

# A negative view on the waveform analysis

Alexander Deisting

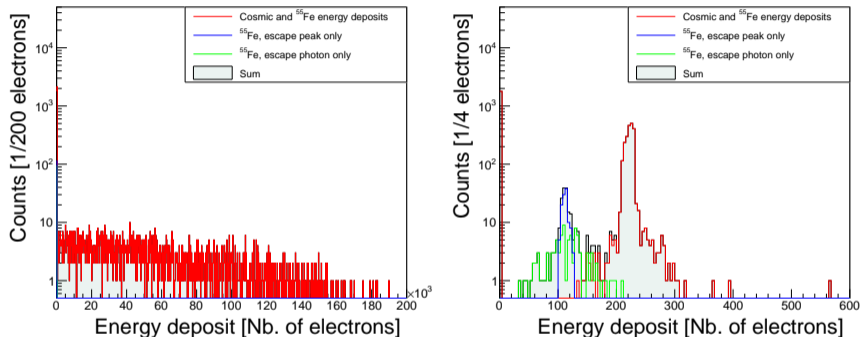


8<sup>th</sup> of November, 2019

## These slides:

- ▶ As Ed showed in his triptych<sup>TM</sup> code, there are quite substantial negative pulses
- ▶ Also the nice anode 1/2 spectra in the waveform RMS analysis – as opposed to the non existing spectra in the anode 1/2 amplitude analysis, showed that there is something going on that our code focusing on positive polarity signals does not catch
- ▶ I put in the negative pulse analysis, but since this involved changing many things, I wanted to make sure:
  - i The actual positive polarity analysis is still working
  - ii The new negative polarity analysis is working
  - iii Bonus: See what effect having negative polarity signals has on the positive polarity pulse analysis
- ▶ To do so I went back to a set of known signals: MC simulation of 11.000 pulses with positive and negative polarity each

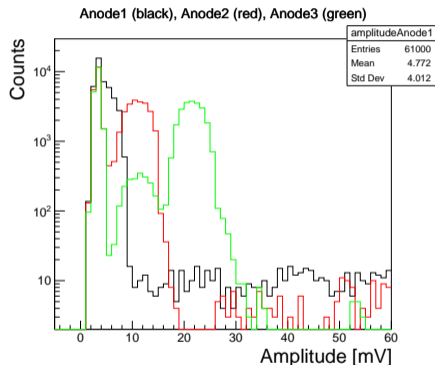
# Reminder: Input spectrum, $^{55}\text{Fe}$ & cosmics [A. Deisting, week 43 slides]



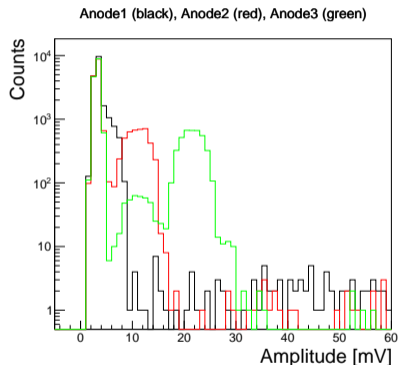
**Figure:** Expected energy deposits of  $^{55}\text{Fe}$  decay radiation and cosmic muons inside a gas volume filled with Ar-CO<sub>2</sub> (98-2). This is the result of a toy Monte-Carlo using the approximate layout of a quarter of the HPTPC, the heed package in Garfield++, and information from the ESTAR and XCOM databases. The right plot is a zoomed view of the left plot.

# Amplitude spectra, all pulses

Previously shown:



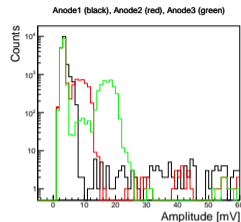
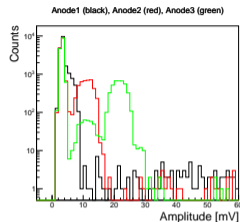
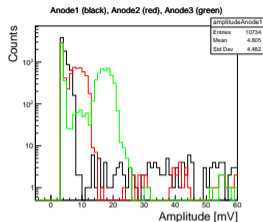
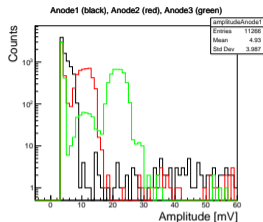
Current analysis:



Noise periods:  $10 \times 10^{-6} \mu\text{s}$  and  $5 \times 10^{-6} \mu\text{s}$

► Seems as if the standard analysis is not broken

# Amplitude spectra



positive amplitude  $>$   
negative amplitudes

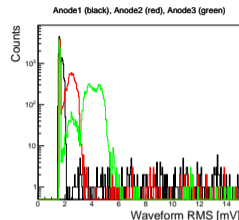
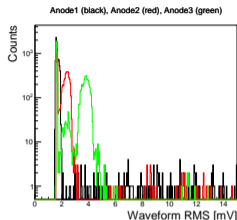
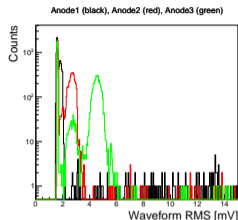
negative amplitude  $>$   
positive amplitudes

all positive amplitudes

all negative amplitudes

- ▶ The negative (positive) amplitude pulses mainly contribute to the noise peak when one looks only for positive (negative) pulses
- ▶ Even while adding noise, and mixing positive and negative amplitude spectra, the input distributions are recovered when requiring a simple cut as e.g. “positive amplitude  $>$  negative amplitudes”

# Waveform RMS



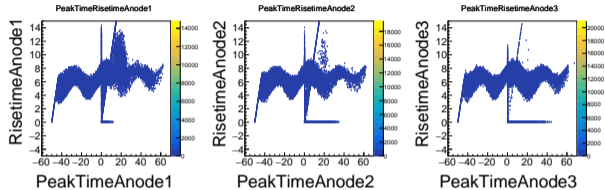
positive amplitude  $>$  negative  
amplitudes

negative amplitude  $>$  positive  
amplitudes

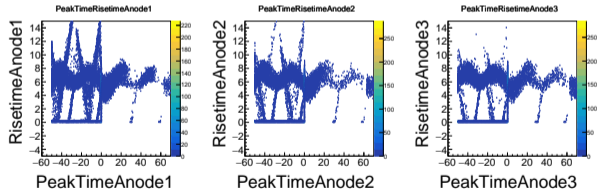
all amplitudes

- ▶ As observed previously: The waveform RMS spectra seem to have better energy resolution when the amplitude spectra
- ▶ A reasonable distribution is recovered while requiring a simple cut as e.g. “positive amplitude  $>$  negative amplitudes” during the analysis of data containing positive and negative polarity pulses.
- ▶ Broader/double peaks appear in the RMS spectrum when no amplitude is selected

Previously shown:



Current analysis:



## Summary:

---

- ▶ Amplitude analysis works for negative, positive and mixtures of negative and positive amplitudes
- ▶ The time analysis still works for positive polarity pulses ...
- ▶ ... however in case of the negative polarity pulses, there still seems to be a bug

(I intended to run this on data as well for today, but I had some rapTORR issues. To be worked out.)



# Backup

## Anode parameter overview, positive polarity pulses

Parameter	anode 1	anode 2	anode 3
channel ID	1	2	3
<i>gain</i>	"CR112"	"CR112"	"CR112"
$V_{\min}$ [mV]	-200.0	-200.0	-200.0
$V_{\max}$ [mV]	1800.0	1800.0	1800.0
preamp clipping value [mV]	1600.0	1600.0	1600.0
DC offset [mV]	-20	30	-12
<i>trigger threshold</i> [mV]	20	20	20
<i>trigger polarity falling</i>	false	false	false
<i>pulse polarity falling</i>	false	false	false
<i>channelName</i>	"Anode_1"	"Anode_2"	"Anode_3"
$G_{\text{preamp}}$	12	12	12
$G_{\text{meshes}}$	10000	25000	50000

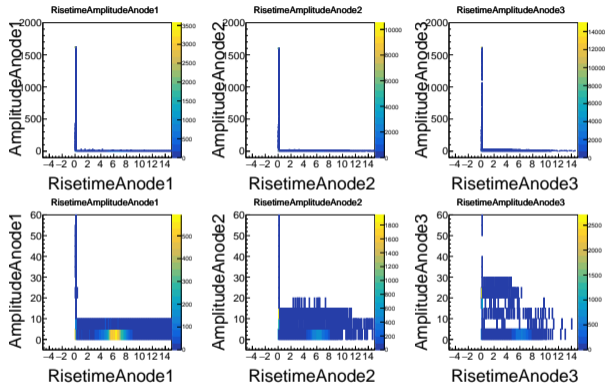
Parameters in italics are currently not used for anything.

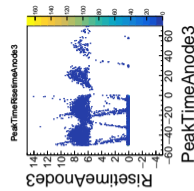
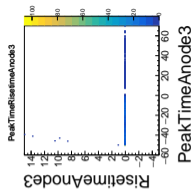
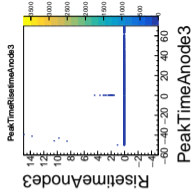
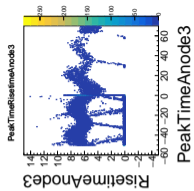
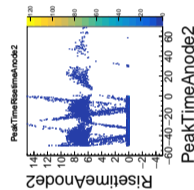
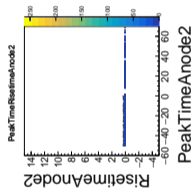
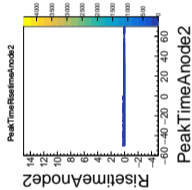
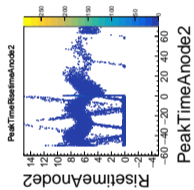
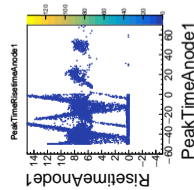
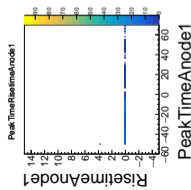
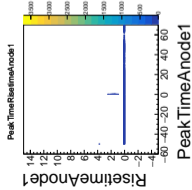
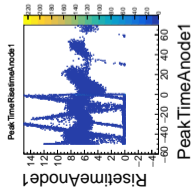
## Anode parameter overview, negative polarity pulses

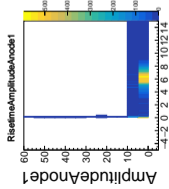
Parameter	anode 1	anode 2	anode 3
channel ID	1	2	3
<i>gain</i>	"CR112"	"CR112"	"CR112"
$V_{\min}$ [mV]	-200.0	-200.0	-200.0
$V_{\max}$ [mV]	1800.0	1800.0	1800.0
preamp clipping value [mV]	1600.0	1600.0	1600.0
DC offset [mV]	-20	30	-12
<i>trigger threshold</i> [mV]	20	20	20
<i>trigger polarity falling</i>	false	false	false
<i>pulse polarity falling</i>	false	false	false
<i>channelName</i>	"Anode_1"	"Anode_2"	"Anode_3"
$G_{\text{preamp}}$	12	12	12
$G_{\text{meshes}}$	8000	20000	40000

Parameters in italics are currently not used for anything.

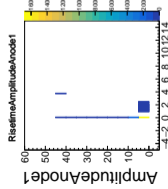
# Time checks



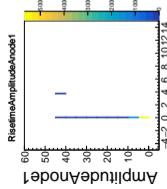




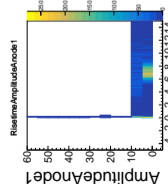
RisettimeAnode1



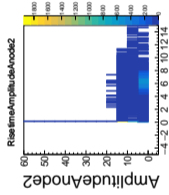
RisettimeAnode1



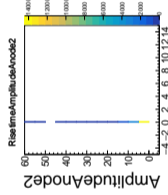
RisettimeAnode1



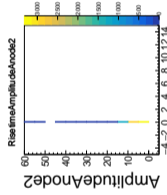
RisettimeAnode1



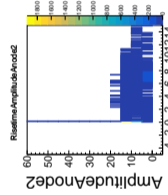
RisettimeAnode2



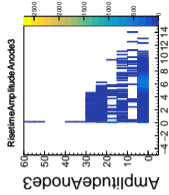
RisettimeAnode2



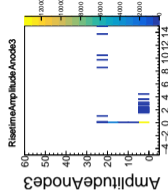
RisettimeAnode2



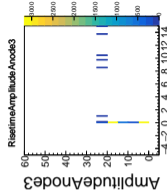
RisettimeAnode2



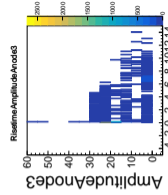
RisettimeAnode3



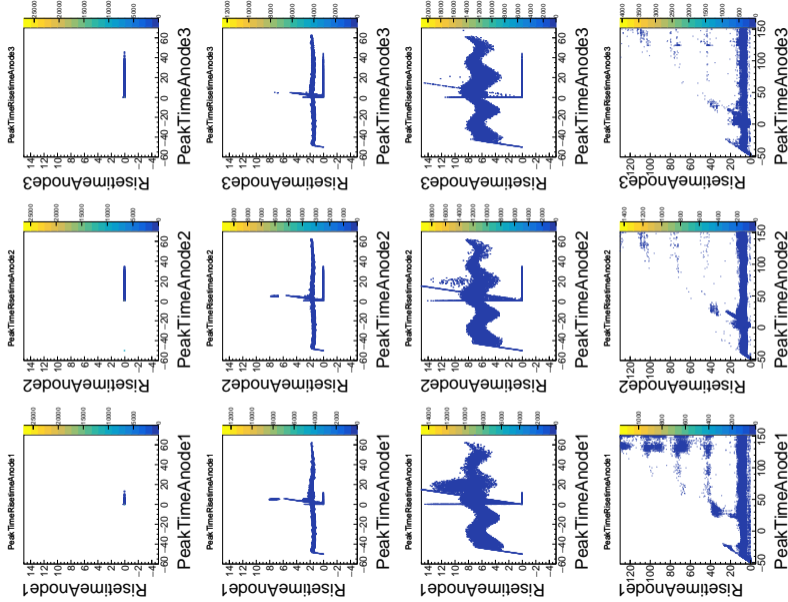
RisettimeAnode3



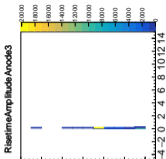
RisettimeAnode3



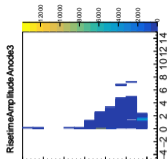
RisettimeAnode3



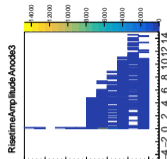




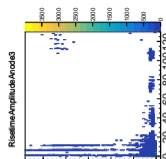
RisetimeAnode3



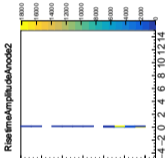
RisetimeAnode3



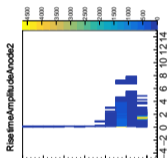
RisetimeAnode3



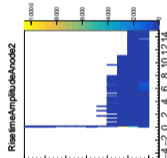
RisetimeAnode3



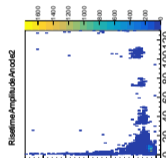
RisetimeAnode2



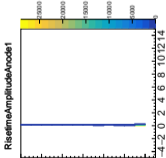
RisetimeAnode2



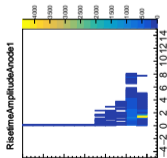
RisetimeAnode2



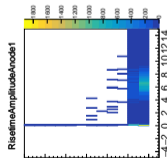
RisetimeAnode2



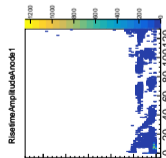
RisetimeAnode1



RisetimeAnode1



RisetimeAnode1



RisetimeAnode1

## ToDo's

Week 46 – Waveform RMS: Data ( $^{55}\text{Fe}$ ) vs background (no source)

## Closing remarks

---

- ▶ There is a good indication that the peaks in the RMS are due to the source
- ▶ The nice spectra from anode 2 and anode 1 may result from negative polarity signals  
⇒  Check for negative signals in the analysis code
- ▶ Amplitude and RMS correlation looks as if we could calibrate that
- ▶ I don't have yet Ed's triptych code to check the different regions as well  
⇒  Do triptych's for the three signal regions
- Still pending: sin-fit correction on pre-selected waveforms

# ToDo's

## Week 36 – Waveform analysis

- ▶ A fraction of the August data - pure Ar and Ar-CO<sub>2</sub> at various pressures has been analysed using exponential smoothing as well as a `sin` fit to subtract periodic noise

### On the noise-fit:

- Check more advanced fit-functions than just a plane `sin` fit
  - Check whether the fit range can be changed to improve the overall results
  - Do detailed  $\chi^2$  cuts and checks with the current data set to see where the fit improves the analysis
- ⇒ A more detailed presentation will follow

# ToDo's

## Week 30 – Hardware report

## General HPTPC news:

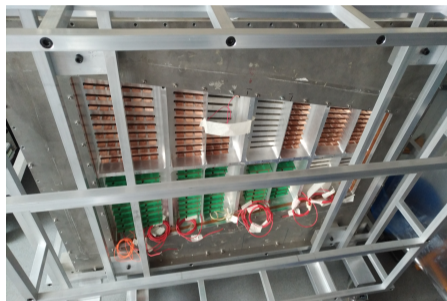
- ▶ Over the last week-end we took data until the DAQ decided to not take data anymore. (This cost us the highest voltage setting)
- ▶ On Monday we had once more Imperial man-power  $\Rightarrow$  See the slides on the OROC holder test
- ▶ Evacuation over the rest of Monday, and the full Tuesday
- ▶ Right now we are back to about  $\text{Ar-CO}_2$  (88-12) at 1 atm
- ▶ Data taking at these voltages is still pending

## Coming up:

- ▶ **Power outage in the HPTPC lab on Mon, T111 probaly more than one day**
- ☑ Couple test pulses in one of the meshes and readout the other preamplifiers
- ☑ Look into changing the T111 and T133 clean room filters

# OROC updated

- ▶ All shortening cards are in place
- ▶ The OROC is back in its box and awaits testing
- ▶ Everything needed for the HV distribution boxes I have in hand
- ▶ The following list of OROC ToDo's applies:
  - Put cooper shims
  - Do the HV distribution network
  - HV tests in air
  - Pulser test in air
  - Construction of the field line termination plane
- ▶ We ordered pre-mixed Ar-CO<sub>2</sub> (85-15) – as soon as this arrives and the tests in air are done, we will start flushing



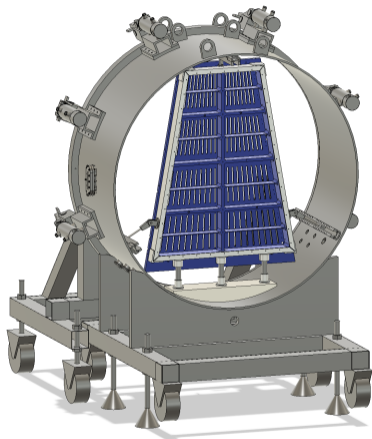


## OROC holder tests

- ▶ The production of the OROC holder for the vessel is progressing well
- ▶ In the pictures you see a fitting test to the vessel
- ▶ The full holder should be completed in the next weeks

## Readout Electronics update

- Two digitiser board with the APV assembled
- Components for the attenuation/protection board arrived at Imperial and the assembly is under-way
- First iteration of the firmware written (@Imperial)
- Felix integration tests with the digitiser board only (@Imperial)
- Commissioning with a small detector ...



## Bias T sparking

- ▶ At about 3.5 kV on anode 3 sparking in the anode 3 bias T occurred
- ▶ Previously higher voltages had been achieved, all components in the anode 3 bias T are rated for high voltage
- ▶ It turns out that the *signal-to-preamp* (which should be at ground potential) leg of bias T discharged against the nearest ground
- ▶ The same could be observed at the anode 2 bias T for high voltages
- ▶ As a measure the anode 3 *signal-to-preamp* leg has been connected to ground via a 3 M $\Omega$  resistor to provide a path to ground in case there is some charging up
- ▶ We observed once more discharges in a bias T, but could not determine whether it was in the anode 3 one
- We'll keep an eye....
- Check that there are no effects on the signal readout with the extra resistor

- ▶ All shortening cards which we could put, are in place
- ▶ The OROC is back in its box and awaits testing
- ▶ We have the aircon again running in the MWPC and MPGD lab
- ▶ Copper shims should arrive tomorrow or next week
- ▶ A slightly reduced list of the usual OROC ToDo's still applies:
  - ✓ Put cooper shims
  - ✓ Do the HV distribution network
  - ✓ HV tests in air
  - ✓ Pulser test in air
  - Construction of the field line termination plane

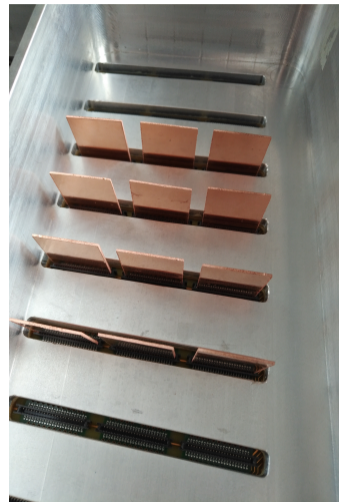
# ToDo's

## Hardware report week 29

## A word on high frequency noise

- ▶ There is 1.35 MHz noise in the data
- ▶ However: It turns out that is only there at certain instances and permanently, uncorrelated with changes to the HPTPC → Possibly something else in the building is responsible
- ▶ In other news: The signal from the preamp at anode 2 looks different, because the evaluation board there has a different capacitor
- ▶ Hence:
  - ✓ Test the response of the modified preamp to test pulses on the test input
  - ☒ Do the same, but with pulses coupled into the real input using a capacitor
  - ☒ Drive the preamps into saturation (check first on the old scope as a safety measure)
  - ☒ Do the same test with a non modified preamp
  - ✓ Drive **only** standard preamps into saturation (check first on the old scope as a safety measure)

- ▶ The first samples for the copper shims arrived, they fit and the rest is ordered
- ▶ Annora measured dust counts and I did a clean of the floor in T111 – We will see whether this helped
- ▶ The next OROC step:
  - ✓ Put shortening cards (copper shims)
  - ✓ Mount the OROC back to the lid of the test box (copper shims can also be placed after this)
  - ✓ Do the HV distribution network
  - ✓ HV tests in air
  - ✓ Pulser test in air
  - Construction of the field line termination plane



# ToDo's

## Hardware report week 26

- ▶ Jocelyn fixed the burst disk:
  - ▶ We discovered where was a puncture in the burst disk
  - ▶ It was replaced with a 5 barG burst disk
- ▶  $^{55}\text{Fe}$  is in the vessel now → We will explore with one fill if we can see a peak related to this source
- Filling the vessel again
- Decision on a gas mixture / pressure for a diffusion measurement will be made when we have first insights from the light gain analysis
- Talk to FIKE what they think is the best solution for our use case (possibly another pressure relive valve)



# ToDo's

## Waveform report week 25

# Update on the toy Monte-Carlo to model the energy deposited in the detector

- ▶ Using the known decay energies the  $^{241}\text{Am}$  sources ( $\alpha$ ,  $\gamma$ ) and the  $^{137}\text{Cs}$  ( $\beta$ ,  $\gamma$ ) decay energy spectra are modelled
  - ▶  $^{137}\text{Cs}$   $\beta$ -spectrum: Based on interpolated IAEA data
  - ▶  $\gamma$ - and  $\alpha$ -spectra: Approximated using Gaußians with an arbitrary width
- ▶ The  $\gamma$  absorption as well as charged particle ranges are extracted from ESTAR, PSTAR and XCOM
- ▶  $\gamma$ s and  $\alpha$ s are assumed to deposit their full energy in the detector
- ☑ Currently cosmic  $\mu$  are added
- ☑ Garfield/heed will be used for the energy deposit of charged particles
- ☑ Furthermore the detector geometry is being put in
- ☑  $^{55}\text{Fe}$  is added