Particle Detectors Principles

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The 'Real' World of Particles

E. Wigner:

"A particle is an irreducible representation of the inhomogeneous Lorentz group" Spin=0,1/2,1,3/2 ... Mass>0

Annals of Mathematics Vol. 40, No. 1, January, 1939

ON UNITARY REPRESENTATIONS OF THE INHOMOGENEOUS LORENTZ GROUP*

By E. WIGNER (Received December 22, 1937)

1. Origin and Characterization of the Problem

It is perhaps the most fundamental principle of Quantum Mechanics that the system of states forms a linear manifold, in which a unitary scalar product is defined. The states are generally represented by wave functions in such a way that φ and constant multiples of φ represent the same physical state. It is possible, therefore, to normalize the wave function, i.e., to multiply it by a constant factor such that its scalar product with itself becomes 1. Then, only a constant factor of modulus 1, the so-called phase, will be left undetermined in the wave function. The linear character of the wave function is called the superposition principle. The square of the modulus of the unitary scalar product (ψ, φ) of two normalized wave functions ψ and φ is called the transition probability from the state ψ into φ , or conversely. This is supposed to give the probability that an experiment performed on a system in the state φ , to see whether or not the state is ψ , gives the result that it is ψ . If there are two or more different experiments to decide this (e.g., essentially the same experiment,

The 'Real' World of Particles

W. Riegler:

"...a particle is an object that interacts with your detector such that you can follow it's track,

it interacts also in your readout electronics and will break it after some time,

and if you a silly enough to stand in an intense particle beam for some time you will be dead ..."

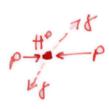
The 'Real' World of Particles

Elektro-Weak Lagrangian

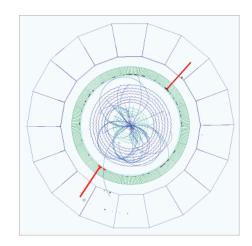
$$\begin{split} L_{GSW} &= L_0 + L_H + \sum_l \left\{ \frac{g}{2} \, \overline{L}_l \gamma_\mu \, \overline{\tau} L_l \, \overline{A}^\mu + g' \bigg[\, \overline{R}_l \gamma_\mu R_l + \frac{1}{2} \, \overline{L}_l \gamma_\mu L_l \, \bigg] B^\mu \right\} + \\ &+ \frac{g}{2} \sum_q \, \overline{L}_q \gamma_\mu \, \overline{\tau} L_q \, \overline{A}^\mu + \\ &+ g' \bigg\{ \frac{1}{6} \sum_q \, \bigg[\overline{L}_q \gamma_\mu L_q + 4 \, \overline{R}_q \gamma_\mu R_q \, \bigg] + \frac{1}{3} \sum_{q'} \, \overline{R}_{q'} \gamma_\mu R_{q'} \bigg\} B^\mu \end{split}$$

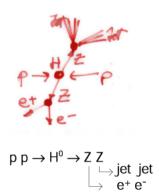
$$L_{H} = \frac{1}{2} (\partial_{\mu} H)^{2} - m_{H}^{2} H^{2} - h \lambda H^{3} - \frac{h}{4} H^{4} + \frac{g^{2}}{4} (W_{\mu}^{+} W^{\mu} + \frac{1}{2 \cos^{2} \theta_{W}} Z_{\mu} Z^{\mu}) (\lambda^{2} + 2\lambda H + H^{2}) + \sum_{l,q,q'} (\frac{m_{l}}{\lambda} \bar{l} l + \frac{m_{q}}{\lambda} \bar{q} q + \frac{m_{q'}}{\lambda} \bar{q}' q') H$$

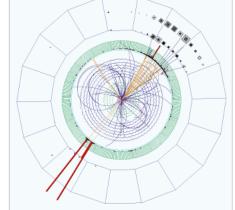
Higgs Particle



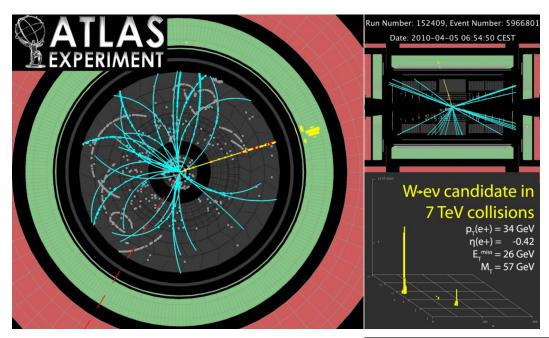
$$p \ p \to H^0 \\ \hookrightarrow \gamma \gamma$$

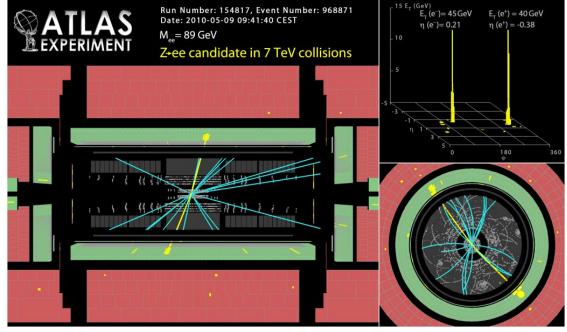




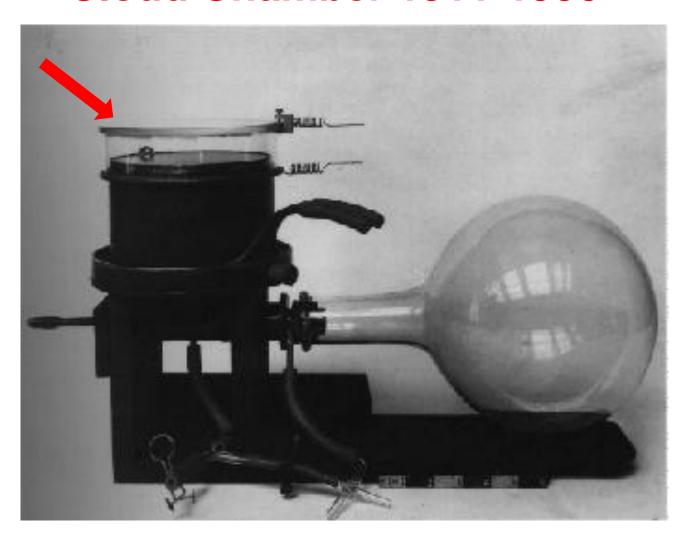


2010 ATLAS W, Z particles





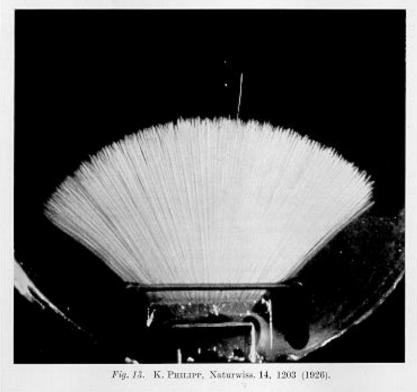
Cloud Chamber 1911-1950



Wilson Cloud Chamber 1911

Cloud Chamber

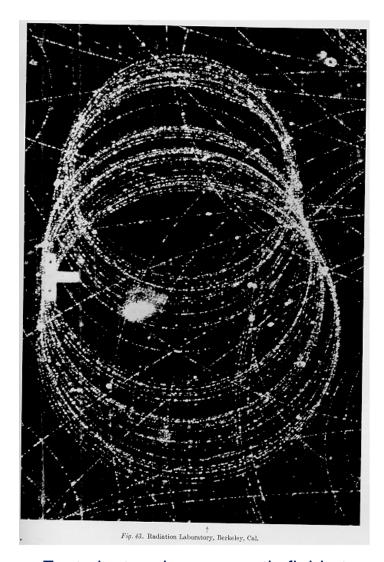
Alpha Particles, Philipp 1926



Charges particles ionize the 'air' along their track.

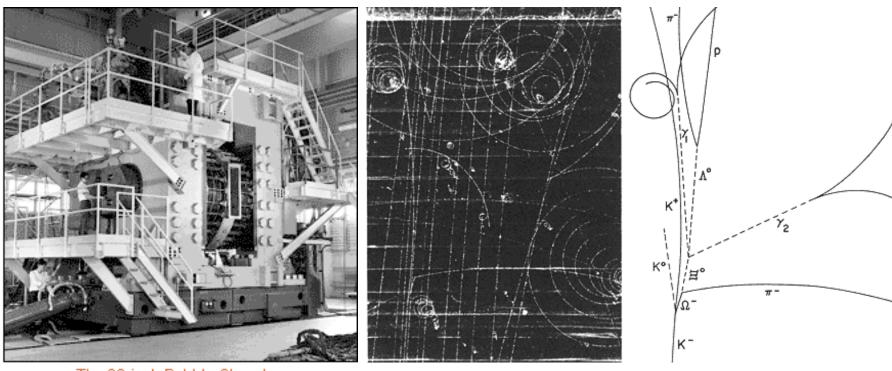
These charges will act as condensation points if the gas is supersaturated → photos of particle tracks.

The alpha particles in this picture lose energy – the fact that they all have the same range shows that all alphas have the same original energy.



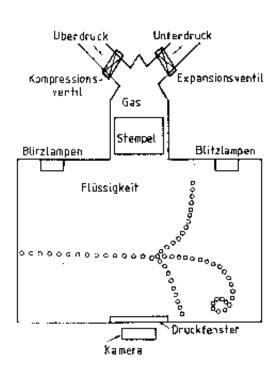
Fast electron in a magnetic field at the Bevatron, 1940

Bubble Chamber (1950-1984)



The 80-inch Bubble Chamber

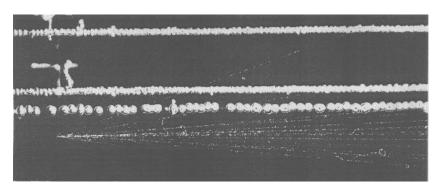
Bubble Chamber

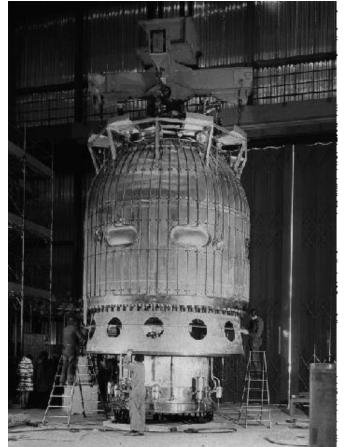




A superheated liquid starts boiling when a charged particle passes the liquid and transfers energy to the atoms of the liquid.

Small bubbles → photography



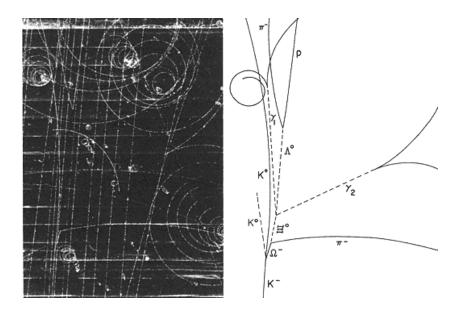


Bubble Chambers

The excellent position ($5\mu m$) resolution makes the Bubble chamber almost unbeatable for reconstruction of complex decay modes.

The drawback of the bubble chamber is the low rate capability (about 10/second). However, LHC has 10⁹ collisions/second!

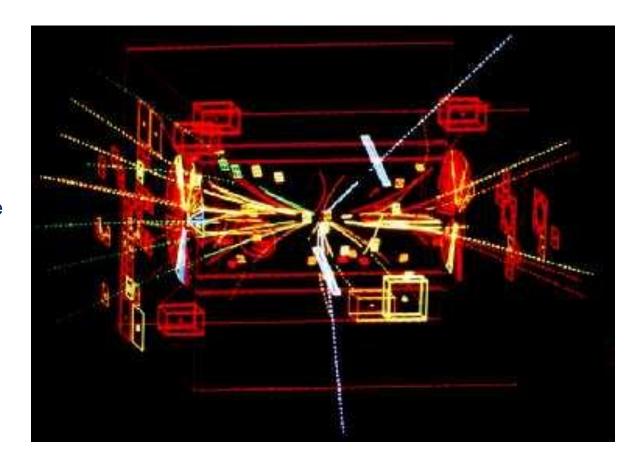
That's why electronics detectors took over in the 70ties.



W, Z-Discovery 1983/84

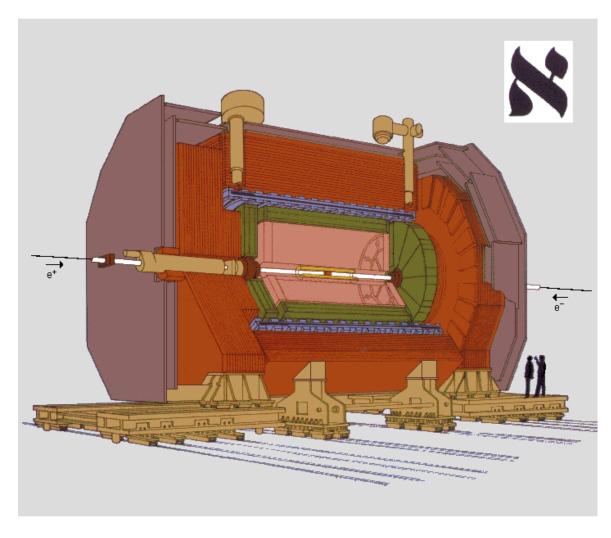
UA1 used a very large wire chamber.

Can now be seen in the CERN Microcosm Exhibition



This computer reconstruction shows the tracks of charged particles from the proton-antiproton collision. The two white tracks reveal the Z's decay. They are the tracks of a high-energy electron and positron.

LEP 1988-2000



- Vertex Detector
- Inner Tracking Chamber
- Time Projection Chamber
- Electromagnetic Calorimeter
- Superconducting Magnet Coil
- Hadron Calorimeter
- Muon Chambers
- Luminosity Monitors

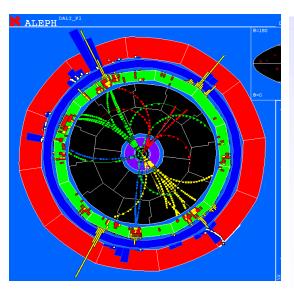
The ALEPH Detector
All Gas Detectors

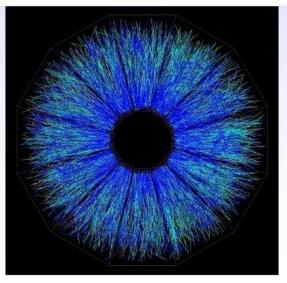
Increasing Multiplicities in Heavy Ion Collisions

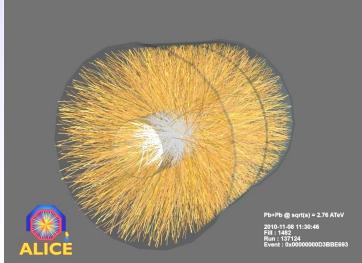
e+ e- collision in the ALEPH Experiment/LEP.

Au+ Au+ collision in the STAR Experiment/RHIC Up to 2000 tracks

Pb+ Pb+ collision in the ALICE Experiment/LHC
Up to 10 000 tracks/collision

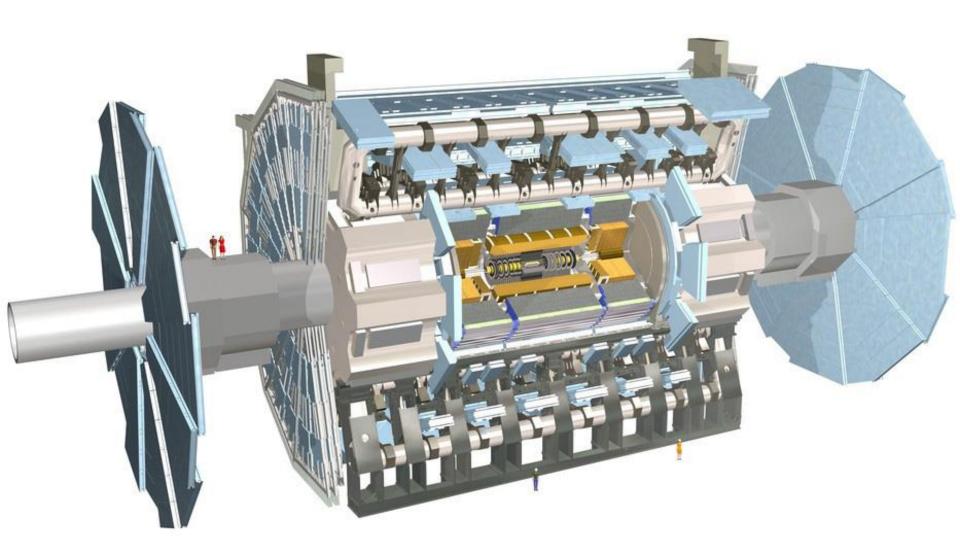






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ATLAS at LHC



The ATLAS detector uses more than 100 million detector channels. Measuring 1 billion collisions per second.

$$E = Ma^{2}$$

$$E = Mb^{2}$$

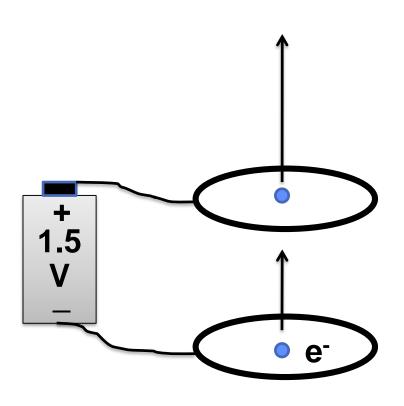
$$E = mc^{2} = Energy = Mess$$

$$\vdots$$

$$E = e_o \cdot 1 V$$

1 Electron Volt - Evergy on Electron goins as it traverses a Polatical Difference of 1V

Build your own Accelerator



$$E_{kin}$$
= 1.5eV =

2 615 596 km/h



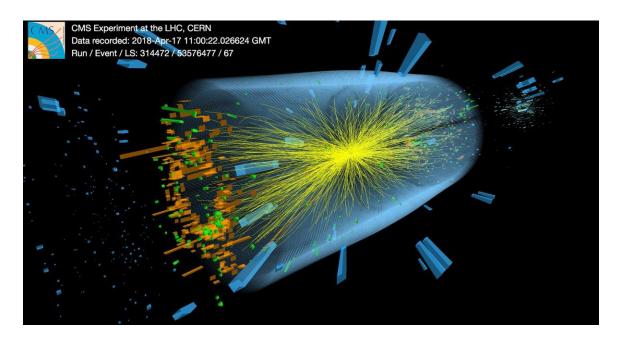
Visible Light: 2=500mm, hv ~2.5 eV

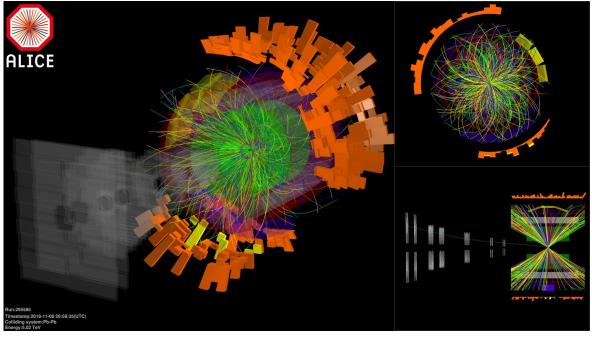
Exciled Shobs in Alons: 1-100 keV "X-Rays"

Nuclear Physics: 1-50 MeV

Particle Physics: 1-1000 GeV (LHC 14 TeV)

Higher Measures Energy: 10 20 eV (Casnic Roys)

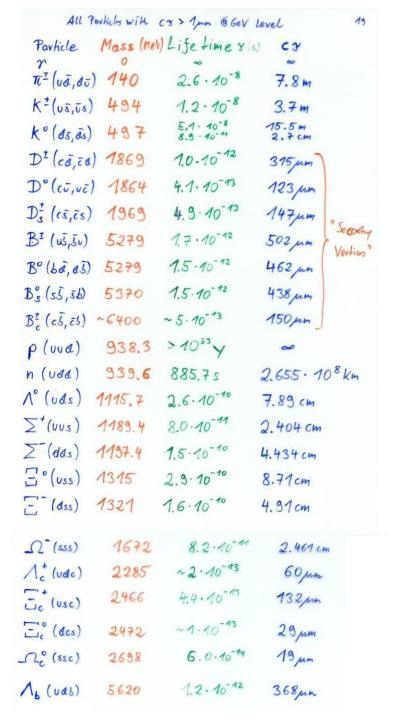




7, W , Z, g, e, M, 3, Ve, Vm, Yz, Tt, To, y, fo (660), g(20), w (782), y' (858), fo (380), Qo (380), \$\phi(1020), h_1(1170), b_1(1235), a, (1260), f2 (1270), f, (1285), y (1295), T (1300), a2 (1320), 10 (1370), 1, (1420), w (1420), y (1440), a, (1450), g (1450), 10 (1500), 12 (1525), w (1650), W3 (1670), TC2 (1670), \$ (1680), 93 (1690), 9 (1700), fo (1710), TT (1800), \$ (1850), \$ (2010), a4 (2040), 14 (2050), 12 (2300), 12 (2340), Kt, Ko, Ko, Ko, Ko (892), K. (1270), K. (1400), K* (1410), K. (1430), K. (1430), K. (1680). K, (1770), K, (1780), K, (1820), K, (2045), Dt. Do, D' (2007), D" (2010) , D, (2420), D," (2460), D," (2460) , D, D, D, Ds. (2536) 1, Ds. (2573) 1, B1, B0, B4, B0, B1, Me (15), J/4(15), X (1P), X (1P), X (1P), W (25), Y (3770), W (4040), Y (4160), V (4415), Y (15), X (1P), X (1P), X (1P), Y (25), X (2P), X52 (2P), T (3S), T (4S), T (10860), T (11020), p, n, N (1440), N (1520), N (1535), N (1650), N (1675), N (1680), N (1700), N (1710), N (1720), N (2130), N (2220), N (2250), N (2600), A (1232), A (1600), A (1620), A (1700), A (1905), A (1910), A (1920), A (1930), A (1950), $\Delta(2420)$, Λ , $\Lambda(1405)$, $\Lambda(1520)$, $\Lambda(1600)$, $\Lambda(1670)$, $\Lambda(1690)$, Λ (1800), Λ (1810), Λ (1820), Λ (1830), Λ (1890), Λ (2100), Λ (2110), Λ (2350), Σ^{+} , Σ° , Σ^{-} , Σ (1385), Σ (1660), Σ (1670), $\sum (1750), \sum (1775), \sum (1915), \sum (1940), \sum (2030), \sum (2250), \equiv 0, \equiv 0, = 0$ \equiv (1530), \equiv (1690), \equiv (1820), \equiv (1950), \equiv (2030), Ω , Ω (2250), $\Lambda_{c}^{+}, \Lambda_{c}^{+}, \Sigma_{c}(2455), \Sigma_{c}(2520), \Xi_{c}^{+}, \Xi_{c}^{0}, \Xi_{c}^{+}, \Xi_{c}^{0}, \Xi_{c}$ Ξc(2780), Ξc(2815), Ωc, Λ, Ξ, Ξ, Ξ, tt

There are Many move

All particles that can possibly leave a track in the detector:



From the 'hundreds' of Particles listed by the PDG there are only ~27 with a life time cs > ~ 1 pm i.e. they can be seen as 'tracks' in a Detector.

~ 13 of the 27 have cs < 500 pm i.e. only mm range at GeV Energies.

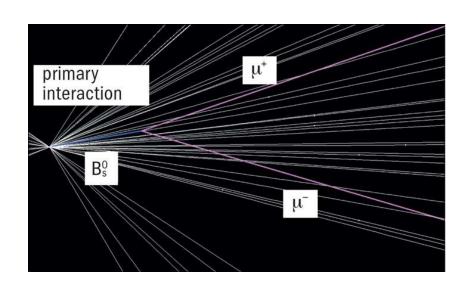
→ "short" Ivochs measured with Emulsions or Verlex Detectors.

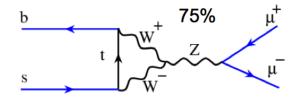
From the ~14 remaining posticles e^{\pm} , μ^{\pm} , γ , π^{\pm} , K° , p^{\pm} , n

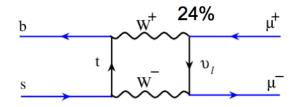
are by far the most frequent ones

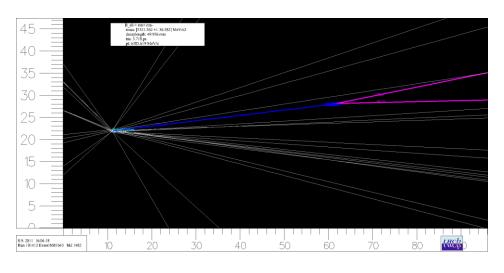
A porticle Delector null be able to identify and measure Energy and Momenta of Hese 8 porticles.

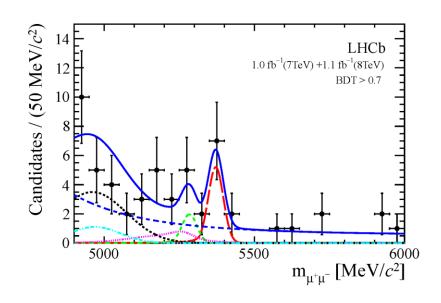
LHCb B decay, displaced Vertex











$$e^{\pm}$$
 $m_{e} = 0.511 \, \text{MeV}$
 m^{\pm} $m_{m} = 105.7 \, \text{MeV} \sim 200 \, \text{me}$
 γ $m_{T} = 0$, $Q = 0$
 π^{\pm} $m_{\pi} = 139.6 \, \text{MeV} \sim 270 \, \text{me}$
 k^{\pm} $m_{k} = 493.7 \, \text{MeV} \sim 1000 \, \text{me}$
 p^{\pm} $m_{p} = 938.3 \, \text{MeV} \sim 2000 \, \text{me}$
 $m_{k} = 497.7 \, \text{MeV} = 0$
 $m_{k} = 939.6 \, \text{MeV} = 0$

The Difference in

Mass, Charge, Interaction

is the key to the Identification

Bosics

Invariant Mass:

$$LAB$$
:
$$\frac{m_1, \vec{p}_1, \vec{p}_2}{m_2, \vec{p}_2, \vec{p}_2} = \frac{m_2, \vec{p}_2, \vec{p}_2}{m_2, \vec{p}_2, \vec{p}_2}$$

Reblivity:
$$\tilde{\alpha} = \begin{pmatrix} a_o \\ \tilde{a} \end{pmatrix}$$
 $\hat{b} = \begin{pmatrix} b_o \\ \tilde{b} \end{pmatrix}$ $\hat{a}\hat{b} = a_o b_o - \bar{a}\bar{b}$

$$E = mc^2 \gamma , \ \vec{p} = m\vec{v} \gamma$$

$$\tilde{p} = \begin{pmatrix} E \\ \tilde{c} \\ \tilde{p} \end{pmatrix}, \ \tilde{p}_A = \begin{pmatrix} E_A \\ \tilde{c} \\ \tilde{p}_A \end{pmatrix}, \ \tilde{p}_z = \begin{pmatrix} E_Z \\ \tilde{c} \\ \tilde{p}_z \end{pmatrix}$$

$$\tilde{p} = \tilde{p}_A + \tilde{p}_Z \quad Exergy + Monelon Conservation$$

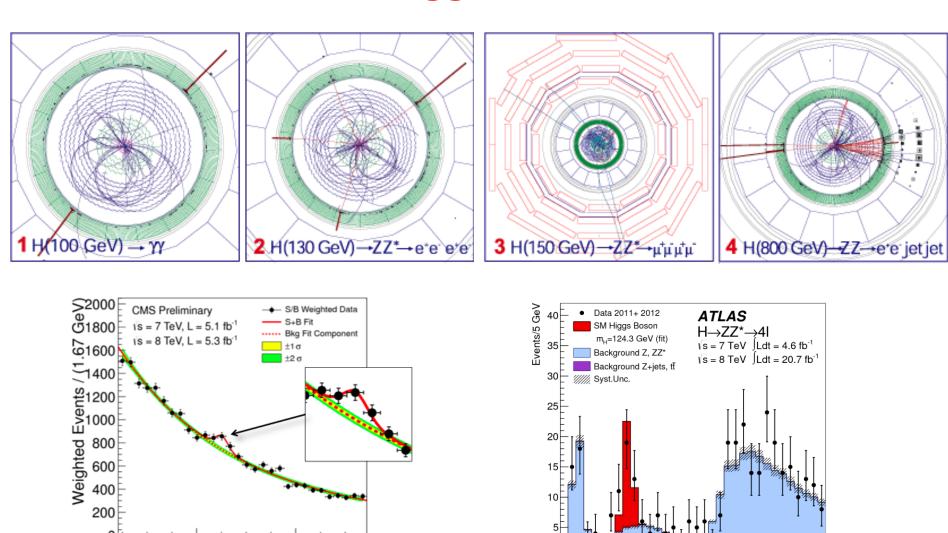
$$\tilde{p}^{2} = (\tilde{p}_{A} + \tilde{p}_{z})^{2} \rightarrow \tilde{p} \tilde{p} = \tilde{p}_{A} \tilde{p}_{A} + \tilde{p}_{z} \tilde{p}_{z} + 2 \tilde{p}_{z} \tilde{p}_{z}$$

$$M^{2} = (\tilde{p}_{A} + \tilde{p}_{z})^{2} \rightarrow \tilde{p} \tilde{p} = \tilde{p}_{A} \tilde{p}_{A} + \tilde{p}_{z} \tilde{p}_{z} + 2 \tilde{p}_{z} \tilde{p}_{z}$$

$$M^2c^2 = m_1^2c^2 + m_2^2c^2 + 2\left(\frac{E_1E_2}{c^2} - \rho_1\rho_2 \cos\theta\right)$$

- · Measuring Momenta ont Energies OR
- · Measuring Momenta and identifying Porticles gives the Mass of the original Particle

Simulated Higgs Boson at CMS



Particle seen as an excess of two photon events above the irreducible background.

100

150

120

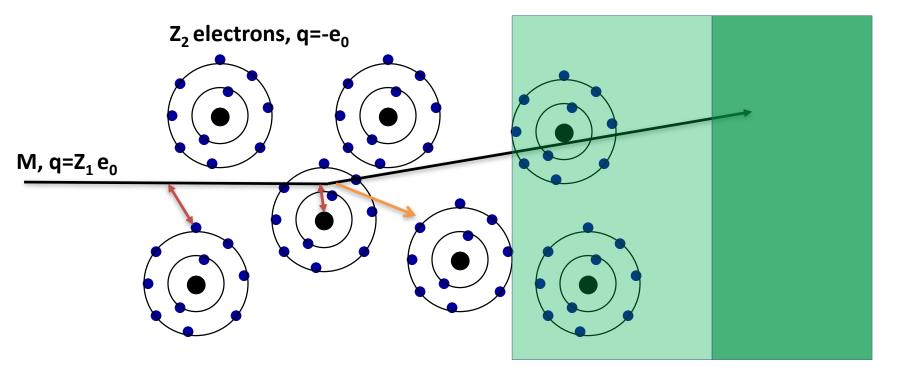
140

 $m_{\gamma\gamma}$ (GeV)

250 m₄₁ [GeV]

200

Electromagnetic Interaction of Particles with Matter

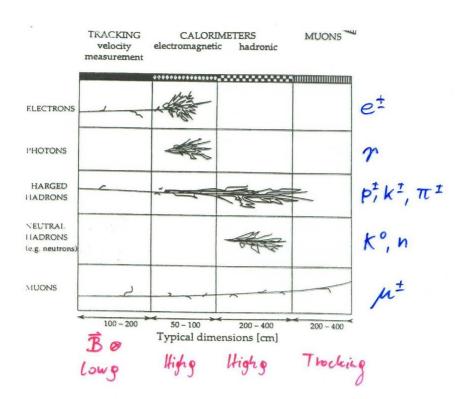


Interaction with the atomic electrons. The incoming particle loses energy and the atoms are excited or ionized.

Interaction with the atomic nucleus. The particle is deflected (scattered) causing multiple scattering of the particle in the material. During this scattering a Bremsstrahlung photon can be emitted.

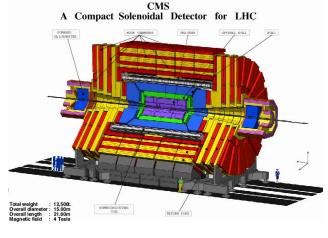
In case the particle's velocity is larger than the velocity of light in the medium, the resulting EM shockwave manifests itself as <u>Cherenkov Radiation</u>. When the particle crosses the boundary between two media, there is a probability of the order of 1% to produced and X ray photon, called Transition radiation.

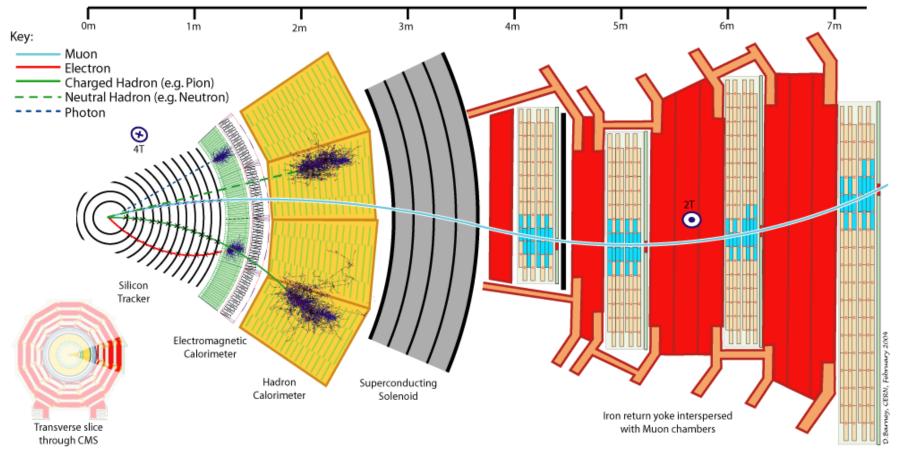
Task of a Particle Detector

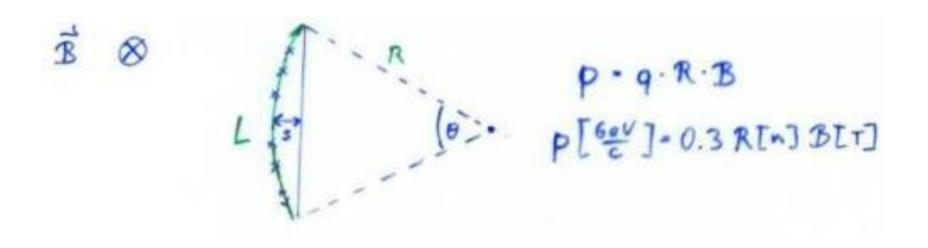


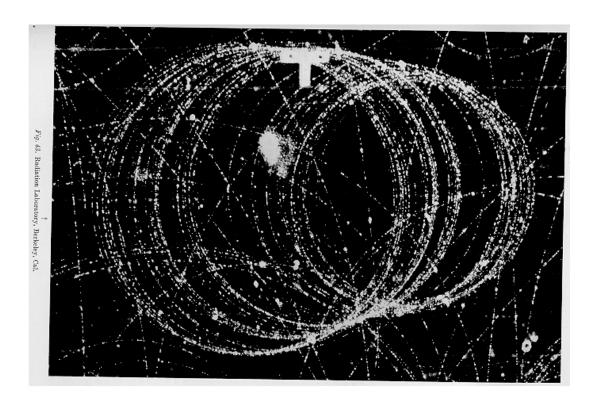
- · Electrons ionite and show Bremsstrahly ove to the small mass
- · Photons don't ionize but show Peir Production in high & Makerial. From Ken on equal to ex
- · Charged Hodrons ionite and show Hadron Shower in derse polerial.
- · Neutral Hodrors don't ionize and show Hadron Shower in Bense Moderial
- · Myons ionix and don't shower

Electron, Muon, Photon Pion, Charged Kaon, Proton Neutral Kaon, Neutron









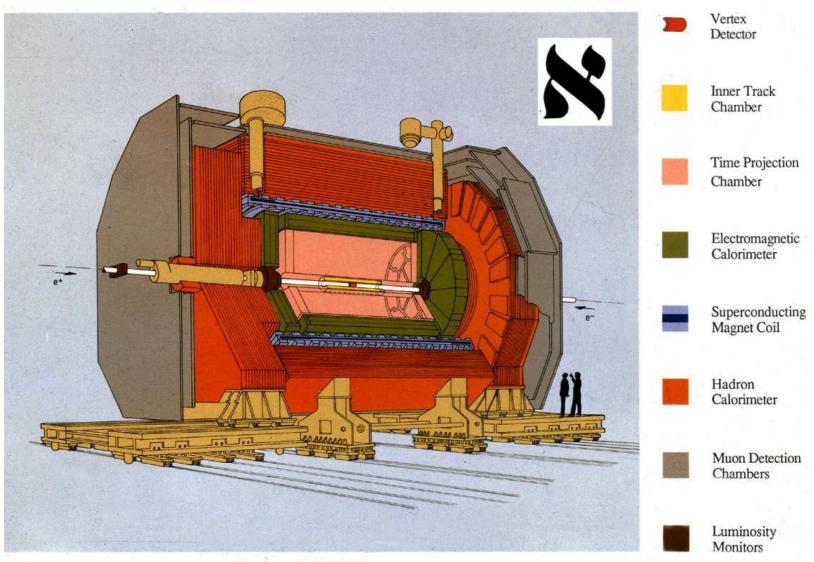
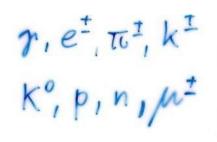
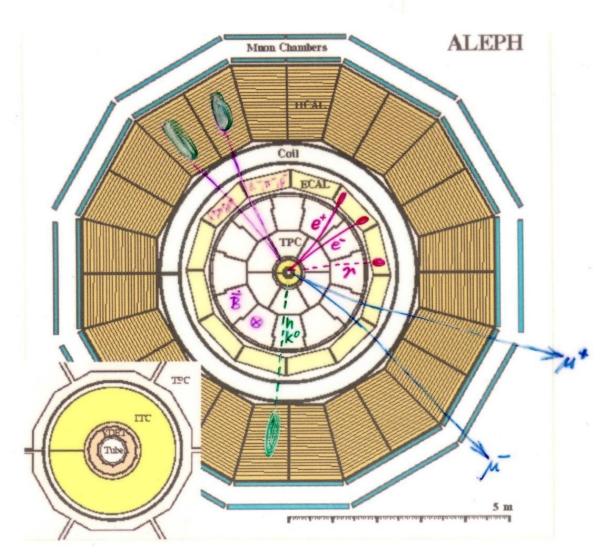


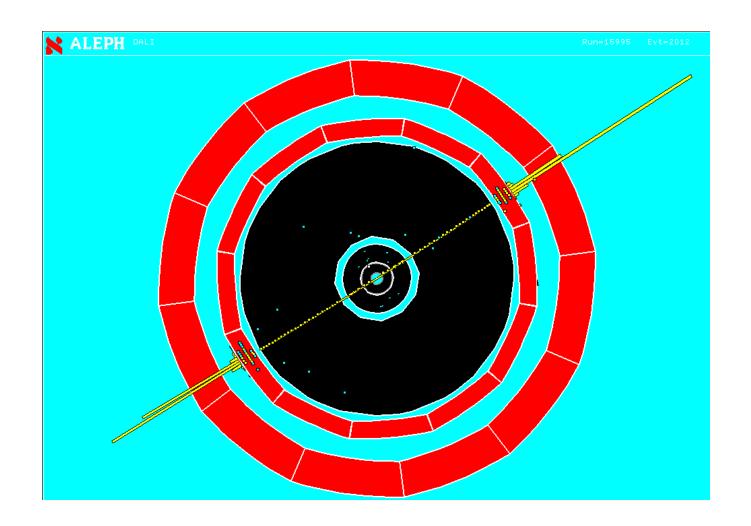
Fig. 1 - The ALEPH Detector





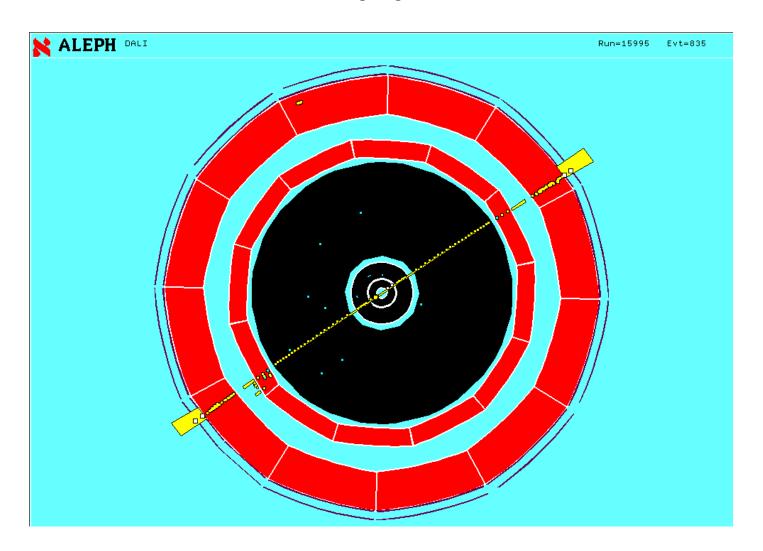
 $Z \rightarrow e^+ e^-$

Two high momentum charged particles depositing energy in the Electro Magnetic Calorimeter

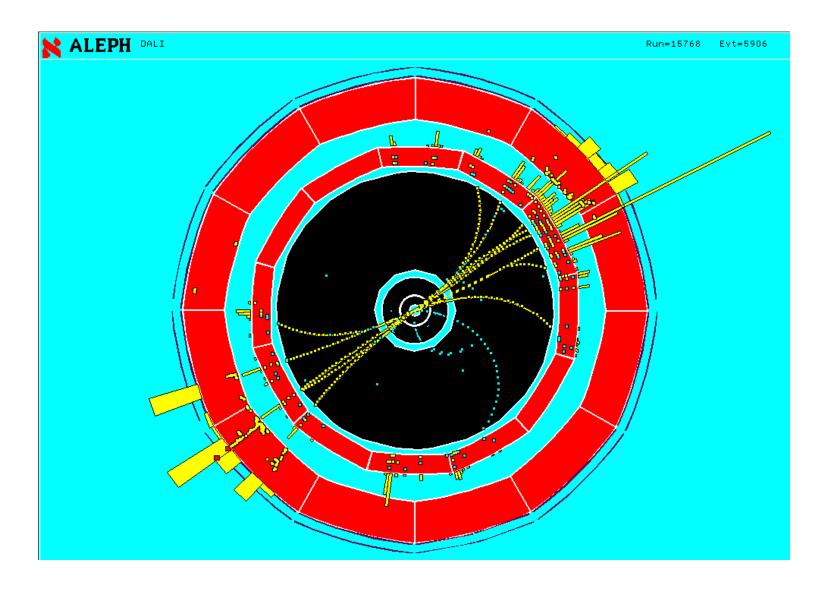


 $Z \rightarrow \mu^+ \mu^-$

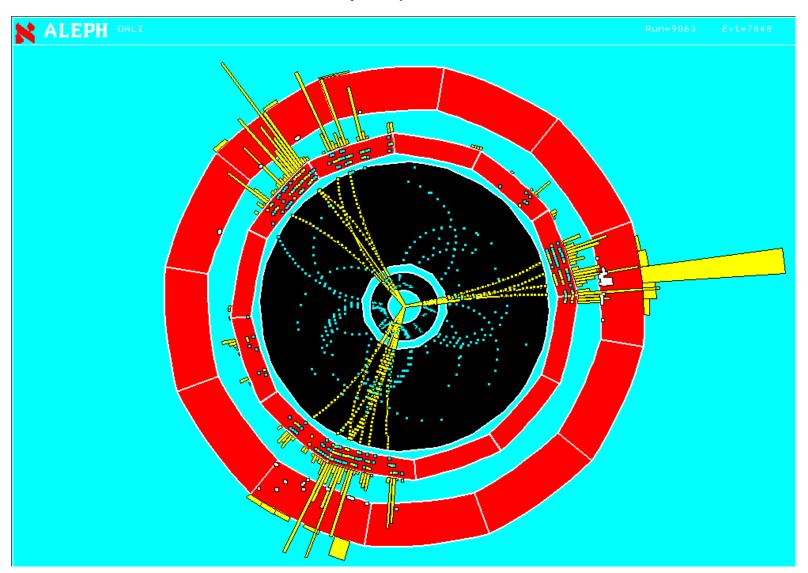
Two high momentum charged particles traversing all calorimeters and leaving a signal in the muon chambers.



 $Z \rightarrow q \overline{q}$ Two jets of particles



 $Z \rightarrow q \overline{q} g$ Three jets of particles

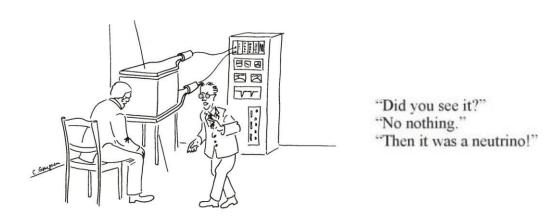


Interaction of Particles with Matter

Any device that is to detect a particle must interact with it in some way \rightarrow almost ...

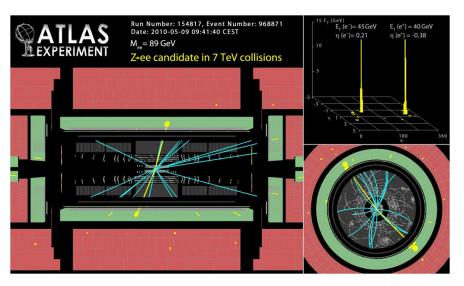
In many experiments neutrinos are measured by missing transverse momentum.

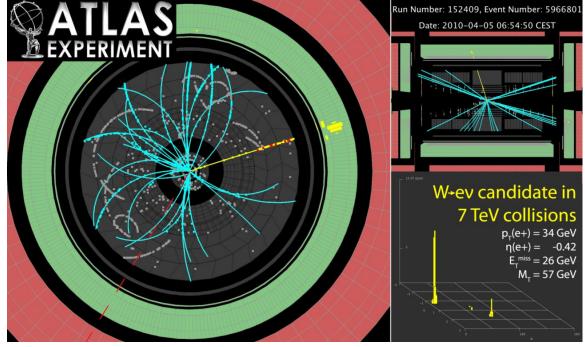
E.g. e^+e^- collider. $P_{tot}=0$, If the Σ p_i of all collision products is $\neq 0 \rightarrow$ neutrino escaped.

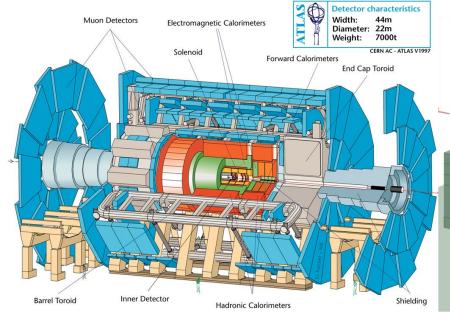


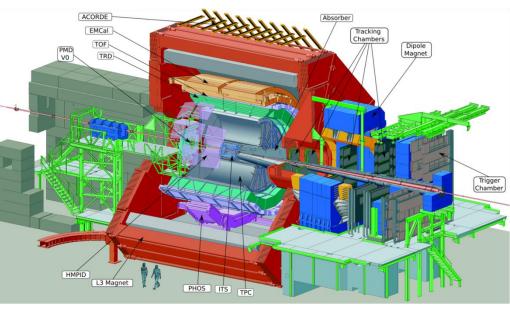
Claus Grupen, Particle Detectors, Cambridge University Press, Cambridge 1996 (455 pp. ISBN 0-521-55216-8)

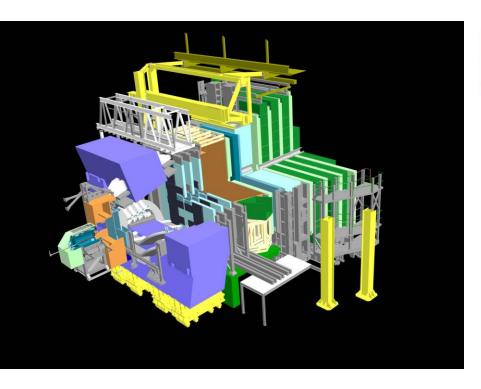
ATLAS W, Z particles!

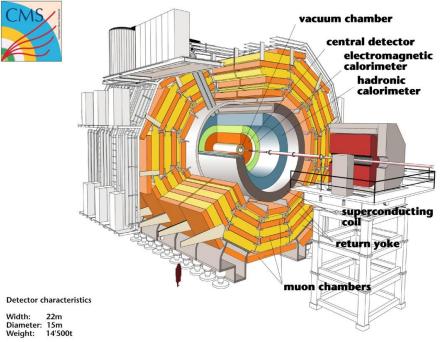












Tracker:

Charged particles are bent in the magnetic field.

Positive and negative particles are bent in opposite directions.

The bending radius measures the particle's momentum.

The task of the tracker is to measure the curved particle track.

When particles traverse material they are deflected, which reduces the accuracy of the momentum measurement.

The beampipe and the tracker must be built with as little (thin) material as possible !!

Neutral particles are not bent in the magnetic field and do not leave a trace in the tracker.

Calorimeter:

The calorimeter fully absorbs a particle and measures it's energy.

Calorimeters are built from very heavy materials in order to absorb the particles in the shortest possible distance.

Electrons and photons are absorbed over very short distances due to there small mass \rightarrow Electromagnetic Calorimeter.

The hadrons need much more material thickness to be absorbed → Hadron Calorimeter.

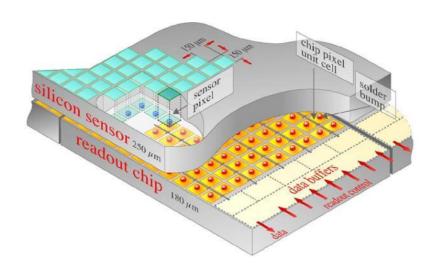
Muon system

Muons are the only particles that manage to go through the hadron calorimeter.

ATLAS Silicon Pixel Detector



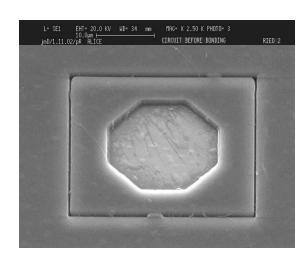
Silicon Pixel Detectors



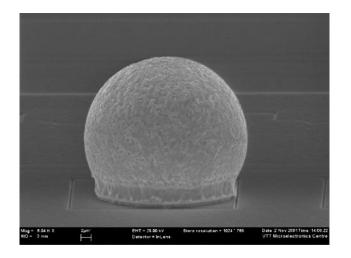
ATLAS: 1.4x10⁸ pixels

40 000 000 'images' per second.

W. Riegler/CERN



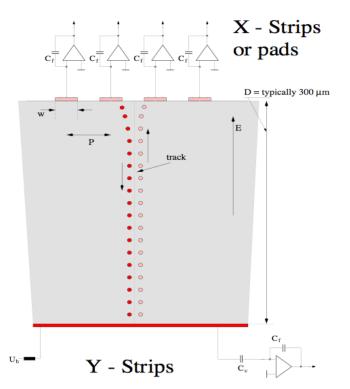
41



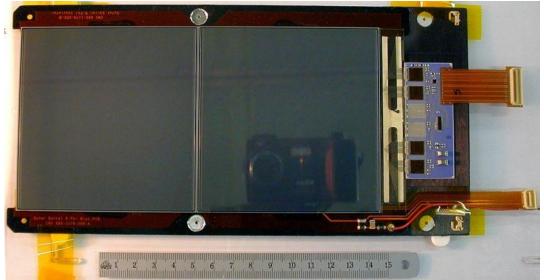
Silicon Strip Detectors

Every electrode is connected to an amplifier → Highly integrated readout electronics.

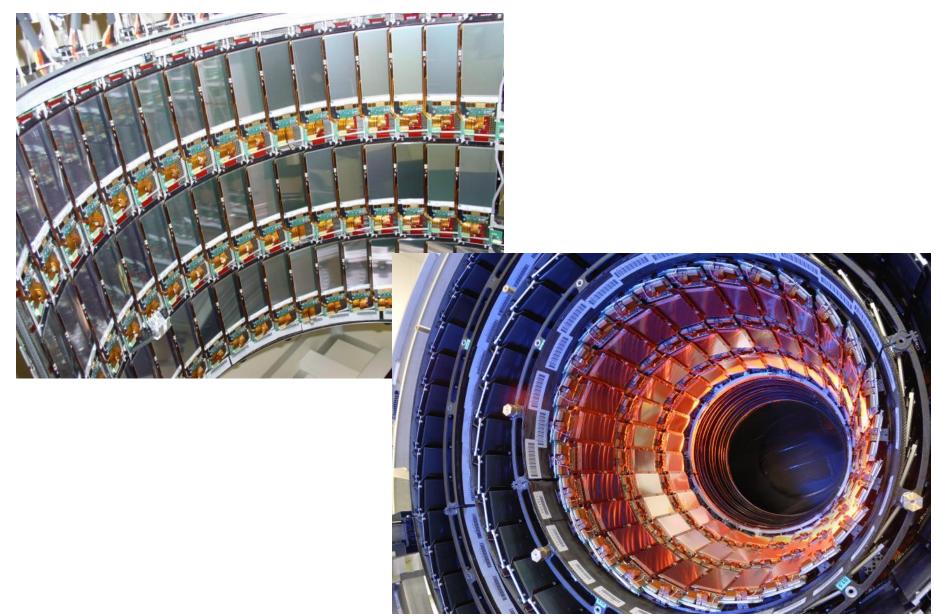
Two dimensional readout is possible.



CMS Outer Barrel Module



Silicon Strip Detectors

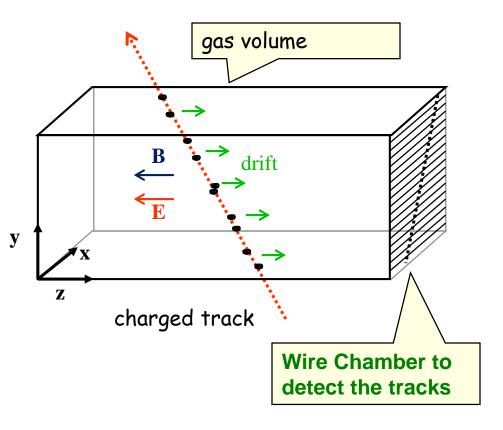


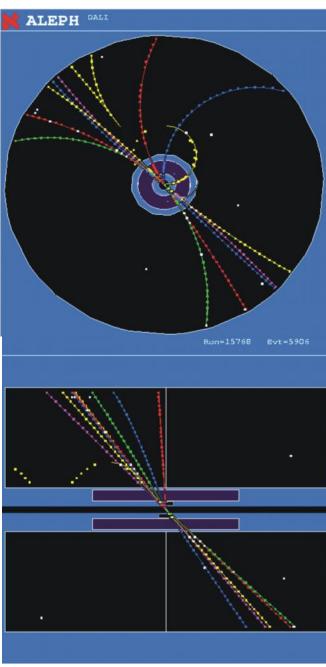
43

Time Projection Chamber (TPC):

Gas volume with parallel E and B Field.
B for momentum measurement. Positive effect: Diffusion is strongly reduced by E//B (up to a factor 5).

Drift Fields 100-400V/cm. Drift times 10-100 μs . Distance up to 2.5m !





W. Riddler/CERN

ALICE TPC: Construction Parameters

Largest TPC:

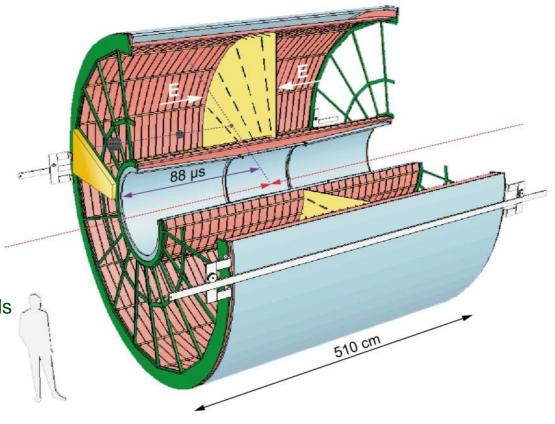
- Length 5m
- Diameter 5m
- Volume 88m³
- Detector area 32m²
- Channels ~570 000

High Voltage:

Cathode -100kV

Material X₀

 Cylinder from composite materials from airplane industry (X₀= ~3%)



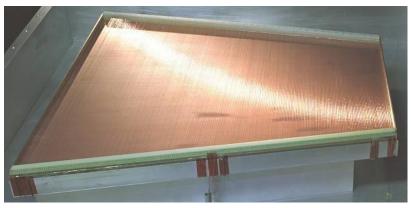
ALICE TPC: Pictures of the Construction

Precision in z: 250µm

End plates 250μm

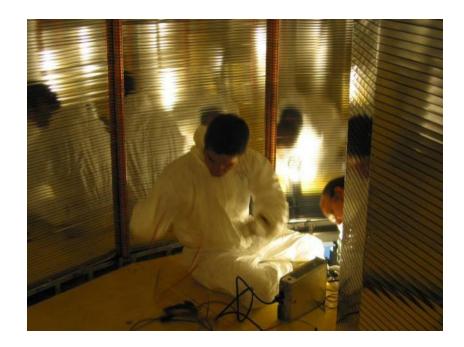






Wire chamber: 40µm





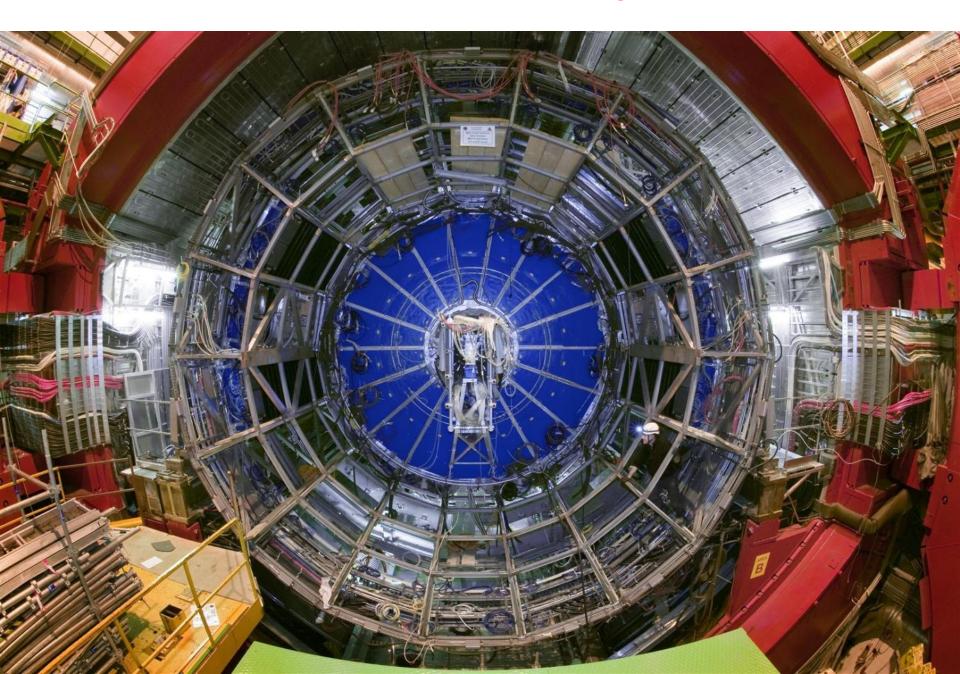
ALICE TPC Construction

My personal contribution:

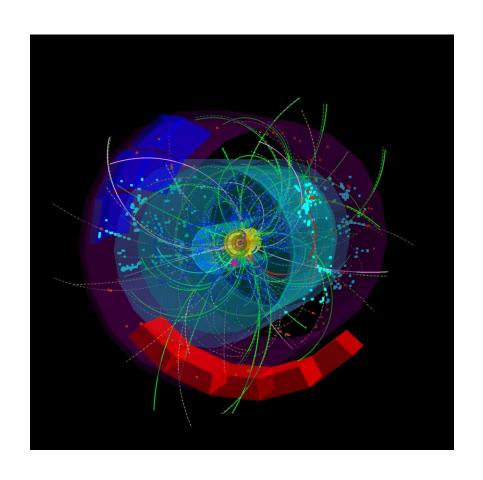
A visit inside the TPC.

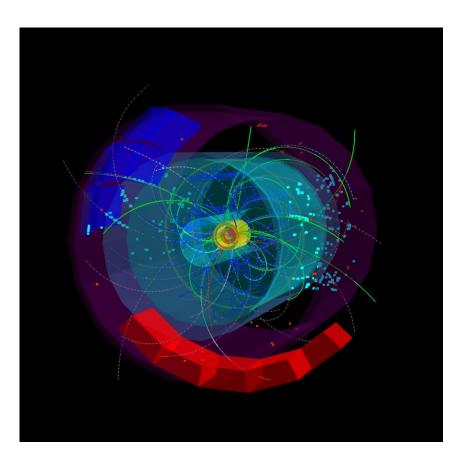


TPC installed in the ALICE Experiment

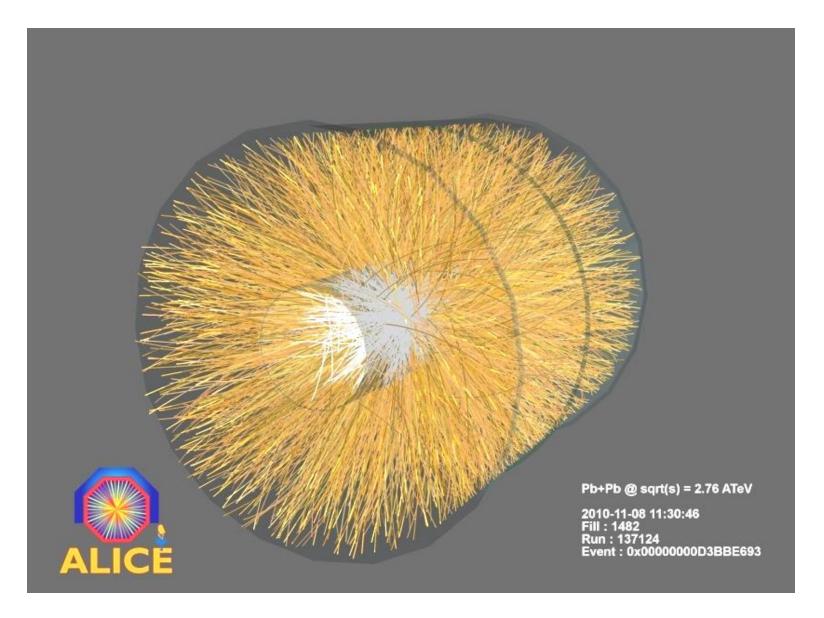


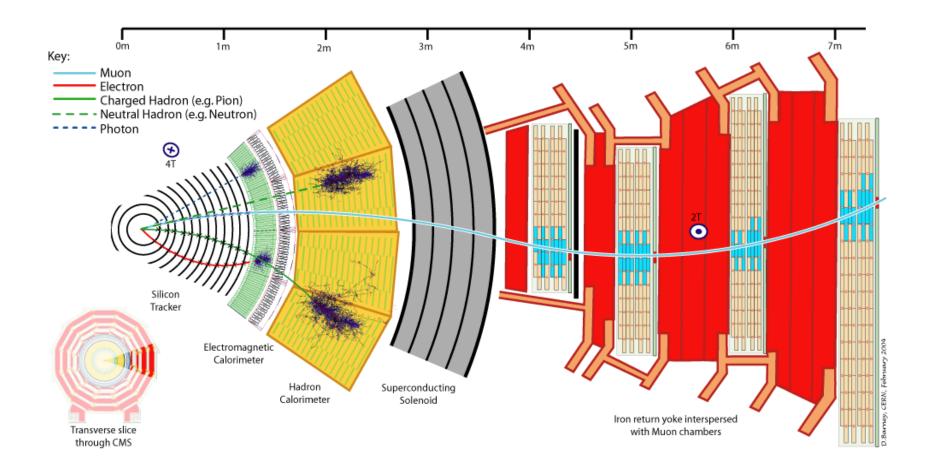
First 7 TeV p-p Collisions in the ALICE TPC in March 2010!





First Pb Pb Collisions in the ALICE TPC in Nov 2010!





Calorimeter

Homogeneous

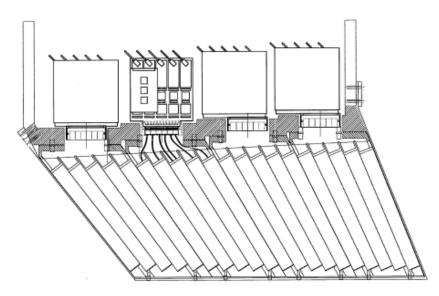
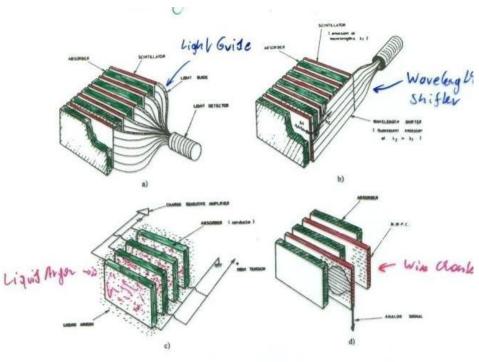


Fig. 2. Longitudinal drawing of module 2, showing the structure and the front-end electronics layout.

Sampling



The Geiger counter reloaded: Drift Tube

Atlas Muon Spectrometer, 44m long, from r=5 to11m.

1200 Chambers

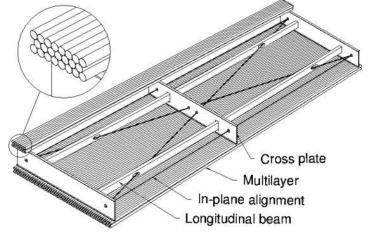
6 layers of 3cm tubes per chamber.

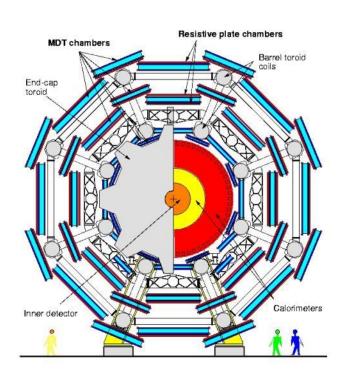
Length of the chambers 1-6m!

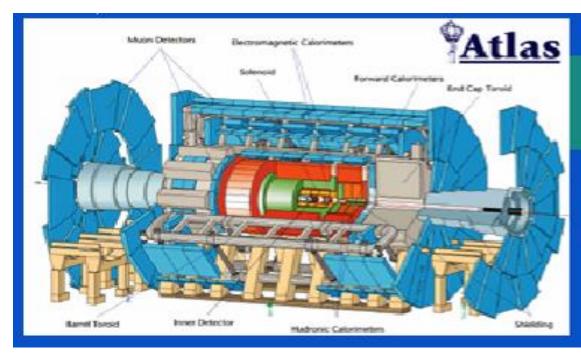
Position resolution: 80μm/tube, <50μm/chamber (3 bar)

Maximum drift time ≈700ns

Gas Ar/CO₂ 93/7







Drift Tube

ATLAS MDT R(tube) =15mm

Calibrated Radius-Time correlation

TDC

Amplifier Discriminator

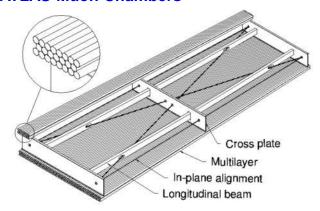
Primary electrons are drifting to the wire.

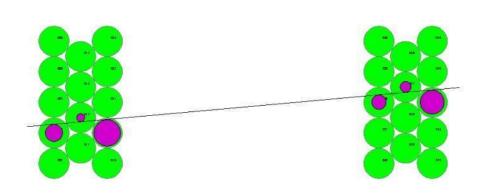
Electron avalanche at the wire.

The measured drift time is converted to a radius by a (calibrated) radius-time correlation.

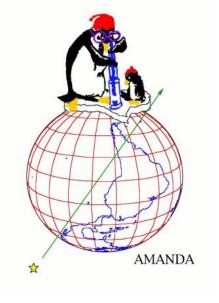
Many of these circles define the particle track.

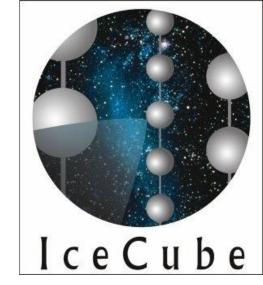
ATLAS Muon Chambers





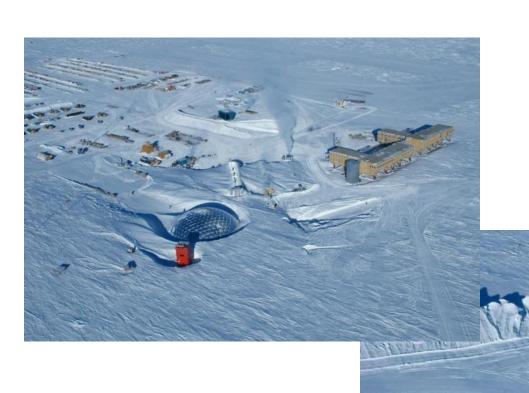
ATLAS MDTs, 80µm per tube





Antarctic Muon And Neutrino Detector Array

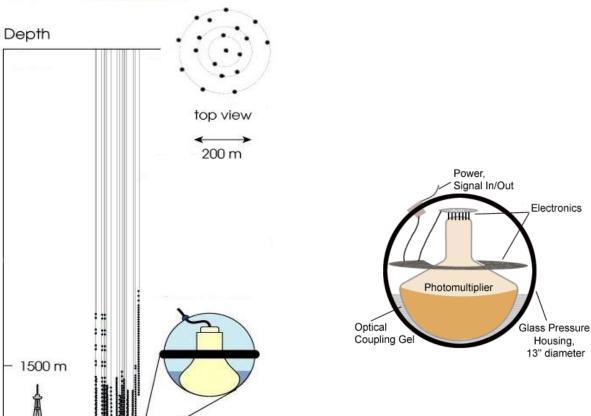
AMANDA & ICE Cube



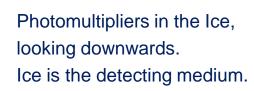
South Pole



AMANDA-II







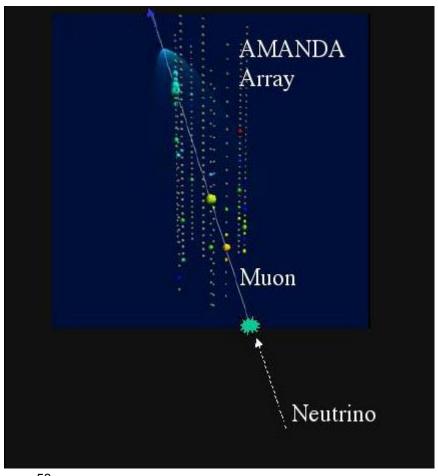


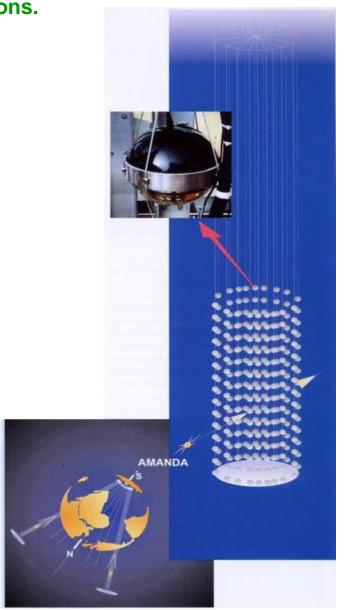
2500 m

- 2000 m

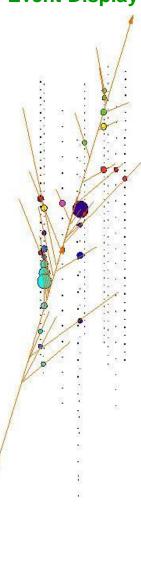
Look for upwards going Muons from Neutrino Interactions. Cherekov Light propagating through the ice.

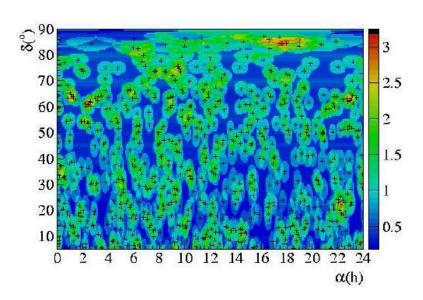
→ Find neutrino point sources in the universe!





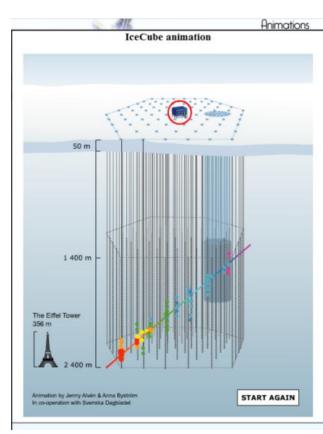
Event Display





Up to now: No significant point sources but just neutrinos from cosmic ray interactions in the atmosphere were found.

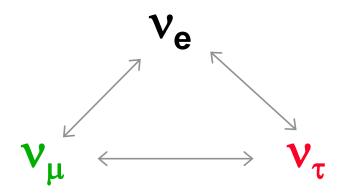
→ Ice Cube for more statistics!



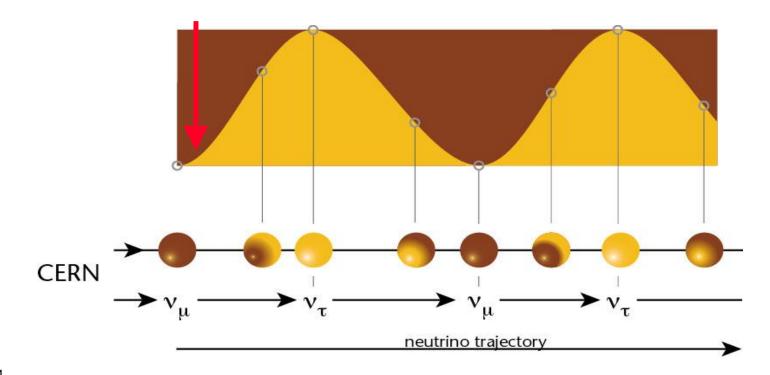
CERN Neutrino Gran Sasso

(CNGS)

If neutrinos have mass:



Muon neutrinos produced at CERN. See if tau neutrinos arrive in Italy.



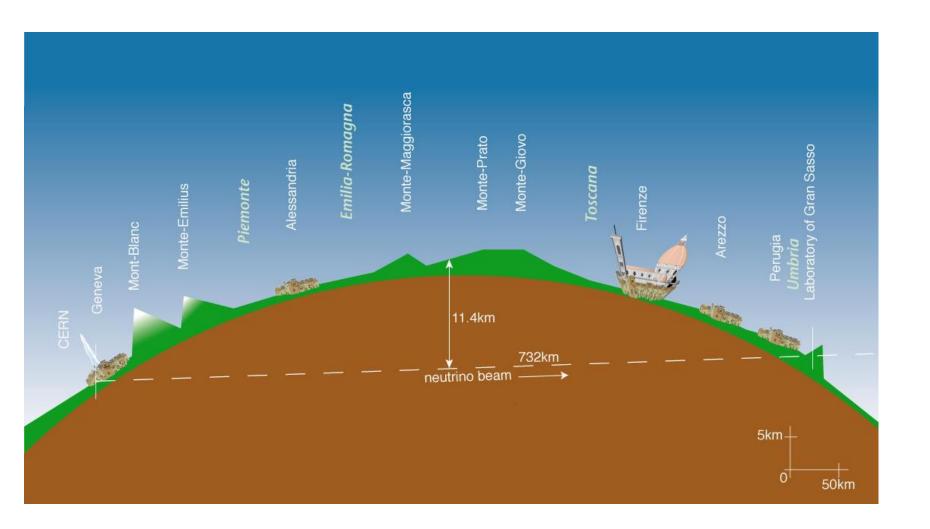
CNGS Project

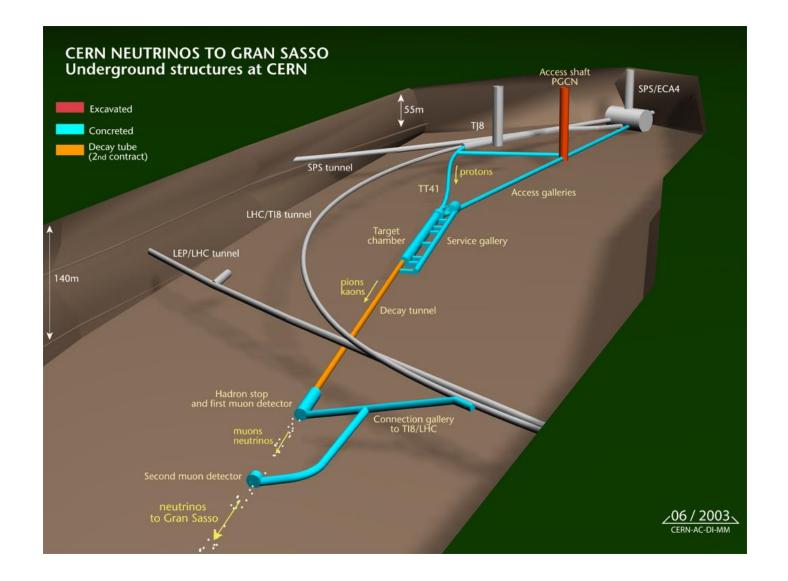
CNGS (CERN Neutrino Gran Sasso)

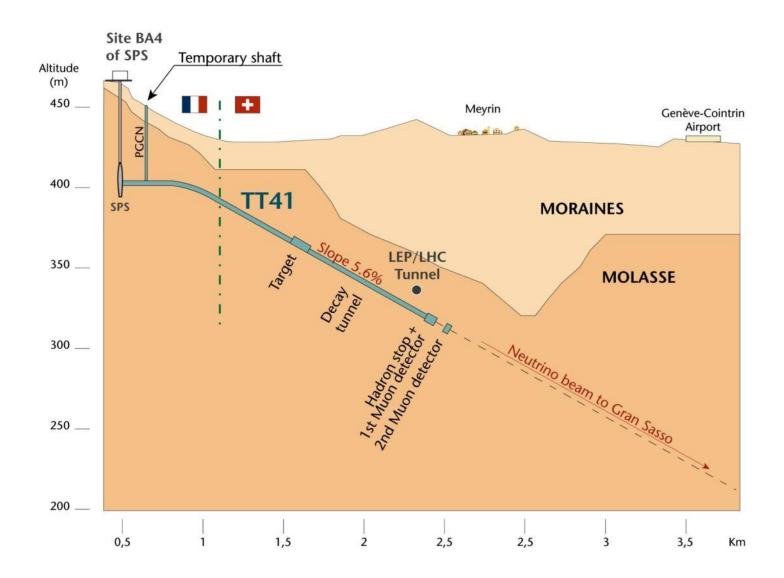
- A long base-line neutrino beam facility (732km)
- send v_{μ} beam produced at CERN
- detect v_{τ} appearance in OPERA experiment at Gran Sasso



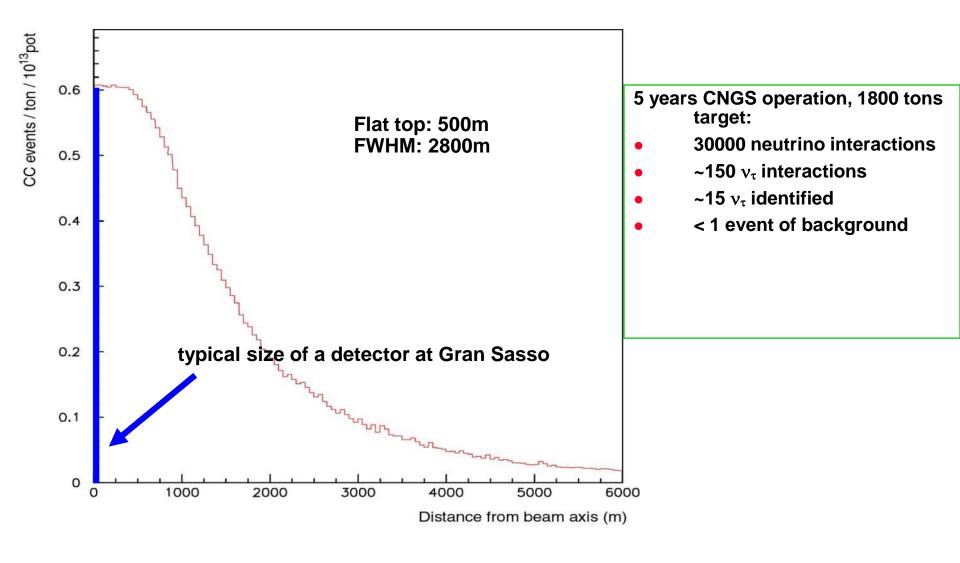
direct proof of v_{μ} - v_{τ} oscillation (appearance experiment)







Radial Distribution of the v_{μ} -Beam at GS



E₆Gschwendtner, CERN

W. Riegler/CERN

Neutrinos at CNGS: Some Numbers

 2×10^{19}

For 1 year of CNGS operation, we expect:

protons on target

protons on target	2 X 10
pions / kaons at entrance to decay tunnel	3 x 10 ¹⁹
pions / kaons at entrance to decay turner	3 X 10.~

10 ¹⁹
1

$$v_{\mu}$$
 in 100 m² at Gran Sasso 3 x 10¹⁴

$$V_{\mu}$$
 events per day in OPERA ≈ 2500

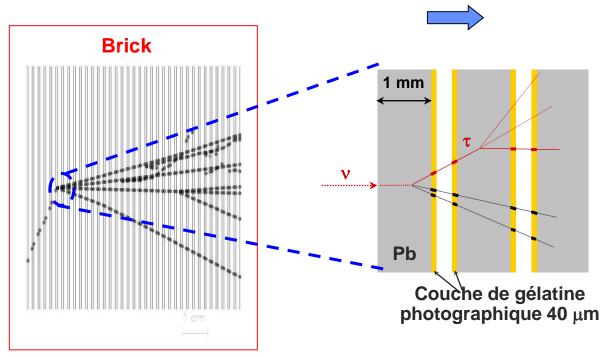
$$V_{\tau}$$
 events (from oscillation) ≈ 2

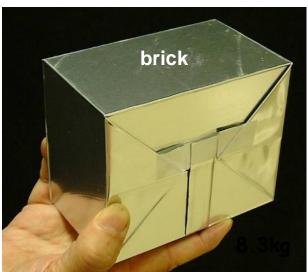
Basic unit: brick

56 Pb sheets + 56 photographic films (emulsion sheets)

Lead plates: massive target

Emulsions: micrometric precision

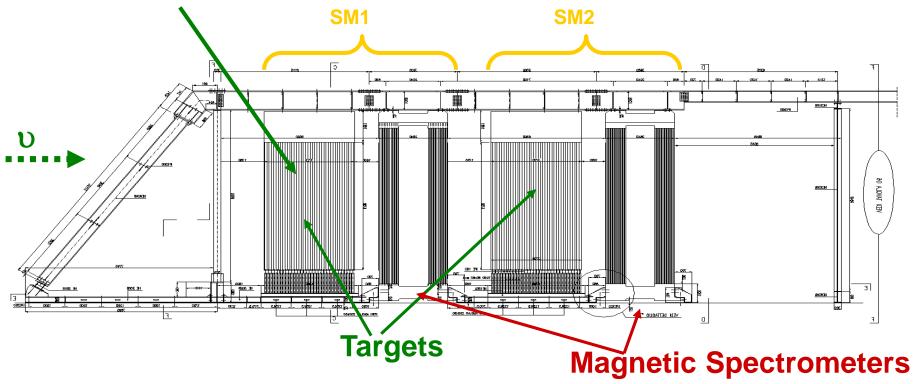




10.2 x 12.7 x 7.5 cm³



31 target planes / supermodule In total: 206336 bricks, 1766 ton



First observation of CNGS beam neutrinos: August 18th, 2006

Second Super-module



Scintillator planes 5900 m² 8064 7m long drift tubes

Details of the first spectrometer)PERA

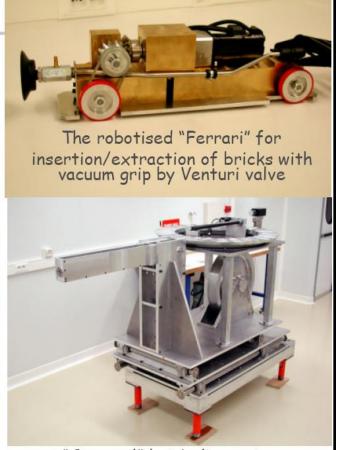


3050 m² Resistive Plate Counters 2000 tons of iron for the two magnets

The Brick Manipulator System (BMS) prototype: a lot of fun for children and adults!

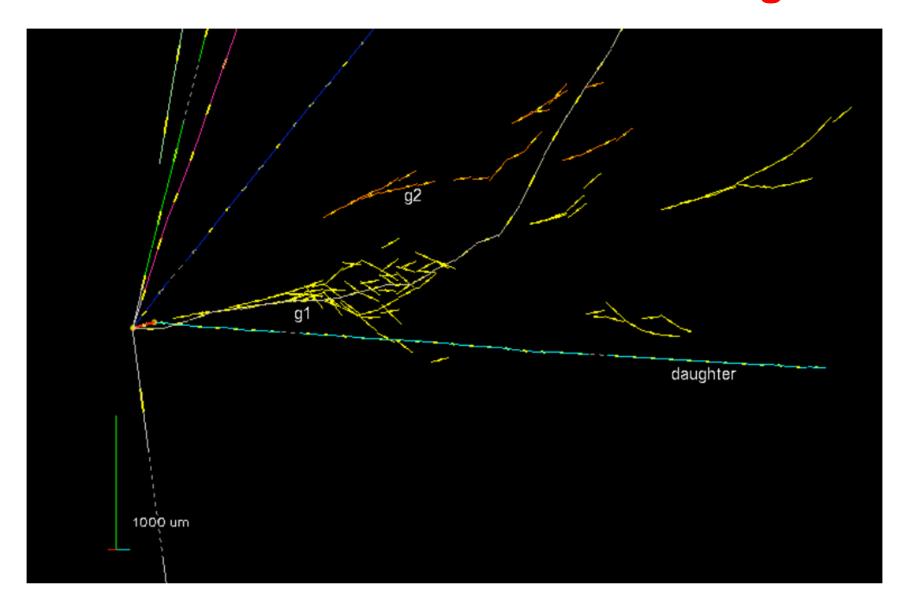


Tests with the prototype wall



"Carousel" brick dispensing and storage system

First Tau Candidate seen a few weeks ago!

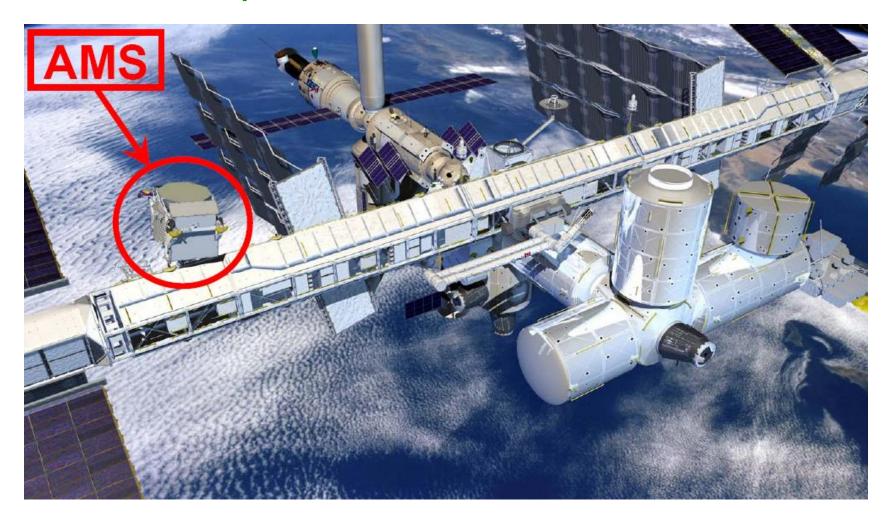


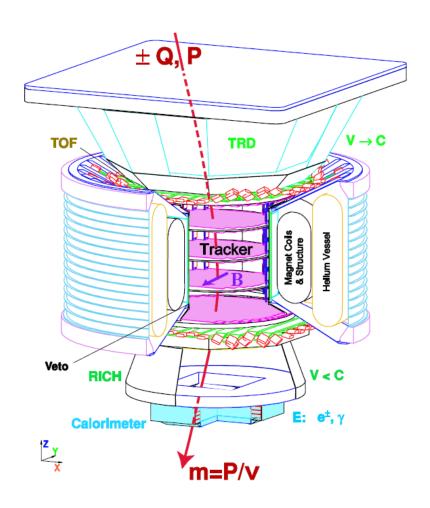


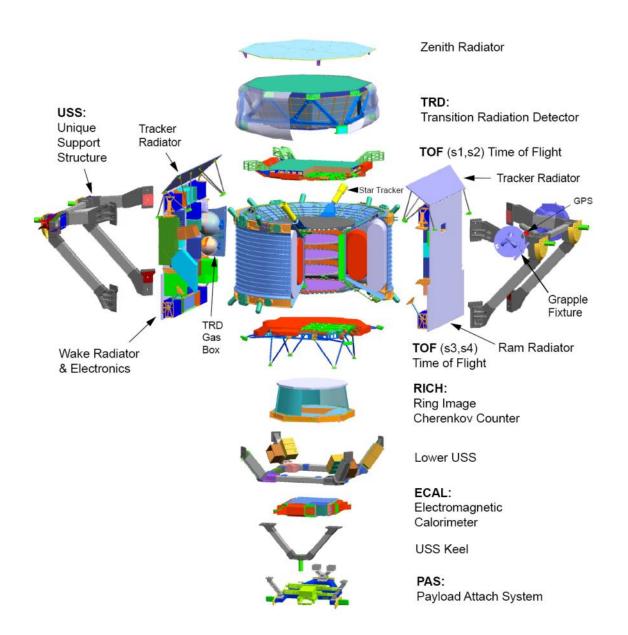
Alpha Magnetic Spectrometer

Try to find Antimatter in the primary cosmic rays. Study cosmic ray composition etc. etc.

Installed on the space station.







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Summary

There are just a few basic principles for particle detection ...

... but there is an infinite number if ingenious ways to use these principles in order to force nature to reveal it's secrets!