

A photograph of Kelvin-Helmholtz clouds, which are wave-like cloud formations that occur in the atmosphere. The clouds are dark and swirling, with a lighter, yellowish-green hue in the center of each wave. The background is a clear, light blue sky.

# OBSERVATION OF A STRONG DIOCOTRON INSTABILITY IN AN ANTI-PROTON PLASMA RING AND “TAILORING” OF PURE ELECTRON PLASMAS IN THE STRONG DRIVE REGIME

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*Ingmari Tietje (TU Berlin)  
On behalf of the AEGIS collaboration (CERN)*



## OUTLINE

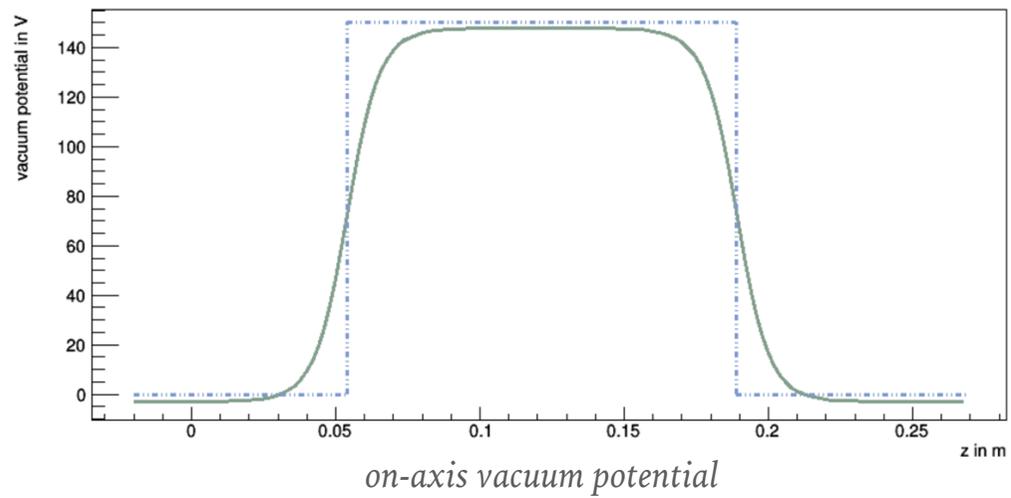
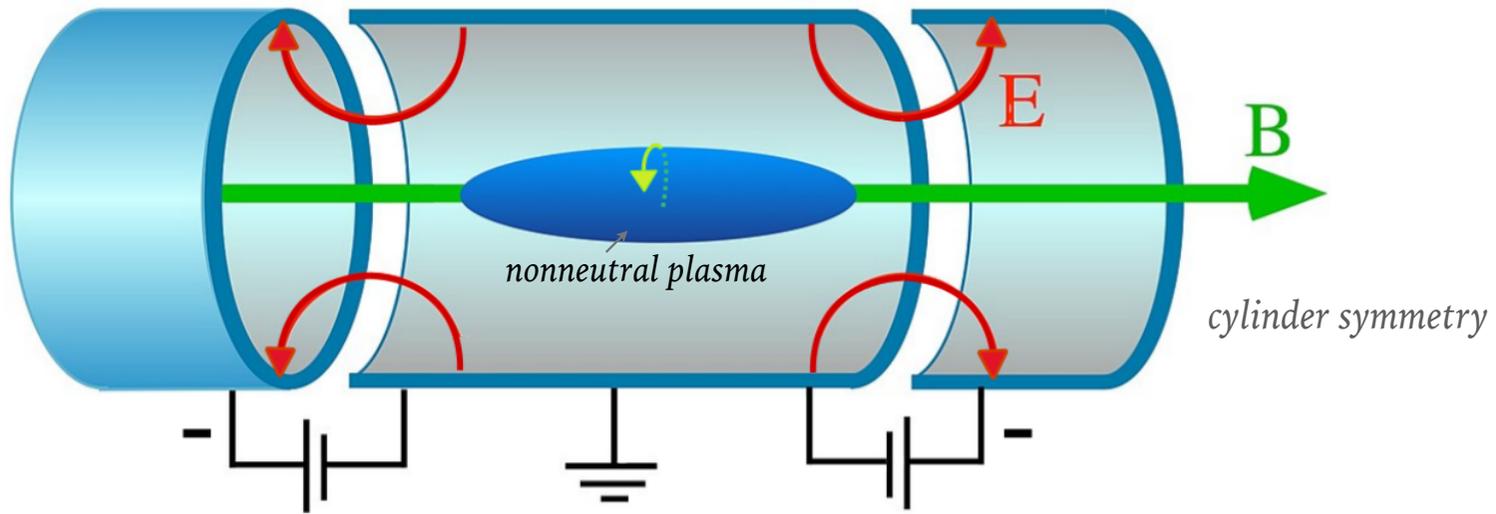
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- Properties of nonneutral plasmas.
- Strong Drive Evaporative Cooling (SDREVC) of electron plasmas (pioneered by the ALPHA collaboration)
  - Strong drive.
  - Strong drive + simultaneous evaporation of electrons.
- Diocotron instability
  - Fluid analogy of nonneutral plasma physics.
  - Hollow antiproton ring diocotron evolution.
- Plasma fluctuations.
- Summary & Outlook.

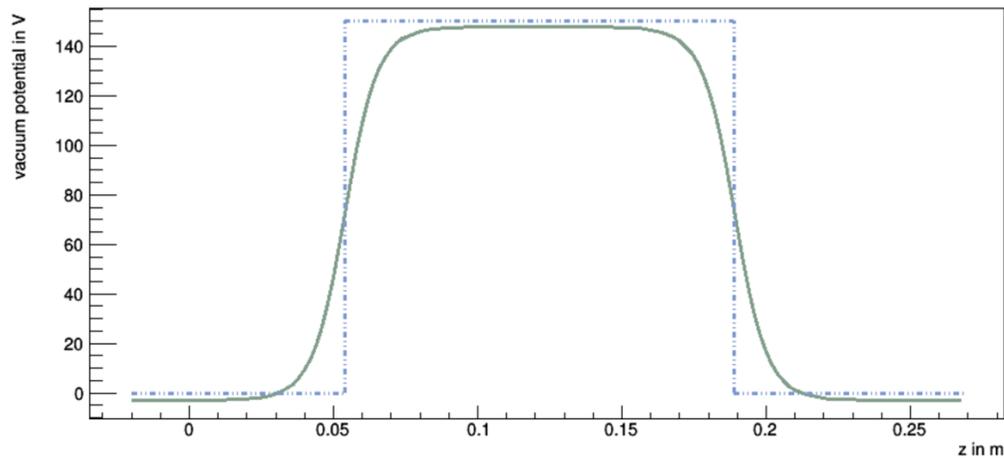
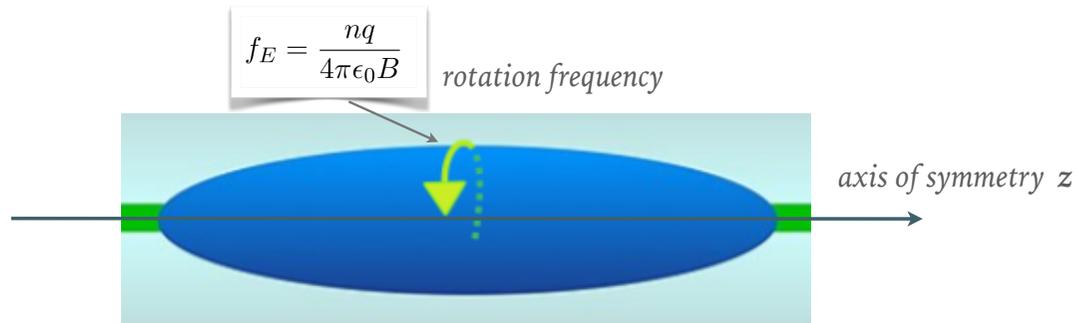
# NONNEUTRAL PLASMAS

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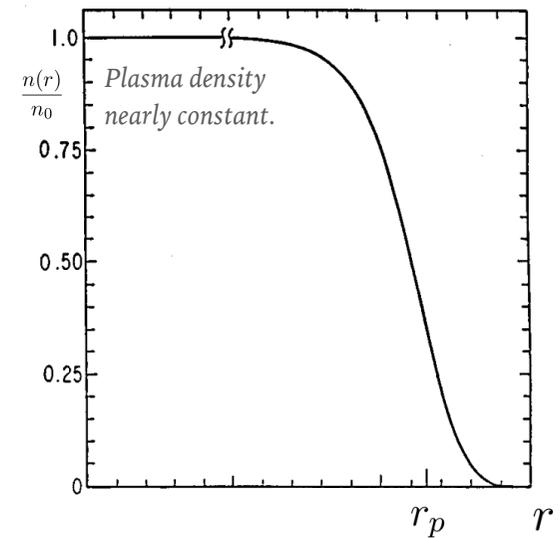
# PENNING-MALMBERG TRAP



# NONNEUTRAL PLASMAS

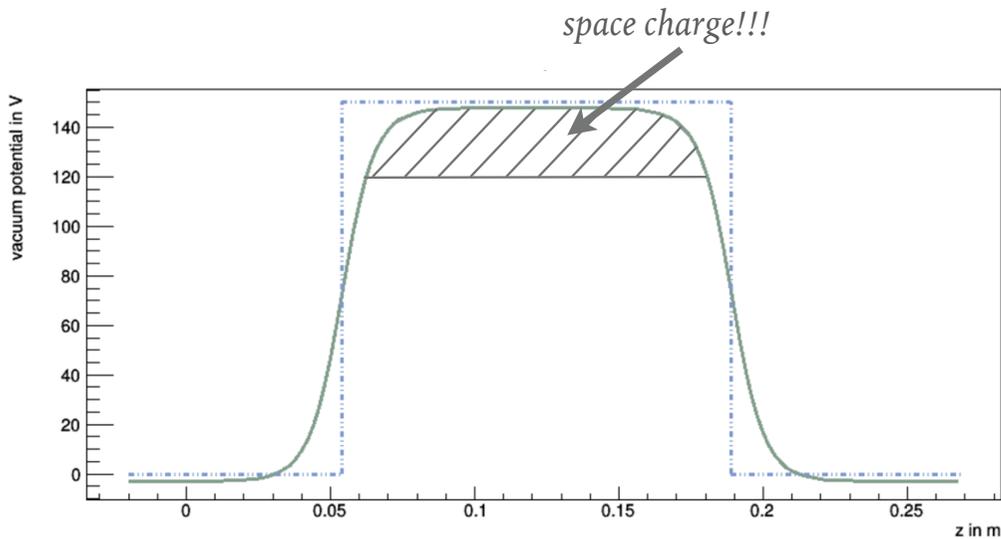
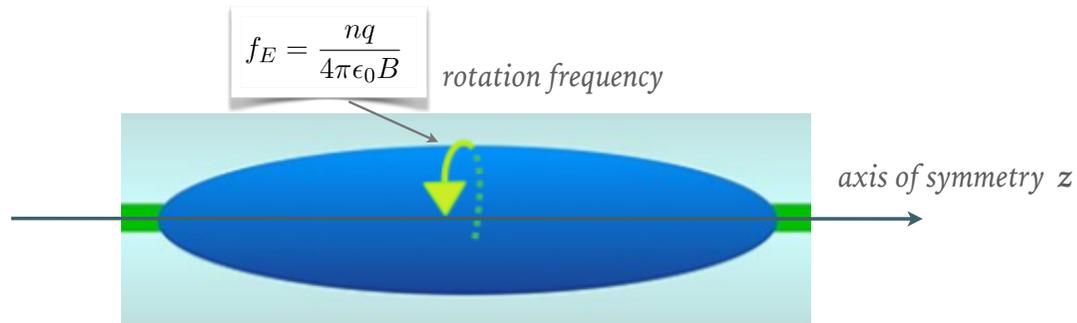


on-axis vacuum potential

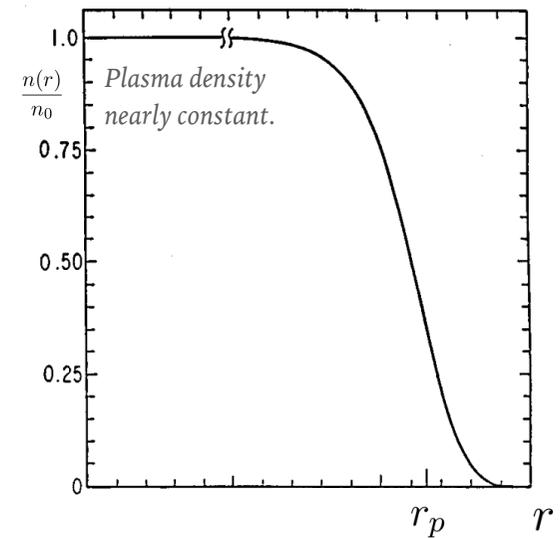


adapted from D. H. E. Dubin and T. M. O'Neil, Rev. Mod. Phys. 71, 87-172 (1999)

# NONNEUTRAL PLASMAS



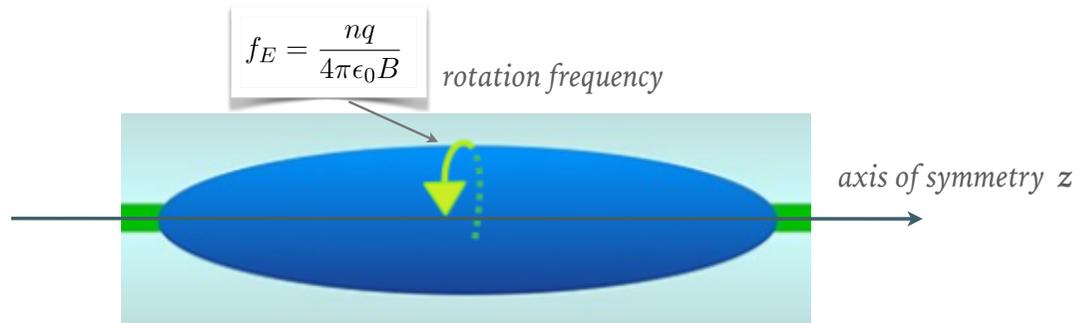
on-axis vacuum potential



$$\lambda_D = \sqrt{\frac{\epsilon_0 k_B T}{nq^2}}$$

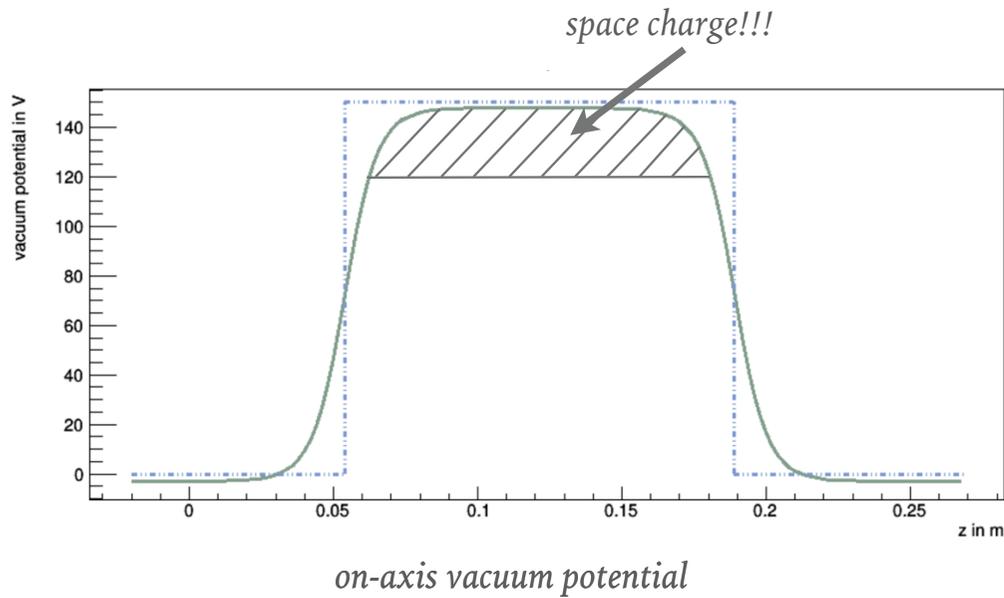
adapted from D. H. E. Dubin and T. M. O'Neil, Rev. Mod. Phys. 71, 87-172 (1999)

# NONNEUTRAL PLASMAS



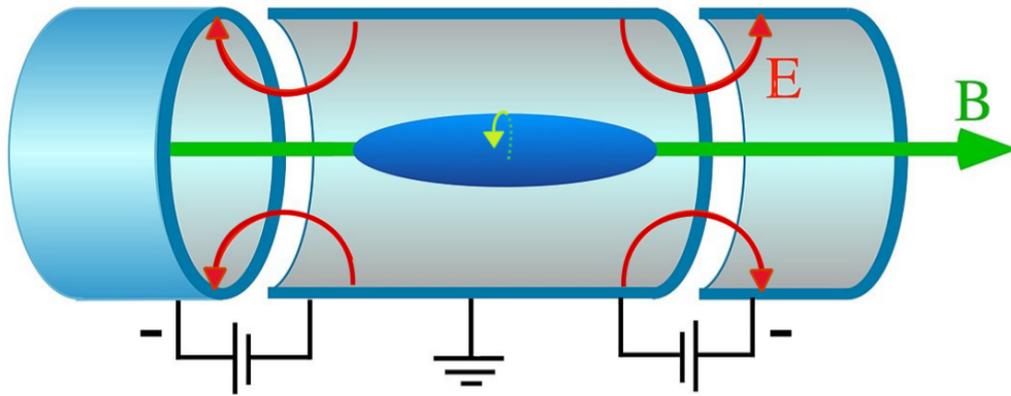
*AEgIS experiment mostly operates in the regime of:*

*Cold-fluid guiding-center approximation!*

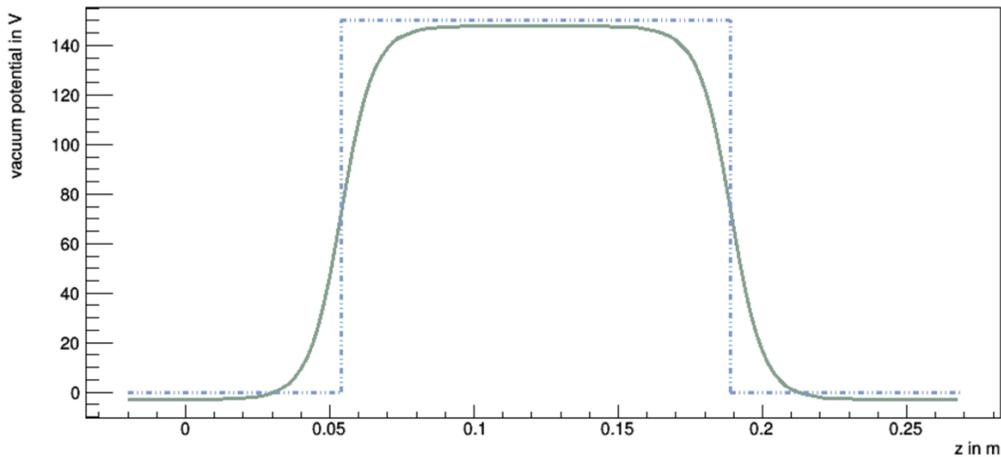


$$\frac{\omega_p^2(r)}{\omega_c^2} = \frac{4\pi n_0(r) m c^2}{B_0} \ll 1$$

# PROPERTIES OF NONNEUTRAL PLASMAS



*Penning-Malmberg trap*



*on-axis vacuum potential*

- ▶ A nonneutral plasma ...
  - ▶ ... significantly alters the vacuum potential of the confining trap.
  - ▶ ... self-shields over the order of a Debye length:  $\lambda_D = \sqrt{\frac{\epsilon_0 k_B T}{nq^2}}$
  - ▶ ... rotates at a uniform rate about the z-axis of the trap, when in thermal equilibrium.
  - ▶ ... assumes a density, which is proportional to the frequency it is rotating at.  $f_E = \frac{nq}{4\pi\epsilon_0 B}$
- ▶ The regime AEGIS is working in can be adequately described within the cold-fluid guiding-center approximation.

$$\frac{\omega_p^2(r)}{\omega_c^2} = \frac{4\pi n_0(r) m c^2}{B_0} \ll 1$$

# **STRONG DRIVE EVAPORATIVE COOLING**

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*of electron plasmas*

# STRONG DRIVE

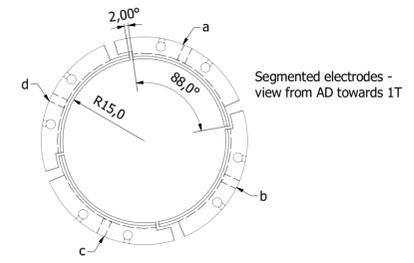
$$f_{rw} \equiv f_E = \frac{nq}{4\pi\epsilon_0 B}$$

- A strongly driven nonneutral plasma 'locks' to its drive frequency. Plasma density proportional to applied drive frequency.

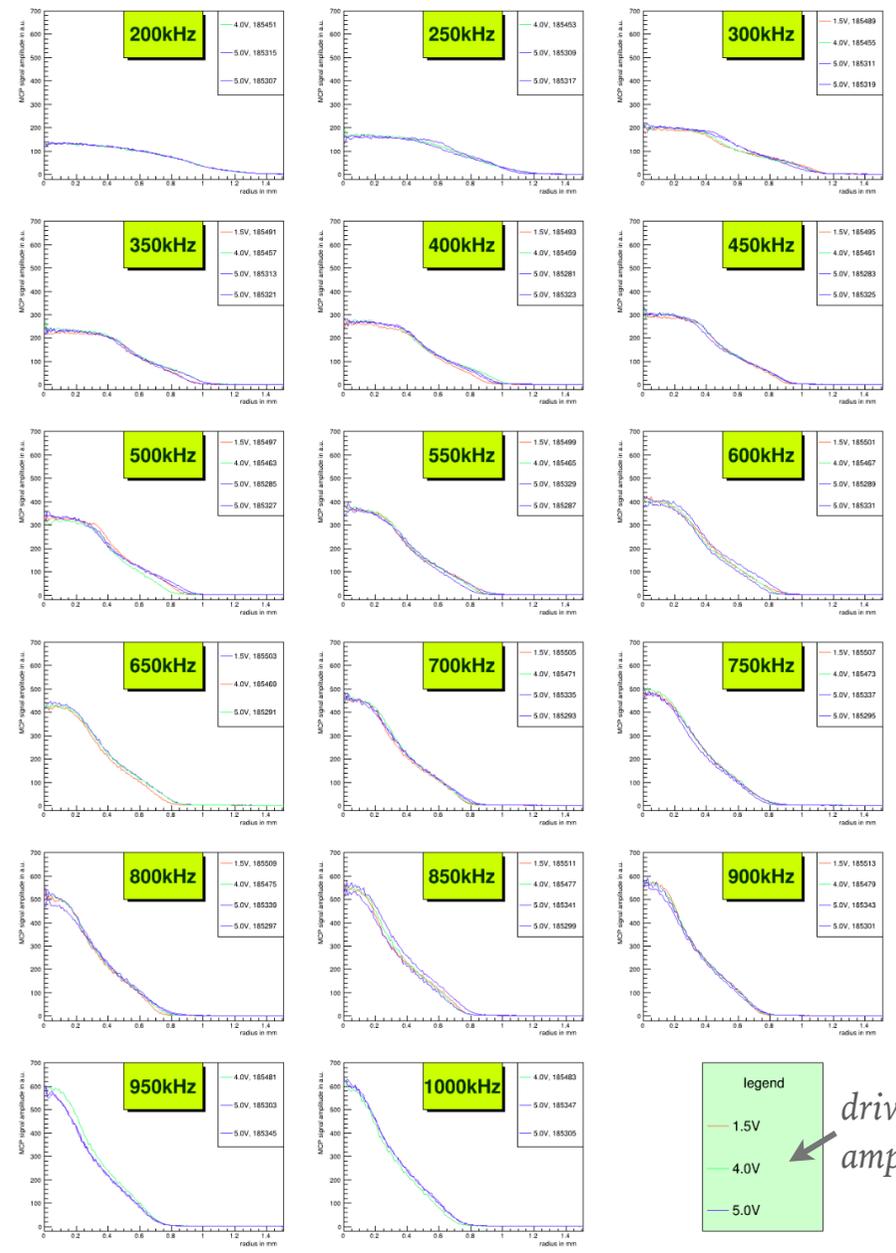
→ Plasma compression.

*J. R. Danielson and C. M. Surko, Phys. Rev. Lett. 94, 035001 (2005).*  
*J.R. Danielson and C. M. Surko, Physics of Plasmas 13, 055706 (2006)*

- The rotating wall technique provides the drive.

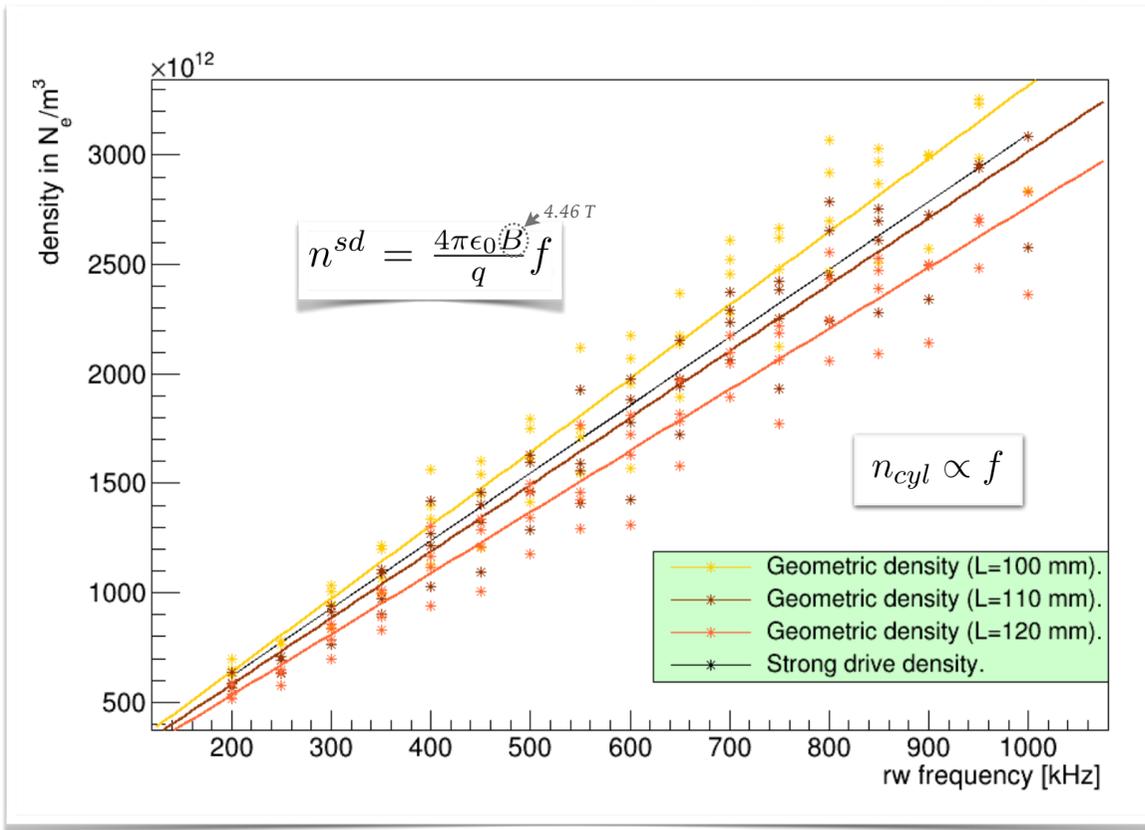


$$\phi_j(\theta, t) = a_{rw} \cos[m_\theta(\theta_j - 2\pi f_{rw}t)], \theta_j - \frac{\Delta\theta}{2} < \theta < \theta_j + \frac{\Delta\theta}{2}$$

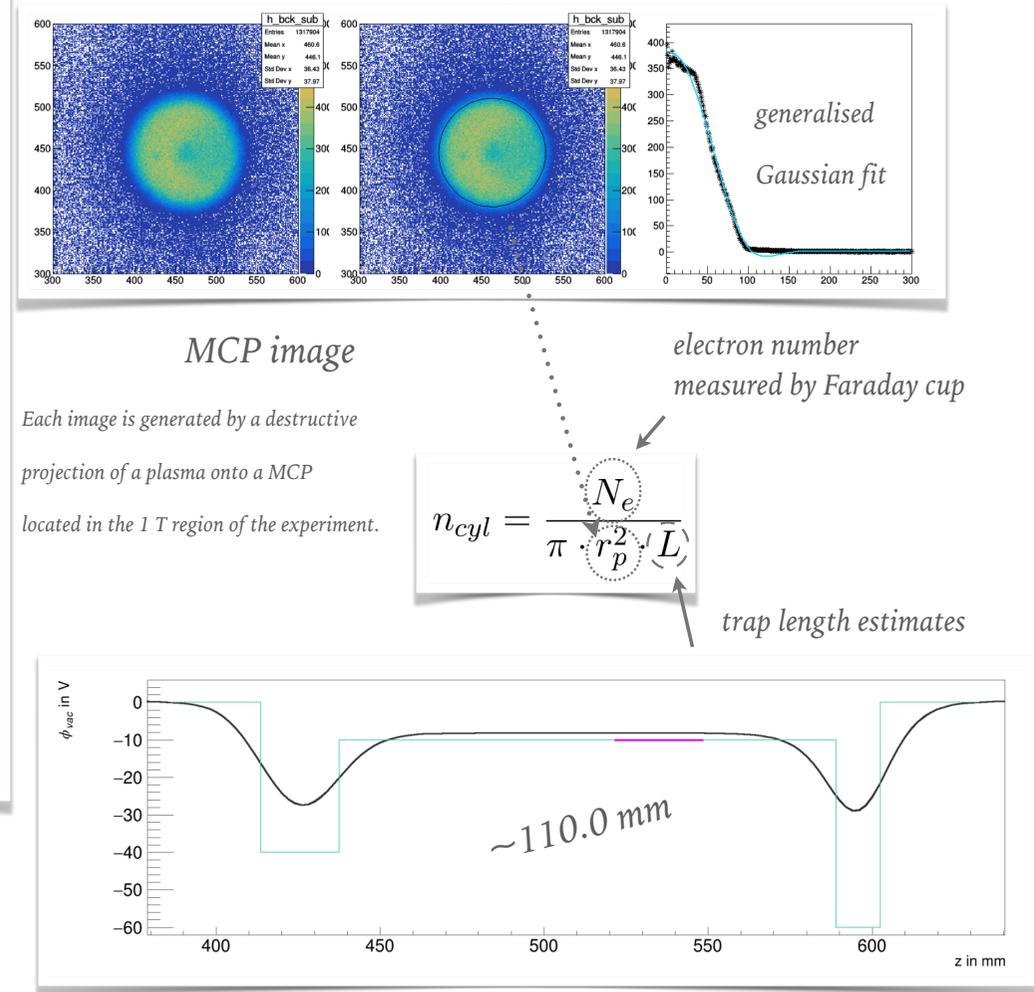


legend  
 — 1.5V  
 — 4.0V  
 — 5.0V  
 ← drive amplitude

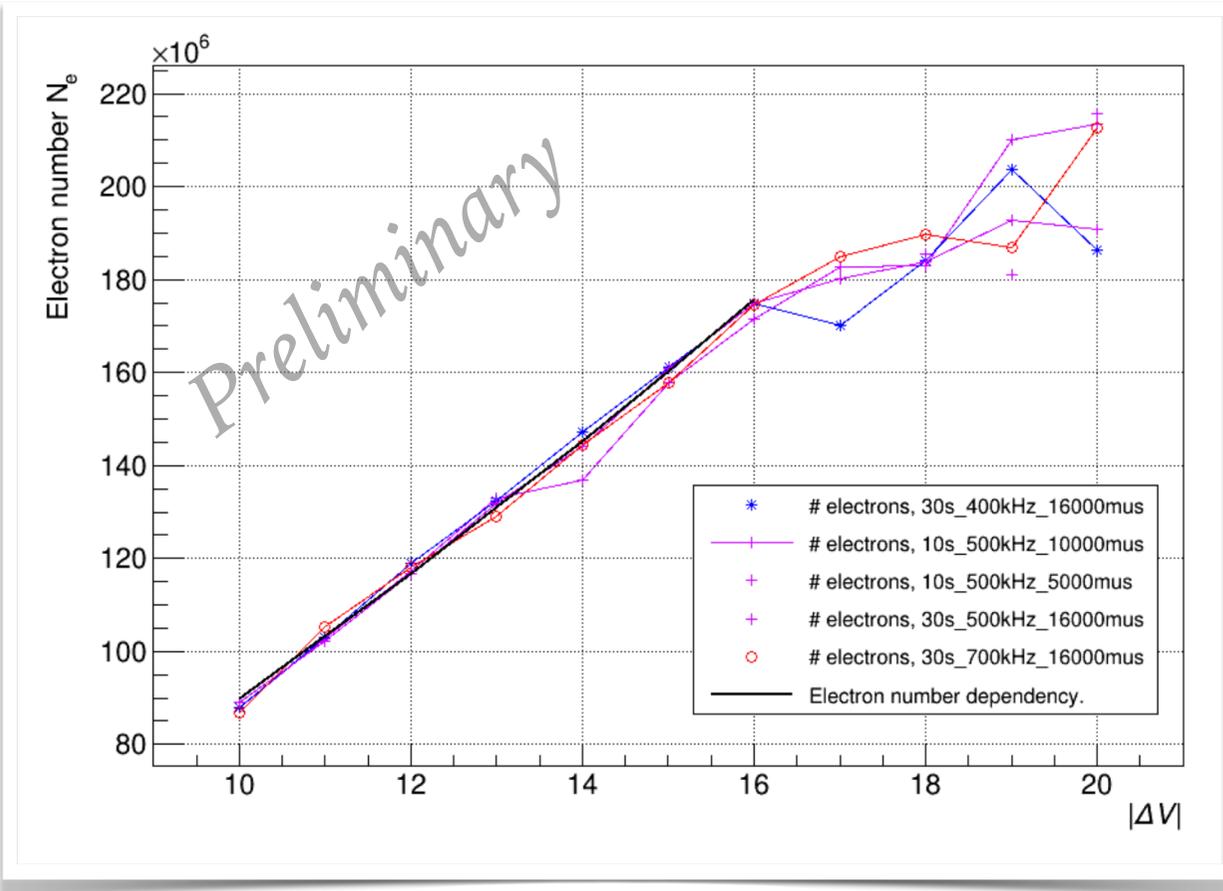
# STRONG DRIVE



The density inferred from measurement clearly indicates, that the plasma is strongly driven.



# STRONG DRIVE EVAPORATIVE COOLING



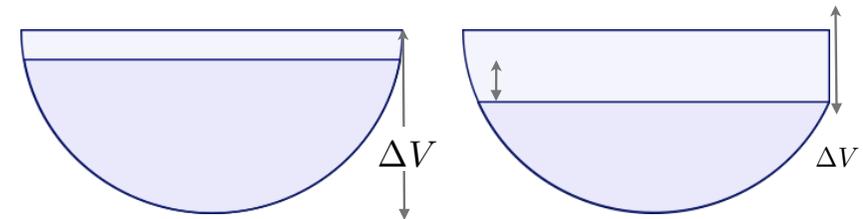
## Enhanced Control and Reproducibility of Non-Neutral Plasmas

M. Ahmadi,<sup>1</sup> B. X. R. Alves,<sup>2</sup> C. J. Baker,<sup>3</sup> W. Bertsche,<sup>4,5</sup> A. Capra,<sup>6</sup> C. Carruth,<sup>7,\*</sup> C. L. Cesar,<sup>8</sup> M. Charlton,<sup>3</sup> S. Cohen,<sup>9</sup> R. Collister,<sup>6</sup> S. Eriksson,<sup>3</sup> A. Evans,<sup>10</sup> N. Evetts,<sup>11</sup> J. Fajans,<sup>7</sup> T. Friesen,<sup>2</sup> M. C. Fujiwara,<sup>6</sup> D. R. Gill,<sup>6</sup> J. S. Hangst,<sup>2</sup> W. N. Hardy,<sup>11</sup> M. E. Hayden,<sup>12</sup> C. A. Isaac,<sup>3</sup> M. A. Johnson,<sup>4</sup> S. A. Jones,<sup>2,3</sup> S. Jonsell,<sup>13</sup> L. Kurchaninov,<sup>6</sup> N. Madsen,<sup>3</sup> M. Mathers,<sup>14</sup> D. Maxwell,<sup>3</sup> J. T. K. McKenna,<sup>6</sup> S. Menary,<sup>14</sup> T. Momose,<sup>15</sup> J. J. Munich,<sup>13</sup> K. Olchanski,<sup>6</sup> A. Olin,<sup>6,16</sup> P. Pusa,<sup>1</sup> C. Ø. Rasmussen,<sup>2</sup> F. Robicheaux,<sup>17</sup> R. L. Sacramento,<sup>8</sup> M. Sameed,<sup>3,4</sup> E. Sarid,<sup>18</sup> D. M. Silveira,<sup>8</sup> C. So,<sup>6,10</sup> G. Stutter,<sup>2</sup> T. D. Tharp,<sup>19</sup> J. E. Thompson,<sup>14</sup> R. I. Thompson,<sup>6,10</sup> D. P. van der Werf,<sup>3,20</sup> and J. S. Wurtele<sup>7</sup>

(ALPHA Collaboration)

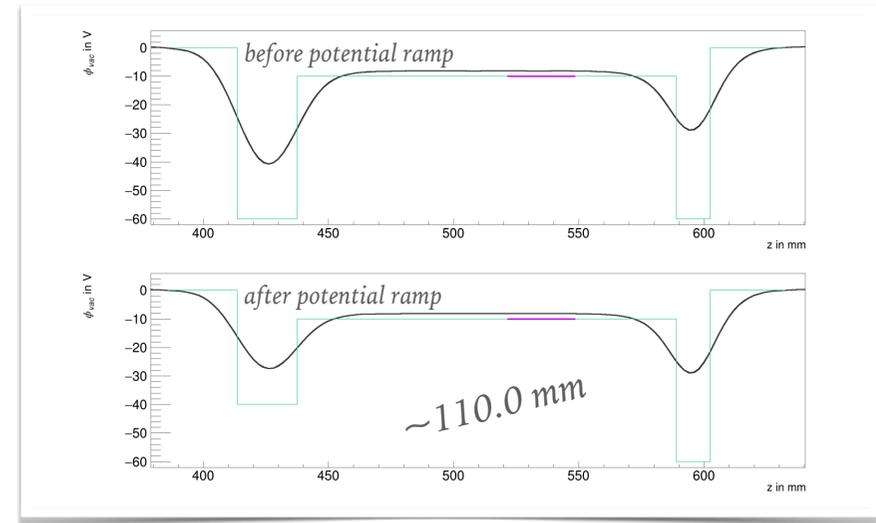
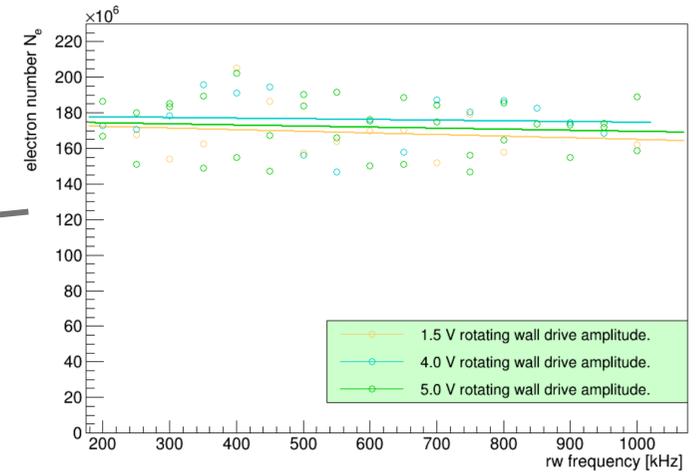
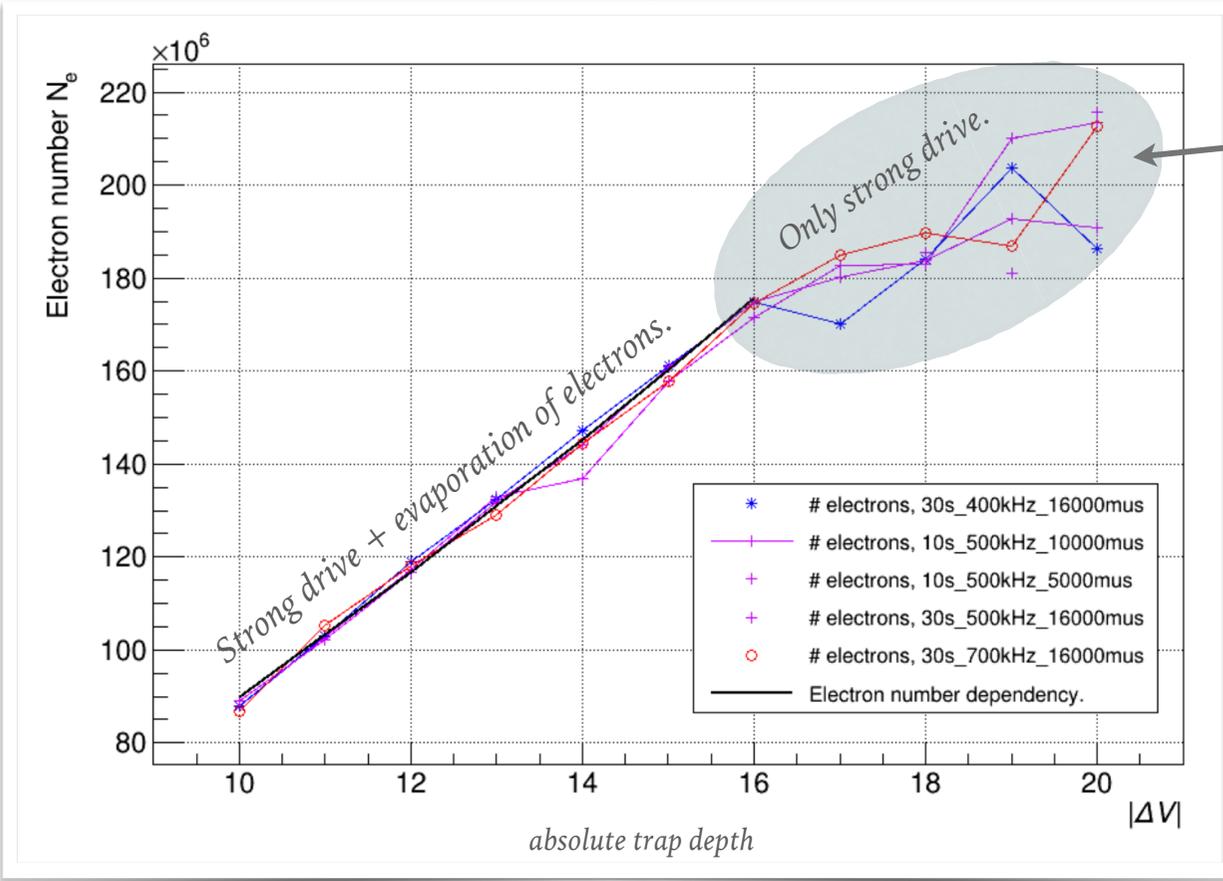
C. Carruth, "Methods for plasma stabilization and control to improve antihydrogen production.",

PhD thesis (UC Berkeley, 2018)

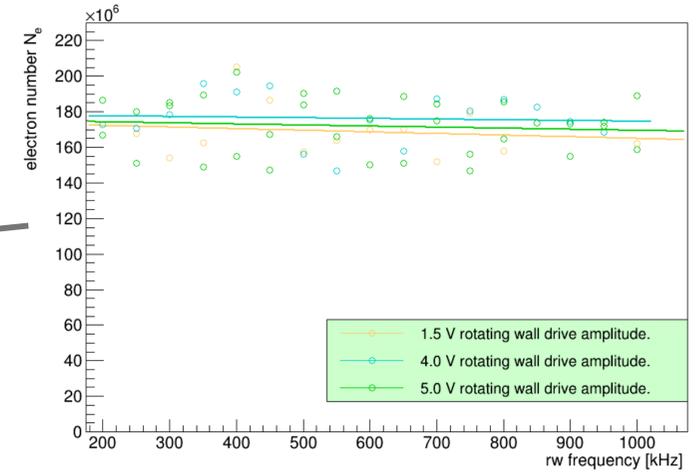
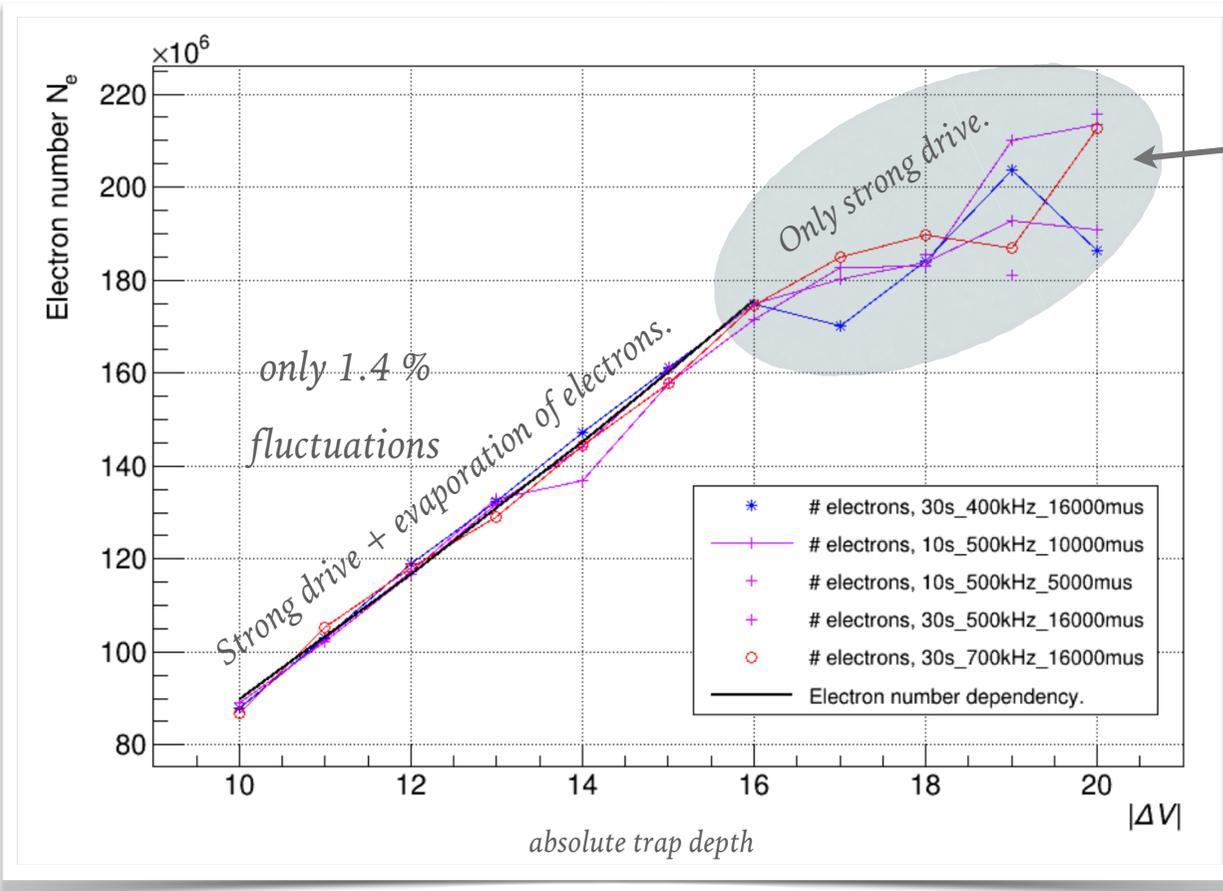


$$\frac{mv_{\parallel}^2}{2} - e\phi > -e\Delta V$$

# STRONG DRIVE EVAPORATIVE COOLING



# STRONG DRIVE EVAPORATIVE COOLING

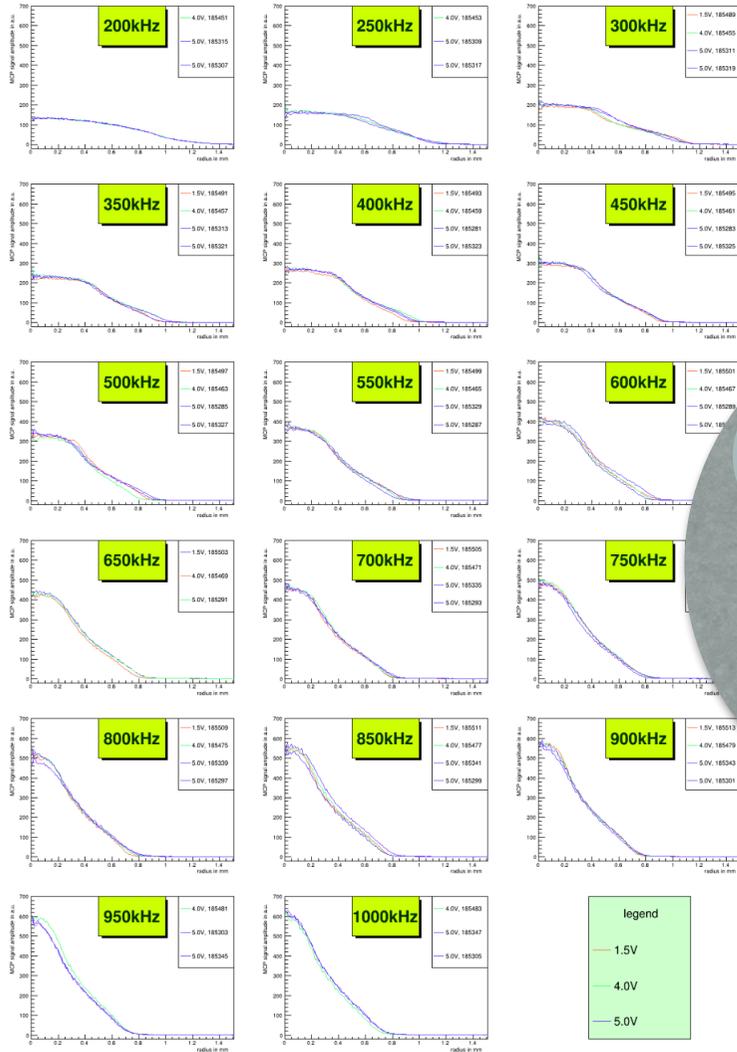


$$N = \frac{N_0}{\phi_{sc}^0} \cdot \frac{L_0}{L} \cdot \phi_{sc}$$

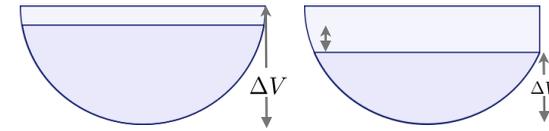
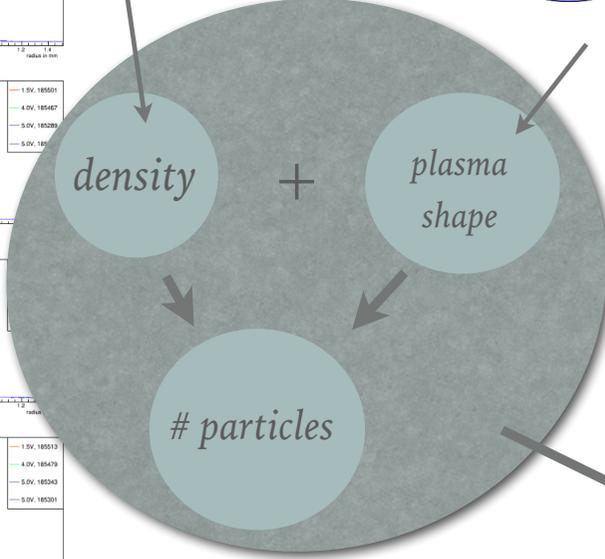
$$N_0 = \pi r_{p0}^2 \cdot L_0 \cdot \frac{4\pi\epsilon_0 B}{q} \cdot \cancel{f} \cdot \cancel{f_0}$$

$$= \pi r_{p0}^2 \cdot L_0 \cdot \frac{4\pi\epsilon_0 B}{q} \cdot f_0,$$

# STRONG DRIVE EVAPORATIVE COOLING



$$f_{rw} \equiv f_E = \frac{nq}{4\pi\epsilon_0 B}$$

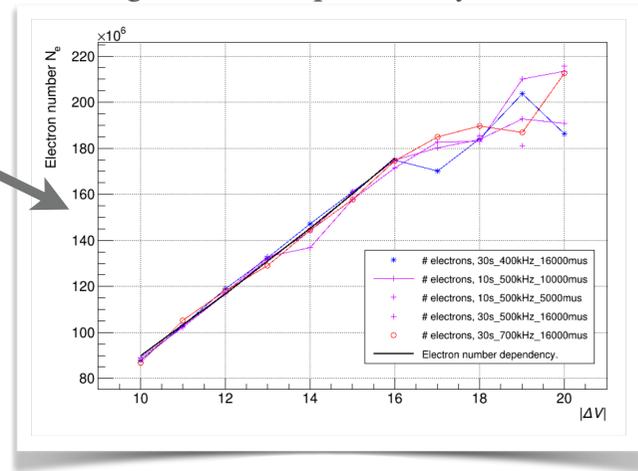


$$\frac{mv_{\parallel}^2}{2} - e\phi > -e\Delta V \quad (\text{Spilling condition})$$

Plasma shape unique in given trap geometry.

S. A. Prasad and T. M. O'Neil, Phys. Fluids 22, 278–281 (1979).

Strong drive + evaporation of electrons.



only 1.4 %  
fluctuations in  
particle number

# DIOCOTRON INSTABILITY

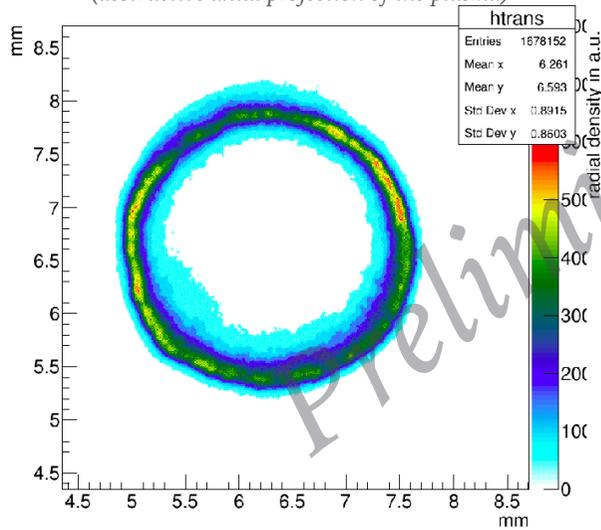
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*(Kelvin-Helmholtz like instability)*

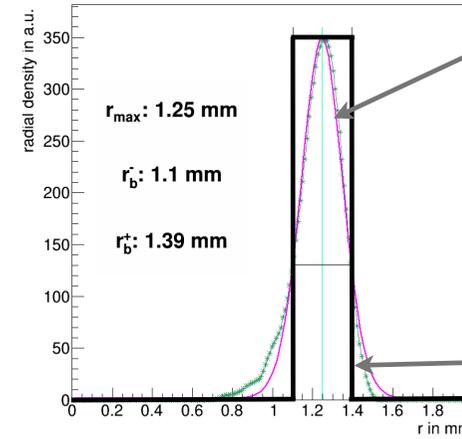
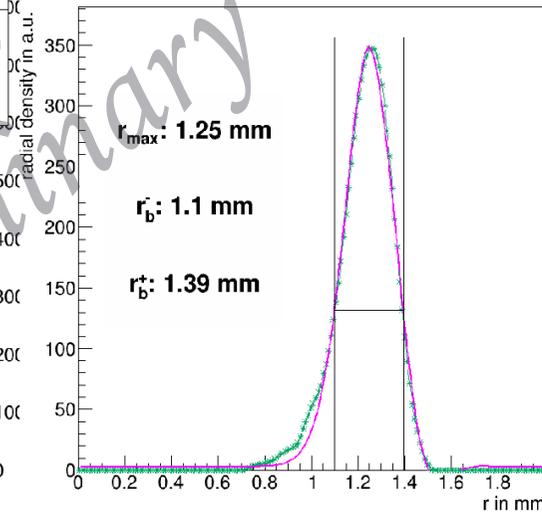
# INITIAL DENSITY PROFILE

processed MCP image

(destructive axial projection of the plasma)



radial projection of plasma



Generalised Gaussian fit.

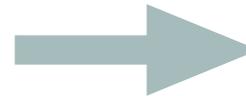
Step-profile approximation.

$$n^0(r) = \begin{cases} 0 & a \leq r < r_b^- \\ \hat{n}_0 = \text{const.} & r_b^- < r < r_b^+ \\ 0 & r_b^+ < r < r_w \end{cases}$$

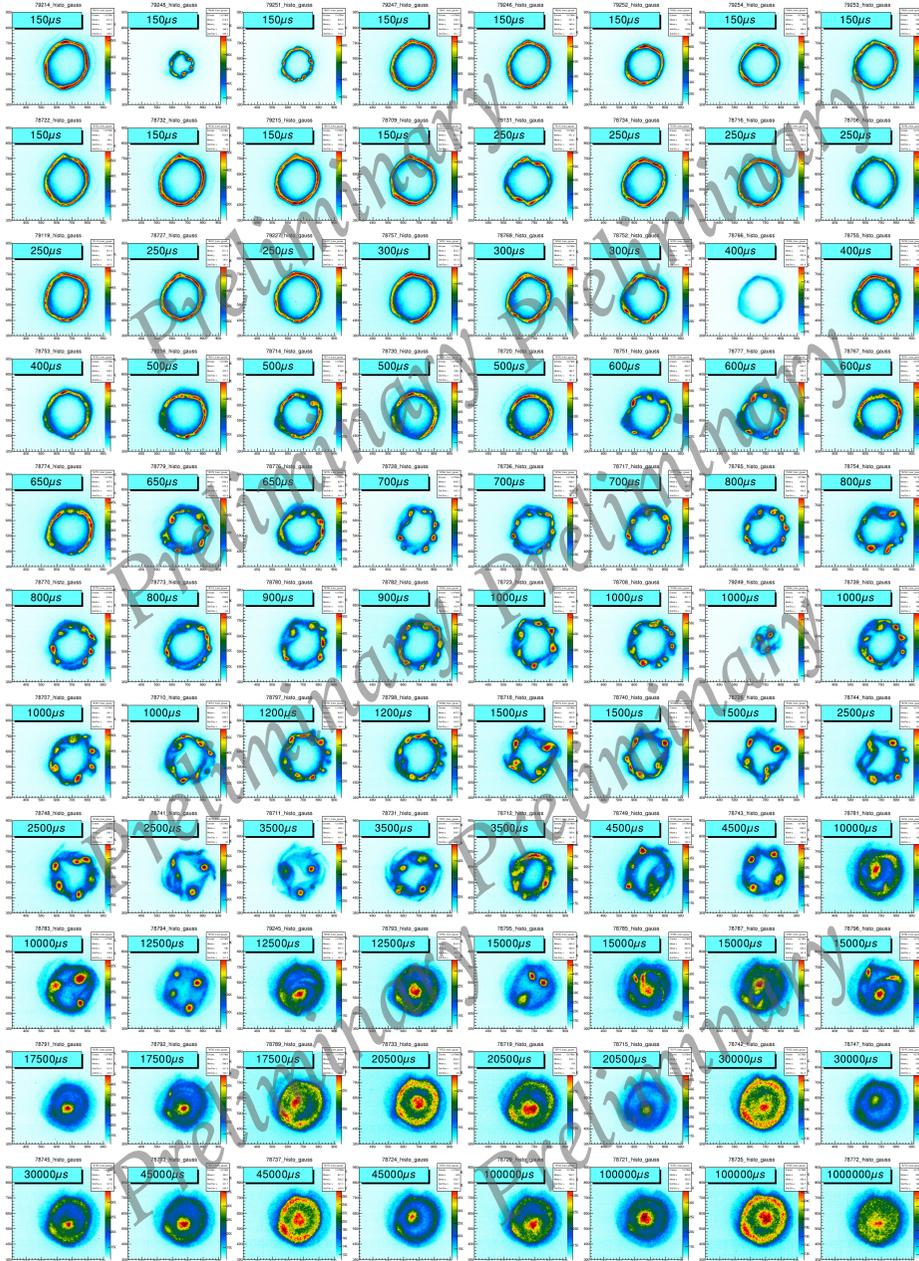


Antiproton ring rotates  
at all times!

$$f_E = \frac{nq}{4\pi\epsilon_0 B}$$



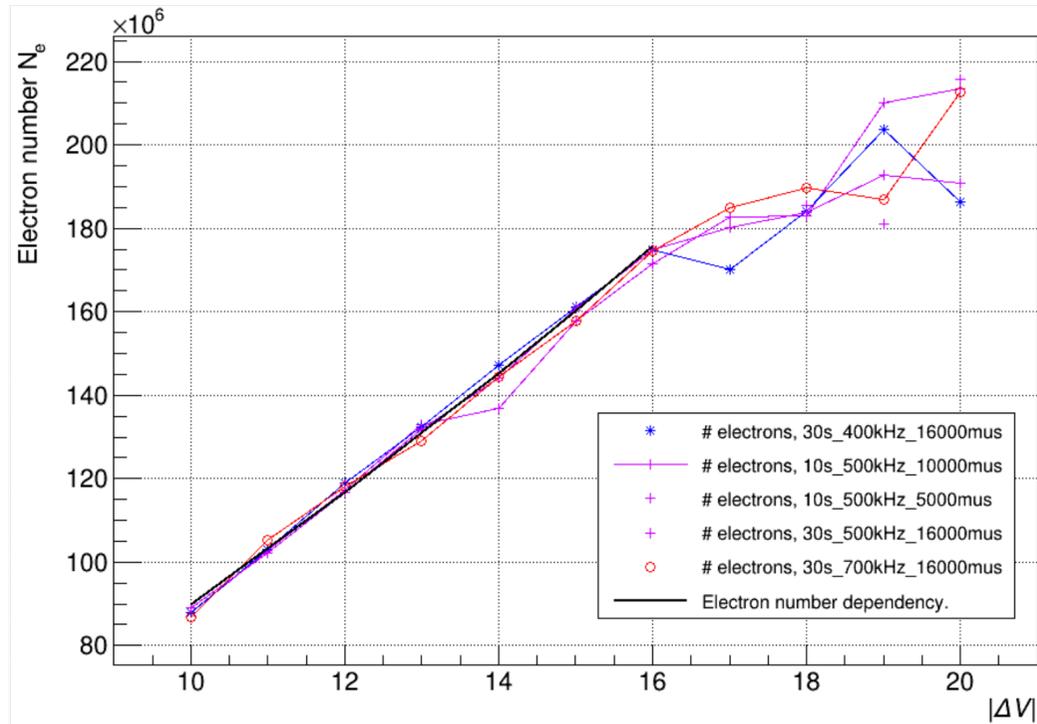
Shear in velocities!



## DIOCOTRON INSTABILITY IN AN ANTIPROTON RING

- Stages of the evolution:
  - Linear Stage.
    - Growth of instabilities.
      - Fastest growing mode determines the number of vortices.
  - Vortex formation = end of the linear stage.
  - Nonlinear Stage.
    - Radial transport of plasma bulk material through nonlinear dynamics.
      - Vortex interaction.
      - Filamentation.
      - Vortex mergers.
  - Final stage.
    - Stable density distribution.

# FLUCTUATIONS



- Fluctuations in antiproton ring size.
  - Due to fluctuations in electron number and density as well as antiproton number fluctuations in the antiproton decelerator cycle.
  - The electron plasma fluctuations affect the antiproton ring size, since the antiproton rings are created through centrifugal separation and ejection of the electron core.
- Fluctuations can be minimized by the SDREVC technique (Strong Drive EVaporative Cooling) in the future.
  - The shallow potential wells of Penning-Malmberg traps permit the precise control of electron plasma particle number by strongly driving the plasma whilst evaporating particles through ramping the potential well.



## SUMMARY & OUTLOOK

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- Summary:
  - Successful implementation of the SDREVC technique into the AEgIS apparatus.
  - Experimental observation of the temporal evolution of a step-profile-like annular antiproton layer.
- Outlook:
  - Incorporation of the SDREVC technique into the regular AEgIS protocols.
    - Stabilisation of charged particle transfers from the trapping to the antihydrogen production region:
      - e.g. the ballistic transfer of antiprotons from the 5T trap region to the antihydrogen production trap.
    - Stabilisation of antiproton ring radii for diocotron instability investigations.

**THANK YOU VERY MUCH FOR YOUR  
ATTENTION!**

