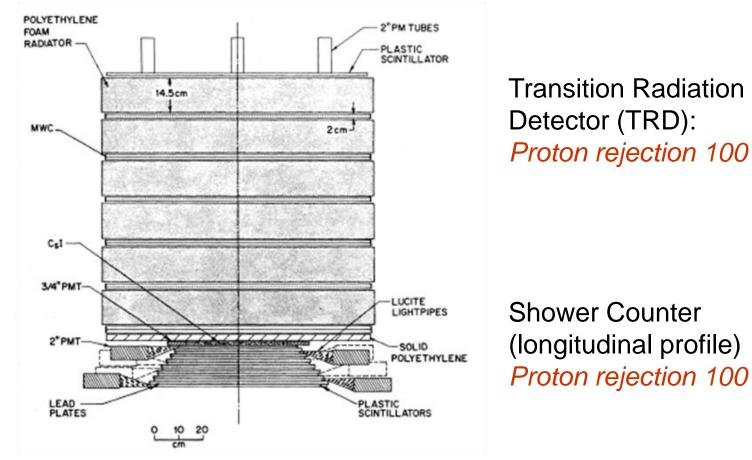
TRD development for hadron identification in multi-TeV energy range.

Mike Cherry Louisiana State University, Baton Rouge, LA USA

With thanks to A. Romaniouk and the ATLAS TR group

Forward Multiparticle Spectrometer meeting, CERN M. Cherry, 4/17/2020 TR used in cosmic ray experiments to separate e, p:

TREE balloon instrument (Chicago 1970s) measured cosmic ray $e^+ + e^-$



TRD 2011 Bar

Transition Radiation Detector (TRD)

Identifies e^{\pm} by transition radiation and Nuclei by dE/dX

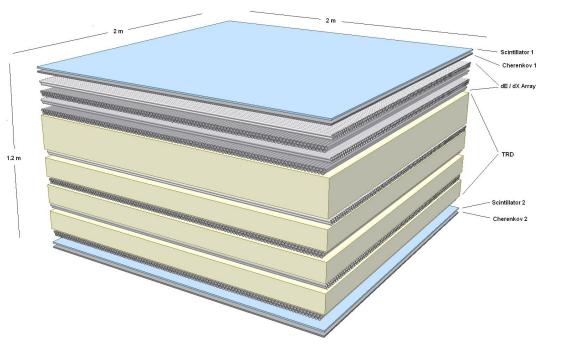
AMS – 02

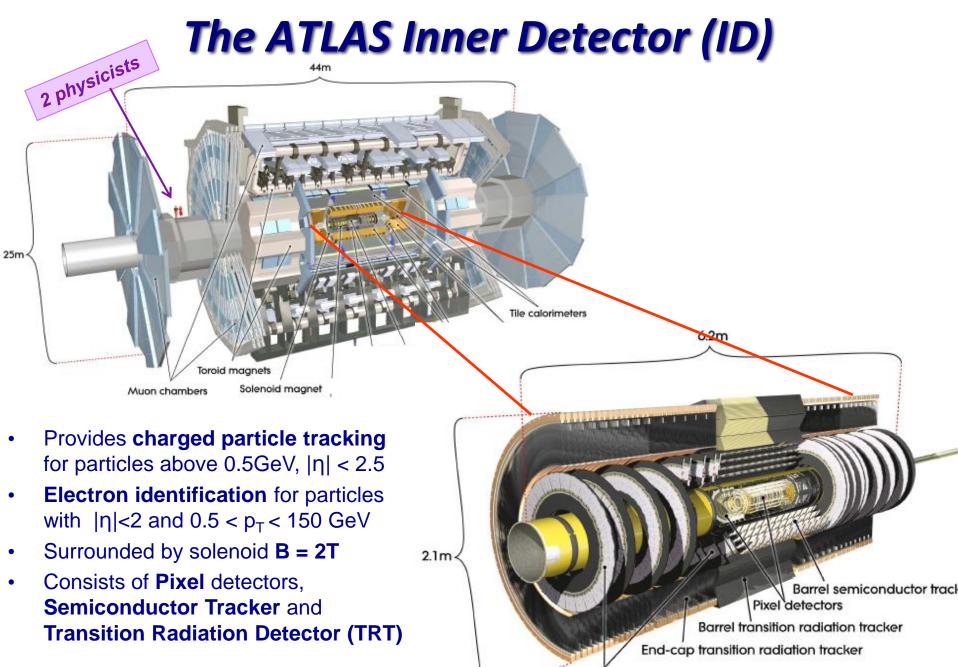
Identifies e[±] vs protons on Intl. Space Station



TRACER --*Transition Radiation Array for Cosmic Energetic Radiation*

Measured energy spectra of cosmic ray nuclei on balloon





End-cap semiconductor tracker

Transition Radiation Detector (TRD)

→ O. Busch



- Parameters :
- Radial position: 2.9 < r < 3.7 m</p>
- |η| < 0.9, 0 <φ< 2π</p>
- 522 modules (18 super-modules) $\rightarrow \sim 675 \text{ m}^2$
- ~25 m³ Xe/CO2 (85:15)
- 1.15 M readout channels
 - Gain Calibration with Krypton source

 \rightarrow J. Stiller

- ≈ 25% X₀
- weight ~ 30 t
- total power: up to 65 kW
- Detector Control System

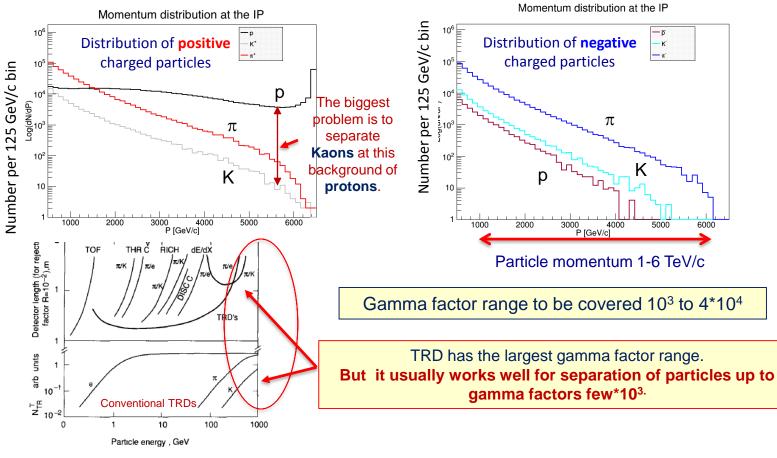
Supermodule Stack (Sector) Chamber (Layer)

ALICE designed to operate at 10 kHz Pb-Pb, few hundred kHz p-p, up to 8000 particles per event

TRDs for the third Millennium, Bari 2011

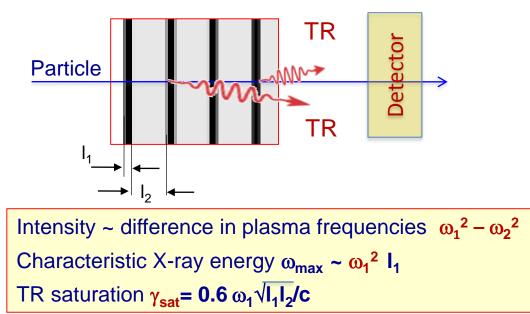
Hadron identification at energies above 1 TeV.

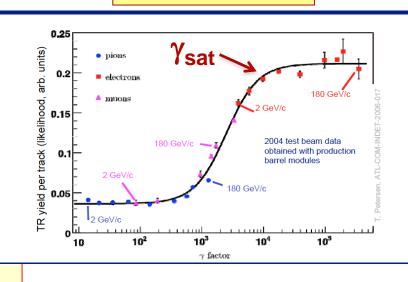
Expected hadron spectrum in FHS experiment.



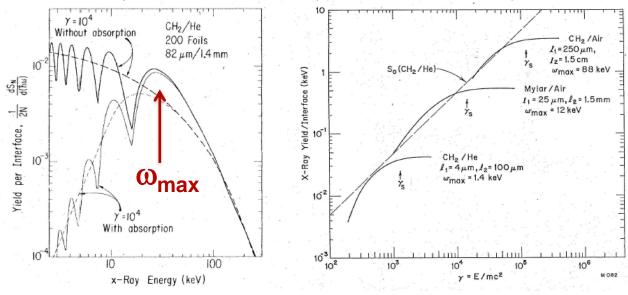
https://doi.org/10.1016/0168-9002(93)90846-A

TR: Tuning performance with detector parameters





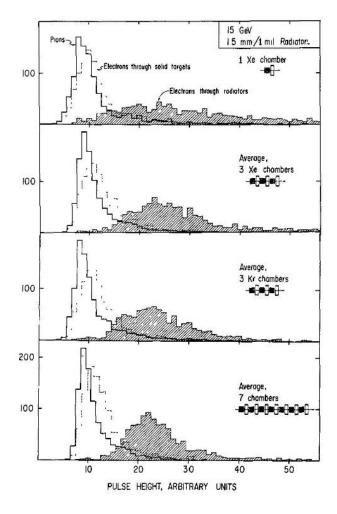
Modifying parameters I_1 , I_2 and ρ (ω_1) one obtains different gamma factor dependences.

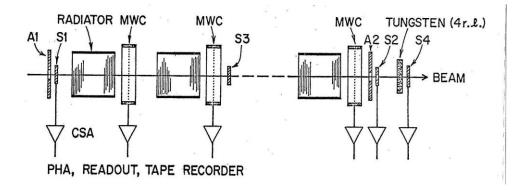


In X-ray energy range

Particle id at accelerators: Yerevan, BNL, SLAC, CERN, DESY, ...

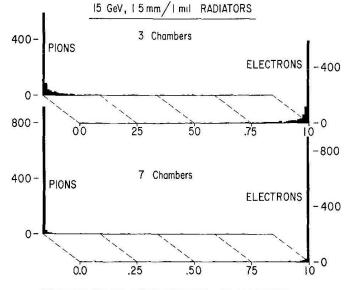
Separating e⁻/π at SLAC, 1973:





Prob. of "e- or π -like" distribution in 7 detectors: $P_1 = P_e^1 \cdot P_e^2 \cdot \cdot \cdot P_e^7$ $P_2 = P_\pi^1 \cdot P_\pi^2 \cdot \cdot \cdot P_\pi^7$

Likelihood of event being e or π : $L_e = P_1/(P_1+P_2)$ $L_{\pi} = P_2/(P_1+P_2) = 1 - L_e$



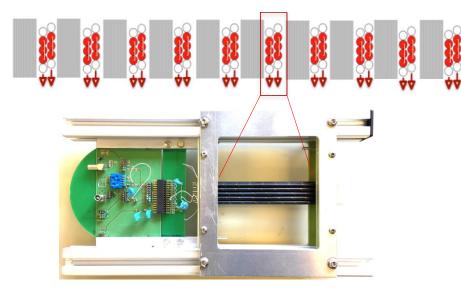
PROBABILITY OF INTERPRETATION AS ELECTRON

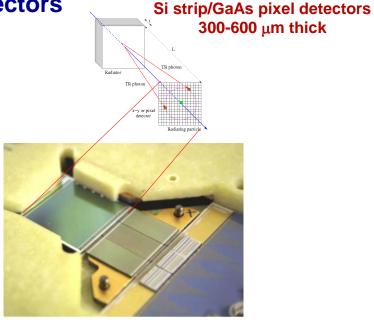
SAS_TRD: Test beams 2017, 2018

Major **goal: to verify MC models of TR production** as function of gamma factors: spectra and angular distributions

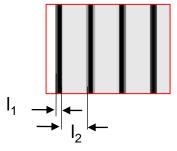
Two types of the detectors

Straw based with Xe gas



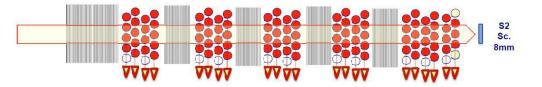


Few types of radiators: Different parameters I_1 and I_2 will be used



Tests in 2017

Straw based prototype





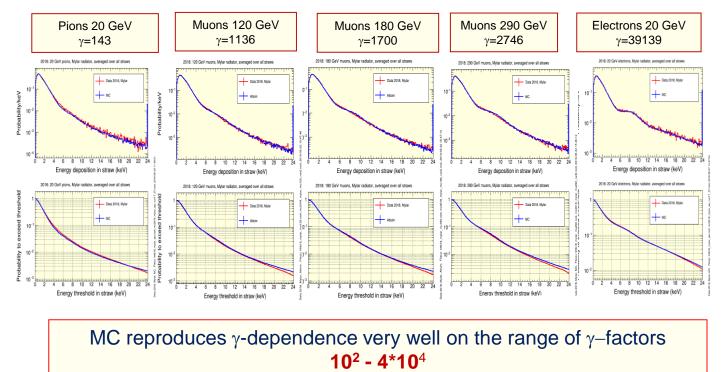
Anatoli Romaniouk, TRD testbeam 2017

External detectors (not shown here): scintillator multiplicity counter, Cherenkov counter, pre-shower and leadglass calorimeter were used to select and purify the identification of the beam particle sort.

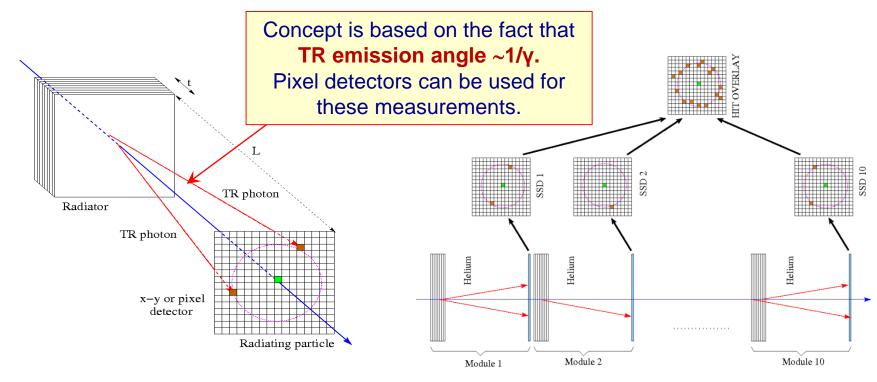
- Prototype with 4 mm diameter straws and Xe-٠ based gas mixture (71.8% Xe, 25.6% CO₂, 2.6% O_2)
- Four or five straws in one layer were jointly connected to readout. Note: everywhere in the text below "straw number" really means "straw layer number".
- Each straw layer (column on the picture) was shifted in vertical direction with respect to each other to minimize fluctuations of the active gas thickness crossed by the beam particles.
- 8 mm scintillator to trigger beam particles. •
- Beam particles: 20 GeV pions, 20 GeV electrons, 120 GeV muons, 180 GeV muons, 290 GeV muons
- Radiators:
- 30 foils in each radiator block (55 foils in very first block)
 - ✓ Mylar 50 µm foils, 3 mm gap
 - Polyethylene (PE) 270 µm/3.3mm \checkmark
 - ✓ Polypropylene (PP): 62 µm foil / 2.2 mm gap (NB: Compound foils in this radiator made from four 15 µm foils)

Data/MC comparison: Mylar radiator, different particles.

Differential (top row) and Integral (bottom row) energy spectra in straws. Mylar radiator 50 um 15 foils

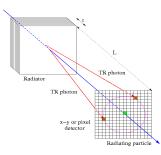


Angle TRD: concept

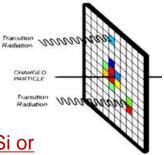


Advantages:

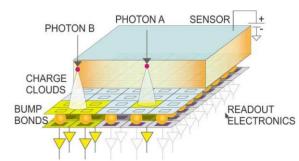
- Combination angle and TR energy information would significantly improve identification power of such kind detectors.
- This approach would allow to minimize material budget
- It combines tracking and TRD functions in one detector



II. TRD based on high granular semiconductor technology.



<u>TimePix3 front-end chip attached to Si or</u> <u>GaAs sensors.</u>





Timepix3 front-end hybrid pixel readout chip:

- · Various sensor materials possible.
- Simultaneous per-pixel measurement of a time-of-arrival (ToA) and an energy (time-over-threshold technique).
- Time resolution of 1.56ns
- Spatial resolution of ~16µm
- + 256 x 256 pixel matrix with 55 x 55 μm2 pitch
- throughput of up to 40 Mhits/s/cm2

TimePix4 with improved time measurements is coming soon (see later)

XTR energy [keV]

XTR energy [ke/

0.0005

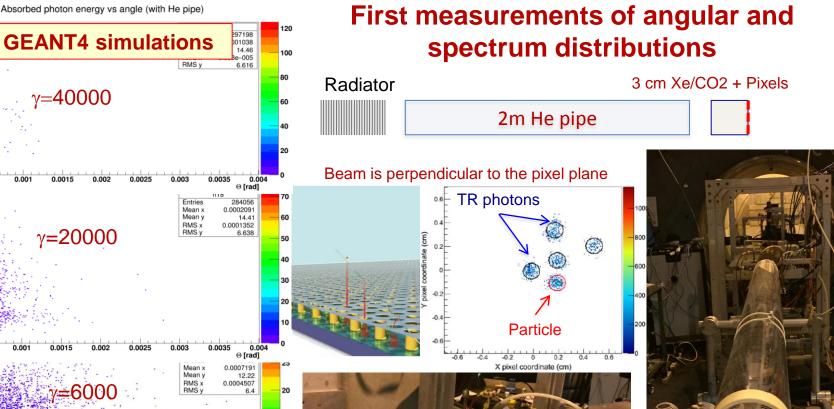
0.0005

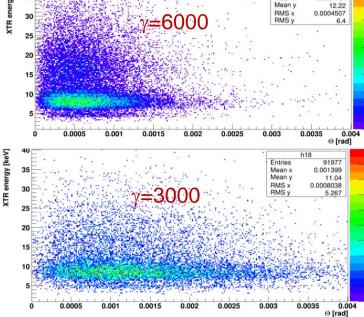
0.001

0.001

0.0015

0.0015





First detailed studies of TR energy/angle spectra with silicon Timepix3 detector –

Schioppa et al. NIM A 936, 523 (2019) Dachs et al., NIM A 958, 162037 (2020) Alozy et al, NIM A 961, 163681 (2020)

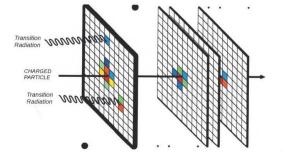
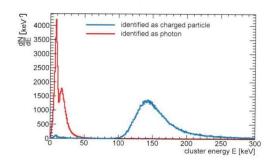
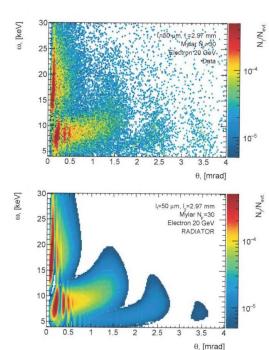


Figure 3: Sketch of a Transition Radiation event in the Timepix3 system.





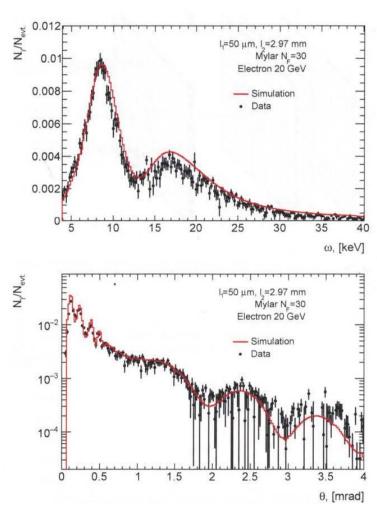
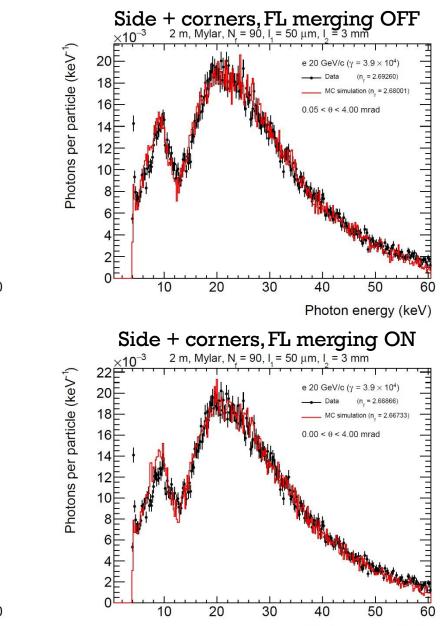


Figure 5: Energy versus angle distribution of TR photons for Mylar radiator consisting of 30 foils of 50 μ m thick sparated by 2.97 mm: data (top) and simulation (bottom)

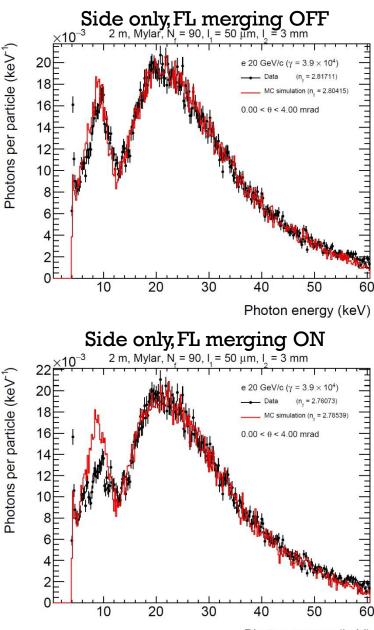
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Figure 4: Distributions of energy for all the clusters identified as the charged particle (blue) and as transition radiation (red) in events with an electron trigger.

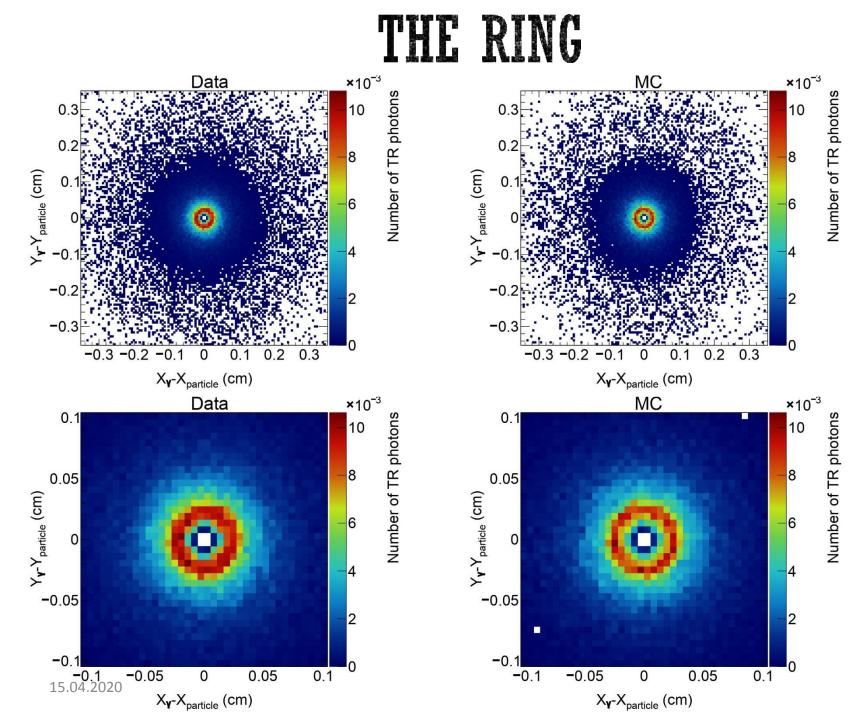
ENERGY SPECTRA

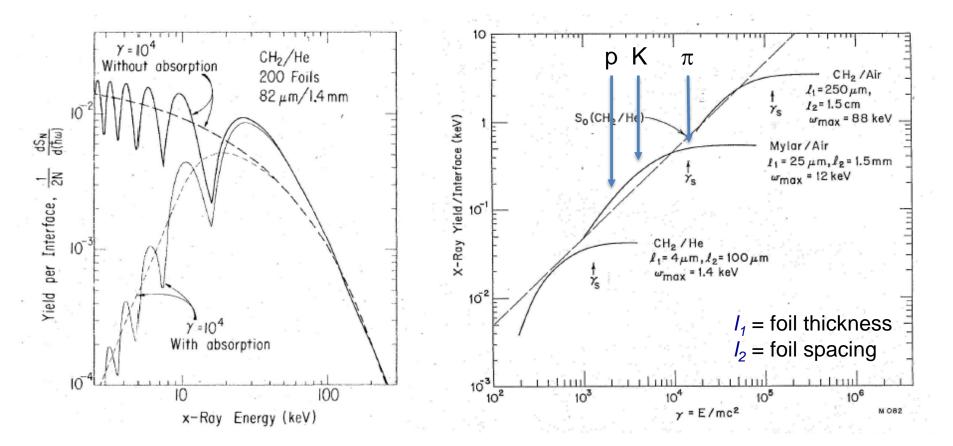


Photon energy (keV)



Photon energy (keV)



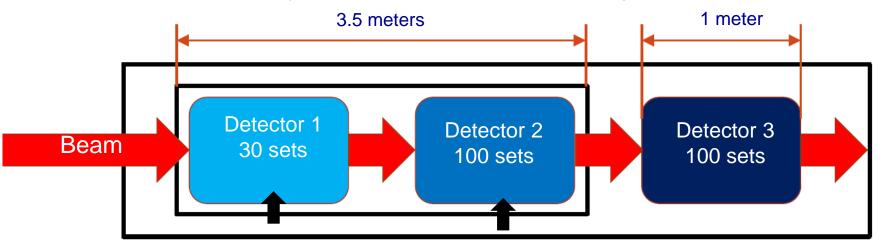


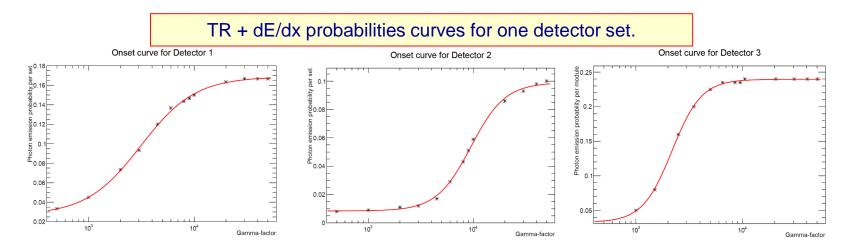
TR saturation $\gamma_s \sim 0.6 \omega_1 \sqrt{I_1 I_2} / c$ Characteristic X-ray energy $\omega_{max} = \omega_1^2 I_1 / 2\pi c$

| E (TeV) | γ_{p} | γ_{K} | $\gamma_{\pi\pm}$ |
|---------|--------------|--------------|-------------------|
| 1 | 1.07É+03 | 2.02E+03 | 7.16E+03 |
| 2 | 2.14E+03 | 4.04E+03 | 1.43E+04 |
| 6 | 6.40E+03 | 1.21E+03 | 4.30E+04 |

SAS_TRD: beam composition measurements

No optimization just some assumptions based on TR simulations (no angular information used) Let's suppose we have 3 types of the detectors with different gamma factors dependences





Anatoli Romaniouk, XSCRC2017: Cross sections for Cosmic Rays @ CERN, 31.03.2017

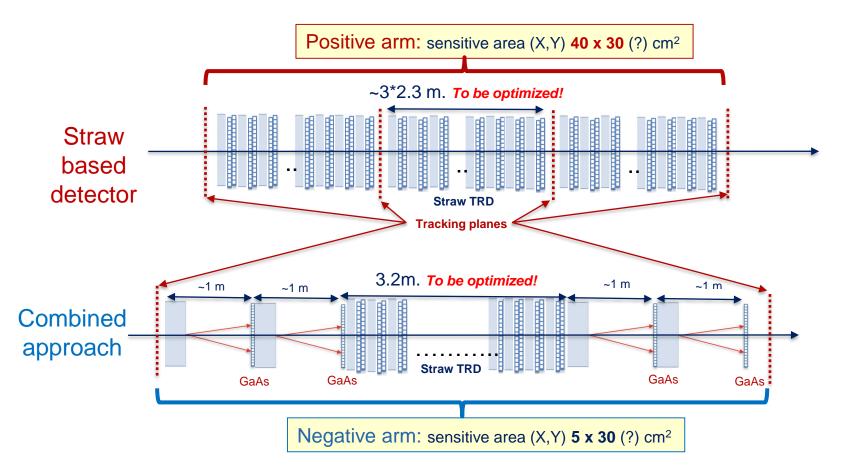
Recent design work for Forward Scattering Spectrometer led by ATLAS TR group:

- Measurement of angular distribution
- Development of straw tubes, Si strip/GaAs pixel detectors

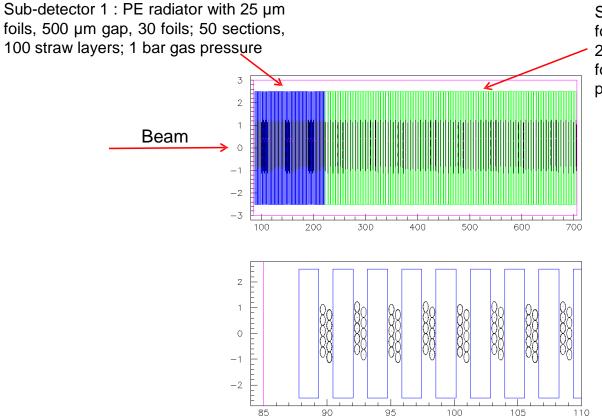
Final design depends on details of configuration (e.g., real estate available), but

- we know how to tune detectors to obtain desired particle id and energy measurement
- "Measurements of Compton Scattered Transition Radiation at High Lorentz Factors", G.L. Case et al., NIM A <u>524</u>, 257 (2004).
- "Measuring the Lorentz Factors of Energetic Particles with Transition Radiation", M.L. Cherry, NIM A <u>706</u>, 39 (2013).
- "First Measurements of the Spectral and Angular Distribution of Transition Radiation Using a Silicon Pixel Sensor on a Timepix3 Chip", E. J. Schioppa et al., NIM A <u>936</u>, 523 (2019).
- "Identification of Particles with Lorentz factor up to 10⁴ with Transition Radiation Detectors Based on Micro-strip Silicon Detectors", J. Alozy et al., NIM A <u>927</u>, 1 (2019).
- "Transition Radiation Measurements with a Si and a GaAs Pixel Sensor on a Timepix3 Chip", F. Dachs et al., NIM A <u>958</u>, 162037 (2020).
- "Studies of the Spectral and Angular Distributions of Transition Radiation Using a Silicon Pixel Sensor on a Timepix3 Chip", J. Alozy et al., NIM A <u>961</u>, 163681 (2020).

TRD for FHS concept (2 concepts)



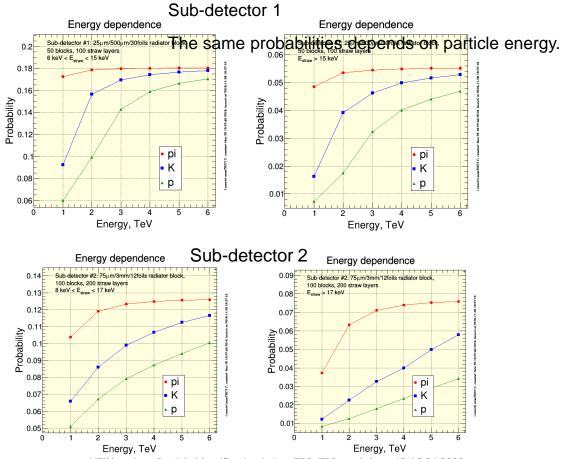
Detector configuration



Sub-detector 2 : PE radiator with 75 µm foils, 3 mm gap, 12 foils; 100 sections, 200 straw layers; 1.5 bar gas pressure for better absorption of high-energy TR photons

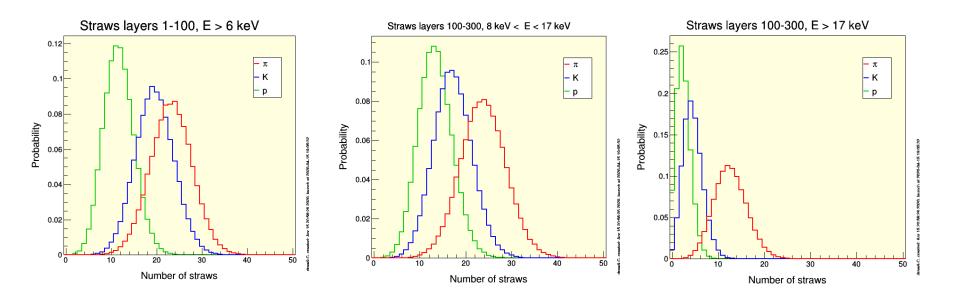
V.Tikhomirov. Particle identification in LargTRD. TRD workshop, 15-16.04.2020

Dependencies vs particle energy



V.Tikhomirov. Particle identification in LargTRD. TRD workshop, 15-16.04.2020

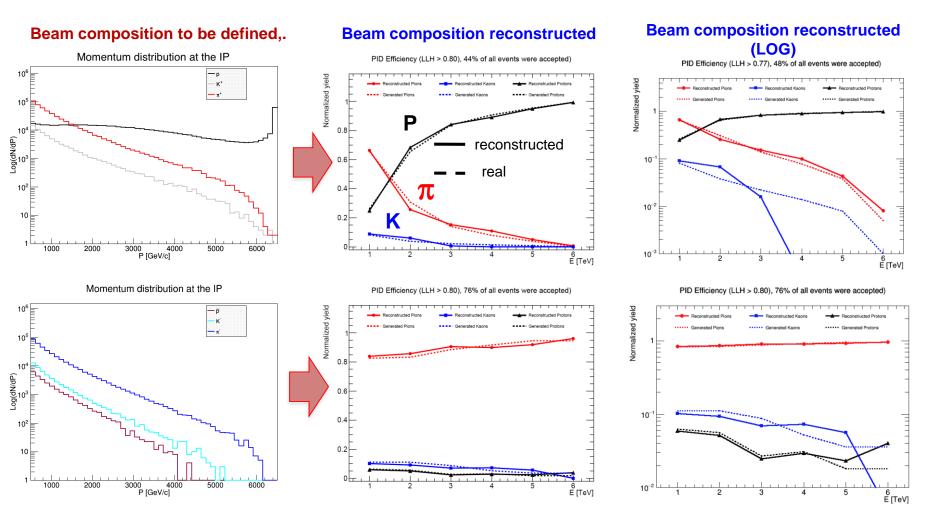
Method NOS



Method NOS (Number Of Straws): probability for different particle sort is estimated from number of straws in sub-detectors with certain energy deposition. When event probability is $P_i = P_{i1} * P_{i2} * P_{i3}$, where i is one of particle sort (π , K, p).

SAS_TRD: Simulation algorithm

We cannot identify particles with 100% probability however we can try to reconstruct composition selecting some fraction of events in which likelihood for any particle > X. This particle considered to be identified for this event.



Anatoli Romaniouk, XSCRC2017: Cross sections for Cosmic Rays @ CERN, 31.03.2017

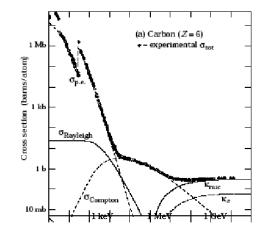
Compton Scatter TRD

TR x-rays emitted with angle $\sim 1/\gamma$

 \rightarrow difficult to spatially separate from ionization signal

Two Paths to take

- 1. Don't separate:
- Layer with thin gas (xenon) detector can detect photons well below ~30 keV
- Detection length sufficiently thin to keep ionization signal not much larger than TR signal
- Maximal efficiency requires keeping ω 's to be low and radiator foils to be thin
- 2. Separate via Compton Scattering
- Employ 250 μ m Al foils to push TR x-ray energies > ~50 keV where Compton scattering begins to dominate
- Al radiator foils can then Compton scatter TR photons, separating them from ionization deposition
- Detect scattered high-energy photons with scintillator (CsI) efficiently

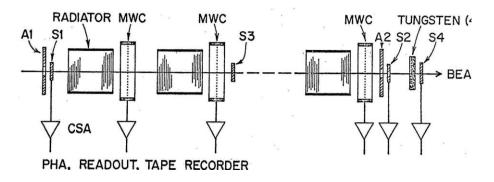


See

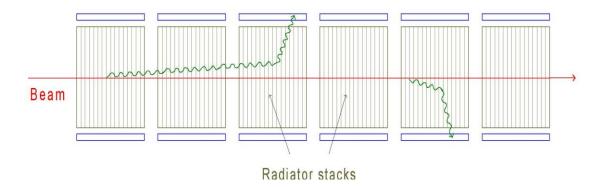
Measurement of Compton Scattered Transition Radiation at High Lorentz Factors, G. Case et al., hep-ex/0209038

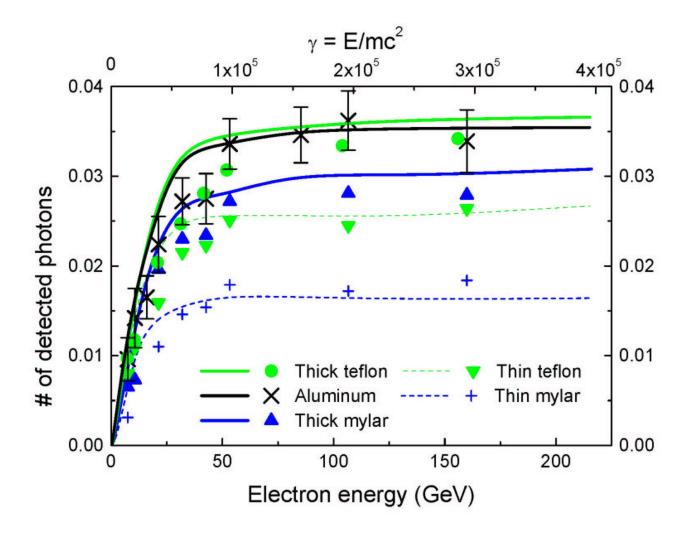
Compton Scattered Transition Radiation from Very High Energy Particles, M. Cherry & G. Case, astro-ph/02060663

"Standard" configuration – SLAC test w/plastic foils/foam, Xe



Compton scatter configuration – CERN test w/AI honeycomb, Nal





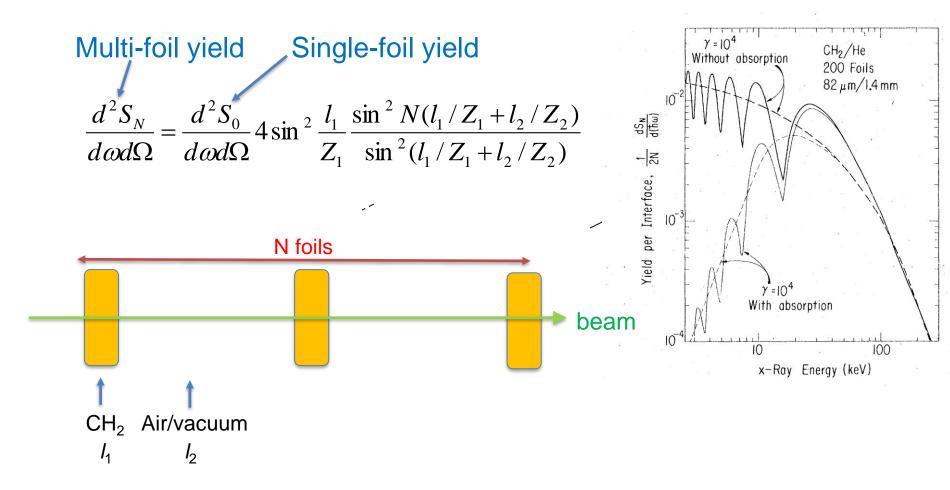
Measured 35-500 keV Compton scattered intensity

Mylar, 50 foils, $l_1 = 125$ and 250 μ m, $l_2 = 3.3$ mm Teflon, 50 foils, $l_1 = 125$ and 250 μ m, $l_2 = 3.3$ mm Al, 37 foils, $l_1 = 150 \mu$ m, $l_2 = 5.1$ mm Can one improve situation with more complicated radiators?

Composite radiators

Radiators with varying foil thickness, gap lengths?

If analytic solution does not exist, one must add field amplitudes in phase.

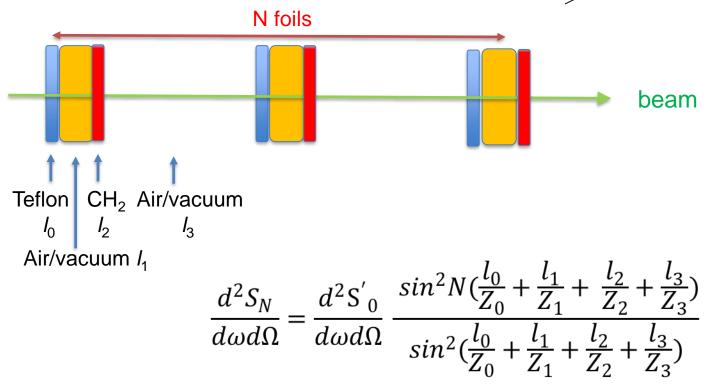


Here Z_i is the formation zone for medium *i*:

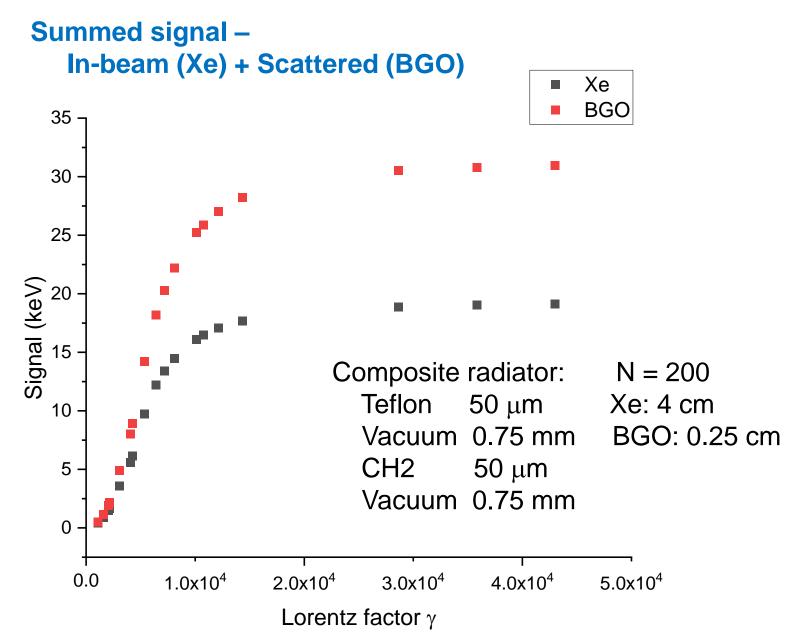
$$Z_i = \frac{4c}{\omega} / (\frac{1}{\gamma^2} + \frac{{\omega_i}^2}{\omega^2} + \theta^2)$$

Multi-foil yield Single-foil yield $\frac{d^2 S_N}{d\omega d\Omega} = \frac{d^2 S_0}{d\omega d\Omega} 4 \sin^2 \frac{l_1}{Z_1} \frac{\sin^2 N(l_1/Z_1 + l_2/Z_2)}{\sin^2(l_1/Z_1 + l_2/Z_2)}$

Example: Multi-layer foil

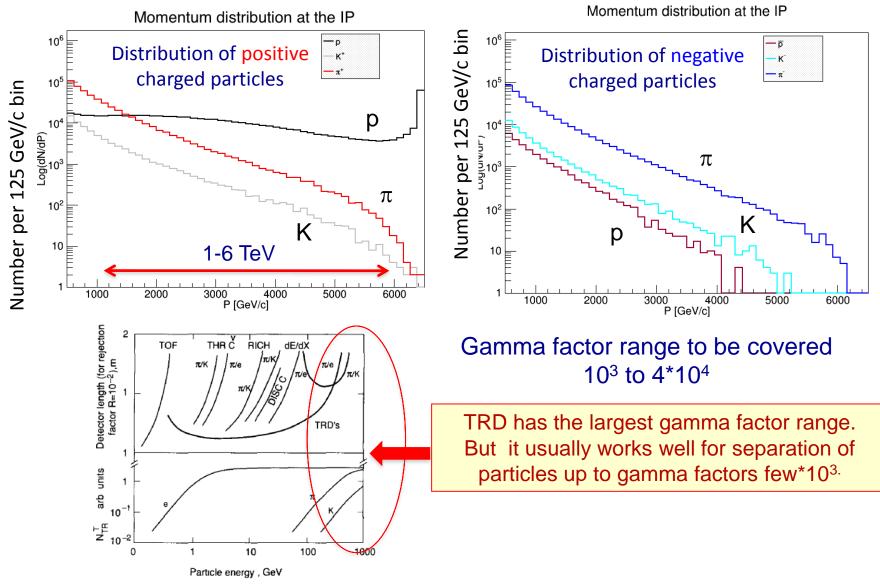


Multi-layer foil has similar modulation properties, potentially increases X-ray yield beyond single-layer foil.



What TRD is for in this experiment?

It is assumed that electrons and muons are identified using calorimeters and muon systems.



Anatoli Romaniouk, XSCRC2017: Cross sections for Cosmic Rays @ CERN, 31.03.2017

We have demonstrated:

- Ability to simulate details of spectra accurately
- Performance of multiple detectors (proportional chambers, straw tubes, Si/GaAs pixel detectors)
- Ability to tune radiator/detector specs to specific cases and reconstruct particle composition:

Once amount of space, rates, etc. are specified, an appropriate p-K-pi id system can be designed.