

# STAR Femtoscopy

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Warsaw University  
of Technology

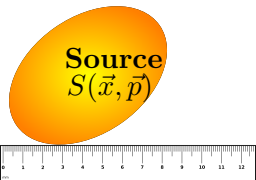


Faculty  
of Physics

WARSAW UNIVERSITY OF TECHNOLOGY

HIRG weekly Meeting 2019

Warsaw, 10.12.2019



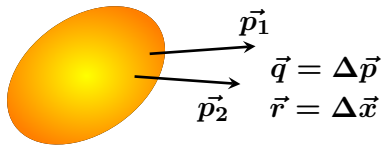
Impossible to examine the particle  
emitting source directly

$$\text{size} \sim 10^{-15} \text{ m}$$

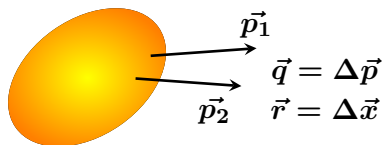
$$\text{lifetime} \sim 10^{-23} \text{ s}$$

**Femtoscscopy** measures space-time characteristics  
of the source through momentum correlations

# Femtoscropy



# Femtoscscopy



$$C(\vec{p}_1, \vec{p}_2) = \frac{P_{12}(\vec{p}_1, \vec{p}_2)}{P_1(\vec{p}_1)P_1'(\vec{p}_2)}$$

experiment

$$C(\vec{q}) = \frac{A(\vec{q})}{B(\vec{q})}$$

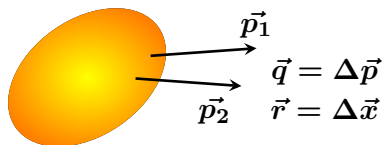
$A(\vec{q})$  - correlated  
 $B(\vec{q})$  - uncorrelated

theory (models)

$$C(\vec{q}) = \int d^3r S(\vec{q}, \vec{r}) |\Psi(\vec{q}, \vec{r})|^2$$

$S(\vec{q}, \vec{r})$  - emission function  
 $\Psi(\vec{q}, \vec{r})$  - pair wave function

# Femtoscscopy



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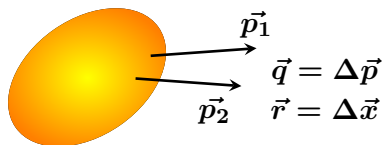
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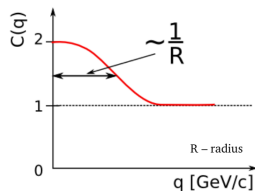
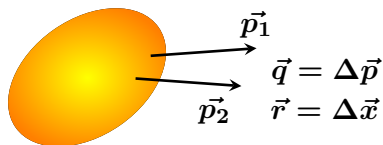
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# Femtoscscopy



calculate size ( $R$ ) of the source

$$C(\vec{p}_1, \vec{p}_2) = \frac{P_{12}(\vec{p}_1, \vec{p}_2)}{P_1(\vec{p}_1)P_1'(\vec{p}_2)}$$

experiment

theory (models)

$$C(\vec{q}) = \frac{A(\vec{q})}{B(\vec{q})}$$

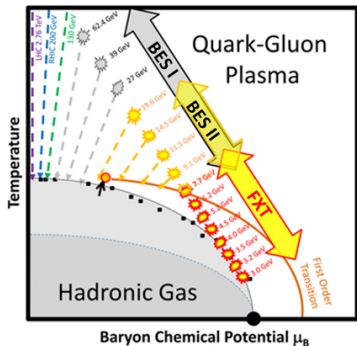
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# Introduction

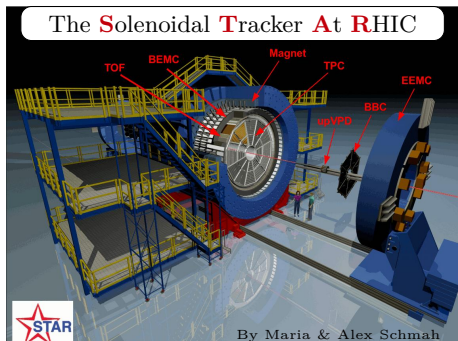
- **Kaon femtoscopy (geometry):**  
provides complementary information to pions
  - ▶ less affected by resonance decays
  - ▶ contain strange quark
  - ▶ heavier than pions
- **Proton femtoscopy (geometry):**
  - ▶ provide additional information about source characteristics
- **Non-identical particle femtoscopy (geometry + dynamics):**
  - ▶ examination of asymmetry in emission process between two kinds of particles ( $\pi K$ ,  $\pi p$ ,  $Kp$ )



Nucl. Phys. A 967, 808 (2017)



# The STAR experiment



## Time Projection Chamber

PID:  $dE/dx$

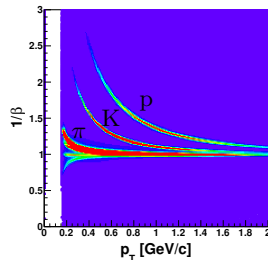
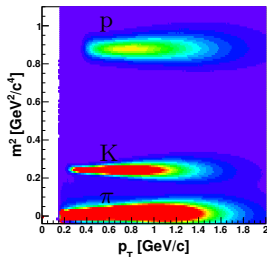
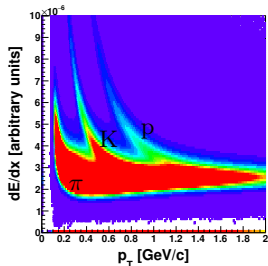
Tracking

$$0 < \phi < 2\pi, |\eta| < 1$$

## Time-Of-Flight

Time resolution  $< 80$  ps

PID:  $m^2$  &  $1/\beta$

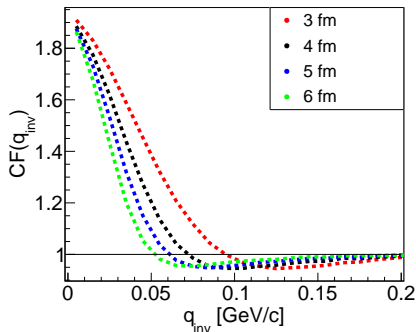
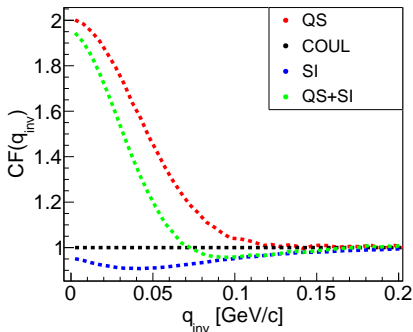


# Neutral kaon femtoscopy

# Neutral kaon correlation function

The shape of correlation function depends on:

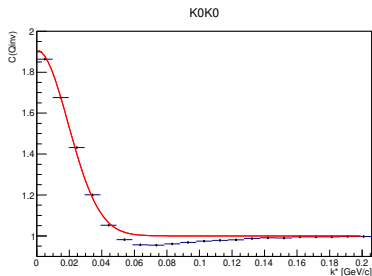
- Quantum Statistical effects (QS)
- Final State Interactions (FSI):
  - ▶ ~~Coulomb Interaction (COUL)~~
  - ▶ Strong Interaction (SI)



# Fitting procedure

The QS correlation function (Gaussian):

$$C(q_{inv}) = 1 + \lambda \exp(-R_{inv}^2 q_{inv}^2)$$



$\lambda$  — the correlation strength

$R_{inv}$  — the size of the particle emitting source

# Fitting procedure

The QS correlation function (Gaussian):

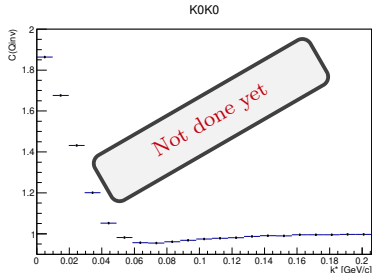
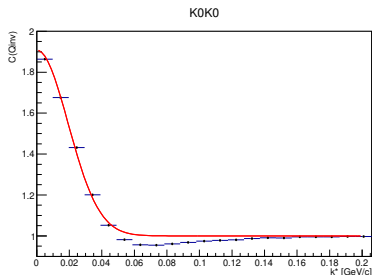
$$C(q_{inv}) = 1 + \lambda \exp(-R_{inv}^2 q_{inv}^2)$$

Final State Interaction  
Lednicky & Lyuboshitz model:

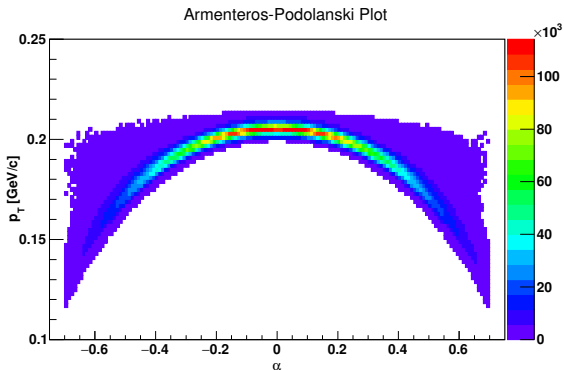
R. Lednicky and V.L. Lyuboshitz, Sov. J.  
Nucl. Phys. 35, 770 (1982)

$\lambda$  — the correlation strength

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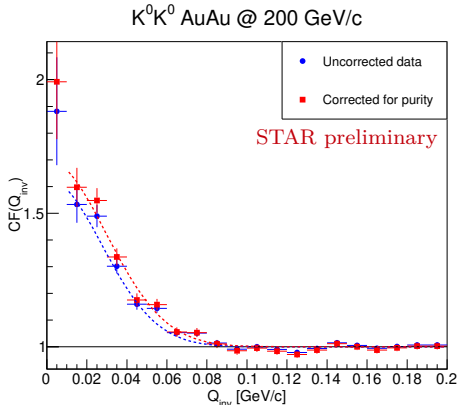


# Armenteros-Podolanski plot



signal region is symmetric  
other particles, e.g.  $\lambda$  &  $\bar{\lambda}$ , are not noticeable

# $K^0 K^0$ correlation functions, $\sqrt{s_{NN}} = 200$ GeV



## Before purity correction

Radius [fm]	$\lambda$
$5.08 \pm 0.19$	$0.630 \pm 0.051$

## After purity correction

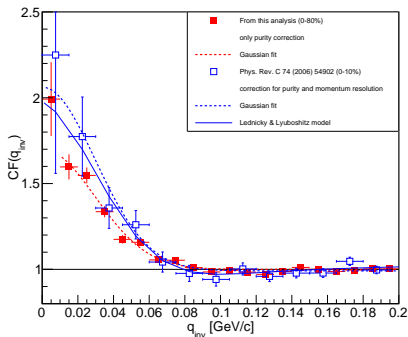
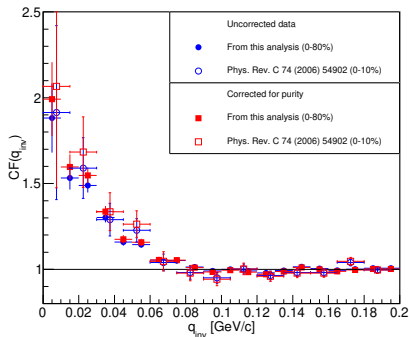
Radius [fm]	$\lambda$
$4.72 \pm 0.20$	$0.701 \pm 0.056$

Smaller values of source's radii after corrections (larger statistical uncertainties)  
and different values of  $\lambda$  parameter (larger for correlations after corrections)

# $K^0 K^0$ CFs, $\sqrt{s_{NN}} = 200$ GeV, comparison

Neutral kaon interferometry in Au+Au collisions at  $\sqrt{s_{NN}} = 200$  GeV

Phys. Rev. C 74 (2006) 54902

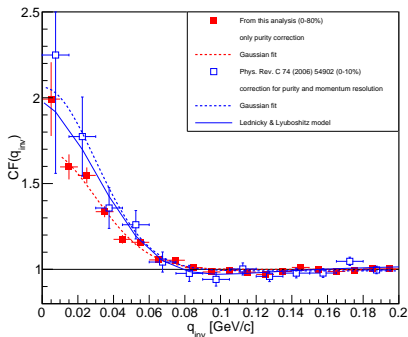
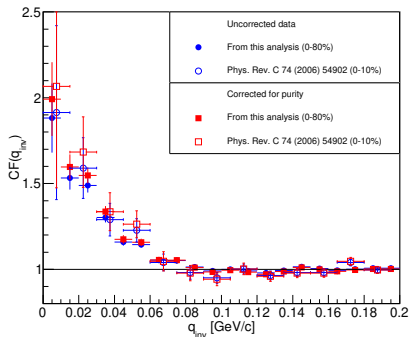


	Radius [fm]
<b>This analysis</b>	$4.72 \pm 0.20$
<b>STAR results</b>	$5.02 \pm 0.61$



# $K^0 K^0$ CFs, $\sqrt{s_{NN}} = 200$ GeV, comparison

Neutral kaon interferometry in Au+Au collisions at  $\sqrt{s_{NN}} = 200$  GeV  
Phys. Rev. C 74 (2006) 54902



- similar source sizes are obtained
- shape of the correlation functions before and after applying purity correction is similar

# Neutral kaon femtoscopy — future plan

- Next step:
  - ▶ Analysis in centrality bins: central (0-10%) and non-central (10-70%)
- Future plan:
  - ▶ Lower BES energies (39 GeV and 54 GeV)
  - ▶ Analyze  $K^0 K^{ch}$  pairs

# Proton femtoscopy

# Proton correlation function

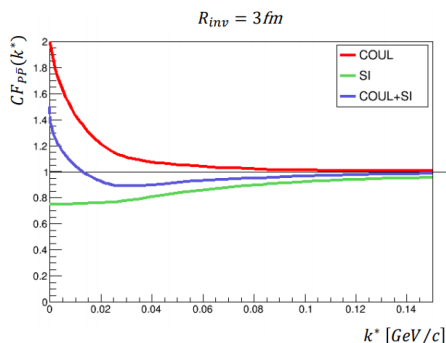
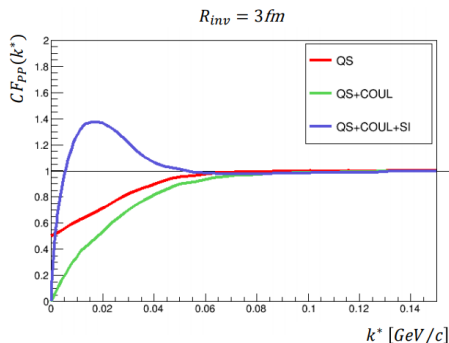
The shape of CF depends on:

proton-proton

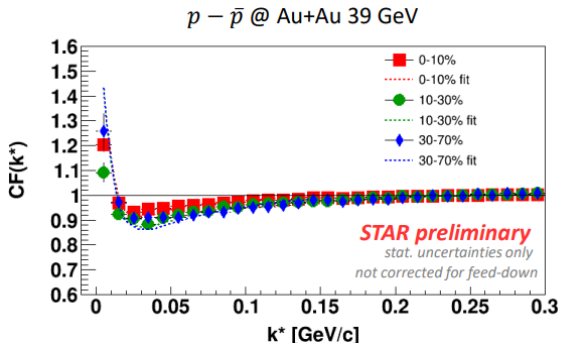
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proton-antiproton

- ~~Quantum Statistical effects (QS)~~
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# $p\bar{p}$ correlation function



Clear centrality dependence

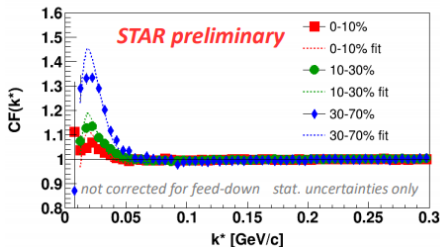
$$R(0-10\%) > R(10-30\%) > R(30-70\%)$$

centrality	$R_{inv} p - \bar{p}$ [fm]
0-10%	$3.69 \pm 0.12 \pm 0.14$
10-30%	$2.69 \pm 0.10 \pm 0.12$
30-80%	$2.56 \pm 0.09 \pm 0.12$

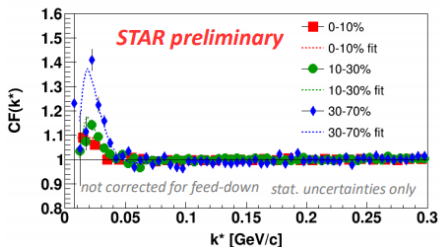
QM 2018

# $pp/\bar{p}\bar{p}$ correlation function

$p - p$  @ Au+Au 39 GeV



$\bar{p} - \bar{p}$  @ Au+Au 39 GeV

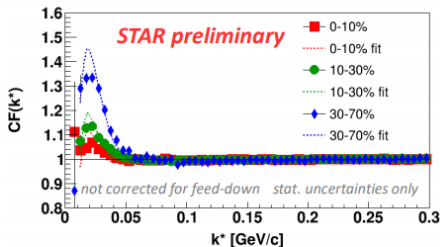


centrality	$R_{inv} p - p$ [fm]	$R_{inv} \bar{p} - \bar{p}$ [fm]	$R_{inv} p - \bar{p}$ [fm]
0-10%	$4.00 \pm 0.15 \pm 0.02$	$3.83 \pm 0.20 \pm 0.03$	$3.69 \pm 0.12 \pm 0.14$
10-30%	$3.61 \pm 0.13 \pm 0.17$	$3.68 \pm 0.15 \pm 0.11$	$2.69 \pm 0.10 \pm 0.12$
30-80%	$2.72 \pm 0.07 \pm 0.07$	$2.95 \pm 0.11 \pm 0.08$	$2.56 \pm 0.09 \pm 0.12$

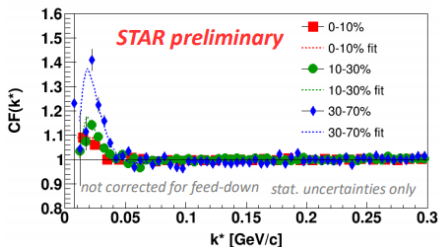
No significant difference between  $p - p$  and  $\bar{p} - \bar{p}$  correlation functions

# $pp/\bar{p}\bar{p}$ correlation function

$p - p$  @ Au+Au 39 GeV



$\bar{p} - \bar{p}$  @ Au+Au 39 GeV

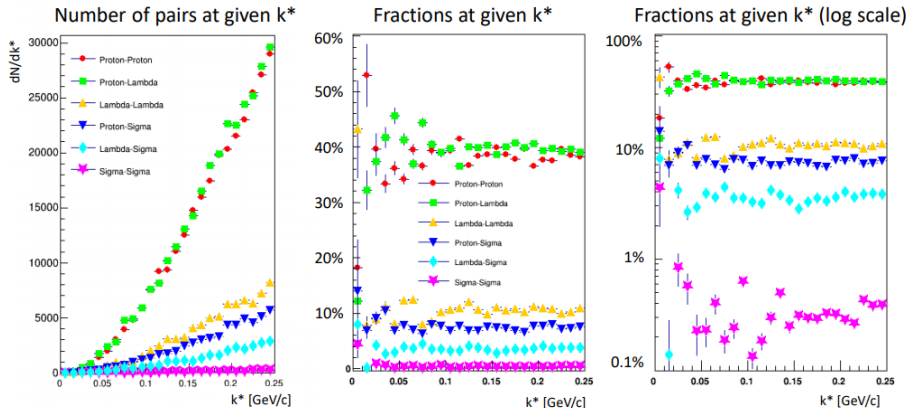


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Radii from  $p - p$  and  $\bar{p} - \bar{p}$  systems differ from radii from  $p - \bar{p}$  system  $\rightarrow$  residual correlations contaminate correlation functions

# Proton femtoscopy — future plan

- Next step:
  - ▶ Correction of residual contamination





# Non-identical particle femtoscopy

# Asymmetry

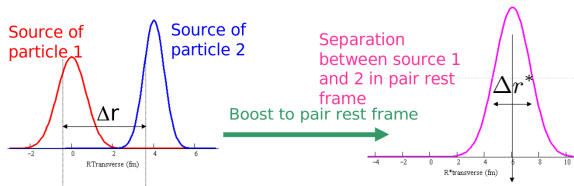
$$C(\vec{q}) = \int |\Psi(\vec{q}, \vec{r})|^2 S(\vec{r}) d^3r$$

known
unknown

R. Lednicky, et al.  
Phys. Lett. B373, 30 (1996)

$$S(\vec{r}) = \exp\left(-\frac{(r_{out} - \mu_{out})^2}{\sigma_{out}^2} - \frac{r_{side}^2}{\sigma_{side}^2} - \frac{r_{long}^2}{\sigma_{long}^2}\right)$$

$\mu_{out}$  — asymmetry in the *outward* direction  
 assumption:  $\sigma_{side} = \sigma_{out}$ ,  $\sigma_{long} = 1.3\sigma_{out}$



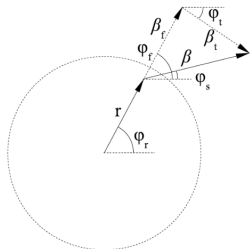
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 assumption:  $\sigma_{side} = \sigma_{out}$ ,  $\sigma_{long} = 1.3\sigma_{out}$

$$\beta_{particle} = \beta_f + \beta_t$$

$\beta_f$  — collective (flow) velocity

$\beta_t$  — thermal (random) velocity

A. Kisiel  
Phys. Rev. C81, 064906 (2010)

Emission asymmetry arises in a system where both thermal and collective velocities exist and are comparable in magnitude

# Spherical harmonics (SH)

SH representation of 3D correlation function as a set of 1D plots

$$C(\mathbf{q}) = \sum_{l,m} C_l^m(q) Y_l^m(\theta, \phi)$$

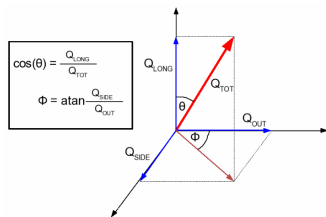
$$C_l^m(q) = \int_{\Omega} C(q, \theta, \phi) Y_l^m(\theta, \phi) d\Omega$$

$\Omega$  - full solid angle

$Y_l^m(\theta, \phi)$  - spherical harmonic function

$q = |\mathbf{q}|$  - pair relative momentum

$\theta$  and  $\phi$  - polar and azimuthal angle



P. Danielewicz and S. Pratt.  
Phys. Lett B618, 60 (2005)  
Phys. Rev. C75, 034907 (2007)

Z. Chajecki and M. Lisa  
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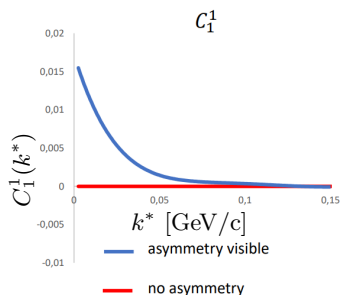
$Y_l^m(\theta, \phi)$  - spherical harmonic function

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$\theta$  and  $\phi$  - polar and azimuthal angle

$C_0^0 \rightarrow$  sensitive to the size of the emitting source  
(shapes same as correlation function)

$C_1^1 \rightarrow$  sensitive to the spacetime emission asymmetry



P. Danielewicz and S. Pratt.  
Phys. Lett B618, 60 (2005)  
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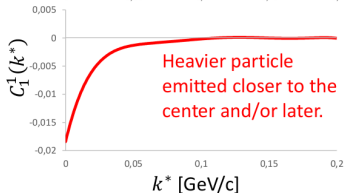
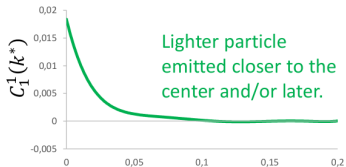
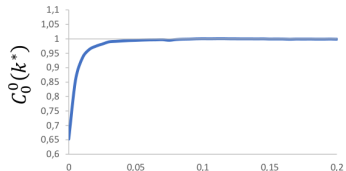
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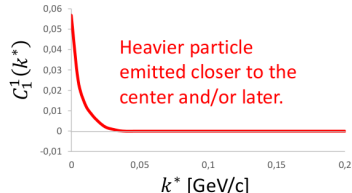
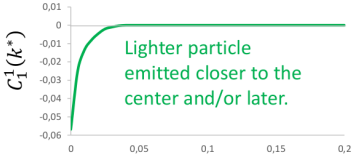
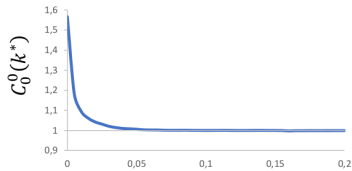
A. Kiesel  
Phys. Rev. C81, 064906 (2010)

# Which particle...?

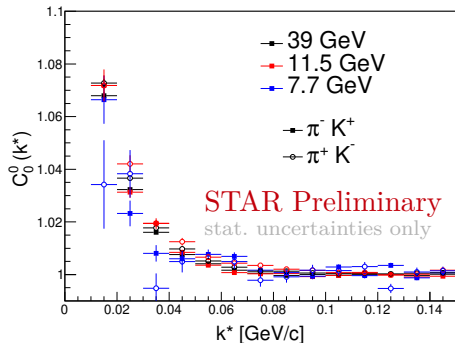
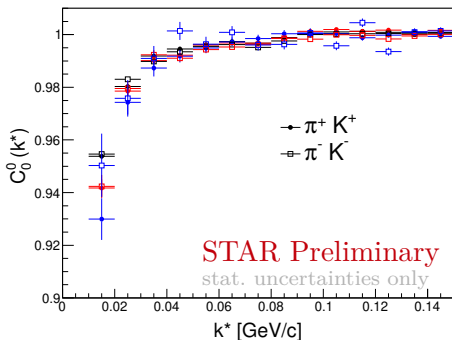
## Like-sign particle combinations



## Unlike-sign particle combinations



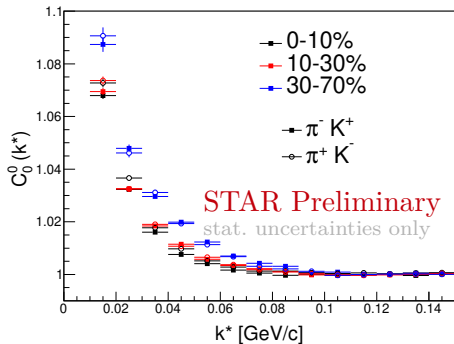
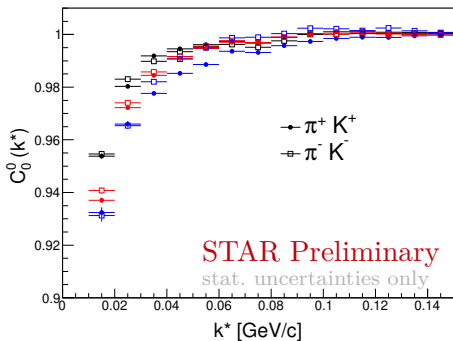
# Energy dependence



- Visible energy dependence

- Higher statistics for low energies are needed (BES-II)

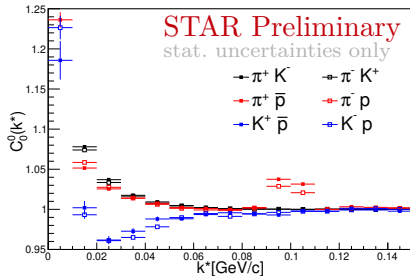
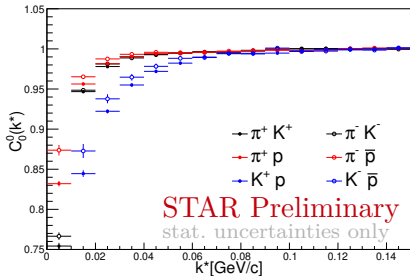
# Centrality dependence, $\sqrt{s_{NN}} = 39$ GeV



- Visible centrality dependence

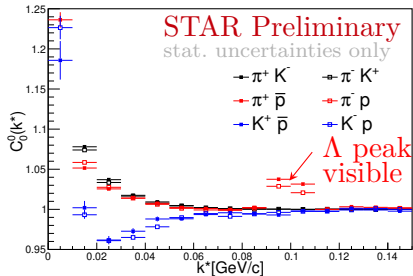
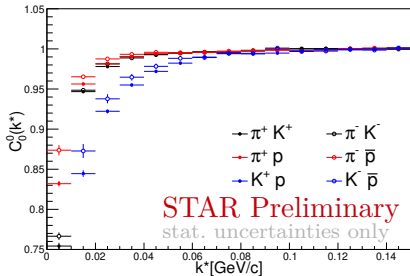


# System dependence, $\sqrt{s_{NN}} = 39$ GeV, 0-10%



- Like-sign pairs are dominated by Coulomb
- Kp  $\rightarrow$  strongest correlation

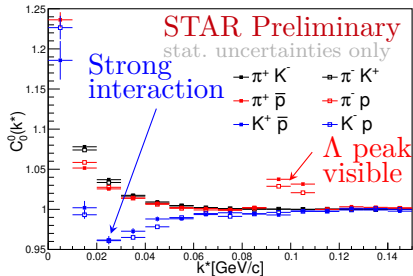
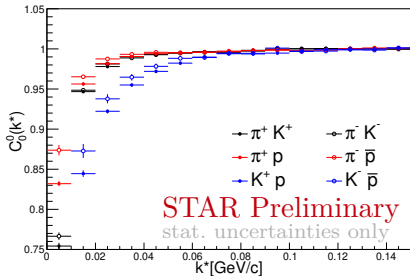
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- $\Lambda$  peak is visible in pion-proton

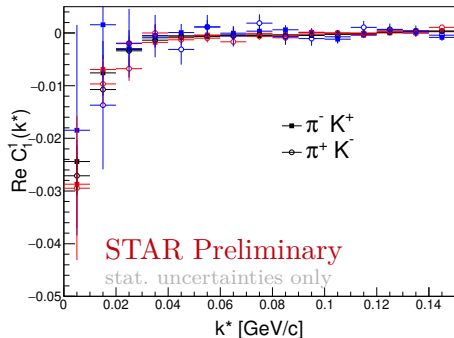
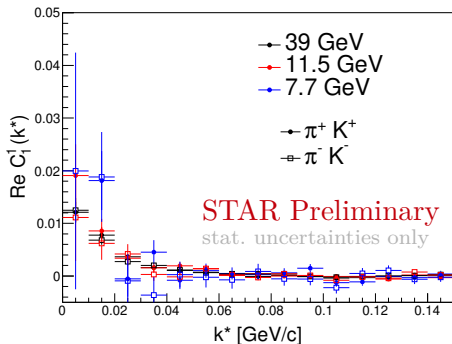
# System dependence, $\sqrt{s_{NN}} = 39$ GeV, 0-10%



- Like-sign pairs are dominated by Coulomb
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- Unlike-sign CFs are more complicated
- $\Lambda$  peak is visible in pion-proton
- Strong interaction is not negligible in Kp

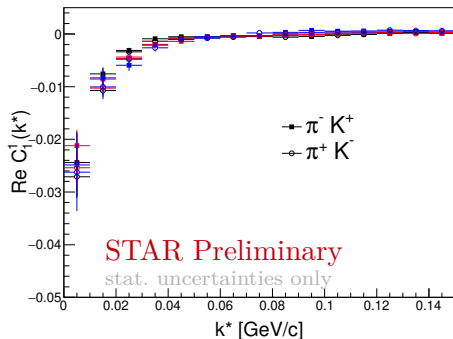
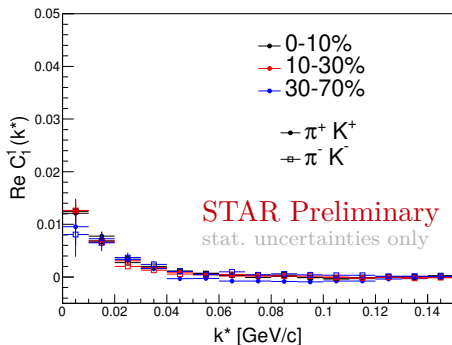
# Source dynamics — energy dependence



- Clear signal of emission asymmetry

- Visible energy dependence

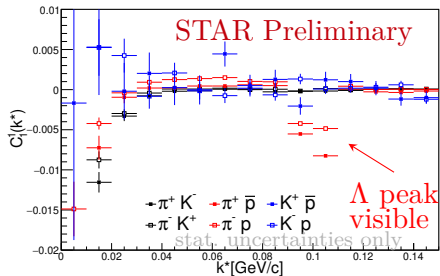
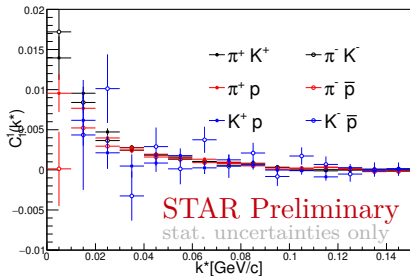
# Source dynamics — centrality dependence



- Clear signal of emission asymmetry

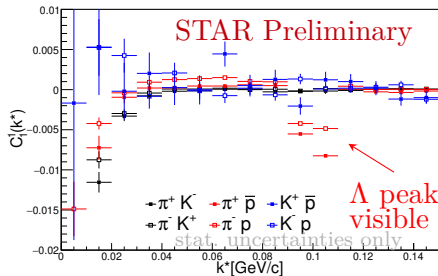
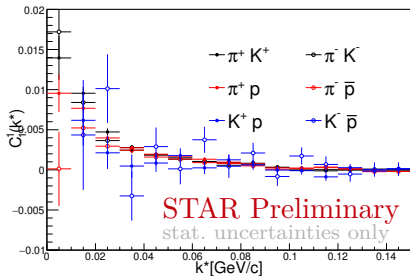
- Visible centrality dependence

# Source dynamics — system dependence



- Visible signal of emission asymmetry

# Source dynamics — system dependence



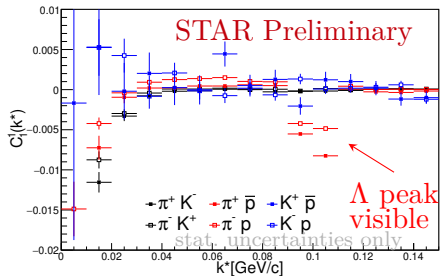
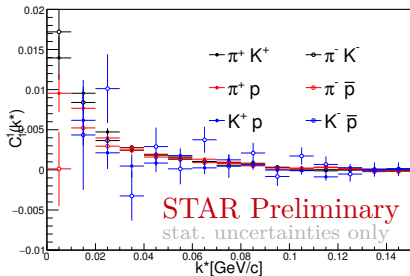
- Visible signal of emission asymmetry

- Expected ordering of particles — confirmed

Lighter particle is emitted closer to the center and/or later.

R. Lednicky, et al., Phys. Lett. B272, 20 (1996)  
STAR, Phys. Rev. Lett. 91, 262302 (2003)  
A. Kisiel, Phys. Rev. C81, 064906 (2010)

# Source dynamics — system dependence



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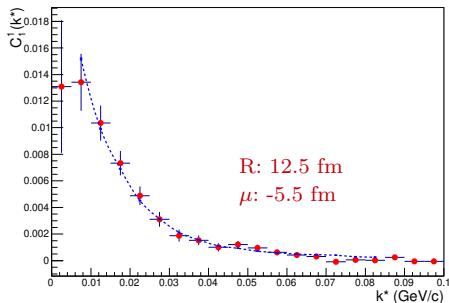
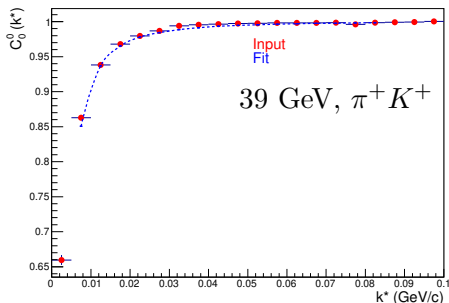
- Expected ordering of particles — confirmed

We are sensitive to collective effects



# Non-identical particle femtoscopy — future plan

- Next step:
  - ▶ Obtain source parameters using CorrFit



- **Kaon femtoscopy (geometry):**
  - ▶ neutral kaon correlations at  $\sqrt{s_{NN}} = 200$  GeV in minimum-bias events (0-80%)
  - ▶ purity correction is done
  - ▶ source sizes using Gaussian fit are obtained
  - ▶ comparison with published data (from 2006):
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- **Non-identical particle femtoscopy (geometry + dynamics):**

- ▶ Strong interaction not negligible in kaon-proton
- ▶ Clear signal of emission asymmetry for particles with different masses
- ▶ Asymmetry does not disappear for low energies