 ± 0.024

Work in 2021

RAA: 2.5σ
 0.940 ± 0.024

w/ method C

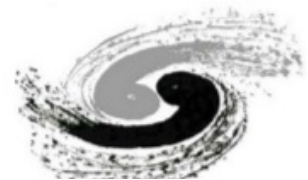
0.943 ± 0.024 w/ new IBD yields

w/ method C

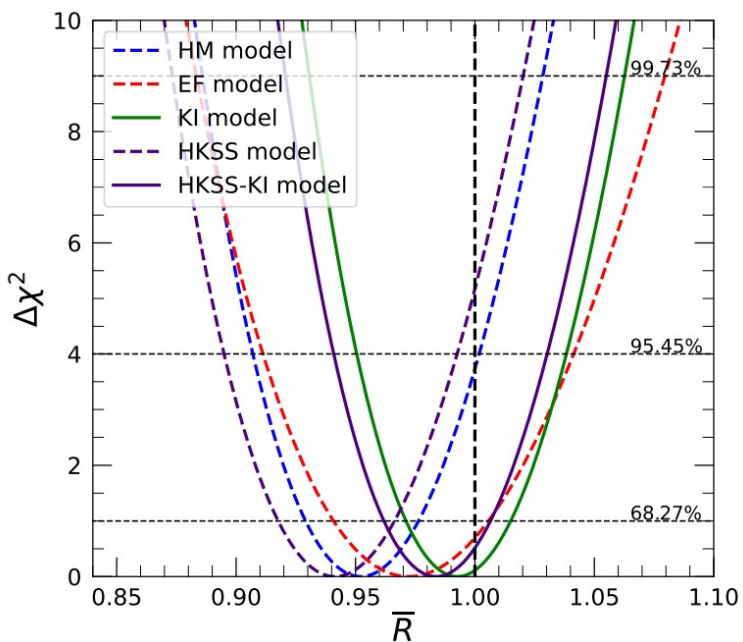
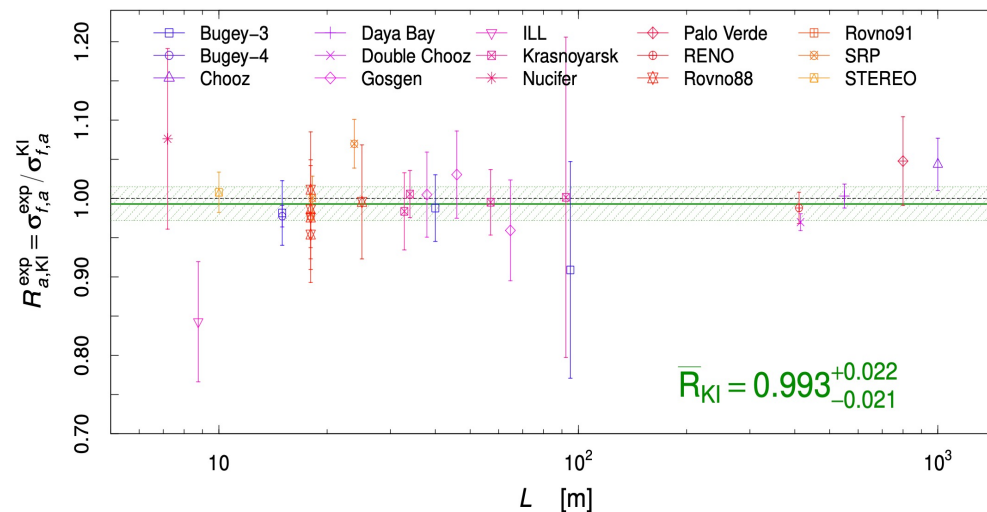
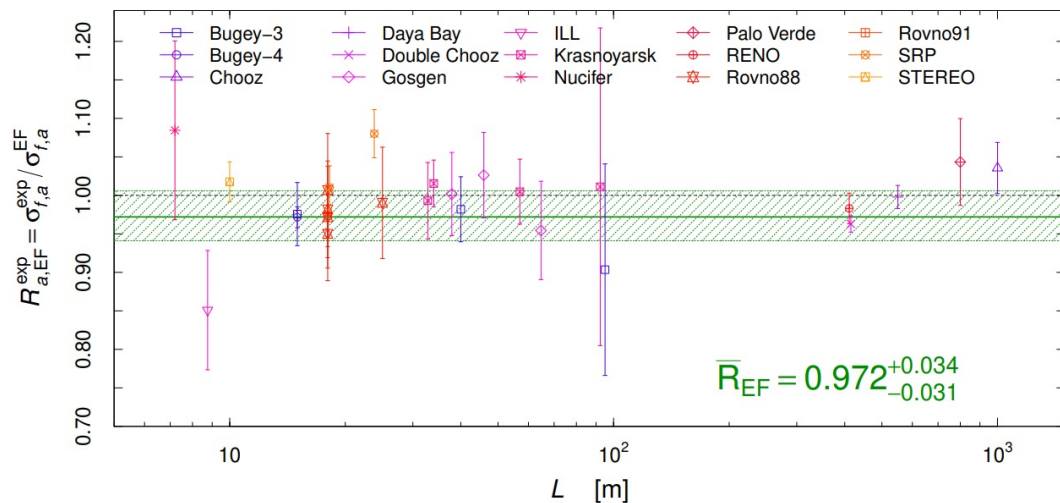
0.952 ± 0.024
 RAA: 1.9σ

RAA: 2.3σ

RAA: 2.8σ (2017) \rightarrow 1.9σ (2021)



Fit of reactor rates



Model	\bar{R}	RAA
HKSS	$0.941^{+0.025}_{-0.023}$	2.3σ
EF	$0.972^{+0.034}_{-0.031}$	0.9σ
KI	$0.993^{+0.022}_{-0.021}$	0.3σ
HKSS-KI	$0.982^{+0.022}_{-0.021}$	0.8σ

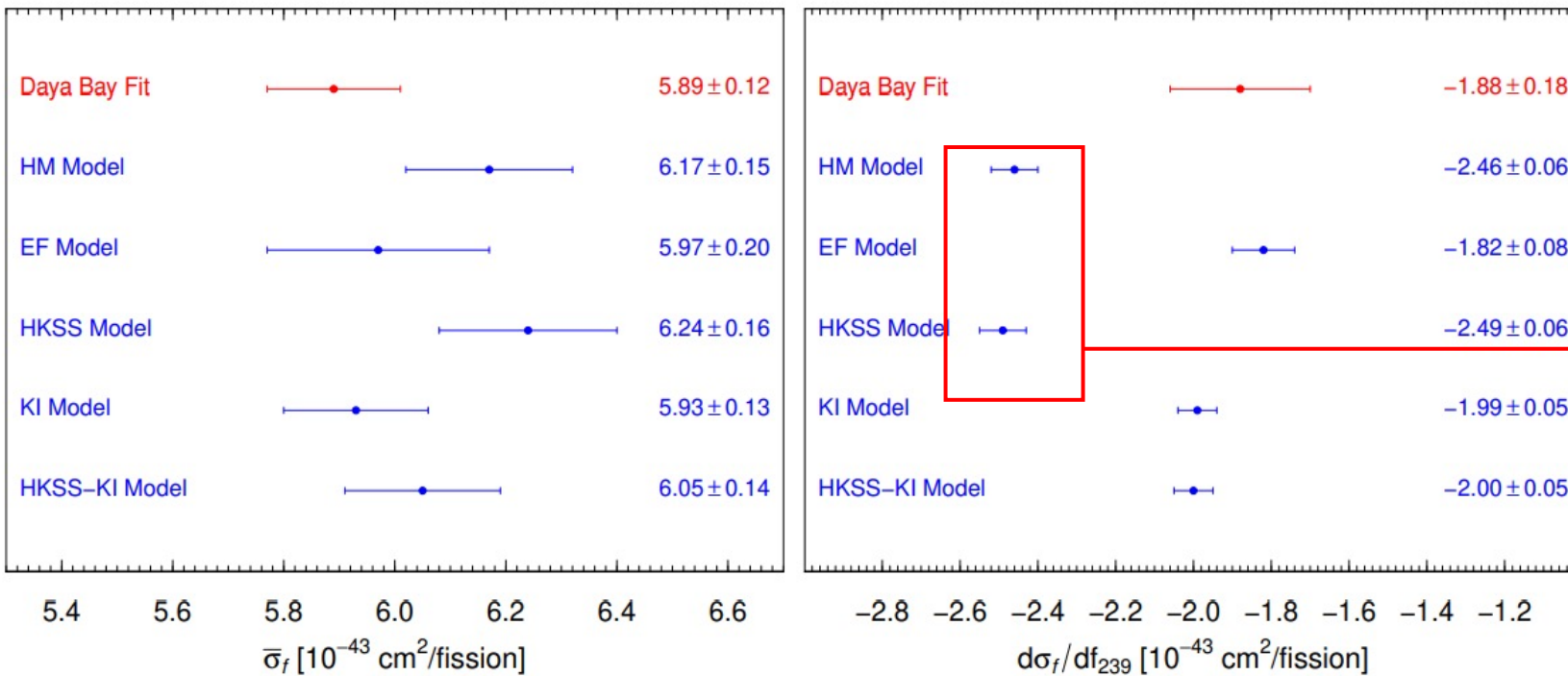
These 3 models give RAA less than 1σ . (No anomaly)



Fit of reactor fuel evolution data

To compare the fuel evolution data with the different model predictions, we first fit the evolution data with a linear function describing the IBD yield as a function of f_{239}

$$\sigma_{f,a}^{\text{lin}} = \bar{\sigma}_f + \frac{d\sigma_f}{df_{239}} (f_{239}^a - \bar{f}_{239}),$$



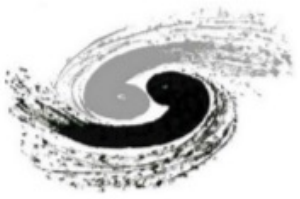
The **HM** and **HKSS** models are disfavored by the evolution data

Fig(b) fit of slope

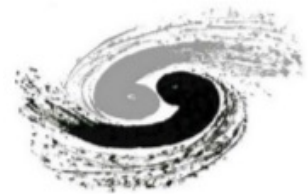
- 3.1σ for HM model
- 3.2σ for HKSS model
- EF, KI and HKSS-KI models give values of $\bar{\sigma}_f$ and $d\sigma_f/df_{239}$ that agree with the fit of the evolution data within the uncertainties.

(a) Fitting with evolution data of Daya Bay

(b) When using RENO data, we have **the similar results**.



Part 5: Best-fit model



Statistic test

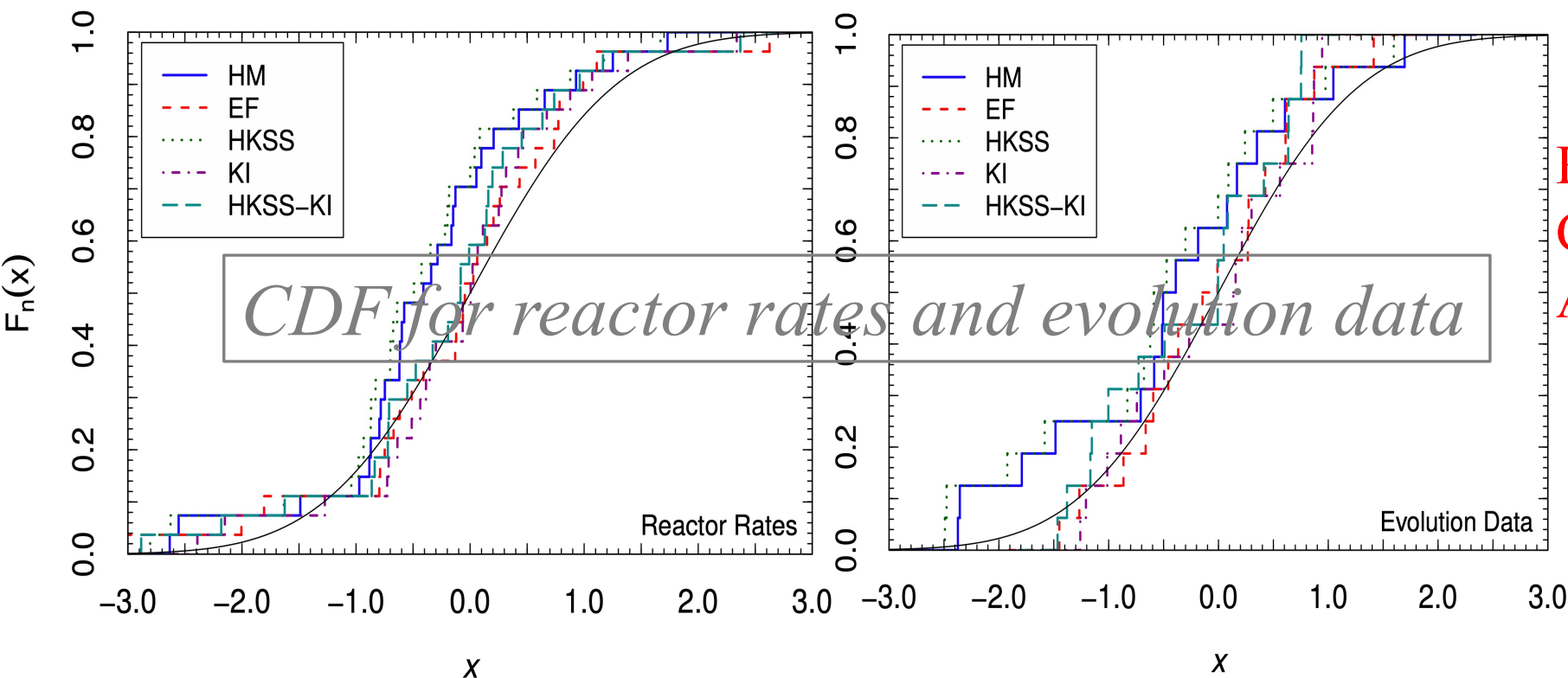
- χ^2 test: only shows the **size** of deviation not show the sign
rejects none of the five models

only size of deviation
 χ^2 test

$$x_a^{\text{mod}} = \sum_b (V^{\text{tot}})_{ab}^{-1/2} (\sigma_{f,b}^{\text{exp}} - \sigma_{f,b}^{\text{mod}})$$

Shapiro-Wilk test

sign test **positive or negative deviations**



Kolmogorov-Smirnov test
Cramer-von Mises test
Anderson-Darling test

Journal of the Royal Statistical Society Series B 64, 281 (2002).

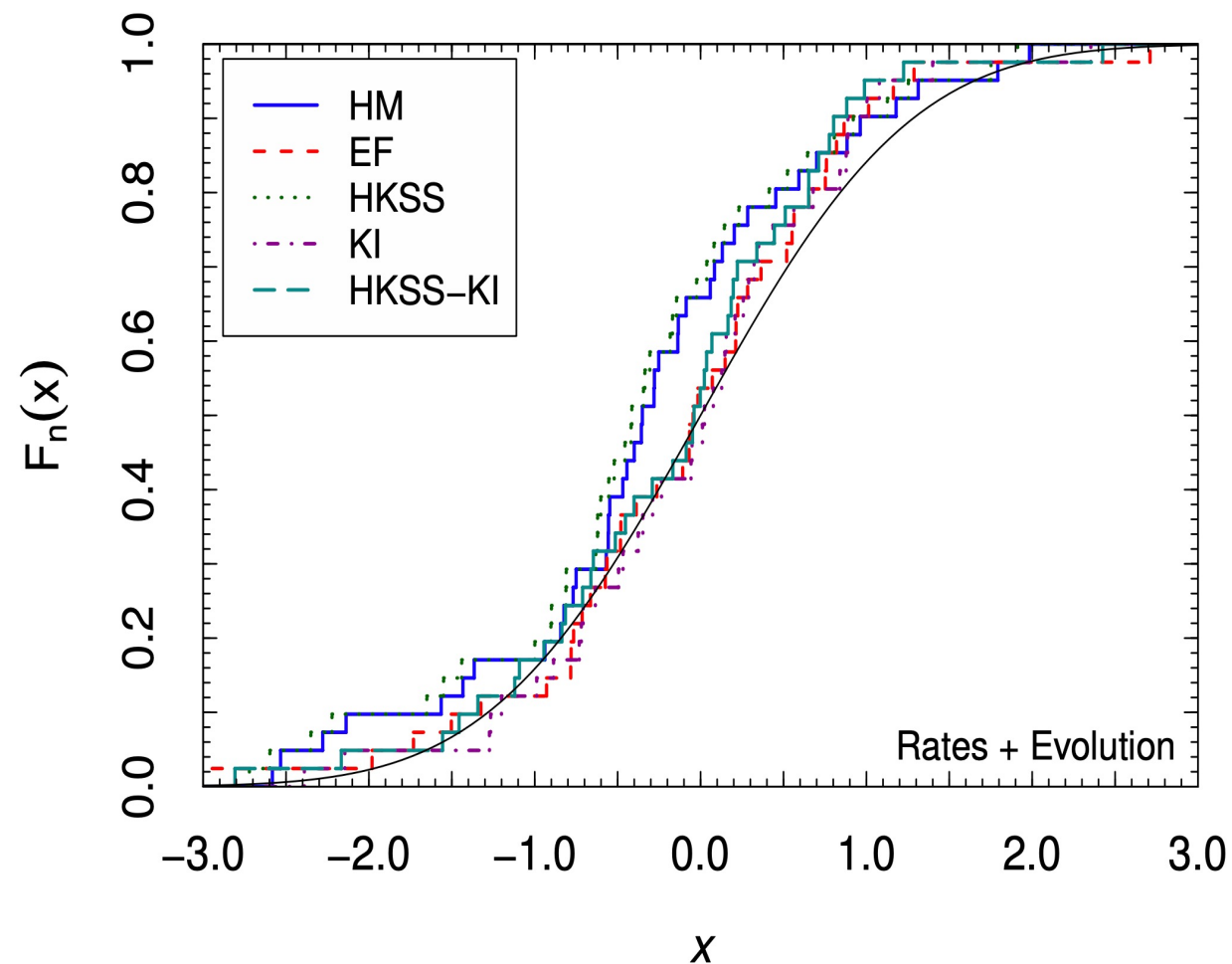
Z_K, Z_C, Z_A test

more powerful, based on likelihood ratio



Statistic test

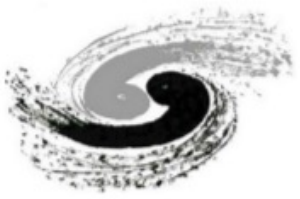
p-value=0.05 → confidence level 95%



rates + evolution data

Test	HM	EF	HKSS	KI	HKSS-KI
χ^2	0.21	0.46	0.14	0.78	0.60
SW	0.37	0.28	0.38	0.69	0.58
sign	0.03	0.38	0.01	0.38	0.38
KS	0.08	0.81	0.04	0.82	0.56
CVM	0.06	0.78	0.03	0.76	0.47
AD	0.07	0.72	0.03	0.76	0.47
Z_K	0.001	0.22	0.0002	0.15	0.03
Z_C	0.11	0.60	0.04	0.86	0.62
Z_A	0.14	0.44	0.06	0.77	0.40

EF model is the best **summation** model;
KI model is the best **conversion** model.

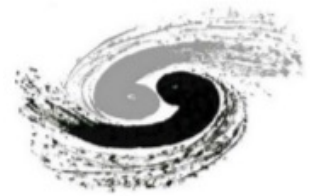


Part 6: Summary

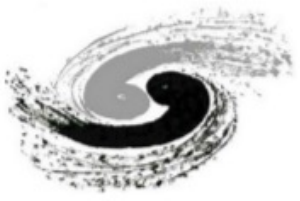
Summary



- Updated IBD yields including high energy regions.
 - Comparison of different fitting method
- HM model**
RAA: $2.8\sigma \rightarrow 1.9\sigma$
- With improved fitting method (**Method C**), the RAA seems smaller for all models (for HM, $2.5\sigma \rightarrow 1.9\sigma$) avoiding the PPP successfully.
- As for the best-fit model, **EF** model is the best **summation model**, and **KI** model is the best **conversion model**.
 - The KI measurement can pull down the **rate deficit**, which implies the reactor antineutrino anomaly might be caused by **mis-normalization** in ILL measurements. (need other experiments to confirm)
 - Shape anomaly (“**5 MeV Bump**”) is still not solved.



Thanks!



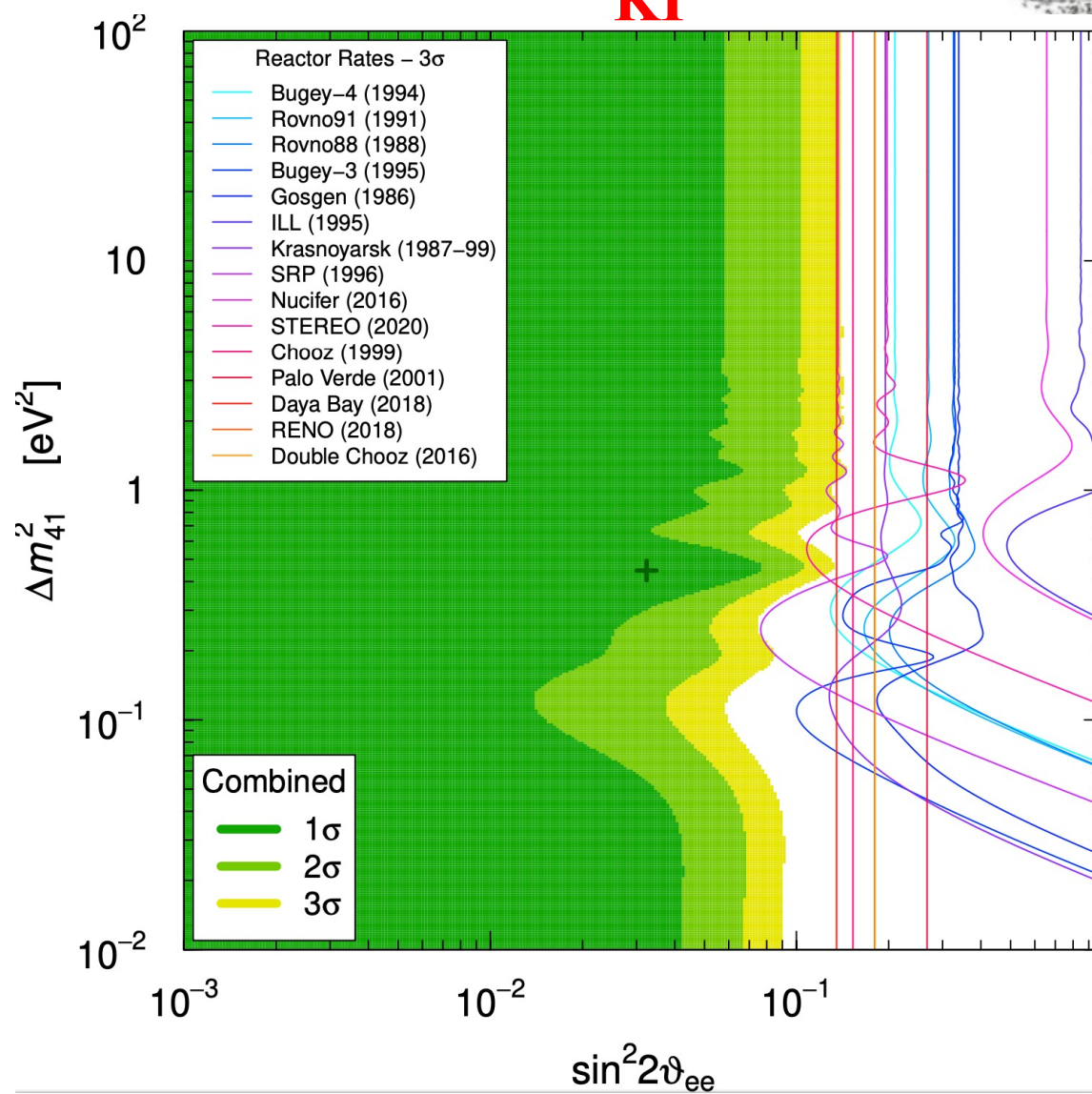
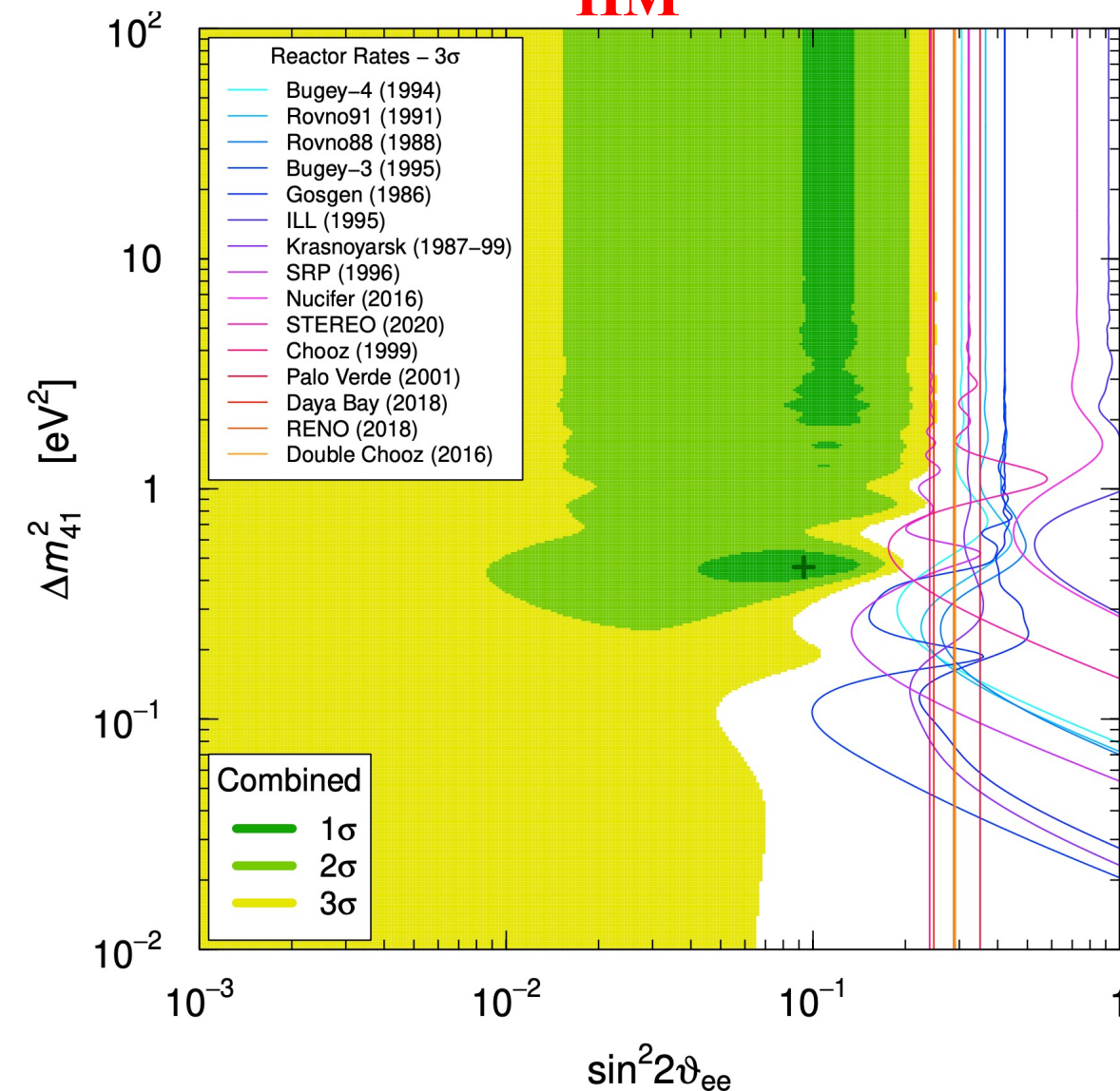
Backup



Oscillation

HM

KI



Oscillation

