

Dark Matter Theory

Eric Kuflik

ICHEP 2020



האוניברסיטה העברית בירושלים
THE HEBREW UNIVERSITY OF JERUSALEM

New Ideas in Dark Matter Theory

Eric Kuflik

ICHEP 2020



האוניברסיטה העברית בירושלים
THE HEBREW UNIVERSITY OF JERUSALEM

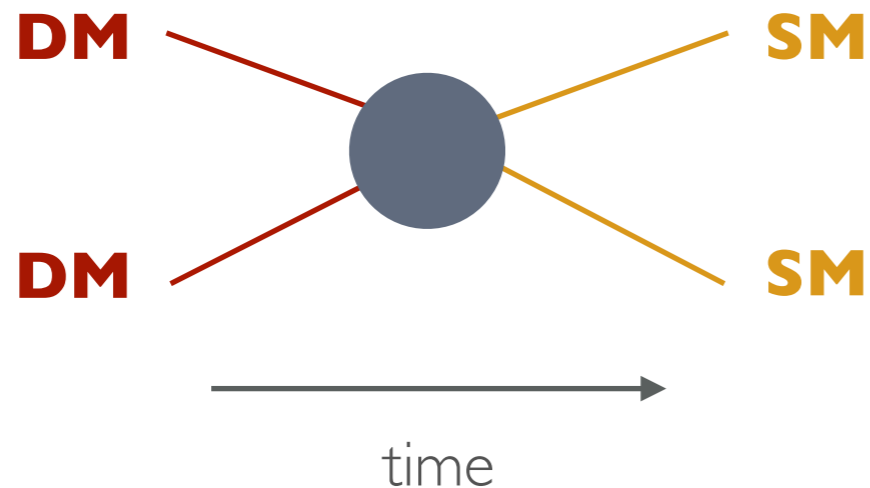
Why?

Past 40 years

WIMP, glorious WIMP*

***Also axions**

WIMP



$$\langle \sigma_{\text{ann}} v \rangle = \frac{\alpha^2}{m_{\text{DM}}^2}$$

Correct relic abundance for

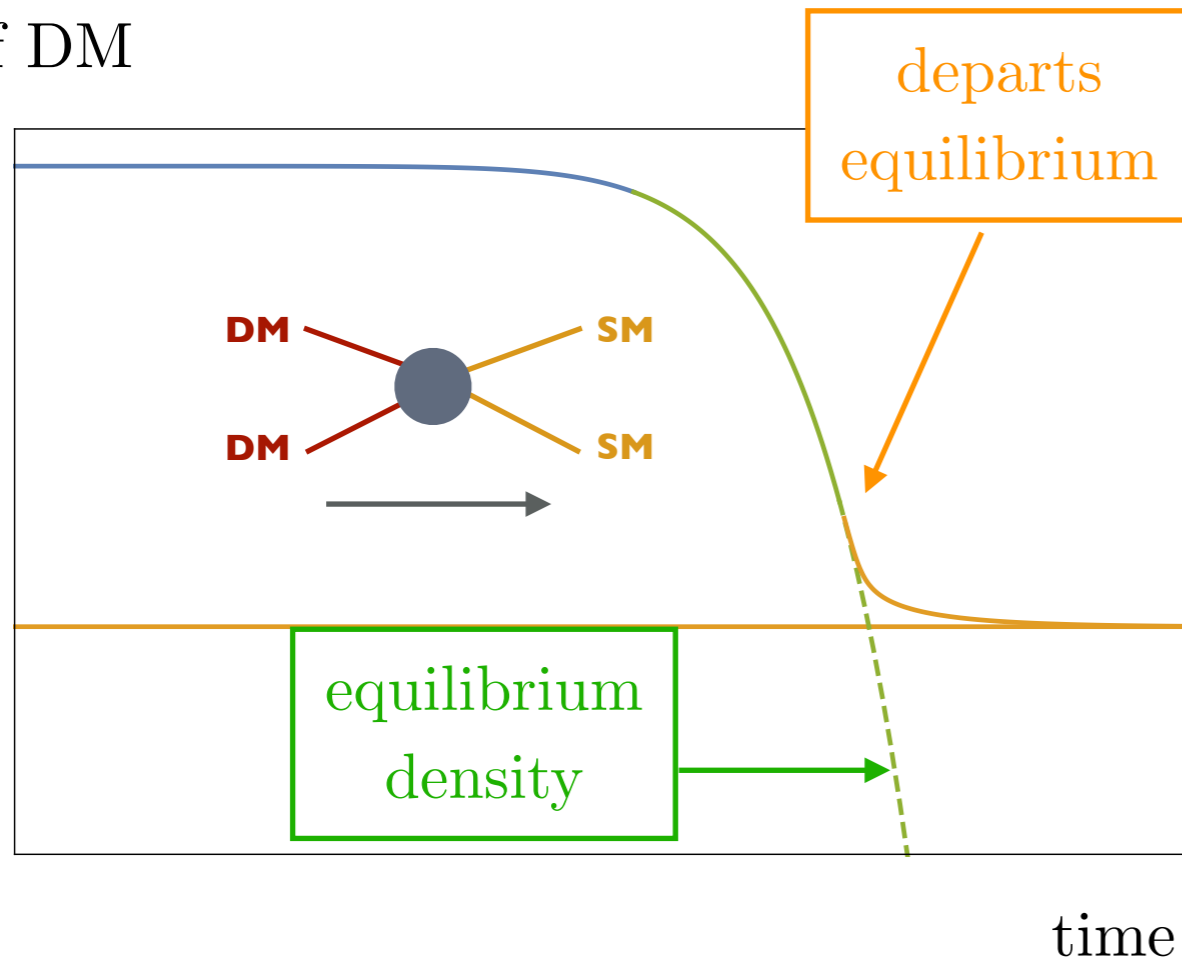
$$m_{\text{DM}} = \alpha \times 30 \text{ TeV}$$

For Weak coupling, Weak scale emerges

Weakly Interacting Massive Particle (WIMP)

WIMP

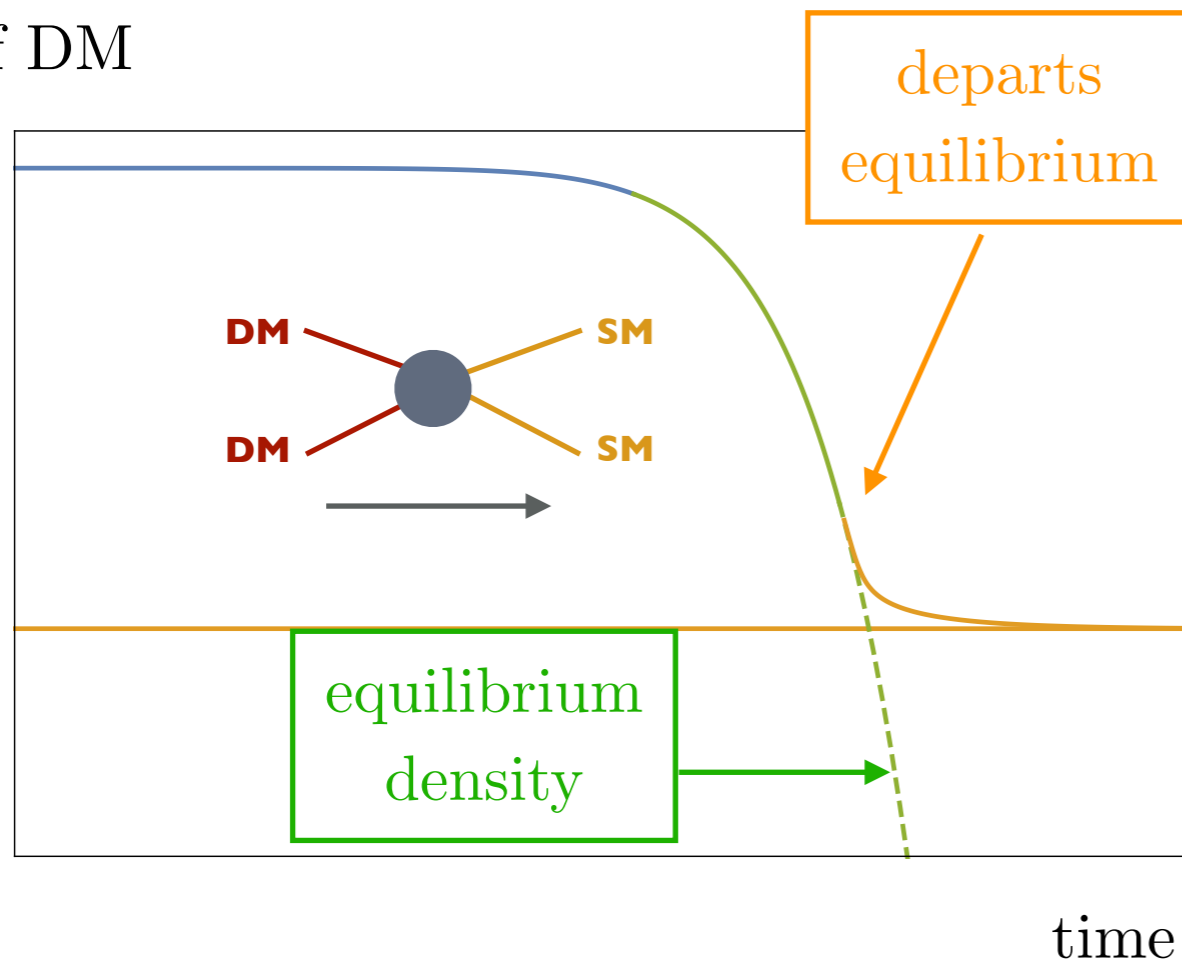
Amount
of DM



Thermal Relic:
Simple and Predictive

WIMP

Amount
of DM



Thermal Relic:
Simple and Predictive

Guiding principle in cosmology

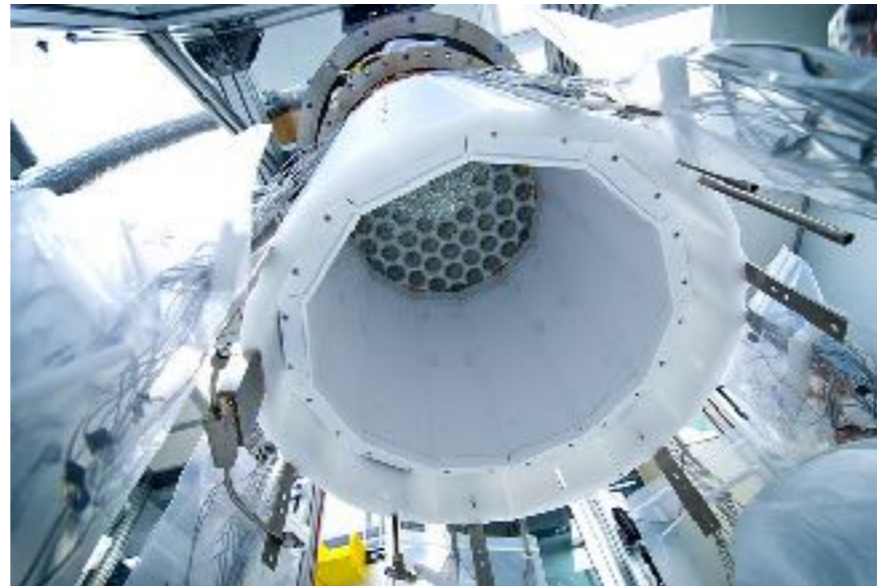
Searching for WIMPs

Direct Production



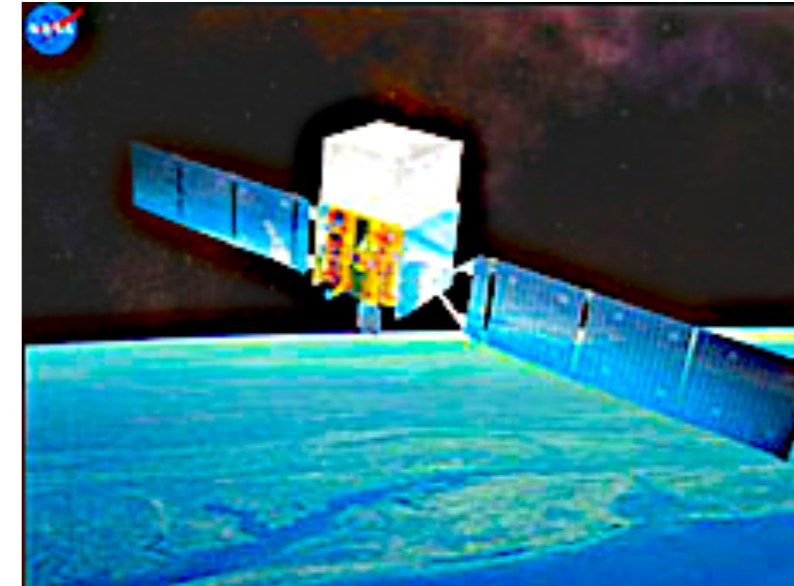
e.g. LHC

Direct Detection



e.g. LUX

Indirect Detection



e.g. FERMI

Experiments are getting increasingly sensitive...
but we still haven't found it

Status in 2019*

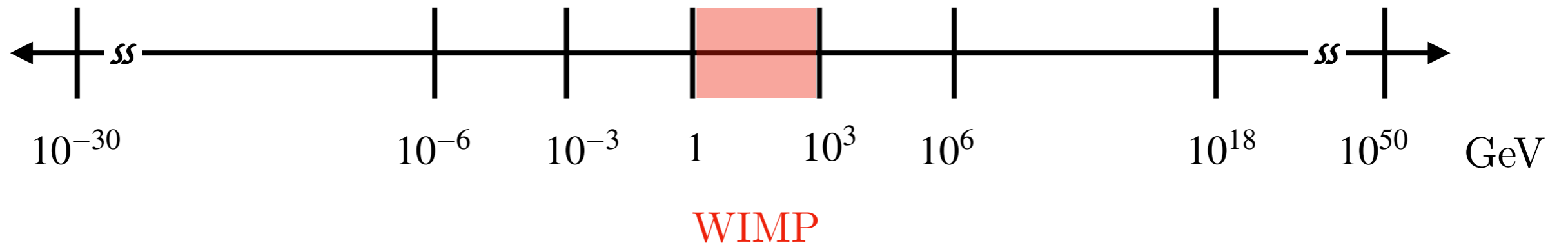
Dominant paradigm being challenged.

Great opportunity for new ideas!

*2020 never happened

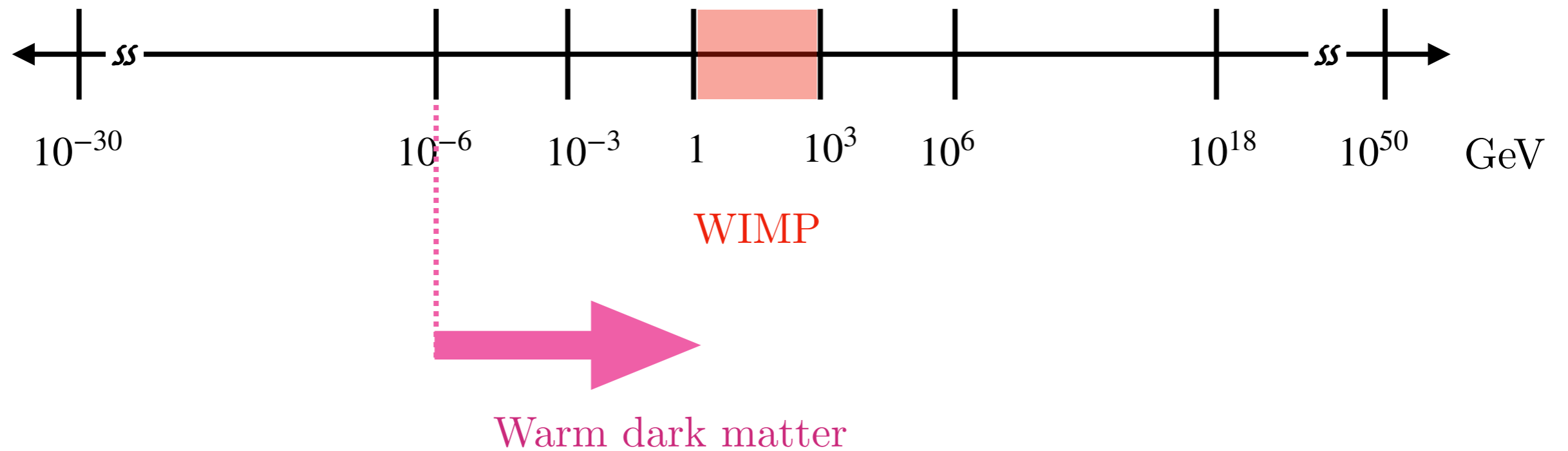
Beyond the WIMP

← dark matter mass →



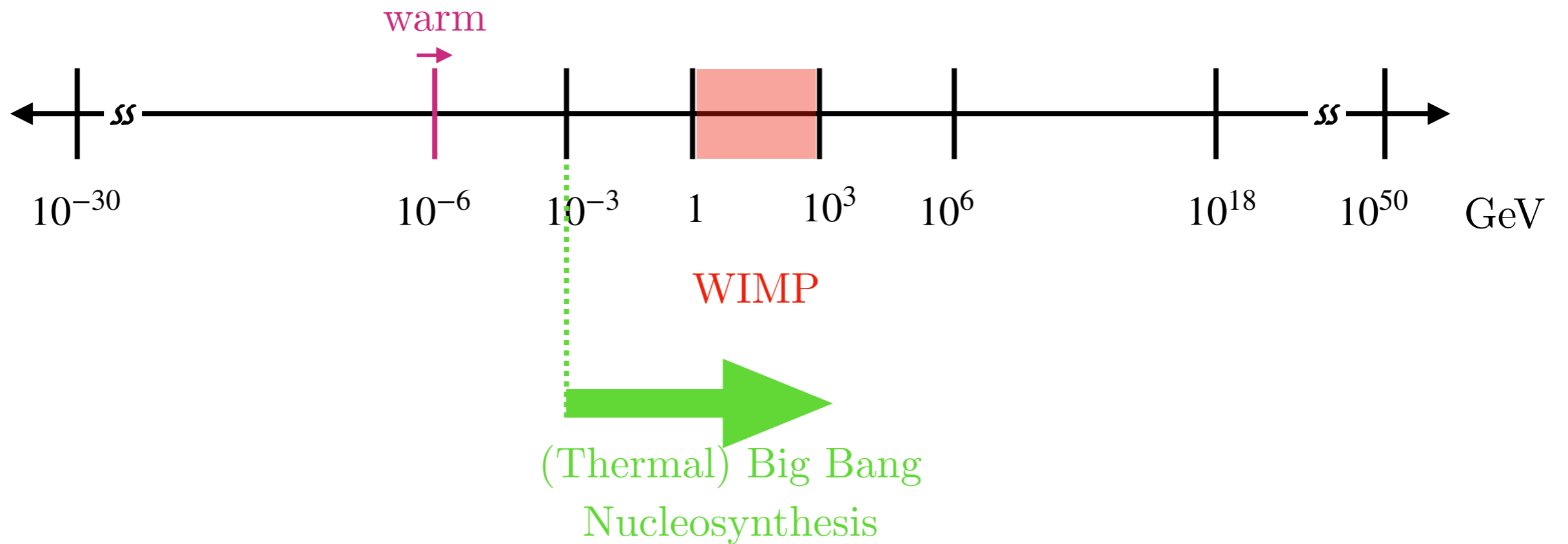
Beyond the WIMP

← dark matter mass →



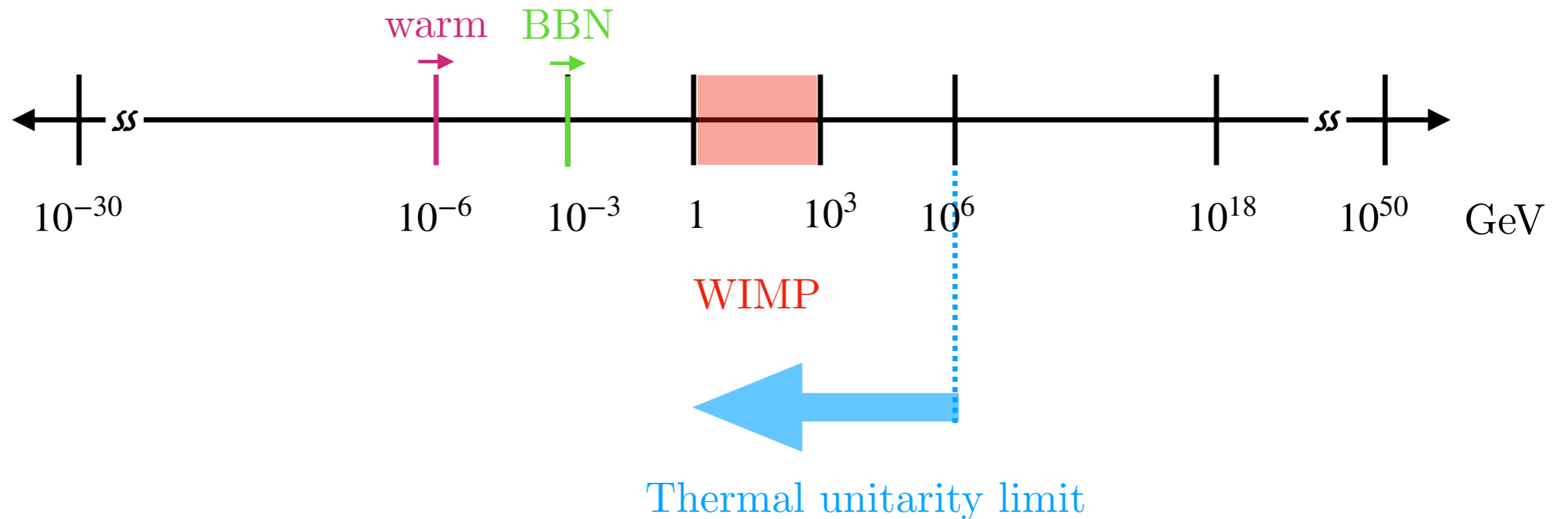
Beyond the WIMP

← dark matter mass →



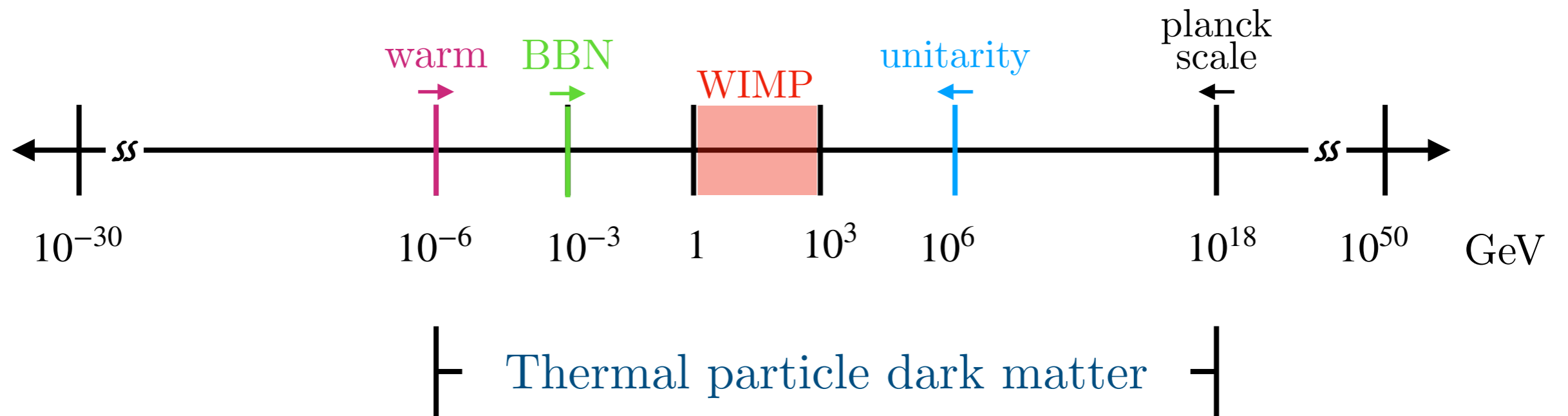
Beyond the WIMP

← dark matter mass →



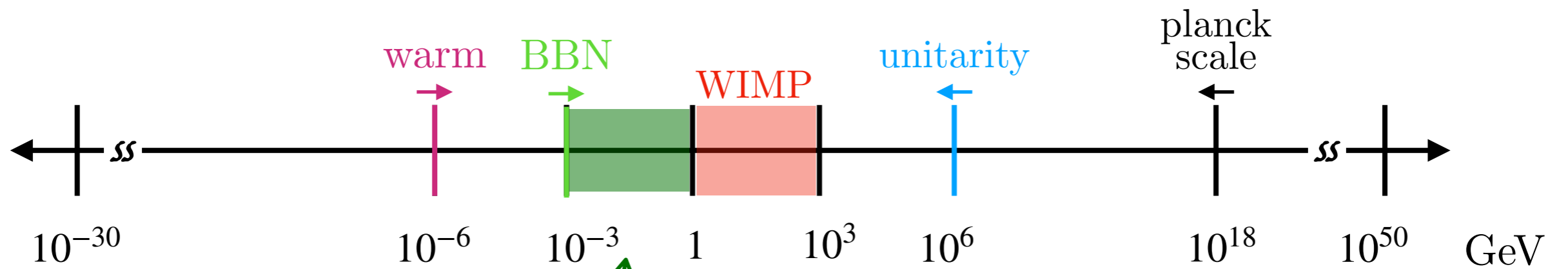
Beyond the WIMP

← dark matter mass →



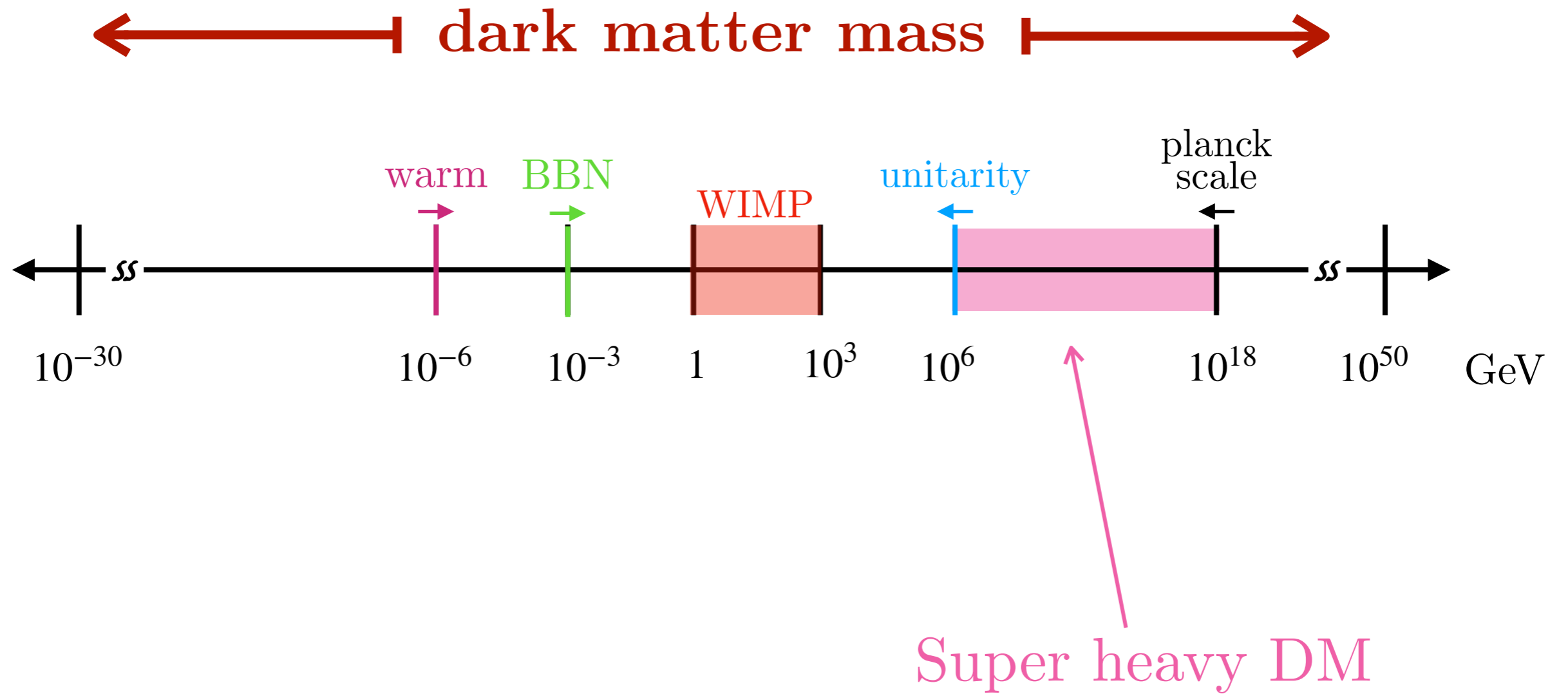
Beyond the WIMP

← dark matter mass →



Light dark matter

Beyond the WIMP

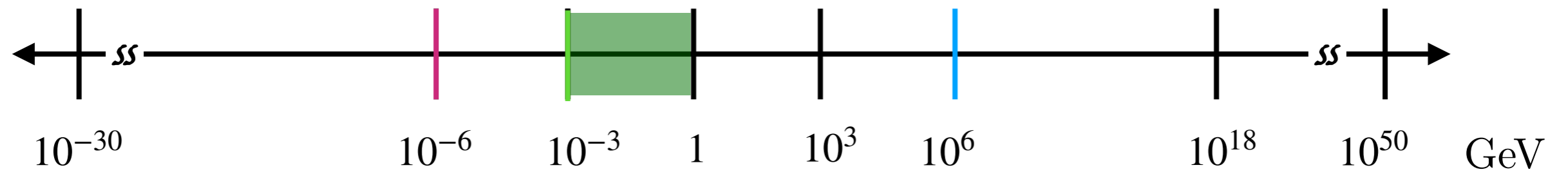


New Theory Ideas

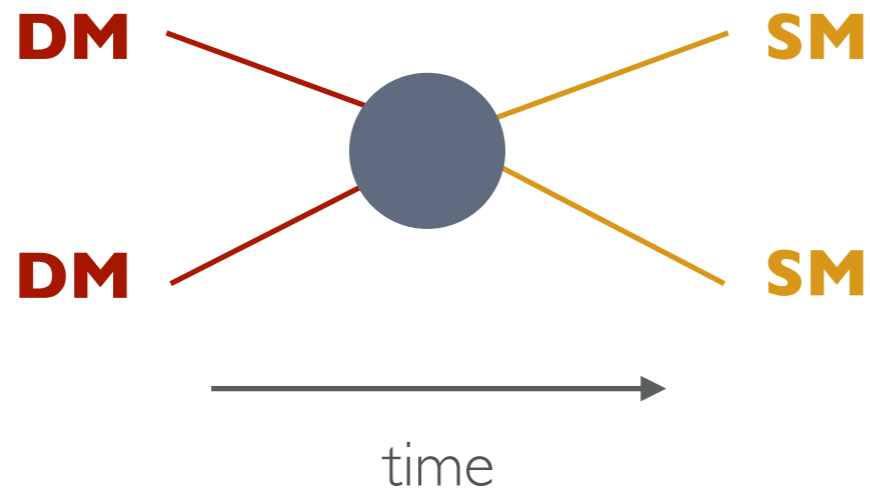
-
- Weakly coupled WIMPs [Pospelov, Ritz, Voloshin 2007; Feng, Kumar 2008]
- SIMPs [Hochberg, **EK**, Volansky, Wacker, 2014; + Murayama, 2015]
- ELDERs [**EK**, Perelstein, Rey-Le Lorier, Tsai, 2016]
- Forbidden dark matter [Griest, Seckel 1991; D'Agnolo, Ruderman, 2015]
- Co-scattering dark matter [D'Agnolo, Pappadopulo, Ruderman, 2017]
- Sub-MeV thermal dark matter [Berlin, Blinov 2017]
- Super heavy thermal dark matter [Kim, **EK** 2019]
- Zombies [Kramer, **EK**, Levi, Outmezguine, Ruderman, 2020]

... are abundant

Light dark matter



Ex. 1: Weakly Coupled $2 \rightarrow 2$



$$\langle \sigma_{\text{ann}} v \rangle = \frac{\alpha^2}{m_{\text{DM}}^2}$$

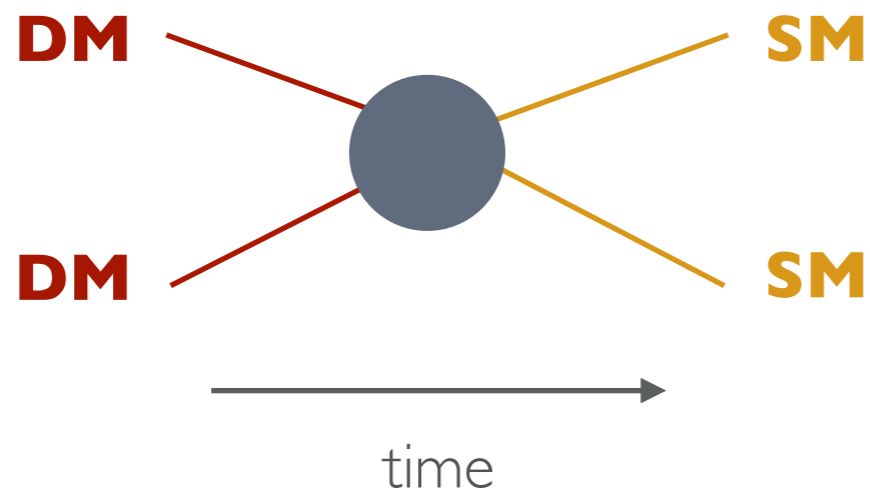
$$m_{\text{DM}} = \alpha \times 30 \text{ TeV}$$

$$\alpha \ll 1$$

[Pospelov, Ritz, Voloshin 2007

Feng, Kumar 2008]

Ex. 2: Forbidden Channels



$$\langle \sigma_{\text{ann}} v \rangle = \frac{\alpha^2}{m_{\text{DM}}^2}$$

$$m_{\text{DM}} < m_{\text{SM}}$$

$$m_{\text{DM}} = \alpha \times 30 \text{ TeV} \times e^{-x_f \Delta}$$

freezeout
temp

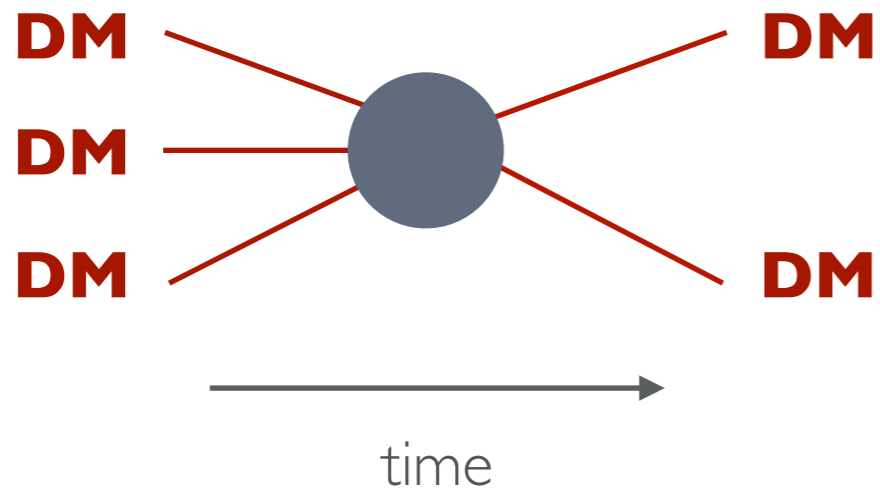
$$x = m/T$$

mass
difference

Forbidden @ T=0;

Proceeds via Boltzmann tail

Ex. 3: SIMPs



$$\langle \sigma v^2 \rangle_{3 \rightarrow 2} = \frac{\alpha^3}{m_{\text{DM}}^5}$$

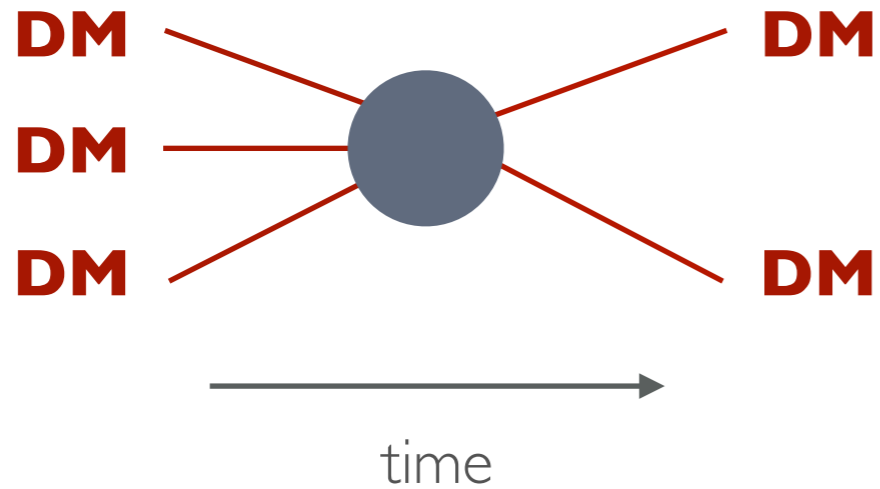
$$m_{\text{DM}} = \alpha \times 100 \text{ MeV}$$

For strong coupling, strong scale emerges

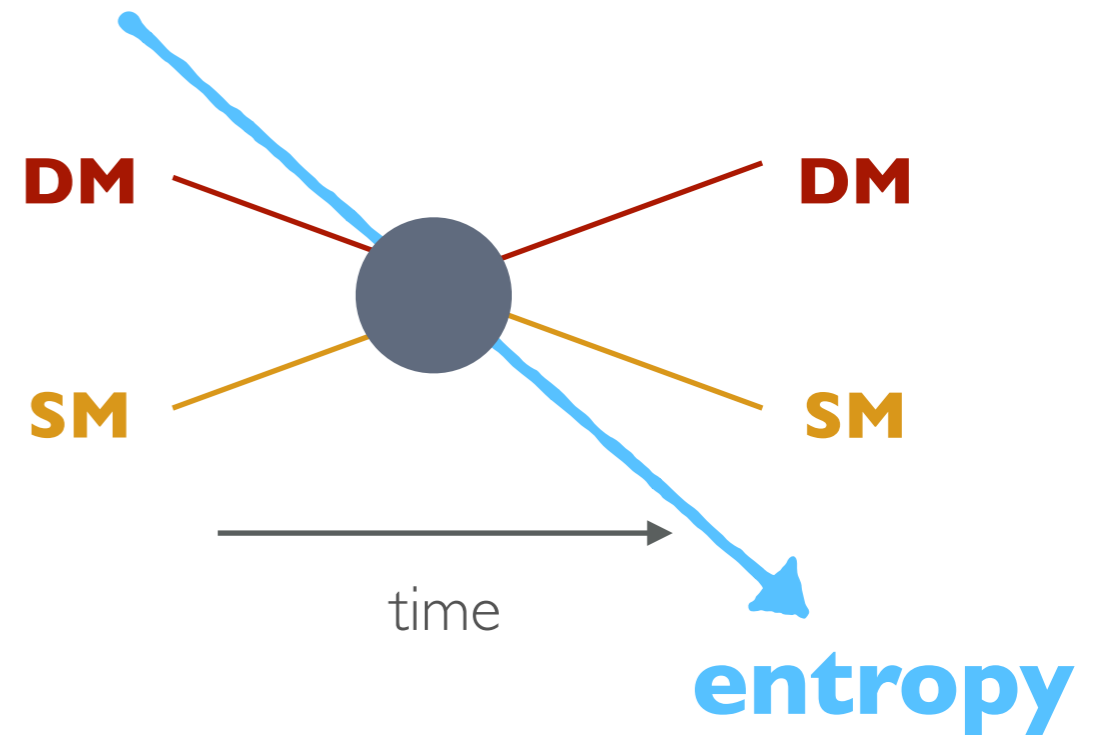
Strongly Interacting Massive Particle (SIMP)

Ex. 3: SIMPs

Pumps heat



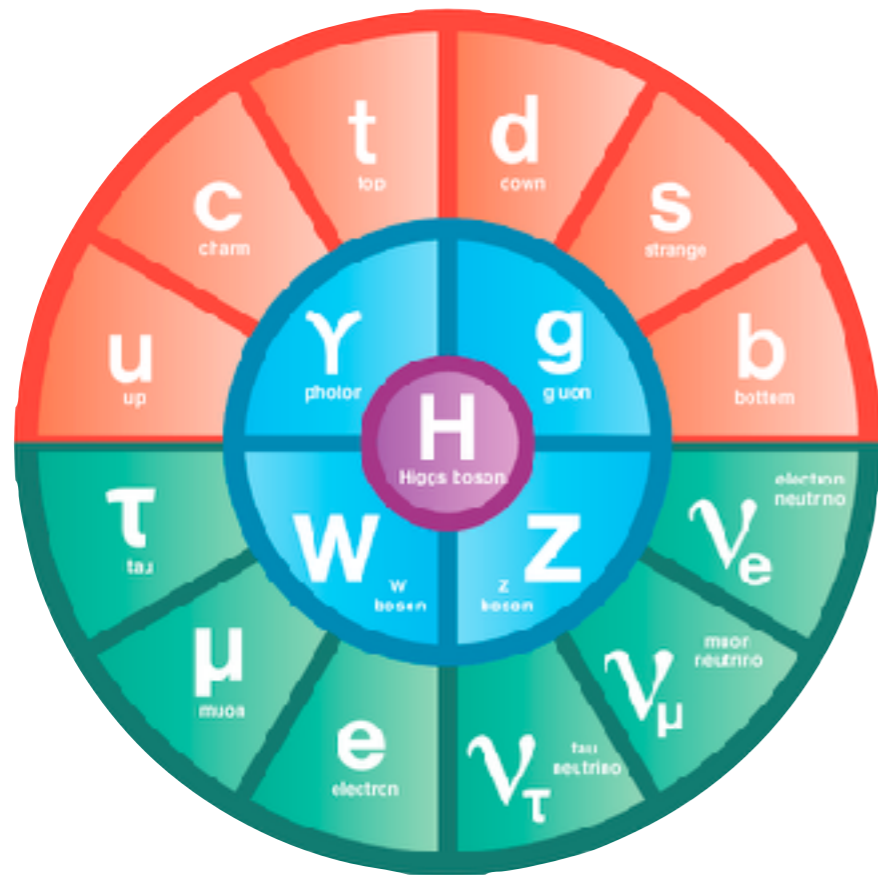
Shed the heat



Generic.

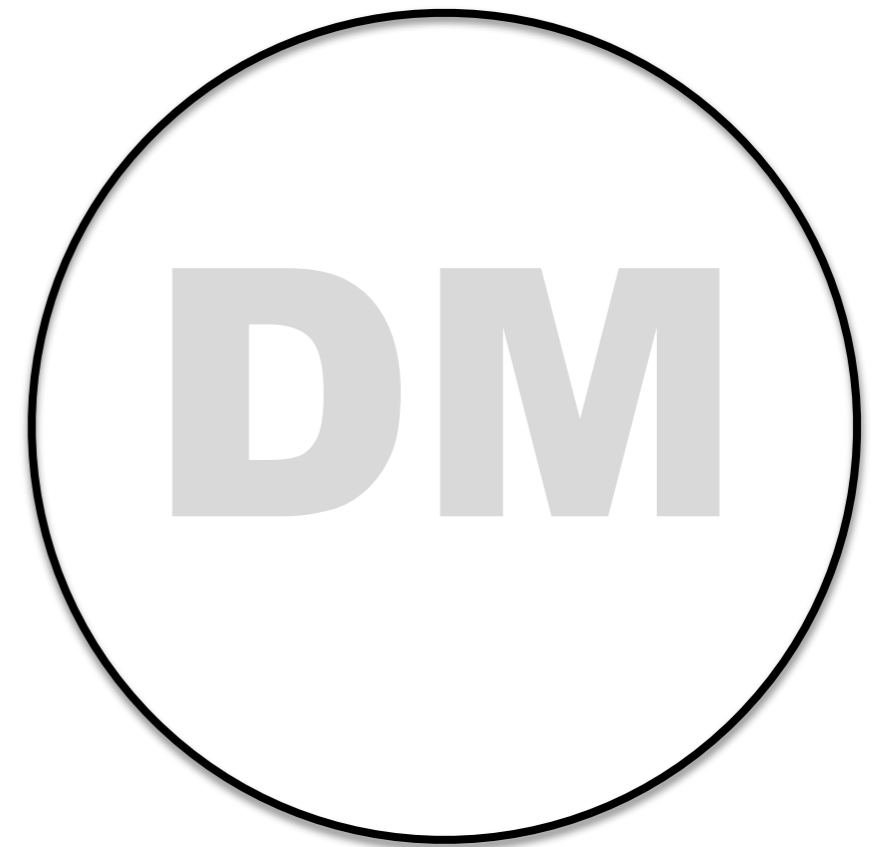
Dark Sectors

Visible sector



SM is a
particle zoo.

Dark sector



Why not in the
dark sector too?

QCD-like sector

Think Standard Model!

Dark matter from strongly coupled gauge theories

e.g. $SU(3)_{\text{dark}} \times U(1)_{\text{dark}}$

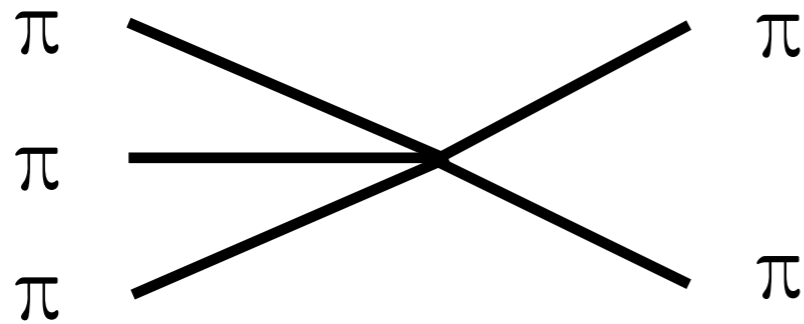
$Sp(N_c), SU(N_c), SO(N_c)$

kinetically mixed
hidden photon (V)

QCD-like theories, pions = dark matter.

Many processes, many dark matter mechanisms.

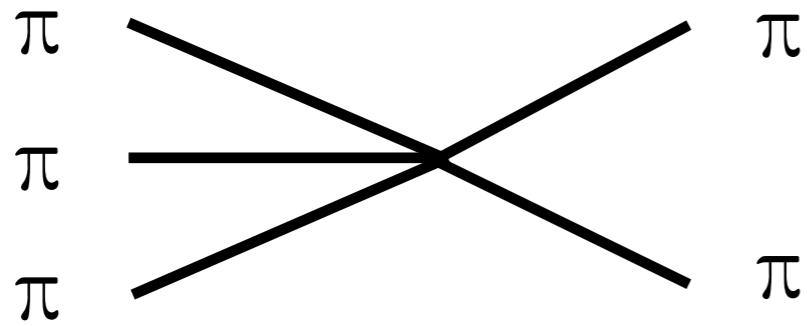
Processes



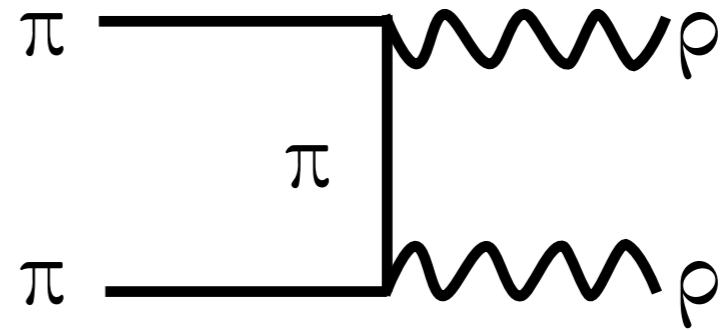
$3 \rightarrow 2$ processes

(From the Wess-Zumino-Witten term. In QCD describes $K \rightarrow \pi \pi \pi$)

Processes



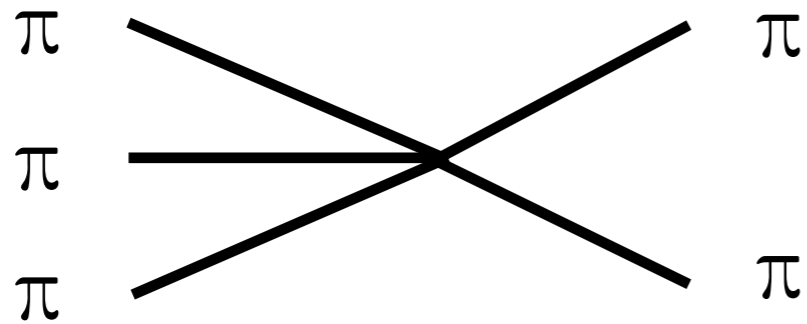
3 \rightarrow 2 processes



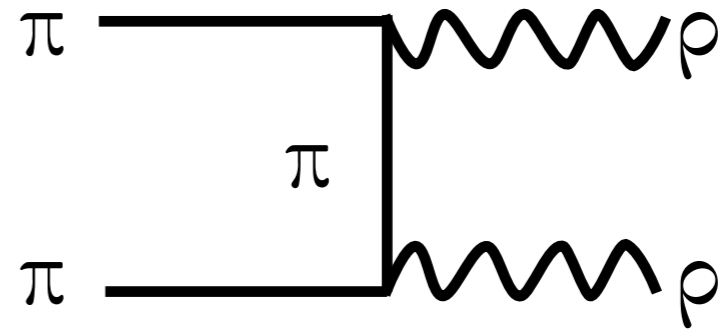
forbidden

$$m_\rho \gtrsim m_\pi$$

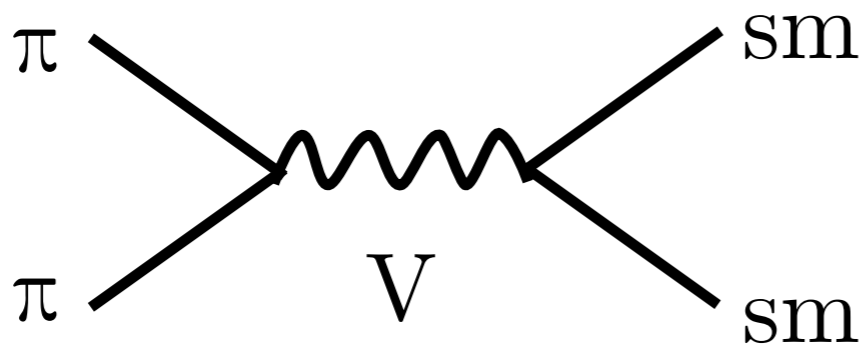
Processes



3→2 processes

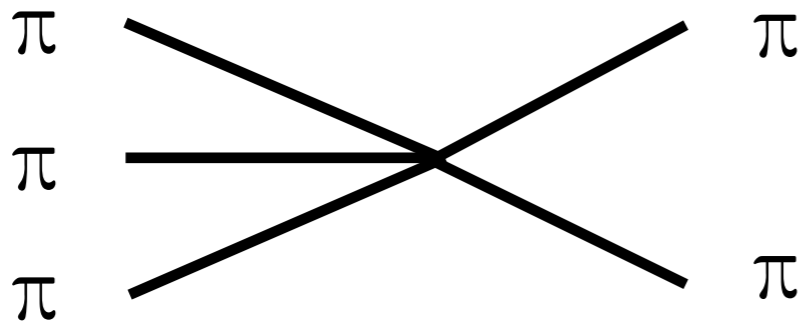


forbidden

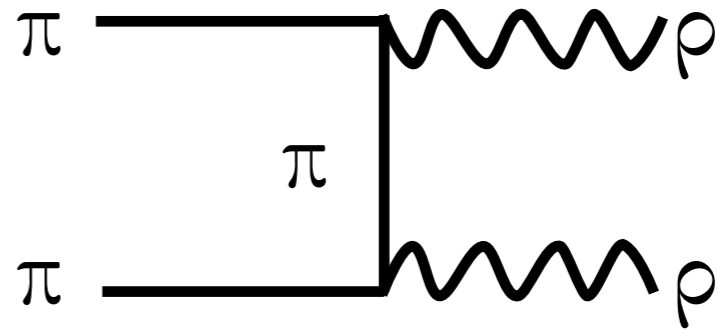


2→2 annihilations

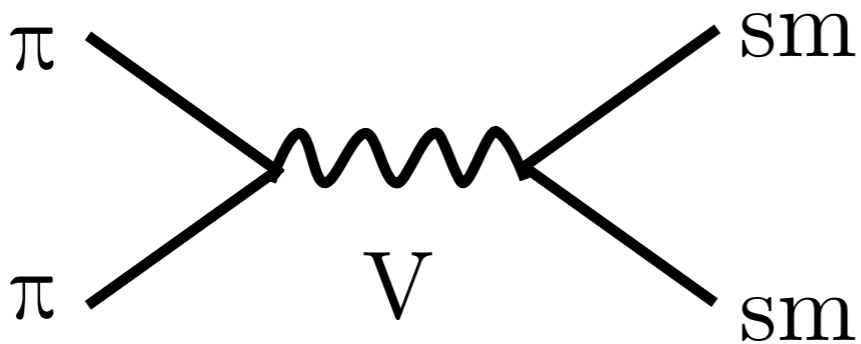
Processes



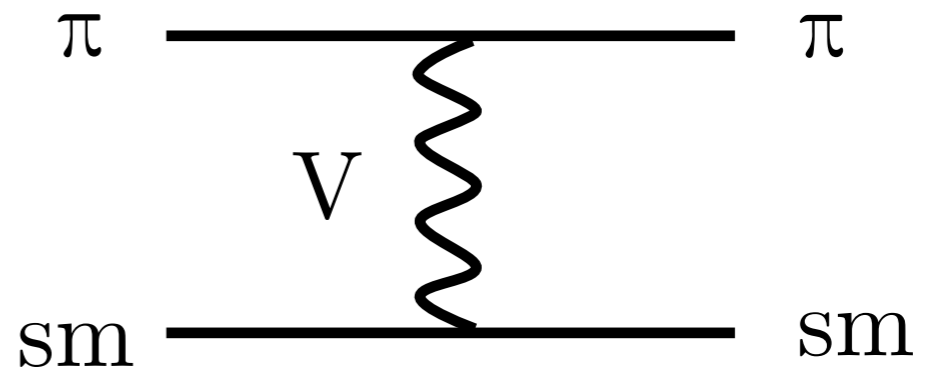
3→2 processes



forbidden



2→2 annihilations



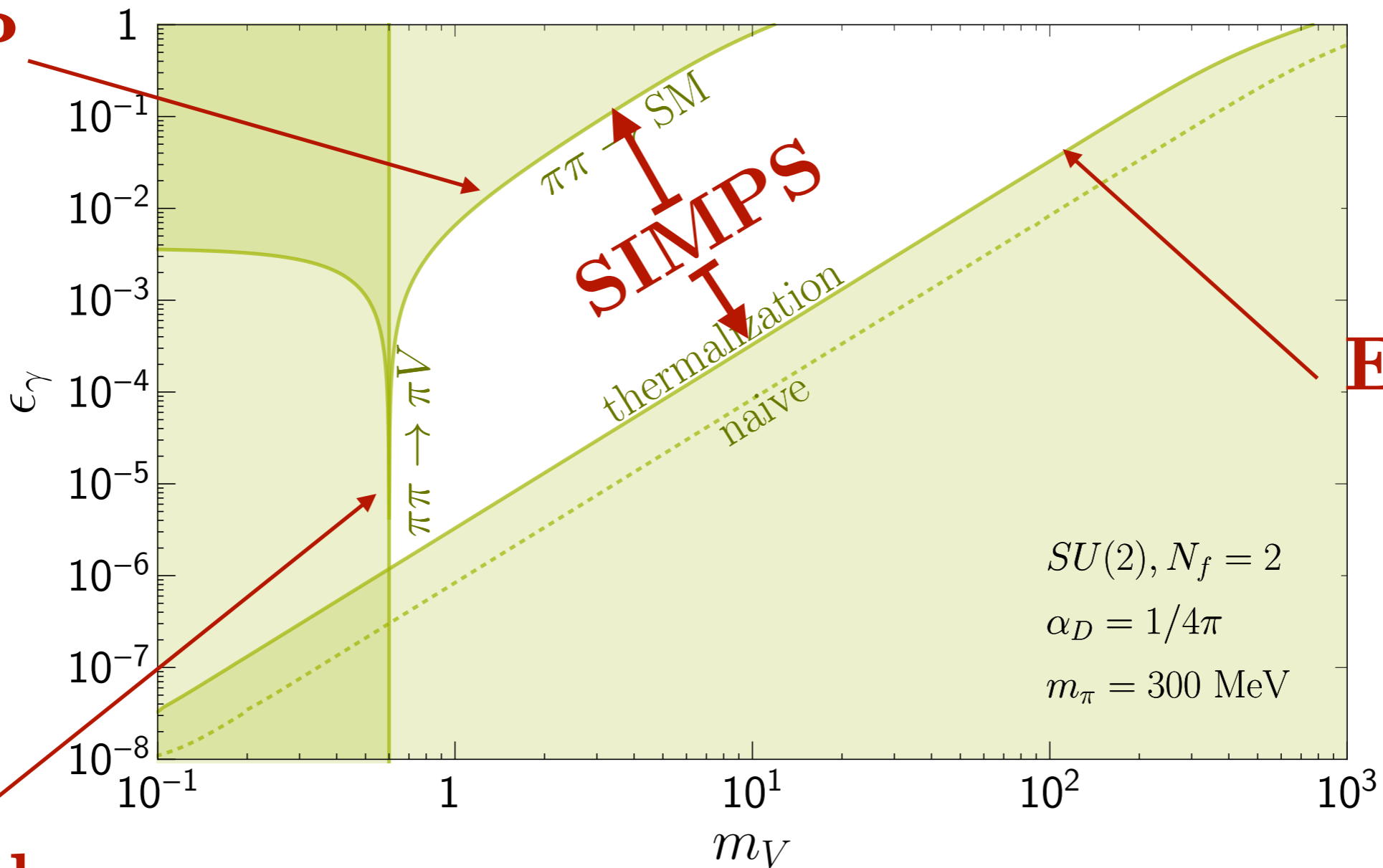
elastic scattering

Predictive

Kinetically mixed U(1) mediator

WIMP

Coupling of V to
the EM current



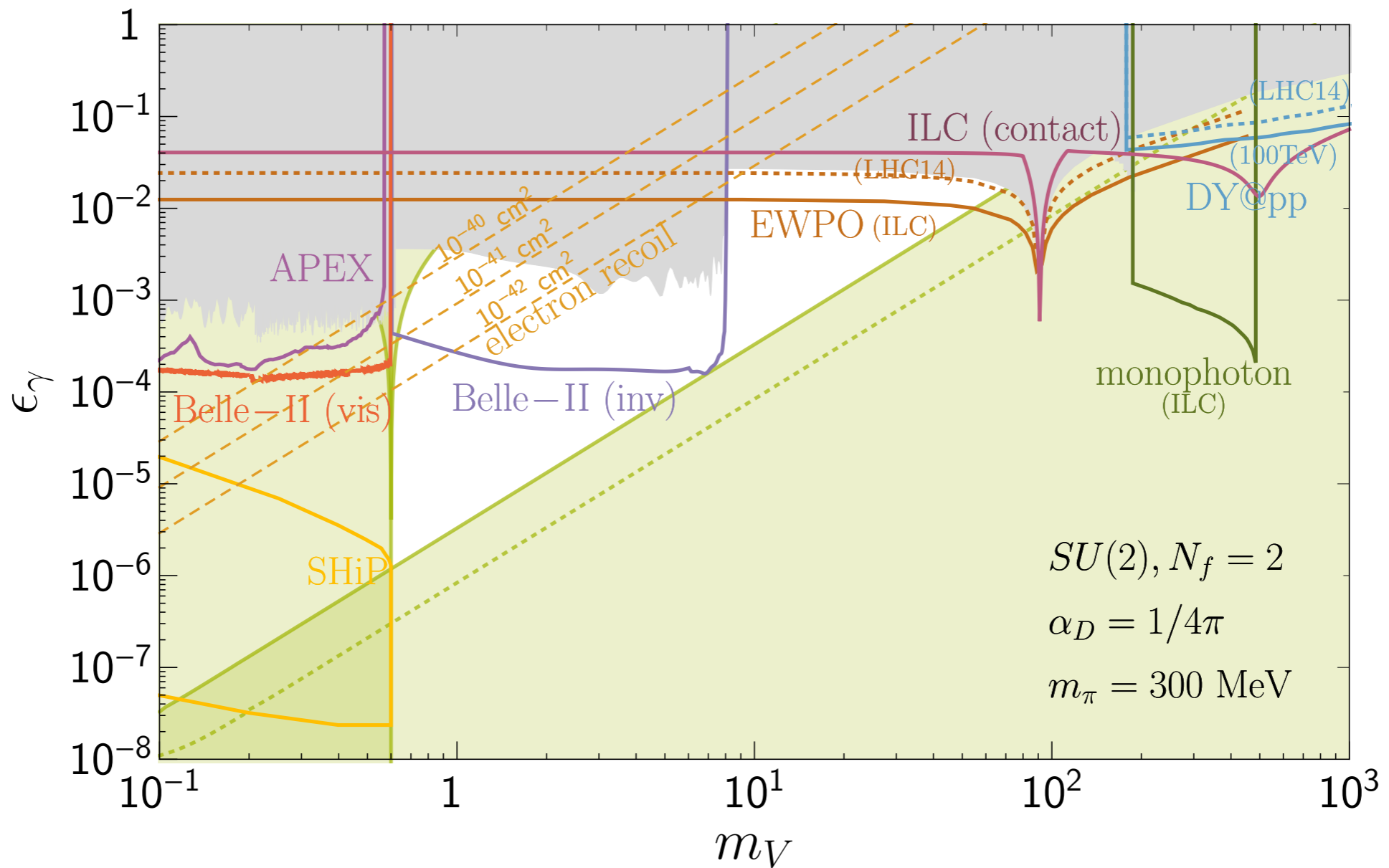
ELDER

Forbidden

Predictive

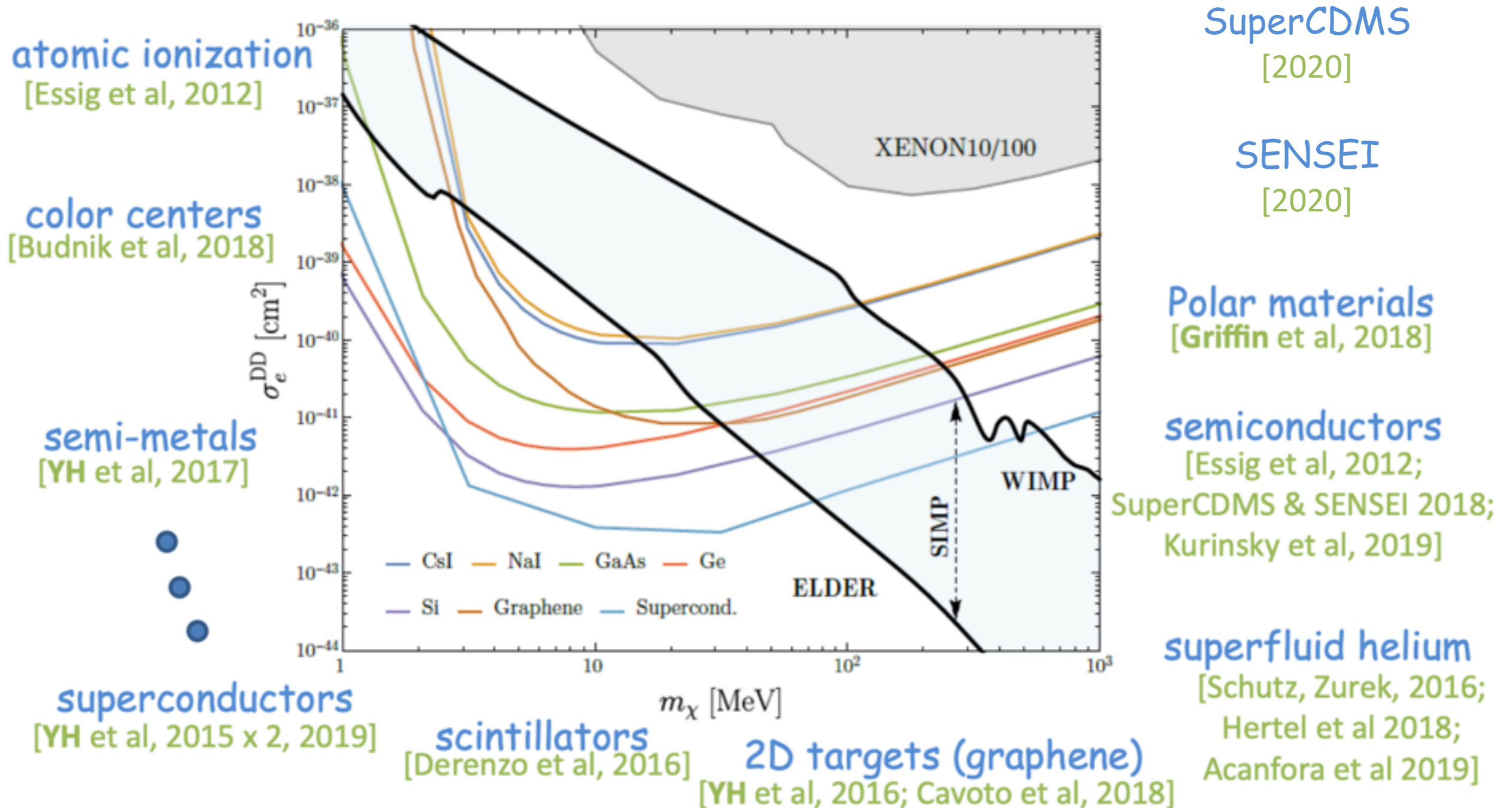
Kinetically mixed U(1) mediator

Coupling of V to
the EM current



Future probes

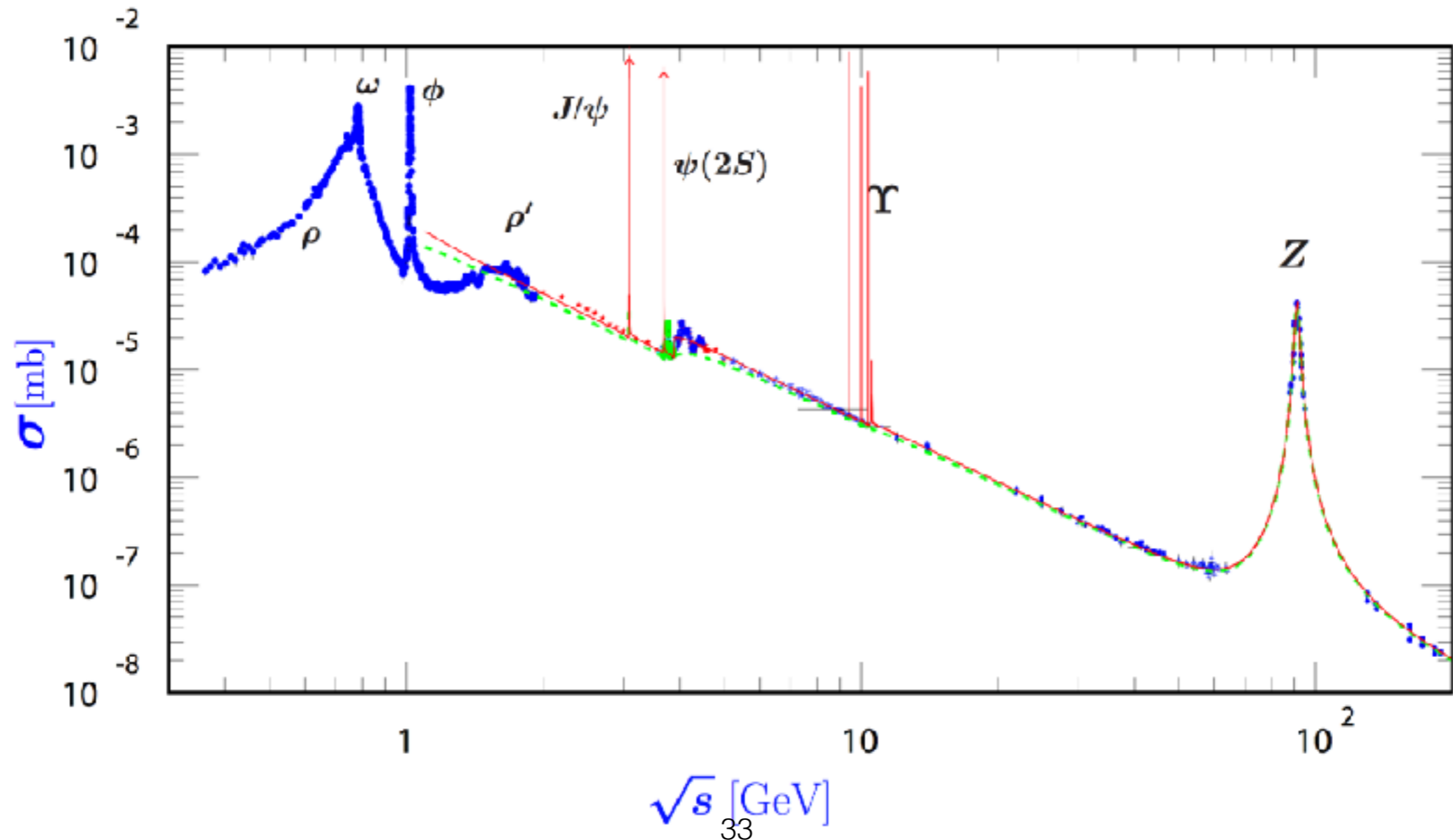
Direct Detection



Dark Spectroscopy

$$e^+e^- \rightarrow \text{resonances}$$

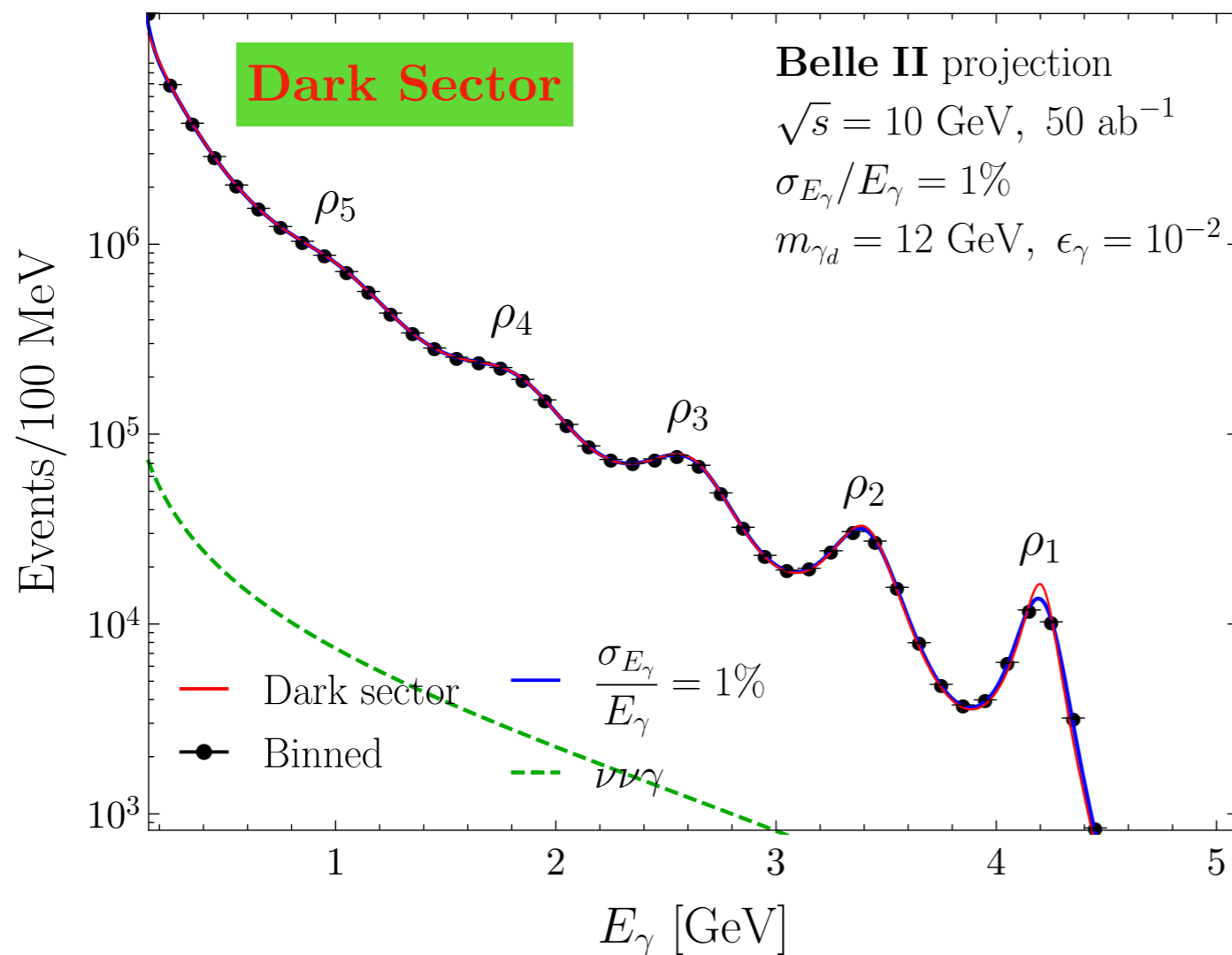
center of mass energy traces the **QCD** resonance structure



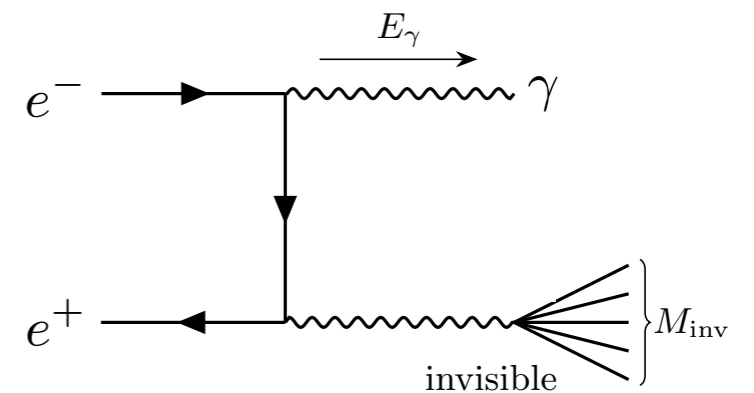
Dark Spectroscopy

$$e^+e^- \rightarrow \gamma + \text{invisible resonances}$$

mono-photon energy also traces the resonance structure

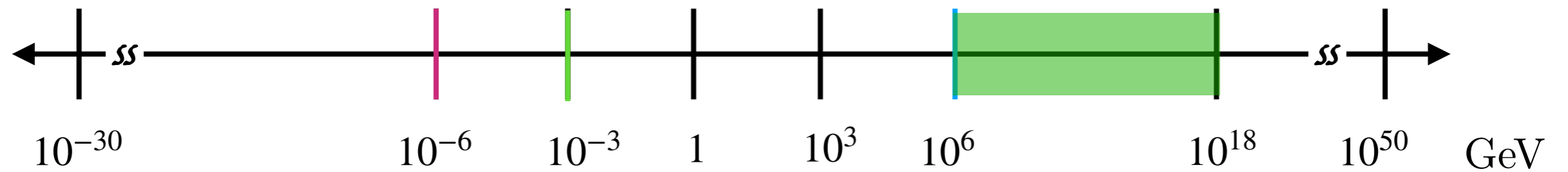


$$E_\gamma = \frac{\sqrt{s}}{2} \left(1 - \frac{M_{\text{res}}^2}{s} \right)$$

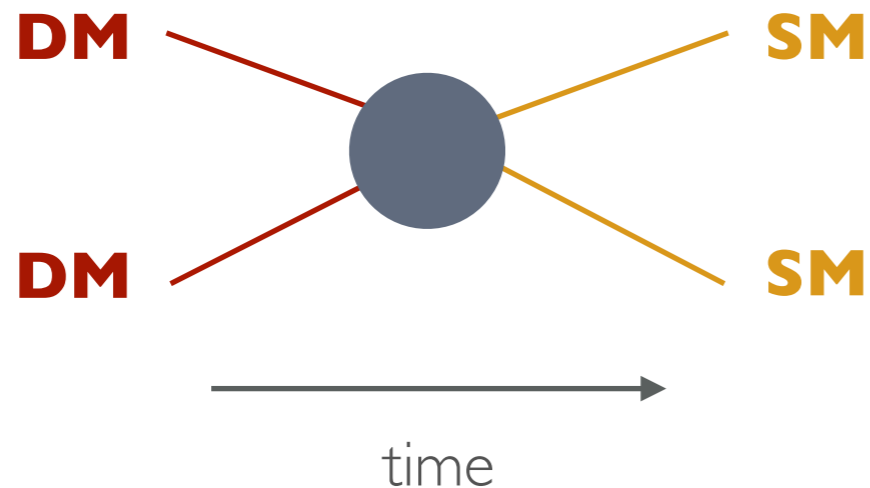


[Hochberg, EK, Murayama, 2016, 2017]

Super heavy dark matter



Unitarity Bound



$$\langle \sigma_{\text{ann}} v \rangle = \frac{\alpha^2}{m_{\text{DM}}^2}$$

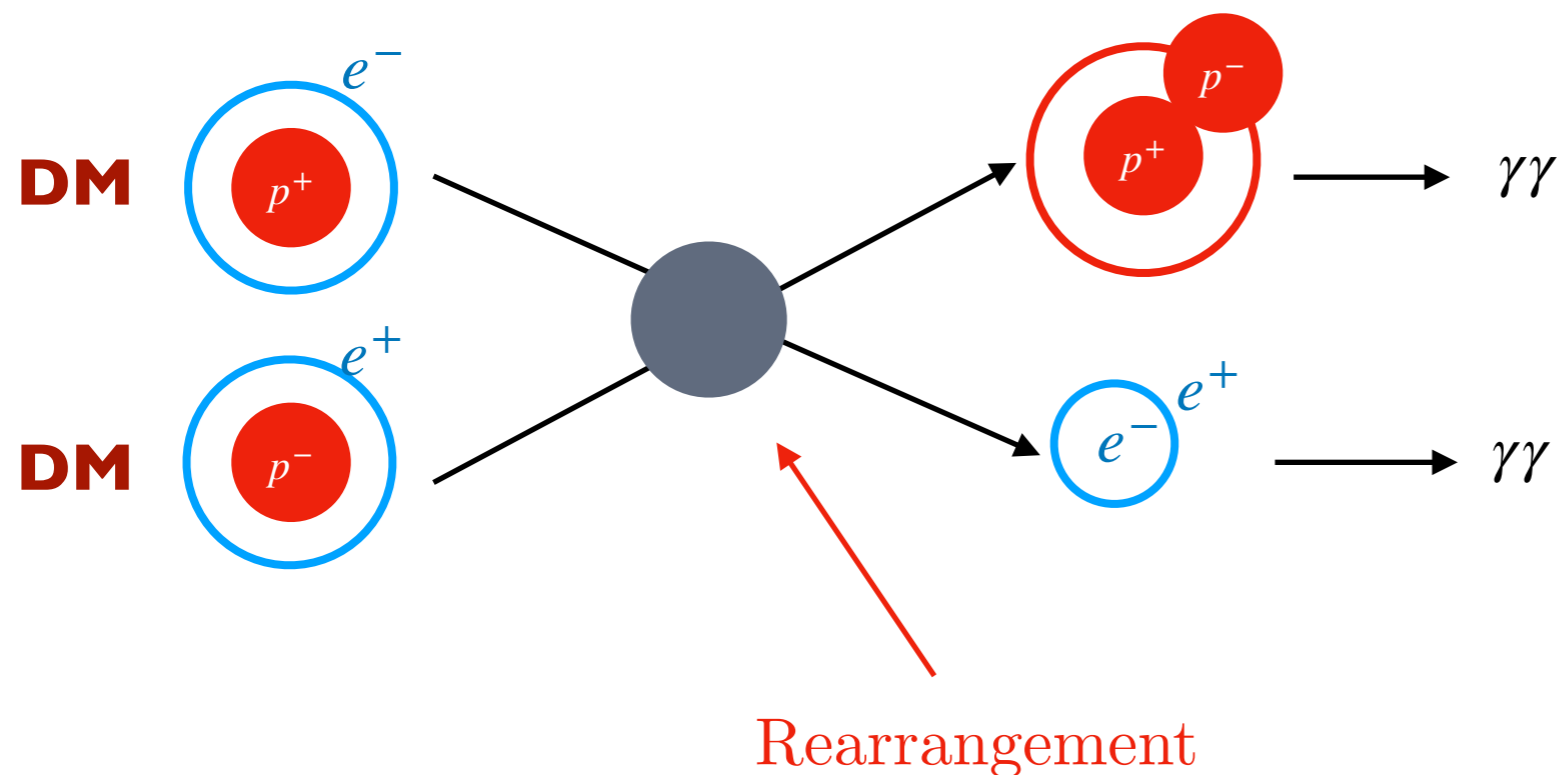
Correct relic abundance for $m_{\text{DM}} = \alpha \times 30 \text{ TeV}$

For perturbative couplings $\alpha < 4\pi$

$$m_{\text{DM}} \lesssim 300 \text{ TeV}$$

Ex. 1: Composite Interactions

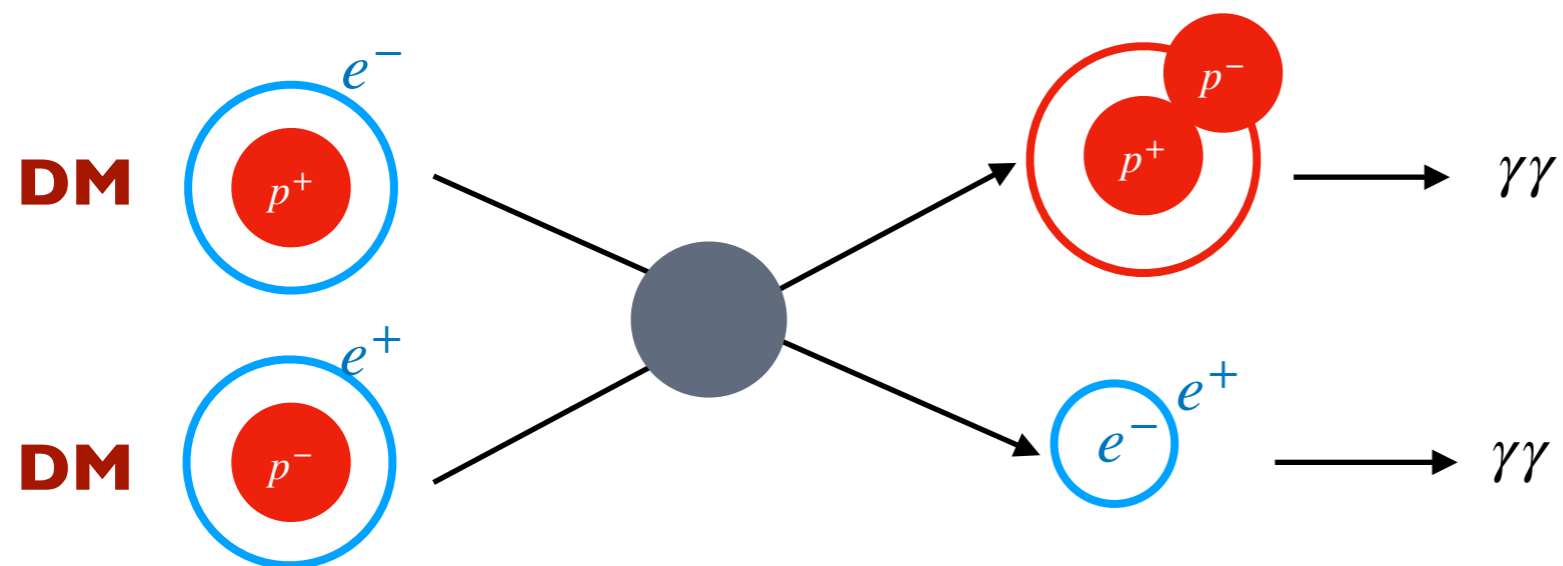
dark hydrogen anti-hydrogen annihilation



Harigaya, Ibe, Kaneta, Nakano, Suzuki (2016); J. Smirnov, J. F. Beacom, (2019); Contino, Mitridate, Podo, Redi, (2019); Gross, Mitridate, Redi, Smirnov, Strumia (2019); Geller, Iwamoto, Lee, Shadmi, Telem (2018)

Ex. 1: Composite Interactions

dark hydrogen anti-hydrogen annihilation



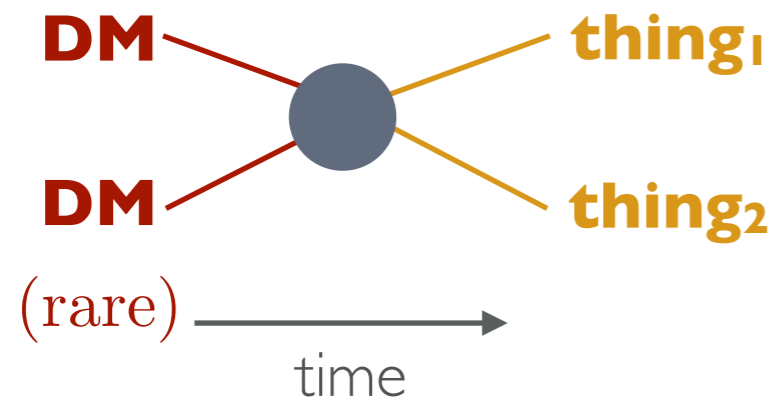
$$\sigma_{\text{ann}} \simeq 4\pi a_0^2$$

Bohr Radius

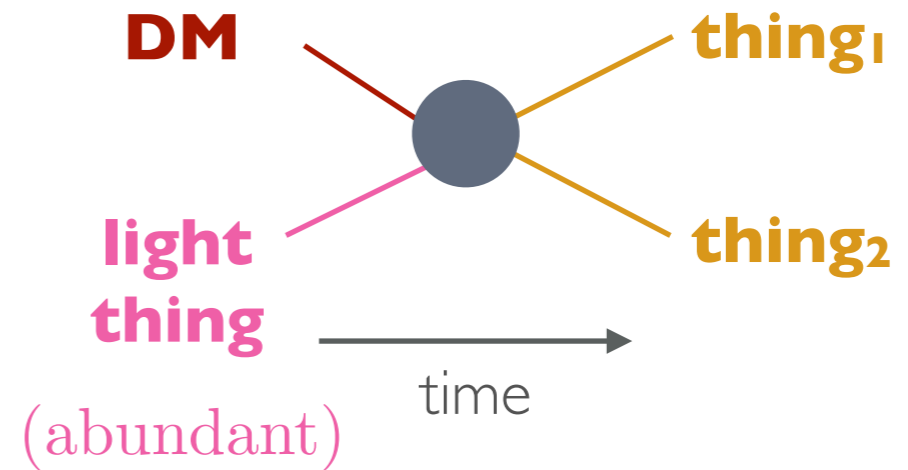
$$a_0 \simeq \frac{1}{\alpha m_e}$$

Predicts much heavier DM

Compare Processes



vs.



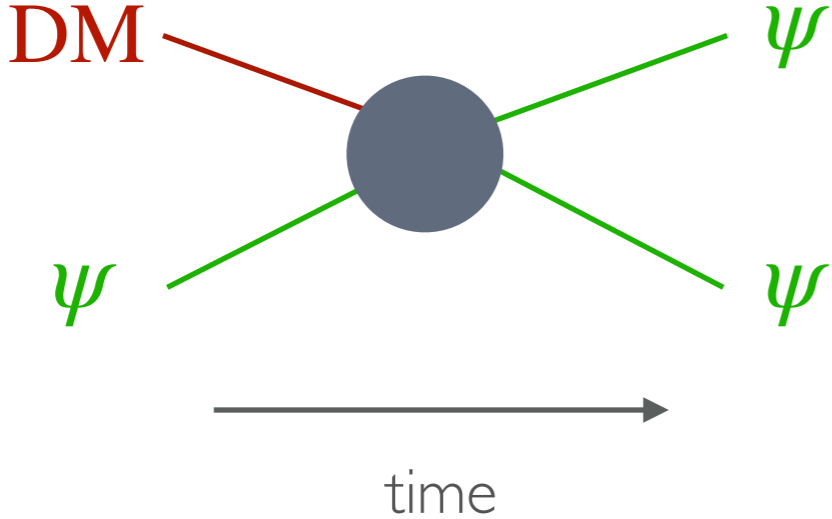
$$\Gamma_{\text{ann}} = n_{\text{DM}} \langle \sigma_{\text{ann}} v \rangle \propto e^{-m_{\text{DM}}/T}$$

$$\Gamma_{\text{ann}} = n_{\text{light}} \langle \sigma_{\text{ann}} v \rangle$$

Less efficient

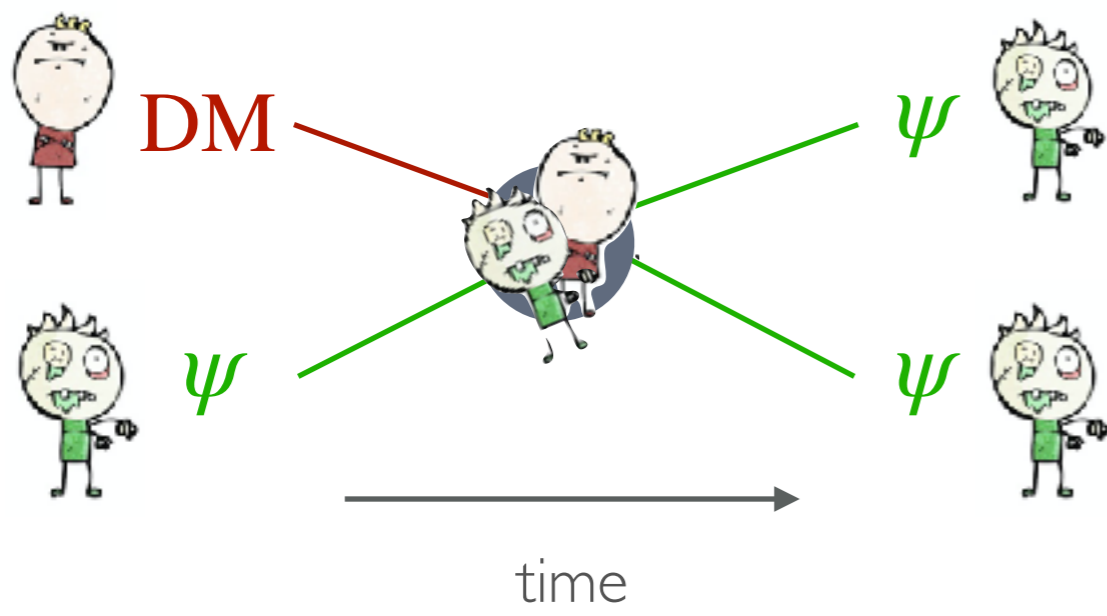
Much more efficient!

Ex. 2: Zombies



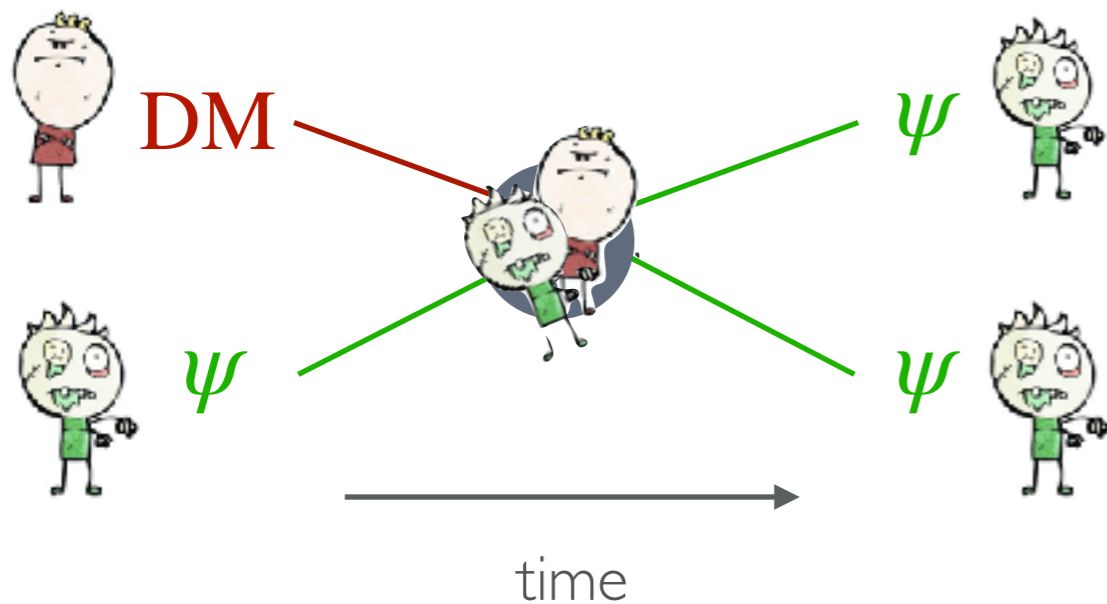
$$m_{DM} > m_{\psi}$$

Ex. 2: Zombies



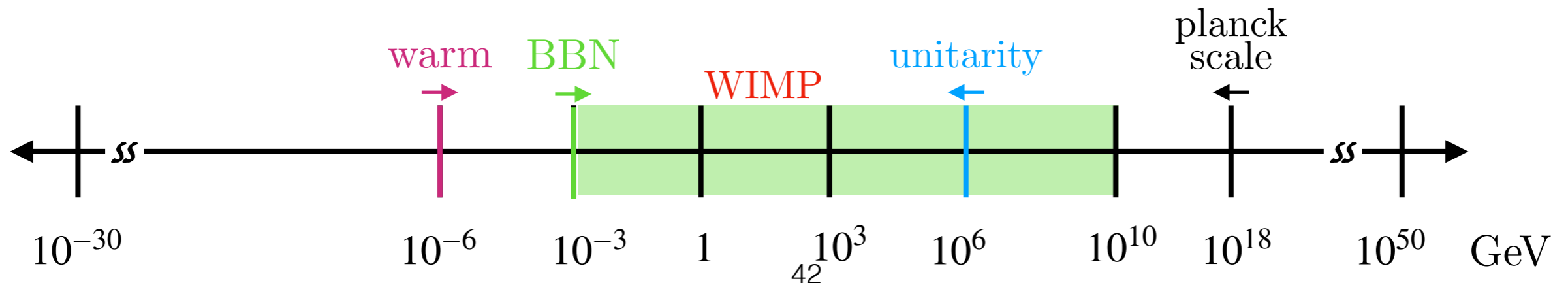
$$m_{DM} > m_{\psi}$$

Ex. 2: Zombies

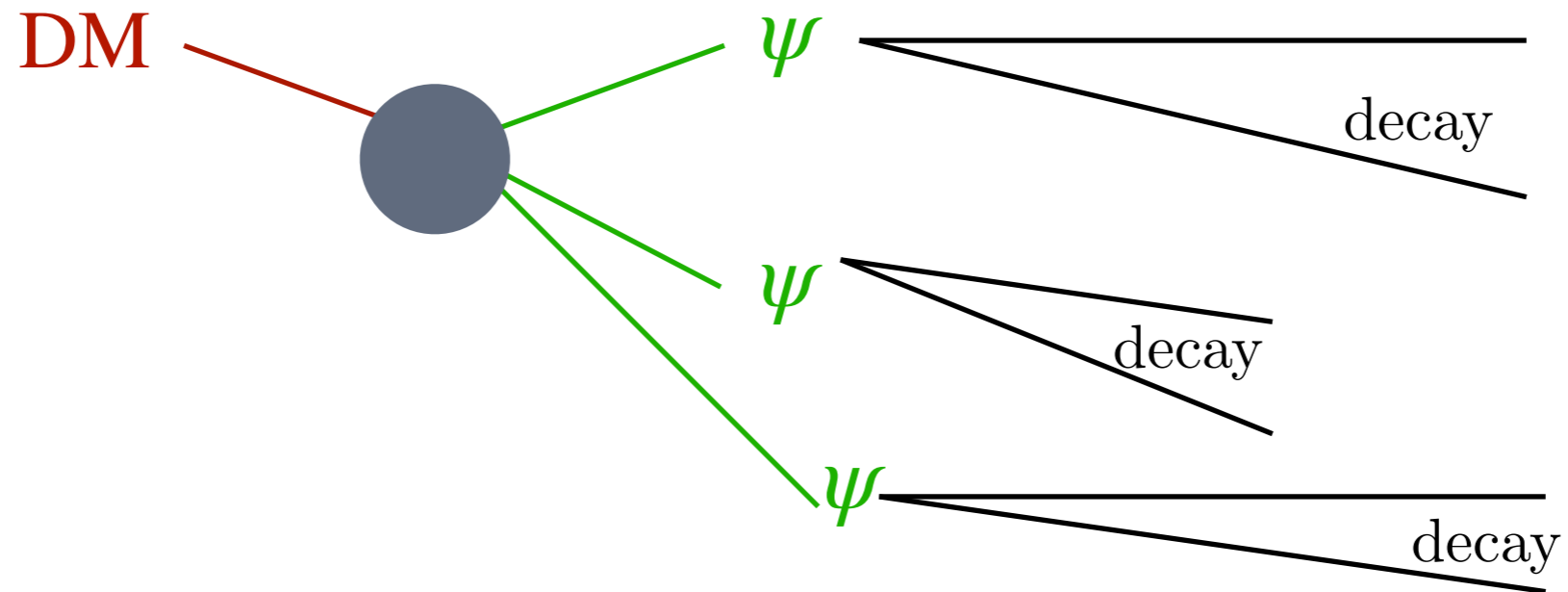


$$\langle \sigma_{\text{zombie}^2} \rangle = \frac{\alpha^2}{m_{\text{DM}}^2}$$

$$m_{\text{DM}} = \alpha^{2/3} \times 10^6 \text{ TeV} \quad (m_{\text{DM}} = 2m_{\psi})$$



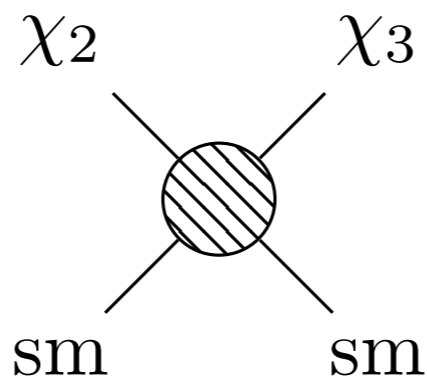
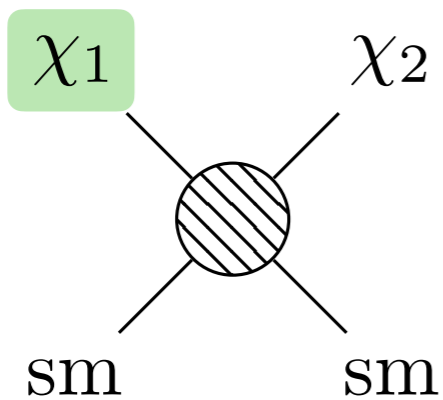
Ex. 2: Zombies



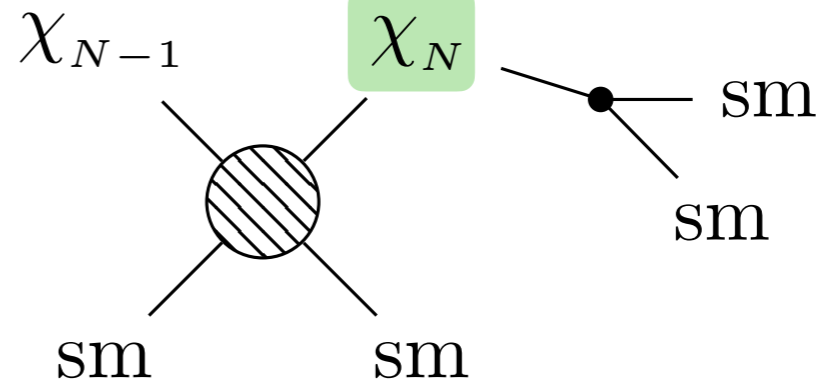
Metastable DM with strong indirect detection signal

Ex. 3: Chain DM

DM candidate



...

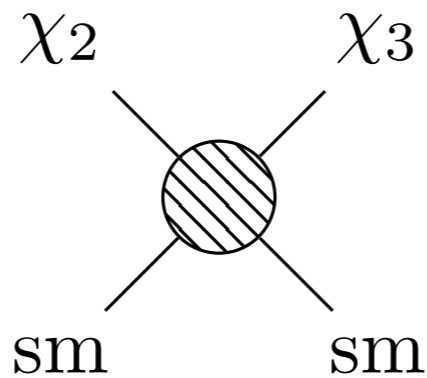
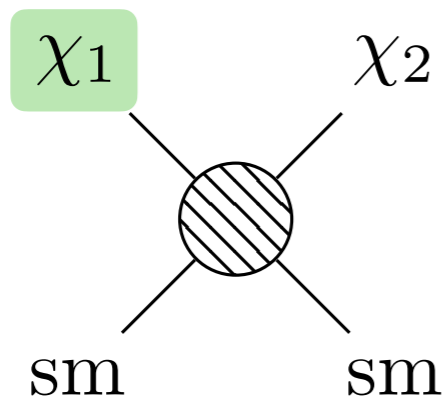


decays

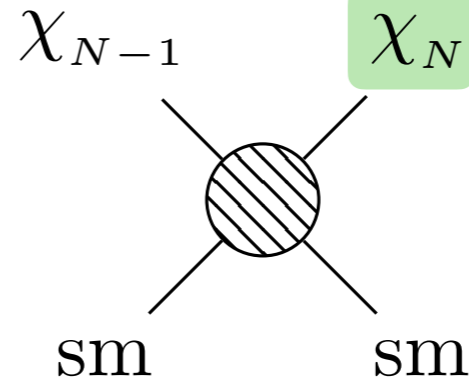
Very efficient because the SM particles are abundant

Ex. 3: Chain DM

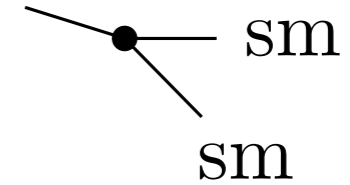
DM candidate



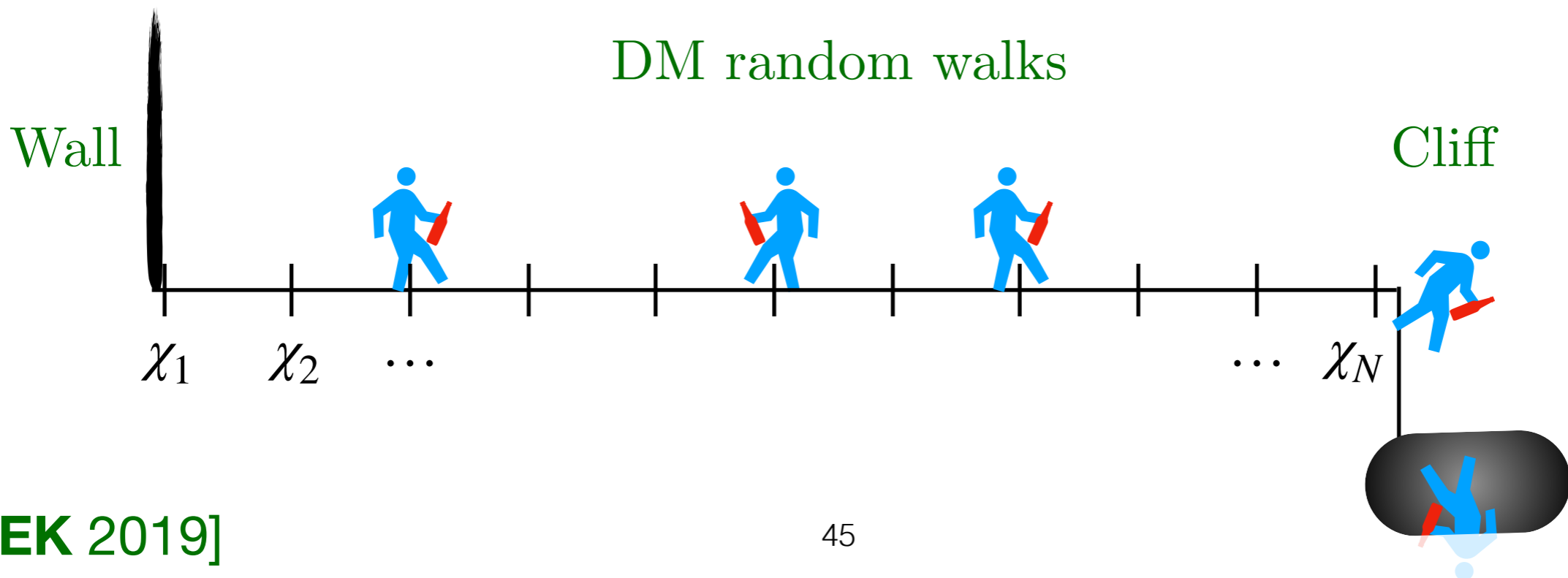
...



decays

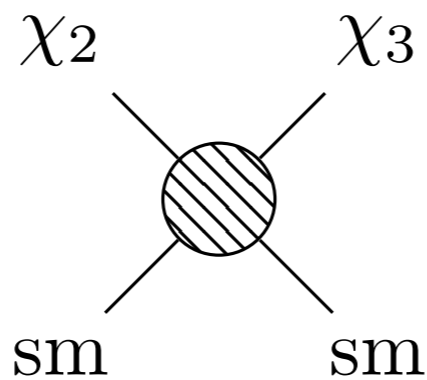
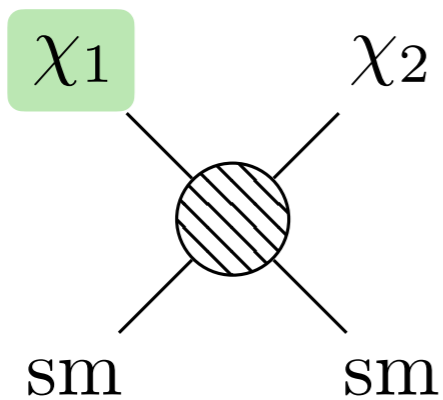


DM random walks

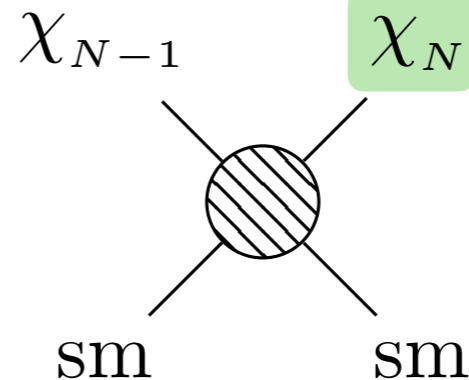


Ex. 3: Chain DM

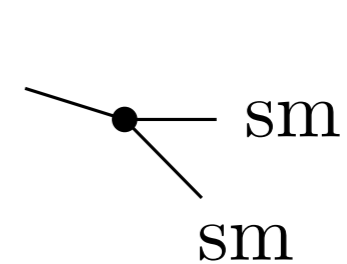
DM candidate



...



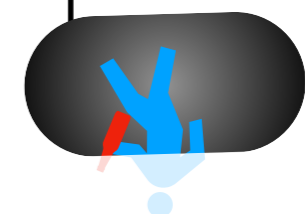
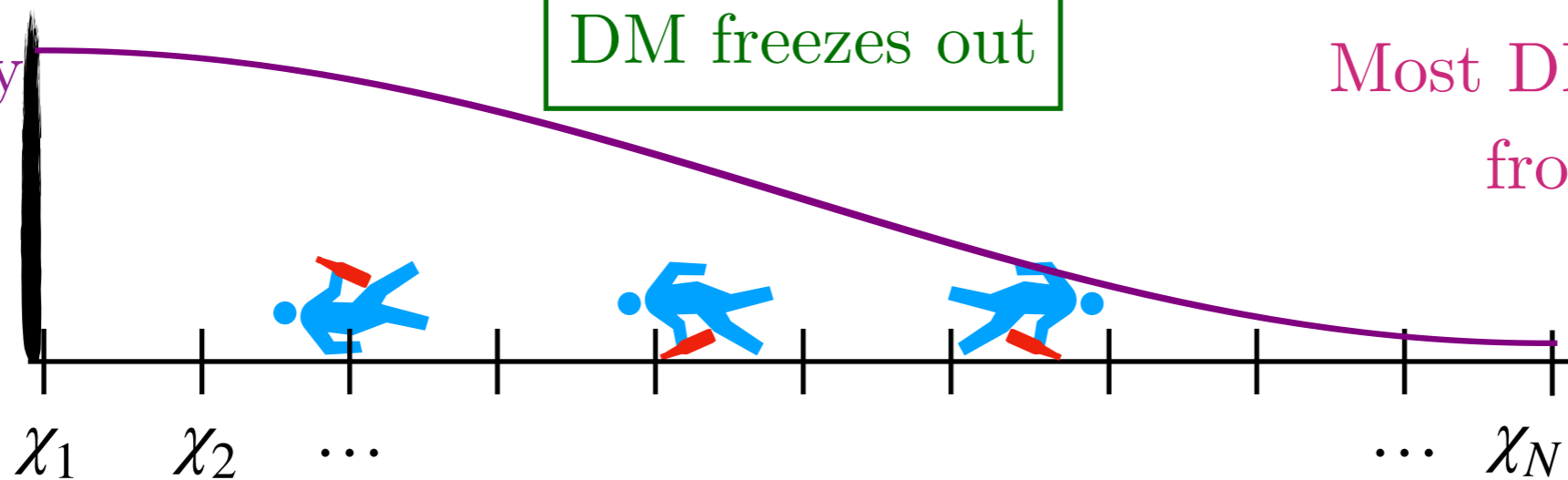
decays



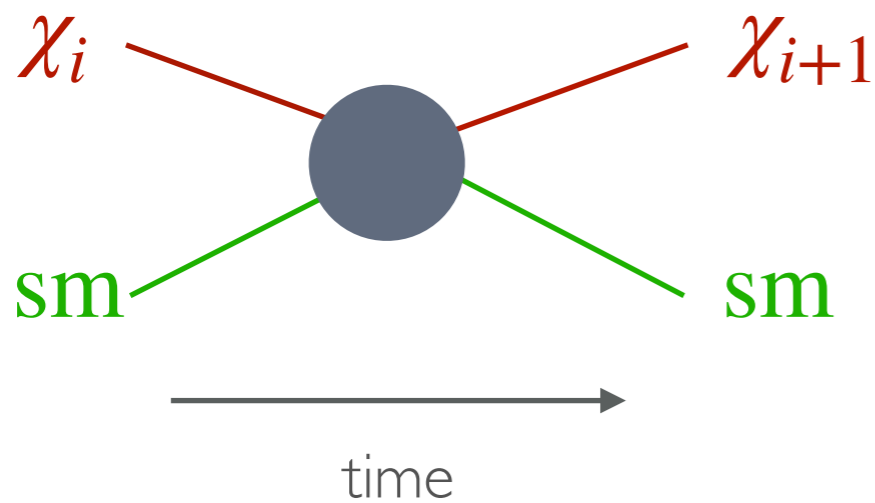
DM freezes out

Most DM settles away from the cliff

DM density

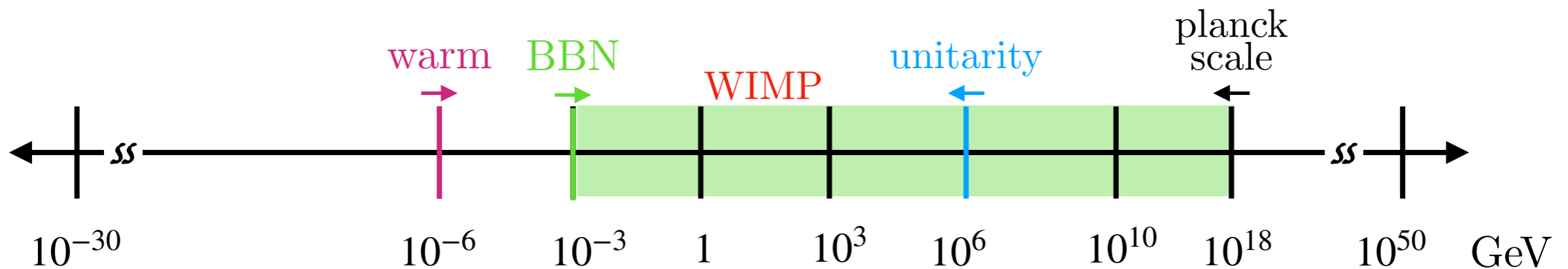


Ex. 3: Chain DM



$$\langle \sigma_{\text{chain} \nu} \rangle = \frac{\alpha^2}{m_{\text{DM}}^2}$$

$$m_{\text{DM}} \simeq \alpha^2 \times 10^{16} \text{ GeV}$$



Outlook

- Lots of activity for thermal dark matter.
- Many different interactions, processes, and their relative importance throughout the cosmological history.
- Novel dark matter frameworks.
- Generic.
- Lots of discovery potential for experiments.
- Much more to do.

Thank you!