

# Muon $g - 2$ and the $B$ -physics anomalies in RPV supersymmetry and the discovery prospect at LHC and future colliders

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arXiv:2106.15647

SUSY 2021

Washington University in St. Louis

August 27, 2021

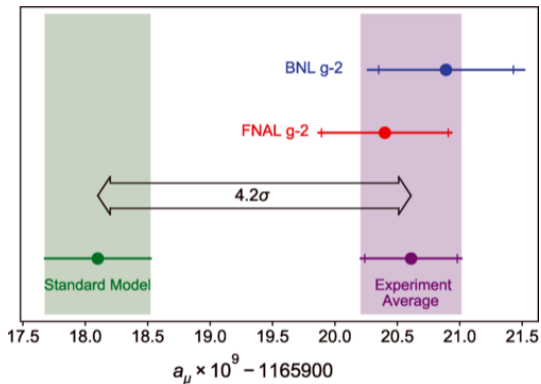
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# Motivation

- The recent experimental results of muon  $g - 2$  (from the Fermilab) and the lepton flavor universality violation in rare B-meson decays (from the LHCb, Belle, BaBar) could be the hints ( $> 3\sigma$  anomalies) of new physics beyond the Standard Model.
- Under the minimal RPV supersymmetric framework, assuming the mass of third generation sfermions lighter than the other two generations (called "RPV3", [Altmannshofer, Dev, Soni \(PRD 2017\)](#)), muon  $g - 2$  and the  $B$ -physics anomalies could be addressed simultaneously and also could be tested at LHC and beyond.

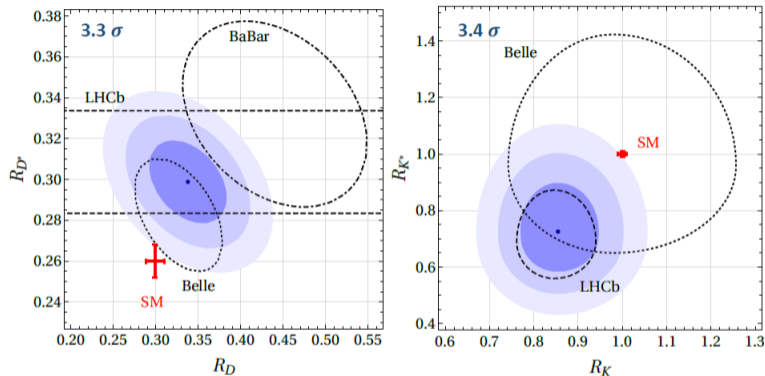
# muon $g - 2$ anomaly

- $\Delta a_\mu = a_\mu^{\text{exp}} - a_\mu^{\text{SM}} = (251 \pm 59) \times 10^{-11}$  has a significance of  $4.2\sigma$ .
- Could be the signal of new physics beyond the SM where some new couplings to muon could be detectable by LHC or future colliders.



B. Abi et al. (PRL 2021)

# B-physics anomalies



Altmannshofer, Dev, Soni, Sui (PRD 2020)

- $R_{D^{(*)}} = \frac{\text{BR}(B \rightarrow D^{(*)} \tau \nu)}{\text{BR}(B \rightarrow D^{(*)} \ell \nu)}$  (with  $\ell = e, \mu$ ),  $R_{K^{(*)}} = \frac{\text{BR}(B \rightarrow K^{(*)} \mu^+ \mu^-)}{\text{BR}(B \rightarrow K^{(*)} e^+ e^-)}$
- Also imply possible new couplings to leptons.

# Explanation of anomalies in RPV3 SUSY

- The  $LQD$  and  $LLE$  part of the RPV SUSY Lagrangian which contains the  $\lambda'$  and  $\lambda$  couplings respectively and are relevant for the  $R_{D^{(*)}}$ ,  $R_{K^{(*)}}$  and  $(g-2)_\mu$  anomalies.

$$\begin{aligned} \mathcal{L}_{LQD} = & \lambda'_{ijk} (\tilde{\nu}_{iL} \bar{d}_{kR} d_{jL} + \tilde{d}_{jL} \bar{d}_{kR} \nu_{iL} + \tilde{d}_{kR}^* \bar{\nu}_{iL}^c d_{jL} \\ & - \tilde{e}_{iL} \bar{d}_{kR} u_{jL} - \tilde{u}_{jL} \bar{d}_{kR} e_{iL} - \tilde{d}_{kR}^* \bar{e}_{iL}^c u_{jL}) + \text{H.c.} \end{aligned} \quad (1)$$

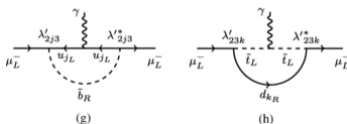
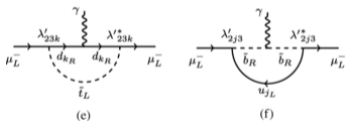
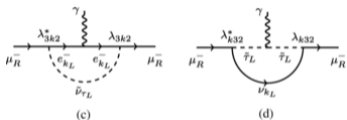
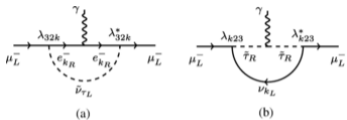
$$\mathcal{L}_{LLE} = \frac{1}{2} \lambda_{ijk} [\tilde{\nu}_{iL} \bar{e}_{kR} e_{jL} + \tilde{e}_{jL} \bar{e}_{kR} \nu_{iL} + \tilde{e}_{kR}^* \bar{\nu}_{iL}^c e_{jL} - (i \leftrightarrow j)] + \text{H.c.} \quad (2)$$

- Following previous discussions ([Kim, Kyaee, Lee \(PLB 2001\)](#); [Altmannshofer, Dev, Soni, Sui \(PRD 2020\)](#)), in RPV3 framework,  $(g-2)_\mu$  correction can be written as:

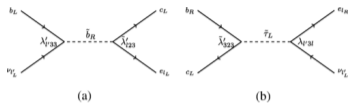
$$\Delta a_\mu = \frac{m_\mu^2}{96\pi^2} \sum_{k=1}^3 \left( \frac{2(|\lambda_{32k}|^2 + |\lambda_{3k2}|^2)}{m_{\tilde{\nu}_\tau}^2} - \frac{|\lambda_{3k2}|^2}{m_{\tilde{\tau}_L}^2} - \frac{|\lambda_{k23}|^2}{m_{\tilde{\tau}_R}^2} + \frac{3|\lambda'_{2k3}|^2}{m_{\tilde{b}_R}^2} \right) \quad (3)$$

# Explanation of anomalies in RPV3 SUSY

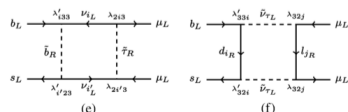
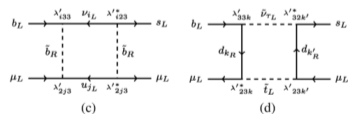
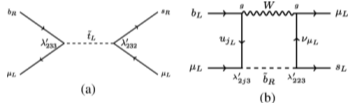
$(g - 2)_\mu$  Kim, Kyae, Lee (PLB 2001)



$R_{D^{(*)}}$  Deshpande, He (EPJC 2017); Altmannshofer, Dev, Soni (PRD 2017) etc.



$R_{K^{(*)}}$  Das, Hati, Kumar, Mahajan (PRD 2017); Trifinopoulos (EPJC 2018) etc.



# Parameter space

- Parameters  $(\lambda_{232}, \lambda'_{233}, \lambda'_{223}, \lambda'_{232}, m_{\tilde{b}_R}, m_{\tilde{b}_L}, m_{\tilde{\nu}_\tau}, m_{\tilde{\tau}_L})$ 
  - $\lambda_{232} = -\lambda_{322} \neq 0 \Leftrightarrow$  contribute to muon  $g - 2$ , other  $\lambda_{3ij}$  couplings cannot be large at the same time due to the constraints of  $\tau \rightarrow \mu\mu\mu$ ,  $\mu \rightarrow e\gamma$  etc.
  - $\lambda'_{2ij} \neq 0 \Leftrightarrow$  include  $\mu$  and free of  $m_{\tilde{\nu}_\tau}$ .
  - $\lambda'_{3ij} = 0$ , otherwise combined with  $\lambda_{32k}$  or  $\lambda_{3k2}$ , well measured meson decays  $(\bar{d}_i d_j) \rightarrow \mu \ell_k$  or  $\tau \rightarrow \mu K$  and  $\tau \rightarrow \mu \eta$  decays will prevent  $\lambda'_{3ij}$  to be large.
  - $m_{\tilde{\tau}_R}$  not involved with this choice of couplings.
  - $m_{\tilde{t}_L}$  can only influence  $\text{BR}(B_s \rightarrow \mu^+ \mu^-)$  and the Wilson coefficients  $(C'_9)^\mu$  and  $(C'_{10})^\mu$  that describe the  $R_{K^{(*)}}$  anomaly. But we can assume a relatively larger value to make the influence small.



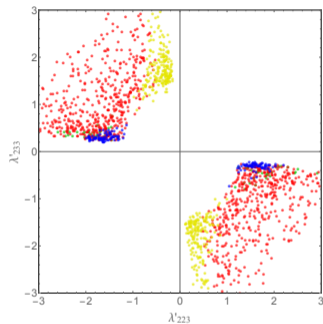
# Parameter space

- 8-D parameter space  $(\lambda_{232}, \lambda'_{233}, \lambda'_{223}, \lambda'_{232}, m_{\tilde{b}_R}, m_{\tilde{b}_L}, m_{\tilde{\nu}_\tau}, m_{\tilde{\tau}_L})$ 
  - $m_{\tilde{b}_R} = m_{\tilde{b}_L}$  for simplicity.
  - $m_{\tilde{\tau}_L}$  has opposite contribution for  $(g-2)_\mu$ . The influence is not important as long as  $m_{\tilde{\tau}_L} \gtrsim O(2 \text{ TeV})$ . Here we choose 4 TeV.
- $\Rightarrow$  6-D parameter space  $(\lambda_{232}, \lambda'_{233}, \lambda'_{223}, \lambda'_{232}, m_{\tilde{b}}, m_{\tilde{\nu}_\tau})$
- In a sense,  $(\lambda', m_{\tilde{b}})$  and  $(\lambda, m_{\tilde{\nu}_\tau})$  are orthogonal in our scenario since  $(\lambda, m_{\tilde{\nu}_\tau})$  can only influence  $(g-2)_\mu$  anomaly and 4-lepton constraint while on the other hand,  $(\lambda', m_{\tilde{b}})$  can only influence  $R_{D^{(*)}}$  and  $R_{K^{(*)}}$  anomalies and other constraints (The influence to  $(g-2)_\mu$  is very small because  $m_{\tilde{b}}^2 \gg m_{\tilde{\nu}_\tau}^2$  as we will see from Fig(d)(e)(g)).
- So, we can plot the constraints and anomalies in two 2-D spaces:  $(\lambda', m_{\tilde{b}})$  and  $(\lambda, m_{\tilde{\nu}_\tau})$

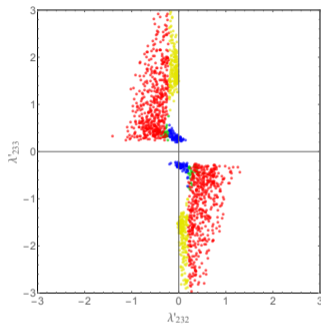
# Numerical scan

- We scan the 6-D parameter space  
( $\lambda_{232}, \lambda'_{233}, \lambda'_{223}, \lambda'_{232}, m_{\tilde{b}_R} = m_{\tilde{b}_L}, m_{\tilde{\nu}_\tau}, m_{\tilde{\tau}_L} = 4 \text{ TeV}$ )
- $m_{\tilde{\nu}_\tau} \in [0.7, 1.2] \text{ TeV}$  (also tried  $m_{\tilde{\nu}_\tau} \in [1.2, 3] \text{ TeV}$  but no solution found in this region)
- $|\lambda_{232}| \in [2.5, 3.5]$  (also tried  $|\lambda_{232}| \in [1, 2.5]$  but no solution found in this region)
- $m_{\tilde{b}} \in [1.2, 10] \text{ TeV}$
- $|\lambda'_{233}| \in [0.01, 3]$
- $|\lambda'_{223}| \in [0.01, 3]$
- $|\lambda'_{232}| \in [0.01, 3]$
- 30 million attempts  $\Rightarrow$  1570 solutions (red+yellow+blue+green points)

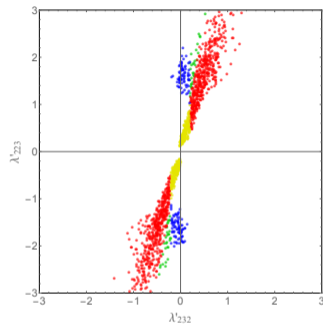
# Numerical scan



(a)



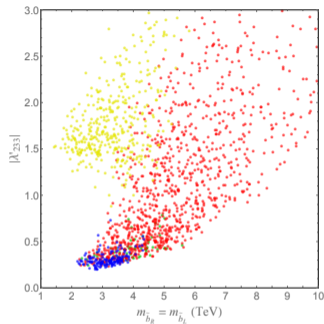
(b)



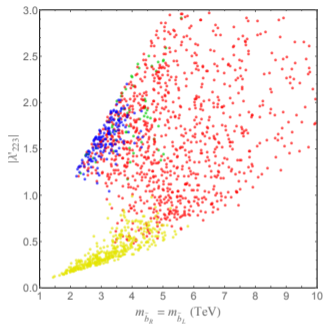
(c)

- Separate to 3 characteristically regions according to Fig(c). Yellow:  $|\lambda'_{232}| < 0.2$  and  $|\lambda'_{223}| < 1$ ;  
Blue:  $|\lambda'_{232}| < 0.2$  and  $|\lambda'_{223}| > 1$ ; Red:  $|\lambda'_{232}| > 0.2$  and  $1.5 < \frac{\lambda'_{223}}{\lambda'_{232}} < 5.5$
- From Fig(c), Red:  $\frac{\lambda'_{223}}{\lambda'_{232}} \sim 3 \Leftarrow B_s - \bar{B}_s$  mixing. Green: crossover region from Red to Blue.
- Yellow+Blue:  $|\lambda'_{232}|$  small or even zero. Fig(a-c):  $|\lambda'_{233} \lambda'_{223}|$  small  $\Leftarrow B_s - \bar{B}_s$  mixing.

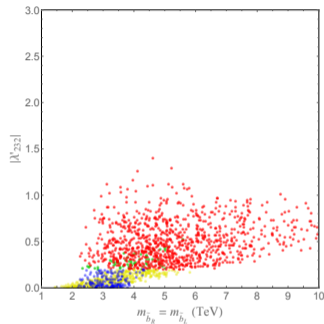
# Numerical scan



(d)



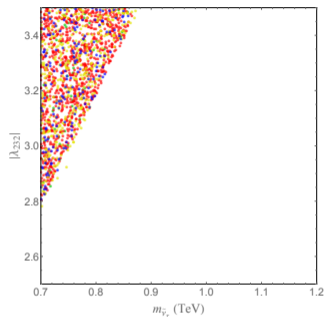
(e)



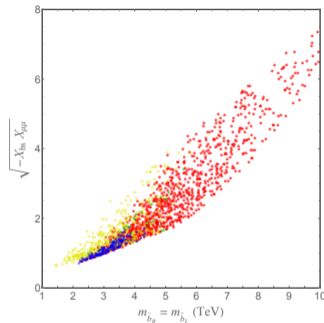
(f)

- Fig(a-c): the absolute sign of  $\lambda'$  not important, only the relative sign matters.
- Fig(e):  $\frac{|\lambda'_{223}|}{(m_{\bar{b}}/1 \text{ TeV})} \lesssim 0.57 \Leftrightarrow D^0 \rightarrow \mu^+ \mu^-$ ; Fig(d):  $\frac{|\lambda'_{233}|}{(m_{\bar{b}}/1 \text{ TeV})} \lesssim 1.0$
- $\Rightarrow$  cannot contribute to  $(g-2)_\mu$  much.
- Fig(h):  $|\lambda'_{232}| \lesssim 1.5 \Leftrightarrow$  from Fig(c),  $|\lambda'_{232}|$  is either small or  $\sim |\lambda'_{223}|/3$

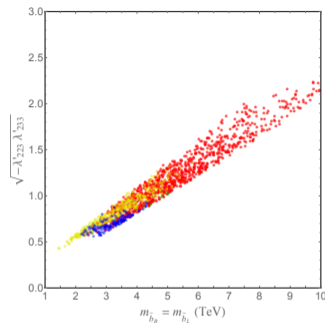
# Numerical scan



(g)



(h)



(i)

- Fig(g): Red, Yellow, Blue and Green points are totally mixed  $\Leftarrow$  Orthogonality of the two 2-D subspaces:  $(\lambda, m_{\tilde{\nu}_\tau})$  and  $(\lambda', m_{\tilde{b}})$ .
- Fig(h)  $\Leftarrow R_{K^{(*)}} \Leftarrow \lambda'_{233} \lambda'_{223} < 0 \Leftarrow$  Fig(a)
- Fig(i):  $\frac{\sqrt{-\lambda'_{223} \lambda'_{233}}}{(m_{\tilde{b}}/1 \text{ TeV})} \sim (0.2, 0.28) \Leftarrow B \rightarrow K \nu \bar{\nu}$

# Numerical scan

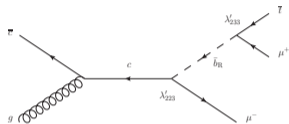
- Fig(g)  $\Rightarrow |\lambda_{232}| \gtrsim 2.78$
- Fig(g)  $\Rightarrow 0.70 \text{ TeV} \lesssim m_{\tilde{\nu}_\tau} \lesssim 0.87 \text{ TeV}$
- Fig(d)  $\Rightarrow |\lambda'_{233}| \gtrsim 0.20$
- Fig(e)  $\Rightarrow |\lambda'_{223}| \gtrsim 0.12$
- Fig(f)  $\Rightarrow |\lambda'_{232}|$  could be very small or even zero
- Fig(d-f)  $\Rightarrow m_{\tilde{b}} \gtrsim 1.44 \text{ TeV}$
- Fig(g)  $\Rightarrow \frac{|\lambda_{232}|}{(m_{\tilde{\nu}_\tau}/1 \text{ TeV})} \gtrsim 4$ ; Fig(d)  $\Rightarrow \frac{|\lambda'_{233}|}{(m_{\tilde{b}_R}/1 \text{ TeV})} \lesssim 1$ ; Fig(e)  
 $\Rightarrow \frac{|\lambda'_{223}|}{(m_{\tilde{b}_R}/1 \text{ TeV})} \lesssim 0.57$ . This means that the sneutrino term gives the main contribution of muon (g-2) as one can see from Eq.(3).

# Benchmark scenarios

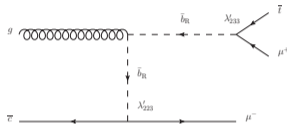
- Density/number of the points in some region  $\propto$  the size of the allowed region of the parameter space.
- Three benchmark scenarios:
  - Red scenario (a subset of the Red region in the scatter plot):  
 $\lambda'_{233} = -\lambda'_{223} = -3\lambda'_{232}$ ,  $m_{\tilde{b}_R} = m_{\tilde{b}_L}$ ,  $m_{\tilde{\tau}_L} = 4$  TeV. Fig(c): choose  $\lambda'_{223} = -3\lambda'_{232}$  to collect as many red points as possible.
  - Yellow scenario (a subset of the Yellow region in the scatter plot):  
 $\lambda'_{233} = -8\lambda'_{223}$ ,  $\lambda'_{232} = 0$ ,  $m_{\tilde{\tau}_L} = 4$  TeV. Fig(a):  $\lambda'_{233} = -8\lambda'_{223}$  to collect as many yellow points as possible.
  - Blue scenario (a subset of the Blue region in the scatter plot):  
 $\lambda'_{223} = -6\lambda'_{233}$ ,  $\lambda'_{232} = 0$ ,  $m_{\tilde{\tau}_L} = 4$  TeV. Fig(a):  $\lambda'_{223} = -6\lambda'_{233}$  to collect as many blue points as possible.
- \* Red scenario,  $\lambda'_{233} = -\lambda'_{223}$  for simplicity.
- \* Yellow & Blue scenario,  $\lambda'_{232} = 0$  for simplicity  $\Rightarrow m_{\tilde{b}_L}, m_{\tilde{t}_L}$  will not appear.

# Collider signals

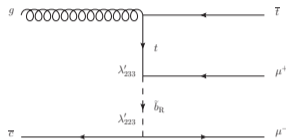
- Signal for  $(\lambda', m_{\tilde{b}})$  space:  $pp \rightarrow \bar{t}\mu^+\mu^-$



(j)



(k)



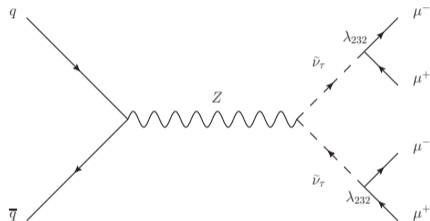
(l)

- $p_T^{t,\mu} > 20 \text{ GeV}$ ,  $|\eta^{t,\mu}| < 2.5$ ,  $\Delta R^{\mu\mu} > 0.4$  and  $\Delta R^{t\mu} > 0.4$ ,  $M_{\mu^+\mu^-} > 0.4 \text{ TeV}$ . Assume  $\mathcal{L} = 3000 \text{ fb}^{-1}$ .  $\sqrt{s} = 14 \text{ TeV}, 27 \text{ TeV}, 100 \text{ TeV}$ .
- Background small.  $pp \rightarrow \bar{t}\mu^+\mu^- X$  (with  $X = j, b, W^+ \rightarrow jj, W^+ \rightarrow \ell^+\nu_\ell$  not detected:  $p_T^{j,b,\ell} < 20 \text{ GeV}$ ,  $E_T^{\text{miss}} < 20 \text{ GeV}$ )
- $pp \rightarrow t\mu^+\mu^-$  is similar but with a larger background.
- Only  $\lambda'_{233}$ ,  $\lambda'_{223}$  and  $m_{\tilde{b}_R}$  contribute to the signal. What can be probed are actually these parameters, a projection of the scenario.



# Collider signals

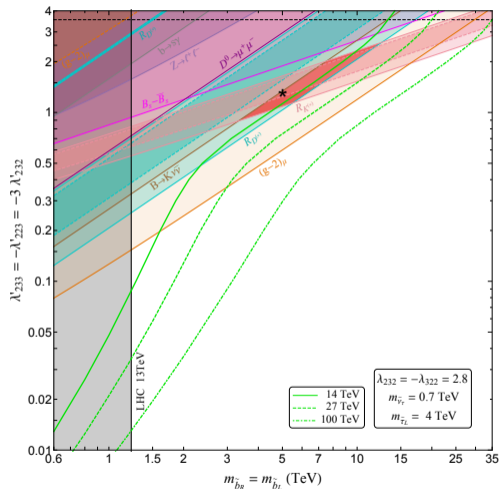
- Signal for  $(\lambda, m_{\tilde{\nu}_\tau})$  space:  $pp \rightarrow \mu^+ \mu^- \mu^+ \mu^-$



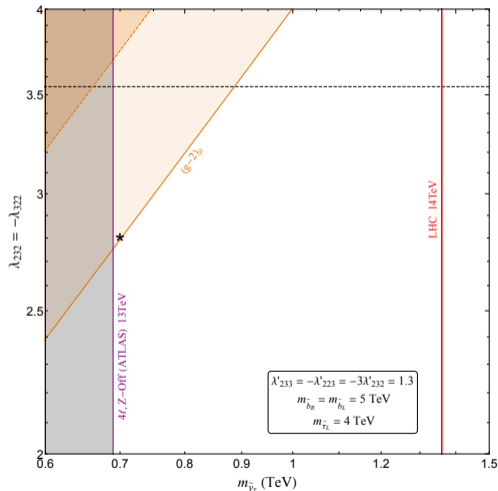
(m)

- $p_T^\mu > 25$  GeV,  $|\eta^\mu| < 2.47$ ,  $\Delta R^{\mu\mu} > 0.2$ ,  $M_{\mu^+\mu^-} > 0.4$  TeV.
- Assume the mass of the lightest neutralino is 100 GeV for the calculation of the branching ratio of  $\tilde{\nu}_\tau$ .  $\text{BR}(\tilde{\nu}_\tau \rightarrow \mu^+ \mu^-)$  is larger than 95% when  $|\lambda_{232}| > 1.2$

# Anomalies and constraints (Red scenario)

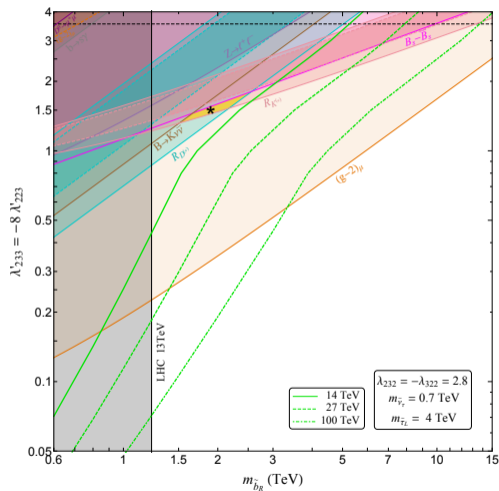


(n)

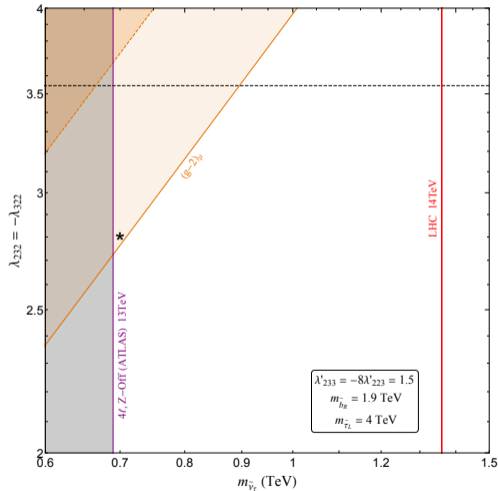


(o)

# Anomalies and constraints (Yellow scenario)

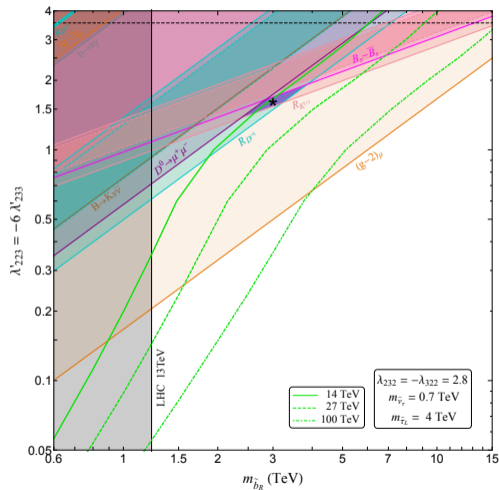


(p)

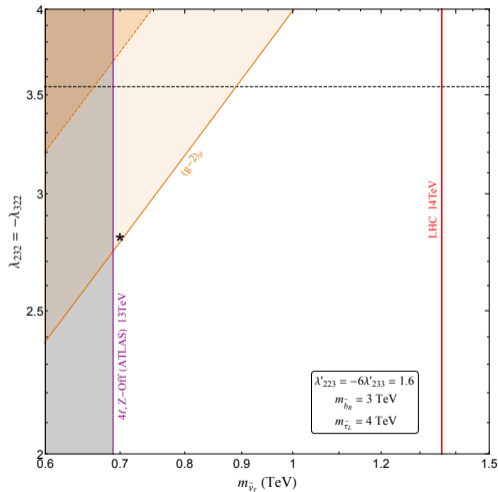


(q)

# Anomalies and constraints (Blue scenario)



(r)



(s)

# Discussions

- The figure on the left uses the value of the black star in the figure on the right and vice versa.
- The cyan, pink and orange shaded regions with solid (dashed) boundaries explain the  $R_{D^{(*)}}$ ,  $R_{K^{(*)}}$  and  $(g-2)_\mu$  anomalies at  $3\sigma$  ( $2\sigma$ ) CL respectively.
- The red, yellow and blue shaded regions are the overlap regions that simultaneously explain all the three anomalies correspond to the red, yellow and blue scenarios.
- The green solid, dashed and dot-dashed lines in Fig(n)(p)(r) are the  $2\sigma$  sensitivities of the  $\sqrt{s} = 14$  TeV, 27 TeV and 100 TeV  $pp$  colliders in the  $\bar{t}\mu^+\mu^-$  channel.
- These green curves bend downward at large  $\lambda'$  region because of the off-shell contribution of  $pp \rightarrow \bar{t}\mu^+\mu^-$  Fig(l).
- The red solid line in Fig(o)(q)(s) are the  $2\sigma$  sensitivities of the LHC 14 TeV in the 4-muon channel.
- Fig(o)(q)(s) are quite similar. Consistent with Fig(g). Consequence of the orthogonality of the two subspaces  $(\lambda, m_{\tilde{\nu}_\tau})$  and  $(\lambda', m_{\tilde{b}})$

# Discussions

- In the yellow and blue scenarios, we can allow a non-zero  $\lambda'_{232}$  (correspond to the yellow and blue points out of the vertical axis in Fig(b,c)). And this will make the  $B_s - \bar{B}_s$  mixing constraint weaker and enlarge the allowed parameter space for  $R_{K^{(*)}}$ .
- We choose the black stars that are very close to the  $3\sigma$  lower bound of  $(g-2)_\mu$  in Fig(o)(q)(s) to show the dependence of  $(g-2)_\mu$  from  $(\lambda', m_{\tilde{b}})$  in Fig(n)(p)(r). Otherwise, the  $3\sigma$  lower bound of  $(g-2)_\mu$  will disappear in Fig(n)(p)(r) because the contribution from  $\tilde{b}$  is much smaller compared with  $\tilde{\nu}_\tau$ .
- $B_s \rightarrow \mu^+ \mu^-$  is always satisfied once  $R_{K^{(*)}}$  is explained. Even if we take the extreme  $3\sigma$  value,  $|(C_{10})^\mu - (C'_{10})^\mu| = 0.89$  (Altmannshofer, Stangl (arXiv:2103.13370)). This implies the RPV contribution  $< 1.4 \times 10^{-10}$  (Becirevic, Fajfer, Kosnik (PRD 2015)) while the current experimental value is  $\text{BR}(B_s \rightarrow \mu^+ \mu^-) = (3.0 \pm 0.4) \times 10^{-9}$
- The lower bound of  $m_{\tilde{\nu}_\tau}$  comes from the recast of the 4-lepton search of ATLAS (ATLAS-CONF-2021-011). The 4-lepton signal in our scenario comes from the pair production of  $\tilde{\nu}_\tau$  with  $\tilde{\nu}_\tau \rightarrow \mu^+ \mu^-$ .

# Summary

- We suggest that in RPV3 SUSY:
  - A  $\tilde{b}$  with mass  $\sim 2 - 12$  TeV and non-zero couplings  $\lambda'_{233}$ ,  $\lambda'_{223}$  and  $\lambda'_{232}$  could explain  $R_{D^{(*)}}$  and  $R_{K^{(*)}}$  anomalies at  $3\sigma$  CL (especially even  $1\sigma$  for  $R_{K^{(*)}}$ ) while having a little bit contribution on  $(g-2)_\mu$  anomaly due to the constraints of  $B \rightarrow K^{(*)}\nu\bar{\nu}$ ,  $B_s - \bar{B}_s$  mixing etc.;
  - A  $\tilde{\nu}_\tau$  with mass  $\sim 0.7 - 1$  TeV and non-zero coupling  $|\lambda_{232}| \gtrsim 2.7$  could explain  $(g-2)_\mu$  anomaly at  $3\sigma$  CL.
  - Both  $(m_{\tilde{b}}, \lambda')$  and  $(m_{\tilde{\nu}_\tau}, \lambda)$  parameter spaces are (partly) testable at HL-LHC through  $\bar{t}\mu^+\mu^-$  or four muon signals.
  - Due to the orthogonality between  $(m_{\tilde{b}}, \lambda')$  and  $(m_{\tilde{\nu}_\tau}, \lambda)$  in the sense of the solutions of anomalies, even if the  $(m_{\tilde{b}}, \lambda')$  solution is ruled out by the signal we proposed or future low energy constraints, the  $(m_{\tilde{\nu}_\tau}, \lambda)$  solution is still valid and vice versa.

## Backup



# Choice of the couplings

- $(g - 2)_\mu \Rightarrow$  candidates from Eq.(3):  $\lambda_{312}, \lambda_{321}, \lambda_{322}, \lambda_{323}$  ( $\lambda'_{2k3}$  terms cannot contribute much). To get enough contribution, we have to let at least two  $\lambda$  couplings to be non-zero otherwise the magnitude of the  $\lambda$  need to be larger than  $\sqrt{4\pi}$ .
- $\checkmark \lambda_{322} \neq 0$ , contributes two times for  $k = 2$  in Eq.(3)
- $\times \lambda_{323} \neq 0$ , need to add another coupling to get enough contribution. But  $\tau \rightarrow \bar{e}\mu\mu \Rightarrow \lambda_{323}\lambda_{312}$  small (propagator  $\tilde{\nu}_\tau$ );  $\tau \rightarrow e\mu\bar{\mu} \Rightarrow \lambda_{323}\lambda_{321}$  small (propagator  $\tilde{\nu}_\tau$ );  $\tau \rightarrow \mu\mu\bar{\mu} \Rightarrow \lambda_{323}\lambda_{322}$  small (propagator  $\tilde{\nu}_\tau$ ).
- ?  $\lambda_{312} \neq 0$  and  $\lambda_{321} \neq 0$ , cannot let  $\lambda_{322} \neq 0$  at the same time due to the constraint of  $\mu \rightarrow e\gamma$
- $\times \lambda'_{3ij} \neq 0$  moreover, because  $(\bar{d}_i d_j) \rightarrow \mu\bar{\mu} \Rightarrow \lambda_{322}\lambda'_{3ij}$  small (propagator  $\tilde{\nu}_\tau$ ).

# Choice of couplings

- $R_{K^{(*)}} \Rightarrow \lambda'_{233} \lambda'_{223} \neq 0$  since  $\lambda'_{3ij} \approx 0$  and  $\lambda_{323} \approx 0$
- $\checkmark$   $\lambda'_{233} \neq 0$  and  $\lambda'_{223} \neq 0$ . Only choice of  $R_{K^{(*)}}$ , also contribute to  $R_{D^{(*)}}$
- $\lambda'_{ij3} \neq 0$  and  $\lambda'_{i3j} \neq 0$  may cause some troubles because
  - $(\bar{u}_j c) \rightarrow \bar{e}_i \mu \Rightarrow \lambda'_{223} \lambda'_{ij3}$  (propagator  $\tilde{b}_R$ ) e.g.  $D^0 \rightarrow \mu^+ \mu^- \Rightarrow \frac{\lambda'_{213}}{m_{\tilde{b}_R}^2}$  small;
  - $(\bar{d}_j b) \rightarrow \bar{e}_i \mu \Rightarrow \lambda'_{233} \lambda'_{i3j}$  (propagator  $\tilde{t}_L$ ) e.g.  $B_s \rightarrow \mu^+ \mu^- \Rightarrow \frac{\lambda'_{232}}{m_{\tilde{t}_L}^2}$  small.
- In our case  $\lambda'_{232} \neq 0$ . But a small  $\lambda'_{232}$  may also be possible (but the cancellation term in  $B_s - \bar{B}_s$  mixing is zero) and  $\lambda'_{232}$  will contribute to  $B_s \rightarrow \mu^+ \mu^-$ ,  $(C'_9)^\mu$  and  $(C'_{10})^\mu$ . But  $\lambda'_{232}$  do not have to be small and prefer the relation  $\lambda'_{223} \approx 3\lambda'_{232}$ .
- Now, non-zero couplings are chosen to be  $\lambda_{232}$ ,  $\lambda'_{233}$ ,  $\lambda'_{223}$  and  $\lambda'_{232}$ .

# Background cross-section

- $pp \rightarrow t\mu^+\mu^-$  has a larger background cross-section because the  $u$  content in proton is much larger than the  $\bar{u}$  content.
- We can look for  $t\mu^+\mu^-$  or even combine both, but the result should be similar because the signal of  $t\mu^+\mu^-$  is nearly the same as  $\bar{t}\mu^+\mu^-$ .

Table 1:  $pp \rightarrow \bar{t}\mu^+\mu^- X$  cross sections (fb)

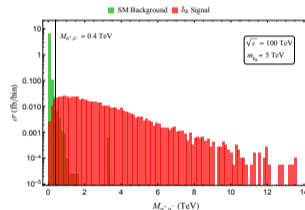
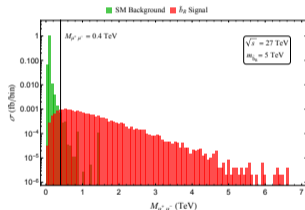
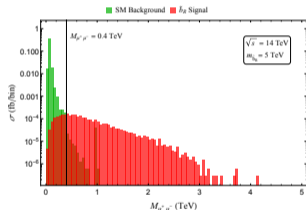
$X$	14 TeV	$M_{\mu^+\mu^-} > 0.15$ TeV	27 TeV	$M_{\mu^+\mu^-} > 0.15$ TeV	100 TeV	$M_{\mu^+\mu^-} > 0.15$ TeV
$j$	0.381	$3.35 \times 10^{-3}$	1.06	$1.05 \times 10^{-2}$	5.83	$7.11 \times 10^{-2}$
$b$	$4.23 \times 10^{-3}$	$3.64 \times 10^{-5}$	$9.47 \times 10^{-3}$	$9.85 \times 10^{-5}$	$3.84 \times 10^{-2}$	$3.92 \times 10^{-4}$
$W^+ \rightarrow jj$	$3.76 \times 10^{-3}$	$2.75 \times 10^{-5}$	$1.49 \times 10^{-2}$	$1.33 \times 10^{-4}$	0.133	$1.58 \times 10^{-3}$
$W^+ \rightarrow e^+\nu_e$	$6.38 \times 10^{-4}$	$5.68 \times 10^{-6}$	$2.53 \times 10^{-3}$	$2.68 \times 10^{-5}$	$2.24 \times 10^{-2}$	$2.28 \times 10^{-4}$
$W^+ \rightarrow \mu^+\nu_\mu$	$6.15 \times 10^{-3}$	$2.67 \times 10^{-3}$	$2.64 \times 10^{-2}$	$1.12 \times 10^{-2}$	0.242	0.120
$W^+ \rightarrow \tau^+\nu_\tau$	$6.34 \times 10^{-4}$	$6.09 \times 10^{-6}$	$2.52 \times 10^{-3}$	$3.08 \times 10^{-5}$	$2.25 \times 10^{-2}$	$2.81 \times 10^{-4}$
Total	0.396	$6.10 \times 10^{-3}$	1.12	$2.20 \times 10^{-2}$	6.29	0.194

$$^a p_T^{j,b,l} < 20 \text{ GeV}, E_T^{\text{miss}} < 20 \text{ GeV}$$

$$^b p_T^{t,\mu} > 20 \text{ GeV}, |\eta^{t,\mu}| < 2.5$$

# Invariant mass distribution (Red scenario)

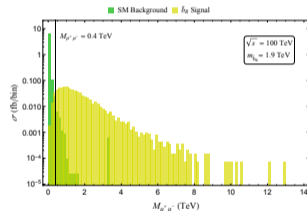
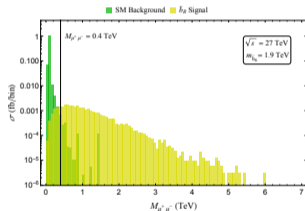
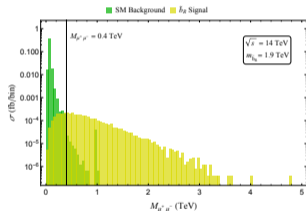
- For the process  $pp \rightarrow \bar{t}\mu^+\mu^-$
- Invariant mass  $M_{\mu^+\mu^-}$  distributions at  $\sqrt{s} = 14$  TeV, 27 TeV, 100 TeV



- We have used  $\lambda'_{233} = -\lambda'_{223} = 1.3$ ,  $m_{\tilde{b}_R} = 5$  TeV for the signal process (the black star in Fig(n)).

# Invariant mass distribution (Yellow scenario)

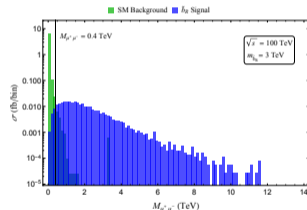
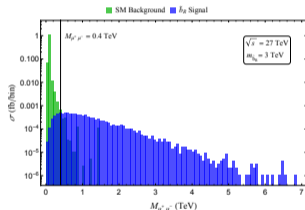
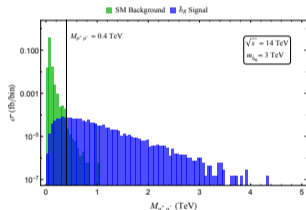
- For the process  $pp \rightarrow \bar{t}\mu^+\mu^-$
- Invariant mass  $M_{\mu^+\mu^-}$  distributions at  $\sqrt{s} = 14$  TeV, 27 TeV, 100 TeV



- We have used  $\lambda'_{233} = -8\lambda'_{223} = 1.5$ ,  $m_{\tilde{b}_R} = 1.9$  TeV for the signal process (the black star in Fig(p)).

# Invariant mass distribution (Blue scenario)

- For the process  $pp \rightarrow \bar{t}\mu^+\mu^-$
- Invariant mass  $M_{\mu^+\mu^-}$  distributions at  $\sqrt{s} = 14$  TeV, 27 TeV, 100 TeV



- We have used  $\lambda'_{223} = -6\lambda'_{233} = 1.6$ ,  $m_{\tilde{b}_R} = 3$  TeV for the signal process (the black star in Fig(r)).

# Anomalies and constraints (Red scenario)

- Since many anomalies and constraints are independent of  $(\lambda_{232}, m_{\tilde{\nu}_\tau})$ , they become just numbers instead of curves in Fig(o).

Anomaly/Constraint	Quantities in Figure(m)	Experimental value/limit
$R_{D^{(*)}}$	$\frac{R_{D^{(*)}}}{R_{D^{(*)}}^{\text{SM}}} = 1.05$	$1.15 \pm 0.04$
$R_{K^{(*)}}$	$(C_9)^\mu = -(C_{10})^\mu = -0.23$	$-0.35 \pm 0.08$
$D^0 \rightarrow \mu^+ \mu^-$	$\text{BR}(D^0 \rightarrow \mu^+ \mu^-) = 2.8 \times 10^{-10}$	$< 7.6 \times 10^{-9}$ (95% CL)
$B \rightarrow K^{(*)} \nu \bar{\nu}$	$R_{B \rightarrow K^{(*)} \nu \bar{\nu}} = \frac{\text{BR}(B \rightarrow K^{(*)} \nu \bar{\nu})}{\text{BR}_{\text{SM}}(B \rightarrow K^{(*)} \nu \bar{\nu})} = 4.6$	$< 5.2$ (95% CL)
$B_s - \bar{B}_s$ mixing	$\Delta M_{B_s} = (20.1 \pm 1.7) \text{ ps}^{-1}$	$(17.757 \pm 0.021) \text{ ps}^{-1}$
$B_s \rightarrow \mu^+ \mu^-$	$< 9.1 \times 10^{-12}$	$(3.0 \pm 0.4) \times 10^{-9}$

# Anomalies and constraints (Yellow scenario)

- Since many anomalies and constraints are independent of  $(\lambda_{232}, m_{\tilde{\nu}_\tau})$ , they become just numbers instead of curves in Fig(q).

Anomaly/Constraint	Quantities in Figure(o)	Experimental value/limit
$R_{D^{(*)}}$	$\frac{R_{D^{(*)}}}{R_{D^{(*)}}^{\text{SM}}} = 1.04$	$1.15 \pm 0.04$
$R_{K^{(*)}}$	$(C_9)^\mu = -(C_{10})^\mu = -0.13$	$-0.35 \pm 0.08$
$D^0 \rightarrow \mu^+ \mu^-$	$\text{BR}(D^0 \rightarrow \mu^+ \mu^-) = 2.6 \times 10^{-12}$	$< 7.6 \times 10^{-9}$ (95% CL)
$B \rightarrow K^{(*)} \nu \bar{\nu}$	$R_{B \rightarrow K^{(*)} \nu \bar{\nu}} = \frac{\text{BR}(B \rightarrow K^{(*)} \nu \bar{\nu})}{\text{BR}_{\text{SM}}(B \rightarrow K^{(*)} \nu \bar{\nu})} = 3.3$	$< 5.2$ (95% CL)
$B_s - \bar{B}_s$ mixing	$\Delta M_{B_s} = (22.4 \pm 1.7) \text{ ps}^{-1}$	$(17.757 \pm 0.021) \text{ ps}^{-1}$
$B_s \rightarrow \mu^+ \mu^-$	$3.0 \times 10^{-12}$	$(3.0 \pm 0.4) \times 10^{-9}$



# Anomalies and constraints (Blue scenario)

- Since many anomalies and constraints are independent of  $(\lambda_{232}, m_{\tilde{\nu}_\tau})$ , they become just numbers instead of curves in Fig(s).

Anomaly/Constraint	Quantities in Figure(q)	Experimental value/limit
$R_{D^{(*)}}$	$\frac{R_{D^{(*)}}}{R_{D^{(*)}}^{\text{SM}}} = 1.03$	$1.15 \pm 0.04$
$R_{K^{(*)}}$	$(C_9)^\mu = -(C_{10})^\mu = -0.13$	$-0.35 \pm 0.08$
$D^0 \rightarrow \mu^+ \mu^-$	$\text{BR}(D^0 \rightarrow \mu^+ \mu^-) = 5.4 \times 10^{-9}$	$< 7.6 \times 10^{-9}$ (95% CL)
$B \rightarrow K^{(*)} \nu \bar{\nu}$	$R_{B \rightarrow K^{(*)} \nu \bar{\nu}} = \frac{\text{BR}(B \rightarrow K^{(*)} \nu \bar{\nu})}{\text{BR}_{\text{SM}}(B \rightarrow K^{(*)} \nu \bar{\nu})} = 1.3$	$< 5.2$ (95% CL)
$B_s - \bar{B}_s$ mixing	$\Delta M_{B_s} = (22.2 \pm 1.7) \text{ ps}^{-1}$	$(17.757 \pm 0.021) \text{ ps}^{-1}$
$B_s \rightarrow \mu^+ \mu^-$	$2.8 \times 10^{-12}$	$(3.0 \pm 0.4) \times 10^{-9}$