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# A novel ML approach for the reconstruction of particle showers with a tracking detector

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- We will present an approach using Convolutional Neural Networks for the reconstruction of particle showers using informations from a high-granularity tracking detector
- ML allows you go beyond calorimetry, also performing tracking
- We are developing this technique to improve the reconstruction of neutrinos at the SND@LHC experiment
- Being able to perform real-time calorimetry adds a lot to the SND@LHC physics case
- All of this is an ongoing exploratory work, for the time being.

# Introduction





Example of a particle shower





- SND@LHC is a newly approved experiment at the Large Hadron Collider (LHC)
- Its objective it is to study neutrinos of all flavours produced at the ATLAS interaction point, measuring their cross-sections in the GeV-TeV range for the first time
- Detector can also probe light dark matter scattering signatures
- The tracker of this detector is built at EPFL







SND@LHC's installation undergoing in TI18 tunnel









# SND@LHC detector layout

- Detector Layout: -target region: Emulsion walls (tungsten plates interleaved with nuclear emulsion films) combined with scintillating fibre (SciFi) tracking planes -Muon Identification system: iron plates interleaved with scintillating bars.
- Emulsion films are taken out for developing every few months while the SciFi layers allow real-time event analysis.



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<u>SND@LHC - Scattering and Neutrino Detector at the LHC</u>







# Motivation





- active layers of a sampling calorimeter
- The goal is to use the information of the SciFi tracker to perform prompt analysis
- Overview of the talk:
  - -Energy reconstruction in EM showers (feasibility study) -How to deal with ghost hits
  - -Energy reconstruction in case of neutrino scattering (EM + HAD showers at SND@LHC)
  - -Future Neutrino flavour tagging

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For real-time event analysis, emulsion walls act as passive materials and SciFi planes behaves as the

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- Shower energy resolution achievable by classical methods (counting hits) is ~22%
- It does not provide flavor tagging, nor use the topological information from the shower, nor from the muon detector
- Feasibility study convinced us to take this direction:
- Objective:

- Measure the energy of EM shower in the energy range 0-100 GeV.

Particle gun electron: - electrons shot on to the center of the first plane







Example of inputs: here the incoming particle is a 64 GeV electron.







- Procedure: SciFi hits (images) -> convolutional neural network (CNN) -> Energy
- Analysis of the detector response demonstrates that the target tracker planes behave as a sampling calorimeter.
- The CNN exhibits a resolution of 5% at E = 100 GeV and it is almost unbiased









Ghost hit problem

- Explain the pb!
- The architecture of the CNN was changed in order to use only the (x,z) and (y,z) projections of the simulated hits on the target tracker.
- This new architecture exhibits an average fractional energy resolution of 5.7% with PG data.
- keep the same resolution with less information!



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100

120

140

80 .

100 -

120

140

100

150

30000 events Linear Regression (Optimised CNN)











- Artificial Neural Networks algorithms are subset of ML
- They are comprised of one input layer, hidden layers, and an output layer.
- Each node of a layer is connected to all the nodes of the previous layer.
- Weights (->) are the parameters of the NN
- CNN are a type of Artificial Neural Networks originally designed for large number of inputs (pixels)
- The convolution operation allows to reduce drastically the number of NN parameters and to keep track of the spatial information between adjacent nodes.



**Deep neural network** Multiple hidden layers Input layeı



Image

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Convolved Feature













CERN



**CMBR Block** 

Loss functions:

$$smooth\_l1\_loss$$

$$l(x,y) = \sum_{n} z_{n}$$

$$l(x,y) = \begin{cases} \frac{0.5 (x_{n} - y_{n})^{2}}{\beta} & if |x_{n} - y_{n}| < \beta \\ |x_{n} - y_{n}| - 0.5\beta & otherwise \end{cases}$$

translation

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 Our architecture is composed of n CMBR blocks where n should be proportional to the number of inputs.

A CMBR block is a succession of 4 different operations: -CoordConv= Convolution operation but invariant under

-MaxPool = take maximum value to reduce dimentionality -BatchNorm = normalisation to avoid very large value -ReLu = Removing negative values to increases the nonlinear properties of the loss function







Objective:

- Measure the energy of an EM shower induced by a neutrino in the energy range 100-5000 GeV (expected at SND@LHC)

Inputs:

 $-\nu_{\rho}$  neutrino with elastic scattering and charged-current deep inelastic scattering

- no simulation of readout electronics;
- hits are defined as a Yes/No signal in each pixel (no amplitude information)
- true (X, Y) positions of the simulated hits



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Example of inputs: here the incoming particle is a 767GeV  $\nu_{e}$ . The shower induced by this particle produced hits (white dots) on the 5 SciFi planes



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- Bias of ~9 GeV compared to 100-5000 GeV range
- Energy resolution around 500 GeV
- Degradation in performance is due to: -Shower generated at any depth in the detector -Shower sampling goes from  $10X_0$  to  $15X_0$ -> only 2 planes have a significant amount of hits on average -Analysis has to be improved
- We separate the data samples (elastic scattering vs) charged-current deep inelastic scattering) during the test of the CNN accuracy
- Elastic events seems to be better reconstructed

# Neutrino energy reconstruction













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- Objective: -as a first step towards flavour tagging we try to performed classification of elastic and inelastic events
- The spatial distribution of hits can be used as a way of discriminating between elastic and inelastic scattering events.
- The CNN used for the first study was modified to provide output label probabilities rather than a predicted energy value.



# Classification of elastic and inelastic events











- for the event to be elastic or inelastic predicted by the CNN
- After training, the prediction accuracy was found to be 94.5%



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This time the loss function will evaluate the difference between the true label, and the probability









- Add information from the scintillating pads of the Muon detector
   –> greatly improve efficiency for not fully contained showers
- Study effect of non-binary SciFi channels response
- Move on to full detector simulation and higher stat samples
- Optimize network structure & parameters to achieve best energy resolution
- Add the possibility of adding tracking for isolated tracks (muons) to aid flavour tagging
- SND@LHC finishing installation -> Use in the LHC Run 3





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- Using CNN algorithms for SND@LHC prompt analysis has many advantages:
  - once trained, it is fast and straightforward to compute the energy
  - It is flexible, as the same architecture can be used for multiple tasks (type of interaction, flavour tagging)
  - It can be very accurate
- Allow to perform real time physic
- It is very greedy: we need feed it a lot with labelled data to achieve acceptable accuracy. For this, an accurate description of the detector geometry and digitalisation of the detector signal is essential.
- It would be a crucial asset for particle detector operating at a high luminosity collider or beam dump
- In case of SND@LHC, after HL-LHC upgrade, emulsions cannot operate and a ML approach would become even more important
- New approach to neutrino physics with a tracking detector









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# <mark>감사해요</mark> (Thank you)





# BACKUP









## Muon upstream

# 5 planes of 82x61 cm made of 10 horizontal bars

3 planes of 82x61cm made of 2 layers (X and Y) of 60 bars each



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## Muon downstream

## SciFi tracker

5 planes of 41x40cm made of 2 layers (X and Y) with a  $250 \mu m$ resolution





planes (smaller dimension of the input images)

# Advantages:

- -It save a lot of computing time
- Optimal size of the sub-image calculated using whole sample: Take 3 sigma around the barycenter of the hits of the planes



•To find the optimal CNN architecture, we worked on a simplified dataset with smaller SciFi

# -That avoid polluting the CNN with extra hits so that the shower features stand out better



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# Downsizing process













- Objective:
  - Measure the energy of an electronic shower in the energy range 0-100 GeV.
- Detector geometry used:
  - 4 planes of 41.6x38.7 cm.

  - Tungsten bricks of 5.6cm ( $10X_0$  radiation lengths). - granularity implemented corresponds to the fibre diameter (250 $\mu$ m)
- Particle gun electron: - shot before the first plane though the center with a random angle within plus/minus 10°.
- Input information for the reconstruction: - (X, Y) truth positions of the hits used (from MC)
  - no simulation of readout electronics
  - hits are defined as a greyscale signal (amplitude of the hit proportional to the numb of hits that fall into the plane pixel).





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- Detector geometry:
  - 5 planes of 41x40 cm.

  - separated by tungsten bricks of 7.8cm (15X0 radiation lengths). - granularity implemented correspond to the resolution of the SciFi ( $250 \mu m$ )
- Neutrino:
  - produced with Genie using the NuEElastic and CCDIS settings
  - energy range 100-5000 GeV
  - shotted in accordance with the alignment of the IP (off center of the plane) with a random angle.
  - shower starting point is spread uniformly over the detector length
- Input information for the reconstruction: - no simulation of readout electronics;
  - hits are defined as a Yes/No signal in each pixel (no amplitude information is used) - (X, Y) truth positions of the hits used (from MC)

