Latest Reactor Neutrino Oscillation Results from the Daya Bay Experiment



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*on behalf of the Daya Bay Collaboration

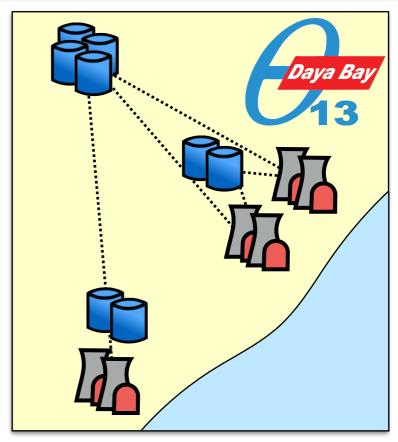










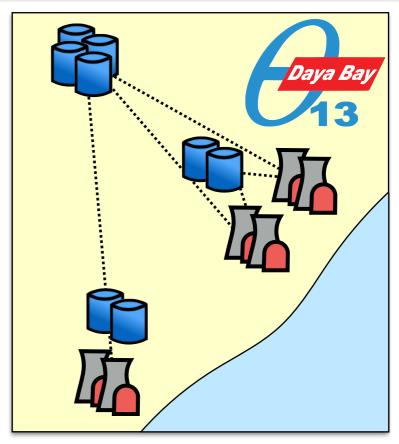


Primarily designed to precisely measure the θ_{13} mixing angle Using reactor antineutrino oscillation at the ~2 km baseline Discovered non-zero value of θ_{13} in 2012





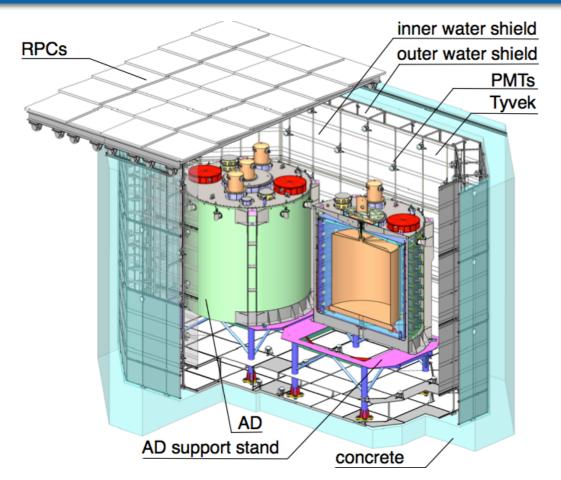


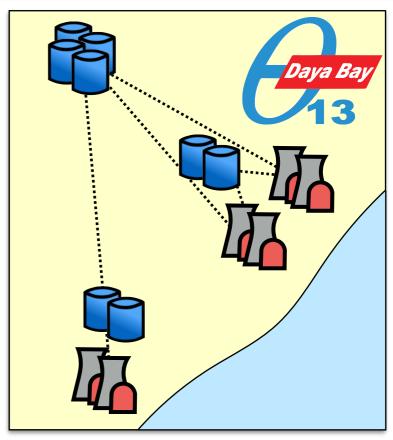


- 6 reactors with 2.9 GW_{th} each → 17.4 GW_{th} total
- One of the most powerful reactor complexes in the world
- Nuclear reactors strong source of pure electron antineutrinos
- $^{\rm L}$ Each Daya Bay reactor emits ~6 \times $10^{20}\,\bar{\nu}_e$ /s isotropically





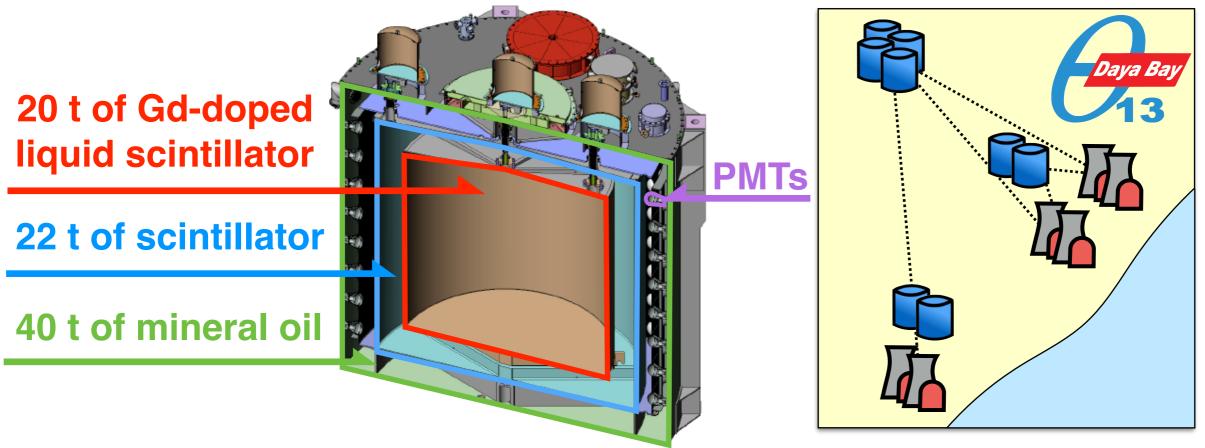




- 8 functionally identical antineutrino detectors (ADs)
- Placed in 3 underground experimental halls at the optimal distance from the reactors
- Submerged in the instrumented ultra pure water pools







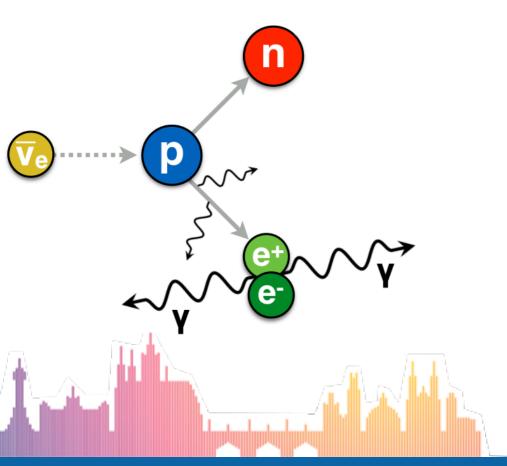
- 8 functionally identical antineutrino detectors (ADs)
- Placed in 3 underground experimental halls at the optimal distance from the reactors
- Submerged in the instrumented ultra pure water pools
- Each AD consists of 3 nested volumes: Gd-doped scintillator

(main target), pure scintillator (γ-catcher), mineral oil (buffer)





- ••• Primary interaction inverse beta decay (IBD): $\bar{\nu}_e + p \rightarrow e^+ + n$
- Prompt signal energy related to incident antineutrino energy $E_{\bar{\nu}} \simeq E_p + 0.78~{\rm MeV}$
- Spatial and time correlation of prompt and delayed signals greatly suppresses the background (background/signal <2% in Daya Bay)

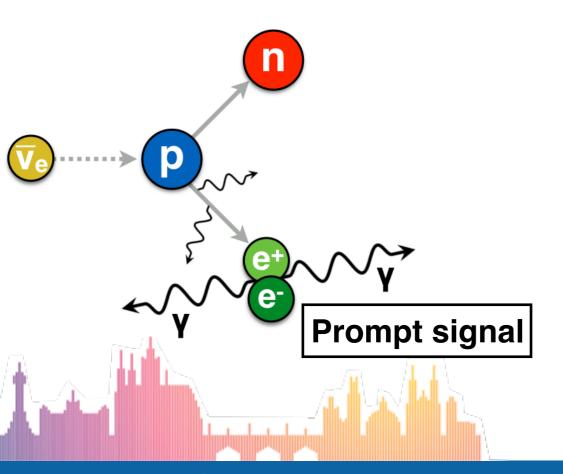






- Prompt signal

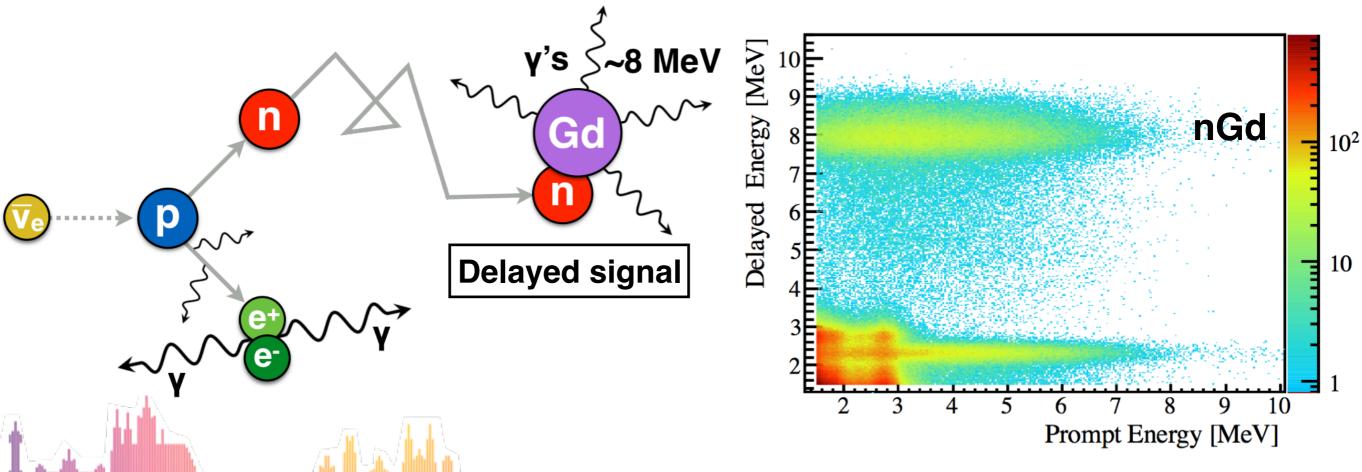
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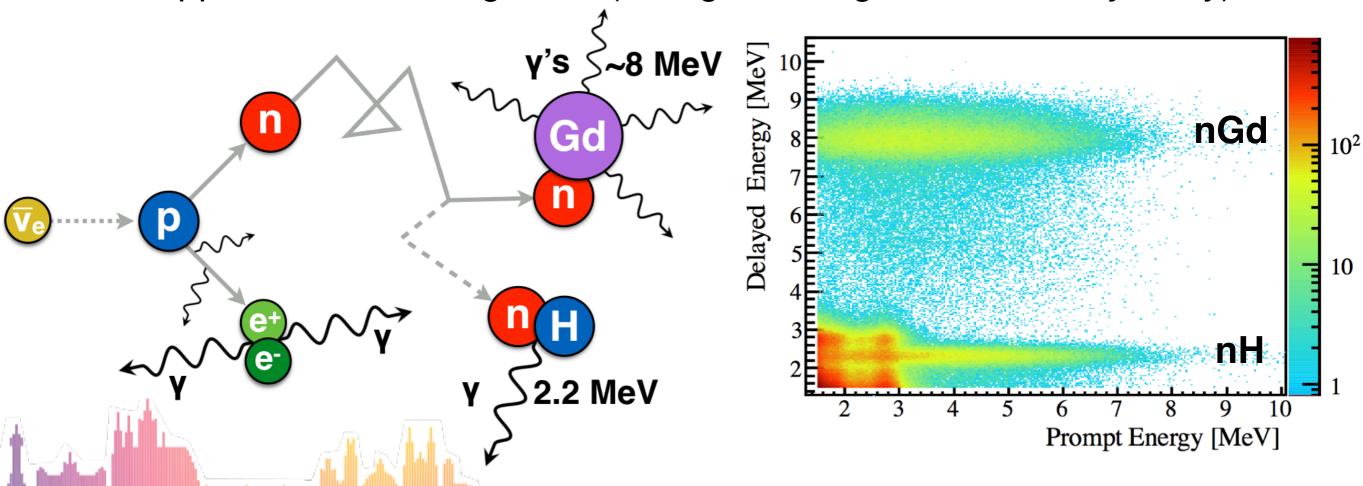
- Prompt signal Delayed signal
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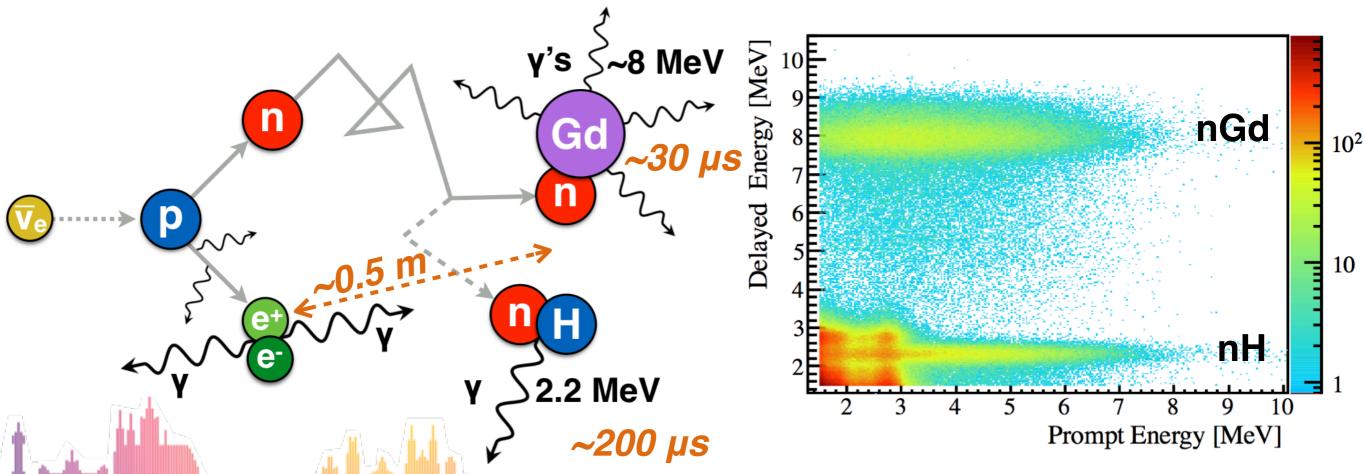
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Neutrino Mixing and Oscillation



Three-neutrino mixing:

Atmospheric, accelerator v

Solar, reactor L~60 km v

Flavor
$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_ au \end{pmatrix} = \begin{pmatrix} 1 \\ 0 \\ 0 \end{pmatrix}$$

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix}$$

Flavor states
$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \begin{pmatrix} c_{13} & 0 & s_{13}e^{-i\delta} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta} & 0 & c_{13} \end{pmatrix} \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

$$\begin{pmatrix}
c_{12} & s_{12} & 0 \\
-s_{12} & c_{12} & 0 \\
0 & 0 & 1
\end{pmatrix}$$

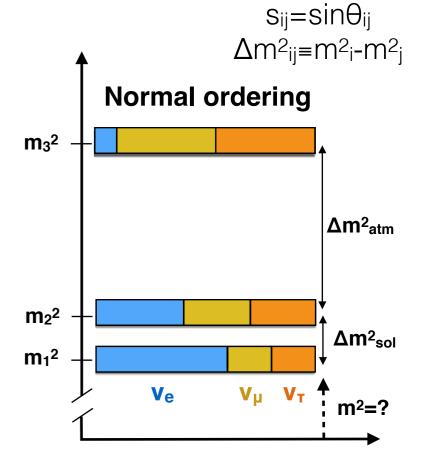
$$egin{pmatrix}
u_1 \\
\nu_2 \\
\nu_3 \end{pmatrix}$$
 Mass states

 $c_{ij} = cos\theta_{ij}$

Reactor L~2 km, accelerator v

Oscillation parameters:

Parameter	Value	Open questions
Δm_{21}^2	7.5×10 ⁻⁵ eV ²	_
$ \Delta m_{31} ^2 \simeq \Delta m_{32} ^2$	2.5×10 ⁻³ eV ²	Ordering?⇔Δm ₃₁ ² ≤0
θ ₁₂	33°	_
θ_{23}	45°?	Maximal?⇔θ ₂₃ ≷45°
θ ₁₃ *	9°	
δсР	? °	Value?



*Last to known angles - discovered by Daya Bay in 2012



Oscillation Measurement Principle



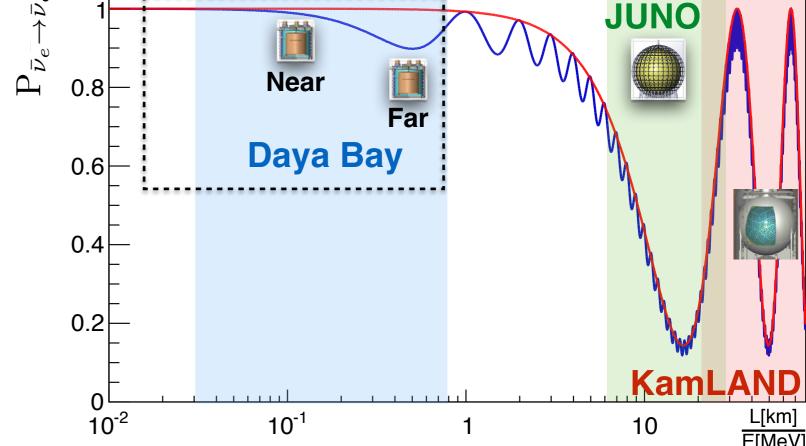
WW Looking at the flux of $\bar{\nu}_{\rho}$ as a function of distance and energy

WW Some $\bar{\nu}_{\rho}$'s disappear due to neutrino oscillation

Medium baseline

Short baseline

$$P_{\bar{\nu}_e \to \bar{\nu}_e}(L, E) = 1 - \frac{\sin^2 2\theta_{12} \cos^4 \theta_{13} \sin^2 \frac{\Delta m_{21}^2 L}{4E}}{4E} - \left[\sin^2 2\theta_{13} \left(\cos^2 \theta_{12} \sin^2 \frac{\Delta m_{31}^2 L}{4E} + \sin^2 \theta_{12} \sin^2 \frac{\Delta m_{32}^2 L}{4E}\right)\right]$$





Oscillation Measurement Principle



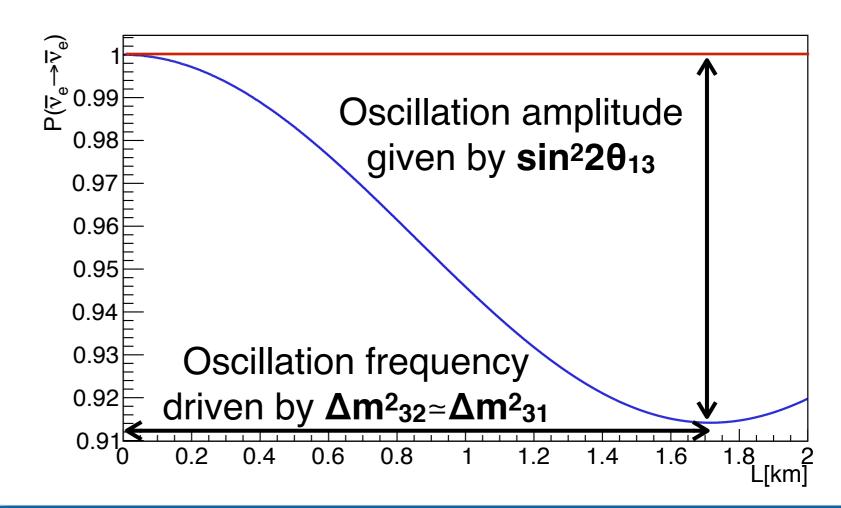
WW Looking at the number of $\bar{\nu}_{\rho}$ as a function of distance and energy

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Medium baseline

Short baseline

$$P_{\bar{\nu}_e \to \bar{\nu}_e}(L, E) = 1 - \left[\sin^2 2\theta_{12} \cos^4 \theta_{13} \sin^2 \frac{\Delta m_{21}^2 L}{4E} \right] - \left[\sin^2 2\theta_{13} \left(\cos^2 \theta_{12} \sin^2 \frac{\Delta m_{31}^2 L}{4E} + \sin^2 \theta_{12} \sin^2 \frac{\Delta m_{32}^2 L}{4E} \right) \right]$$



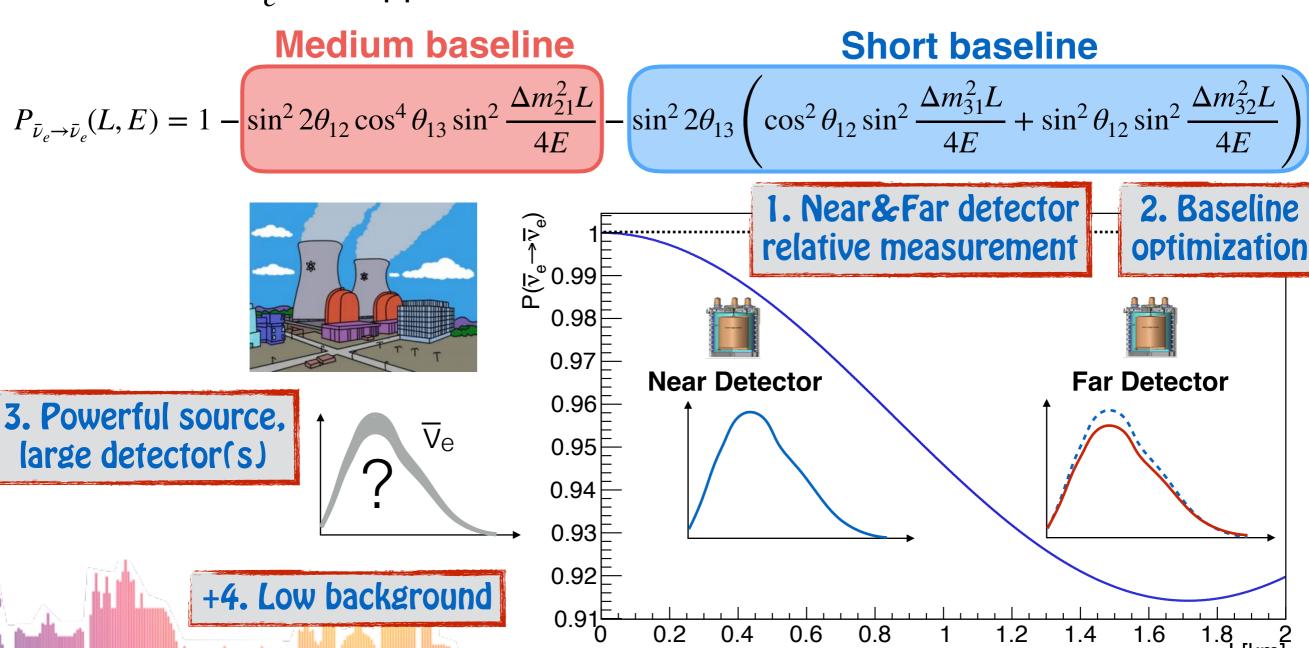


Oscillation Measurement Principle



 $\overline{\mathbf{W}}$ Looking at the number of $\bar{\nu}_e$ as a function of distance and energy

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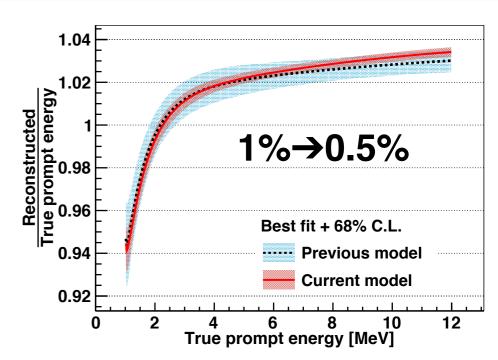


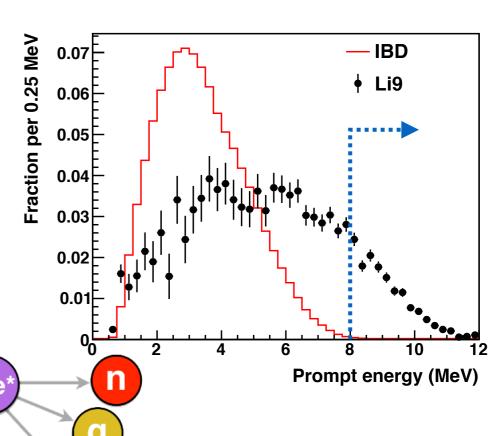


Systematics Improvements



- Absolute energy scale uncertainty < 0.5%</p>
 - Special calibration campaign with various radioactive source encapsulations
 - FADC readout in one AD to determine electronics non-linearity
- Reduction of the uncertainty of the 9Li/8He background
 - Dominant background uncertainty
 - Prompt energy cut to enhance the ⁹Li/IBD - feasible due to large statistics
- Spent nuclear fuel uncertainty 100%→30%
 - Utilize precise spent nuclear fuel history provided by power plant



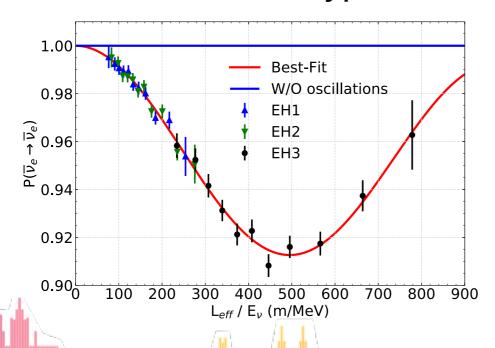


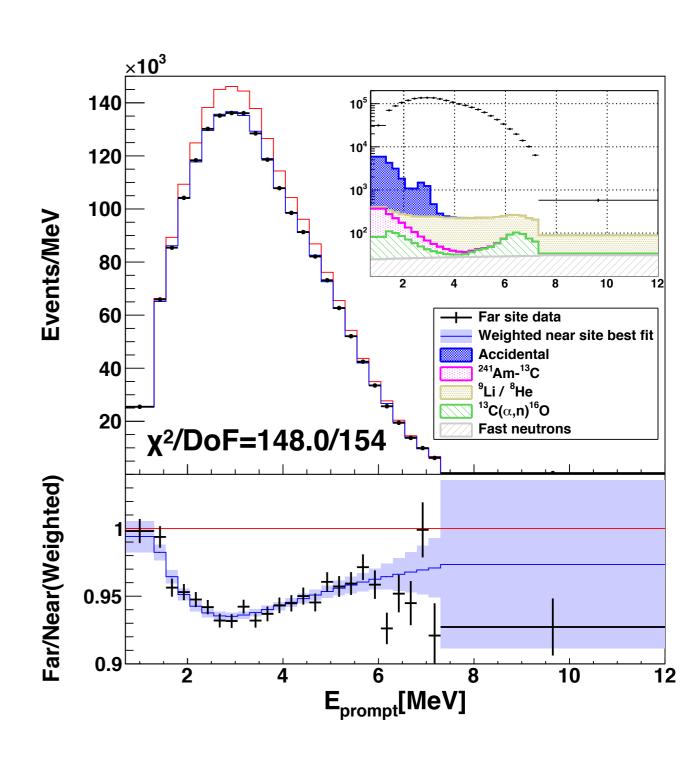


Latest Measurement with nGd



- X Based on 1958 days data set
- More than 3.9 million antineutrino interactions (with 0.5 million in far site)
 - The largest rector antineutrino data set in the world
- \bowtie Consistent with 3- ν hypothesis





PRL 121, 241805 (2018)



Latest Oscillation Result with nGd



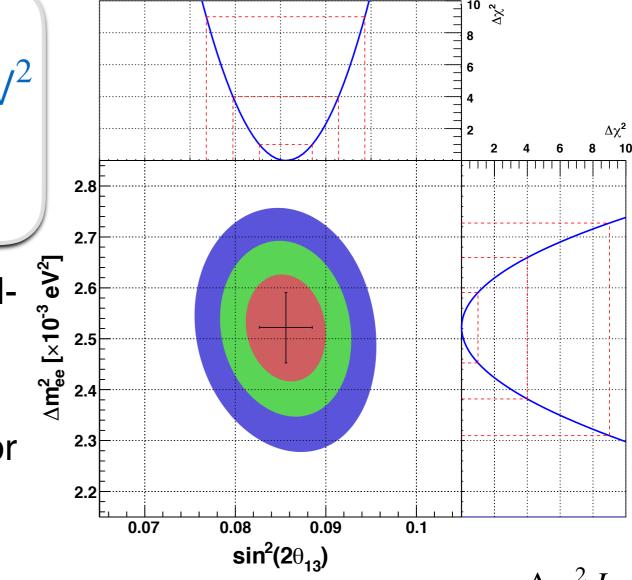
$$\sin^2 2\theta_{13} = 0.0856 \pm 0.0029$$

$$|\Delta m_{ee}^2| = (2.522^{+0.068}_{-0.070}) \times 10^{-3} \text{ eV}^2$$

$$\Delta m_{32}^2 = (2.472^{+0.068}_{-0.070}) \times 10^{-3} \text{ eV}^2 \text{ (Normal ordering)}$$

$$\Delta m_{32}^2 = -(2.575^{+0.068}_{-0.070}) \times 10^{-3} \text{ eV}^2 \text{ (Inverted ordering)}$$

- Measurement of sin²2θ₁₃ with world-leading 3.4% precision
- Measurement of IΔm²eel with 2.8% precision, comparable to accelerator experiments
- Statistics still contributes by ~60% (~50%) to the sin²2θ₁₃ (IΔm²_{ee}I) uncertainty



$$P_{\bar{\nu}_e \to \bar{\nu}_e}(L, E) \simeq 1 - \sin^2 2\theta_{12} \cos^4 \theta_{13} \sin^2 \frac{\Delta m_{21}^2 L}{4E}$$

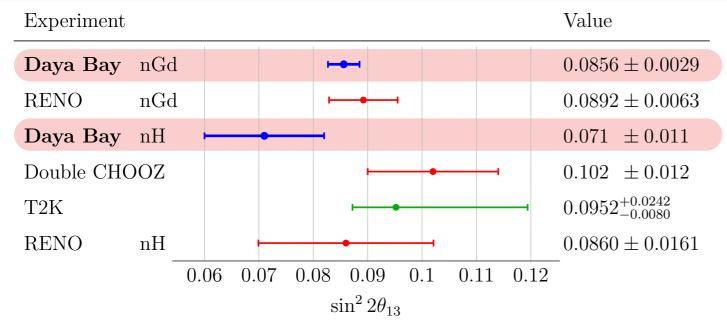
Effective mass $-\sin^2 2\theta_{13} \sin^2 \frac{\Delta m_{ee}^2 L}{4E}$ splitting *PRL 121, 24*

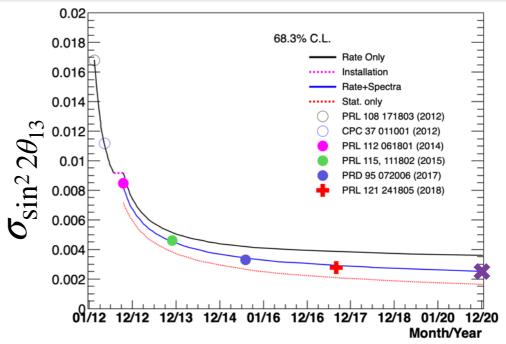
PRL 121, 241805 (2018)



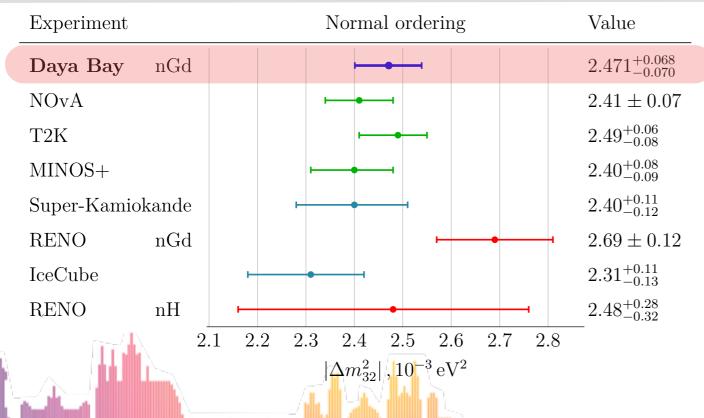
Global Comparison and Prospects

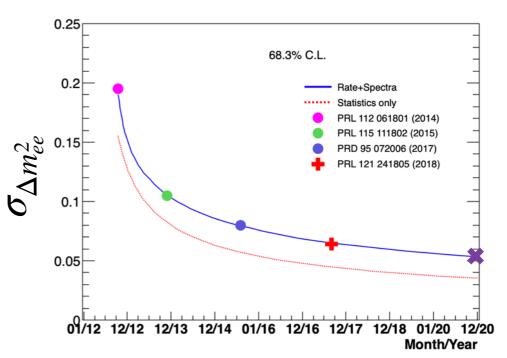






 $\sin^2 2\theta_{13}$ uncertainty 3.4% \rightarrow 2.7%





 Δm^2_{ee} uncertainty 2.8% \rightarrow 2.1%

End of Daya Bay data taking-

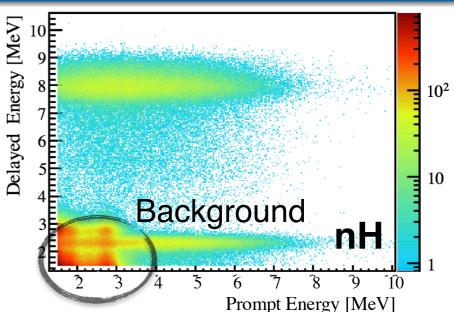


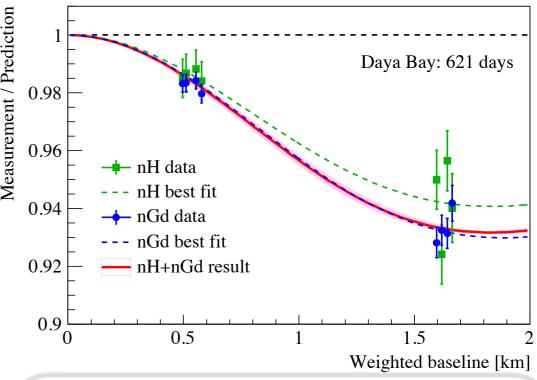


'Latest' Oscillation Result with nH



- nH essentially independent analysis from nGd
 - Different statistical sample
 - Mostly decoupled systematics
- Challenges
 - Large background (bg/signal up to 1) X
 - Larger systematic uncertainties X
- Latest' rate-only analysis with only 621 days data set
- Update under intense preparation
 - Expected to be one of the most precise measurements of θ_{13} in the world





 $\sin^2 2\theta_{13} = 0.071 \pm 0.011$ Assuming $\Delta m_{32}^2 = (2.44 \pm 0.06) \times 10^{-3} \text{ eV}^2 \text{ (NO)}$

PRD 93, 072011 (2016)

See poster by Jinjing Li



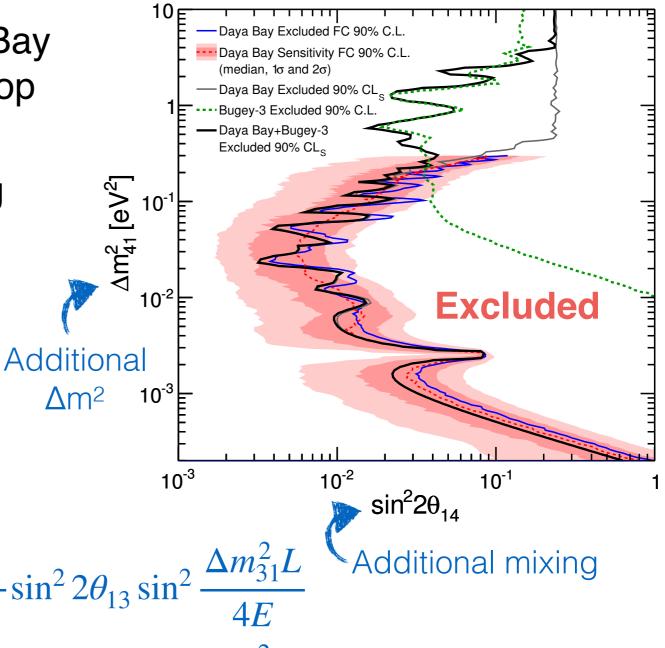


Search for Light Sterile Neutrino



arXiv:2002.00301, accepted by PRL

- Sterile' neutrino signature in Daya Bay
 an additional oscillation mode on top of the known one
- No significant signal observed using 1230-day data set
- Le Placed the most stringent limits on $sin^2 2\theta_{14}$ for $\Delta m^2_{41} < 0.2$ eV²
- Larger Δm²₄₁ by combination with Bugey-3





$$P_{\bar{\nu}_e \to \bar{\nu}_e}(L, E) \approx 1 - \sin^2 2\theta_{13} \sin^2 \frac{\Delta m_{31}^2}{4E}$$
 $-\sin^2 2\theta_{14} \sin^2 \frac{\Delta m_{41}^2}{4E}$

 $\frac{h_{41}E}{4E}$

Additional oscillation mode

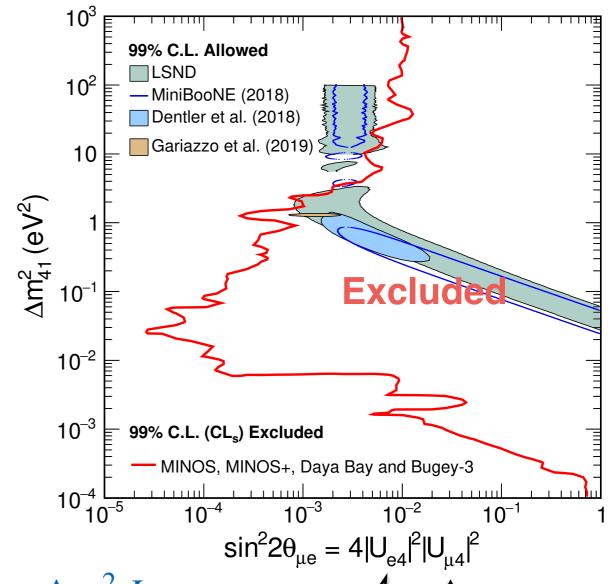


Search for Light Sterile Neutrino



arXiv:2002.00301, accepted by PRL

- $ightharpoonup^{}_{s}$ Combined ν_{s} search with Bugey-3 and MINOS,MINOS+
- LESSIP Excluded LSND and MiniBooNE 99% C.L. allowed regions at 99% CL_s for Δm²₄₁<1.6 eV²
- ▲ Global ν_s (appearance) fits ruled out to >99% CL_s



See poster by **432 Zhuojun Hu**

 $P_{\overline{
u}_{\mu} o \overline{
u}_{e}}^{ ext{short baseline}} \simeq \sin^{2} 2\theta_{\mu e} \sin^{2} \frac{\Delta m_{41}^{2} L}{4E}$

 $\sin^2 2\theta_{\mu e} \equiv \sin^2 2\theta_{14} \sin^2 \theta_{24}$

Daya Bay, Bugey-3 $\bar{\nu}_e
ightarrow \bar{\nu}_e$

 $\begin{array}{c} \textbf{MINOS,} \\ \textbf{MINOS+} \\ \ddot{\nu}_{\mu} \rightarrow \ddot{\nu}_{\mu} \end{array}$



Conclusions



The latest Daya Bay reactor antineutrino oscillation measurement using nGd data sample with >5 years of data yielded:

$$\sin^2 2\theta_{13} = 0.0856 \pm 0.0029$$

 $|\Delta m_{ee}^2| = (2.522^{+0.068}_{-0.070}) \times 10^{-3} \text{ eV}^2$
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- Independent analysis using neutron capture on hydrogen with \sim 2 years of data provides consistent value of θ_{13}
 - New result under intense preparation
- Search for light sterile neutrino has not found any
 - Daya Bay, Bugey-3 and MINOS+ placed strong limits on u_{s} mixing
- Data taking will finish in Dec 2020 final Daya Bay result on θ_{13} will be the standard for the foreseeable future



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Extras

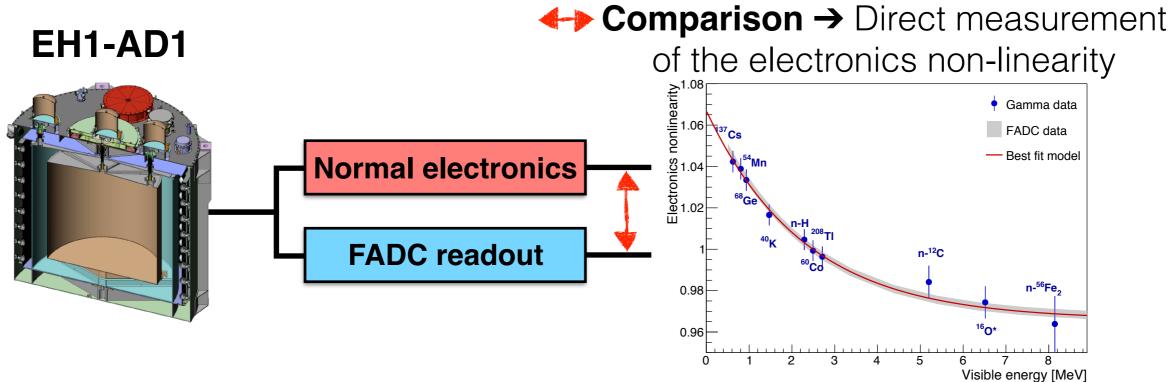




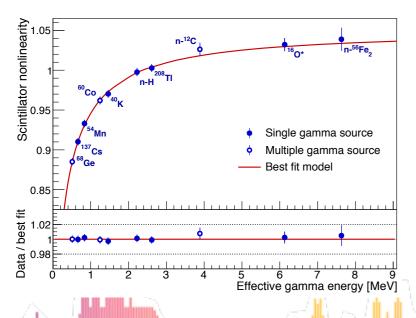


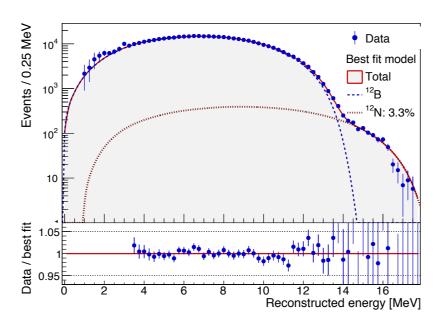
Absolute Energy Scale



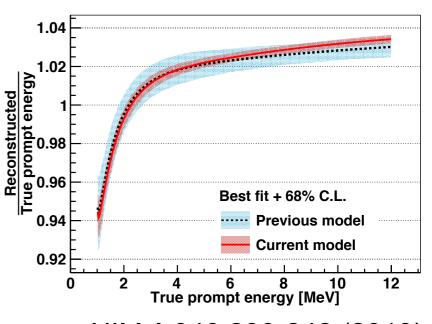








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NIM A 940 230-242 (2019)

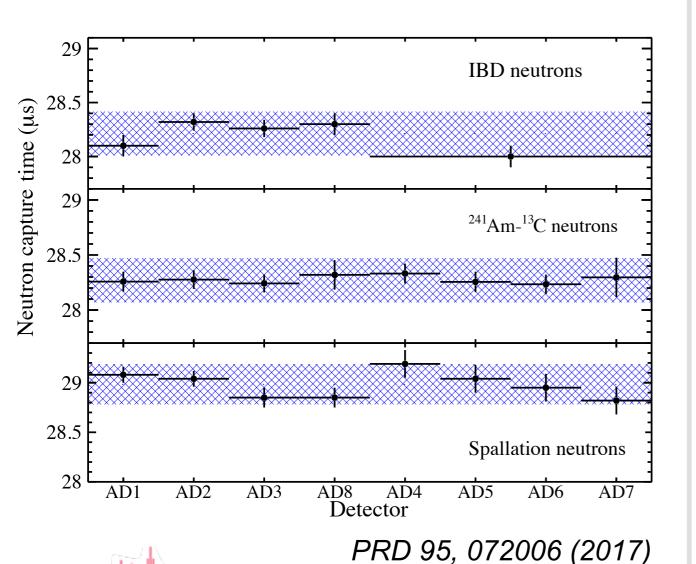


Dominant Relative Uncertainties



sin²2θ₁₃ uncertainty dominated by the 0.13% relative detection efficiency uncertainty

(which is dominated by Gd capture fraction of 0.1%)



Δm²_{ee} uncertainty dominated by the <0.2% relative energy scale uncertainty

Neutron from IBD

- Neutron from muon spallation
- Alpha from natural radioactivity
- Gamma from natural radioactivity
- Neutron from IBD (alternate method)

