

Anomalies, CPT and Leptogenesis

Sarben Sarkar

Dept. of Physics, King's College London, UK

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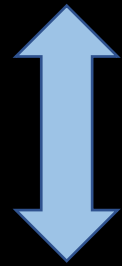
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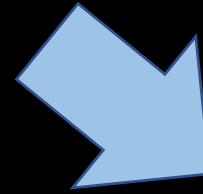
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Matter Antimatter Asymmetry

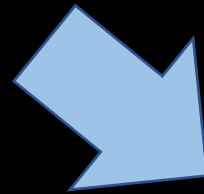
Baryon number asymmetry



Lepton Number asymmetry



Baryogenesis




Leptogenesis

DATA

- ρ_0 component energy density
- ρ_c critical energy density
- $\Omega_{matter} = 0.27 \pm 0.04$
- $\Omega_B = 0.044 \pm 0.004$

DATA

- n_B baryon number density
- n_γ photon number density
- $\frac{n_B}{n_\gamma} = 6.1 \pm 0.3 \times 10^{-10}$ at $T \sim 1 \text{ GeV}$
- matter-antimatter symmetry
 $\frac{n_B}{n_\gamma} = \frac{n_{\bar{B}}}{n_\gamma} \approx 10^{-18}$

How to get Baryogenesis in the Standard Model ?

- Since CPT invariance is assumed need non equilibrium physics
- Amount of CP violation not sufficient
- Beyond-the-Standard-Model physics is required

C(harge conjugation)P(arity)T(ime reversal)Theorem

Let $\Theta = CPT$

Lagrangian density = $\mathcal{L}(x)$

For any Lorentz invariant Hermitian local Lagrangian

$$\Theta \mathcal{L}(x) \Theta^{-1} = \mathcal{L}^\dagger(-x)$$

Our BSM model for Leptogenesis

Spontaneously Broken Lorentz invariance

Spontaneously Broken CPT invariance

Sterile neutrino

String inspired gravitational degrees of freedom

Leptogenesis effective Lagrangian

$$\mathcal{L} = \mathcal{L}_{SM} + i \bar{N} \gamma^\mu \partial_\mu N - \frac{m_N}{2} (\bar{N}^c N + \bar{N} N^c) - \bar{N} \gamma^\mu B_\mu \gamma^5 N - \sum_k y_k \bar{L}_k \tilde{\phi} N + h.c.$$

where N is a right-handed Majorana neutrino with heavy mass m_N

$\bar{N} \gamma^\mu B_\mu \gamma^5 N$ is CPTV and LIV

Features of model

Majorana nature of N leads to acceptable leptogenesis

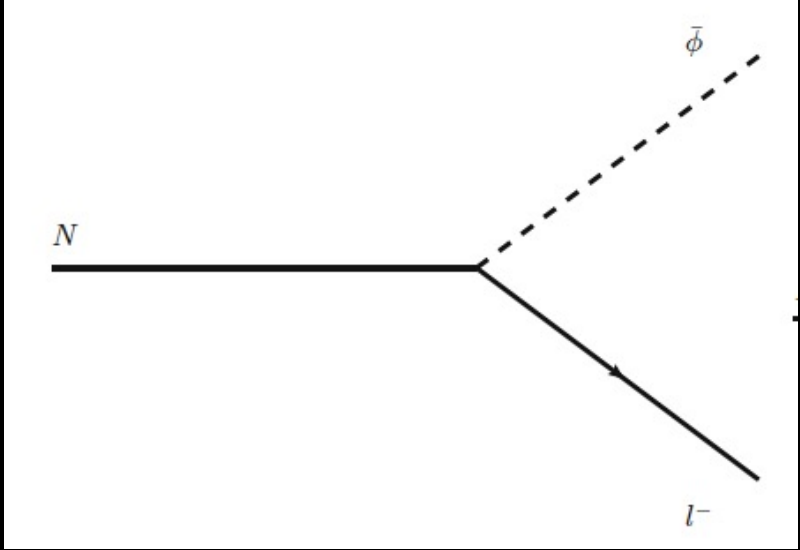
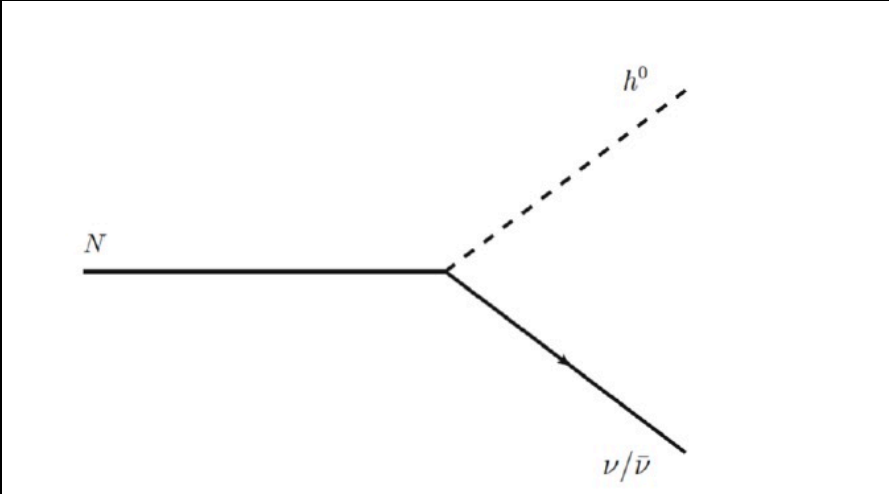
Very heavy N leads to freeze out at TeV scale

B-L conservation converts ΔL to ΔB

Leptogenesis occurs at tree level

CP and CPT violating decays

- Background $B_i = \mathbf{0}, \quad i = 1, 2, 3$
 $B_0(\mathbf{z}) = \Phi f(\mathbf{z})$
 $\mathbf{z} = \frac{m_N}{T}, \text{ and } f(\mathbf{z}) = 1 \text{ or } f(\mathbf{z}) = \mathbf{z}^{-3}$
- Tree-level decays:
 - $N \rightarrow l^- h^+, \nu h^0$ channel 1
 - $N \rightarrow l^+ h^-, \bar{\nu} h^0$ channel 2



CP violation through CPTV: decay rates

Channel 1 $N \rightarrow l^- \bar{\phi}$

Channel 2 $N \rightarrow l^+ \phi$

$$\Gamma_1 = \sum_k \frac{|Y_k|^2}{32\pi^2} \frac{m^2}{\Omega} \frac{\Omega + B_0}{\Omega - B_0},$$

$$\Gamma_1 \neq \Gamma_2$$

$$\Omega = \sqrt{B_0^2 + m^2}.$$

$$\Gamma_2 = \sum_k \frac{|Y_k|^2}{32\pi^2} \frac{m^2}{\Omega} \frac{\Omega - B_0}{\Omega + B_0}.$$

Lepton asymmetry

$$\frac{\Delta L}{n_\gamma} \simeq 10^{-10},$$



$$\frac{B_0}{m} \simeq 10^{-8}$$

Boltzmann
analysis

$$\frac{\Delta L^{tot}}{s} \simeq .017 \frac{B_0}{m_N}$$

at freezeout temperature $T = T_D$

$$\frac{m_N}{T_D} \simeq 1.6$$

$$m_N = O(100) TeV$$

$$\text{For } y \sim 10^{-5}$$

$$B_0 \sim 0.1 MeV$$

$$T_D \sim 60 TeV$$

String inspired microscopic origin of B_μ

1. Gravitational multiplet in bosonic string theory: $g_{\mu\nu}, H_{\mu\nu\lambda}, \phi$

$$2. S = \frac{1}{2\kappa^2} \int \sqrt{g} [R - e^{-4\phi} H_{\mu\nu\lambda} H^{\mu\nu\lambda}]$$

$$3. \bar{\Gamma}_{\mu\nu}^\lambda = \Gamma_{\mu\nu}^\lambda + \sqrt{\kappa^2/3} g^{\lambda\rho} H_{\mu\nu\rho} \text{ to } O(\alpha') \longrightarrow R(\bar{\Gamma})$$

4. Contorsion $H_{\mu\nu\rho}$

$$H_{\mu\nu\rho} = \partial_{[\mu} B_{\nu\rho]}$$

$$e^{-2\Phi} \varepsilon_{bac}^\mu H^{abc} = 4\partial^\mu b(x) = 4B_\mu(x)$$

axion

Fermions in torsion background

- Gravity gauge theory local Lorentz invariance

- Spin connection \longleftrightarrow gauge field

- Spin connection \supset torsion

- Fermions \longrightarrow torsion \longrightarrow gauge D_μ

- Fermion-torsion coupling: $\int B_\mu \bar{\psi} \gamma^\mu \gamma^5 \psi \sqrt{-g} d^4 x$

$$B_d = \frac{1}{3!} \varepsilon^{abc}_d T_{abc} \quad T_{abc} \text{ torsion tensor.}$$

Background for B_μ

- $B_\mu = \partial_\mu b$ (4D)

- $b(x) = \bar{b}(x) + \tilde{b}(x)$

background

fluctuation

- (1) $\frac{d}{dt} \bar{b} = \text{constant}$

- (2) $\partial_\mu \left[\sqrt{g} \left(\frac{8}{3\kappa^2} \partial^\mu \bar{b} - J^{5\mu} \right) \right] = 0$

Background solutions

$$1. B_0 = \text{constant} \quad B_i = 0$$

$$2. \langle J^5 \ 0 \rangle \neq 0 \implies B_0 = \frac{3\kappa^2}{8} \langle J^5 \ 0 \rangle$$

Current constraints

$$|B_0| < .01 \text{ eV}$$

$$|B_i| < 10^{-31} \text{ GeV}$$

What if ?

$$\langle J^5 \quad 0 \rangle = 0 \quad T_D \sim 10^5 \text{ GeV}$$



$$B_0 \sim T^3 \quad T \leq T_D$$

$$B_0(T_D \approx 10^5 \text{ GeV}) = 0.1 \text{ MeV}$$

$$\Rightarrow B_0(T_0 \approx 0.23 \text{ meV}) = O(10^{-57}) \text{ GeV}$$

Technical details

Classical torsionful spacetime manifold

$$d\underline{e}^a + \underline{\omega}^a_b \wedge \underline{e}^b = \underline{\mathfrak{T}}^a$$

$$d\underline{\omega}^a_c + \underline{\omega}^a_b \wedge \underline{\omega}^b_c = \underline{R}^a_c$$

\underline{e}^a is the vierbein

$\underline{\omega}^a_b$ is the spin-connection

fermions of SM on a torsionful manifold

$$S = S_{gravity} - \frac{1}{2} \sum_f \int (\bar{\Psi}_f \underline{\gamma} \wedge * \mathcal{D}\Psi_f - \mathcal{D}\bar{\Psi}_f \wedge * \underline{\gamma} \Psi_f) - \frac{1}{2} \int \mathcal{F} \wedge * \mathcal{F} - \int Tr[G \wedge * G]$$

$$\mathcal{D}\Psi = d\Psi + \frac{1}{4} \underline{\omega}^{ab} \gamma_{ab} \Psi + ie \underline{\mathcal{A}} \Psi + ig \underline{\mathcal{B}} \Psi$$

$$S_{gravity} = \frac{1}{4\kappa^2} \int \epsilon_{abcd} \underline{\mathcal{R}}^{ab} \wedge \underline{e}^c \wedge \underline{e}^d$$

$$\underline{\gamma} = \gamma_a \underline{e}^a \text{ and } \gamma_{ab} = \frac{1}{2} [\gamma_a, \gamma_b]$$

\mathcal{A} and \mathcal{B} the gauge connection 1-forms

Equation of motion with respect to $\underline{\omega}^{ab}$

$$\mathcal{T}_{abc} = -\frac{\kappa^2}{2} \varepsilon_{abcd} \sum_f J_f^{5d}$$

$$J_f^{5d} = i\bar{\Psi}_f \gamma^d \gamma^5 \Psi_f$$

Anomalies

axial anomaly

$$d * \underline{J}^5 = -\frac{\alpha_{em} Q^2}{\pi} \mathcal{F} \wedge \mathcal{F} - \frac{\alpha_s N_f}{2\pi} \text{Tr}[G \wedge G] - \frac{N_f}{8\pi^2} \tilde{R}^{ab} \wedge \tilde{R}^{ab}$$

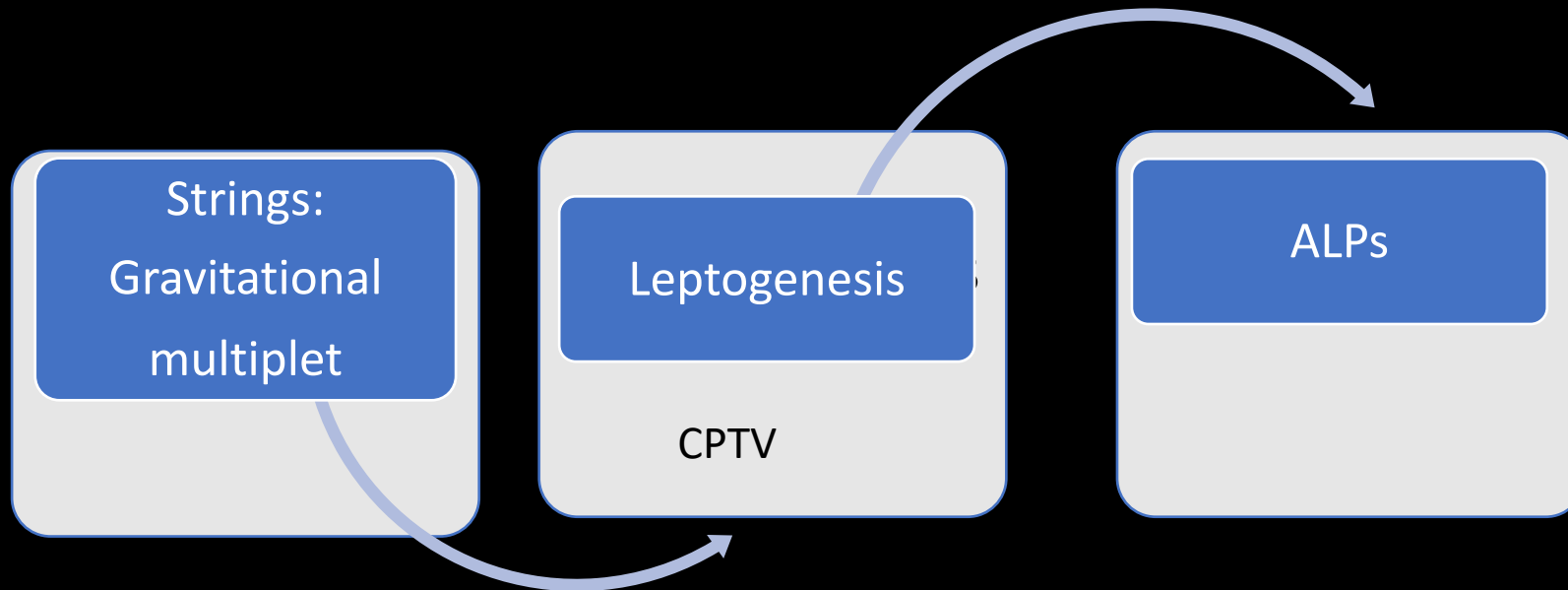
Dynamical torsion

Axial anomaly \longrightarrow

$$-\frac{\alpha_{em}\bar{Q}^2}{\pi f_b} \int \mathbf{b} \mathcal{F} \wedge \mathcal{F} - \frac{1}{2} \int d\mathbf{b} \wedge * d\mathbf{b} - \frac{1}{8\pi^2} \int \left(\Theta + \frac{N_f}{f_b} \mathbf{b} \right) \tilde{R}^{ab} \wedge \tilde{R}_{ab} - \frac{\alpha_s}{2\pi} \int \left(\theta + \frac{N_q}{f_b} \mathbf{b} \right) \text{Tr}[G \wedge G]$$

$$f_b = \kappa^{-1} \sqrt{8/3} = 4 \times 10^{18} \text{ GeV.}$$

CONCLUDING remarks



Viabale scenario

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