



Dark Sectors & Portal Interactions Overview

Gordan Krnjaic

CERN Forward Spectrometer Meeting, April 16, 2020

Open Questions in Fundamental Physics & Cosmology

Neutrino Masses
Matter Asymmetry
Inflation



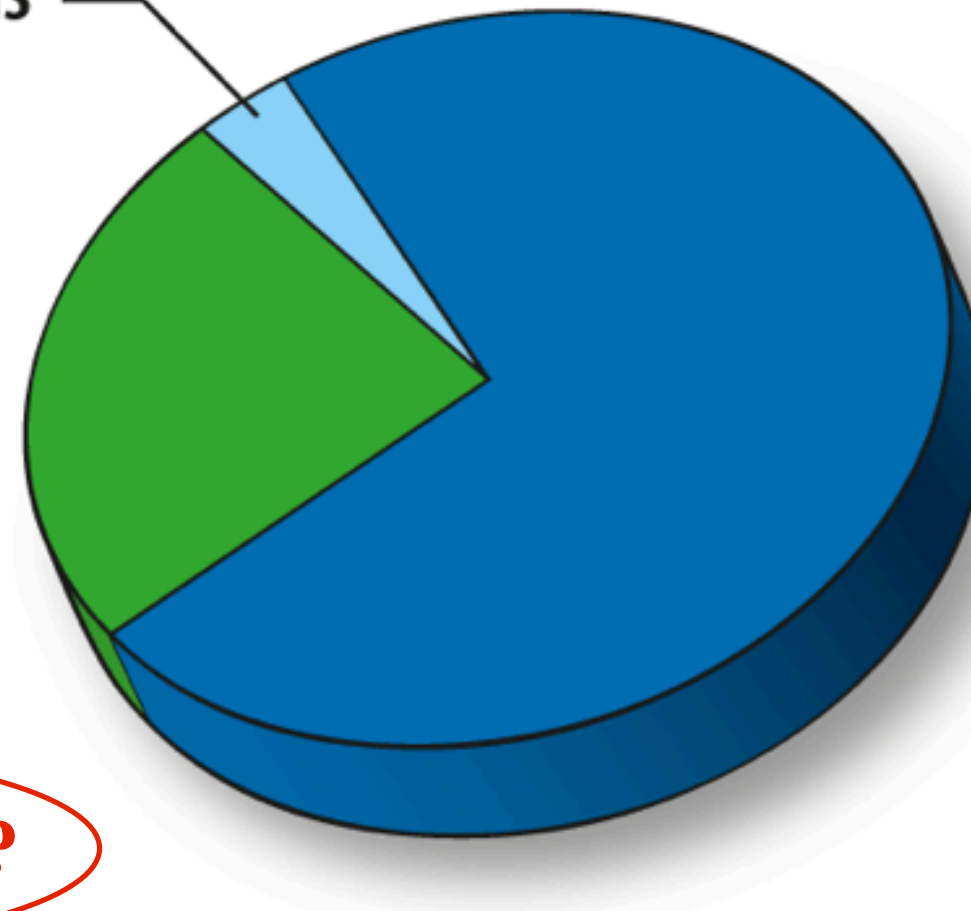
**Accelerated
Cosmic
Expansion**



Atoms
4.6%

Dark
Matter
24%

Dark
Energy
71.4%

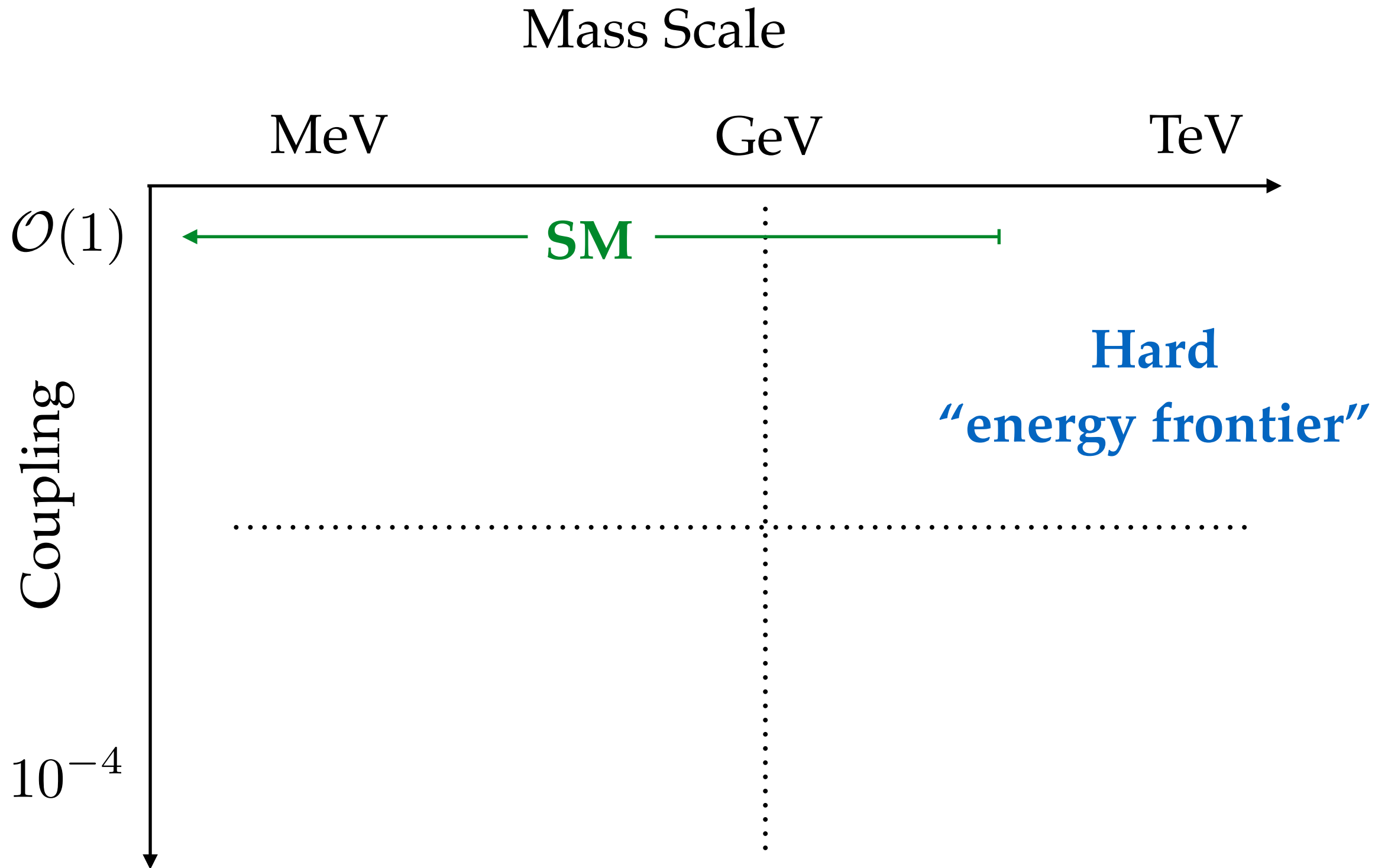


TODAY

What is this stuff?

Also Quantum Gravity

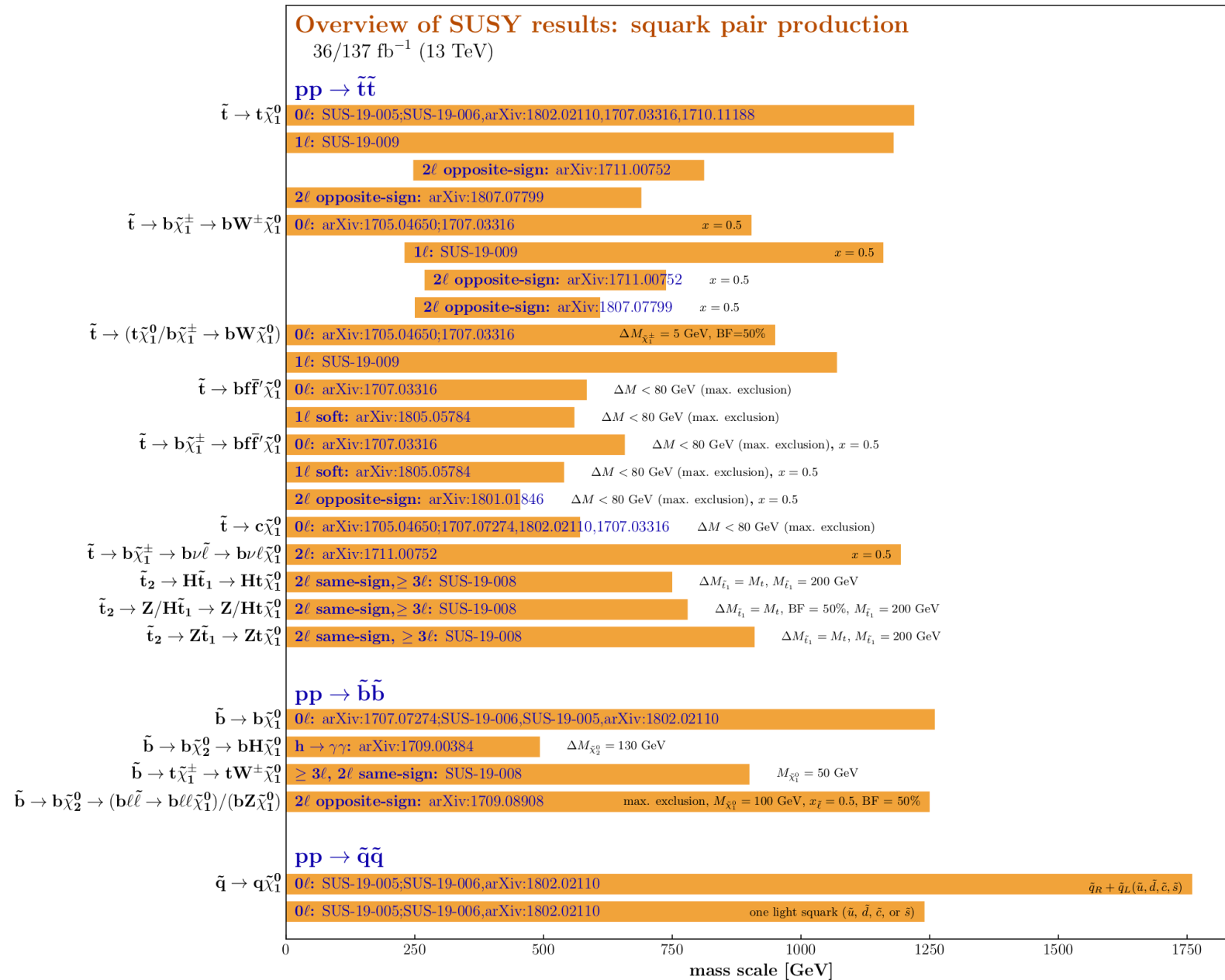
How to look for new physics?



What have we learned on the “energy frontier”?

CMS (preliminary)

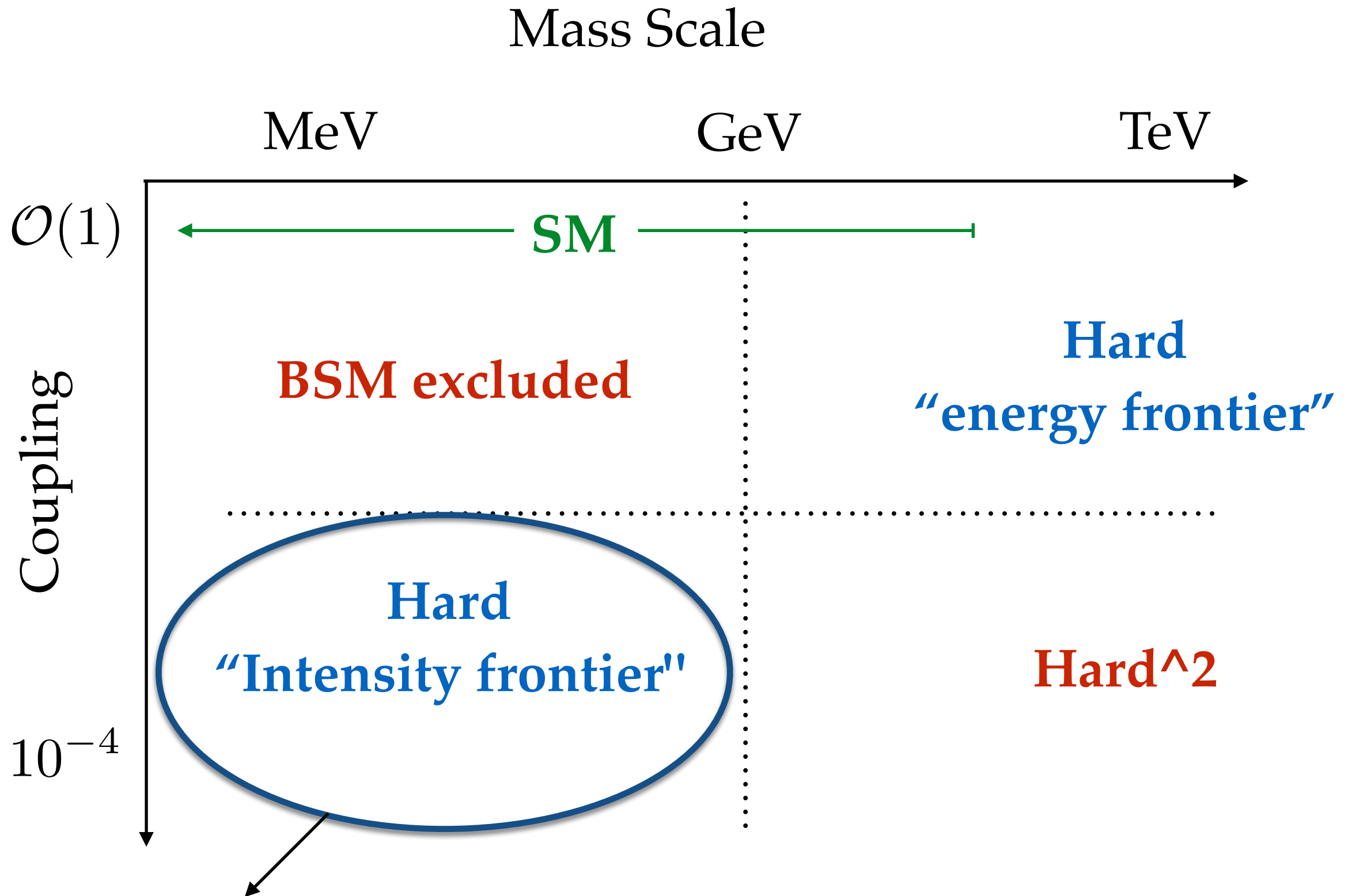
May 2019



Selection of observed limits at 95% C.L. (theory uncertainties are not included). Probe **up to** the quoted mass limit for light LSPs unless stated otherwise. The quantities ΔM and x represent the absolute mass difference between the primary sparticle and the LSP, and the difference between the intermediate sparticle and the LSP relative to ΔM , respectively, unless indicated otherwise.

Null LHC results: no evidence yet of new SM charged particles

How to look for new physics?



BSM: Smaller coupling, lower mass, **SM neutral**

Overview

Part 1) Minimal Single Particle SM Extensions

Part 2) Add Light \sim GeV Dark Matter

Overview

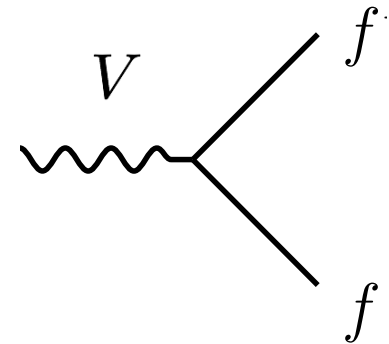
Part 1) Minimal Single Particle SM Extensions

Part 2) Add Light \sim GeV Dark Matter

How to couple single neutral particle to the SM?

Option 1: New gauge force directly coupled to SM current

$$\mathcal{L} \supset g V_\mu J_{\text{SM}}^\mu, \quad J_{\text{SM}}^\mu \equiv \sum_f Q_f \bar{f} \gamma^\mu f$$



Only anomaly free possibilities:

$$U(1)_{B-L}, \quad U(1)_{L_i-L_j}, \quad U(1)_{B-3L_i}$$

Qualitatively similar, but some differences in bounds

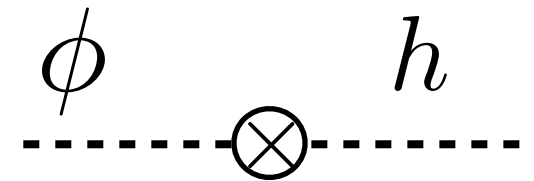
Two parameter family of models: $\{g, m_V\}$

How to couple single neutral particle to the SM?

Option 2: Mass or kinetic mixing with neutral SM particles

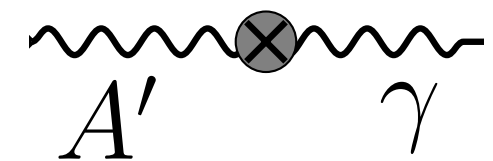
Scalar / Higgs mixing

$$\phi H^\dagger H \rightarrow \phi h$$



Dark / visible photon mixing

$$F'_{\mu\nu} F^{\mu\nu}$$

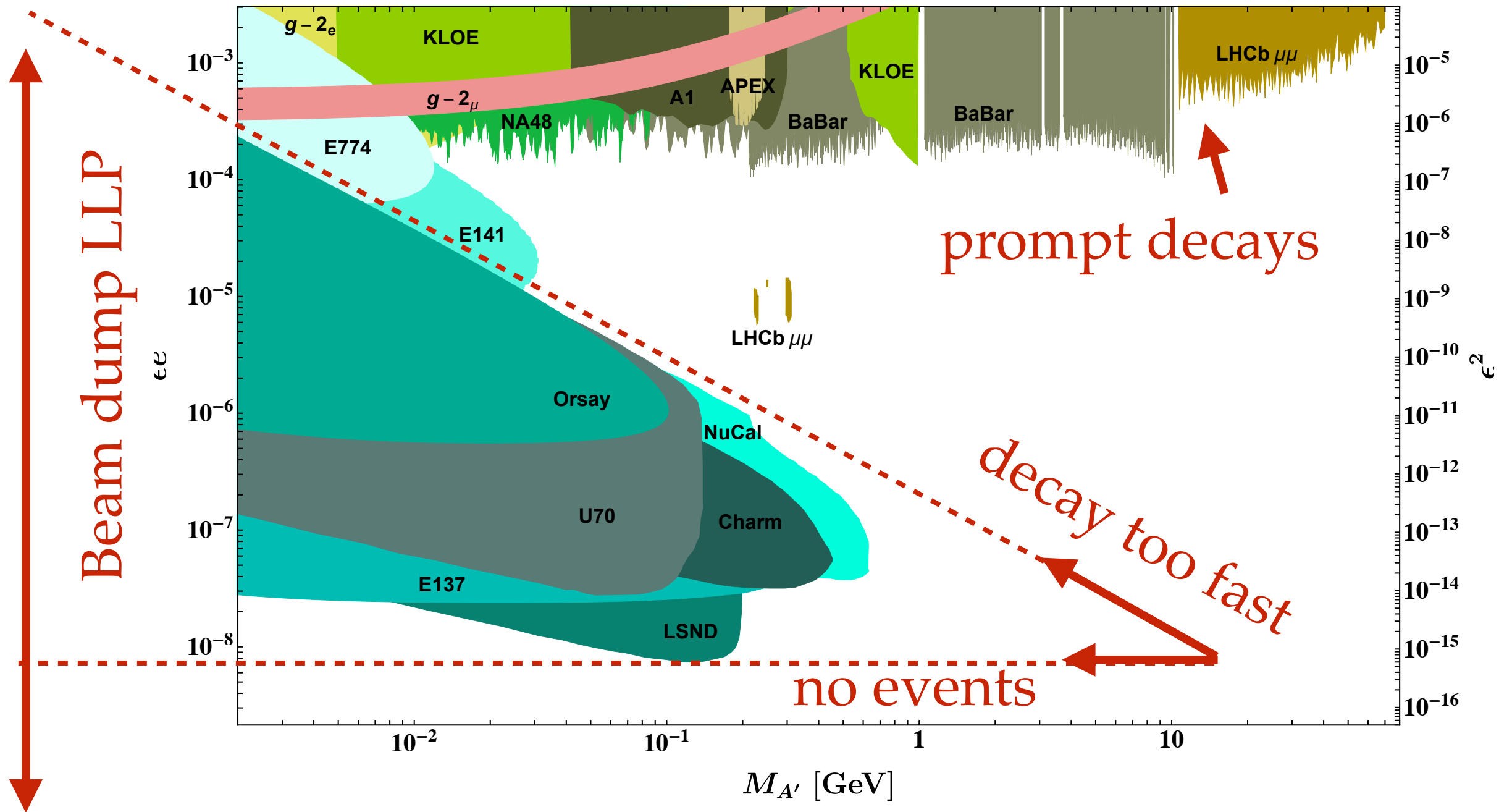


Sterile / active neutrino mixing

$$L H N$$

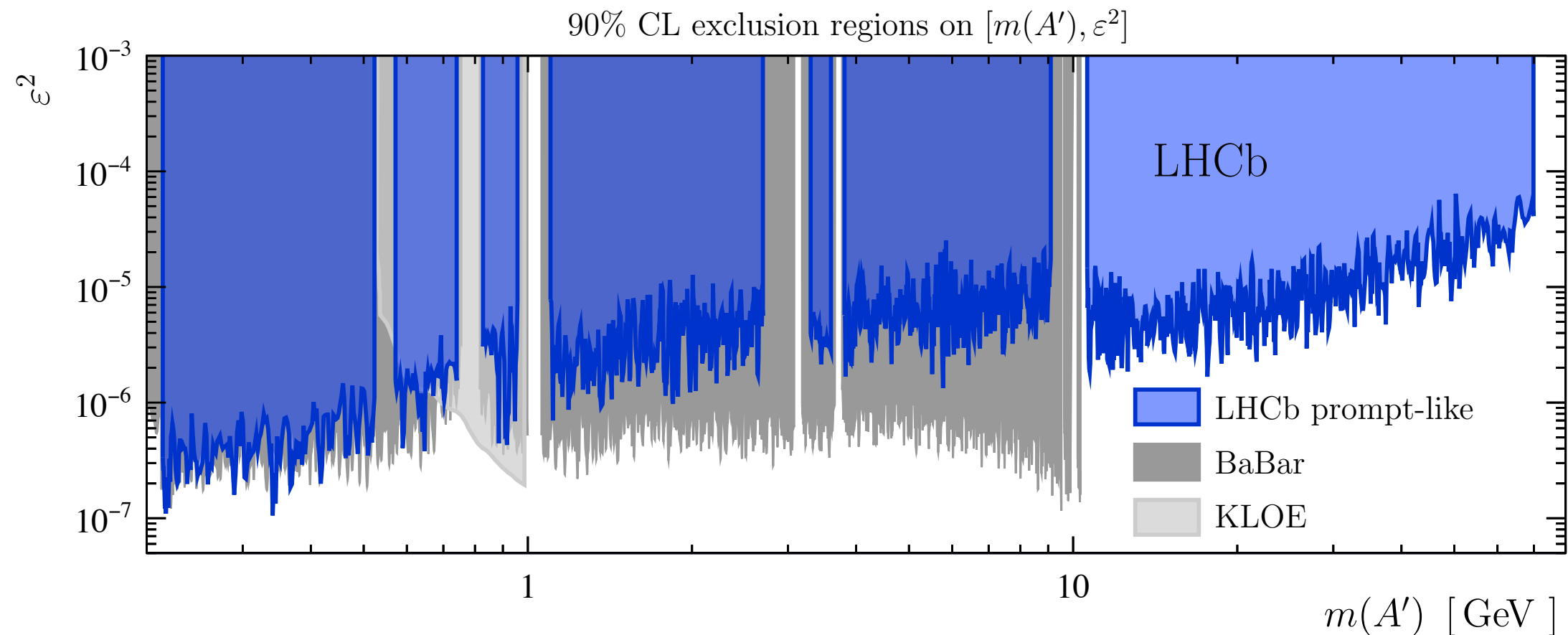


Minimal Kinetically Mixed Dark Photon $\epsilon F'_{\mu\nu} F^{\mu\nu}$



Collider strategy: prompt decays

Resonance searches for visible daughters: BABAR, Belle II, LHCb...



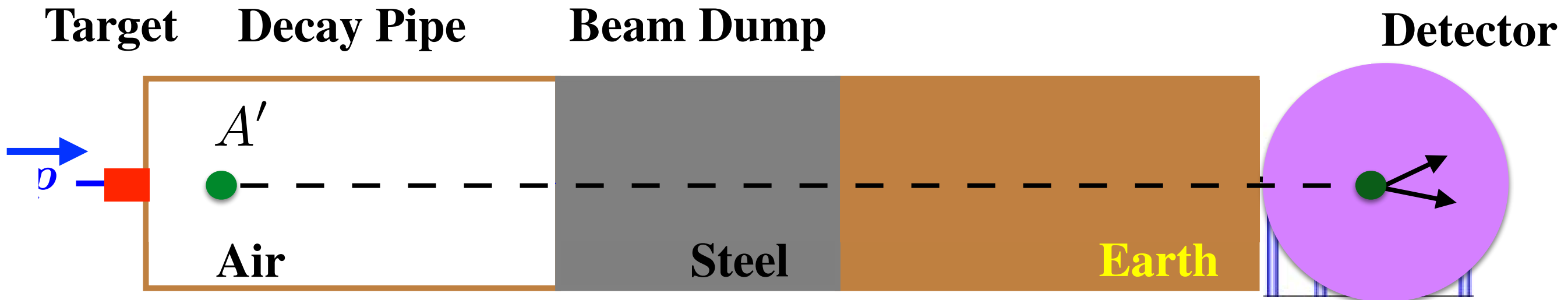
B-factories: continuum production

$$e^+e^- \rightarrow \gamma A' \rightarrow \gamma(e^+e^-)$$

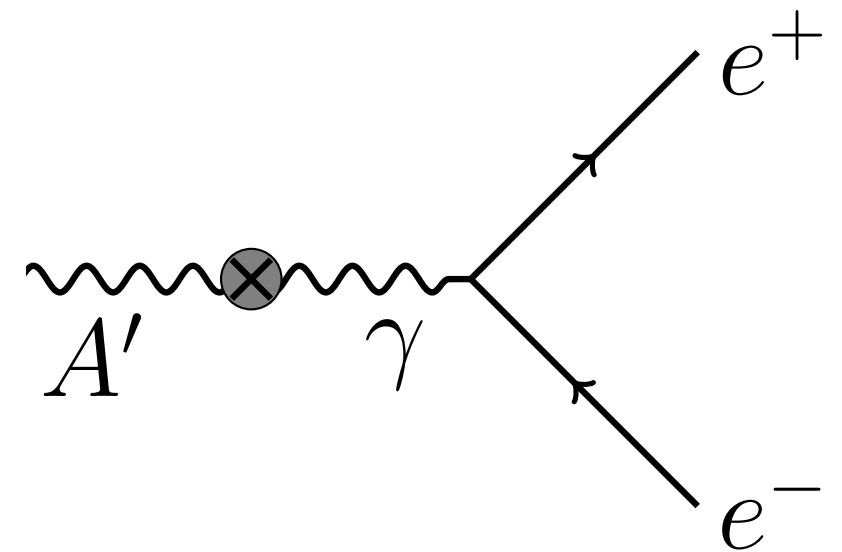
Colliders (also short-er baseline fixed targets)

$$K^+ \rightarrow \pi^+ A' \rightarrow \pi^+(e^+e^-)$$

Beam Dumps: LLP searches

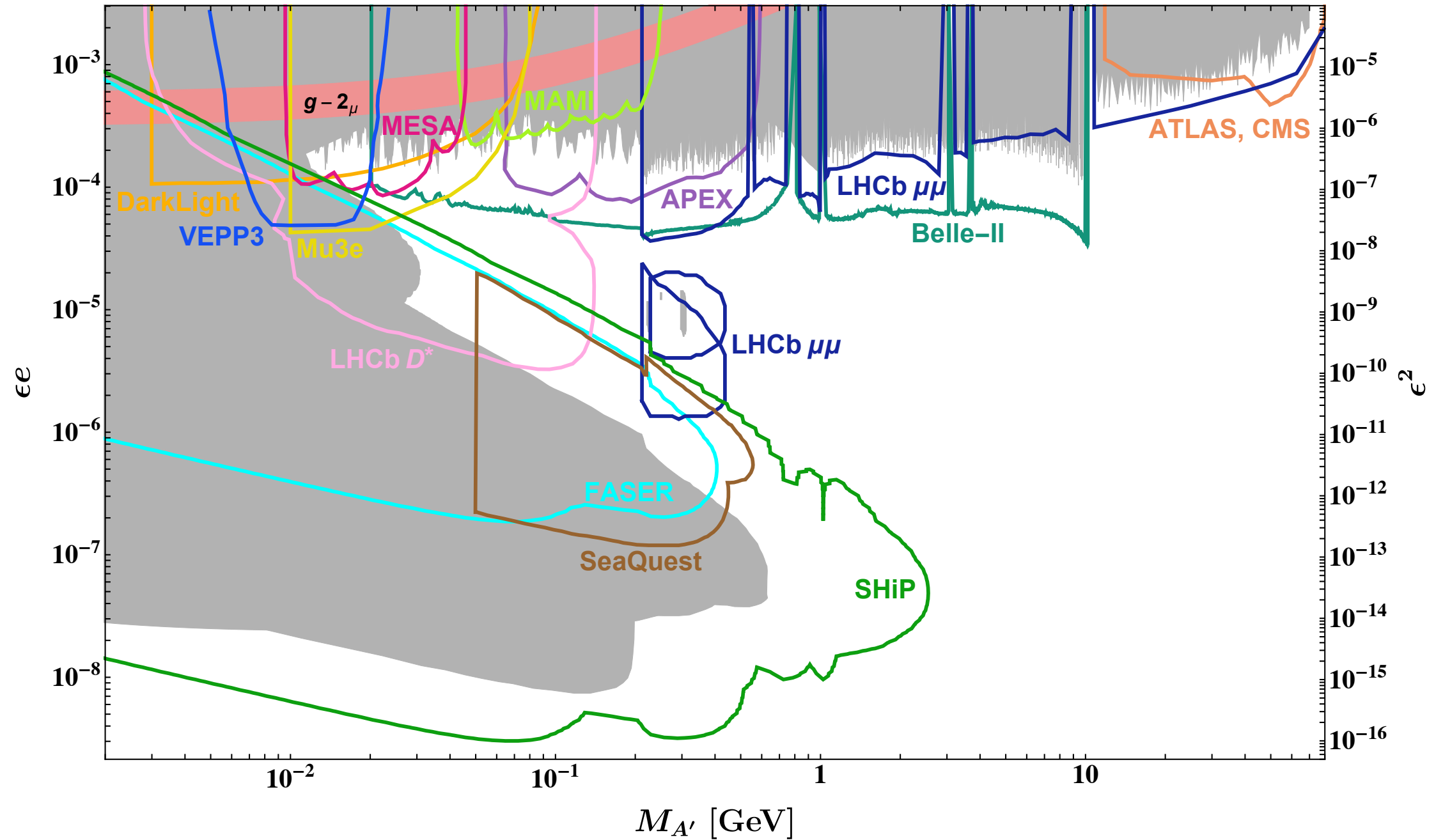


- 1) LLP produced in target
- 2) Passes through shielding
- 3) Decays in detector

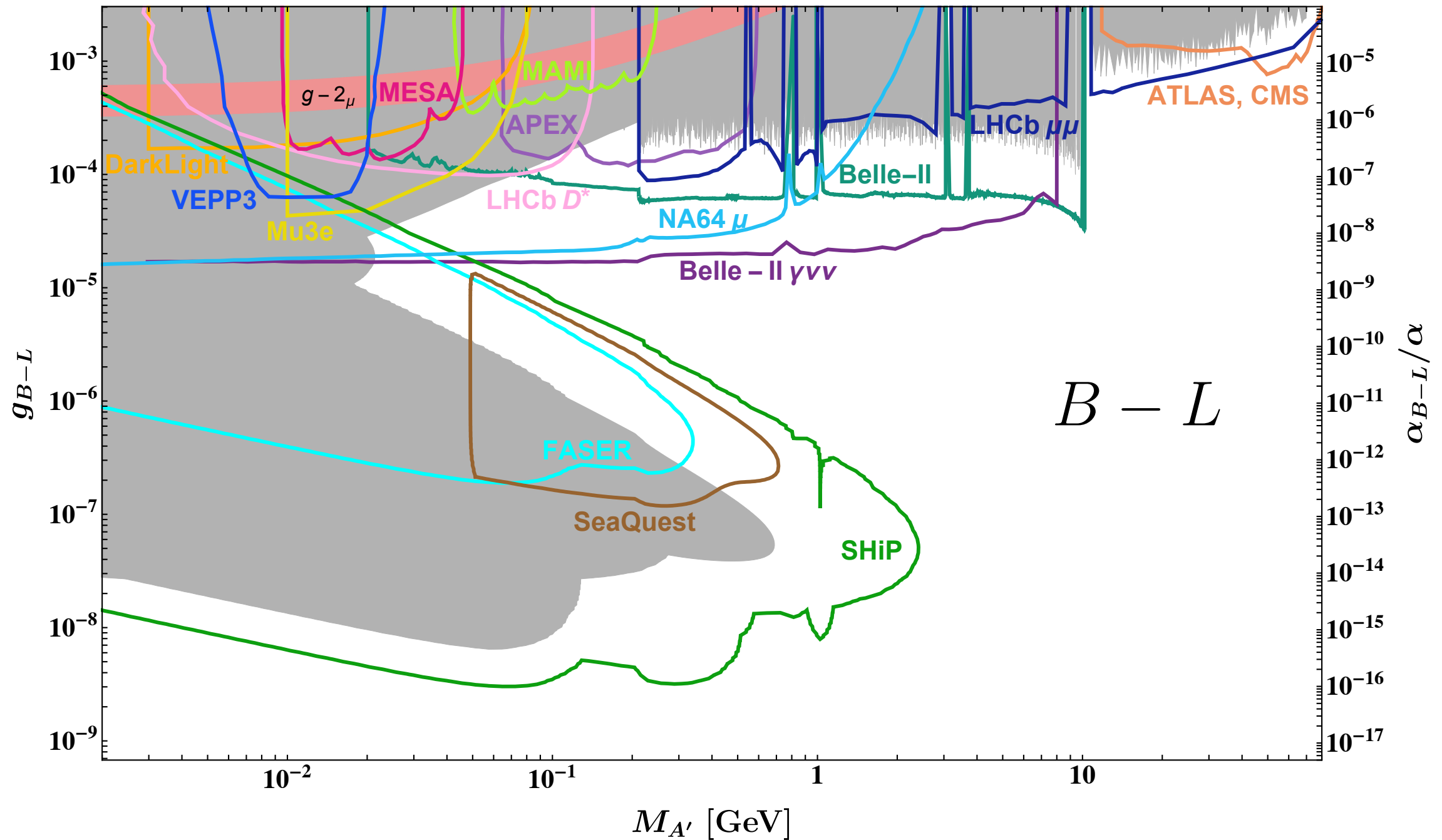


Minimal Kinetically Mixed Dark Photon

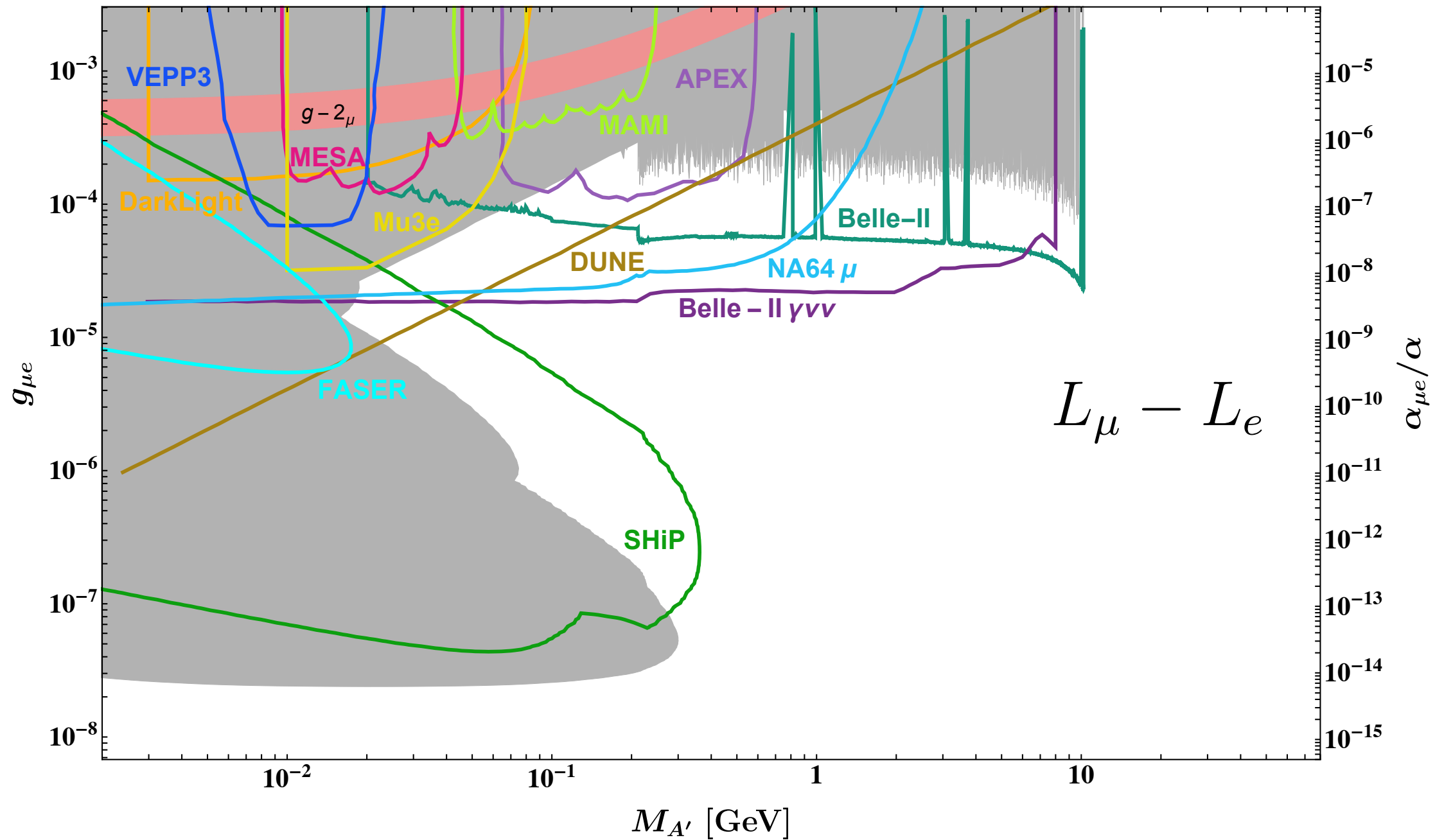
$$F'_{\mu\nu} F^{\mu\nu}$$



Gauged 5th force U(1)

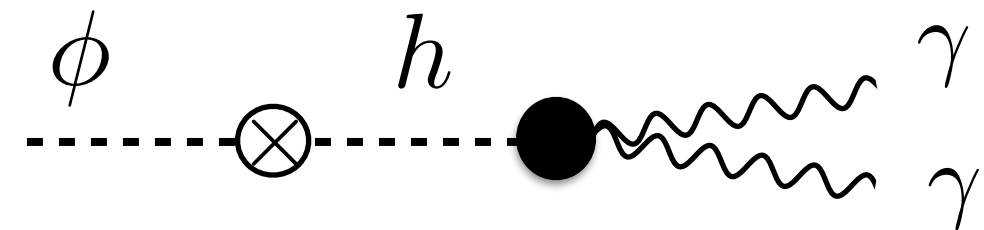
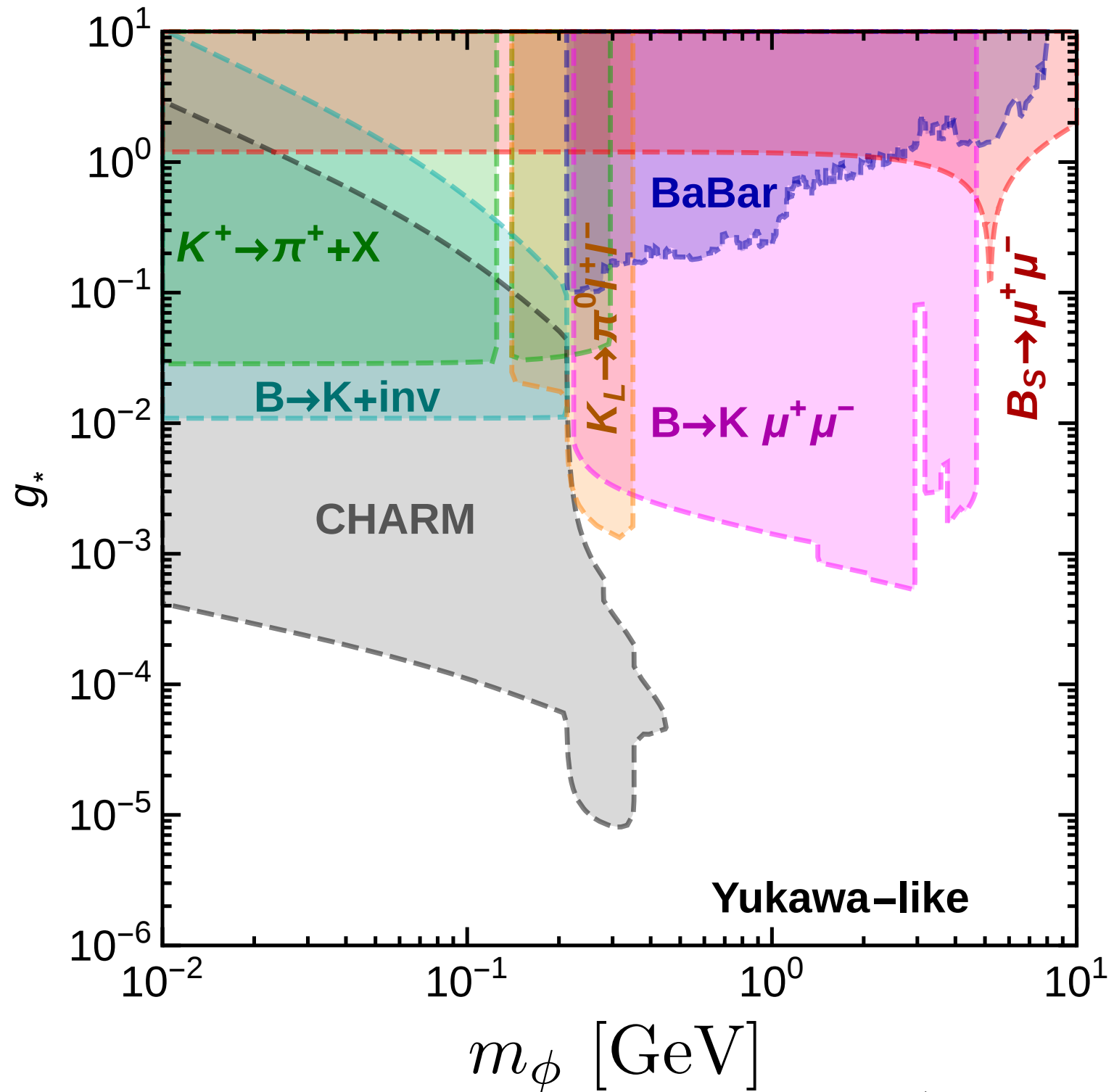


Gauged 5th force U(1)



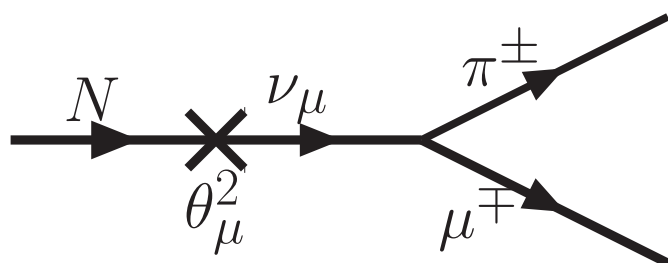
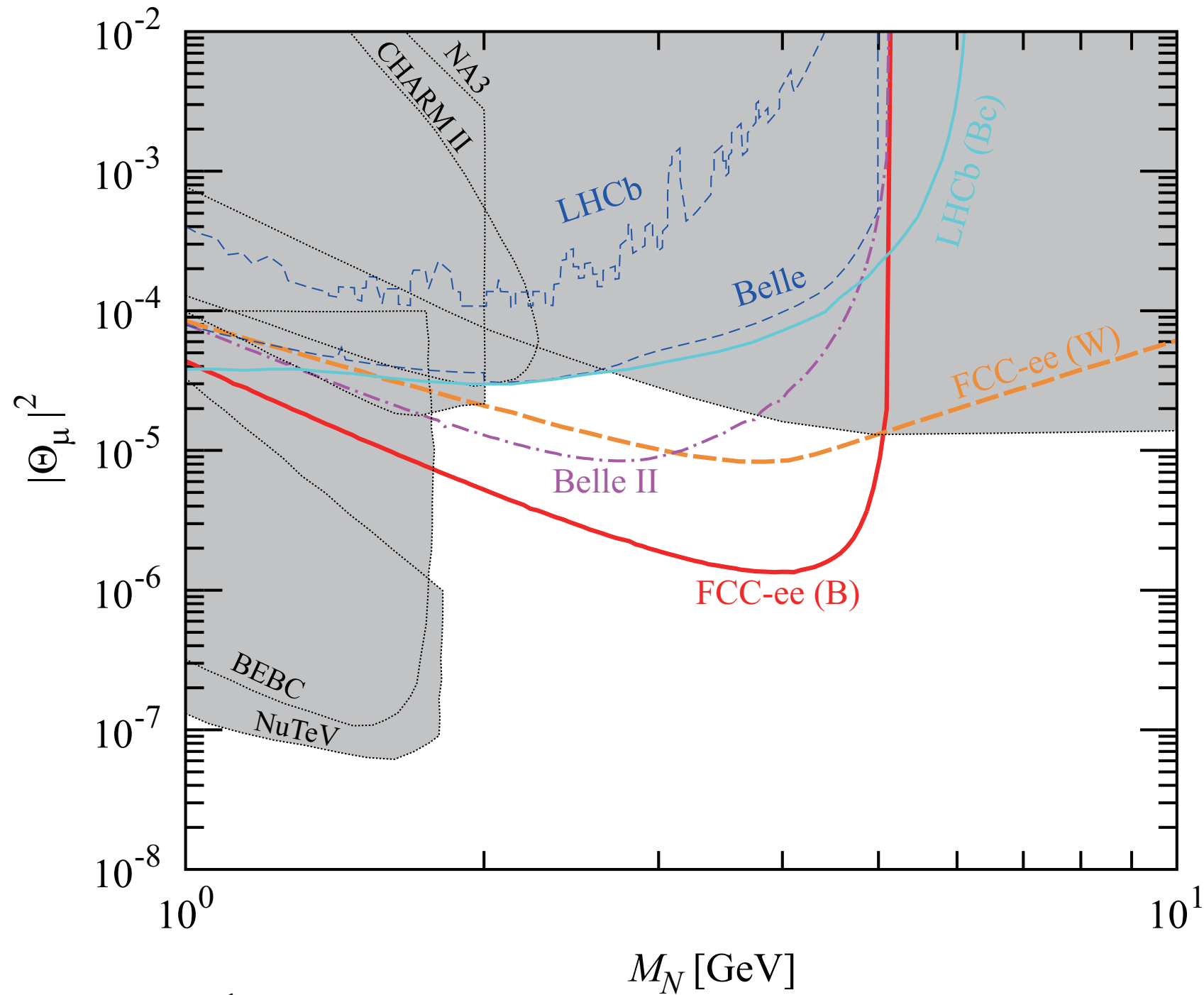
Only scenario that mainly couples to 2nd and 3rd generations

Scalar / Higgs Mixing $\phi H^\dagger H$



Sterile / Active Neutrino Mixing

LHN



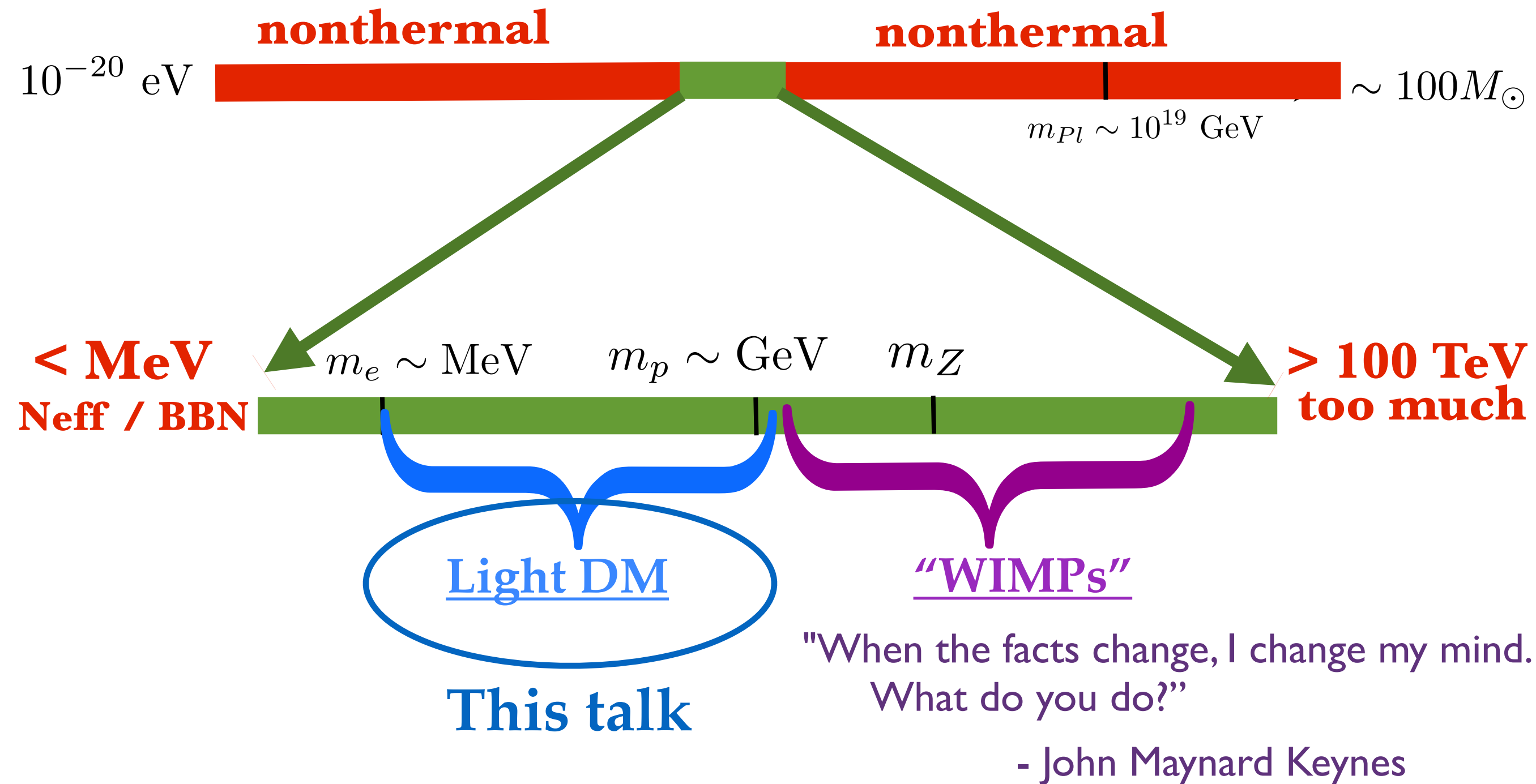
Overview

Part 1) Minimal Single Particle SM Extensions

Part 2) Add Light \sim GeV Dark Matter

Q: What's so great about equilibrium?

A: Narrows Viable Mass Range (!)



Light DM vs. WIMPs

Light DM must be SM neutral

Otherwise would have been discovered at earlier colliders

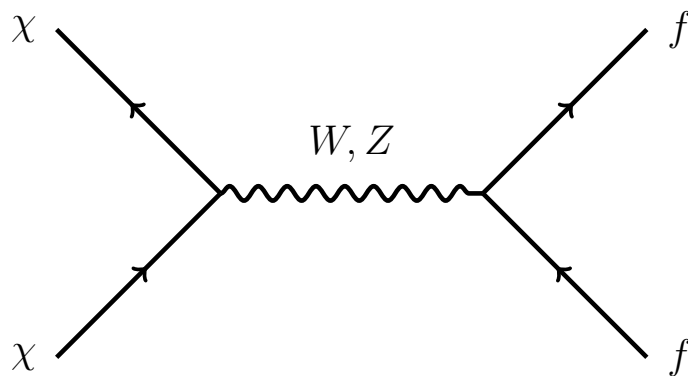
Light DM vs. WIMPs

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Otherwise would have been discovered at earlier colliders

Light DM requires light new force carriers

Overproduced without comparably light, neutral “mediators”



$$\sigma v \sim G_F^2 m_\chi^2 \sim 10^{-29} \text{ cm}^3 \text{ s}^{-1} \left(\frac{m_\chi}{\text{GeV}} \right)^2$$

Always too small if mediator at weak scale

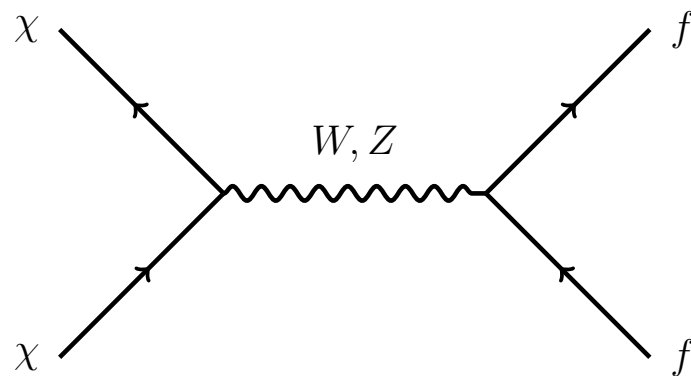
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Annihilation through **renormalizable** interactions

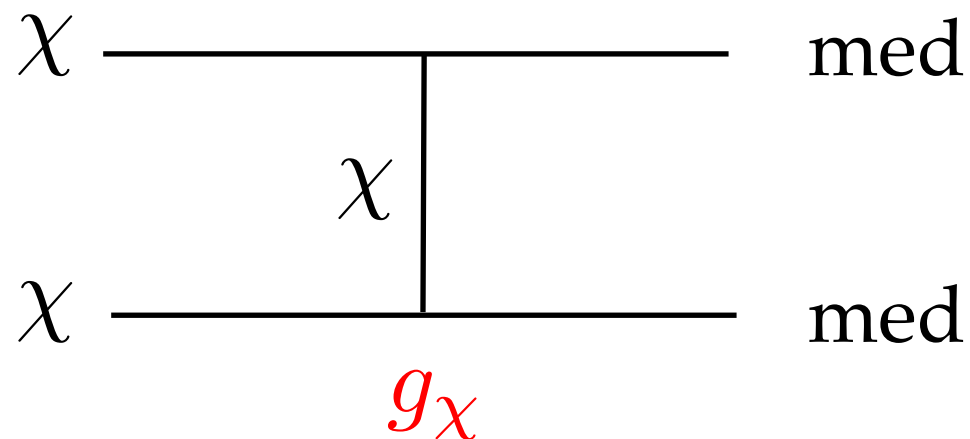
Higher dimension operators have same problem as electroweak forces

Light mediators are not optional!

Who's Heavier: DM or Mediator?

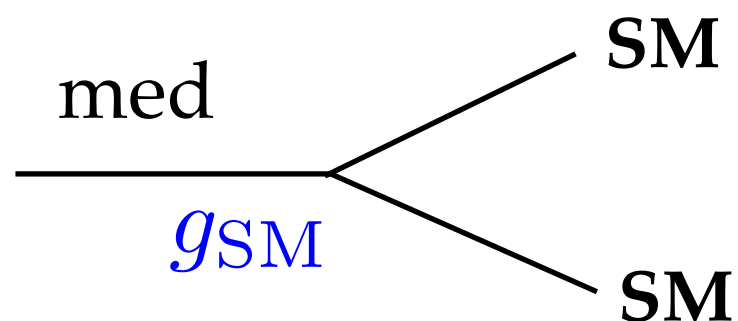
Hidden Annihilation

$$m_\chi > m_{\text{med}}$$



No clear experimental target

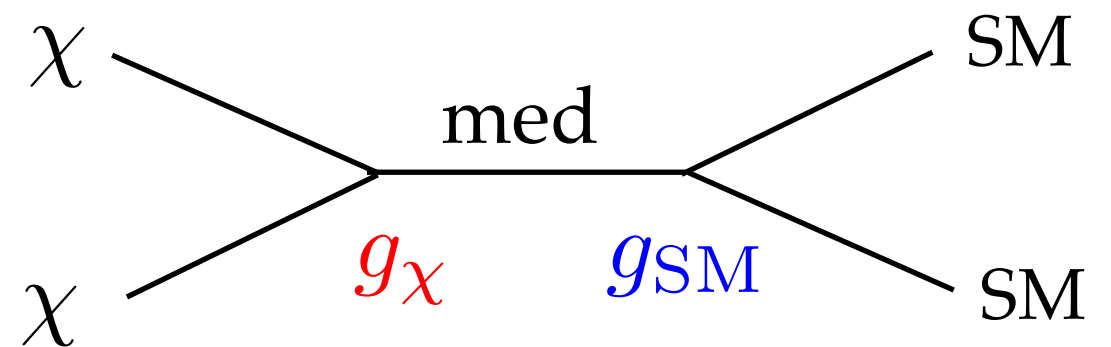
Abundance set by g_χ



Mediator decays **visibly**

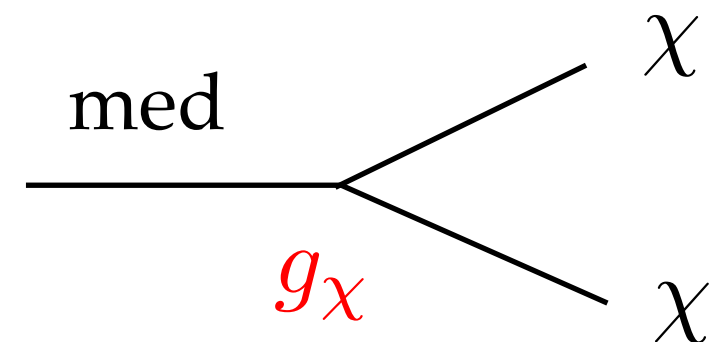
Direct Annihilation

$$m_\chi < m_{\text{med}}$$



Predictive thermal targets

Abundance depends on g_{SM}

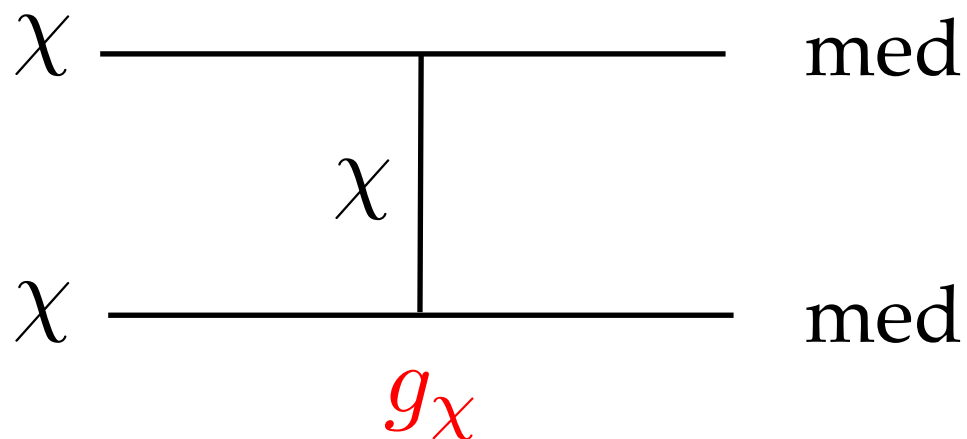


Mediator decays **invisibly***

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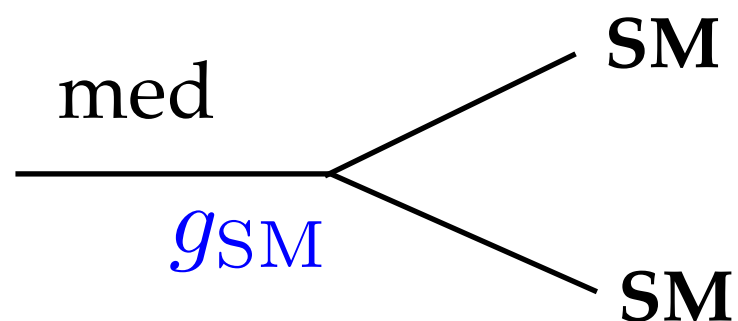
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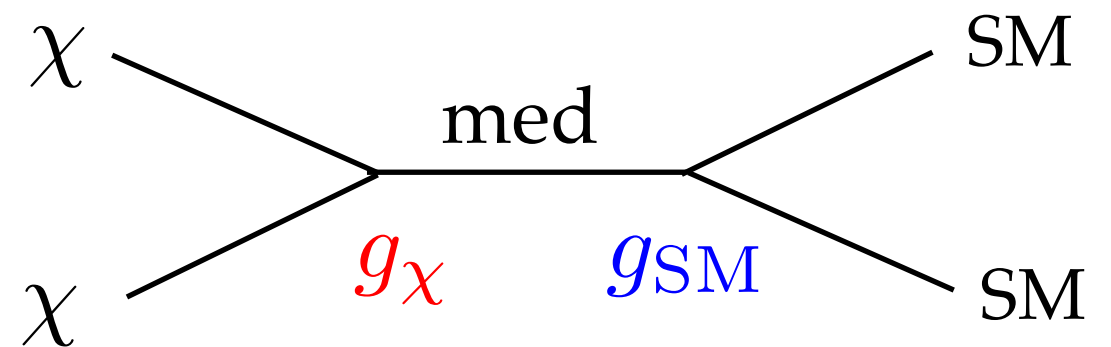
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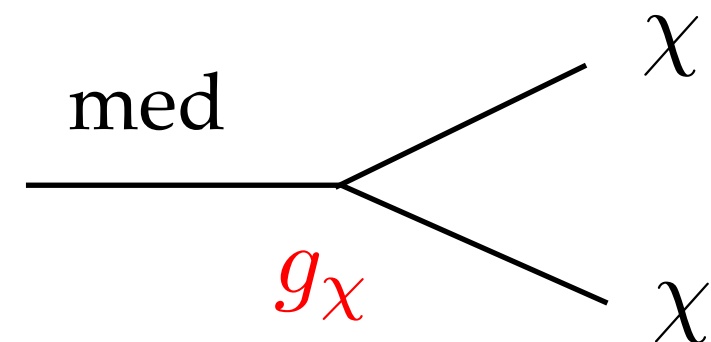
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Predictive thermal targets

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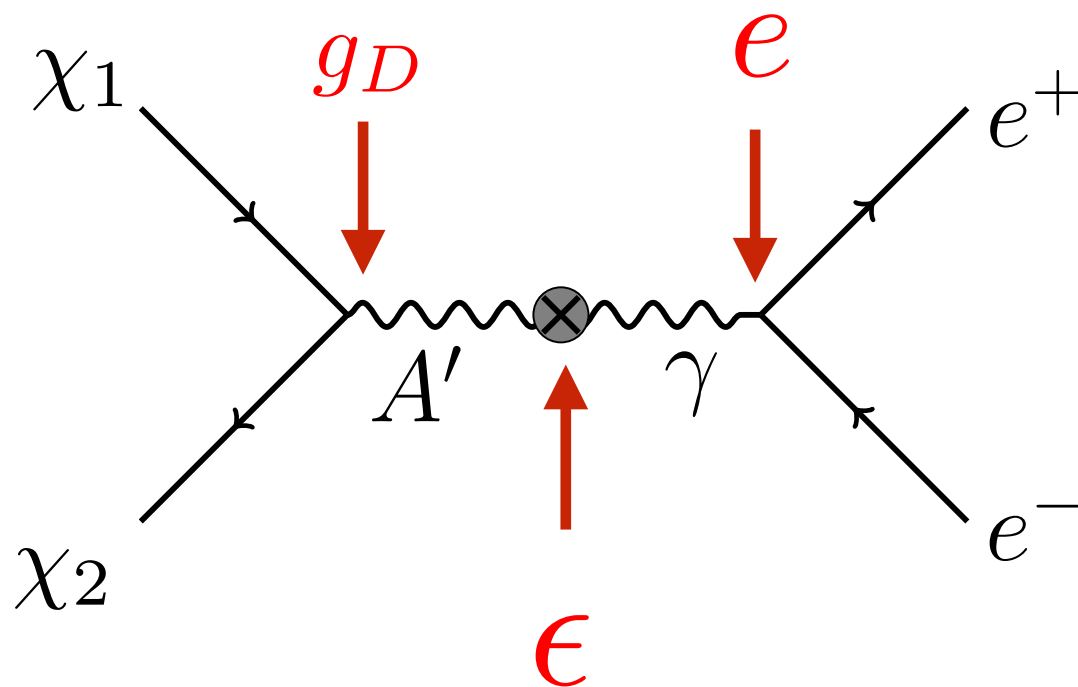
Mediator decays **invisibly***

Representative Model

Dark photon + “pseudo-Dirac” DM current

$$\mathcal{L} \supset g_D A'_\mu \bar{\chi}_2 \gamma^\mu \chi_1 + h.c.$$

Dominant process for relic abundance

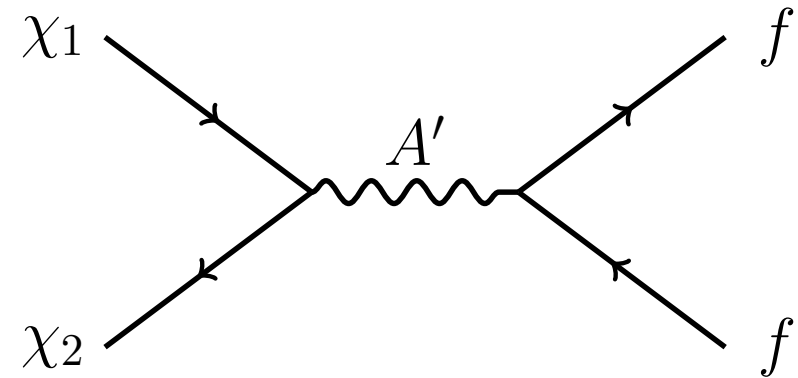


Direct annihilation

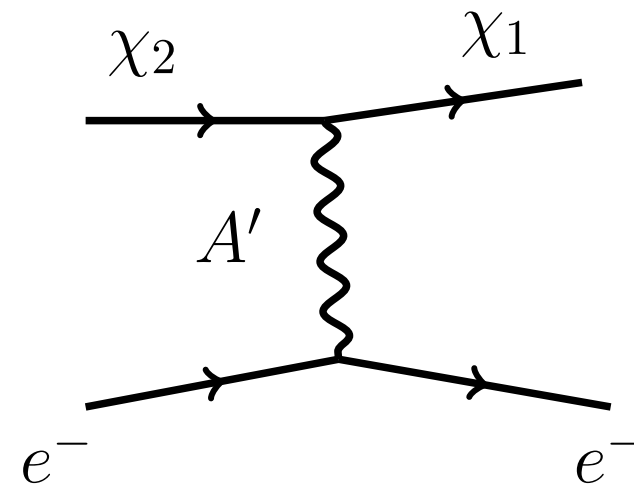
$$m_{A'} > m_1 + m_2$$

Representative Model: Inelastic Dark Matter

Coannihilation

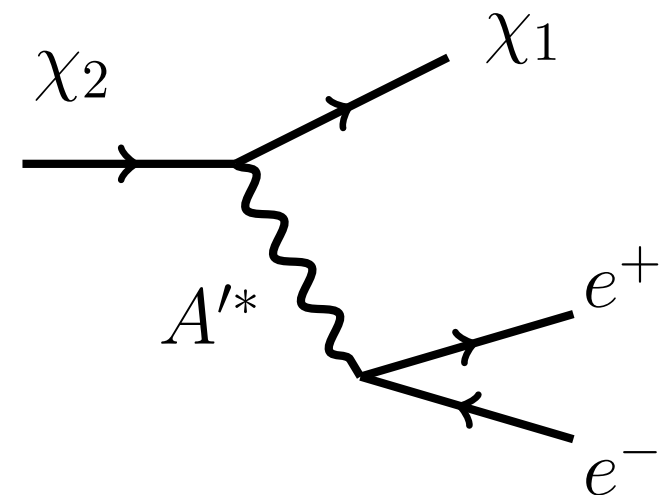


Upscattering +
Downscattering

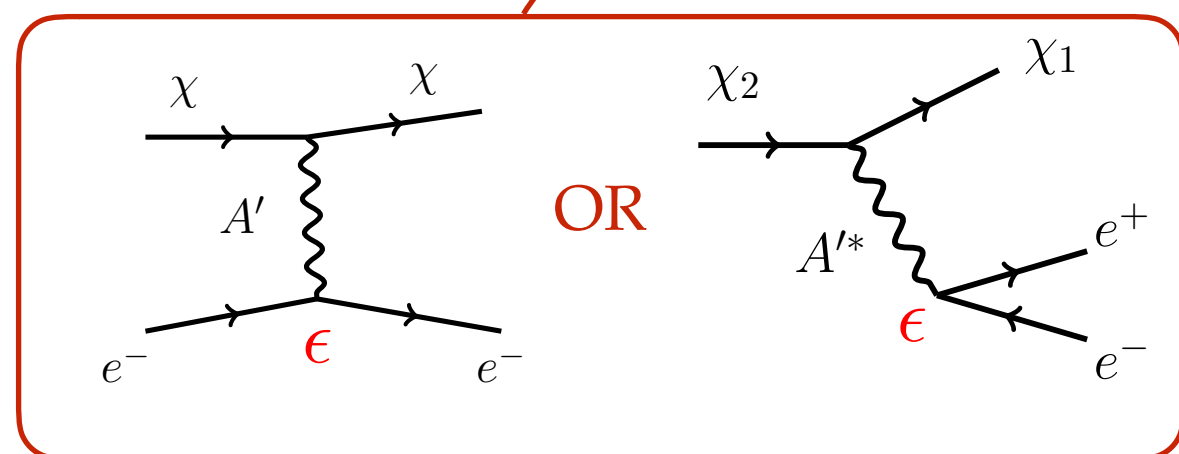
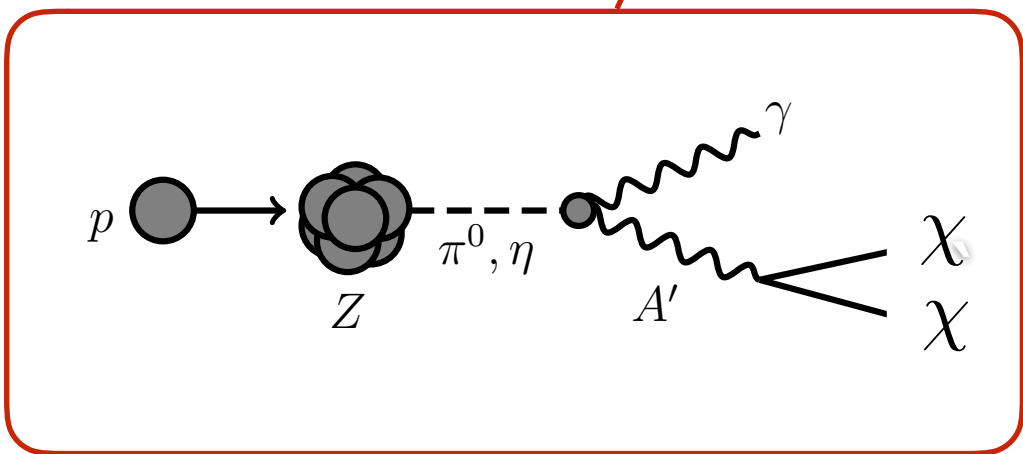
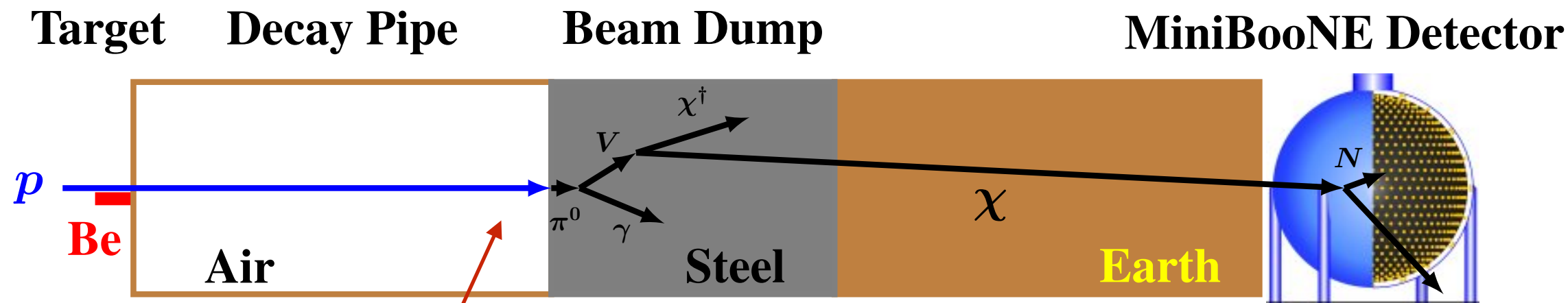


Excited State Decays

$$\Gamma(\chi_2 \rightarrow \chi_1 e^+ e^-) = \frac{4\epsilon^2 \alpha \alpha_D \Delta^5}{15\pi m_{A'}^4}$$



Beam Dump Strategy



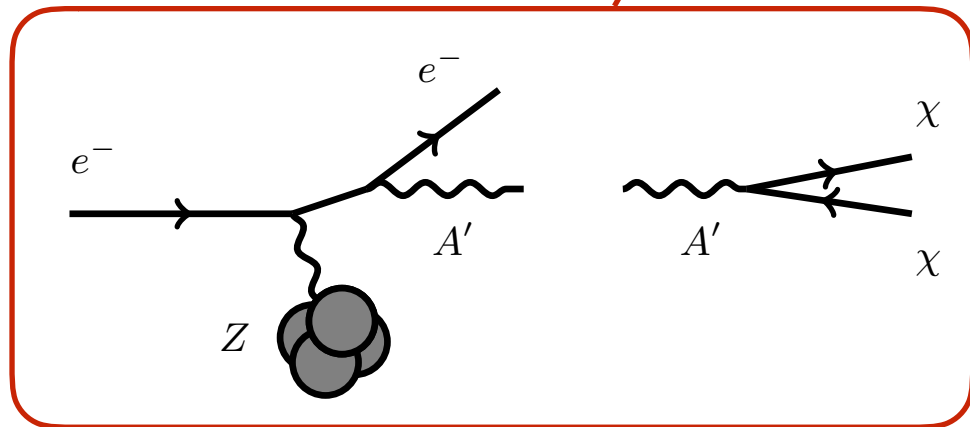
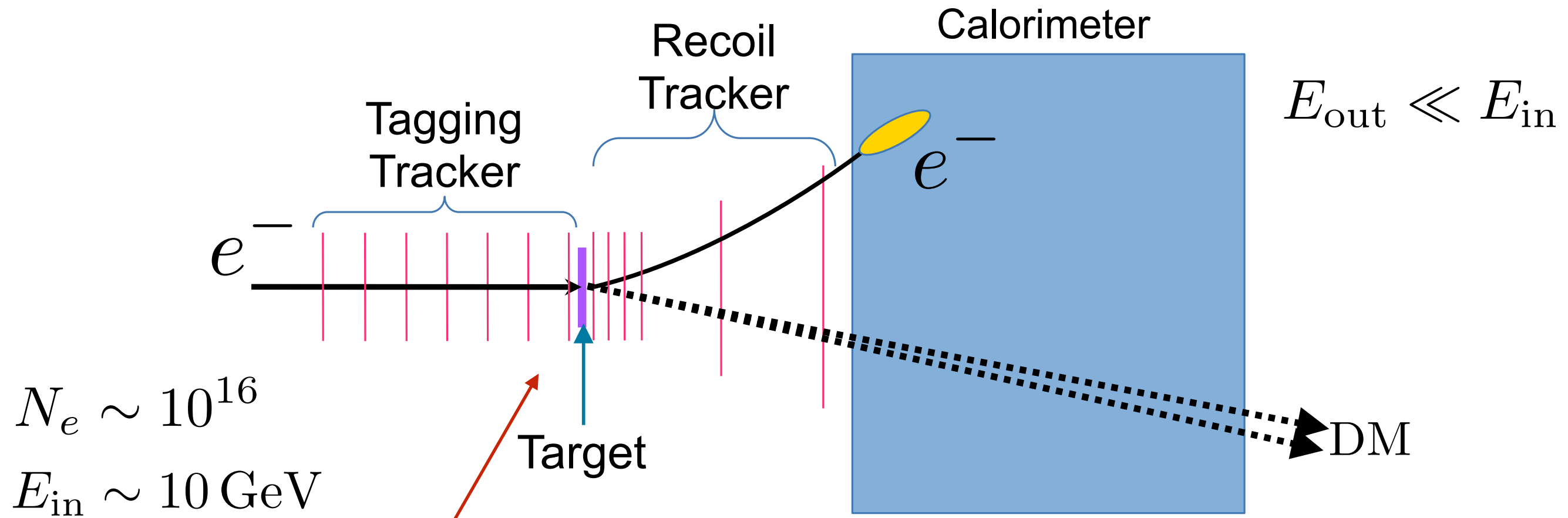
$$N_p \sim 10^{20}$$

$$E_{\text{beam}} \sim 10 \text{ GeV}$$

$$[\text{production}] \times [\text{detection}] \propto \epsilon^4$$

Existing proton beam & neutrino detector
 Relativistic direct detection (no halo)

Missing Momentum Concept



- 1) Measure **each** e- energy in/out
- 2) Trigger on missing momentum
- 3) Veto additional SM activity

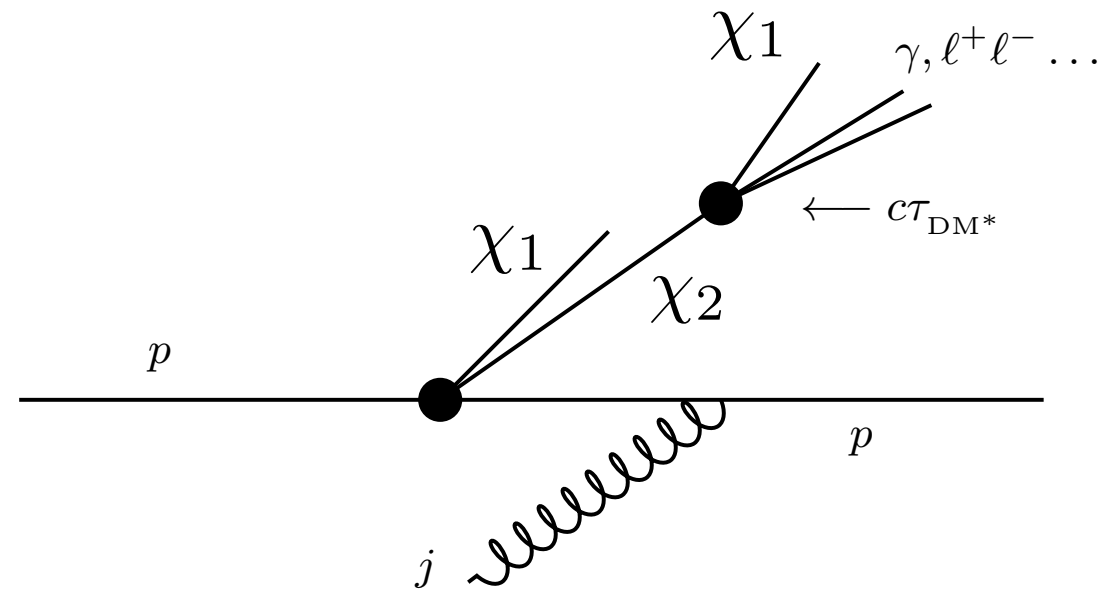
Only measure electron beam — don't require DM to scatter

$$\text{Signal} \propto \epsilon^2$$

Colliders and LLP Displaced Vertices

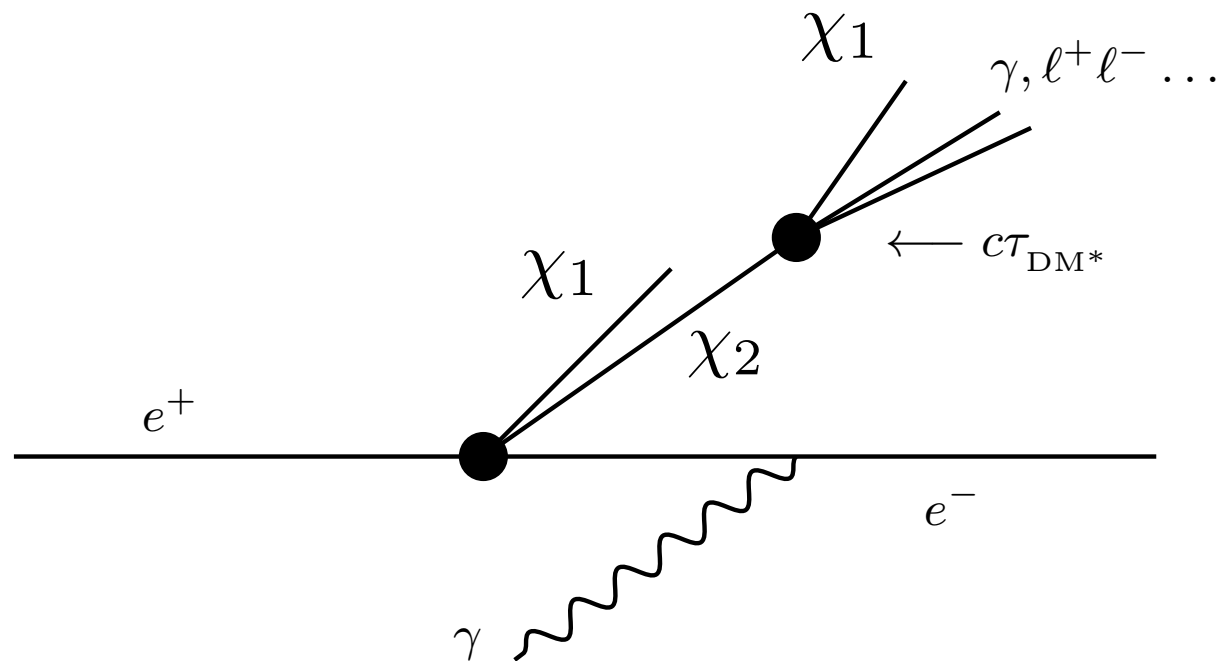
Hadron Collider

$$J + \cancel{E}_T + \ell^+ \ell^-$$

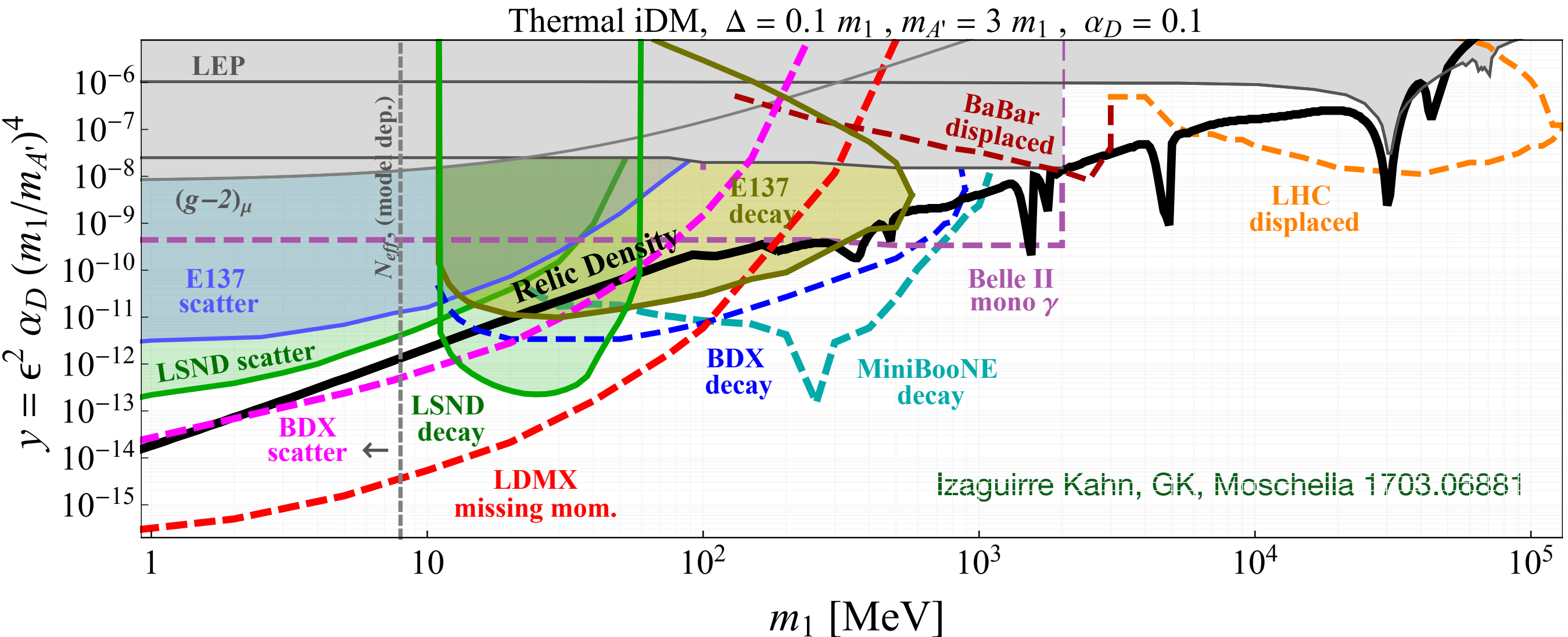


Lepton Collider

$$\gamma + \cancel{E} + \ell^+ \ell^-$$



Testing Thermal DM Production Targets



Broad variety of search strategies required to cover “thermal target”

See also

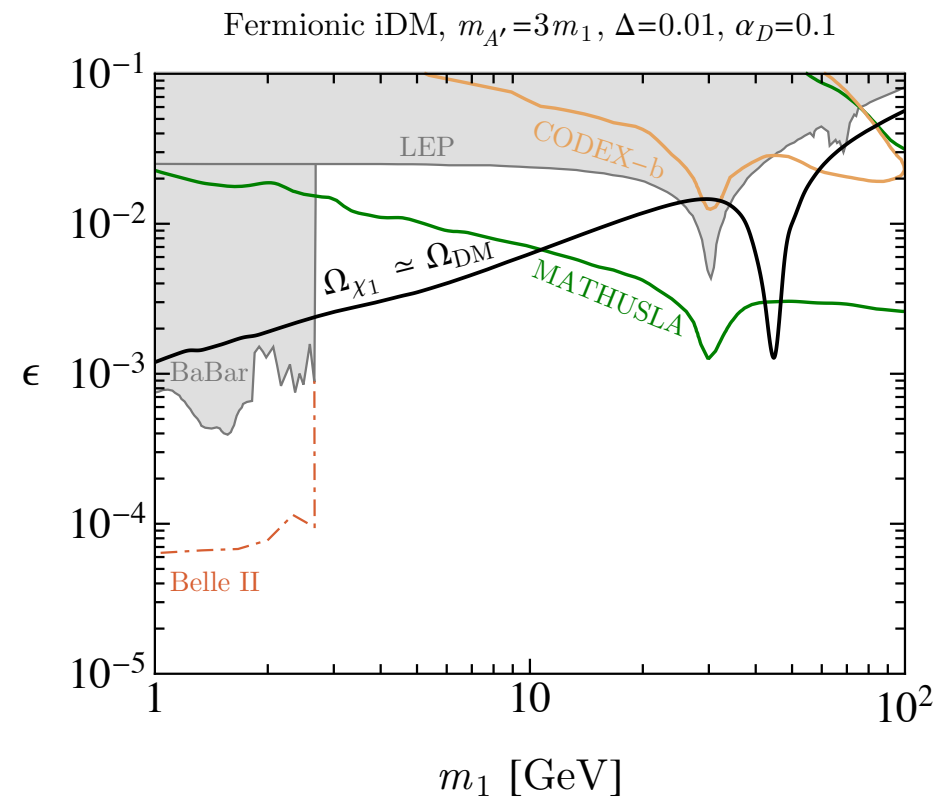
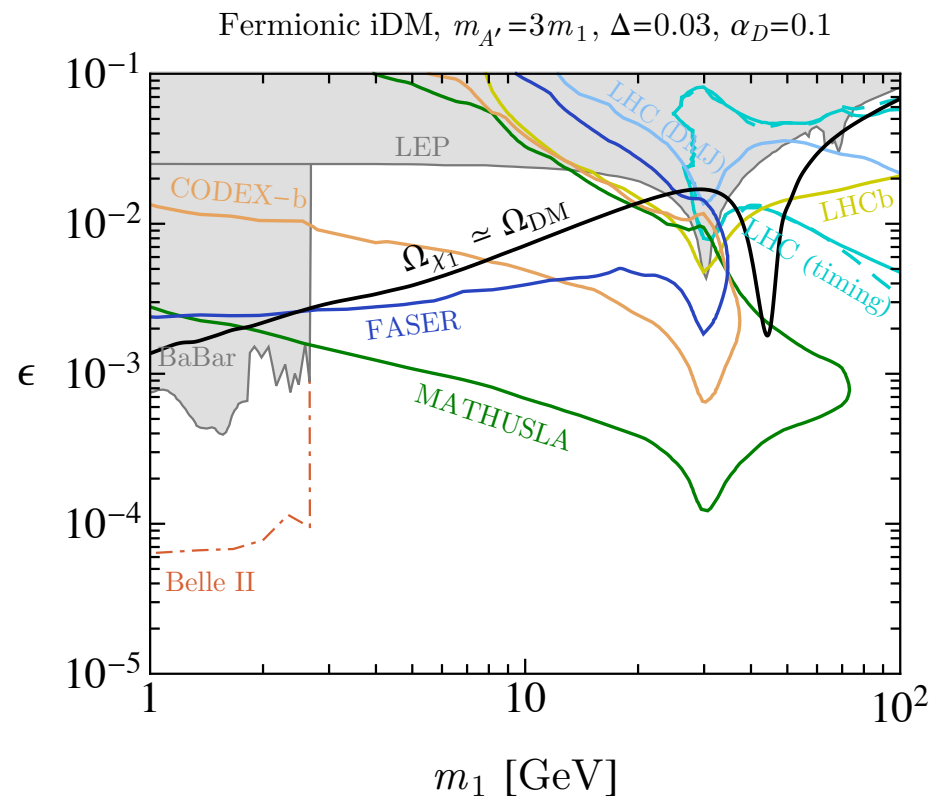
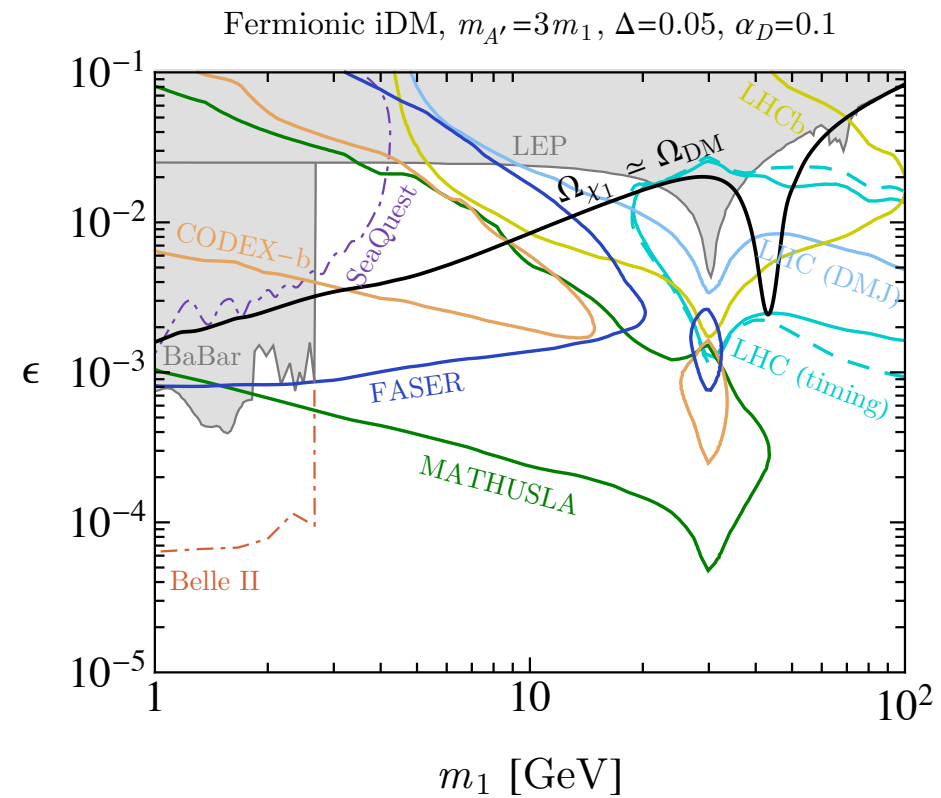
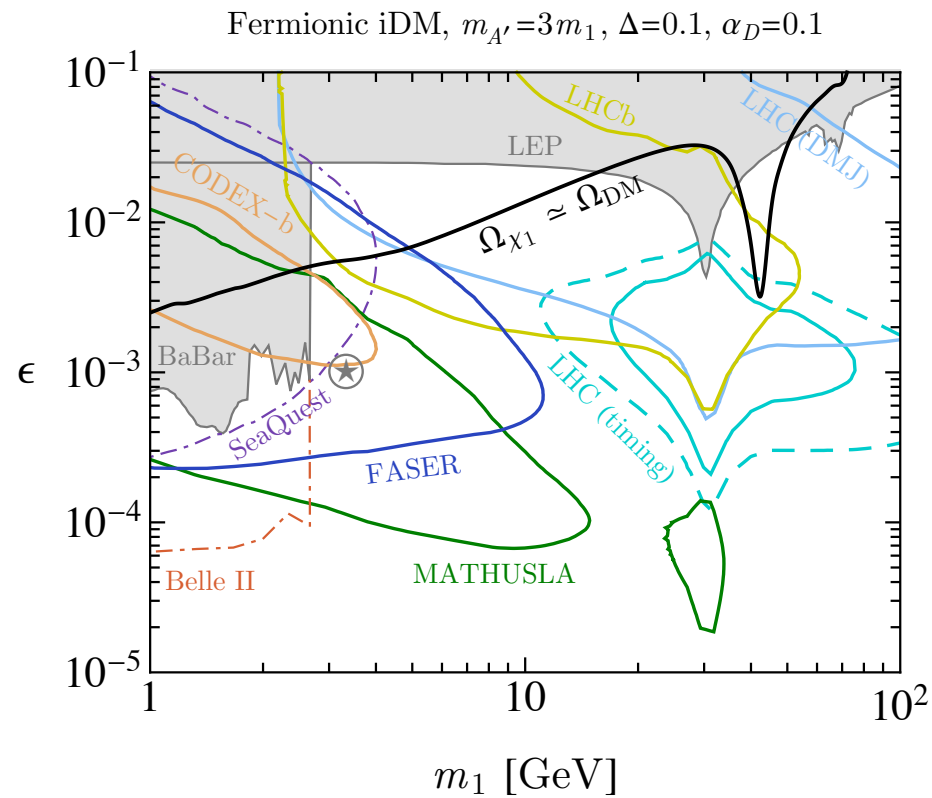
Mohlabeng 1902.05075

deNiverville, Tsai, Liu 1908.07525

Berlin, Kling 1810.01879

and... Yu-Dai Tsai and Felix Kling’s talks

Testing Thermal DM Production Targets



Updated studies: LLP only searches “lifetime frontier”

Berlin, Kling 1810.01879

Concluding Remarks

Broader priors on BSM physics: light weakly coupled states

Minimal single-particle SM extensions

New U(1) forces (e.g. B-L gauge boson)

Mixing with neutral SM states (e.g. sterile neutrino)

Search strategies

Prompt decays at colliders + B-factories

Displaced LLP searches at beam dumps

Adding $< \text{GeV}$ Dark Matter

LLP signatures at colliders from inelastic DM decays

Comprehensively test thermal freeze out via coannihilation