Introduction to Neutrino Physics

Hiroaki SUGIYAMA (Toyama Prefectural Univ.)

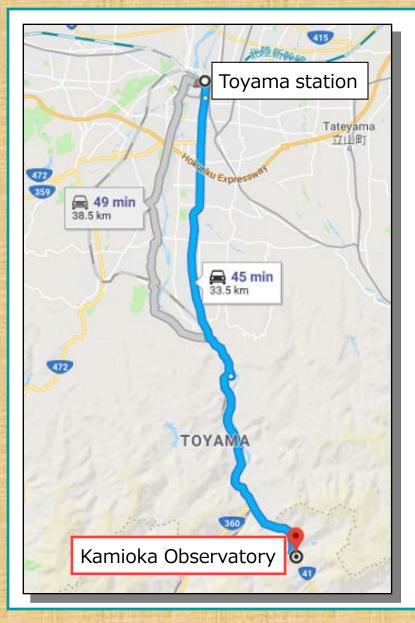
What are Neutrinos

- Elementary particles
- Electrically neutral
- Spin 1/2 (Fermion that obeys the Fermi statistics)
- Three types (at the least) as partners of electron, muon, and tau-lepton
 Image: Constraint of the least)
 Image: Const

higgstan.com

- Very light such that they had been regarded as massless particles for a long time
- Weakly interact

Neutrino and Toyama



- Easy (relatively) to go to Kamioka Observatory (in Gifu prefecture) from Toyama city
- Two Nobel prizes for neutrino experiments in Kamioka Observatry

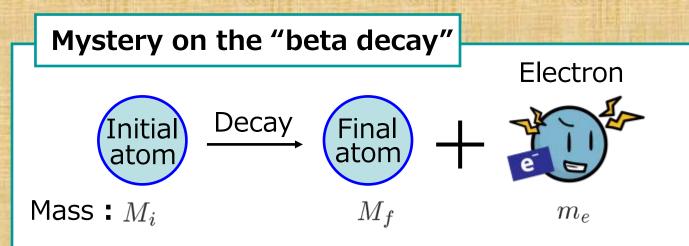




Koshiba

Kajita (Kamiokande) (Super-Kamiokande)

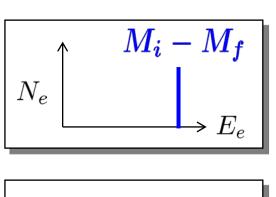
Hello, Neutrino ! - Why is it Necessary ? -



 N_e

Electron energy spectrum

Expectation :



 $> E_e$

Observation :



Mono-energetic

Energy conservation

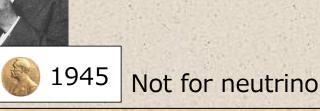
Dear Radioactive Ladies and Gentlemen,

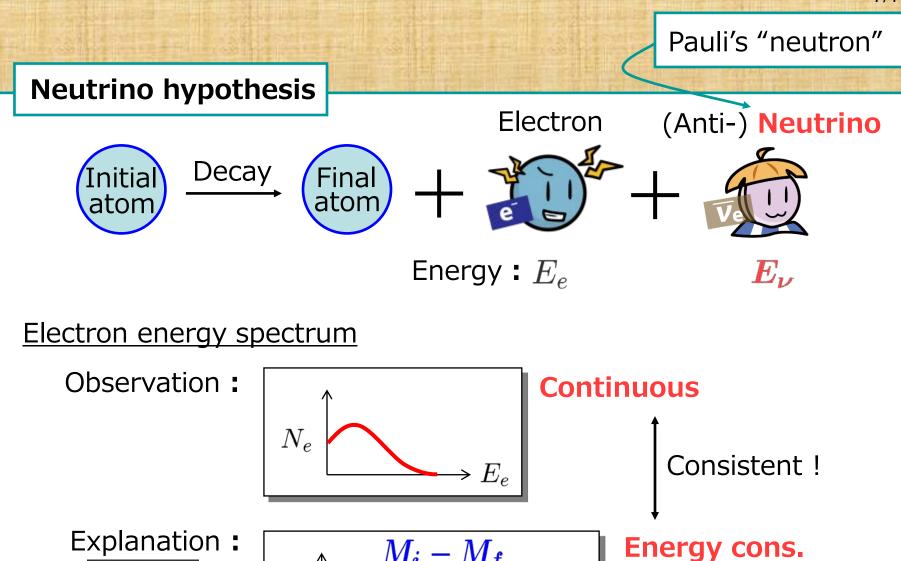
As the bearer of these lines, to whom I graciously ask you to listen, will explain to you in more detail, how because of the "wrong" statistics of the N and Li6 nuclei and the continuous beta spectrum, I have hit upon a desperate remedy to save the "exchange theorem" of statistics and the law of conservation of energy. Namely, the possibility that there could exist in the nuclei electrically neutral particles, that **I wish to call neutrons**, which have spin 1/2 and obey the exclusion principle and which further differ from light quanta in that they do not travel with the velocity of light. The mass of the neutrons should be of the same order of magnitude as the electron mass and in any event not larger than 0.01 proton masses. The continuous beta spectrum would then become understandable by the assumption that in beta decay a neutron is emitted in addition to the electron such that the sum of the energies of the neutron and the (neutrino) electron is constant...

 $\boldsymbol{E_{\nu}} + E_e = M_i - M_f$

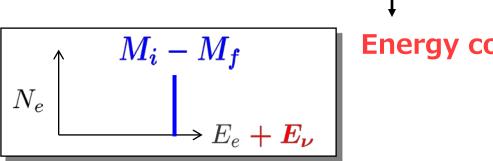
I agree that my remedy could seem incredible because one should have seen those neutrons very earlier if they really exist. But only the one who dare can win and the difficult situation, due to the continuous structure of the beta spectrum, is lighted by a remark of my honoured predecessor, Mr. Debye, who told me recently in Bruxelles: "Oh, It's well better not to think to this at all, like new taxes". From now on, every solution to the issue must be discussed. Thus, dear radioactive people, look and judge. Unfortunately, I cannot appear in Tubingen personally since I am indispensable here in Zurich because of a ball on the night of 6/7 December. With my best regards to you, and also to Mr. Back.

Your humble servant W. Pauli

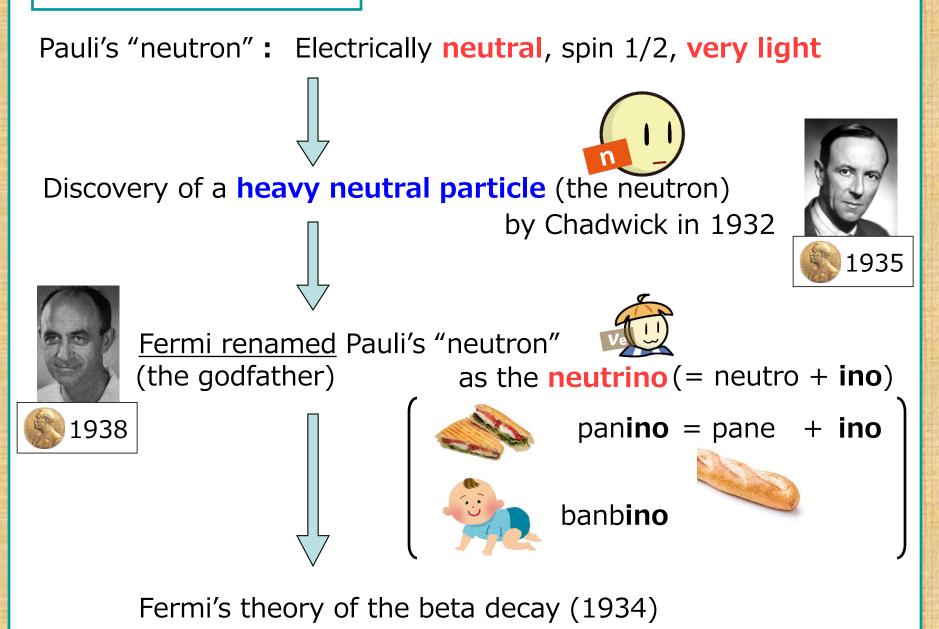




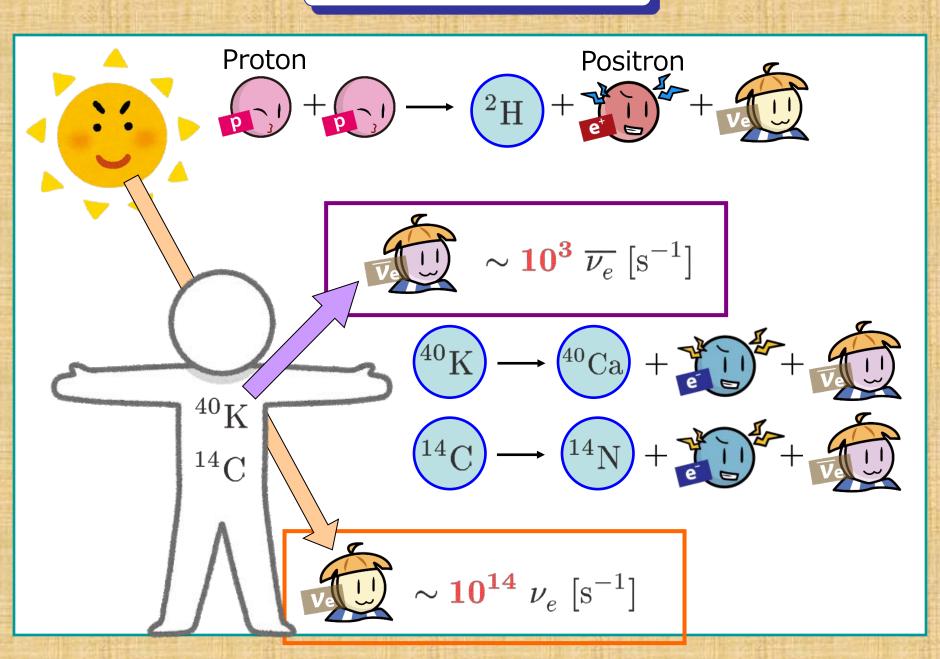




Neutrino's godfather



Neutrinos Around



Discovery of Neutrino

The nuclear reactor is a powerful source of (anti-) neutrinos

+ $(1) \rightarrow (X_1) + (X_2) + 2 \text{ or } 3$

Reines and Cowan proposed an experiment

Electric power $\simeq 0.33 \times$ Thermal power

 $^{235}\mathrm{U}$

1995

to detect the reactor (anti-) neutrino

 $1 [GW_{th}] \Rightarrow 1.6 \times 10^{20} \overline{\nu_{e}} [s^{-1}]$

+620 + 620 + 200 MeV

Proposal : Phys. Rev. 90, 492 (24 Feb. 1953)



1956)

10/17

First result : Phys. Rev. 92, 830 (9 Jul. 1953)

Improved result : Science 124, 103 (

Neutrino Oscillation

The Nobel Prize in Physics 2015



Takaaki Kajita



11/17

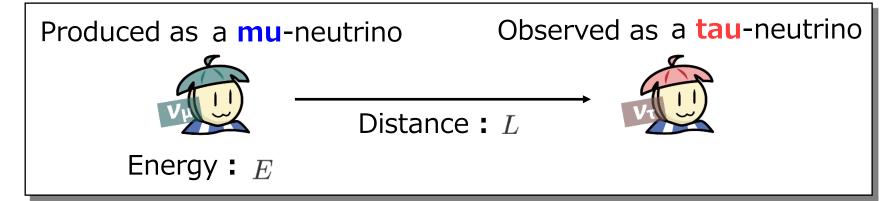
Arthur B. McDonard

"for the discovery of **neutrino oscillations**, which shows that neutrinos have mass"

What is the neutrino oscillation ?



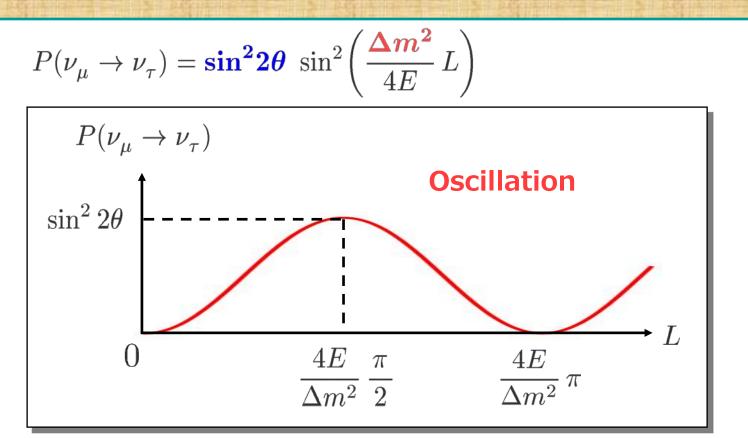
Neutrino oscillation : Oscillation of transition probability (Quantum theory)



Transition probability

$$P(\nu_{\mu} \rightarrow \nu_{\tau}) = \frac{\sin^{2}2\theta}{\sin^{2}\left(\frac{\Delta m^{2}}{4E}L\right)}$$

$$\Delta m^{2} = m_{2}^{2} - m_{1}^{2}$$
Difference of squared masses of neutrinos A parameter (mixing angle)

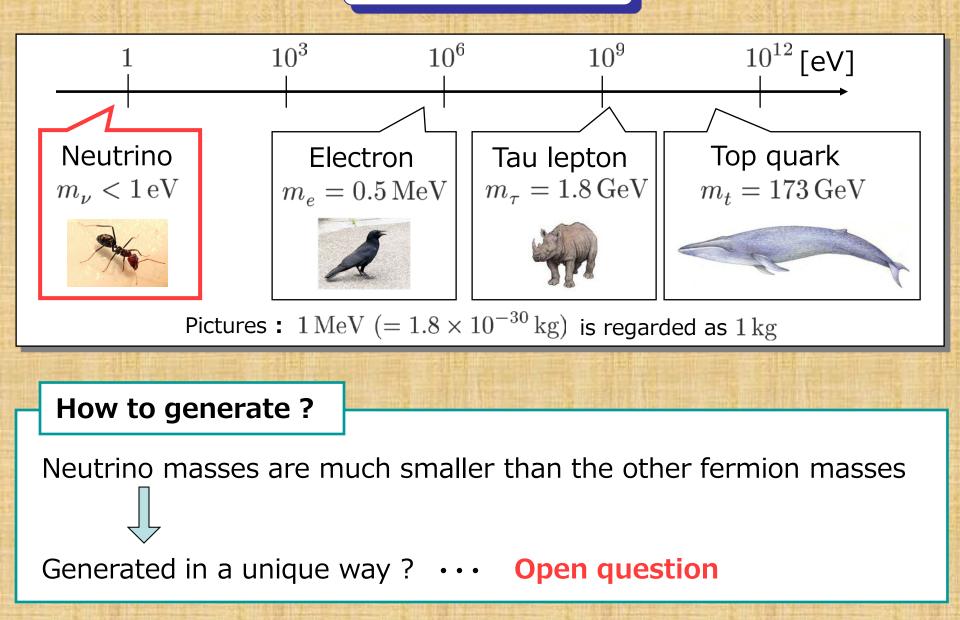


Observation of oscillation

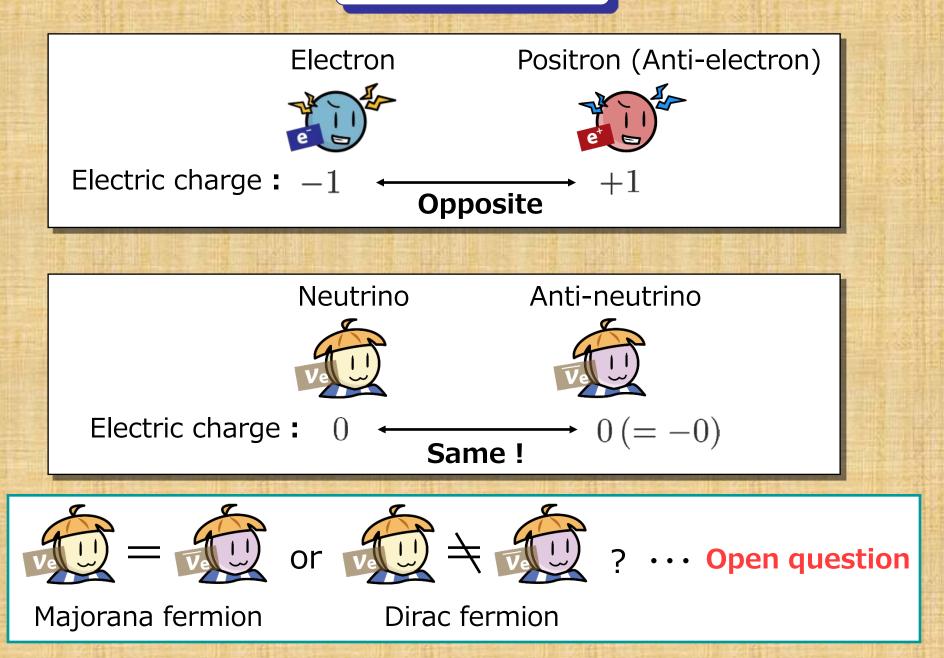
$$P(\nu_{\mu} \rightarrow \nu_{\tau}) \neq 0 \implies \Delta m^{2} \neq 0$$
$$\implies m_{1} \neq 0 \text{ and/or } m_{2} \neq 0$$
Existence of neutrino r

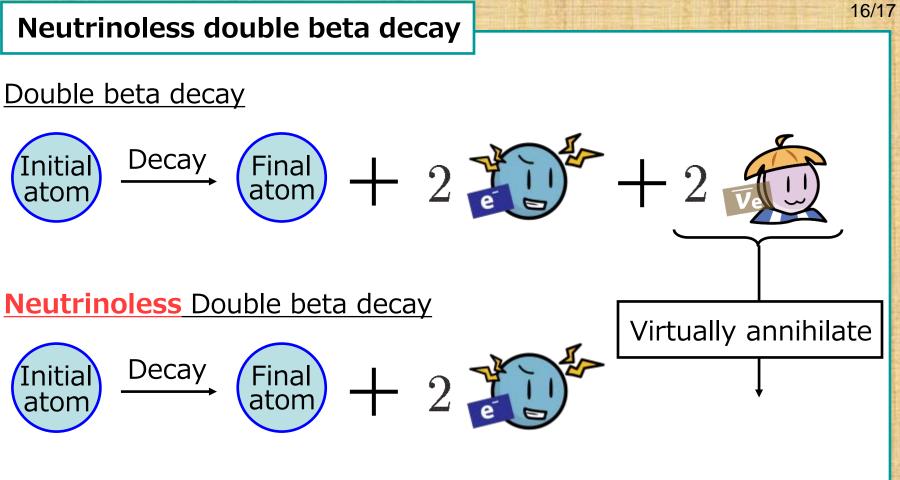
Existence of neutrino mass though neutrinos have been regarded as massless

Neutrino Mass



Anti-neutrino





- Possible only with Majorana neutrinos
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
 ↓
- KamLAND-Zen in Kamioka Obs. searches for the decay

Summary

Neutrinos : Elementary particles, weakly interact, electrically neutral, very light



Open questions :

How are **tiny neutrino masses** generated ?

- \rightarrow Probably, via neutrino-specific mechanism
- → Theoretical studies are necessary

$$\nu = \overline{\nu}) \qquad \qquad (\nu \neq \overline{\nu})$$

Are neutrinos Majorana fermions ? Dirac fermions ?

- \longrightarrow Origin of neutrino masses depends on the answer Majorana fermions \Rightarrow Neutrinos are really unique
- → Searches for **neutrinoless double beta decay** e.g., KamLAND-Zen in Kamioka Observatry



 ν physics guides us to new physics