

# Prospect for Korea Neutrino Observatory (KNO)

**Intae Yu**

**Sungkyunkwan University**

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## Kamiokande (1983-1996)

3000 ton



- Neutrinos from SN1987a.
- Atmospheric neutrino deficit.
- Solar neutrinos.

## Super-Kamiokande (1996- )

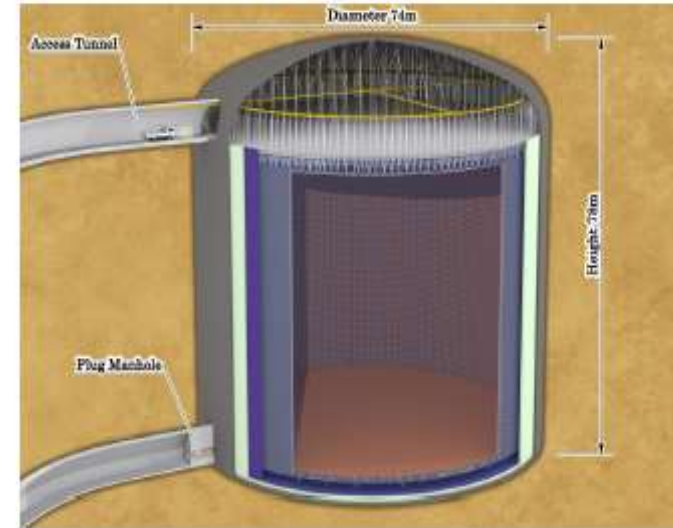
50,000 ton



- Atmospheric neutrino oscillation.
- Solar neutrino oscillation with SNO.
- Far detector for KEK-PS (K2K) and J-PARC beam (T2K): electron neutrino appearance.
- World leading limit on proton lifetime  $> 10^{34}$  years.

## Hyper-Kamiokande ( $\sim 2026-$ )

$2 \times 260,000$  ton

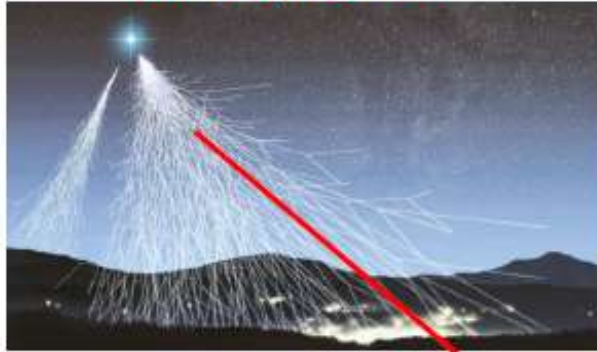


### Physics programme:

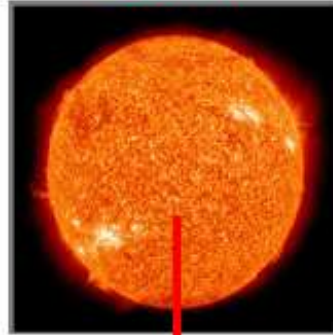
- Neutrino oscillations: Mass Hierarchy, Leptonic CP violation,  $\theta_{23}$  Octant, ...
- Nucleon decay:  $p \rightarrow e^+ \pi^0$ ,  $p \rightarrow K^+ \bar{\nu}$ , ...
- Neutrino astrophysics: Solar neutrinos, Supernova neutrinos, WIMP searches

# Overview of Hyper-K Physics

Atmospheric  $\nu$



Solar  $\nu$



Supernova  $\nu$

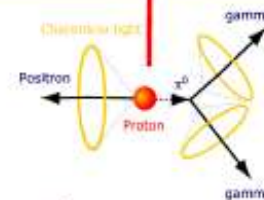
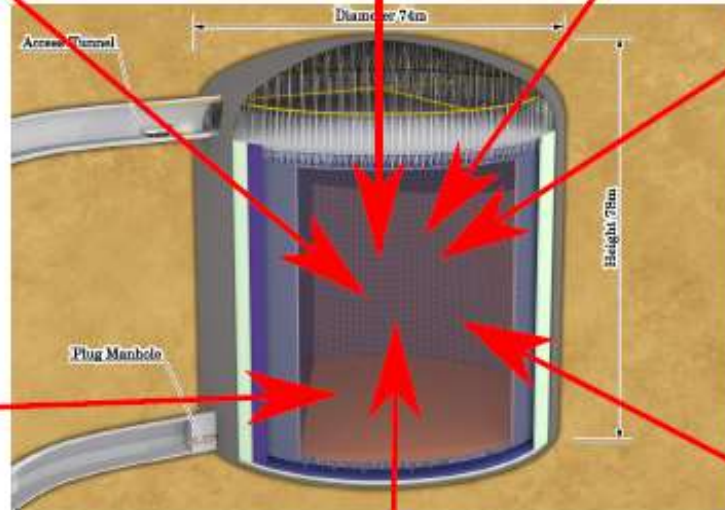


WIMP  $\chi\chi \rightarrow \nu\nu$



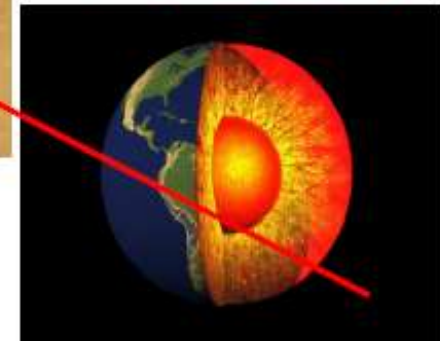
Beam  $\nu$

J-PARC



Nucleon Decay

$\nu$  Tomography

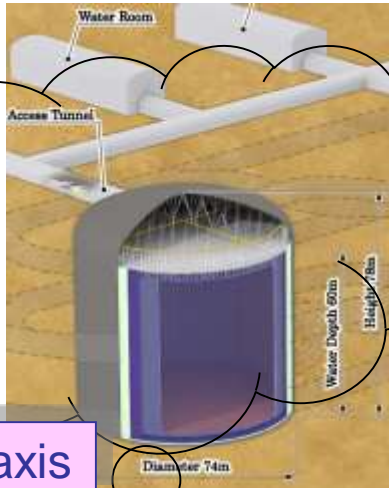


# Status of Hyper-Kamiokande

- Hyper-Kamiokande (HK) proto-collaboration was formed
- Two host institutions: U of Tokyo (ICRR), KEK (IPNS)
- U of Tokyo commitment ensures that the Hyper-K construction will begin in April, 2020.
- MEXT has made an official budget request in August and got a supplementary budget of 3.5B (32M \$) yen in FY 2019.

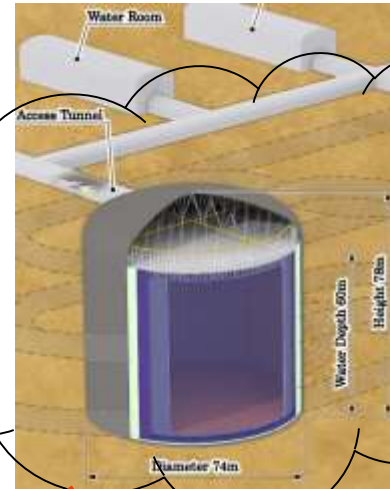
# Neutrino Detector in Korea

**KNO**



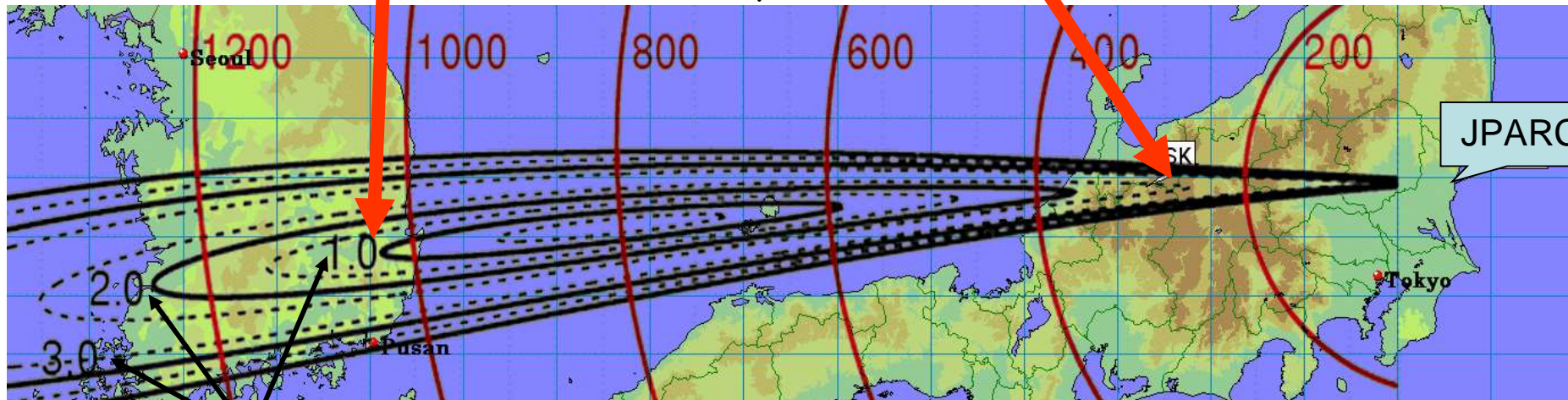
1.5 deg. off axis

**Hyper-K**



2.5 deg. off axis

**The J-PARC  $\nu_\mu$  beam comes to Korea.**



**Off-axis angle**

see hep-ph/0504061

By K. Hagiwara, N. Okamura, K. Senda

# Pros and Cons of KNO

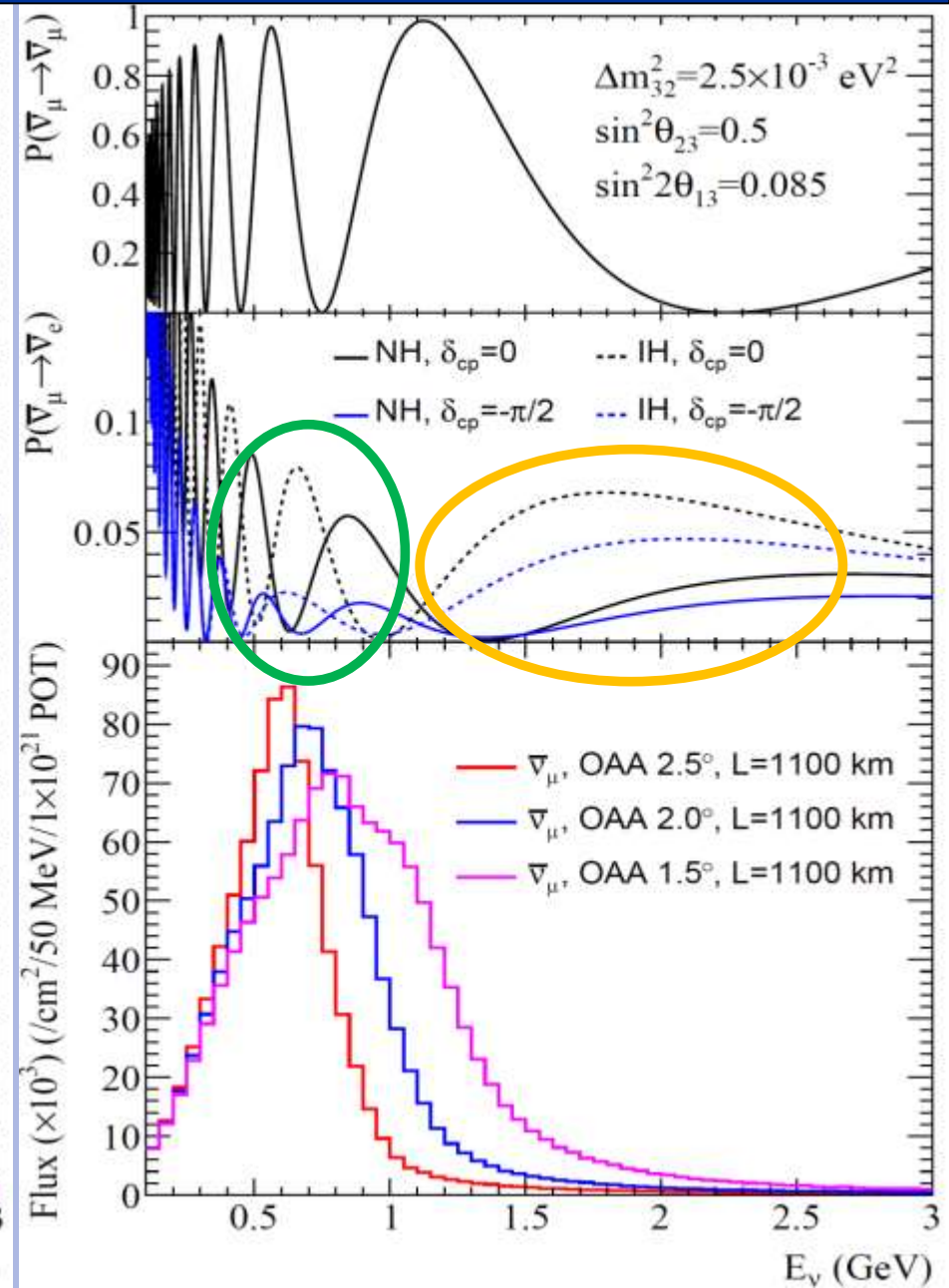
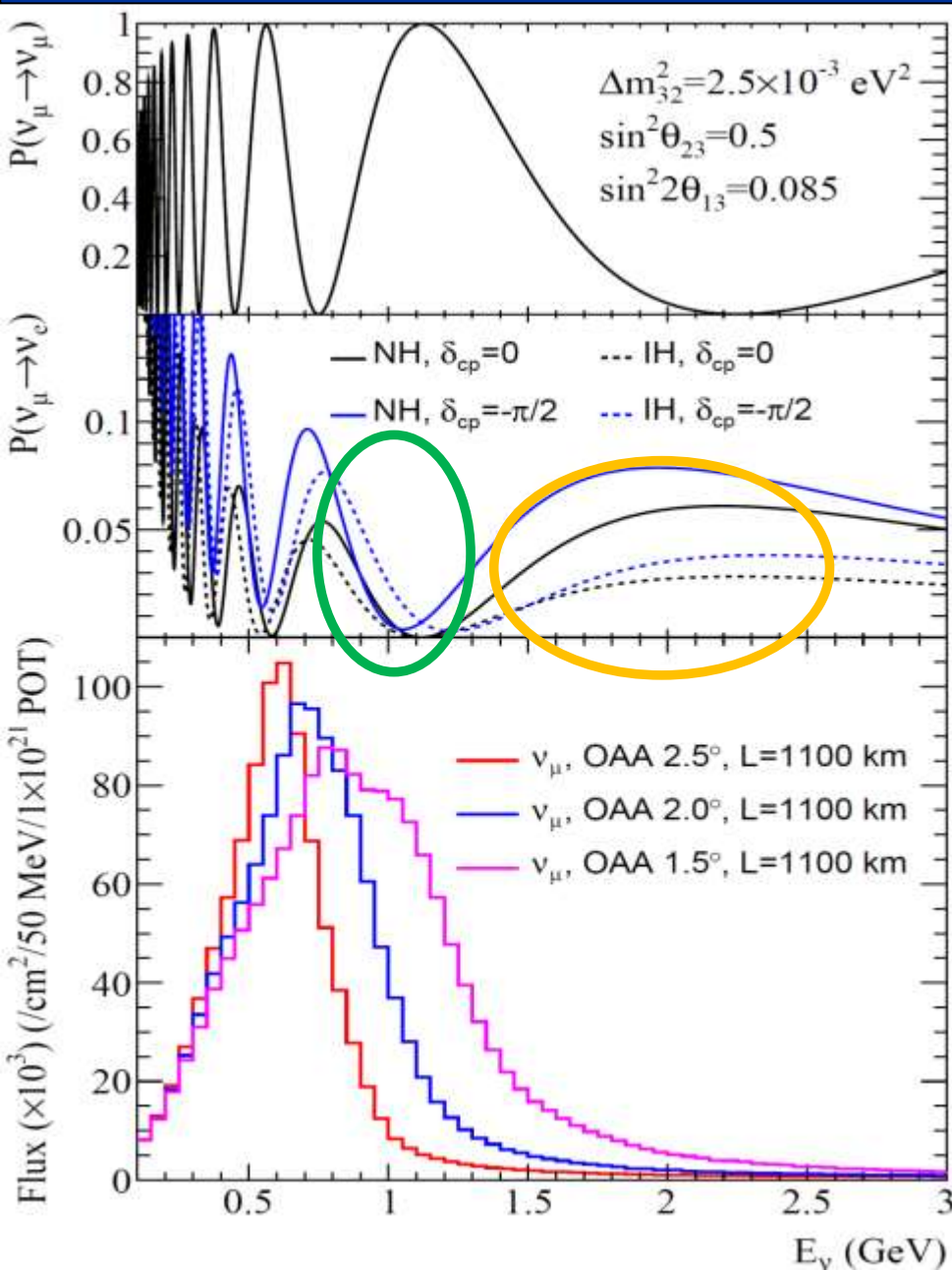
## □ Pros

- 1<sup>st</sup> and 2<sup>nd</sup> oscillation maxima at KNO → more sensitive to leptonic CP violation
- Higher mass density and longer baseline → better determination of neutrino mass hierarchy and better sensitivity to non-standard neutrino interactions
- Larger overburden (~1000 m) → better sensitivity to neutrinos of astronomical origin (solar/SN/galactic..)

## □ Cons

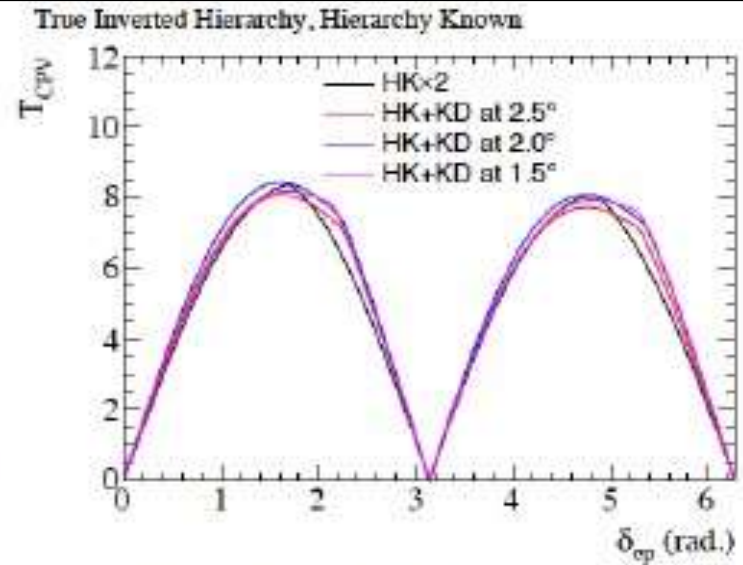
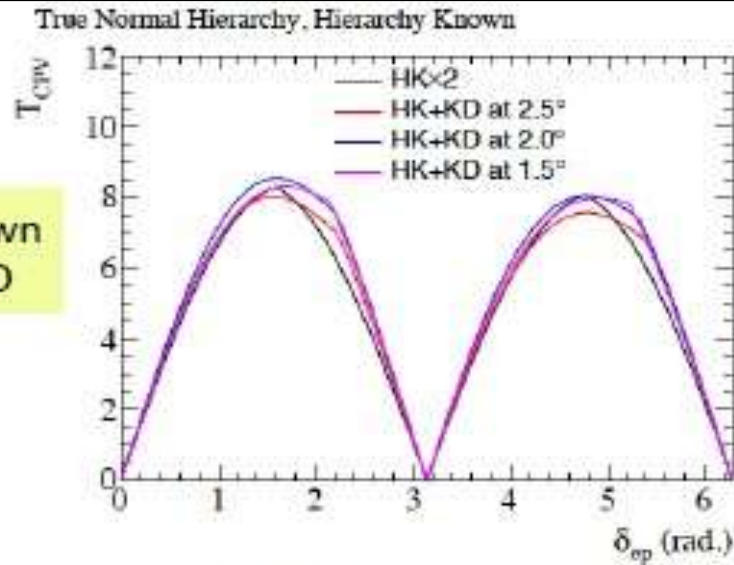
- Neutrino beam flux at KNO is ~ 10 times smaller than HK flux due to longer baseline

# 1<sup>st</sup> and 2<sup>nd</sup> Oscillation Maxima in Korea

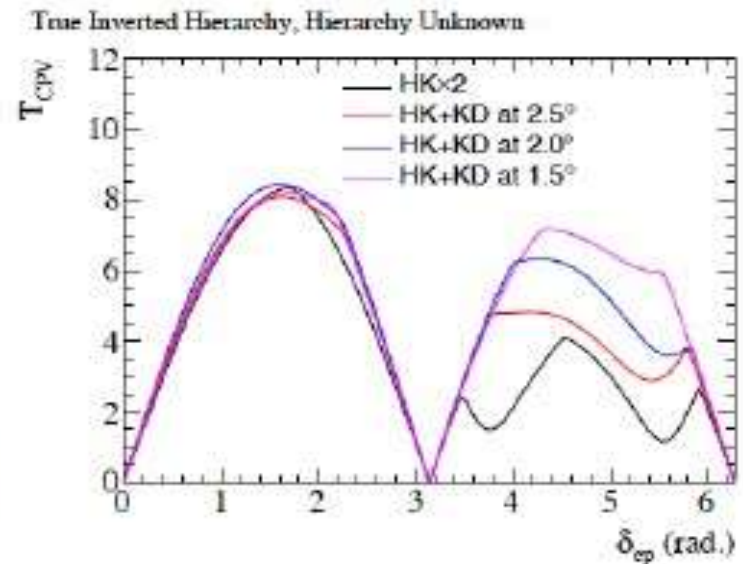
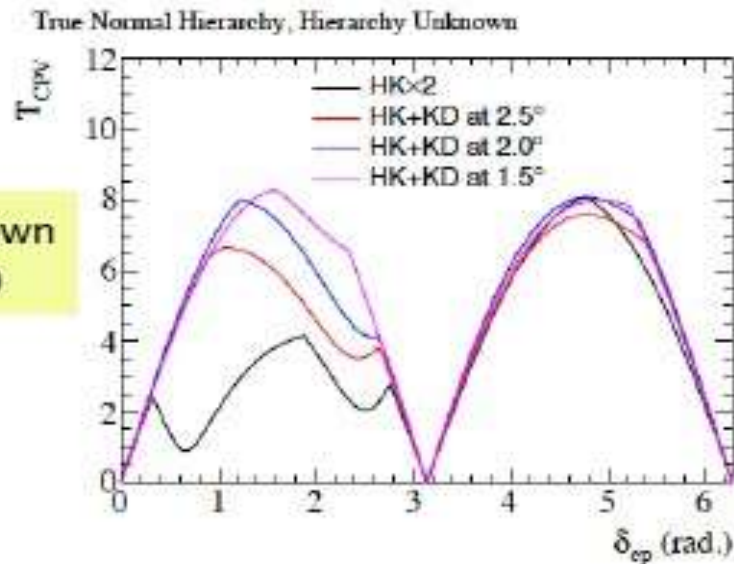


# Physics Potential at KNO: $\delta_{cp}$

Known  
MO



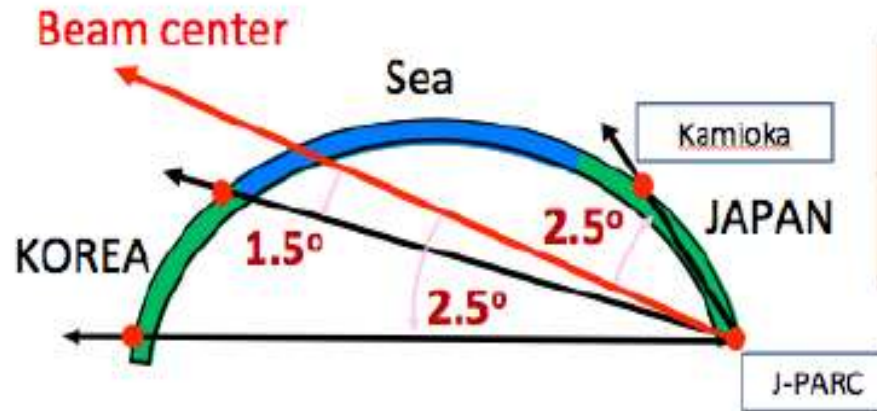
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MO



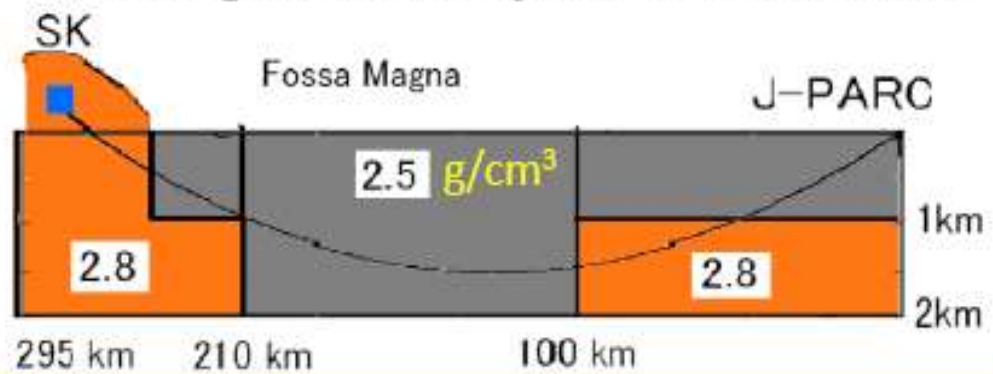
10 years of operation with 1.3 MW beam



# Matter Density Profile



Matter profile along the T2K baseline



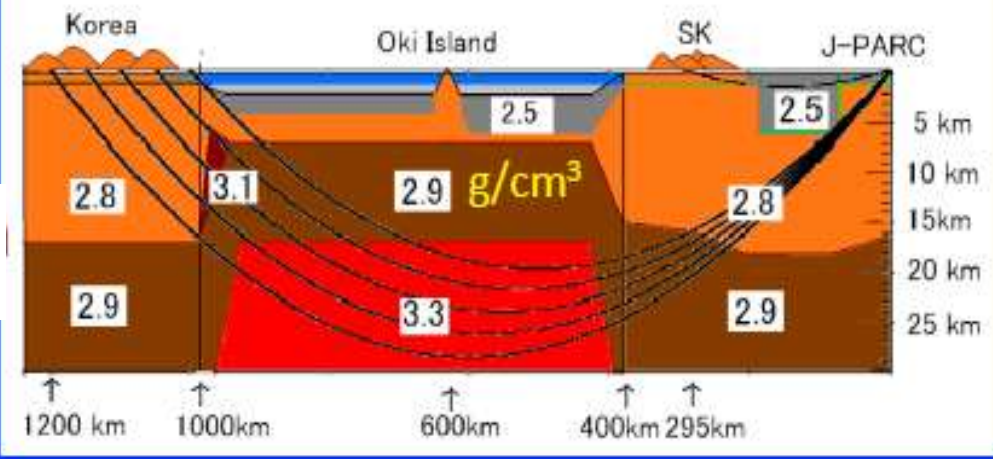
Matter density:

$$r_A = 2\sqrt{2}G_F N_e E_\nu / \Delta m_{31}^2$$

More matter effects

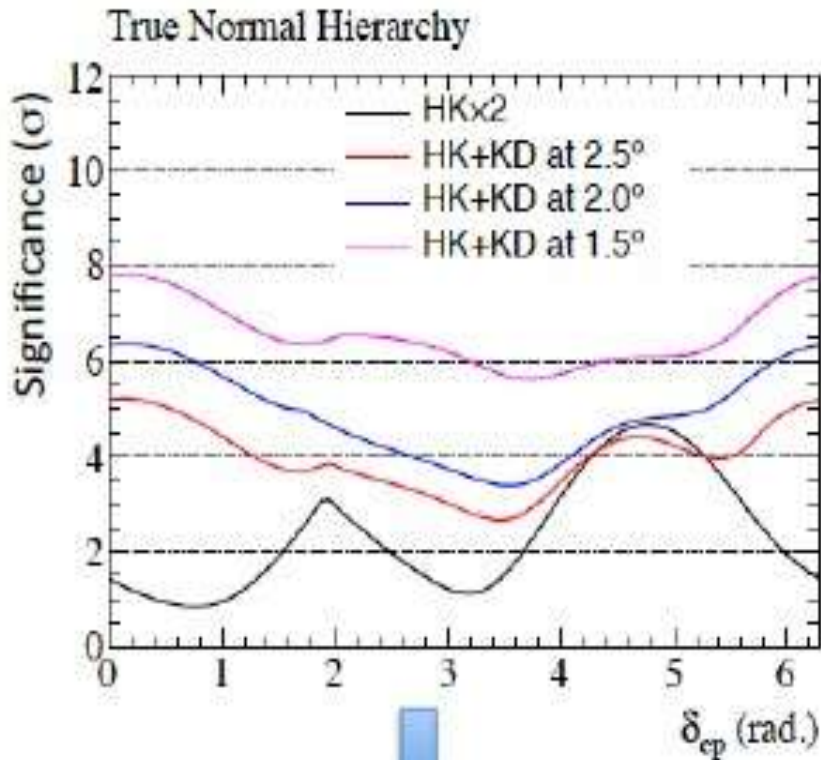
- Better Measurement of Neutrino Mass Ordering
- Longer baseline
  - Higher matter density
  - Higher neutrino energy

Matter profile along the Tokai-to-Korea baseline

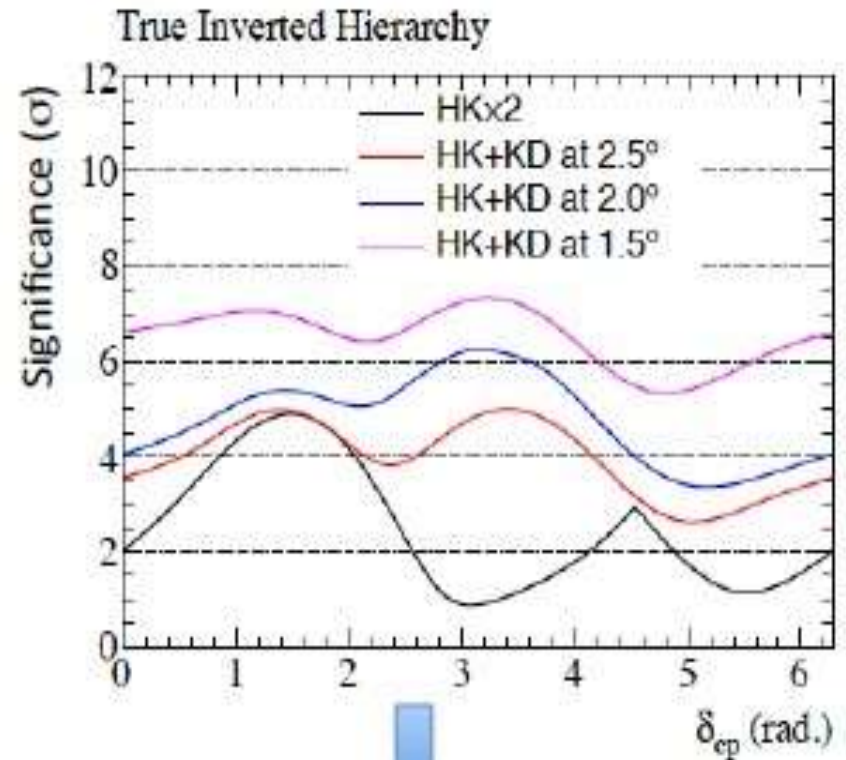


# Physics Potential at KNO: Mass Ordering

10 years of operation with 1.3 MW beam



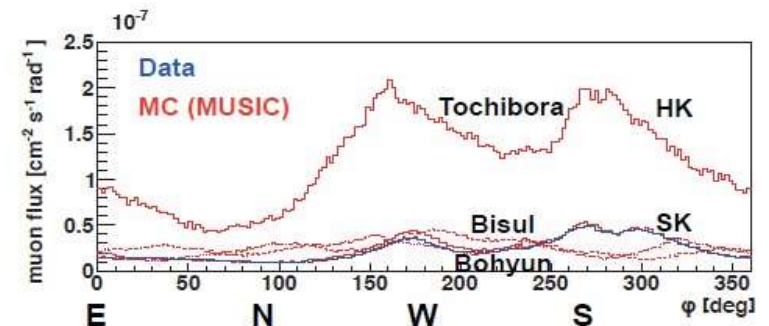
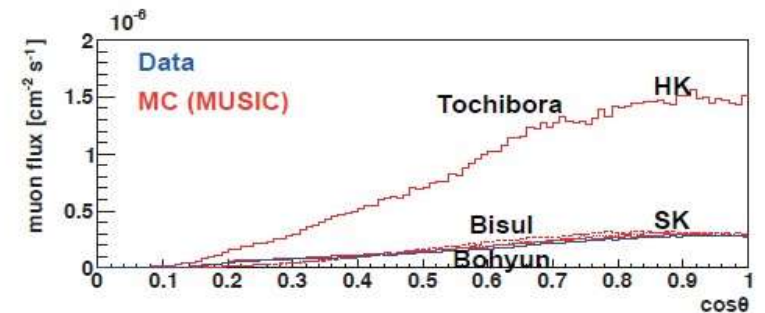
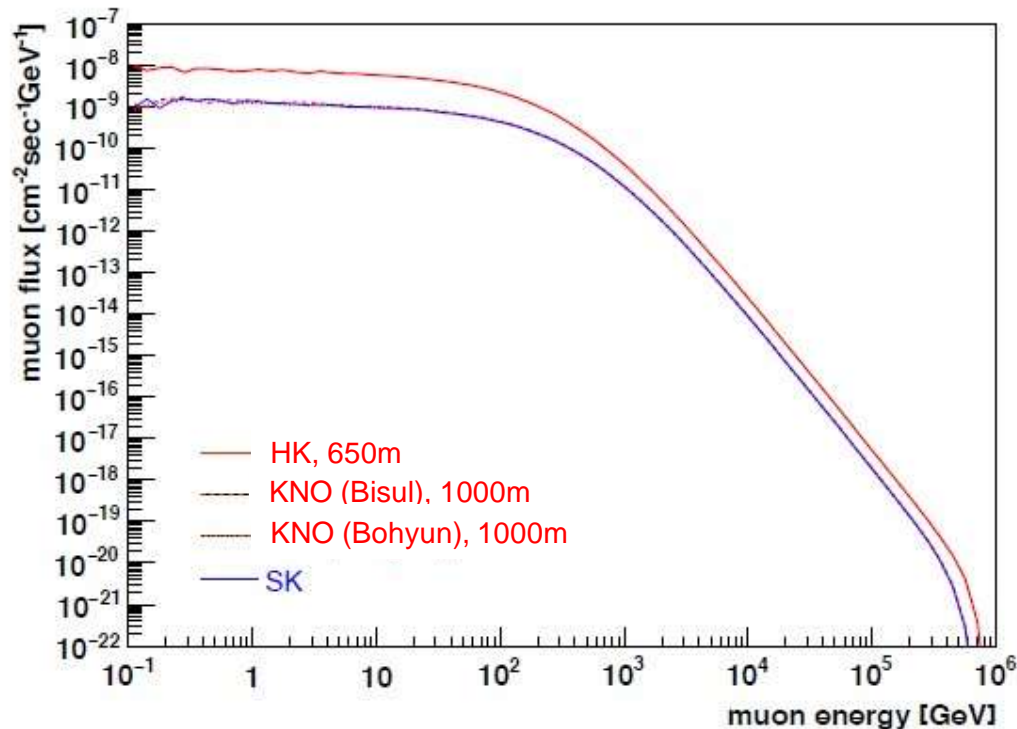
HK+KD 1.5°: 6 ~ 8  $\sigma$  for all  $\delta_{CP}$   
HK x2 : 1 ~ 4.5  $\sigma$  for all  $\delta_{CP}$   
( $< 3 \sigma$  for most cases)



HK+KD 1.5°: 5.5 ~ 7  $\sigma$  for all  $\delta_{CP}$   
HK x2 : 1 ~ 5  $\sigma$  for all  $\delta_{CP}$   
( $< 3 \sigma$  for most cases)

# Cosmogenic Muon Flux

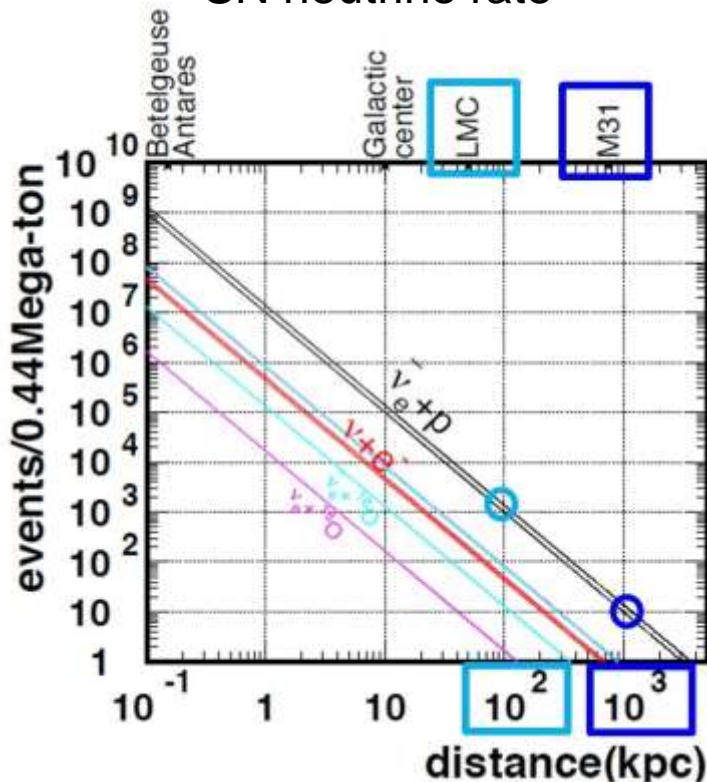
- Overburden of KNO site  $\sim 1000$  m (HK: 650 m)
- Muon flux at KNO is 5 times smaller than HK flux  $\rightarrow$  less cosmogenic backgrounds



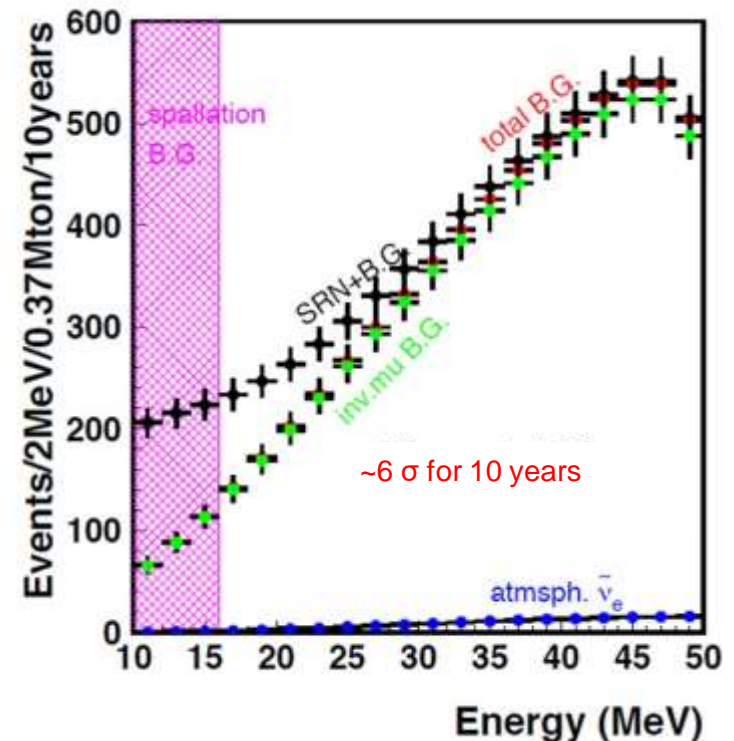
# Neutrinos of Astronomical Origins

- Super Nova Neutrinos (SN)
- Super Nova Relic Neutrinos (SRN)

SN neutrino rate



SRN spectrum



# Neutrinos of Astronomical Origins

- Neutrinos from active galactic nuclei and microquasars
- Neutrinos from interactions of cosmic protons and nuclei in the Galaxy
- Neutrinos from gamma-ray bursts (GRB)
- Neutrinos from clusters of galaxies
- Neutrinos from dark matter decays
- Solar Neutrinos

Multi-Messenger  
Astronomy

# History of KNO/T2HKK

- Oct. 17, 2000: Another far detector using a JHF neutrino beam by S.B. Kim (KOSEF-JSPS Joint Seminar at KIAS)
- 2005/2006/2007: A large Cherenkov detector in Korea using a J-PARC neutrino beam (T2KK) by T. Kajita.  
→ 3 joint workshops supported by KOSEF and JSPS
- 2011: Proposal of 0.5 M ton water Cherenkov Hyper-Kamiokande detector at Kamioka (LOI as arXiv:1109.3262 and arXiv:1412.4673v2)
- 2015: Staged construction of two HK detectors of each 0.26 Mton at Kamioka
- July 10, 2016: The first T2HKK meeting in London  
→ present a proposal to the HK collaboration  
→ T2HKK working group (S. Seo)

# Activities of KNO/T2HKK

- Sep. 2, 2016: First Workshop on T2HKK in Korea (SNU)
- Oct. 20, 2016: Pioneering Symposium at Korean Physical Society meeting (Gwangju)
- Nov. 2016: A white report on T2HKK released. It was published in Prog. Theor. Exp. Phys. 2018, 063C01.
- Nov. 21-22, 2016: International Workshop on 2<sup>nd</sup> Detector in Korea (SNU)
- Nov. 24, 2017: 1<sup>st</sup> KNO Workshop (KNU)
- Aug. 21, 2018: 2<sup>nd</sup> KNO Workshop (KASI)
- Nov. 2, 2018: 3<sup>rd</sup> KNO Workshop (KNU)
- Aug. 25, 2019: KNO Workshop with NUFACT 2019 (KNU)

# Korean Efforts on KNO Realization

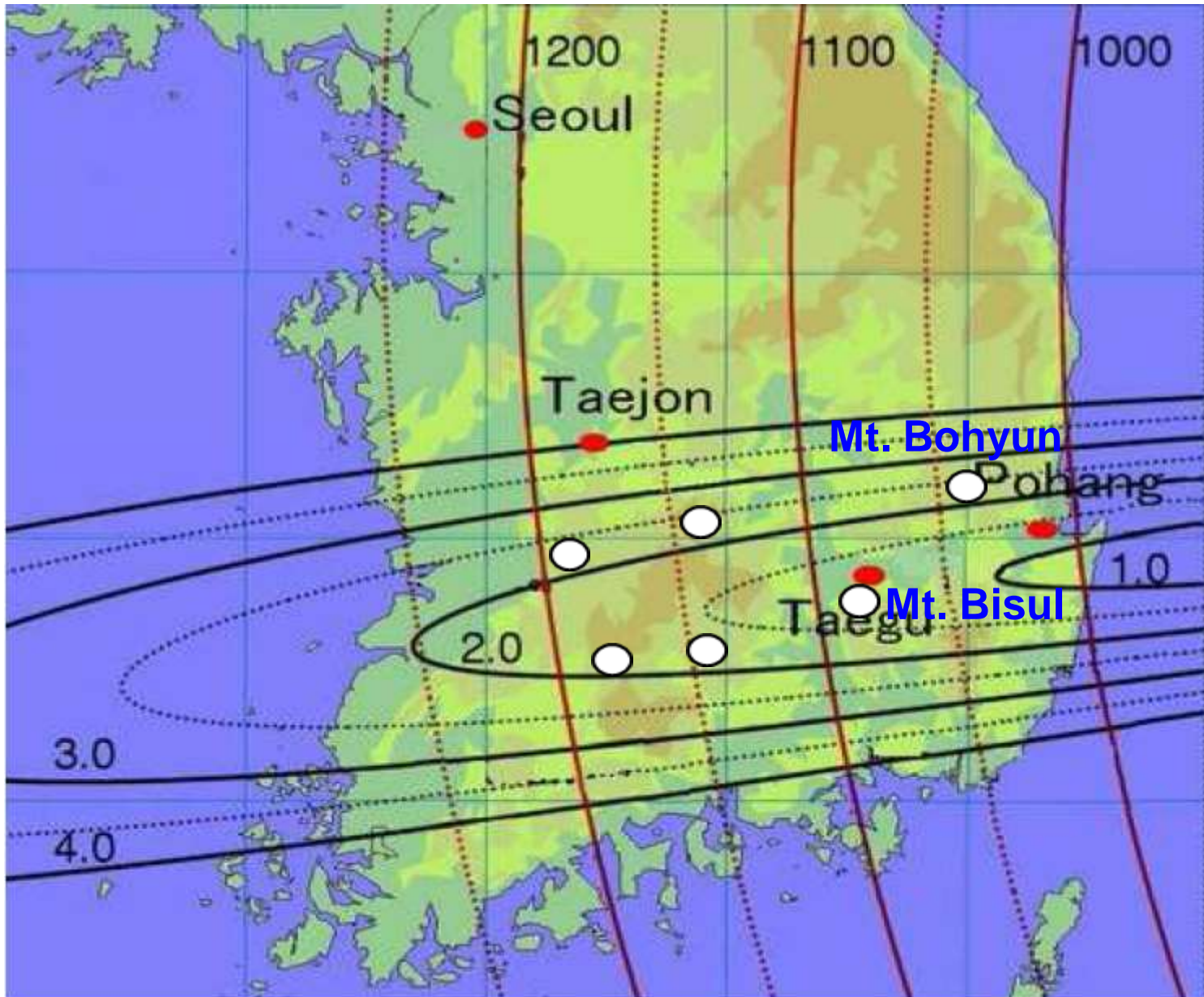
- 2018. 10. 20: Kick-off Meeting for KNO organization including physicists and astronomers
- Five working groups were formed in the meeting
- Each working group has held regular meetings
- Discussions with Korean government have been started
- Detector R&D work is in progress
- Several options for KNO detector are being considered
- Korean efforts are in very early stage



# Working Groups of KNO Organization

- Government Relations Working Group  
contact and discussions with government and funding agency
- Detector R&D Working Group  
photo sensor, water purification, DAQ, and etc
- Science Working Group  
particle physics and astronomy subgroups
- Proposal Working Group  
preparation for KNO proposals
- International Relations Working Group  
foreign support and participation

# KNO Candidate Sites



# List of KNO Candidate Sites

S. B. Kim (SNU)

Site	Height (m)	Baseline (km)	Off-axis angle (degree)	Elements of rock
Mt. Bisul	1084	1088	1.3°	Granite porphyry, Andesitic breccia
Mt. Hwangmae	1113	1140	1.8°	Flake granite, Porphyritic gneiss
Mt. Sambong	1186	1180	1.9°	Porphyritic granite, Biotite gneiss
Mt. Bohyun	1124	1040	2.2°	Granite, Volcanic rocks, Volcanic breccia
Mt. Minjuji	1242	1140	2.2°	Granite, Biotite gneiss
Mt. Unjang	1125	1190	2.2°	Rhyolite, Granite porphyry, Quartz porphyry

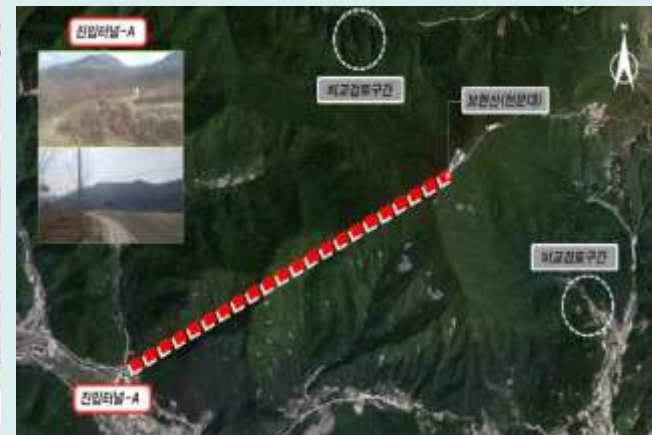
# KNO Candidate Sites – Mt. Bisul and Mt. Bohyun

S. B. Kim (SNU)

## Mt. Bisul

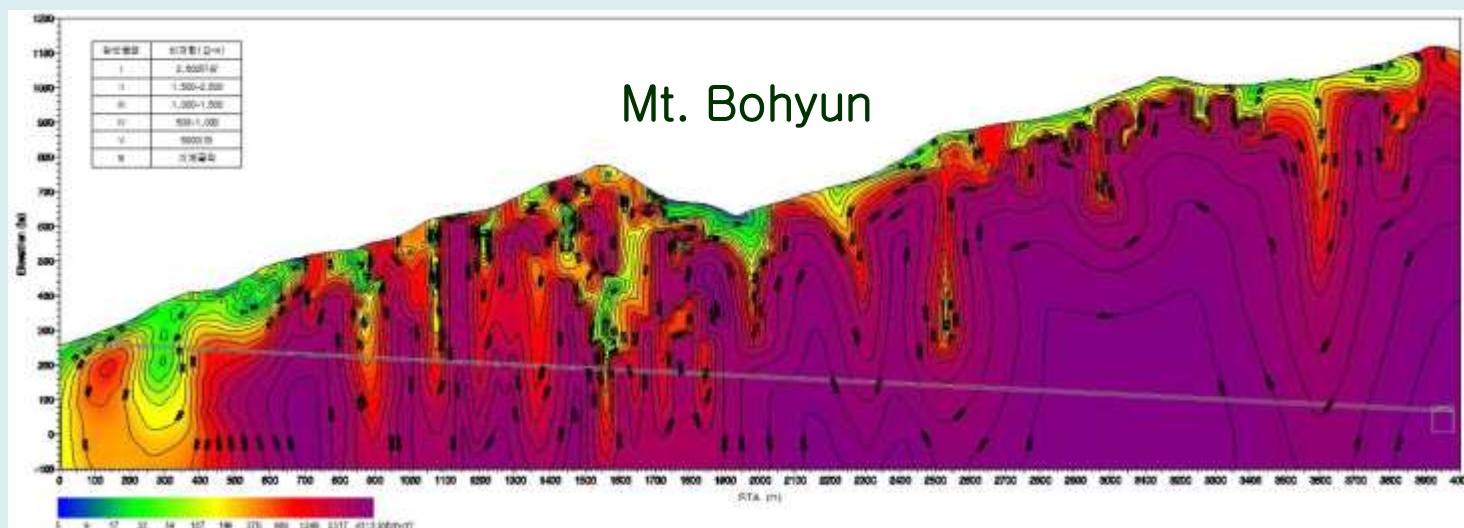
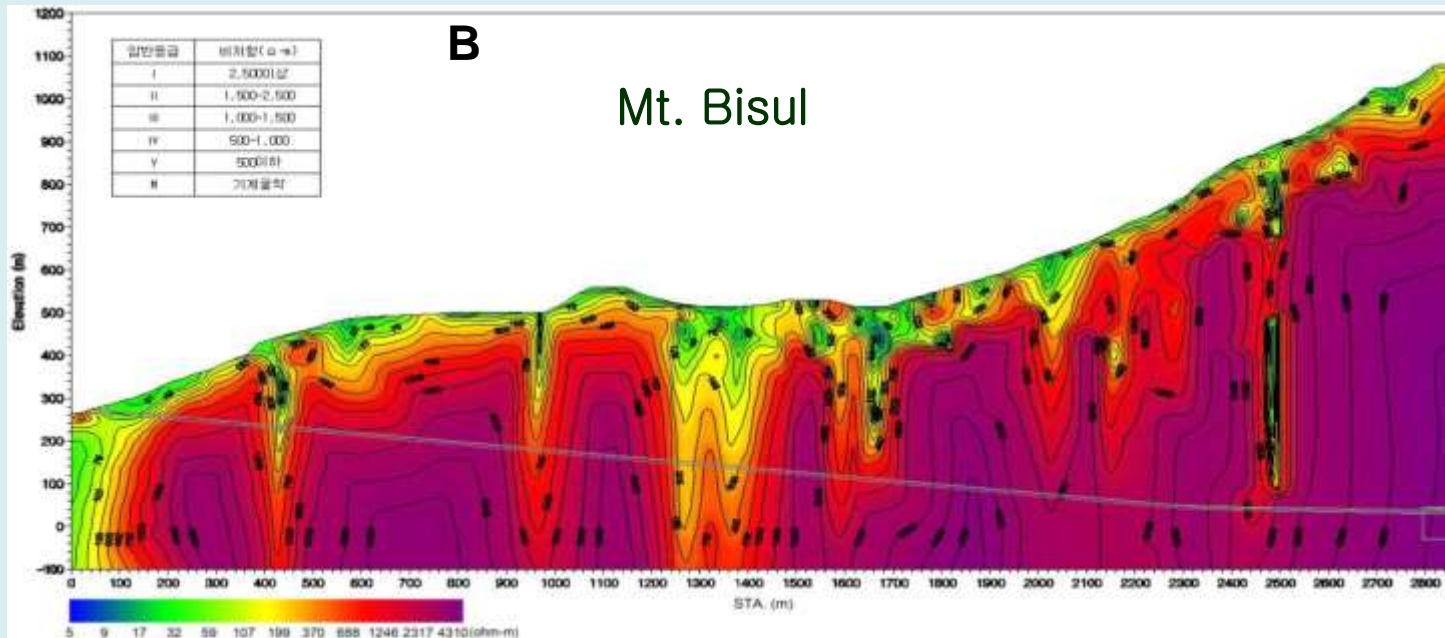


## Mt. Bohyun



# Bedrock Investigation of KNO Sites

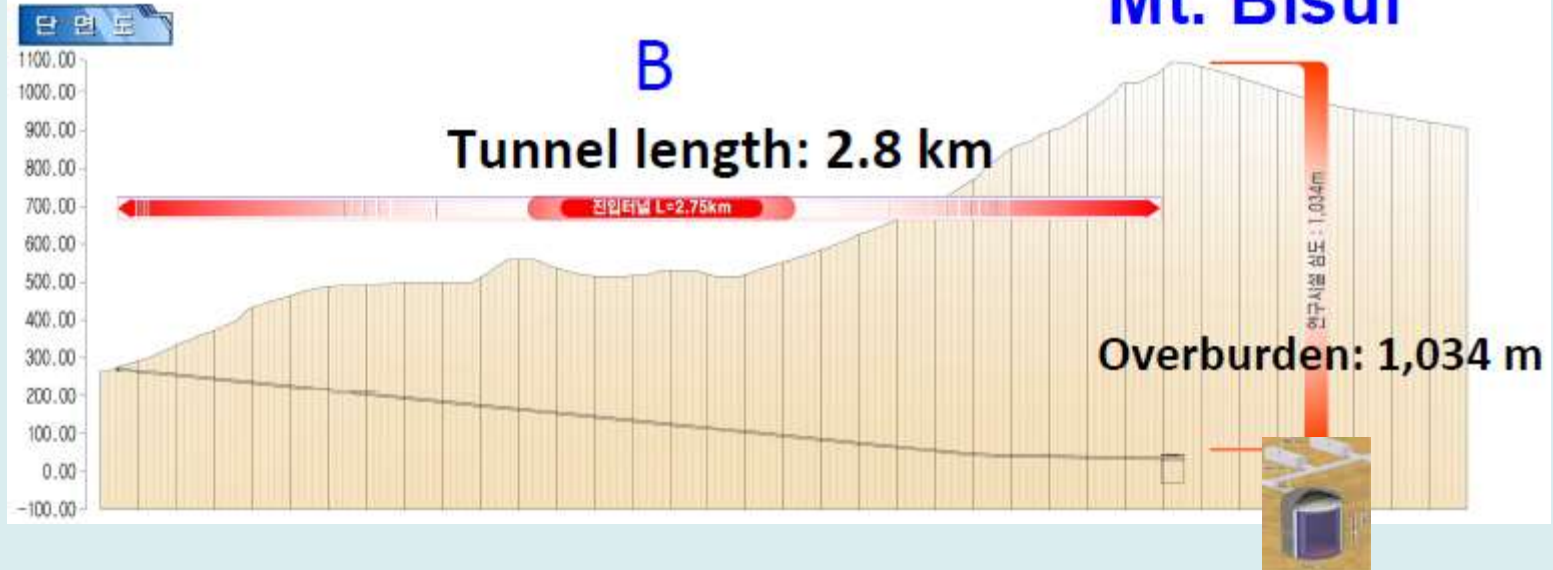
S. B. Kim (SNU)



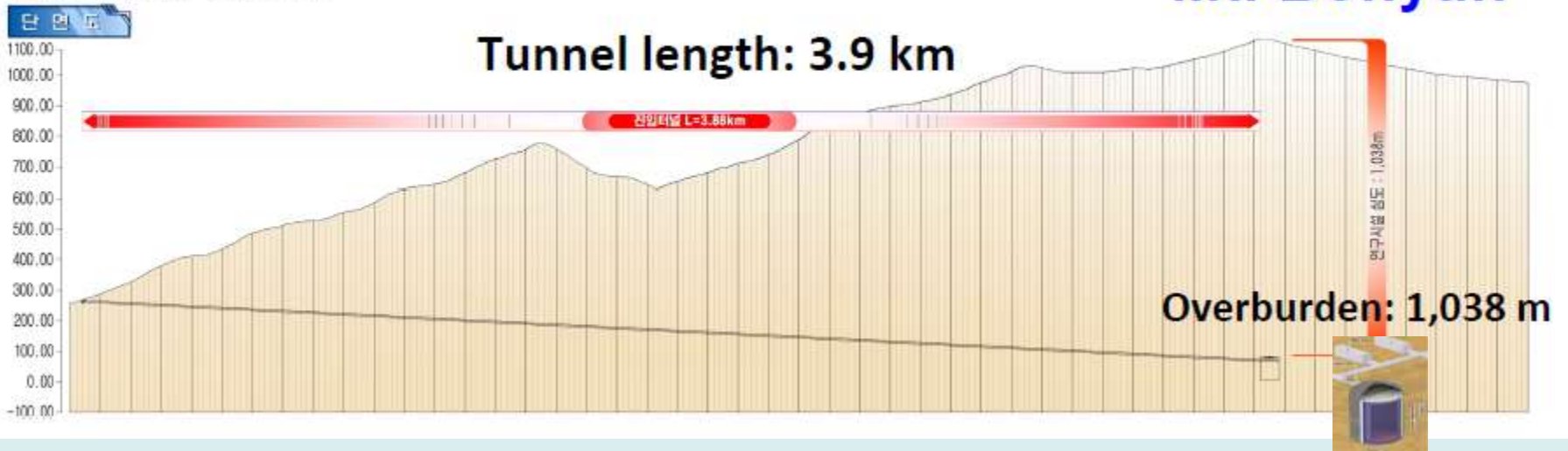
# Conceptual Design of KNO Tunnel

S. B. Kim (SNU)

B구간 진입시 종단면도

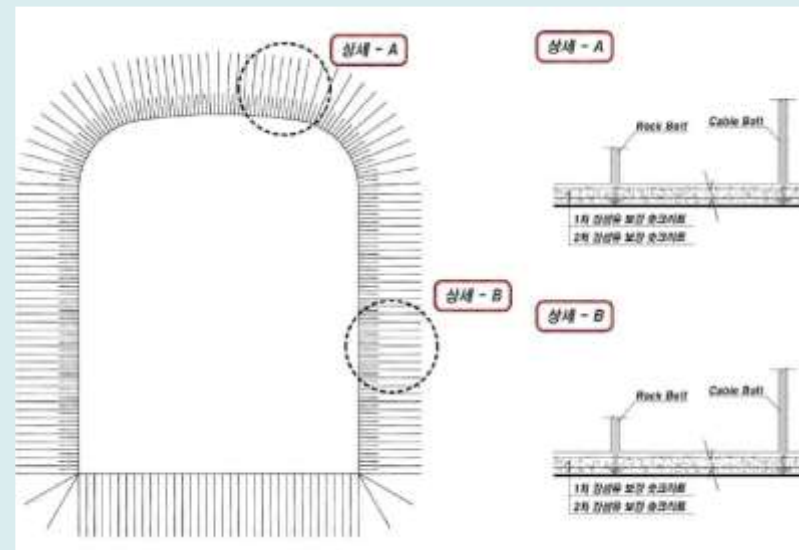
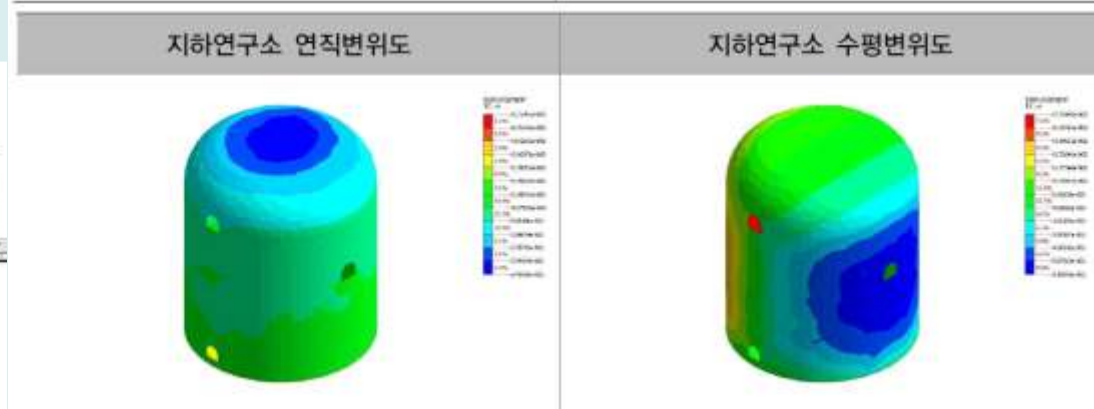
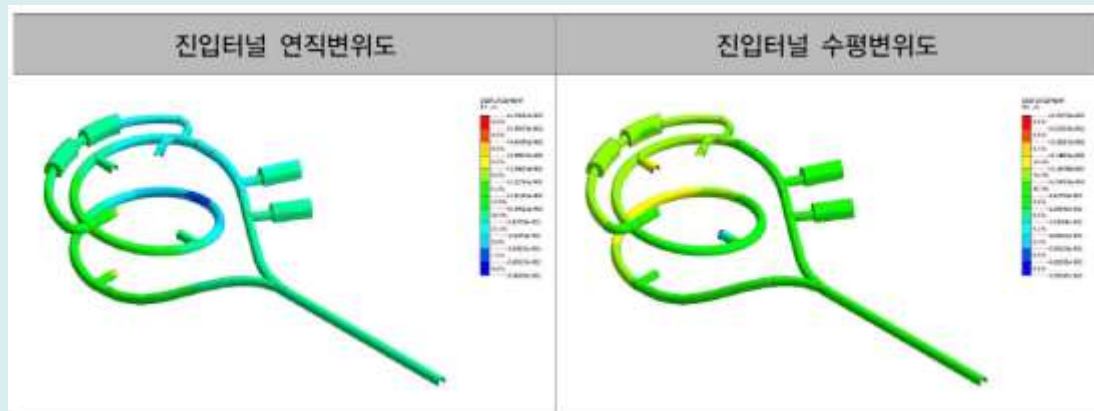
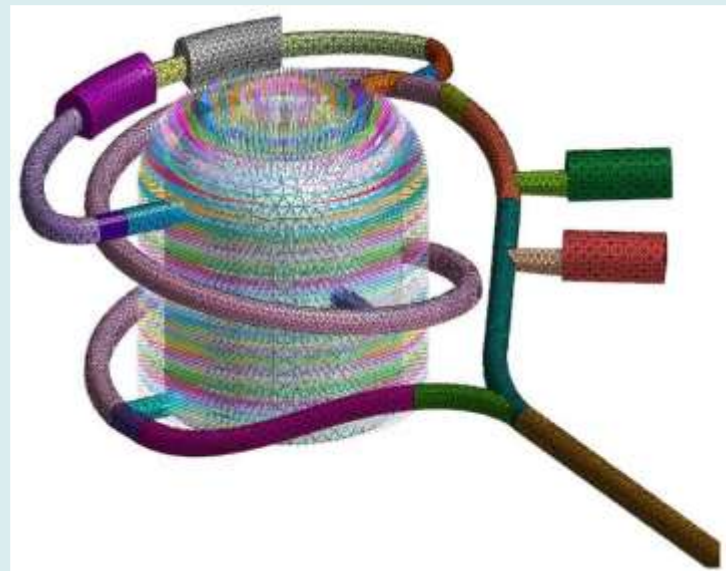


A구간 진입시 종단면도



# Stress Analysis and Reinforcement

S. B. Kim (SNU)



# Activities on Detector R&D

- Two independent approaches of photo sensor R&D
- Development of conventional PMT
  - University of Seoul in collaboration with Korean company MECARO
  - Work on 3 inch PMT first and move to larger PMT
- Development of Silicon PMT
  - Kyungpook National University in collaboration with Russian group
  - Hybrid PMT using photocathode, scintillator, and SiPM



# Activities on Detector R&D

- Development of water purification system
  - Seoul National University in collaboration with Korean company DICOTECH
  - prototype construction of radon vacuum degassifier
  - development of high-sensitive radon measurement device
  
- Frontend electronics R&D
  - Korean company NOTICE sells FADC modules
  - preliminary evaluation in progress

# Activities on KNO Science

- Particle physics subgroup identifies potential KNO physics topics through workshops and seminars
  - organize Korean neutrino meetings
  - carry out sensitivity studies using simplified simulations (published in PTEP 2018)
  
- Astronomy subgroup is preparing for a white paper on KNO astronomy
  - list of potential KNO astronomy topics
  - emphasis on multi-messenger astronomy using neutrinos

# Summary and Prospect

- KNO greatly enhances physics sensitivities in the measurements of leptonic CP violation, mass ordering, proton decay, NSI, and many others
- KNO also serves as a powerful neutrino telescope for multi-messenger astronomy
- KNO organization and working groups are formed and active
- Efforts on detector R&D and science are in progress
- KNO can be a flagship project for Korean HEP for the next 10 years

Thank you

**BACK UP**

# KNO Strategy

