# The SENSEI Experiment: An Ultrasensitive Search for Sub-GeV Dark Matter

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IFIBA - CONICET/FNAL
for the SENSEI\* Collaboration
@ ICHEP 2020

#### The SENSEI Collaboration













#### Fermilab:

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 G. Fernandez-Moroni, D. Rodrigues, M. Sofo-Haro, L. Stefanazzi, J. Tiffenberg

#### **Stony Brook:**

• L. Chaplinsky, Dawa, R. Essig, D. Gift, S. Munagavalasa, A. Singal

#### Tel-Aviv:

• L. Barak, I. Bloch, E. Etzion, A. Orly, S. Uemura, T. Volansky

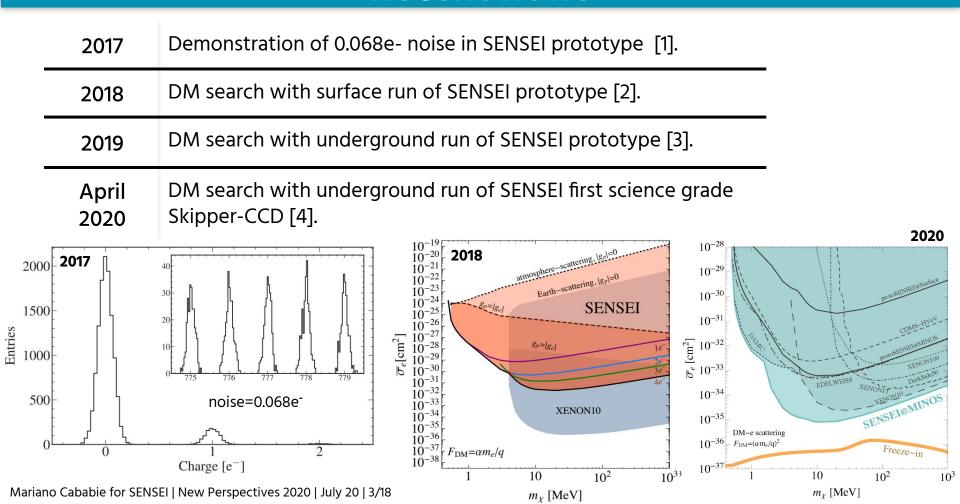
#### U. Oregon:

• T.-T. Yu

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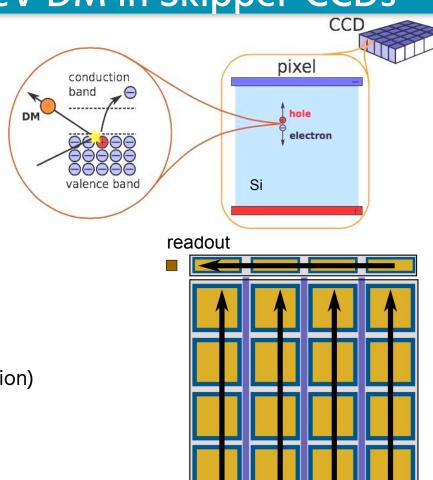


#### Recent news



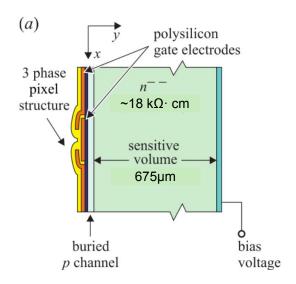
## Electron recoils for sub-GeV DM in Skipper-CCDs

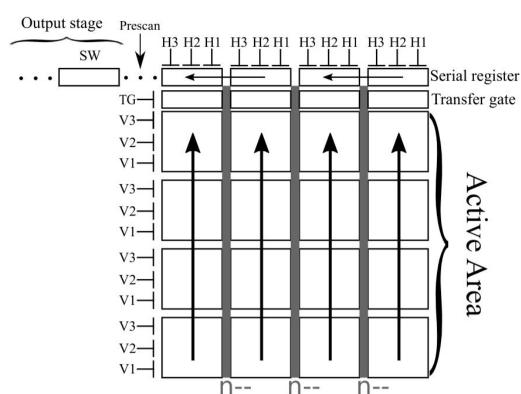
- Benchmark models:
  - ◆ DM-e<sup>-</sup> scattering, DM absorption
- Silicon CCDs as ionization detectors
  - DM-e<sup>-</sup> interaction (or absorption)
  - Energy transfer via electron recoil
  - lonized h<sup>+</sup> are captured by potential well
  - Signal is readout after exposure is finished.
- DM range mass: 1-1000 MeV (~eV on DM absorption)



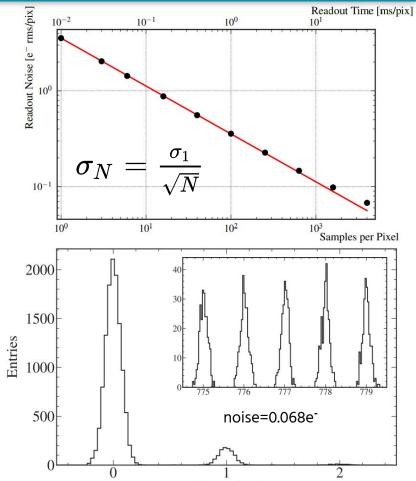
#### **CCD** basics

- CCD = pixelated silicon array
- ◆ ~2g per device of high-resistivity fully-depleted silicon
- >99.9% charge collection and transfer efficiency
- \* ~**5.5Mpixels** of 15x15x675 μm³ each



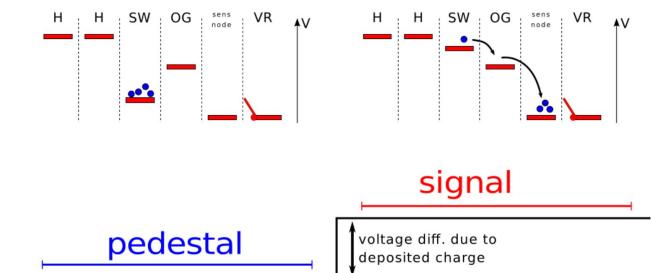


- DM range mass: 1-1000 MeV (~eV on DM absorption)
  - Very small signals
    - Very low energy threshold
- Skipper technology allows to read repeatedly the same pixel to achieve sub-electron noise
- Low energy threshold down to 1.2eV (Si band gap)



Charge [e<sup>-</sup>]

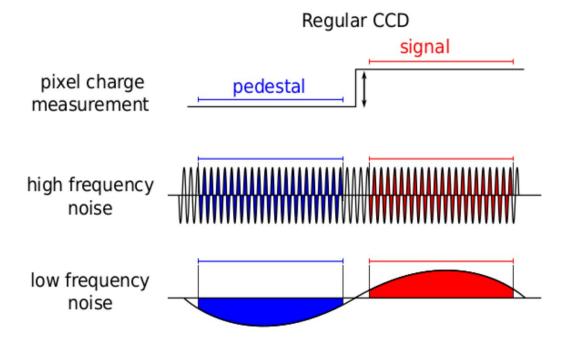
- In a conventional CCD, charge is moved to the sense node and readout once.
   Then it is drained and charge is lost.
- Longer integration reduces noise but cannot reduce 1/f noise.
- Skipper-CCD moves charges towards and backwards the floating sense node to achieve multiple readout



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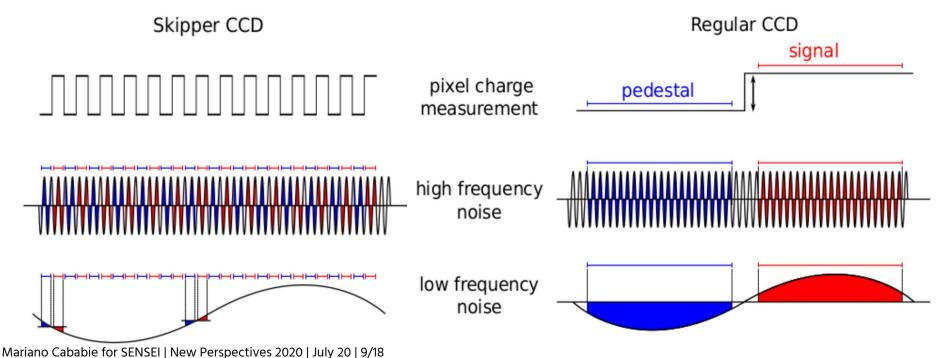
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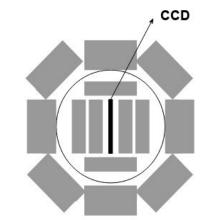


## Our setup: location and shielding

- Setup ~107m below surface at shallow underground MINOS site @FNAL.
- This reduces muon environmental background radiation

Inner (1" each) and outer (2" each) lead bricks
 reduces environmental gamma radiation

 Operated at 135K and high-vacuum regime to reduce dark current without generating CTI





NuMI building

107 m

MINOS Hall

## **Electronics | Inside the Vessel**



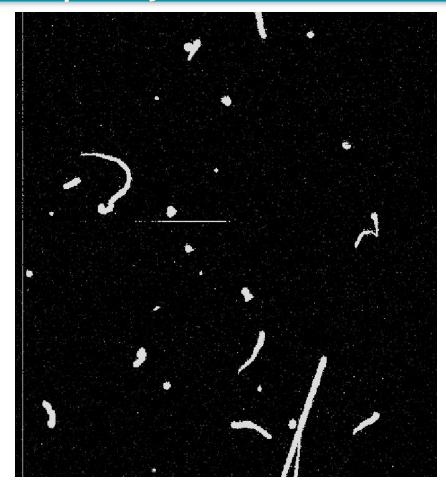
- Shielding design adapted from DAMIC:
  cylindrical vacuum vessel with lead
  "plugs" above and below the CCD
- Operated with specifically designed readout electronics (LTA board)
   (Low Threshold Acquisition)

 LTA boards admit multiple reading of multiple CCDs synchronously which enables scaling



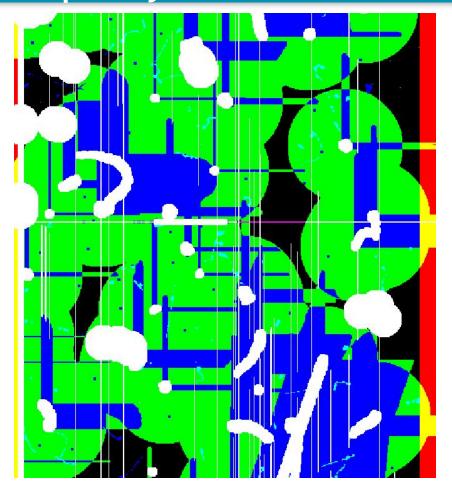
#### Our last result: quality cuts

- Bad pixels/dark spikes
- Serial register hits
- Bleeding (CTI)
- Halo (low energy events near high energy events) (new!)
- Loose clusters (≥2e- analysis)
- Others



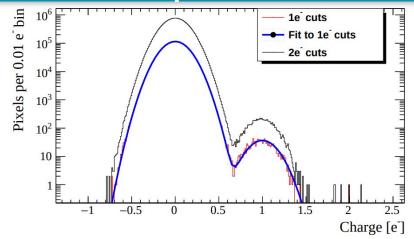
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#### Our last result: data and some specifics

- ♦ Blinded dataset of 22 images, Feb 25 Mar 20
   →20hs exposure + 6hs readout (each)
  - →Total exposure: **19.926 gram-days**
- 300 skipper samples → 0.14e<sup>-</sup> readout noise
- \* x20 more mass than 2019 (x10-15 effectively)



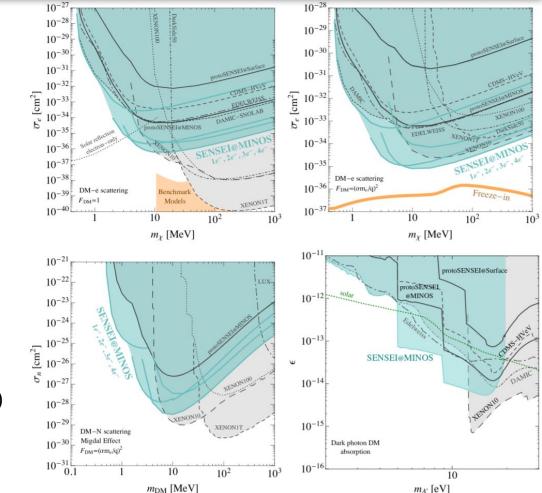
- Background as low as ~3400 events/kg/keV/day (~3 times less than 2019)
- DC rate Single-electron event rate as low as ~1.6x10<sup>-4</sup> e<sup>-</sup>/pixel/day or ~450 events/gram/day (~20 times less than 2019)
- \* x(8-35) more effective exposure (depending on  $e^{-1}$  channel)

#### Our last result: limits on DM

World-leading constraints on DM-e<sup>-</sup> scattering for light mediator (top right) and heavy mediator (top left), up to 10 MeV.

World-leading constraints for
 DM-nucleus scattering (bottom left)
 through light mediator from 600 keV to
 MeV (Migdal Effect [5]).

World-leading constraints for **DM absorption on electrons** (bottom right)
 from 1.2 to 12.8 eV.



# Perspectives

2017	Demonstration of 0.068e- noise in SENSEI prototype [1].
2018	DM search with surface run of SENSEI prototype [2].
2019	DM search with underground run of SENSEI prototype [3].
April 2020	DM search with underground run of SENSEI first science grade Skipper-CCD [4].
2021	???

Perspectives

• We have our science detectors and they work! Next step: production (in progress).

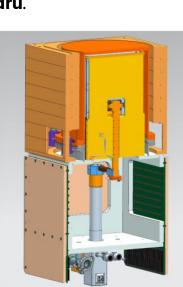
• We are assembling our vessel that will go to SNOLAB.

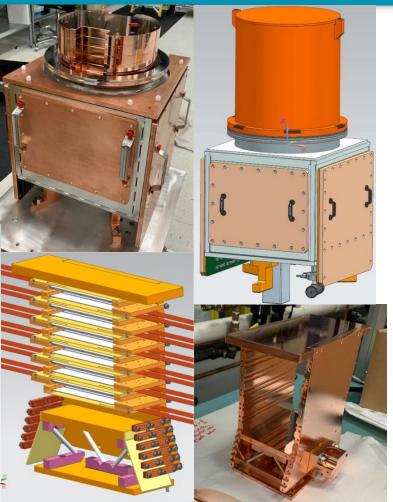
o MINOS (standard shield): 10000 dru

MINOS (extra shield): 3000 dru

SNOLAB (final setup): 5 dru.

 Vessel is at Fermilab, ready for testing prior to travel.





## Perspectives

- \* "Phase 1" system **fully operational** since December @SNOLAB
- ♦ Final mass: 100g (~2g now).
- Deployment in stages, increasing mass.
   Results will be presented gradually.
- SENSEI should be deployed by 2020 end of 2020 beginning of 2021.



Sho Uemura, Kevin Kuk and Guillermo Moroni @SNOLAB December 2019



~2027

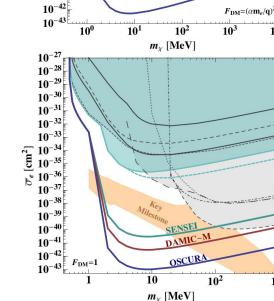
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#### The SENSEI Collaboration



# **THANK YOU!**

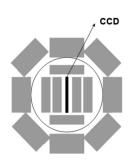
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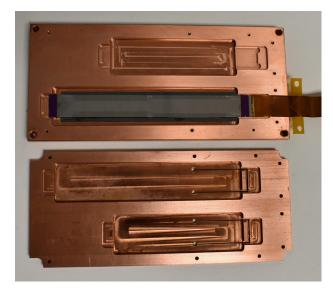


# **BACK UP SLIDES**

## MINOS shielding









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#### References

- [1] Tiffenberg, Javier, et al. "Single-electron and single-photon sensitivity with a silicon Skipper CCD." Physical Review Letters 119.13 (2017): 131802.
- [2] Crisler, Michael, et al. "SENSEI: first direct-detection constraints on sub-GeV dark matter from a surface run." Physical Review Letters 121.6 (2018): 061803.
- [3] Abramoff, Orr, et al. "SENSEI: Direct-detection constraints on sub-GeV dark matter from a shallow underground run using a prototype skipper CCD." Physical review letters 122.16 (2019): 161801.
- [4] Barak, Liron, et al. "SENSEI: Direct-Detection Results on sub-GeV Dark Matter from a New Skipper-CCD." arXiv
- [5] Essig, Rouven, et al. "Relation between the Migdal Effect and Dark Matter-Electron Scattering in Isolated Atoms and Semiconductors." Physical Review Letters 124.2 (2020): 021801.

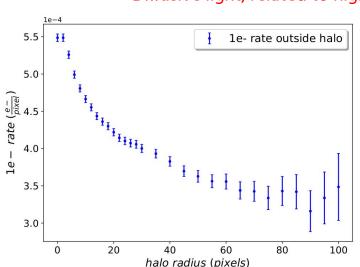
preprint arXiv:2004.11378 (2020).

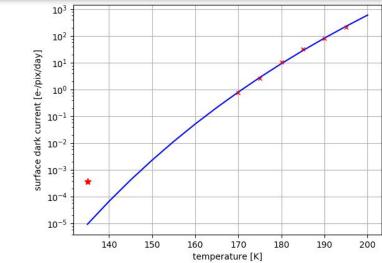
#### Our last result: single electron event rate

- \* A 1e- rate excess is found extrapolating from higher temperatures **assuming only surface DC**.
- Extrinsic or intrinsic sources?

RO stage luminescence, other DC

Diffusive light, related to high energy events

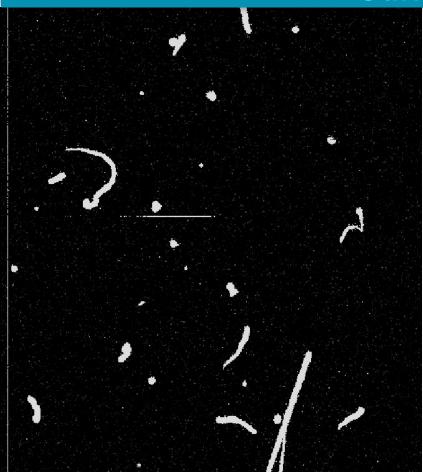


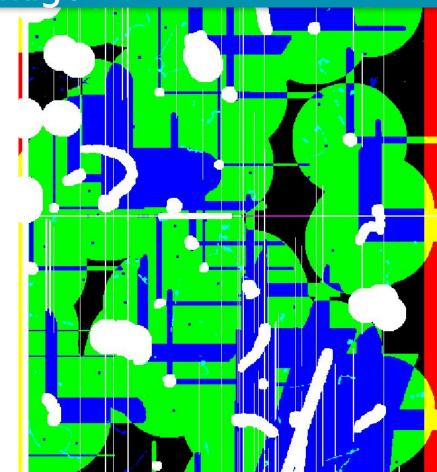


- Spatial correlation between high energy events (>360eV) and 1e- events.
- Low-energy photons? From copper module, CCD or both?
- Can we mask it up to 100%?

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Sample image





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#### 1e- rate vs. shielding

- We have data with and without the outer ring of lead bricks
- Factor of 3 reduction in the rate of high-energy tracks → factor of 3 reduction in the 1e<sup>-</sup> rate
  - There is some mechanism by which ionizing radiation generates charge uniformly in our CCD
  - Better shielding will very likely further reduce our 1e<sup>-</sup> rate

