



# Impact of cross-section modeling on the NOvA oscillation analyses

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# Understanding how cross sections impact NOvA

What features of the neutrino interaction does the detector measure?

What models do we use to estimate those features?

How sensitive is the 3-flavor oscillation analyses to those features?

# What is NOvA

## Long-baseline neutrino experiment

$E \approx 2 \text{ GeV}$  (off-axis narrow band beam)

$L = 810 \text{ km}$

Oscillations governed by  $\Delta m_{32}^2$  ( $\Delta m_{31}^2$ )

## NuMI beam produced at Fermilab

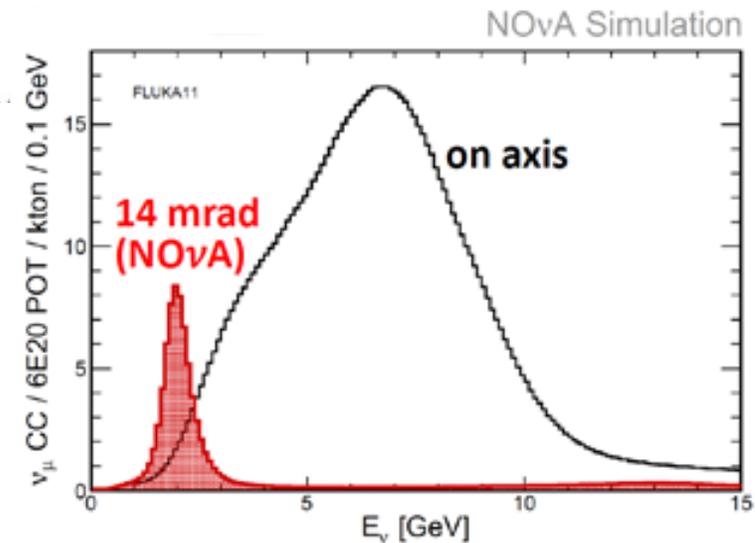
$\nu_\mu$  and  $\bar{\nu}_\mu$  beam modes

$(\bar{\nu}_\mu) \rightarrow (\bar{\nu}_\mu)$  and  $(\bar{\nu}_\mu) \rightarrow (\bar{\nu}_e)$  oscillations

## Two detector experiment

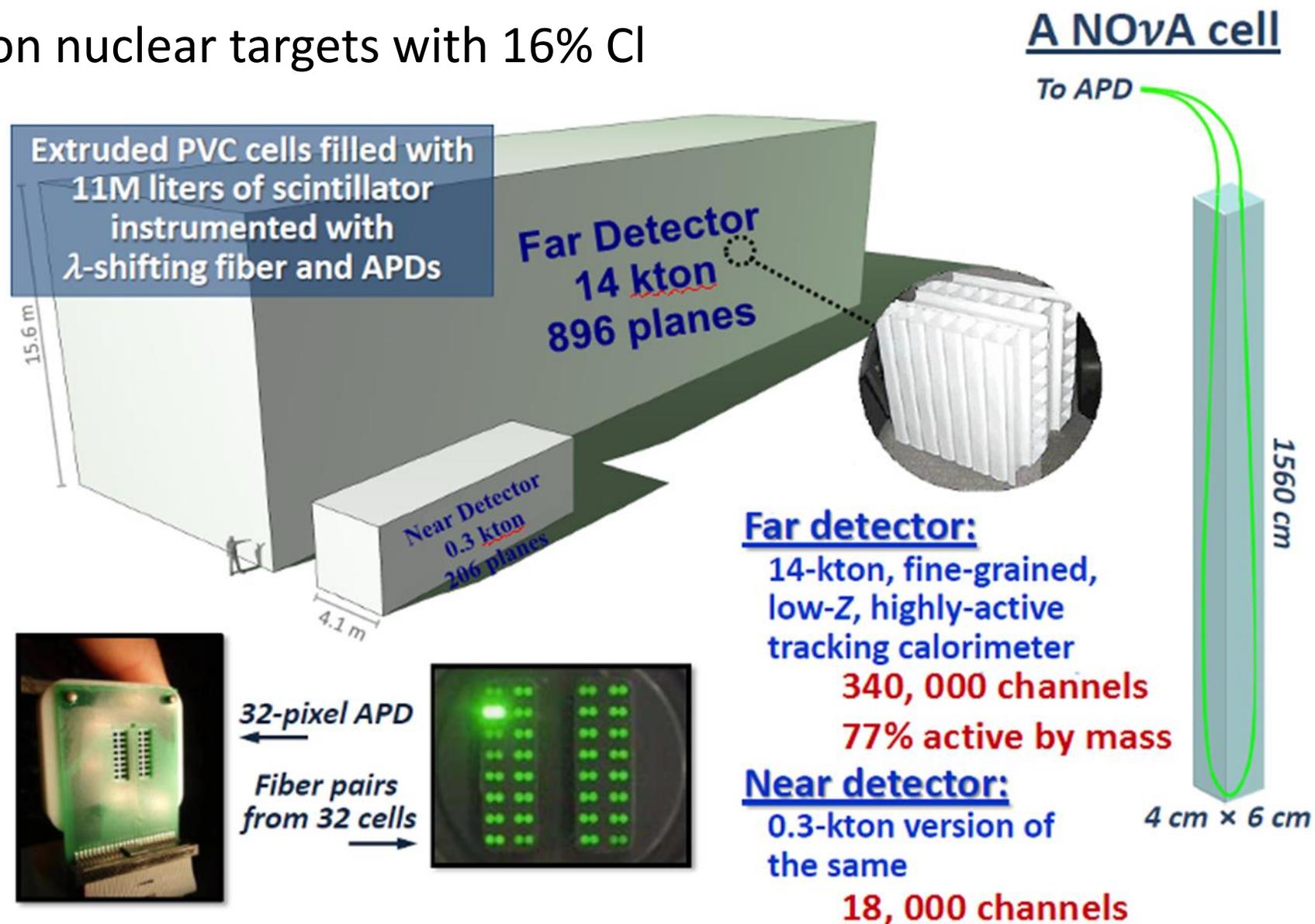
Near detector (Fermilab, IL)  
Measure beam before oscillation

Far Detector (Ash River, MN)  
Measure oscillated beam



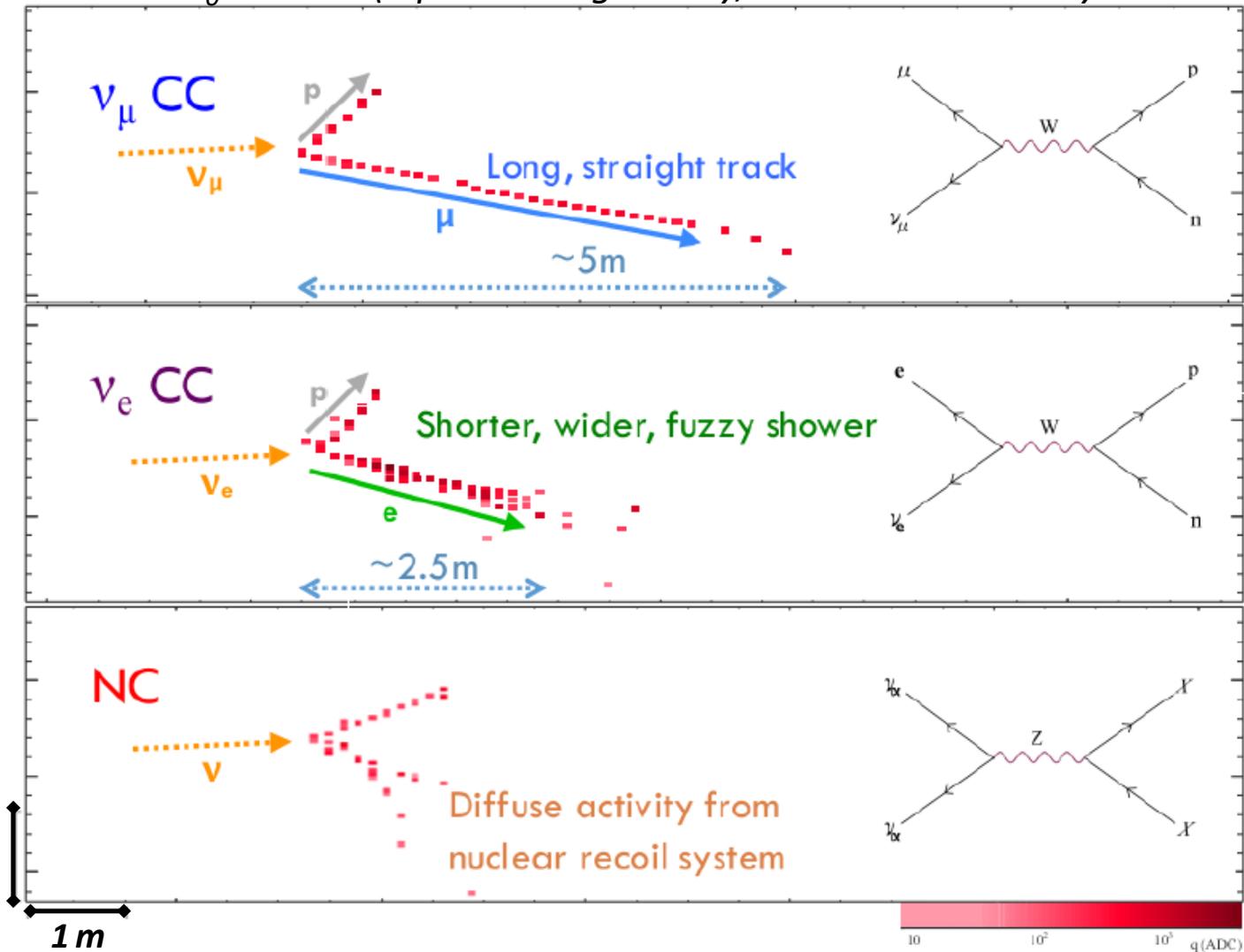
# The NOvA Detectors

Giant hydrocarbon nuclear targets with 16% Cl



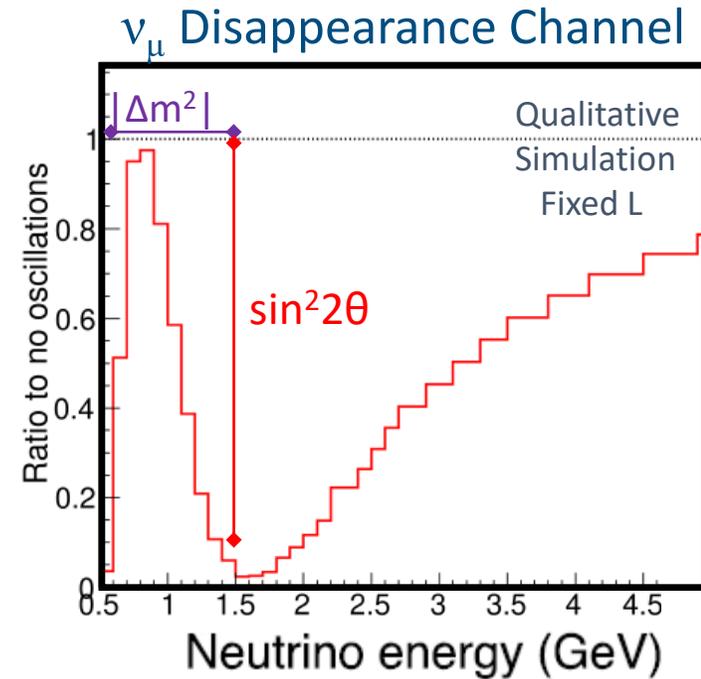
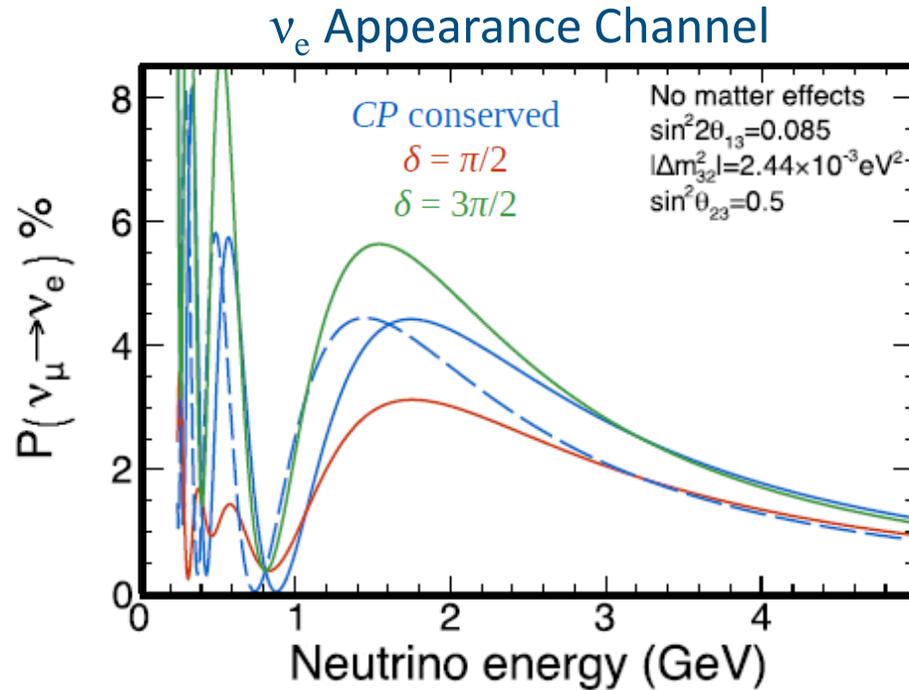
# NOvA Event Topologies

$X_0 = 38$  cm (6 planes longitudinally, 10 cells transversely)

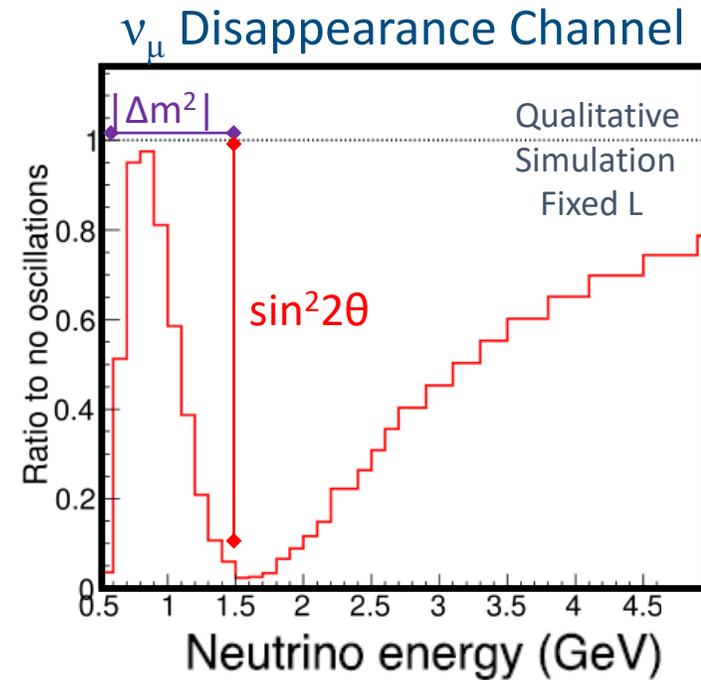
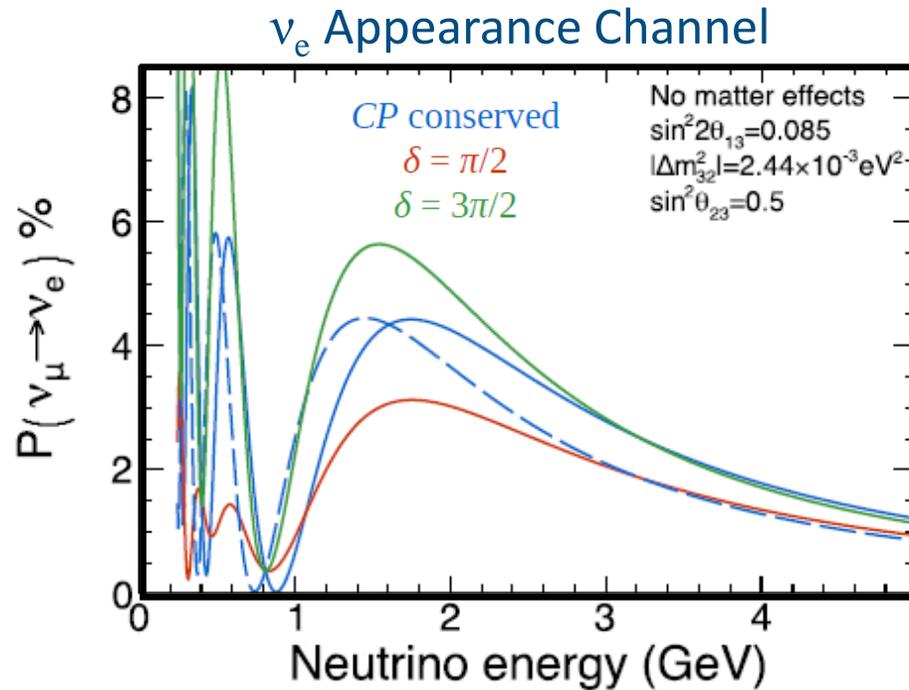


# NOvA Oscillation Measurements

Sensitive to:  $\delta_{CP}$ ,  $\theta_{23}$ , mass splitting (hierarchy /ordering)



# NOvA Oscillation Measurements

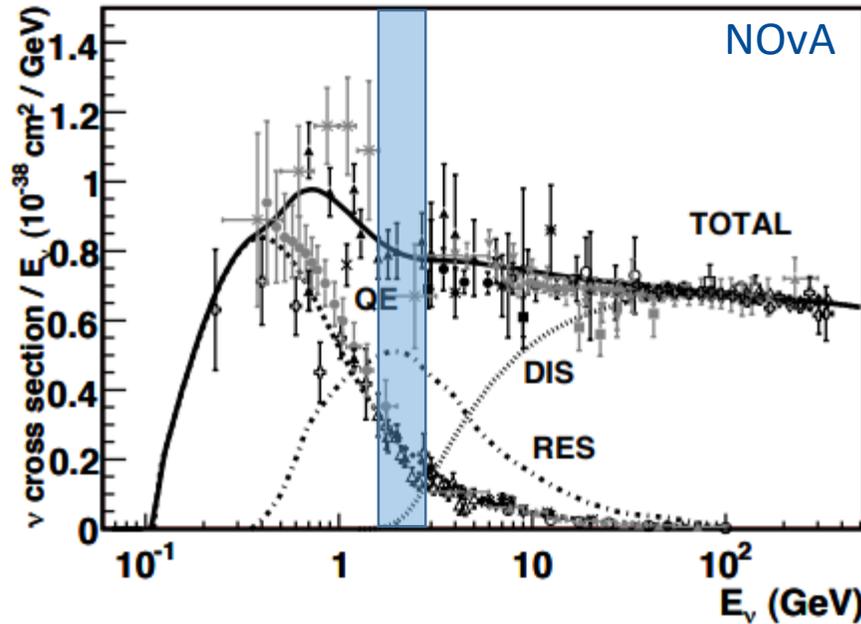


## The key measurements for oscillation results

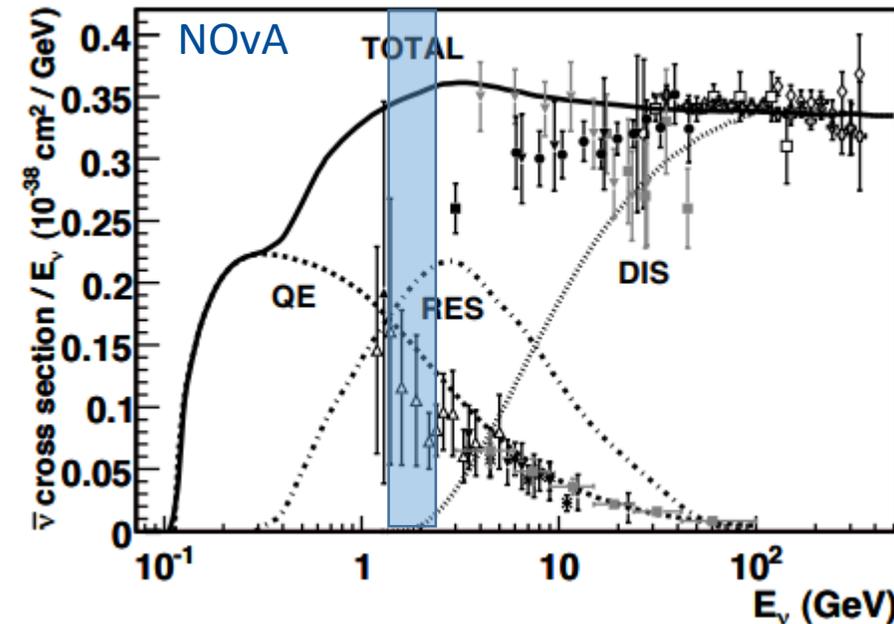
- Neutrino energy
- Event yields

# What processes are relevant

## Neutrinos



## Antineutrinos

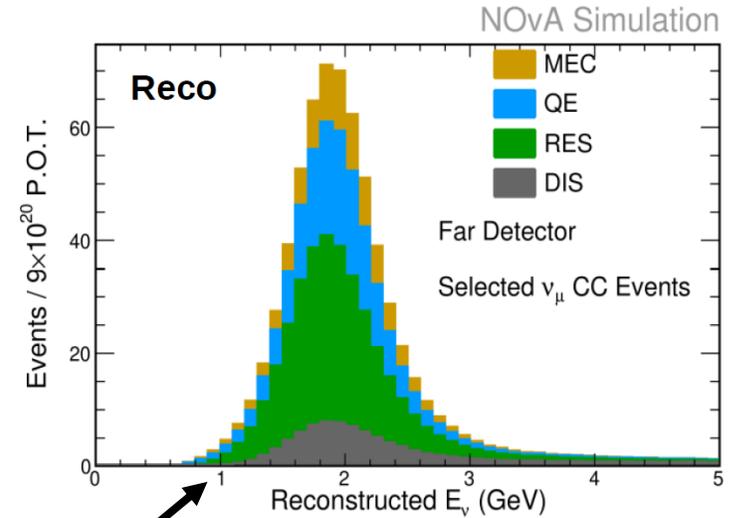
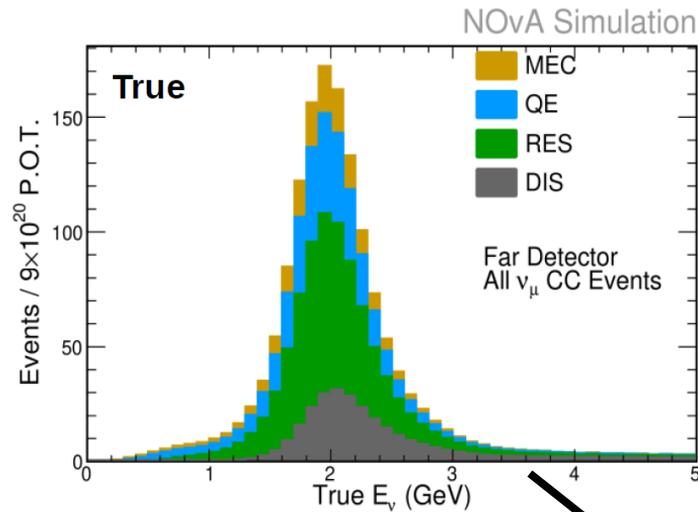


*Rev. Mod. Phys. 84, 1307 (2012)*

NOvA at an energy where resonant production is dominant

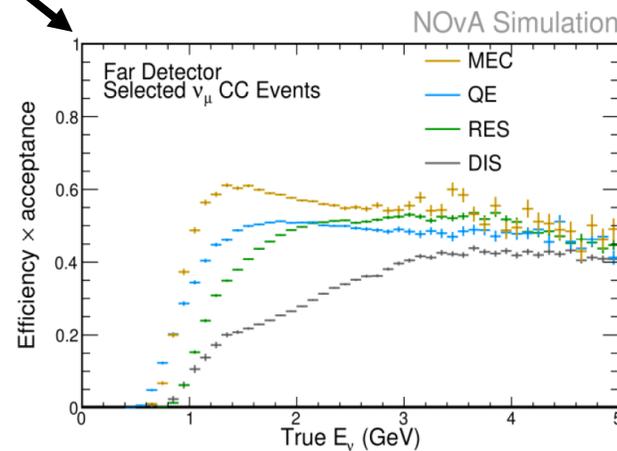
However the mix of QE, 2p2h, RES, and DIS is important and muddles things

# Event Yields



**Different processes have different selection efficiencies**

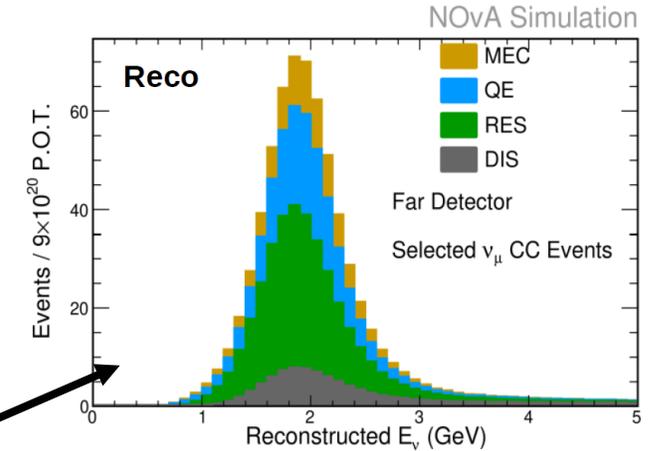
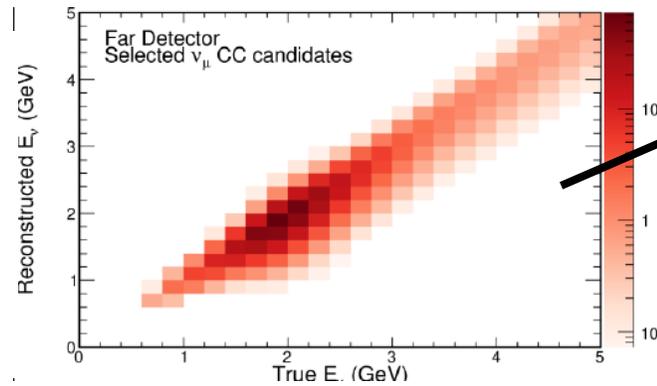
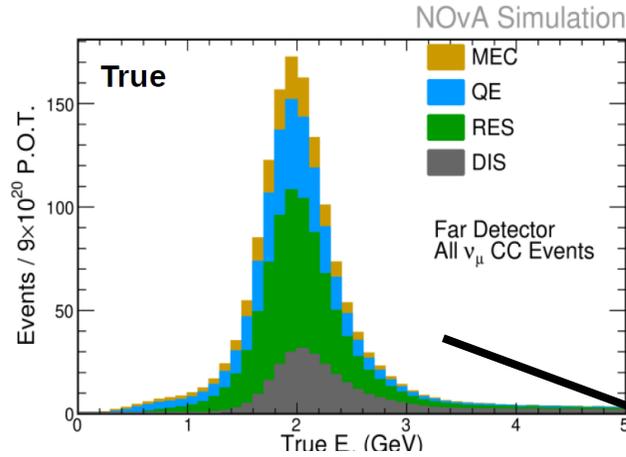
*How much 4-momentum goes to the lepton, how much goes to the hadronic system*



*Ideally apply same selection in ND and FD, so they have same efficiencies, but the mix of processes is not exactly identical in both detectors*

*Flux is different (eg oscillation)  
Containment is different*

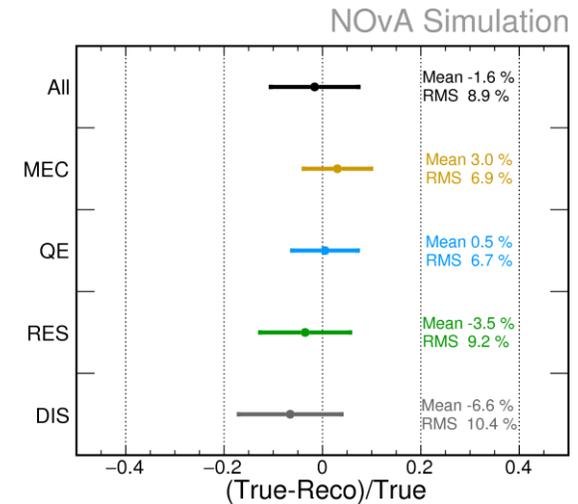
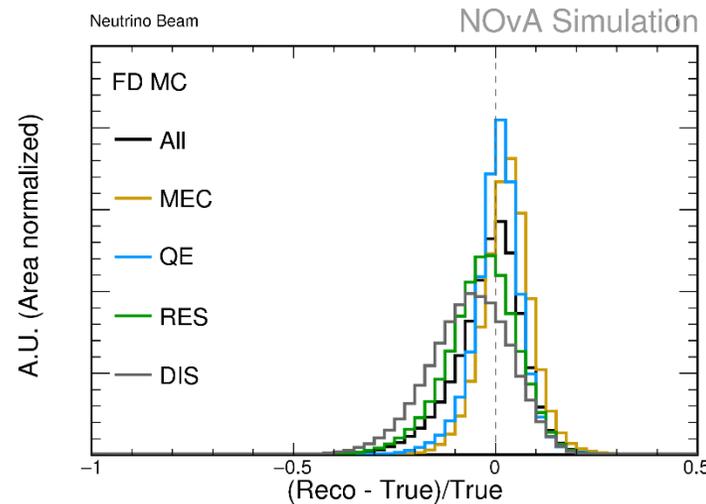
# Neutrino Energy



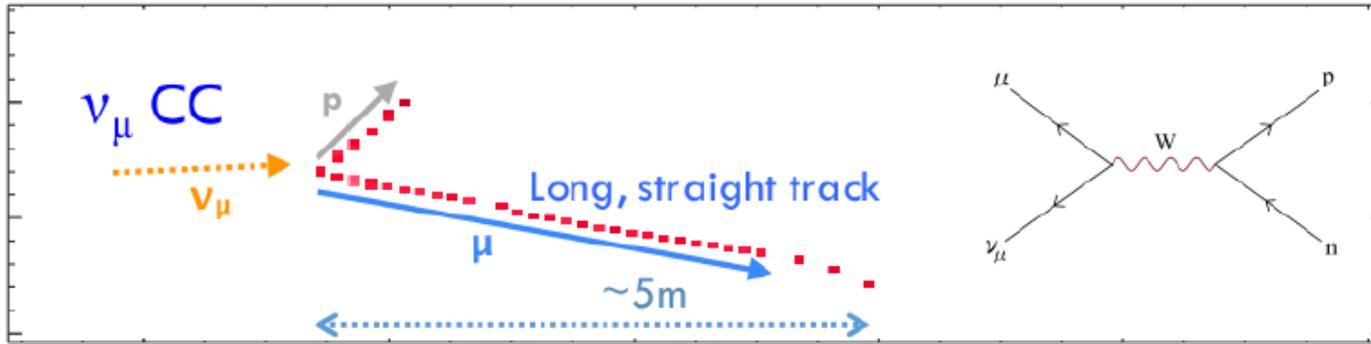
**Different processes have different energy resolutions and energy scales**

*How much 4-momentum goes to the lepton, how much goes to the hadronic system*

*How much 4-momentum is invisible*

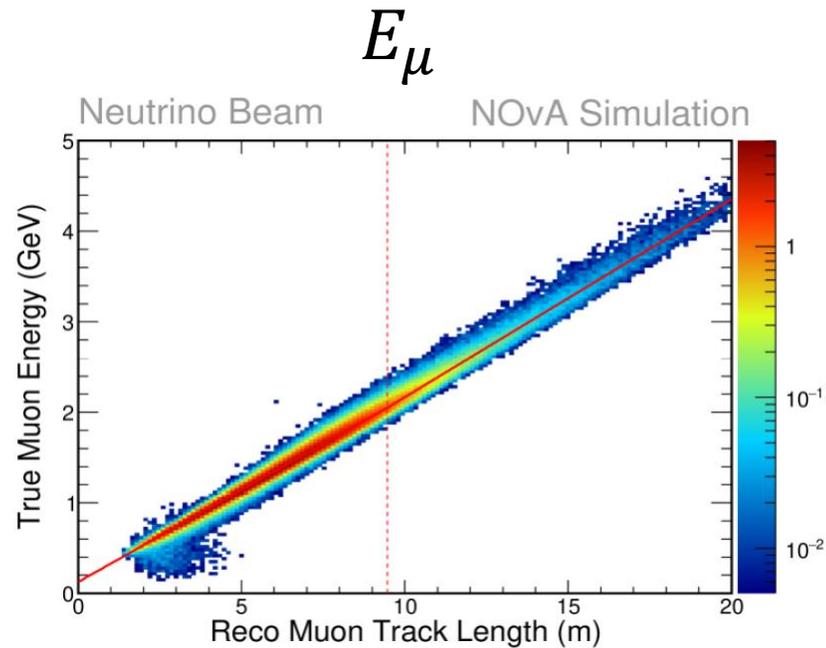


# Event Topology

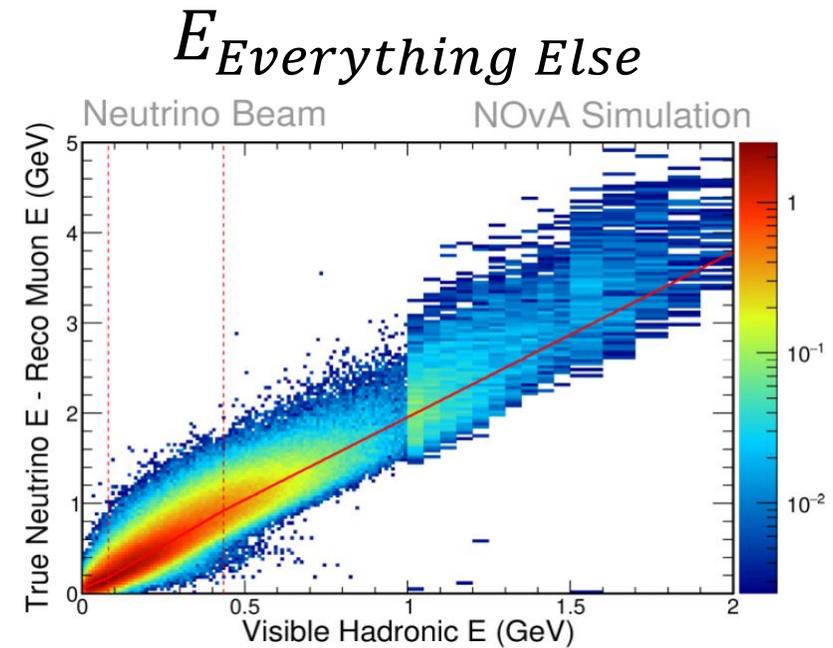


# Energy Measurement

$$E_\nu = E_\mu + E_{Everything\ Else}$$



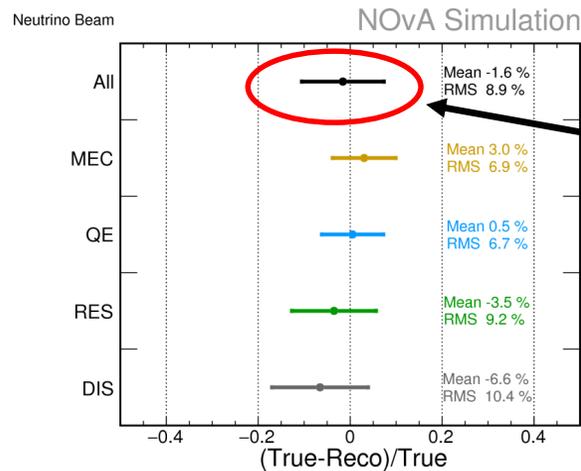
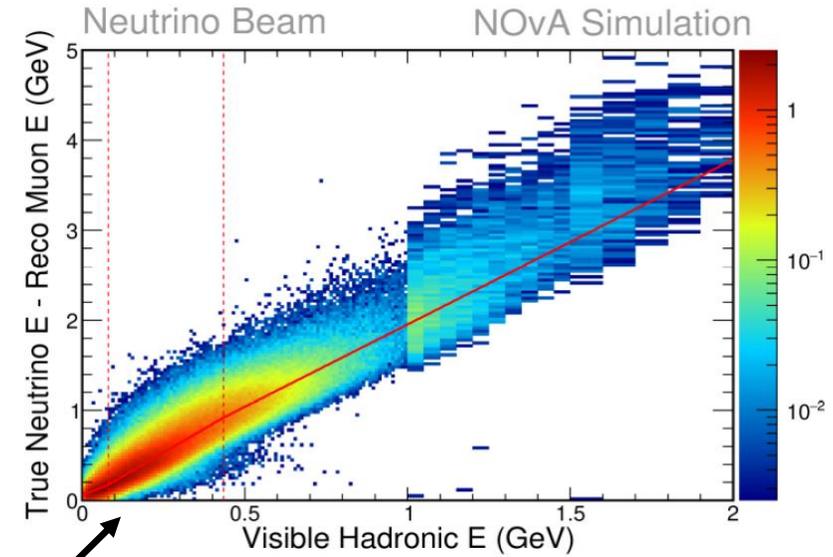
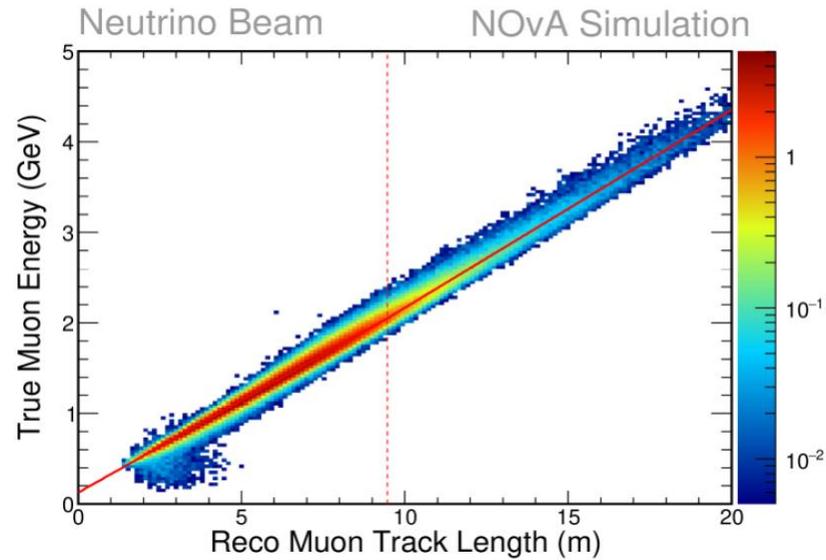
Muon energy by range  
 $\nu$  cross-section independent



Calorimetric sum of non-muon hits  
Modelling dependent

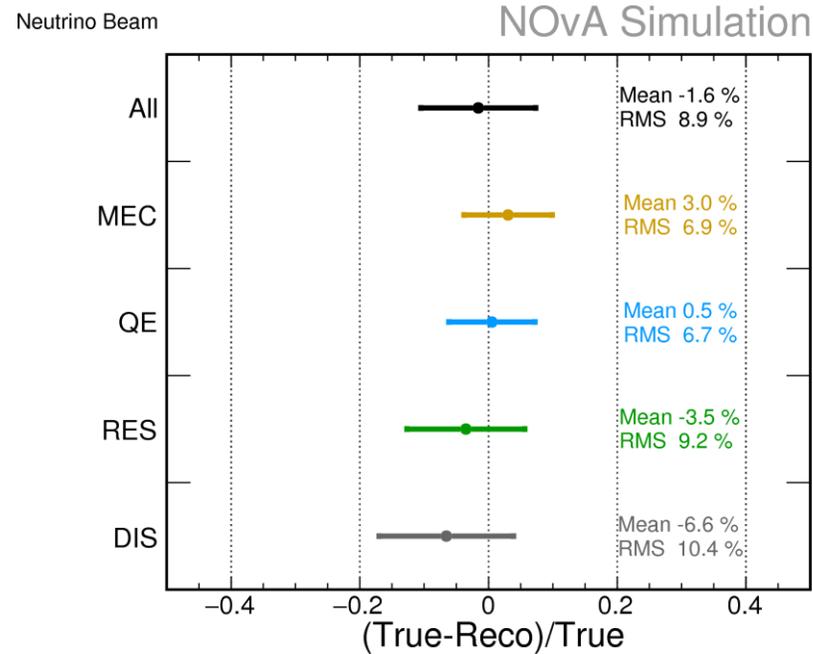
# Energy Measurement

$$E_\nu = E_\mu + E_{Everything\ Else}$$

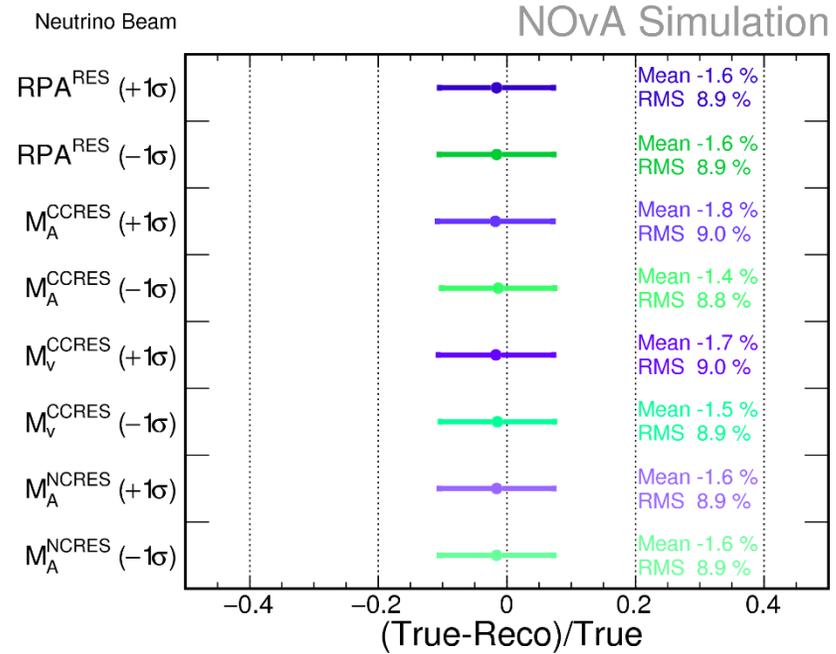


With this method, the average value depends on the relative mix of the processes and how much energy goes to the lepton

# Energy Measurement



These differences in energy scale & resolution would represent our uncertainty if we had complete uncertainty on the relative contribution from each process



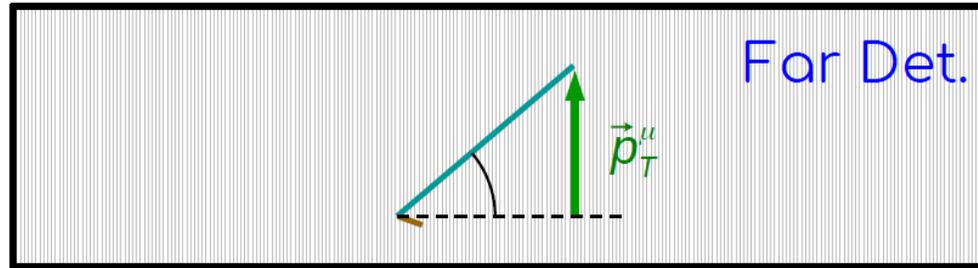
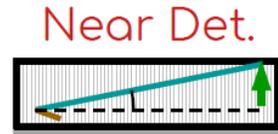
Fortunately, we know the relative contribution from each process fairly well

As determined by our model uncertainties

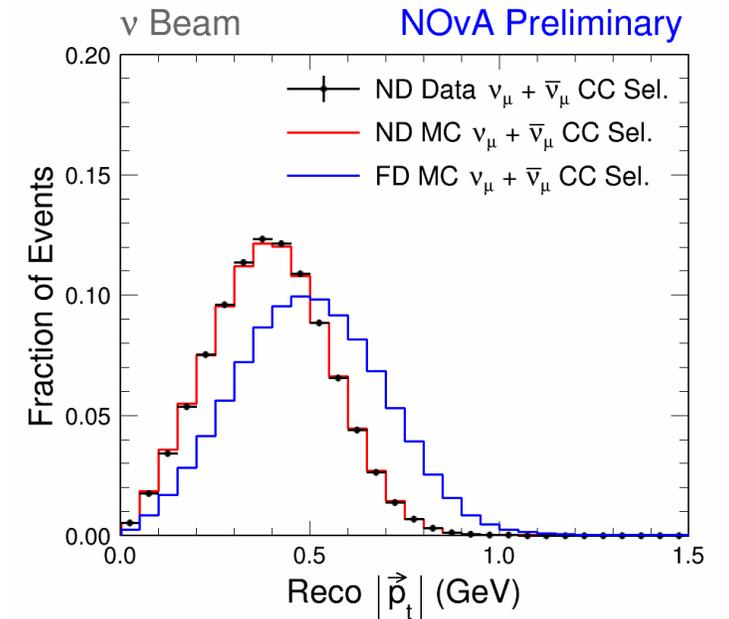
**The primary impact of cross-section modelling on analyses is through event yields not energy reconstruction!**

# Mitigating Acceptance Uncertainties

Acceptance of smaller ND events depends on event kinematics

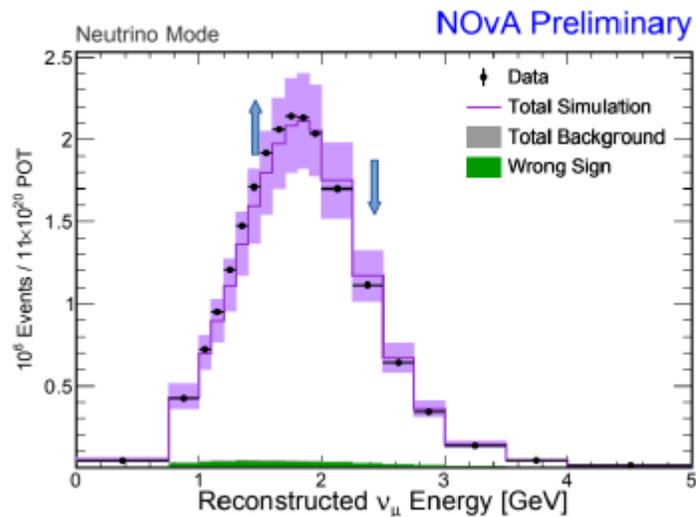
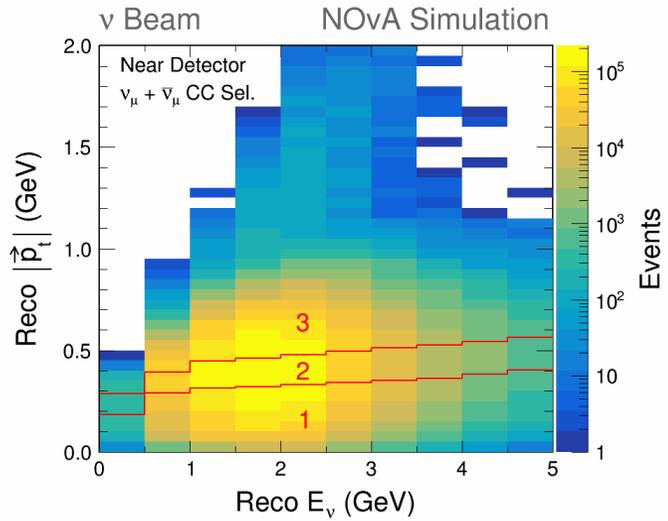


Acceptance correlated with lepton  $p_T$



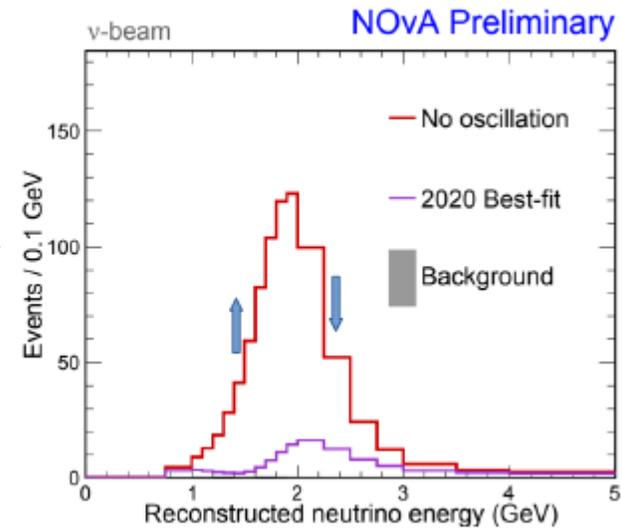
# Mitigating Acceptance Uncertainties

Subdivide samples by lepton  $p_T$  and extrapolate ND to FD



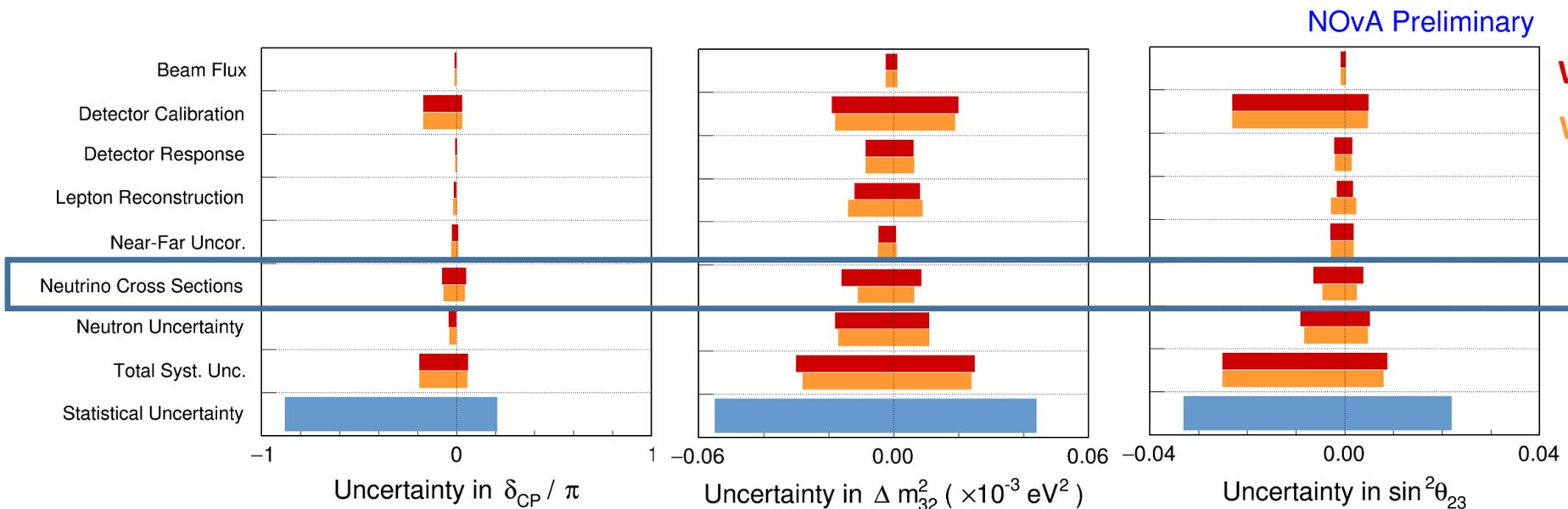
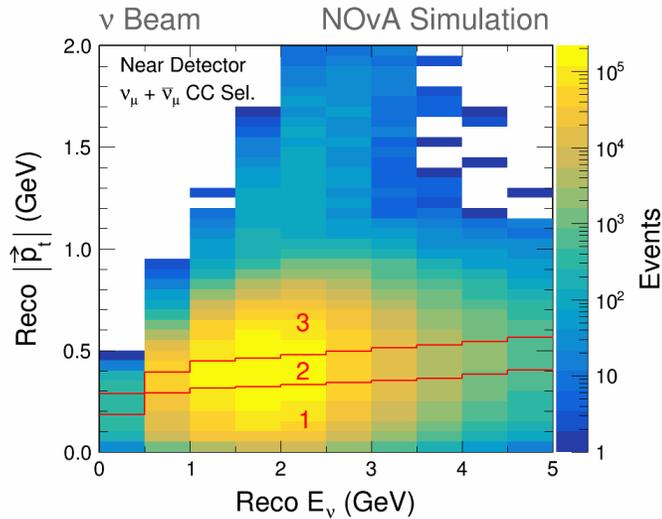
“Extrapolation”

(account for beam divergence, detector size difference, etc. w/ MC)



# Mitigating Acceptance Uncertainties

Subdivide samples by lepton  $p_T$  and extrapolate ND to FD



Without  $p_T$  extrap.

With  $p_T$  extrap.

Up to 30% reduction to impact of cross-section uncertainties on analyses

Stats Limited

Cross sections not dominant systematic

# Cross-section model

Genie 3.0.6

QE: Valencia 1p1h with Z-expansion axial form factor

MEC: Valencia reweighted to ND data

RES: Berger-Sehgal

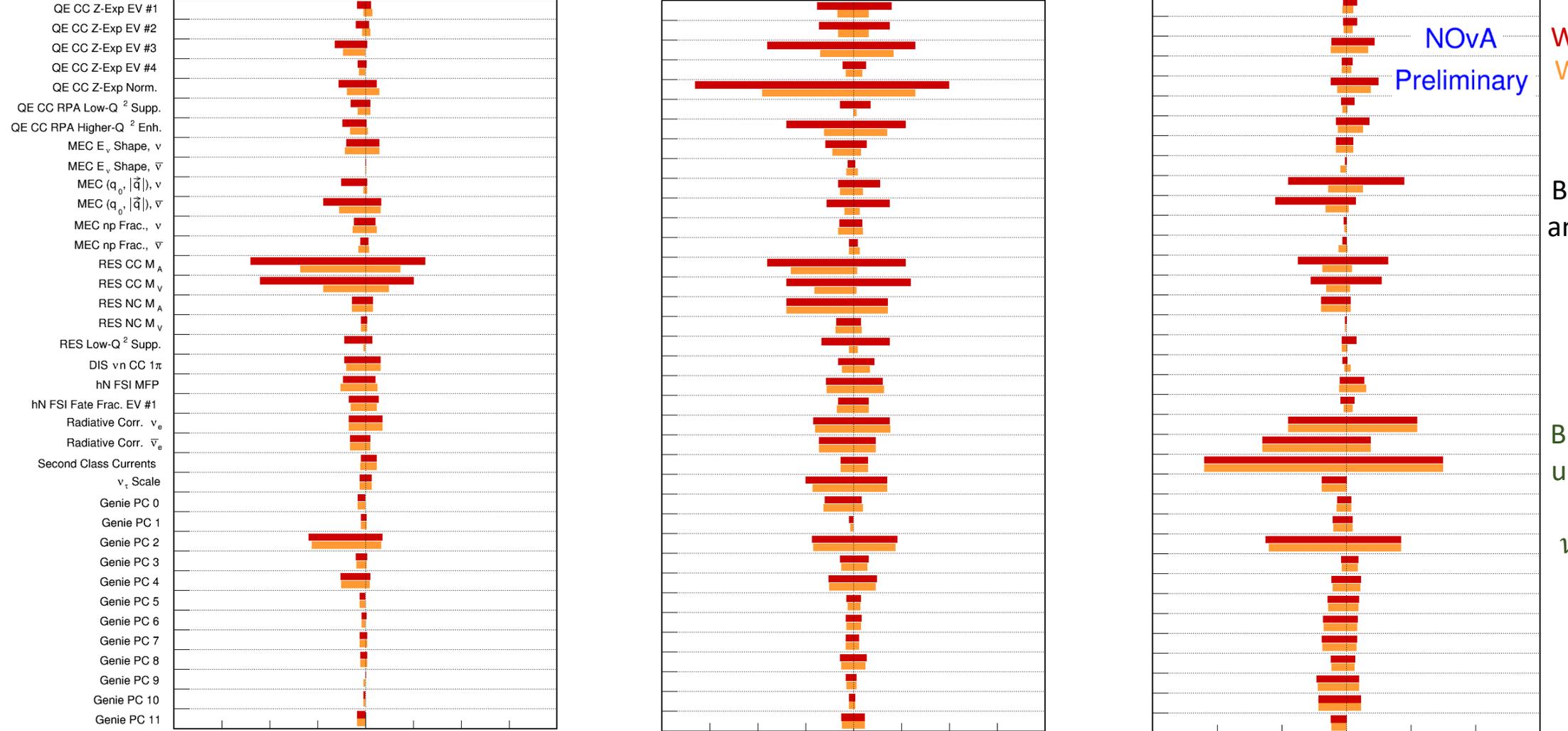
DIS: Bodek-Yang

FSI: hN (semi-classical cascade) reweighted to pion scattering data

# Impact Cross-sections – Joint $\nu_\mu + \nu_e$ Analysis

Constraint mainly from  
 $\nu_\mu$  Disappearance

Constraint mainly from  
 $\nu_e$  Appearance



Without  $p_T$  extrapol.  
With  $p_T$  extrapol.

Biggest cross-section analysis uncertainties affect relative normalization of QE and RES

Biggest cross-section uncertainties for  $\delta_{CP}$  related to  $\nu_e - \nu_\mu$  differences

-0.02      0.00      0.02

-0.004      0.000      0.004

-0.06      0.00      0.06

Uncertainty in  $\Delta m_{32}^2$  ( $\times 10^{-3}$  eV<sup>2</sup>)

Uncertainty in  $\sin^2\theta_{23}$

Uncertainty in  $\delta_{CP} / \pi$

# Summary

Cross-section uncertainties impact NOvA oscillation analyses through

1 - Event Yields

Primary cross-section effect

2 - Energy Estimation

Less dominant effect, mitigated by:

Event energy primarily being leptonic

Calorimetric response to charged-particles

NOvA energies correspond to a mix of production modes

Resonant production is leading cross-section uncertainty

$\nu_e$  appearance prediction depends on observed  $\nu_\mu$  events

$\nu_e - \nu_\mu$  cross-section differences important

NOvA analyses are statistics limited and systematic uncertainties are dominated by calorimetric and neutron systematics (not cross sections)