

THE UNIVERSITY OF



Kavli Institute for Cosmological Physics AT THE UNIVERSITY OF CHICAGO

# Gravitational wave propagation beyond GR: waveform distortions and echoes

Jose Maria Ezquiaga, Wayne Hu, Macarena Lagos, Meng-Xiang Lin arXiv: 2108.10872 mxlin@uchicago.edu 2021.08.26 at SUSY2021



THE UNIVERSITY OF CHICAGO

Kavli Institute for Cosmological Physics At The UNIVERSITY OF CHICAGO

# Motivation

- LIGO/VIRGO gravitational wave (GW) detections give a strong constrain on gravity theory beyond GR
- Future: Einstein Telescope, Cosmic Explorer, LISA, TianQin, BBO, DECIGO, PTA
- It's the best era to test GR!





# Motivation

- We study GW propagation if GWs interact with another field
- Homogenous and isotropic cosmological background (no interaction with scalar or vector DoF in linear theory, only tensor DoF)
- Parametrized equation of motion
- Theoretical Examples: massive bigravity, Yang-Mills theories, Abelian multi-gauge field, Multi-Proca theories (see also Jimenez et al. 2020 for details)





## **Equation of Motion**

velocityMass
$$\left[\hat{I}\frac{d^2}{d\eta^2} + \hat{\nu}(\eta)\frac{d}{d\eta} + \hat{C}(\eta)k^2 + \hat{\Pi}(\eta)k + \hat{M}(\eta)\right] \begin{pmatrix} h \\ s \end{pmatrix} = 0$$
FrictionChiralh: GW fields: Coupled tensor field



for Cosmological Physics



### General Solution: two eigen propagation modes

$$h(\eta, k) = \sum_{A} h_0(k) f_A(\eta, k) e^{-i\phi_A(\eta, k)}$$
$$\phi_A = \int \omega_A d\eta \qquad (\text{GR: } \omega_A = ck)$$

- Constant coefficients: exact analytical solution
- Time-dependent coefficients: WKB approximation
- Two eigenmodes propagate independently in high-k limit.
- The detected signal is the superposition of the two eigenmodes.





## Phenomenon highlights

- Three timescales: mixing, coherence, broadening
- Observational implication:
  - Echoes
  - Phase distortion







# Echoes: from coherence to decoherence







## **Broadening**



Real waveforms: k-dependent group velocity  $\rightarrow$  phase distortion





# Summary

- We study GW propagation if GW field interacts with another tensor field, obtain general WKB solution
- Three timescales: mixing, coherence, broadening
- Observational implication:
  - Echoes
  - Phase distortion
- Other interesting phenomena in the paper (see arxiv:2108.10872)
  - Apparent luminosity distance change
  - Polarization oscillations and amplitude/phase birefringence
  - Broadening prevents decoherence





## BACKUP





# **Typical timescales**

• Dispersion relation:

$$\omega_A(k) = \omega_A(k_0) + \frac{\partial \omega_A}{\partial k}(k - k_0) + \frac{1}{2}\frac{\partial^2 \omega_A}{\partial k^2}(k - k_0)^2 + \cdots$$

- Mixing:  $T_{mix}|\omega_1 \omega_2| \sim 2\pi$ 
  - $T > T_{mix}$ : oscillations due to mixing
- Coherence:  $T_{coh}|v_{g1} v_{g2}| \sim \sigma_A$ ,  $v_{g,A} = \frac{\partial \omega_A}{\partial k}$ 
  - $T > T_{coh}$ : echoes
- Broadening:  $T_{broad,A} \left| \frac{\partial^2 \omega_A}{\partial k^2} \right| \sim \sigma_x^2$ 
  - $T > T_{broad}$ : phase distortions





## Observation: for real binary coalescence signal

	Regime	Observables
0)	$z \ll z_{ m mix}, z_{ m broad}, z_{ m coh}$	Unmodified waveform
1)	$z_{\rm broad} < z \ll z_{\rm mix}, z_{\rm coh}$	Single event with modified phase evolution
2a)	$z_{\rm mix} < z < z_{\rm broad}, z_{\rm coh}$	Single event with $d_L^{\text{GW}} \neq d_L$ and constant phase shift, or
2b)		frequency-dependent amplitude modulation with phase distortions
3)	$z_{\rm mix}, z_{ m broad} < z < z_{ m coh}$	Single event with modified phase evolution
4)	$z_{\rm coh} < z < z_{ m broad}$	Echoes with different arrival times and $d_L^{\text{GW}}$
5)	$z_{ m coh}, z_{ m broad} < z$	Echoes with different arrival times and phase distortions

Main effects: echoes, phase distortions, oscillations and birefringence











 $\beta$ : degree of circular polarization (amplitude birefringence)  $\chi$ : orientation of elliptical polarization (phase birefringence)







Kavli Institute for Cosmological Physics



Observation: apparent luminosity distance (friction mixing)





Kavli Institute for Cosmological Physics

AT THE UNIVERSITY OF CHICAGO



**Observation: birefringence (chiral mixing)** 

