

# Grand Unified Theories

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SUSY 2021

Beijing CAS

August 23 - 28, 2021

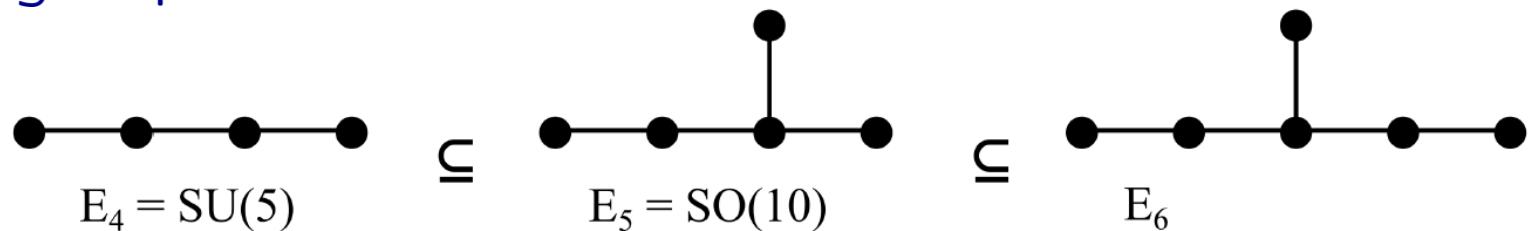
## Plan of the Talk

This is a broad-brush overview of GUTs

Much more in 3 dedicated sessions for GUTs

The hallmark of GUTs is proton decay; I will spend some time on this issue

Chiral nature of SM suggests embedding in exceptional Lie groups:



$\text{SU}(5)$  is the simplest rank-4 group;  $\text{SO}(10)$  has  $\nu_R$  leading to neutrino mass;  $E_6$  has additional vector-like leptons which can be dark matter

See talk on  $E_6$  by B. Bajc on Wednesday

# Grand Unified Theories: Motivations

- Electric charge quantization
  - ◊  $Q_p = -Q_e$  to better than 1 part in  $10^{21}$
- Miraculous cancellation of anomalies
- Quantum numbers of quarks and leptons
- Existence of  $\nu_R$  and thus neutrino mass via seesaw
- Unification of gauge couplings with low energy SUSY
- $b - \tau$  unification and a framework for fermion flavor structure
- Baryon asymmetry of the universe via leptogenesis

Pati, Salam (1973)

Georgi, Glashow (1974)

Georgi, Quinn, Weinberg (1974)

# Structure of Matter Multiplets

## Standard Model

$$Q = \begin{pmatrix} u_1 & u_2 & u_3 \\ d_1 & d_2 & d_3 \end{pmatrix} \sim (3, 2, \frac{1}{6})$$

$$u^c = (u_1^c \quad u_2^c \quad u_3^c) \sim (\bar{3}, 1, -\frac{2}{3})$$

$$d^c = (d_1^c \quad d_2^c \quad d_3^c) \sim (\bar{3}, 1, \frac{1}{3})$$

$$L = \begin{pmatrix} \nu \\ e^- \end{pmatrix} \sim (1, 2, -\frac{1}{2})$$

$$e^c \sim (1, 1, +1)$$

$$\nu^c \sim (1, 1, 0)$$

## SO(10)

$u_r : \{-+++--\}$	$d_r : \{-++-+-\}$	$u_r^c : \{+-+-++\}$	$d_r^c : \{+-+---\}$
$u_b : \{+-+-+-\}$	$d_b : \{+-+-++\}$	$u_b^c : \{+-+-++\}$	$d_b^c : \{-+-+-\}$
$u_g : \{++-+-+\}$	$d_g : \{++--+-\}$	$u_g^c : \{--+-++\}$	$d_g^c : \{--+--\}$
$v : \{---+-+\}$	$e : \{---+-+\}$	$v^c : \{++++++\}$	$e^c : \{+++--\}$

Frist 3 spins refer to color, last 2 are weak spins

$$Y = \tfrac{1}{3}\Sigma(C) - \tfrac{1}{2}\Sigma(W)$$

# Finding Order in Fermion Mass Spectrum

Fermion masses in units of  $m_t$

$$m_t = 1.0$$

$$m_b = 1.67 \times 10^{-2}$$

$$m_c = 3.6 \times 10^{-3}$$

$$m_s = 3.1 \times 10^{-4}$$

$$m_u = 1.3 \times 10^{-5}$$

$$m_d = 2.3 \times 10^{-5}$$

$$m_\tau = 1.0 \times 10^{-2}$$

$$m_3 = 2.9 \times 10^{-13}$$

$$m_\mu = 6.2 \times 10^{-4}$$

$$m_2 = 5.2 \times 10^{-14}$$

$$m_e = 3.0 \times 10^{-6}$$

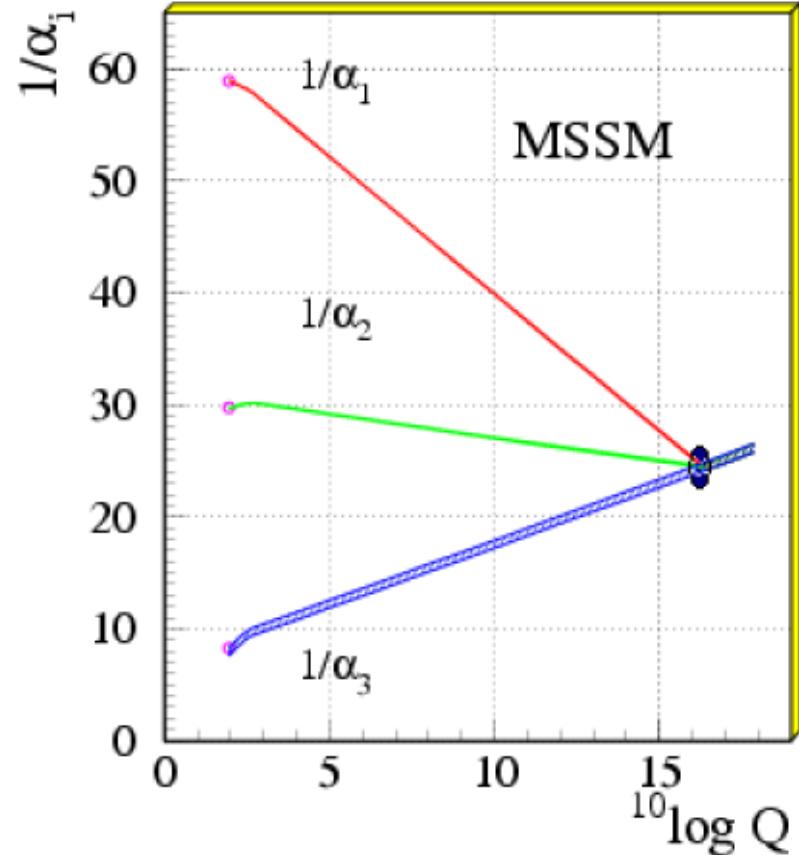
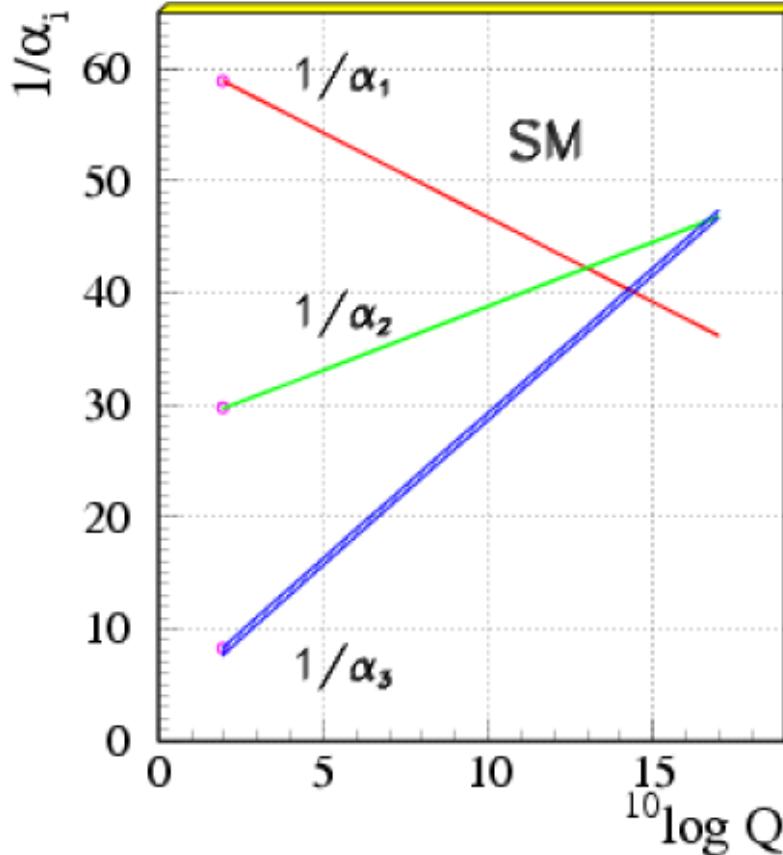
$$m_1 = < m_2$$

$$V_q = \begin{pmatrix} 0.976 & 0.22 & 0.004 \\ -0.22 & 0.98 & 0.04 \\ 0.007 & -0.04 & 1 \end{pmatrix} \quad U_\ell = \begin{pmatrix} 0.85 & -0.54 & 0.16 \\ 0.33 & 0.62 & -0.72 \\ -0.40 & -0.59 & -0.70 \end{pmatrix}$$

$$\text{Im} \left( \frac{V_{ub} V_{cs}}{V_{us} V_{cb}} \right) = 0.34$$

GUTs provide a desirable setup to address these issues

# Gauge Coupling Unification

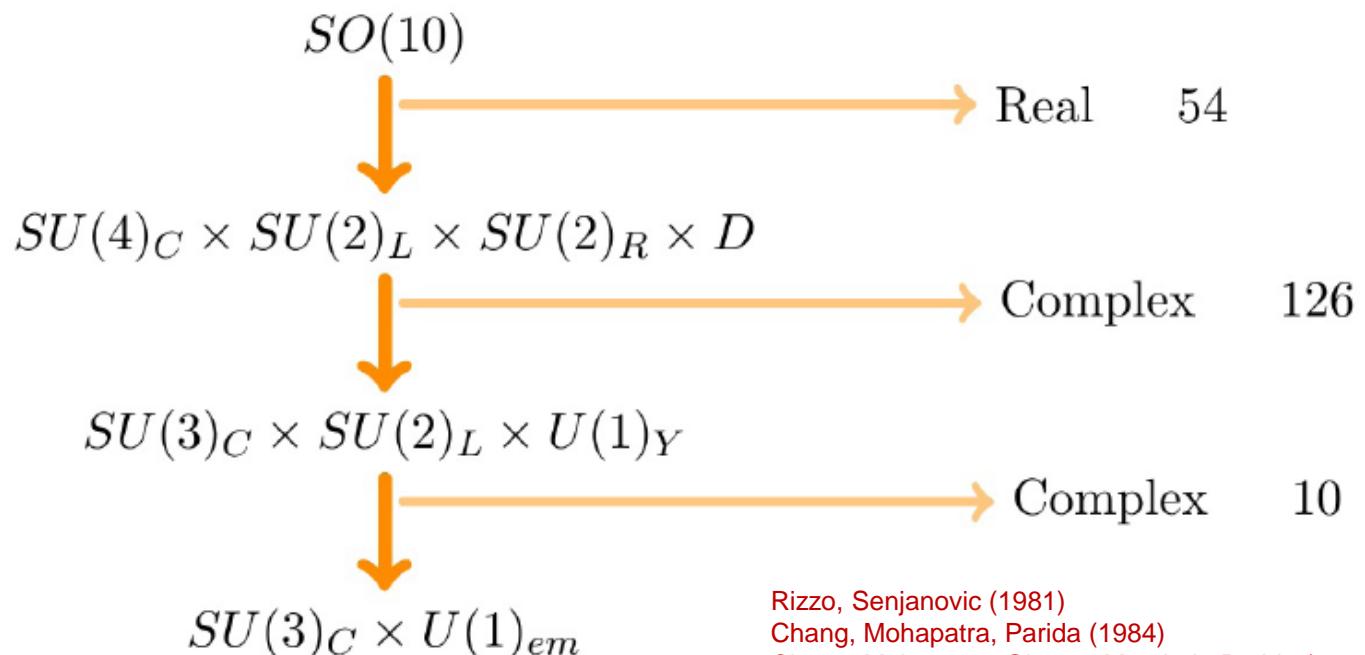


From S. Raby, PDG Review

# Gauge Coupling Unification in Non-SUSY SO(10)

$SO(10)$  admits an intermediate Pati-Salam symmetry

Gauge couplings evolve differently above the Pati-Salam scale

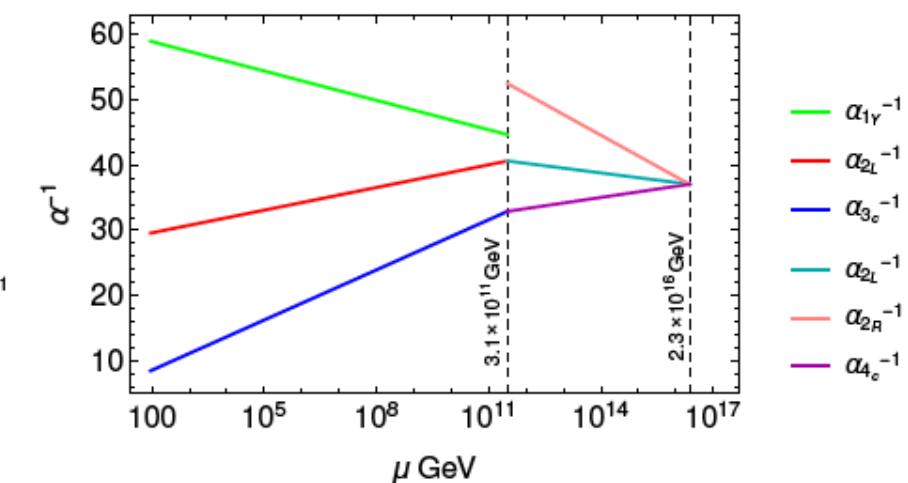
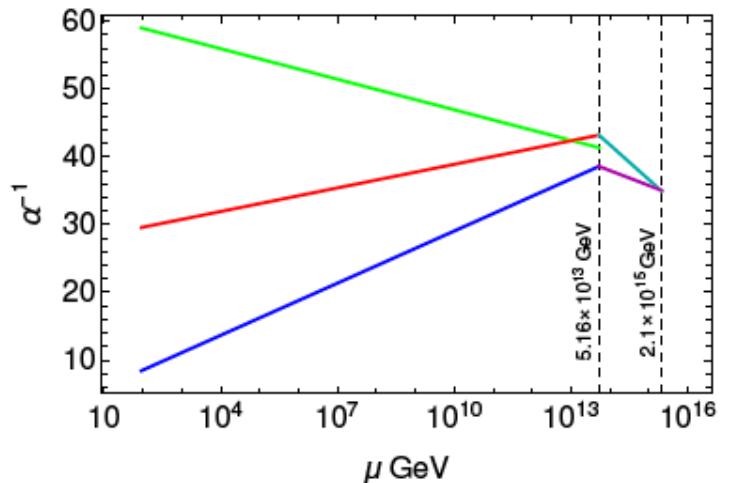


- Rizzo, Senjanovic (1981)  
Chang, Mohapatra, Parida (1984)  
Chang, Mohapatra, Gipson, Marshak, Parida (1985)  
Deshpande, Keith, Pal (1993)  
Bertolini, Di Luzio, Malinsky (2013)  
Babu, Khan (2015)  
Graf, Malinsky, Mede, Susic (2016)  
Babu, Bajc, Saad (2017)  
Babu, Saad (2020)  
Ohlsson, Pernow, Sonnerlind (2020)  
Meloni, Ohlsson, Pernow (2020)

# Unification with Intermediate Scale in SO(10)

Intermediate Pati-Salam symmetry –  $SU(4)_c \times SU(2)_L \times SU(2)_R$

With and without left-right Parity



From Babu, Bajc, Saad (2017)

# Quark-Lepton Unification

Initial ideas of unification arose from

$$SU(4)_c \times SU(2)_L \times SU(2)_R \times P \quad \text{Pati, Salam (1973)}$$

Quarks and leptons are unified, with lepton number identified as the 4th color

Each family of fermions belong to  $\{(4, 2, 1)_L \oplus (4, 1, 2)_R\}$

$$\begin{pmatrix} u_1 & u_2 & u_3 & \nu \\ d_1 & d_2 & d_3 & \ell \end{pmatrix}_L \oplus \begin{pmatrix} u_1 & u_2 & u_3 & \nu \\ d_1 & d_2 & d_3 & \ell \end{pmatrix}_R$$

Gauge bosons do not mediate proton decay, but Higgs bosons do

$$p \rightarrow e^+ e^- e^+$$

Pati (1984)

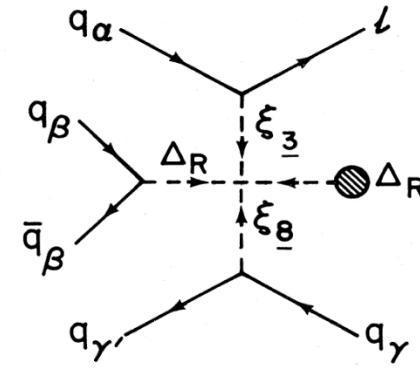


FIG. 1. Mechanism for  $3q \rightarrow lq\bar{q}$ .

# Minimal SU(5) GUT

**Matter multiplets:**

$$\{10 + \bar{5} + 1\}$$

Georgi, Glashow (1974)

$$10 : \frac{1}{\sqrt{2}} \begin{pmatrix} 0 & u_3^c & -u_2^c & u_1 & d_1 \\ -u_3^c & 0 & u_1^c & u_2 & d_2 \\ u_2^c & -u_1^c & 0 & u_3 & d_3 \\ -u_1 & -u_2 & -u_3 & 0 & e^c \\ -d_1 & -d_2 & -d_3 & -e^c & 0 \end{pmatrix}$$

$$\bar{5} : (d_1^c, d_2^c, d_3^c, e, -\nu_e)$$

$$1 : \nu^c$$

**Higgs:**

$$24_H, \quad \{5_H, \bar{5}_H\} \rightarrow \text{Contain color triplets } \{H_C, \bar{H}_C\}$$

**Yukawa Couplings**

$$Y_u^{ij} 10_i 10_j 5_H + Y_d^{ij} 10_i \bar{5}_j \bar{5}_H$$

$$M_\ell = M_d^T \Rightarrow m_b = m_\tau, \quad m_s = m_\mu, \quad m_d = m_e$$

# Doublet-Triplet Splitting in SUSY SU(5)

$$W_{D-T} = \bar{\mathbf{5}}_H (\lambda \mathbf{24}_H + M) \mathbf{5}_H$$

$$\langle \mathbf{24}_H \rangle = \begin{pmatrix} 1 & 0 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 & 0 \\ 0 & 0 & 1 & 0 & 0 \\ 0 & 0 & 0 & -3/2 & 0 \\ 0 & 0 & 0 & 0 & -3/2 \end{pmatrix} V$$

FINE-TUNED TO  $O(M_W)$

$$M_{H_c} = \lambda V + M \sim O(M_{GUT}) \quad M_H = -\frac{3}{2}\lambda V + M$$

## The Good

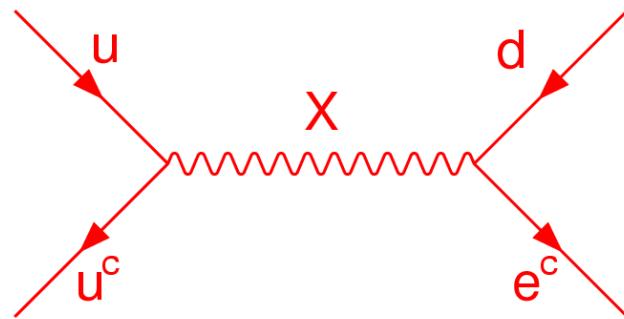
- (1) Predicts unification of couplings
- (2) Uses economic Higgs sector

## The Not so Good

- (1) Unnatural fine tuning
- (2) Typically large proton decay rate

# Nucleon Decay in SUSY GUTs

Gauge boson exchange



$$\Gamma^{-1}(p \rightarrow e^+ \pi^0) = (2.0 \times 10^{35} \text{ yr})$$

$$\times \left( \frac{\alpha_H}{0.01 \text{ GeV}^3} \right)^{-2} \left( \frac{\alpha_G}{1/25} \right)^{-2} \left( \frac{A_R}{2.5} \right)^{-2} \left( \frac{M_X}{10^{16} \text{ GeV}} \right)^4$$

$$(-2\alpha_3^{-1} - 3\alpha_2^{-1} + 3\alpha_Y^{-1})(M_Z) = \frac{1}{2\pi} \left\{ 36 \ln \left( \frac{M_X}{M_Z} \left( \frac{M_\Sigma}{M_X} \right)^{1/3} \right) + 8 \ln \left( \frac{M_{\text{SUSY}}}{M_Z} \right) \right\}$$

$M_\Sigma$ : Heavy color octet mass, uncertain: Threshold effect

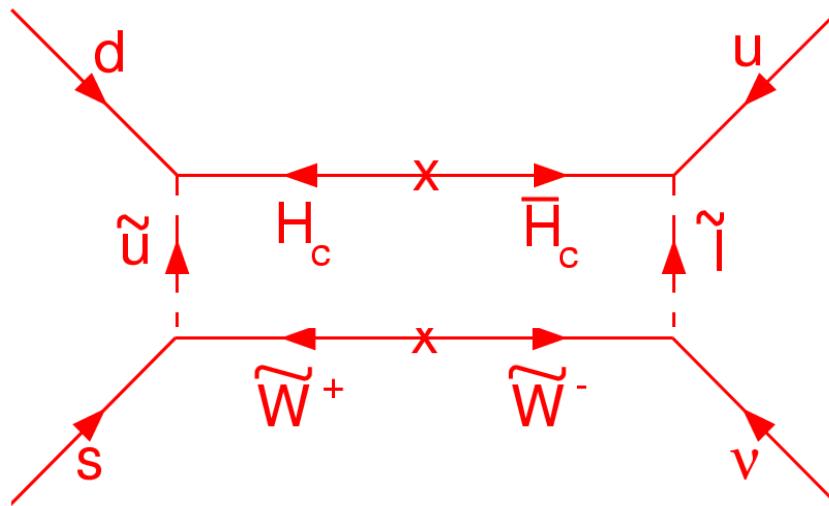
$\frac{M_\Sigma}{M_X} \leq 1.8$  (perturbation theory)

Hisano, Murayama, Yanagida (1993)

Nath, Perez, Phys. Rept. (2007)

Babu, Kearns, et. al. (2013)

# SUSY Proton Decay Modes



Sakai, Yanagida (1982)  
Weinberg (1982)

$$\Gamma_{d=5}^{-1}(p \rightarrow \bar{\nu} K^+) \simeq 1.2 \cdot 10^{31} \text{ yrs} \times \left( \frac{0.012 \text{ GeV}^3}{\beta_H} \right)^2 \left( \frac{7}{\bar{A}_S^\alpha} \right)^2 \left( \frac{1.25}{R_L} \right)^2 \\ \times \left( \frac{M_T}{2 \cdot 10^{16} \text{ GeV}} \right)^2 \left( \frac{m_{\tilde{q}}}{1.5 \text{ TeV}} \right)^4 \left( \frac{190 \text{ GeV}}{M_{\tilde{W}}} \right)^2$$

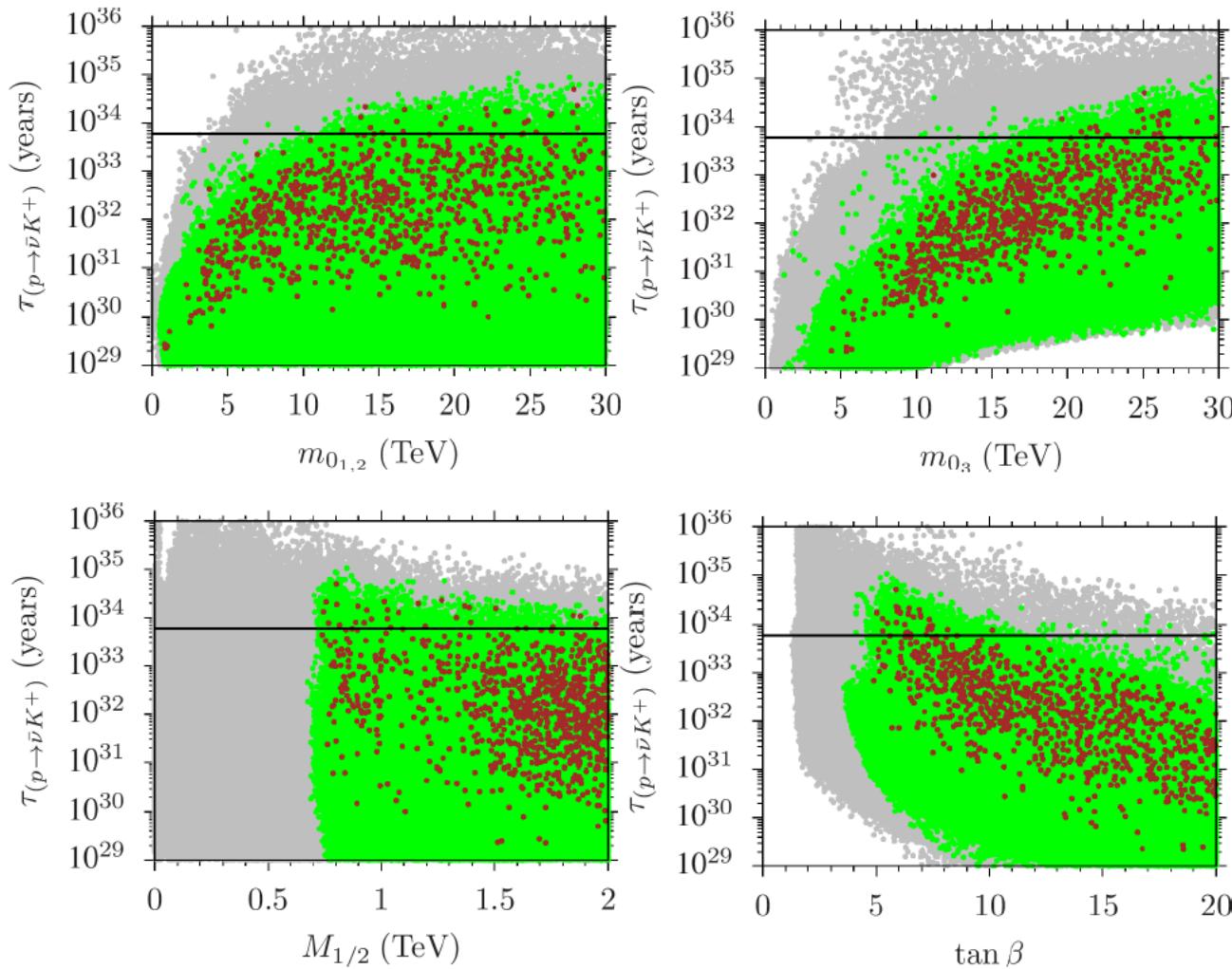
Hisano, Murayama, Yanagida (1993);

Murayama, Pierce (2002);.....

Ellis, Evans, Nagata, Olive, Valesco-Sevilla (2019);

Babu, Gogoladze, Un (2020)

# Proton Decay in SUSY SU(5)



# SUSY SU(5) Benchmark Points

	Point 1	Point 2	Point 3
$m_{01,2}$	<b>12580</b>	16720	21050
$m_{03}$	25750	26140	26110
$M_{1/2}$	815.3	803	1947
$\tan \beta$	6.38	5.93	5.31
$A_0/m_{03}$	-0.746	1.4	-1.12
$\mu$	394.8	<b>360</b>	792.1
$m_A$	14730	26170	20420
$m_t$	<b>173.3</b>	173.3	173.3
$\tan \beta$	6.38	5.93	5.31
$m_h$	124.19	125.35	123.22
$m_H$	14827	26342	20554
$m_A$	14730	26170	20420
$m_{H^\pm}$	14730	26170	20420
$m_{\tilde{\chi}_1^0}, m_{\tilde{\chi}_2^0}$	<b>359.58, 412.35</b>	<b>315.49, 375.83</b>	<b>803.66, 817.58</b>
$m_{\tilde{\chi}_3^0}, m_{\tilde{\chi}_4^0}$	<b>425.25, 769.71</b>	<b>382.82, 671.81</b>	<b>937.27, 1781</b>
$m_{\tilde{\chi}_1^\pm}, m_{\tilde{\chi}_2^\pm}$	<b>418.8, 741.16</b>	<b>377.34, 647.93</b>	<b>845.5, 1731</b>
$m_{\tilde{g}}$	2275	2102	4848
$m_{\tilde{u}_{L,R}},$	12216, 13068	16513, 16845	21103, 21551
$m_{\tilde{t}_{1,2}},$	15255, 21197	10955, 20021	12962, 20630
$m_{\tilde{d}_{L,R}},$	12217, 12158	16514, 16567	21103, 21091
$m_{\tilde{b}_{1,2}},$	21136, 25727	20107, 26129	20659, 26164
$m_{\tilde{\nu}_{1,2}}$	12953	16814	21292
$m_{\tilde{\nu}_3}$	25954	26202	26331
$m_{\tilde{\epsilon}_{L,R}},$	12974, 11610	16809, 16403	21282, 20569
$m_{\tilde{\tau}_{1,2}},$	25224, 25954	25817, 26171	25644, 26277
$\sigma_{SI} (\text{pb})$	$1.28 \times 10^{-8}$	$1.18 \times 10^{-8}$	<b><math>5.13 \times 10^{-9}</math></b>
$\sigma_{SD} (\text{pb})$	$6.11 \times 10^{-5}$	$7.12 \times 10^{-5}$	$4.65 \times 10^{-6}$
$\Omega h^2$	<b>0.122</b>	<b>0.12</b>	<b>0.117</b>
$\tau(p \rightarrow \bar{\nu} K^+) \times 10^{-33}$	6.99	18.56	7.36

## SUSY SO(10) Models

Fermions of each family unified in  $\{16\}_i$  ( $i = 1 - 3$ )

Generally three types of Higgs are needed:

- $SO(10)$  breaking to rank 5 subgroup ( $45_H$  or  $210_H$ )
- Rank reduction ( $16_H$  or  $126_H$ )
- Electroweak symmetry breaking ( $10_H$ )

- (1)  $\{45_H + 16_H + 10_H\}$  models perturbative upto Planck scale
- (2)  $\{210_H + 126_H + 10_H\}$  models have direct couplings, automatic  $R$  parity, and are very predictive

# SUSY SO(10) with Natural Doublet-Triplet Splitting

Higgs sector:  $\{45_H + 10_H + 16_H + \overline{16}_H + 16'_H + \overline{16}'_H\}$

$$W_{D-T} = \lambda(10_H 45_H 10'_H) + M' 10'_H 10'_H$$

$$\langle 45_H \rangle = \begin{pmatrix} a & 0 & 0 & 0 & 0 \\ 0 & a & 0 & 0 & 0 \\ 0 & 0 & a & 0 & 0 \\ 0 & 0 & 0 & 0 & 0 \\ 0 & 0 & 0 & 0 & 0 \end{pmatrix} \otimes i\tau_2 \propto B - L$$

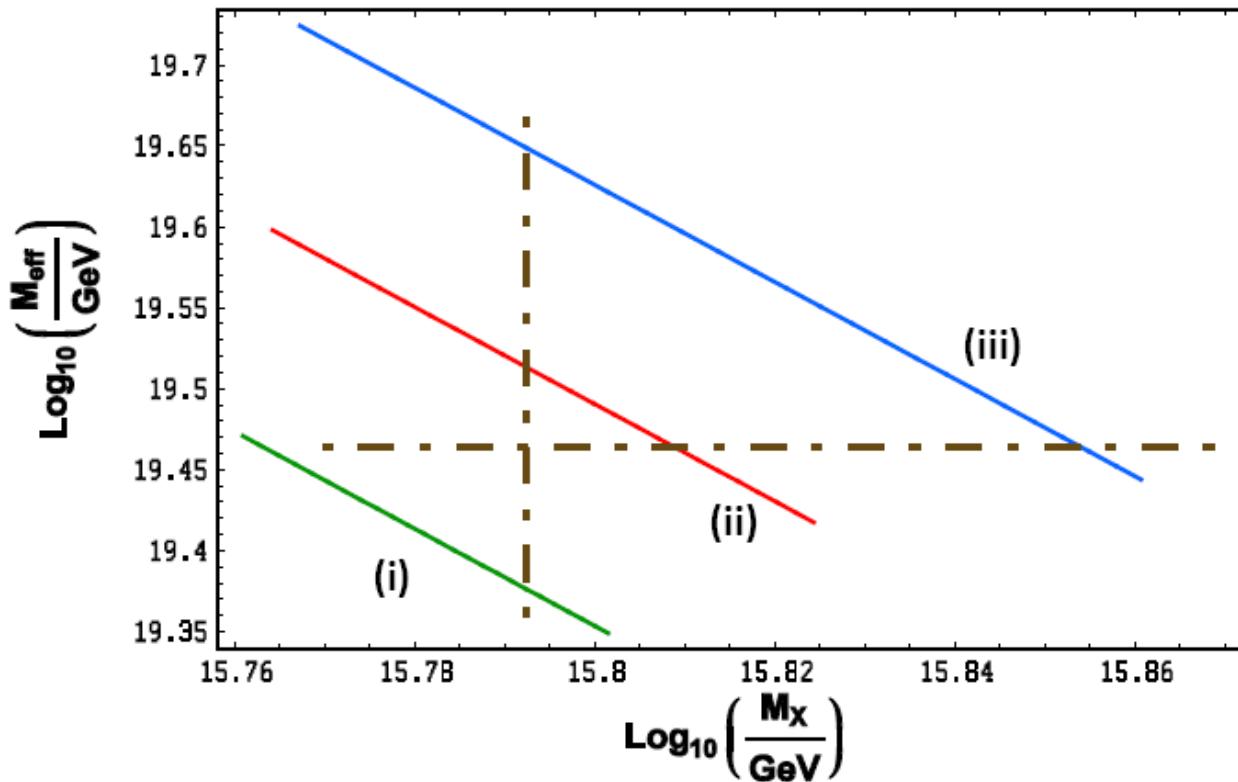
Dimopoulos, Wilczek (1981)  
 Babu, Barr (1993)  
 Barr, Raby (2000)

Adjoint VEV along  $B - L$  gives mass only to color triplets and not to doublets

Superpotential:

$$\begin{aligned} W(A) &= M_A \text{tr} A^2 + \frac{\lambda_A}{M_*} (\text{tr} A^2)^2 + \frac{\lambda'_A}{M_*} \text{tr} A^4 \\ W(A, C, C') &= C \left( \frac{a_1}{M_*} ZA + \frac{b_1}{M_*} C\bar{C} + c_1 S \right) \bar{C}' + C' \left( \frac{a_2}{M_*} ZA + \frac{b_2}{M_*} C\bar{C} + c_2 S \right) \bar{C} \\ W(DT) &= \lambda_1 HAH' + \lambda_{H'} \frac{S^k}{M_*^{k-1}} (H')^2 + \lambda_2 H\bar{C}\bar{C} \end{aligned}$$

Babu, Pati, Tavartkiladze (2010)



Correlation for spectrum with  $\tan\beta = 7$  and  $r=1/1200$ .

(i):  $\alpha_3 = 0.1177$ . (ii):  $\alpha_3 = 0.1184$ . (iii):  $\alpha_3 = 0.1191$ .

See also: Dermisek, Mafi, Raby (2000)

Babu, Pati, Wilczek (2000)

Altarelli, Feruglio (2001)

Bajc, Perez, Senjanovic (2002)

Li, Nanopoulos, Walker (2010)

Babu, Bajc, Tavartkiladze (2012)

Chen, Zhang (2014)

# Minimal SO(10) Yukawa Theory

$$\mathcal{L}_{\text{Yukawa}} = 16 \left( Y_{10} \mathbf{10}_H + Y_{126} \overline{\mathbf{126}}_H \right) \mathbf{16}$$

Two Yukawa matrices determine all fermion masses and mixings, including the neutrinos

$$\begin{aligned} M_u &= \kappa_u Y_{10} + \kappa'_u Y_{126} \\ M_d &= \kappa_d Y_{10} + \kappa'_d Y_{126} \\ M_\nu^D &= \kappa_u Y_{10} - 3\kappa'_u Y_{126} & M_{\nu R} &= \langle \Delta_R \rangle Y_{126} \\ M_l &= \kappa_d Y_{10} - 3\kappa'_d Y_{126} & M_{\nu L} &= \langle \Delta_L \rangle Y_{126} \end{aligned}$$

Model has only 11 real parameters plus 7 phases

Babu, Mohapatra (1993)

Bajc, Melfo, Senjanovic, Vissani (2002)

Fukuyama, Okada (2002)

Bajc, Melfo, Senjanovic, Vissani (2004)

Fukuyama, Ilakovac, Kikuchi, Meljanac, Okada (2004)

Aulakh et al (2004)

Bertolini, Frigerio, Malinsky (2004)

Babu, Macesanu (2005)

Bertolini, Malinsky, Schwetz (2016)

Dutta, Mimura, Mohapatra (2007), (2009)

Bajc, Dorsner, Nemevsek (2009)

Joshipura, Patel (2011)

Dueck, Rodejohann (2013)

Fukuyama, Mimura (2015)

Babu, Bajc, Saad (2018)

Babu, Fukuyama, Khan, Saad (2018)

Babu, Saad (2020)

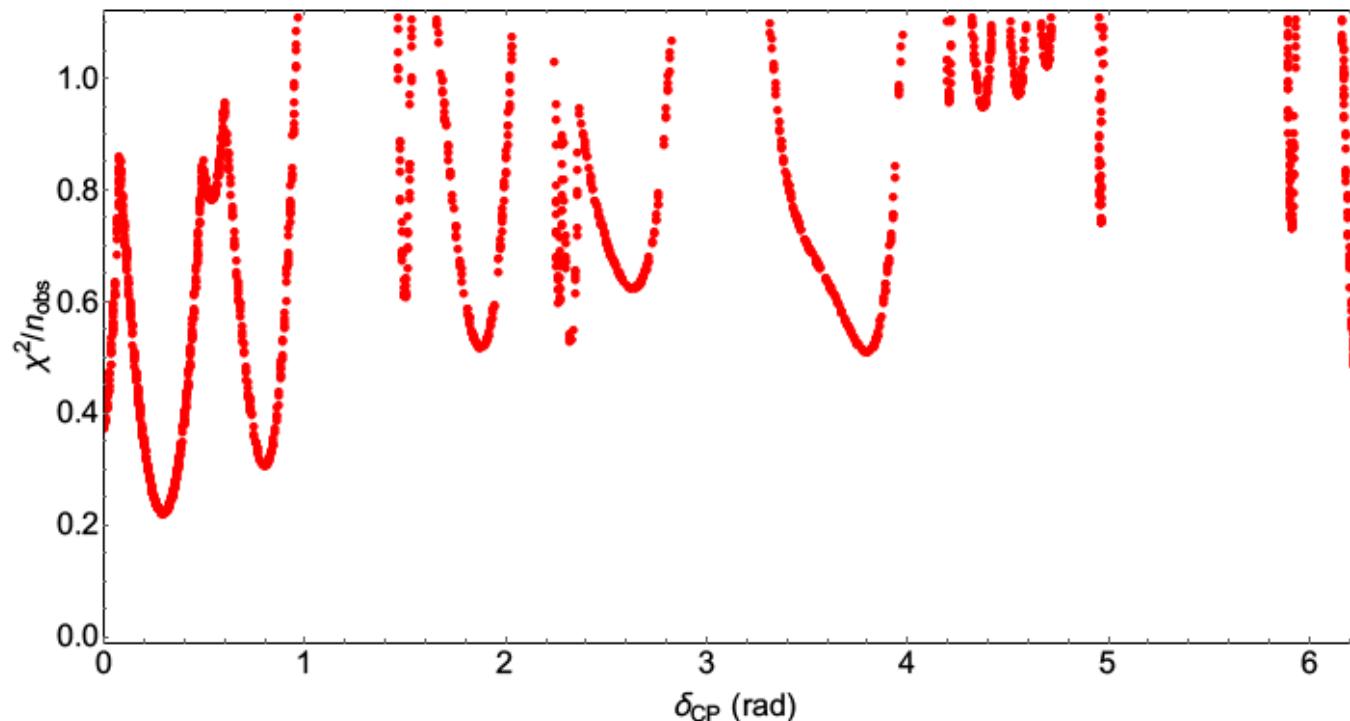
# Recent Global Fit Including CP Phase

$$\mathcal{L}_{\text{Yukawa}} = 16 \left( Y_{10} \mathbf{10}_H + Y_{126} \overline{\mathbf{126}}_H \right) 16$$

Observables (masses in GeV)	SUSY			non-SUSY		
	Input	Best Fit	Pull	Input	Best Fit	Pull
$m_u/10^{-3}$	$0.502 \pm 0.155$	0.515	0.08	$0.442 \pm 0.149$	0.462	0.13
$m_c$	$0.245 \pm 0.007$	0.246	0.14	$0.238 \pm 0.007$	0.239	0.18
$m_t$	$90.28 \pm 0.89$	90.26	-0.02	$74.51 \pm 0.65$	74.47	-0.05
$m_b/10^{-3}$	$0.839 \pm 0.17$	0.400	-2.61	$1.14 \pm 0.22$	0.542	-2.62
$m_s/10^{-3}$	$16.62 \pm 0.90$	16.53	-0.09	$21.58 \pm 1.14$	22.57	0.86
$m_b$	$0.938 \pm 0.009$	0.933	-0.55	$0.994 \pm 0.009$	0.995	0.19
$m_e/10^{-3}$	$0.3440 \pm 0.0034$	0.344	0.08	$0.4707 \pm 0.0047$	0.470	-0.03
$m_\mu/10^{-3}$	$72.625 \pm 0.726$	72.58	-0.05	$99.365 \pm 0.993$	99.12	-0.24
$m_\tau$	$1.2403 \pm 0.0124$	1.247	0.57	$1.6892 \pm 0.0168$	1.688	-0.05
$ V_{us} /10^{-2}$	$22.54 \pm 0.07$	22.54	0.02	$22.54 \pm 0.06$	22.54	0.06
$ V_{cb} /10^{-2}$	$3.93 \pm 0.06$	3.908	-0.42	$4.856 \pm 0.06$	4.863	0.13
$ V_{ub} /10^{-2}$	$0.341 \pm 0.012$	0.341	0.003	$0.420 \pm 0.013$	0.421	0.10
$\delta_{CKM}^o$	$69.21 \pm 3.09$	69.32	0.03	$69.15 \pm 3.09$	70.24	0.35
$\Delta m_{21}^2/10^{-5}(eV^2)$	$8.982 \pm 0.25$	8.972	-0.04	$12.65 \pm 0.35$	12.65	-0.01
$\Delta m_{31}^2/10^{-3}(eV^2)$	$3.05 \pm 0.04$	3.056	0.02	$4.307 \pm 0.059$	4.307	0.006
$\sin^2 \theta_{12}$	$0.318 \pm 0.016$	0.314	-0.19	$0.318 \pm 0.016$	0.316	-0.07
$\sin^2 \theta_{23}$	$0.563 \pm 0.019$	0.563	0.031	$0.563 \pm 0.019$	0.563	0.01
$\sin^2 \theta_{13}$	$0.0221 \pm 0.0006$	0.0221	-0.003	$0.0221 \pm 0.0006$	0.0220	-0.16
$\delta_{CP}^o$	$224.1 \pm 33.3$	240.1	0.48	$224.1 \pm 33.3$	225.1	0.03
$\chi^2$	-	-	7.98	-	-	7.96

# Pinning Down the CP Phase?

When other parameters are marginalized, CP phase is found to have multiple minima in  $\chi^2$



# Fermion Mass Hierarchy From Clockwork

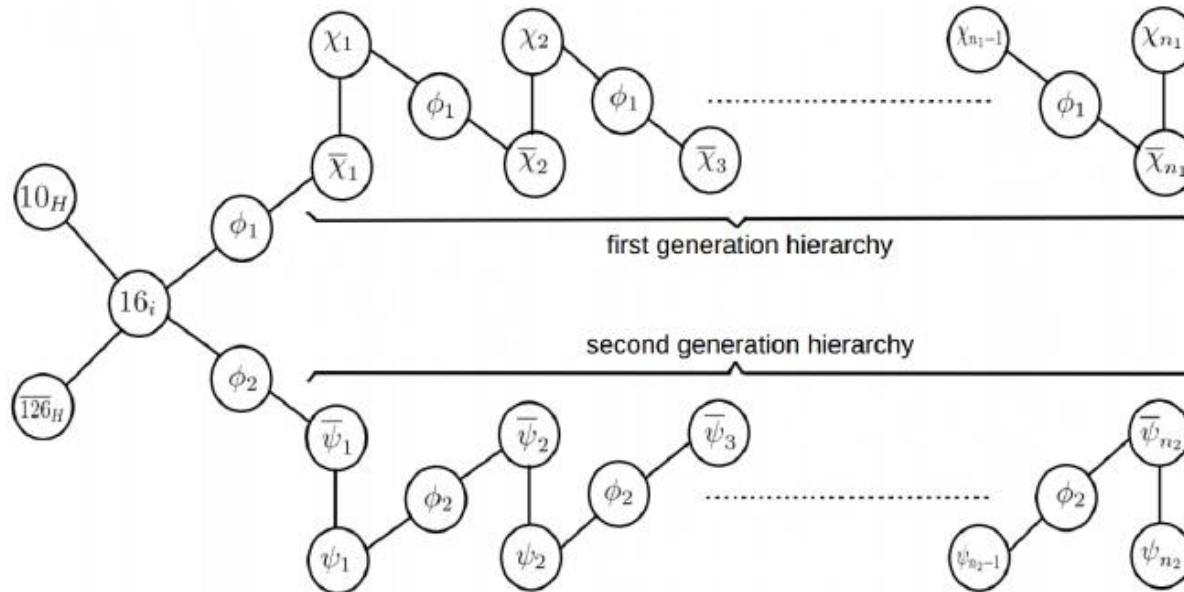
- GUTs alone cannot explain fermions mass hierarchy
- A “clockwork mechanism” can explain hierarchy  
Analogous to Froggatt-Nielsen mechanism
- SM Fermions mix with a chain of vector-like fermions  
Vector fermions have nearest neighbor interactions  
SM fermions are zero modes in this chain
- Small numbers arise naturally  
Recently applied to minimal SO(10)

Choi, Im (2016); Kaplan, Rattazzi (2016); Giudice, Mccullough (2016)

Von Gersdroff (2017); Patel (2017); Alonso et al (2018); Sanino et al (2019),...

# Clockwork SO(10)

Three fermion families mix with two clockwork chains



Babu, Saad (2020)

$$\begin{pmatrix} -q_1 & 1 & 0 & \cdots & 0 \\ 0 & -q_1 & 1 & \cdots & 0 \\ 0 & 0 & -q_1 & \cdots & 0 \\ \vdots & \vdots & \vdots & \ddots & \vdots \\ 0 & 0 & 0 & \cdots & -q_1 & 1 \end{pmatrix}_{n_1 \times (n_1+1)}$$

## Clockwork SO(10) (cont.)

Each chain generates a small parameter  $\epsilon$

$$M_d = 0.80166 e^{2.18422i} \begin{pmatrix} 1.27815 \epsilon_1^2 e^{-3.0004i} & 1.26174 \epsilon_1 \epsilon_2 e^{-2.71759i} & 1.11777 \epsilon_1 e^{1.63891i} \\ 1.26174 \epsilon_1 \epsilon_2 e^{-2.71759i} & 1.61178 \epsilon_2^2 e^{-2.31606i} & 1.20224 \epsilon_2 e^{2.02843i} \\ 1.11777 \epsilon_1 e^{1.63891i} & 1.20224 \epsilon_2 e^{2.02843i} & 1 \end{pmatrix} \text{ GeV}$$

$$M_u = 79.1467 e^{2.27241i} \begin{pmatrix} 1.09514 \epsilon_1^2 e^{3.10025i} & 1.17556 \epsilon_1 \epsilon_2 e^{-2.80578i} & 1.04143 \epsilon_1 e^{1.55072i} \\ 1.17556 \epsilon_1 \epsilon_2 e^{-2.80578i} & 1.24674 \epsilon_2^2 e^{-2.42765i} & 1.12012 \epsilon_2 e^{1.94025i} \\ 1.04143 \epsilon_1 e^{1.55072i} & 1.12012 \epsilon_2 e^{1.94025i} & 1 \end{pmatrix} \text{ GeV}$$

$\epsilon_1 \simeq 0.06$ ,  $\epsilon_2 \simeq 0.33$  arise from length of chains

Neutrino masses and mixings fixed in terms of these

The main test of these models is in proton decay

# Proton Decay in non-SUSY SO(10)

$$\tau_P \approx \frac{\pi}{4} R_L^2 (1 + F + D) \frac{|\alpha|^2}{f_\pi^2} m_p \alpha_G^2 \left[ A_{SR}^2 \left( \frac{1}{M_{(X,Y)}^2} + \frac{1}{M_{(X',Y')}^2} \right)^2 + \frac{4A_{SL}^2}{M_{(X,Y)}^4} \right]^{-1}$$

Here,

$$A_{SL(R)} = \prod_{i=1}^n \prod_{sc}^{Mz \leq m_{sc} < M_G} \left[ \frac{\alpha_i(m_{sc+1})}{\alpha_i(m_{sc})} \right] \frac{\gamma_{L(R)i(sc)}}{b_i(m_{sc+1} - m_{sc})}$$

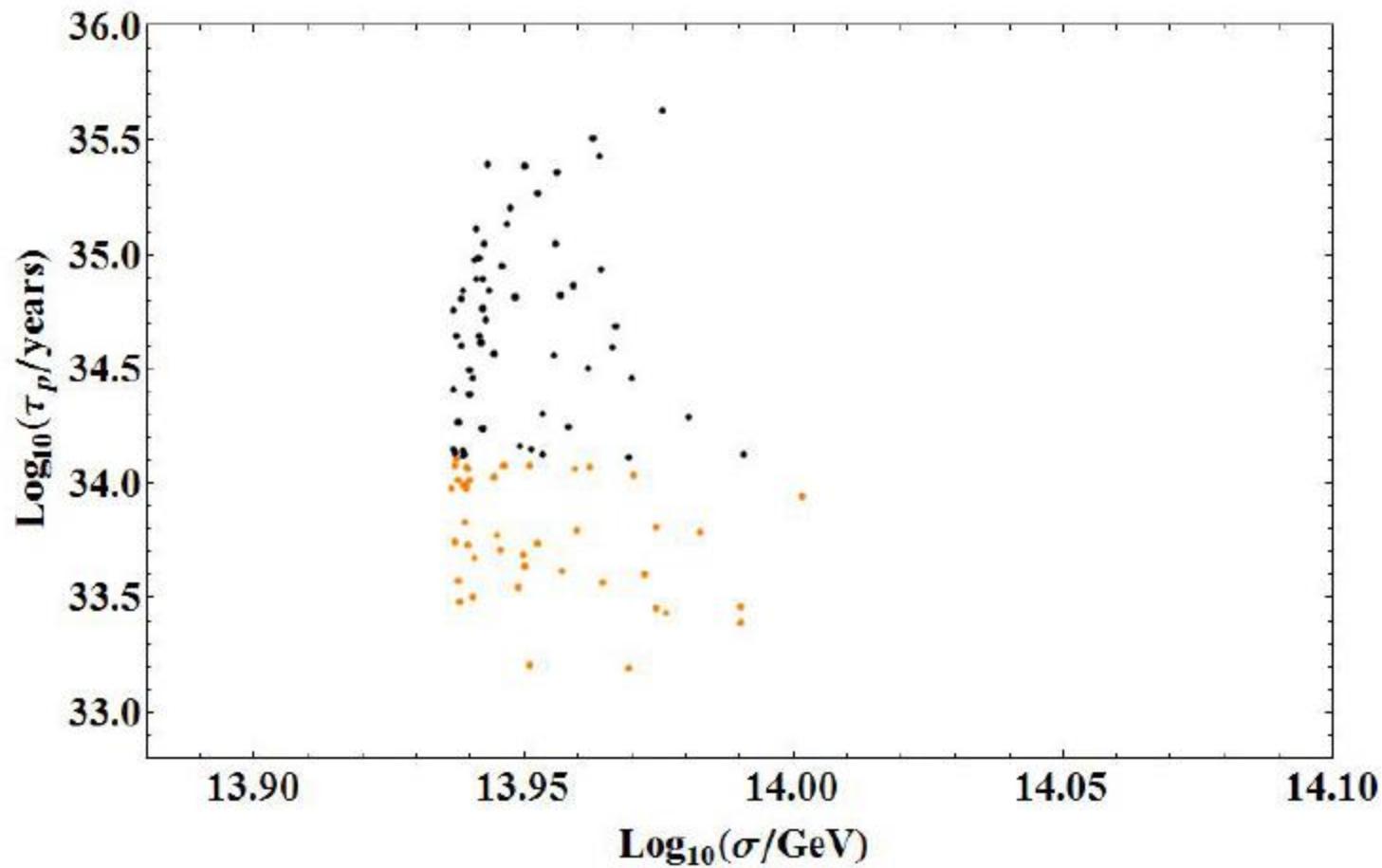
where,

$$\gamma_{L(sm)} = \left\{ \frac{23}{20}, \frac{9}{4}, 2 \right\}; \quad \gamma_{R(sm)} = \left\{ \frac{11}{20}, \frac{9}{4}, 2 \right\}; \quad \gamma_{L/R(ps)} = \left\{ \frac{15}{4}, \frac{9}{4}, \frac{9}{4} \right\}$$

## Proton Life Time

$$\begin{aligned} \Gamma^{-1}(p \rightarrow e^+ \pi^0) &\approx (8.2 \times 10^{34} \text{ yr}) \\ &\times \left( \frac{\alpha_H}{0.0122 \text{ GeV}^3} \right)^{-2} \left( \frac{\alpha_G}{1/34.7} \right)^{-2} \left( \frac{A_R}{3.35} \right)^{-2} \left( \frac{M_X}{10^{16} \text{ GeV}} \right)^4 \end{aligned}$$

# Prediction for Proton Lifetime



Non-SUSY  $SO(10)$

Babu, Khan (2015)

## Prediction of branching ratios

$$\Gamma(p \rightarrow \pi^0 e^+) \rightarrow 47\%$$

$$\Gamma(p \rightarrow \pi^0 \mu^+) \rightarrow 1\%$$

$$\Gamma(p \rightarrow \eta^0 e^+) \rightarrow 0.20\%$$

$$\Gamma(p \rightarrow \eta^0 \mu^+) \rightarrow 0.00\%$$

$$\Gamma(p \rightarrow K^0 e^+) \rightarrow 0.16\%$$

$$\Gamma(p \rightarrow K^0 \mu^+) \rightarrow 3.62\%$$

$$\Gamma(p \rightarrow \pi^+ \bar{\nu}) \rightarrow 48\%$$

$$\Gamma(p \rightarrow K^+ \bar{\nu}) \rightarrow 0.22\%$$

See also: Meloni, Ohlsson, Pernow (2019)

Perez, Murugui, Plascencia (2019)

GUT session talks by Malinksy, Susic, and others

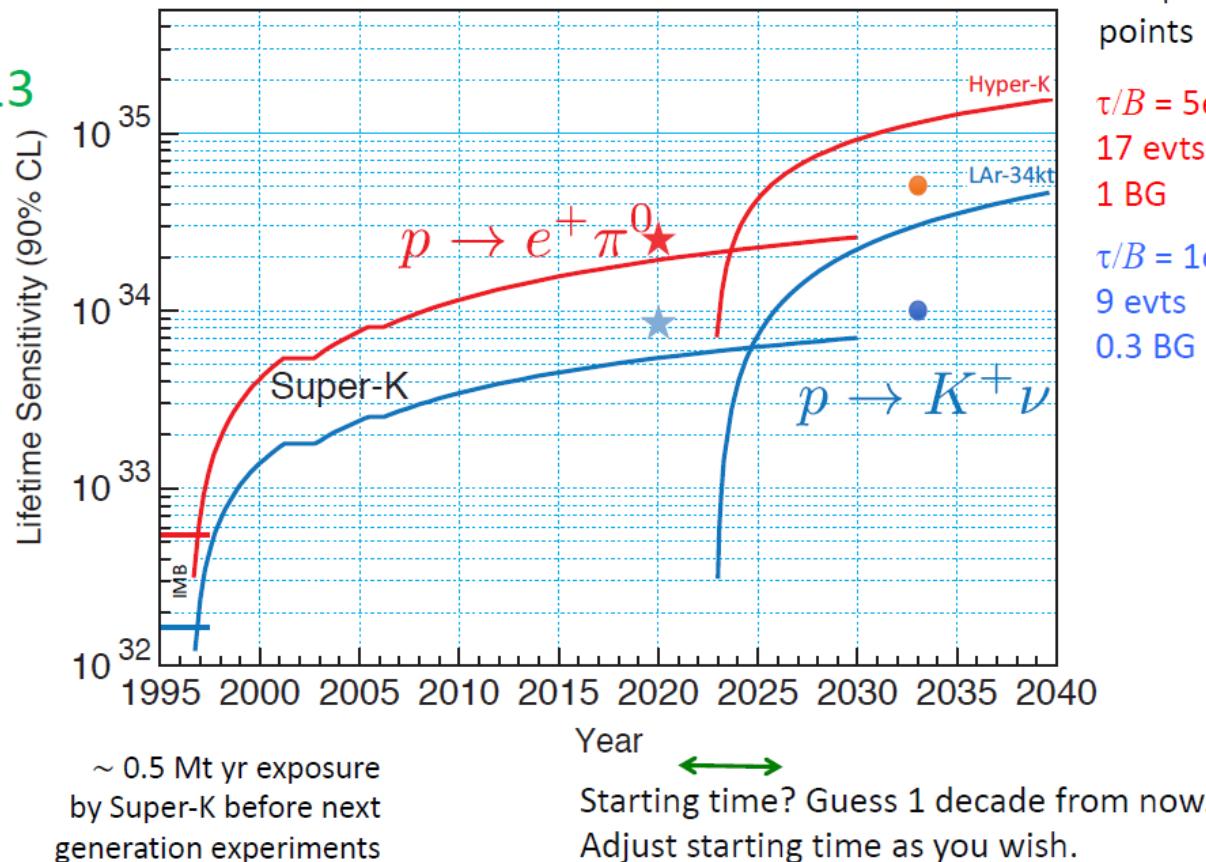
# Experimental Probes of Proton Decay



Retrospective:  
Snowmass 2013

★ = prelim. SK (2020)

10 year sample points  
 $\tau/B = 5e35$   
17 evts  
1 BG  
 $\tau/B = 1e34$   
9 evts  
0.3 BG



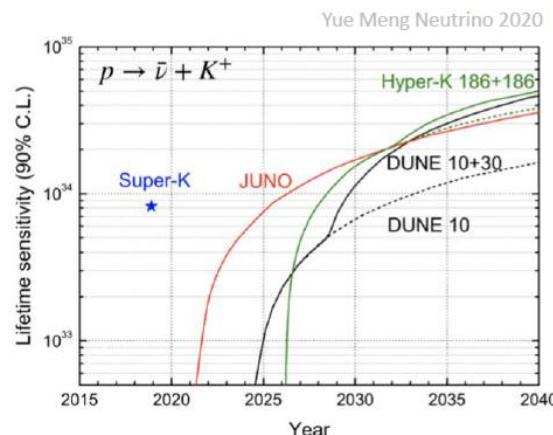
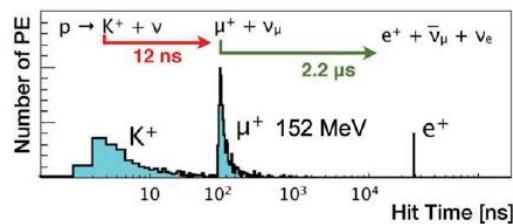
# Experimental Probes of Proton Decay

## JUNO

Primary physic goal is determination of neutrino mass ordering using reactor neutrino oscillation.  
Also geo, supernova, solar, ...

### Proton decay

- Competitive sensitivity to proton decay searches
- Triple coincidence signal



Should enter new territory before DUNE/HK turn on

$6.75 \times 10^{33}$  protons  
85% kaon decay modes x 65% signal efficiency  
Background 0.5 events in 10 years

$$\frac{\tau}{B} > 1.9 \times 10^{34}$$



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# Proton Decay and Long-term Fate of the Universe

Last stage of stellar evolution controlled by proton decay

At large time scale, evolution of white dwarfs and neutron stars driven by proton decay

$$p \rightarrow e^+ \pi^0$$

$\pi^0$  decays to two photons, and  $e^+$  annihilates with  $e^-$

$$p + e^- \rightarrow \gamma + \gamma + \gamma + \gamma$$

White dwarf powered by proton decay generates 400 Watts of power

Proton decay becomes dominant power source at

$$t \approx 30 \times \tau_p \approx 10^{36} \text{ yrs.}$$

# Conclusions

- SUSY GUTs are well motivated BSM candidates
- Proton decay is their hallmark prediction
- SUSY GUTs are under stress from proton decay limits
- Both  $p \rightarrow \bar{\nu}K^+$  and  $p \rightarrow e^+\pi^0$  should be within reach of HyperK, DUNE and JUNO experiments
- Light gauginos and heavy scalars preferred for LHC