

Einstein-Telescope: German Community Meeting

Stabilized High-Power Lasers for 2nd and 3rd Generation Gravitational Wave Detectors

Benno Willke

(for the laser team at AEI and Laser Zentrum Hannover)

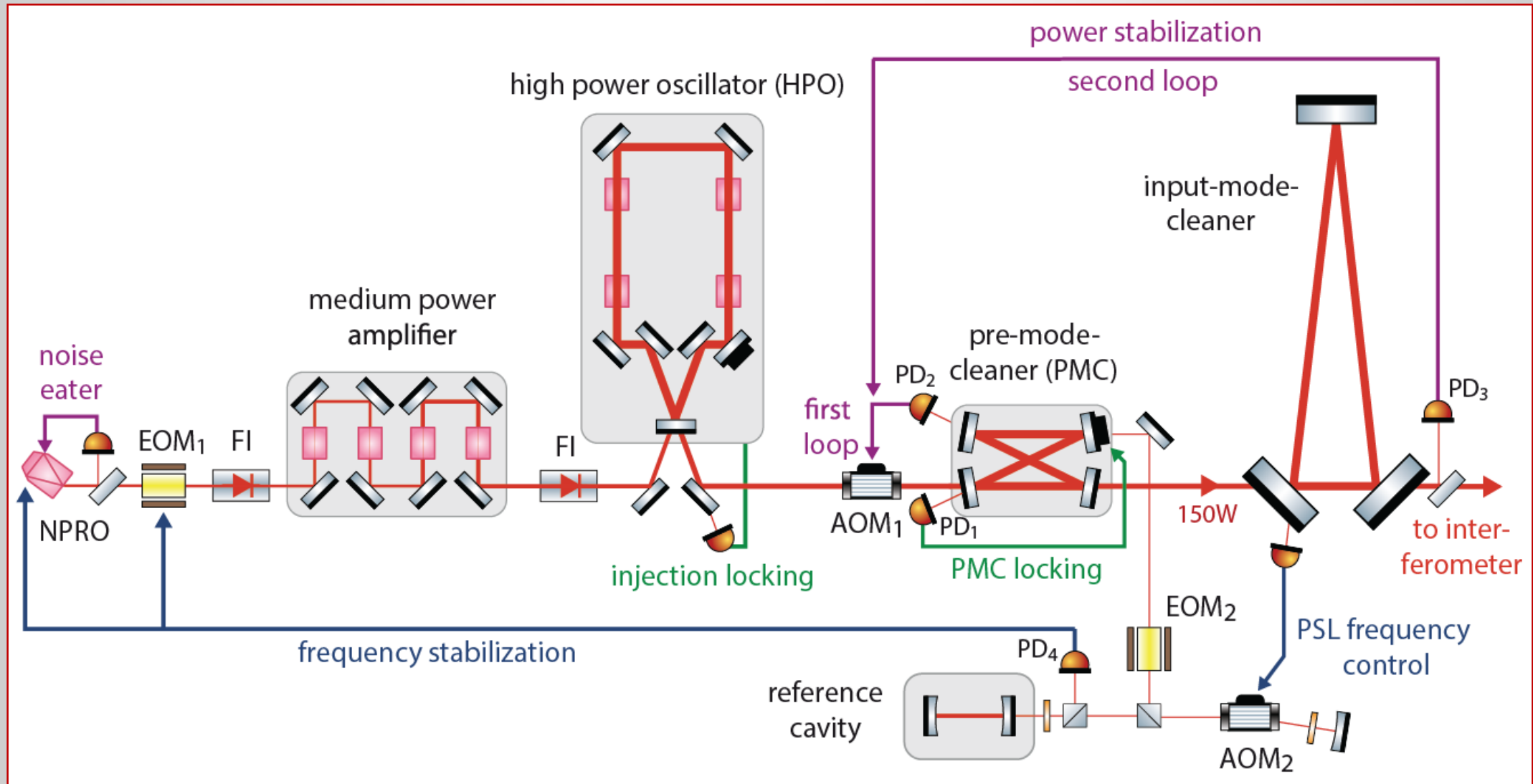
17 Jan 2020 (Hannover)

General Requirements

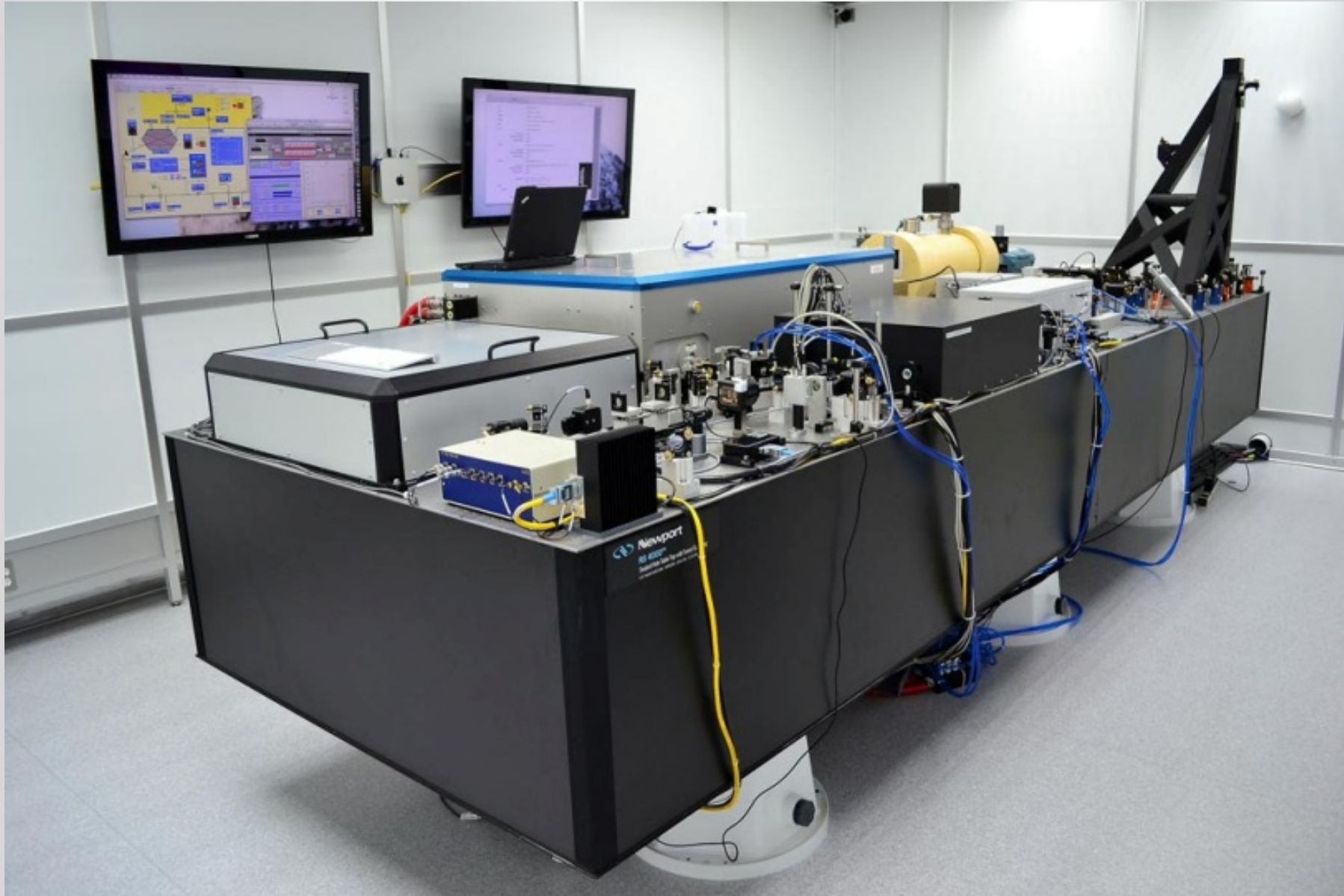
- **Several hundred Watt power** (cw, single-frequency, linearly-polarized)
 - To enable high circulating power in interferometer arms
- **High spatial purity and low beam jitter**
 - Good coupling to input-modecleaner
 - Low shot noise on sensors for laser and input-modecleaner stabilization
- **Low free running noise** (in laser power and frequency)
 - Acceptable stabilization effort (loop gain and cross-couplings)
- **Low-noise sensors** (for laser power and frequency)
 - To achieve required stability for light entering the input modecleaner
- **Fast actuators** (for laser power and frequency) with large range
 - To allow for required loop gain in stabilization control loops
- **High robustness and reliability** with low maintenance requirements



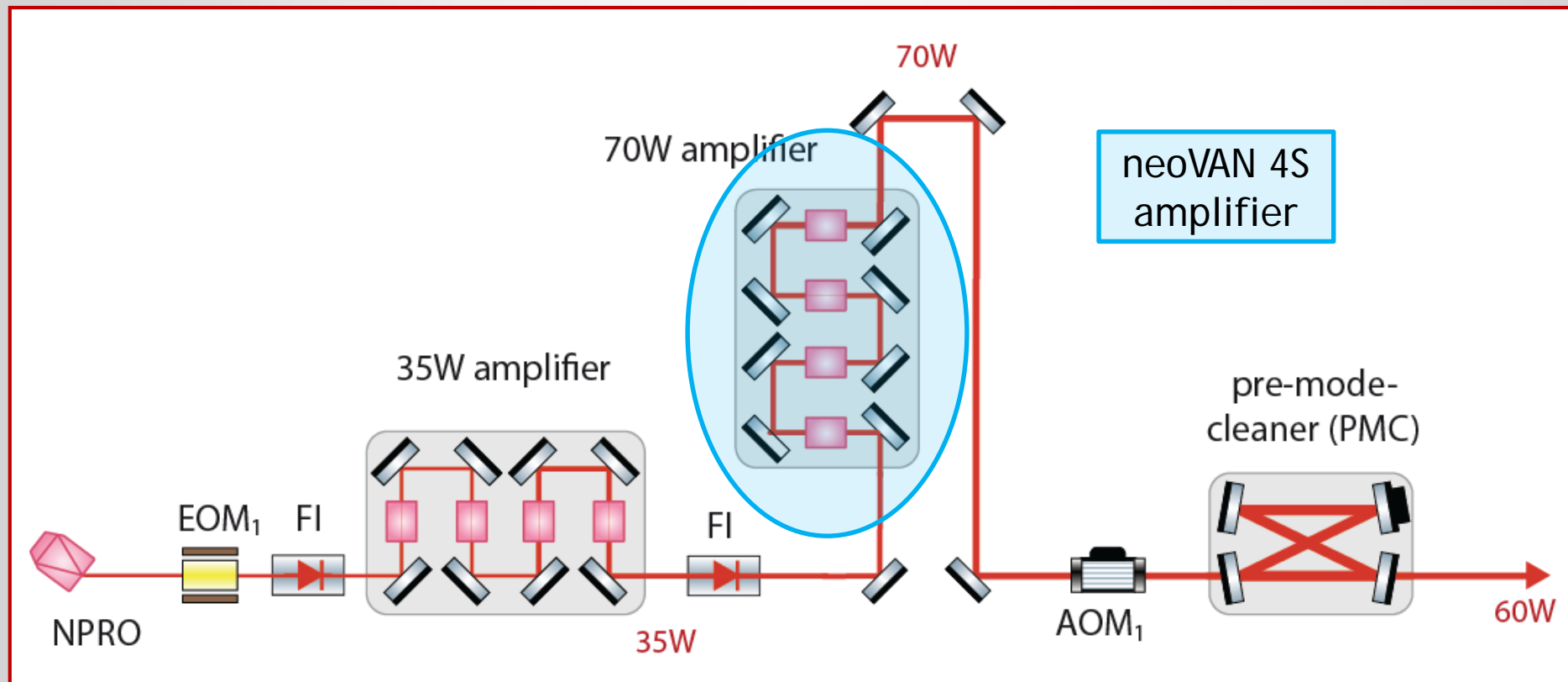
Advanced LIGO Pre-Stabilized High-Power Laser



Advanced LIGO Pre-Stabilized High-Power Laser

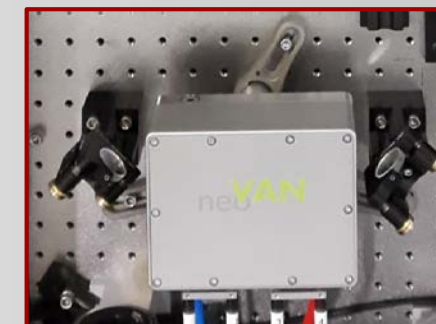
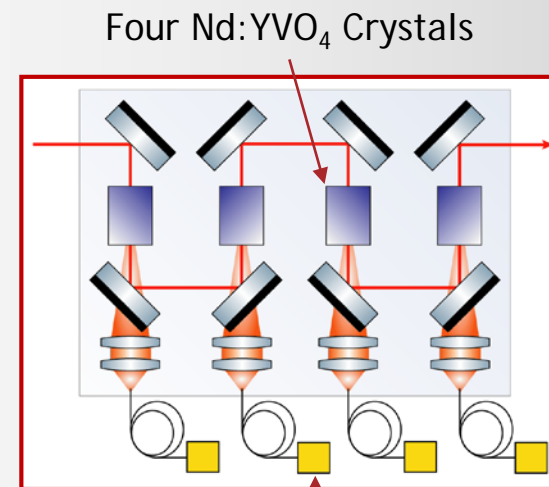
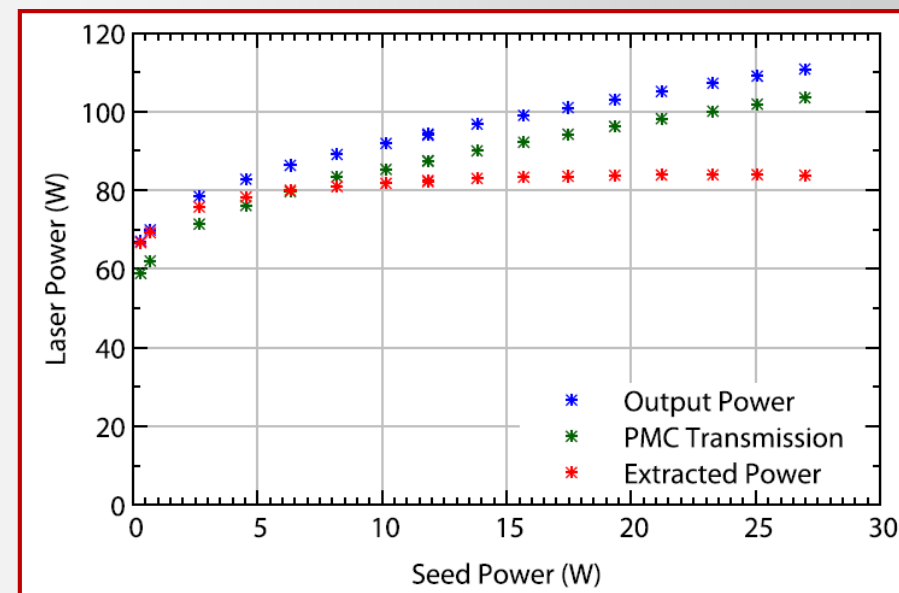
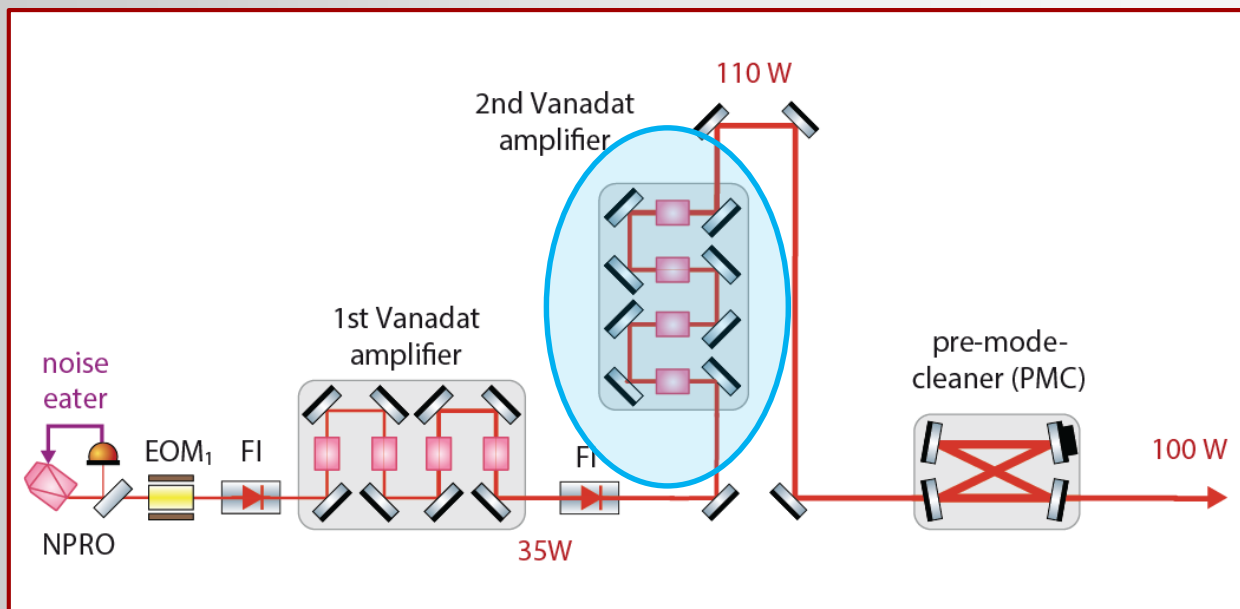


Modified aLIGO PSL for Observing Run O3



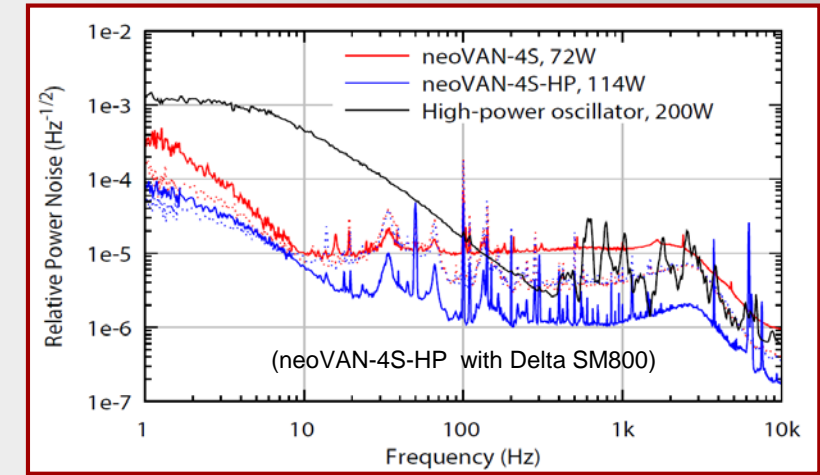
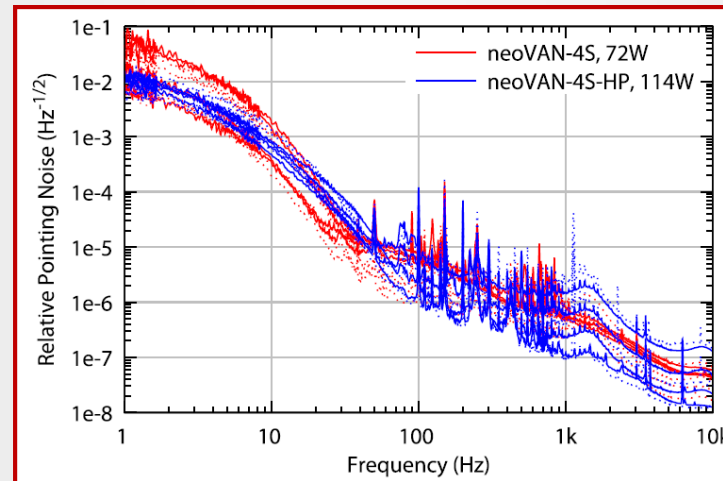
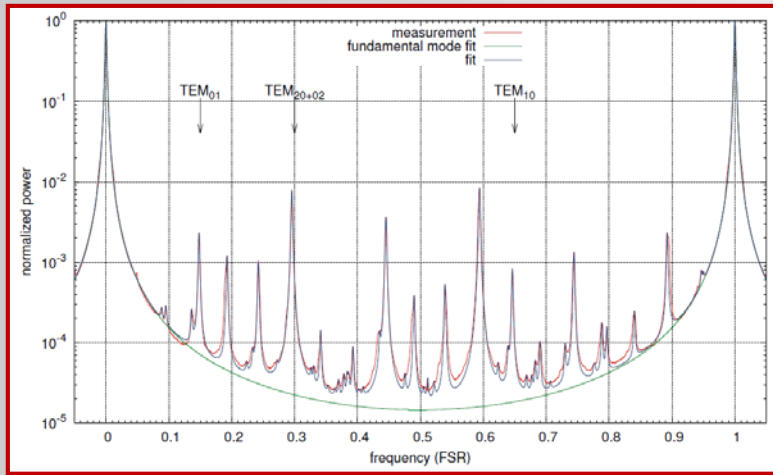
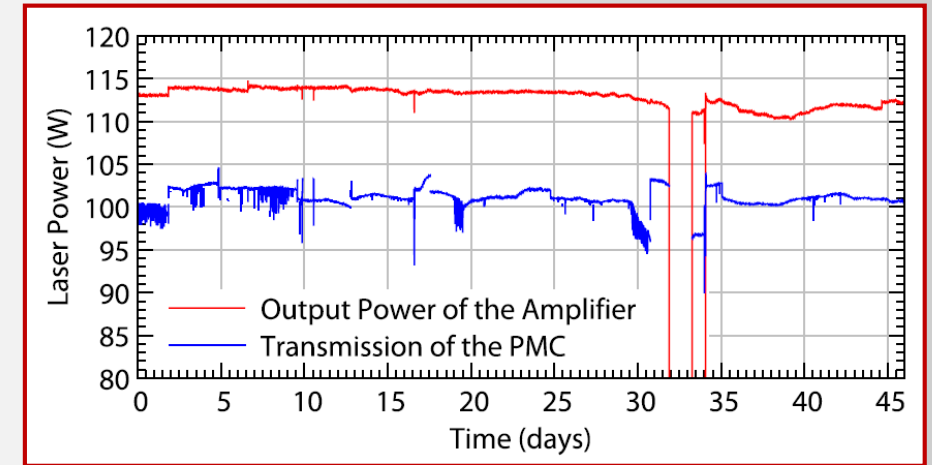
New Generation: neoVAN-4S-HP Amplifier

- First tested for Adv. Virgo at AEI
- Longer pump wavelength
 - 878nm vs. 808nm
 - VBG stabilized pump diodes
- 80W power extraction

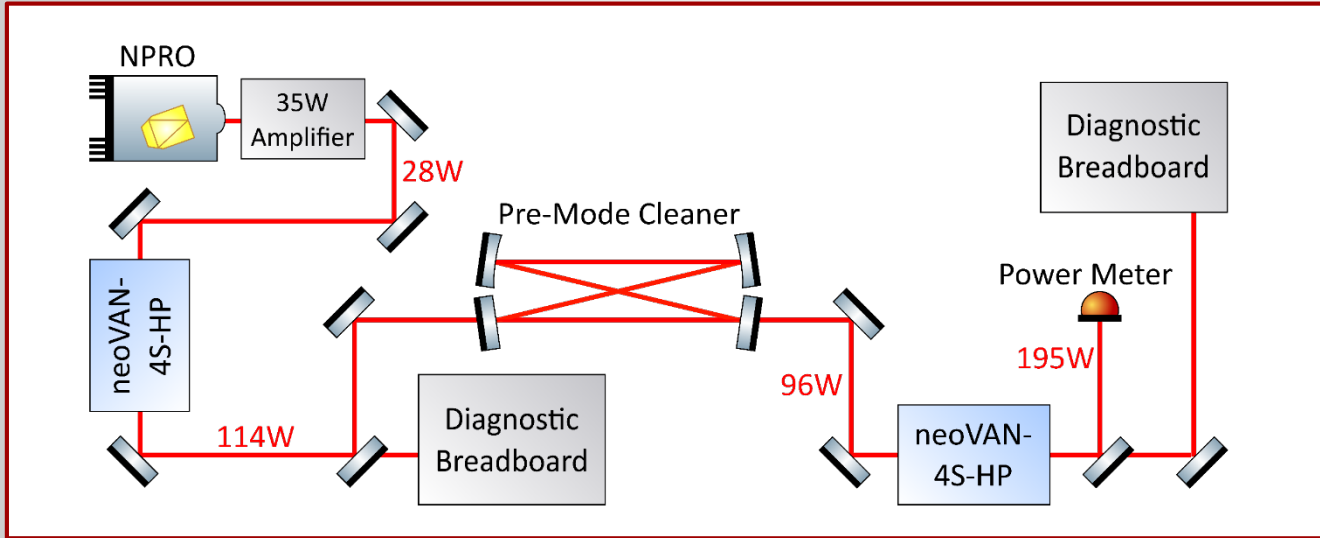


Performance of Nd:Vanadat Amplifier Systems

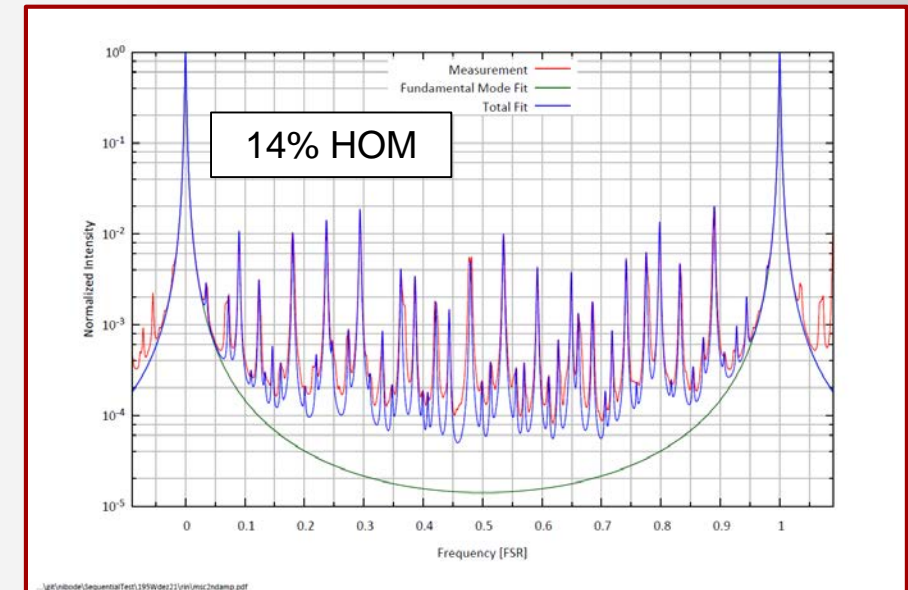
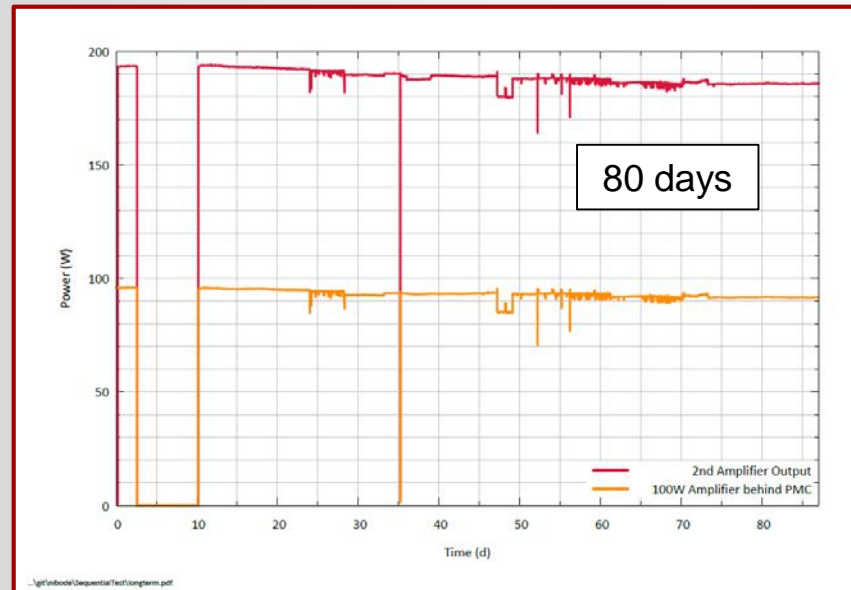
- 114 W output power for 27 W seed
- 3% higher order modes (similar to seed)
- 102W after filter cavity (PMC)
- Reduces beam pointing wrt. HPL
- significantly lower power noise than HPO
- 45 day stable operation with cold-start without realignment



Three Sequential Vanadat Amplifiers

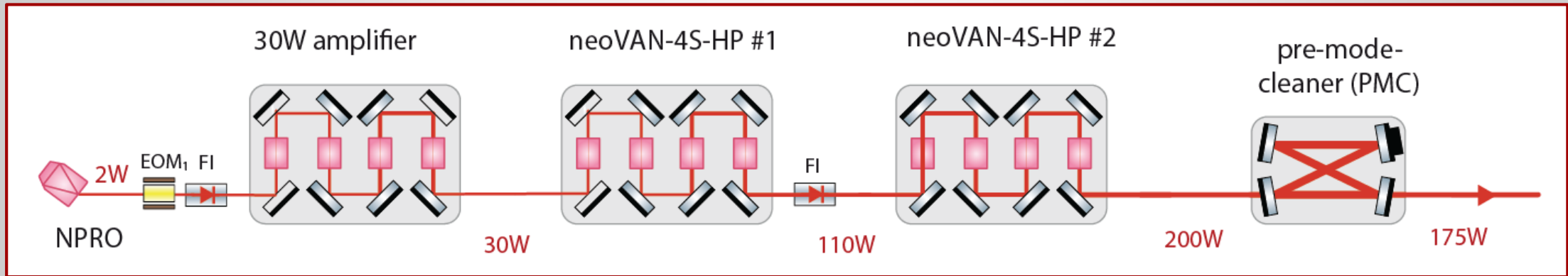


- Sequential neoVAN-4S-HP amplifiers deliver **195W**
- similar noise performance as two-amplifier system
- Large higher-order mode content (14%) indicates limit of power scaling

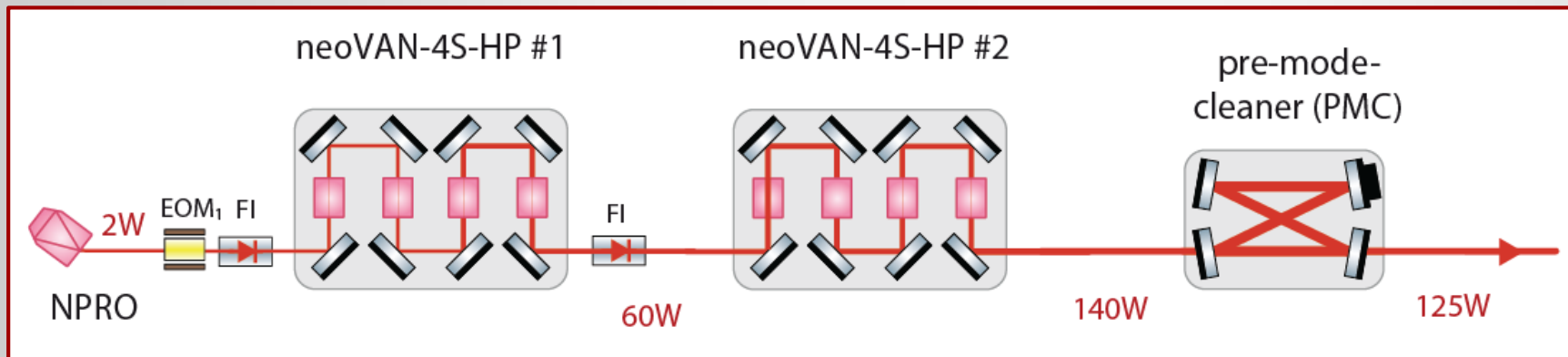


O4 Laser System: Option 1

Test of sequential neoVAN-4S-HP amplifiers at AEI



Advanced LIGO laser system for O4



Requirements For Next Generation GWDs

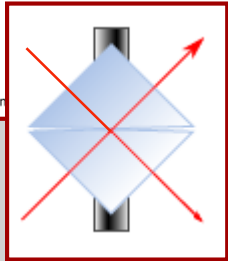
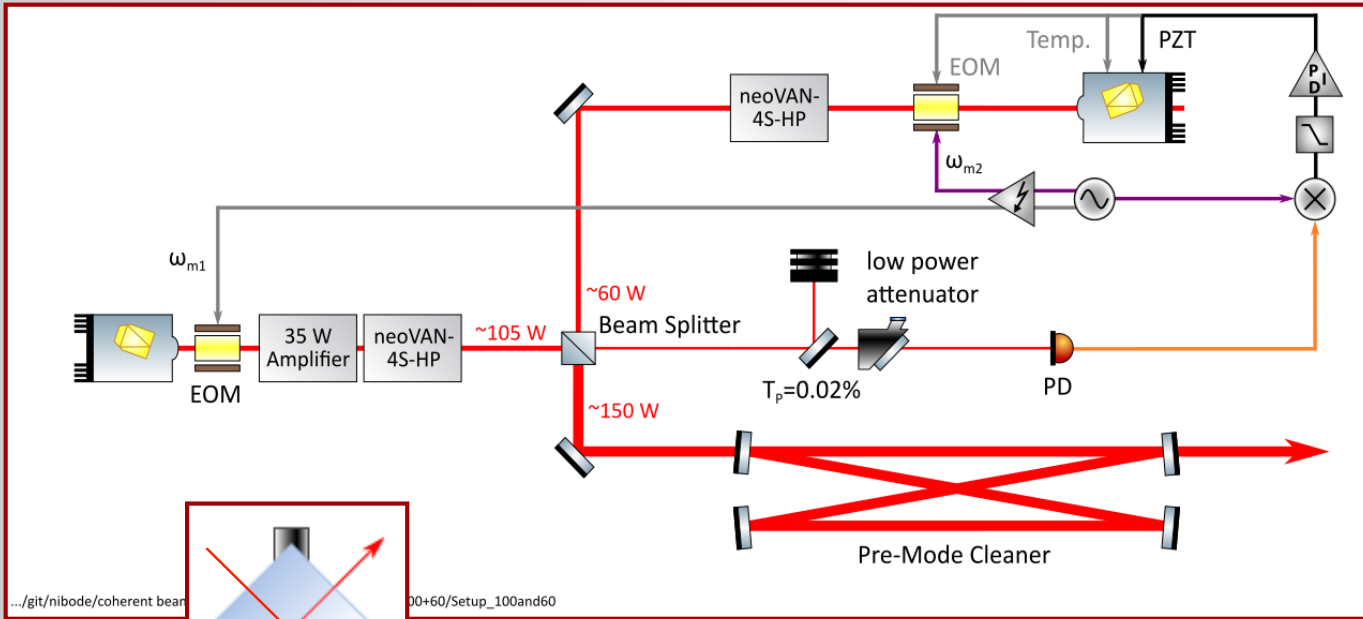
Project	Wavelength	Power	Spatial mode
Einstein Telescope low f	1550 nm	3 W	HG ₀₀
Einstein Telescope high f	1064 nm	500 W	HG ₀₀
Voyager	≥ 1550 nm	200W	HG ₀₀
Cosmic Explorer (II)	≥ 1550 nm	1kW	HG ₀₀
Asia- Australia GWD	1064 nm	200 W	HG ₀₀

- Single-mode, single-frequency, linearly polarized, low free running noise
- Stabilization: factor 10 better than second generation GWDs
- Fast actuators with large range and low cross coupling
- High reliability
- Low maintenance downtime and costs

All numbers are still subject to change !

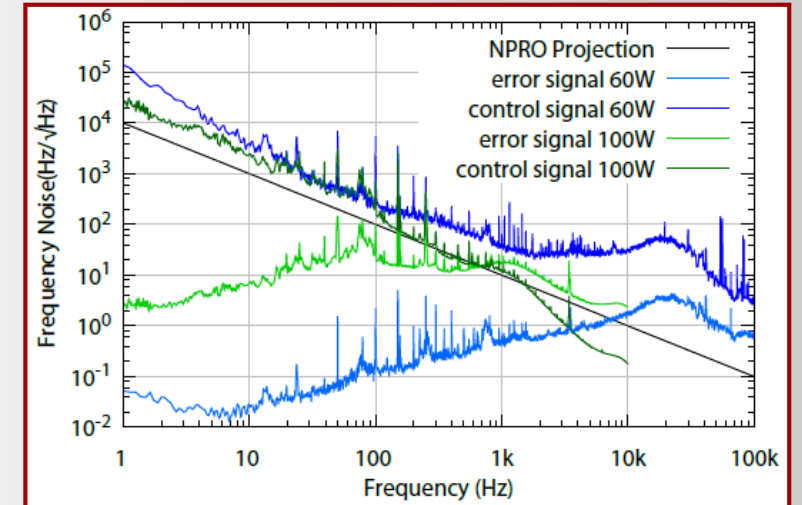
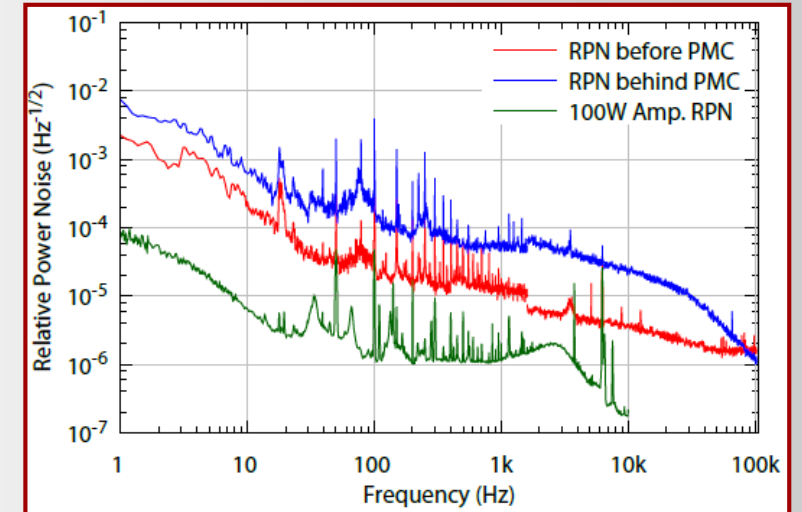
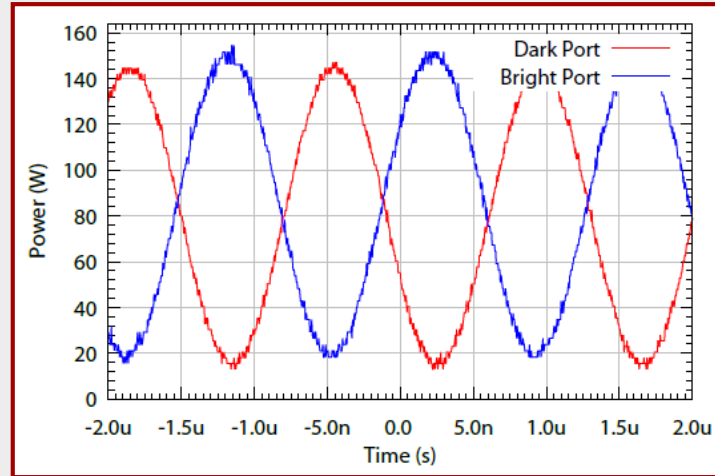


Coherent Beam Combination

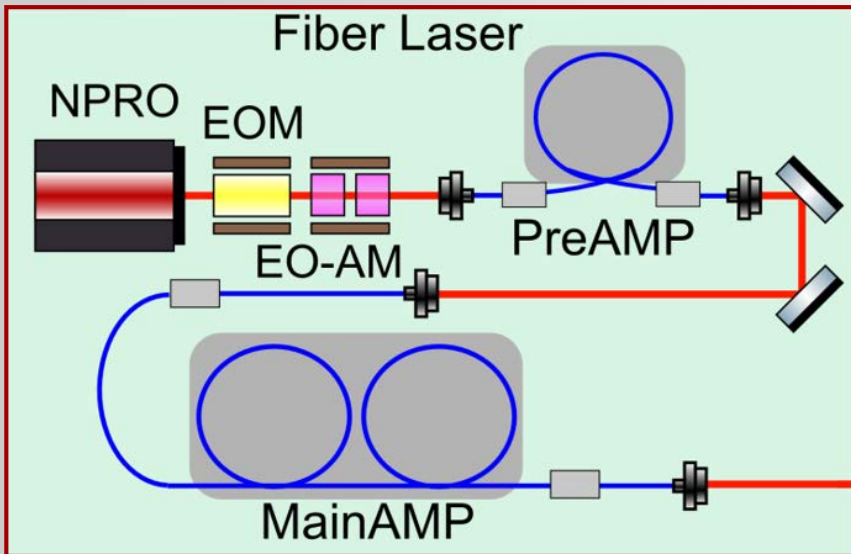


FTIR Actuator:
Variable beam splitter based on the effect of
frustrated total internal reflection

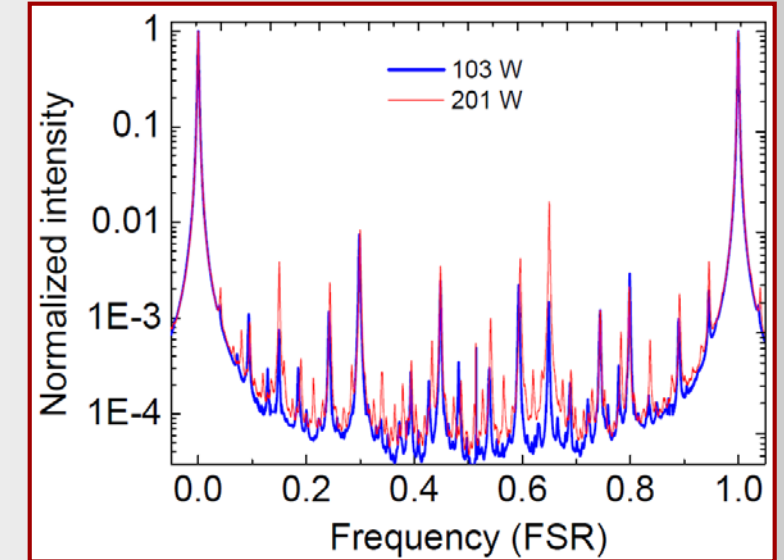
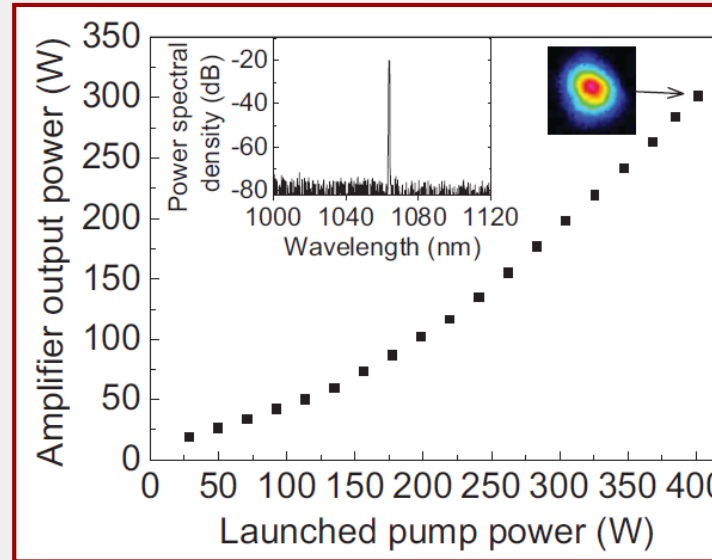
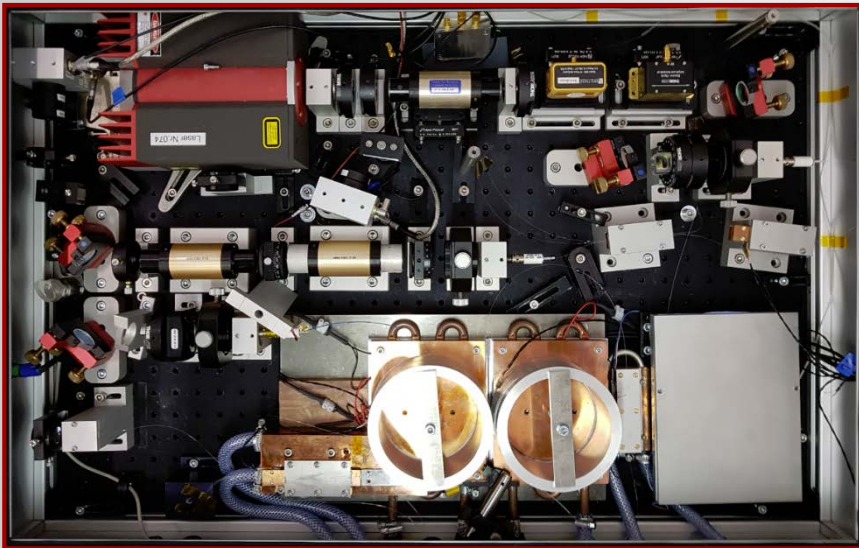
combination of 60W and 105W
⇒ 150W laser beam



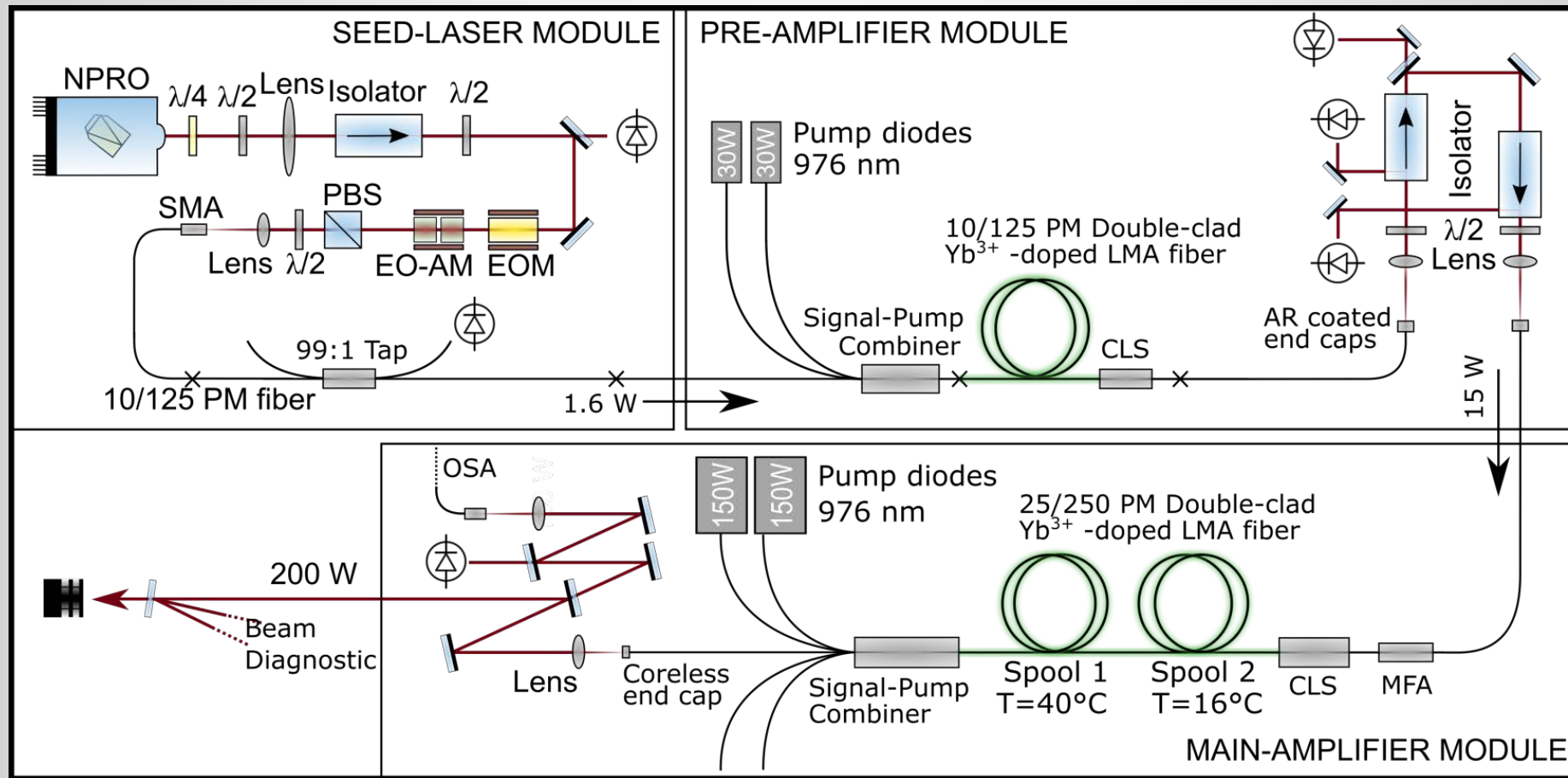
1064nm Fiber Amplifier – Proof of Principle



- two stage fiber amplifier with NPRO seed
- good spatial profile
- output power up to 300W
- no long-term reliable operation

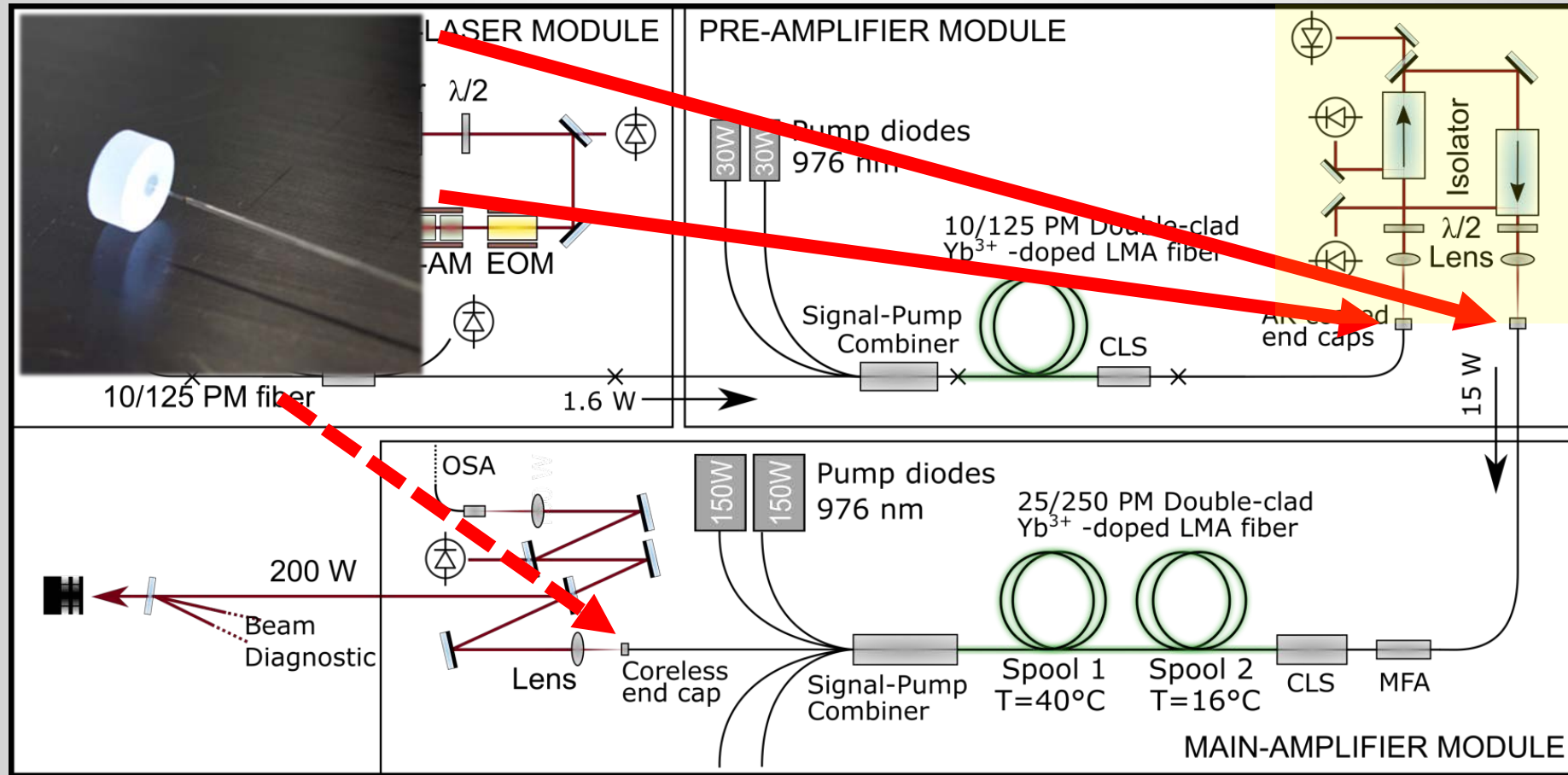


1064nm Fiber Amplifier - Engineering Prototype



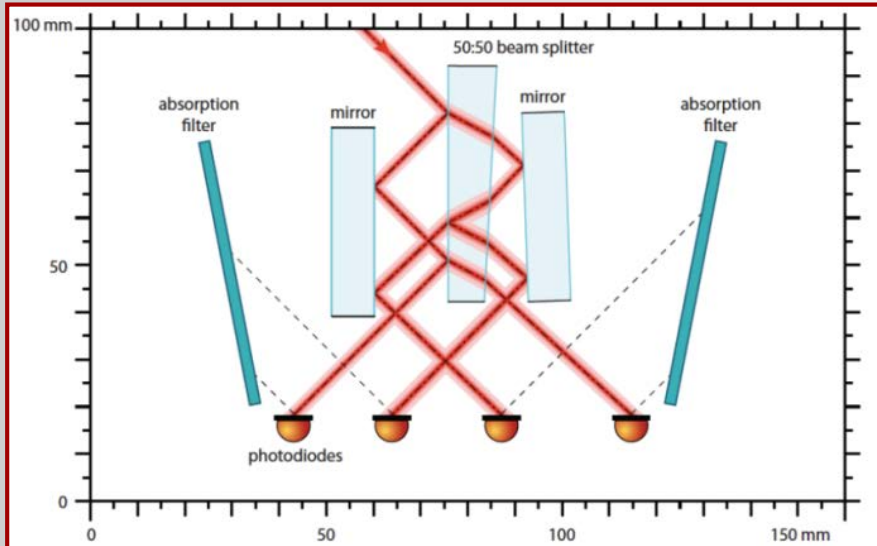
new design with reduced number of splices, reduced reflectivity and higher optical isolation

1064nm Fiber Amplifier Engineering Prototype

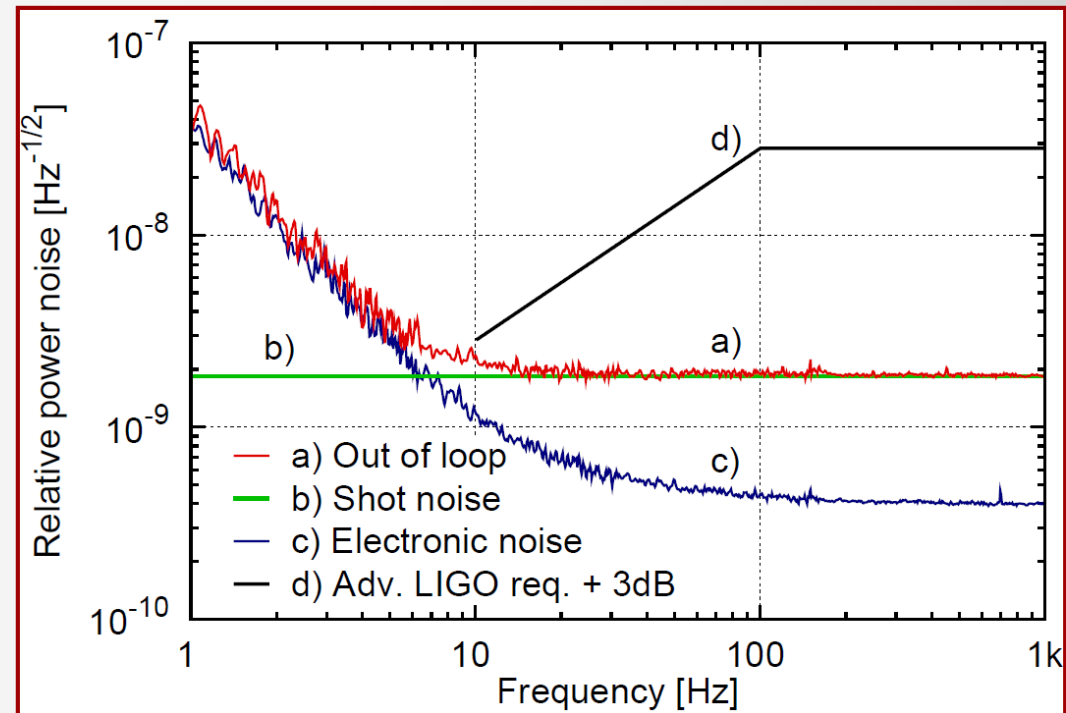
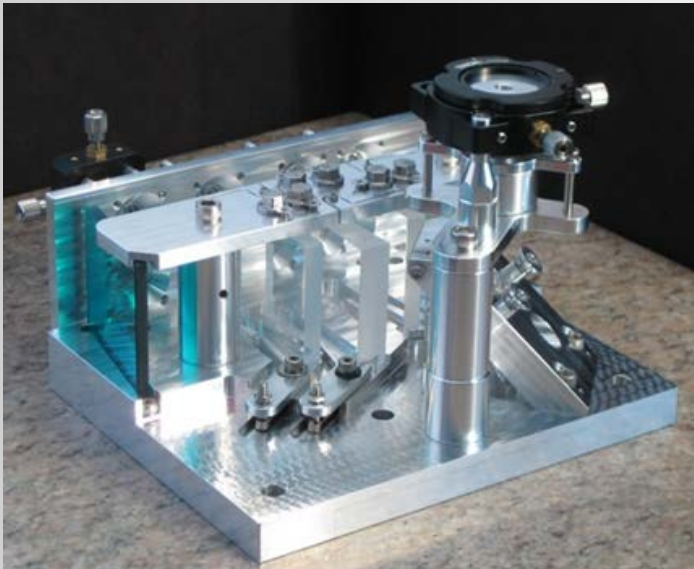


use AR coated fused-silica substrates at fiber-air interface in high-power FI unit and at high power output port

Power Stabilization

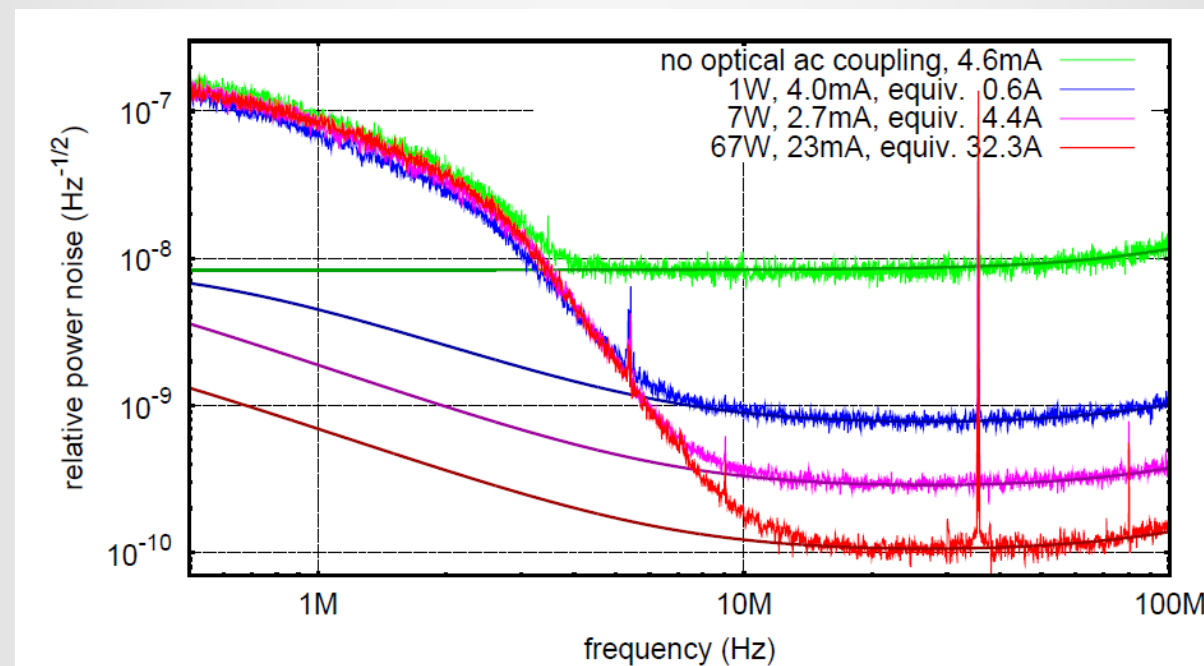
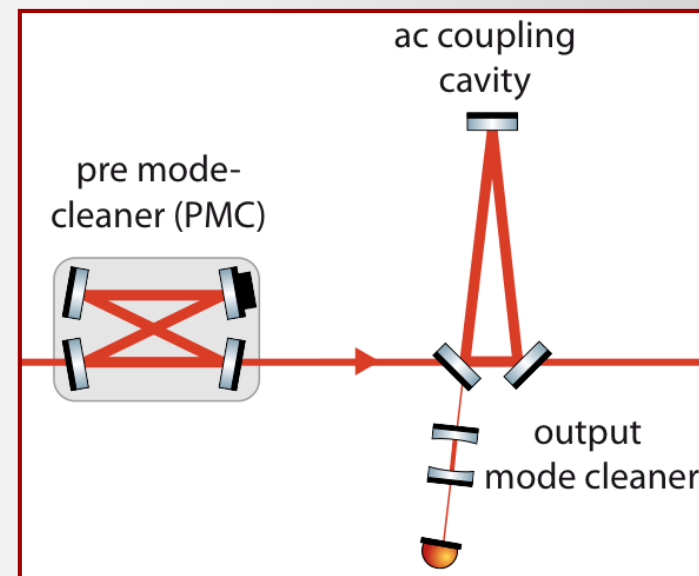


- beam is split onto two sets of four photodiodes (in-loop, out-of-loop)
- all stray beams are steered to absorption filters
- signal conditioning filters to reduce electronic noise coupling
- achieved $2 \times 10^{-9} \text{ } 1/\sqrt{\text{Hz}}$ stability

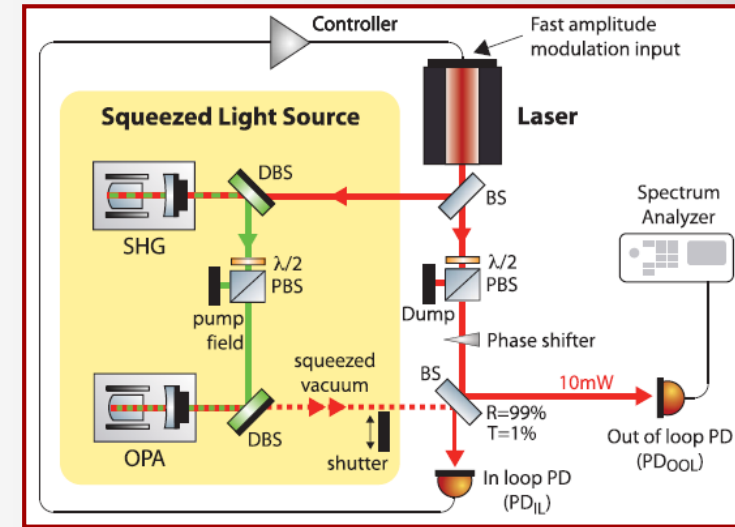
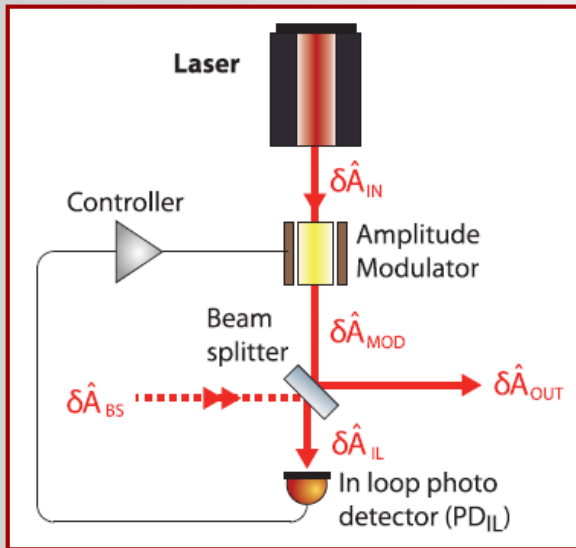


OAC with 32.2A Equivalent Photocurrent

- Power noise of 150W beam in an optical AC coupling scheme with output modecleaner
- Achieved sensing sensitivity
$$RPN = 1 \times 10^{-10} \text{ } 1/\sqrt{\text{Hz}}$$
- This corresponds to shot noise of 32.2A photocurrent



Squeezing Enhanced Power Stabilization



Simulation:

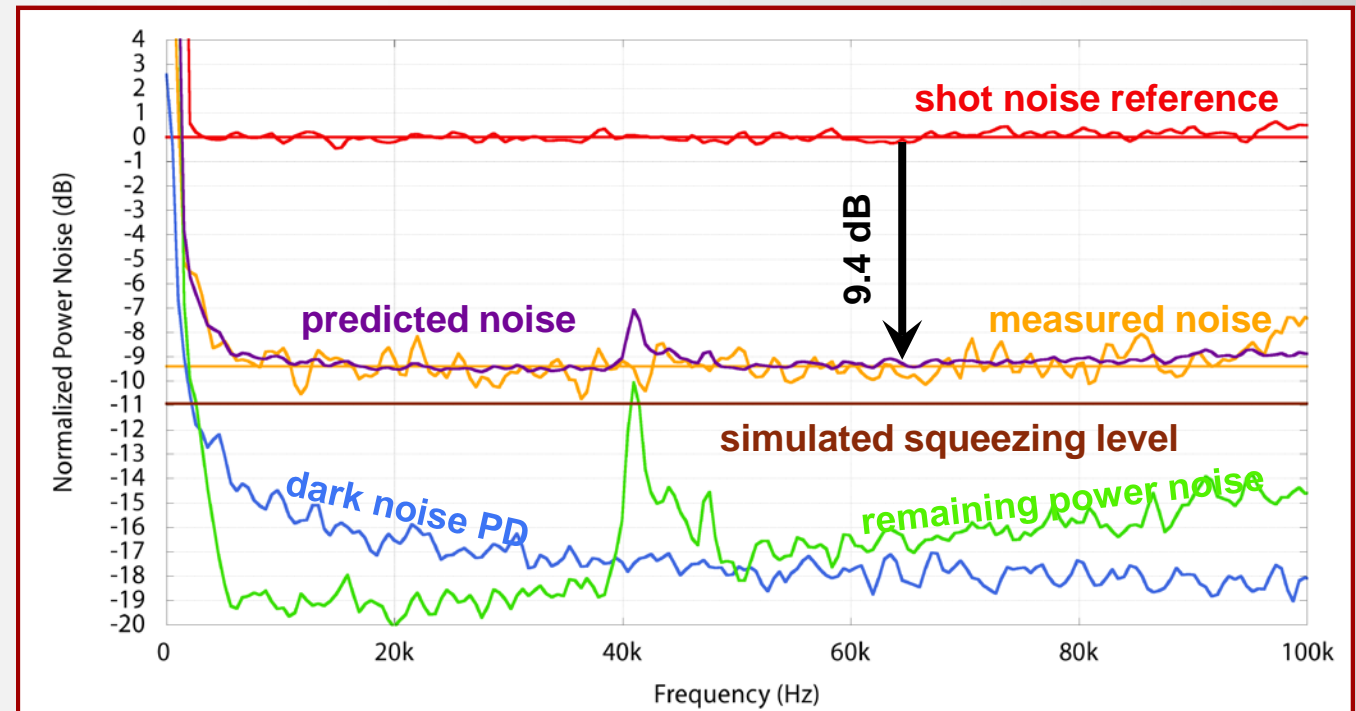
9.5 mW pump, 5.07 % losses, 10.9 dB

Prediction:

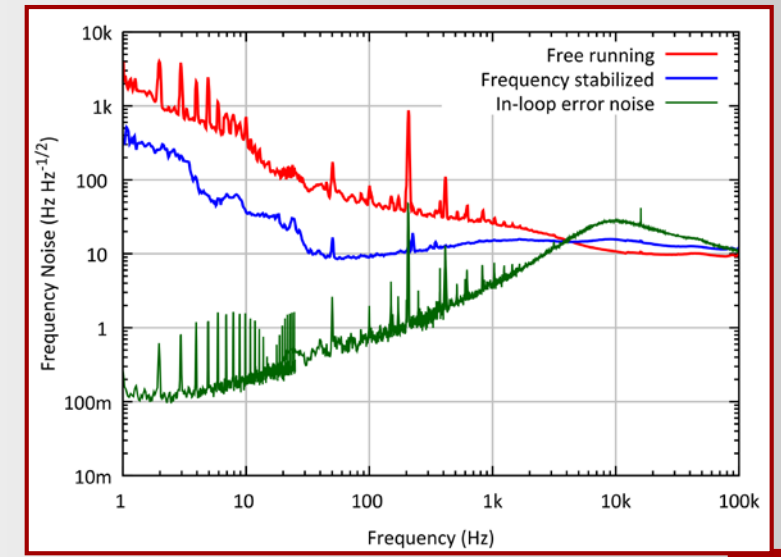
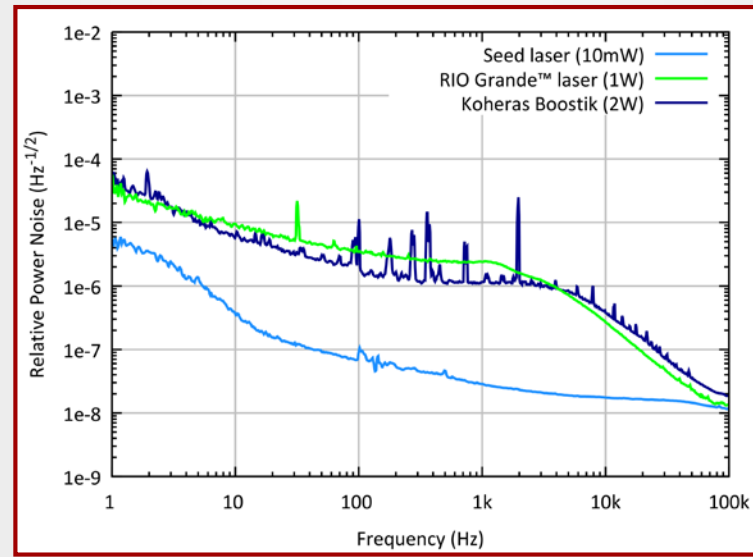
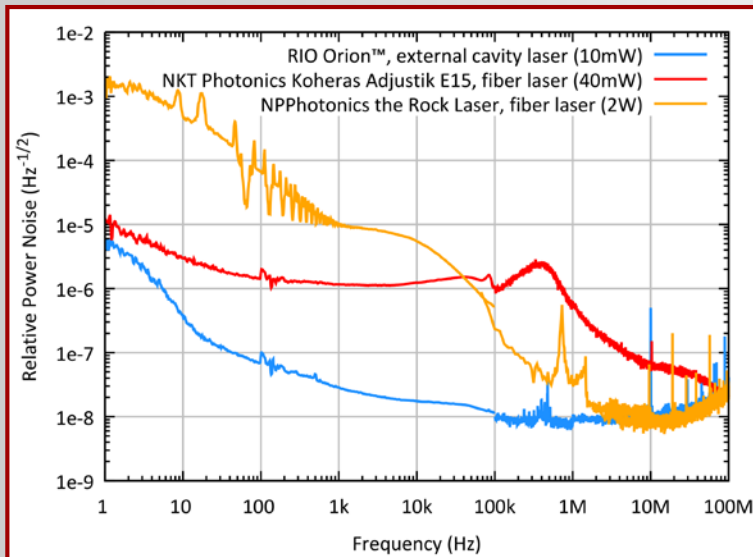
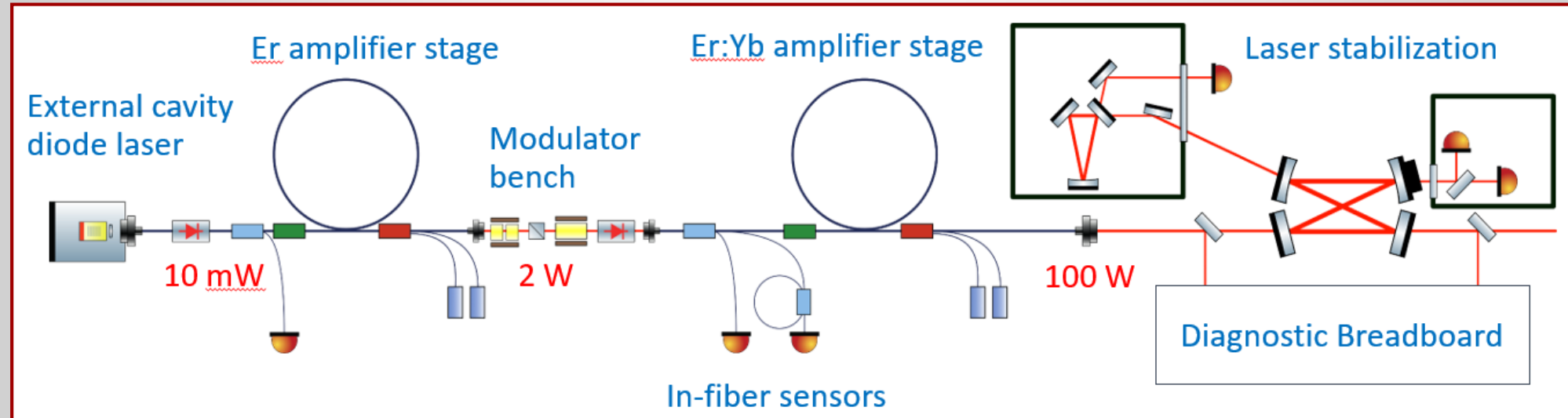
10.9 dB, +PDIL DN, +Stab IL noise

Limitations:

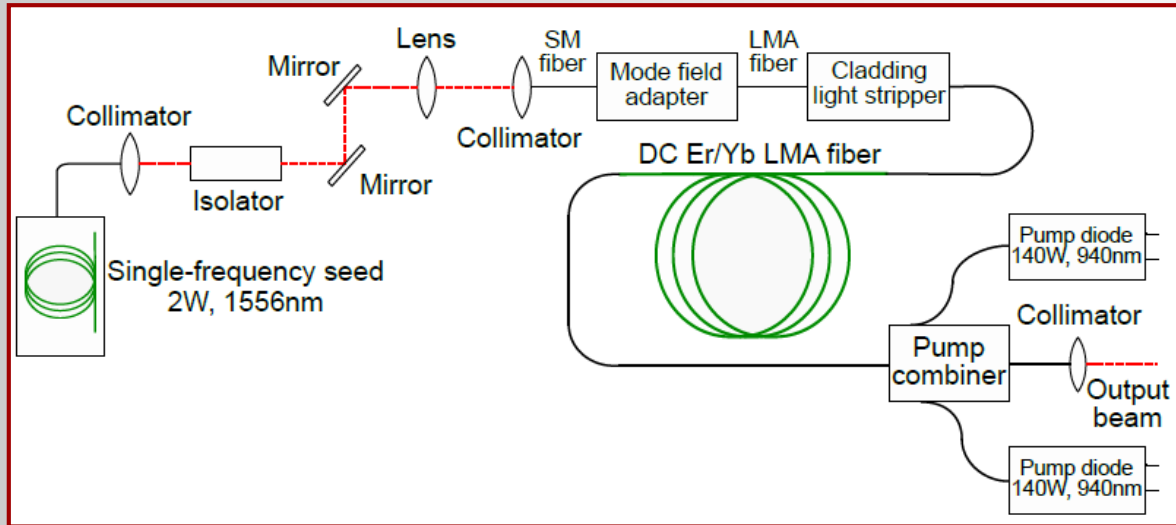
Losses, PD_{IL} DN, Loop Gain



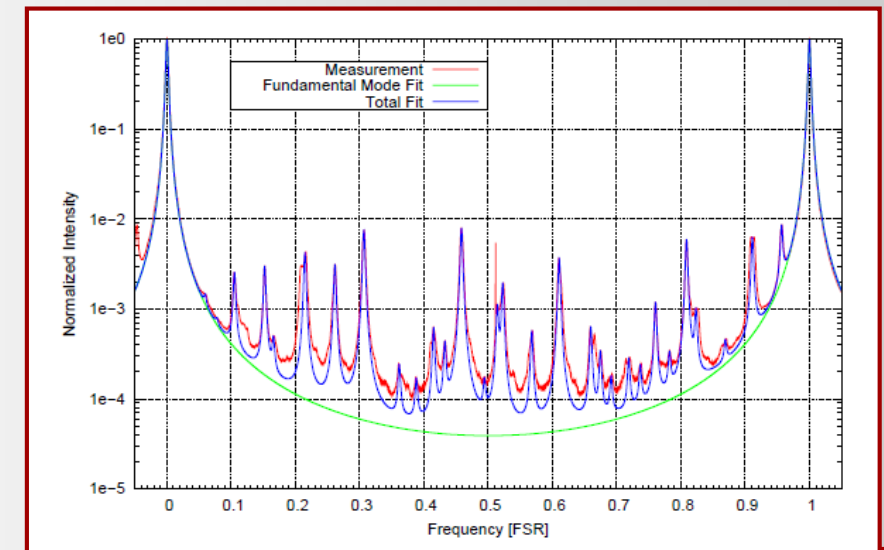
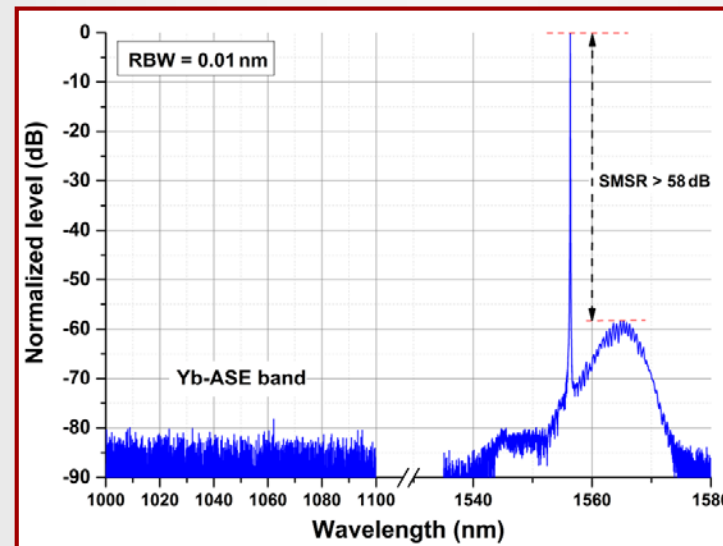
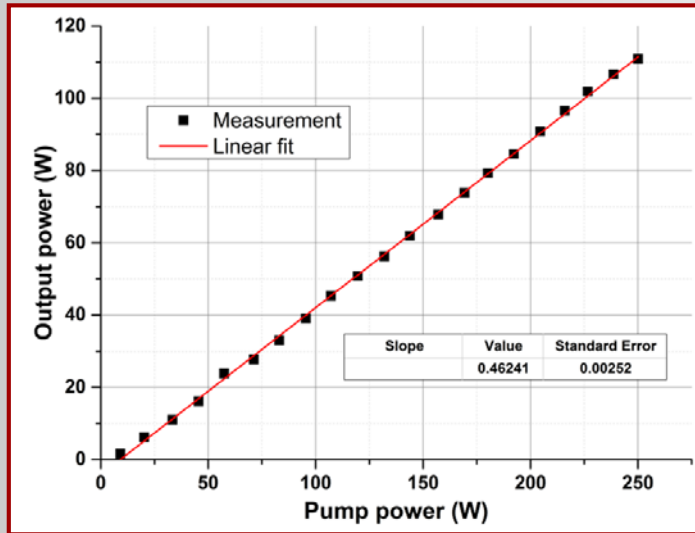
1550nm Laser Stabilization



1550nm LZH Fiber Amplifier

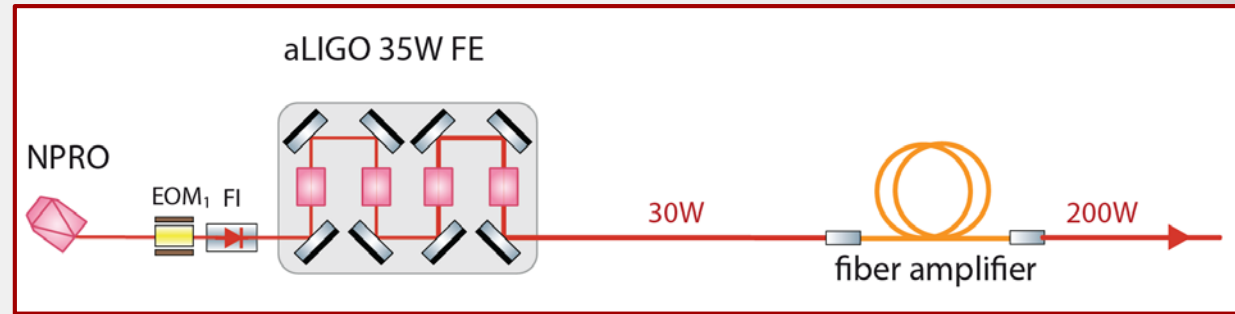


- Off-peak pumping at 940nm
- Linear slope up to 110W
- At 110W no SBS (no excess power noise in MHz region)
- No Yb-band ASE and good suppression of Er-band ASE (~60dB)
- Output power limited by available pump power

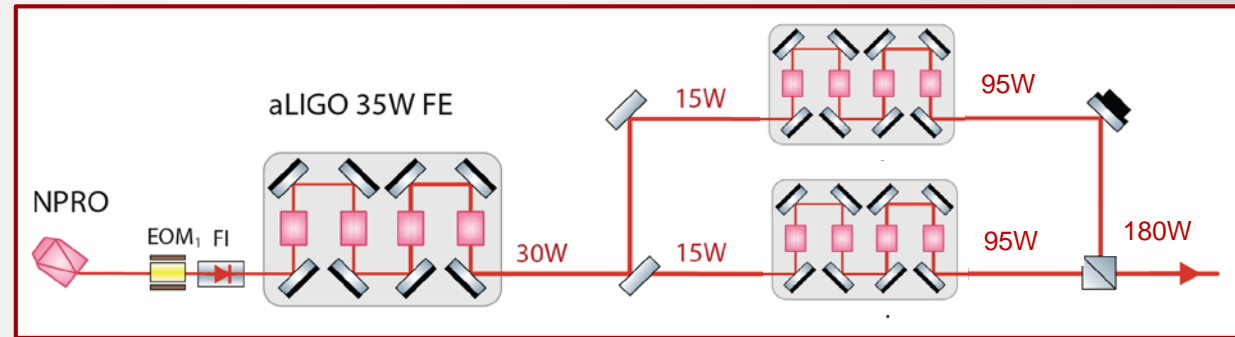


Future Plans

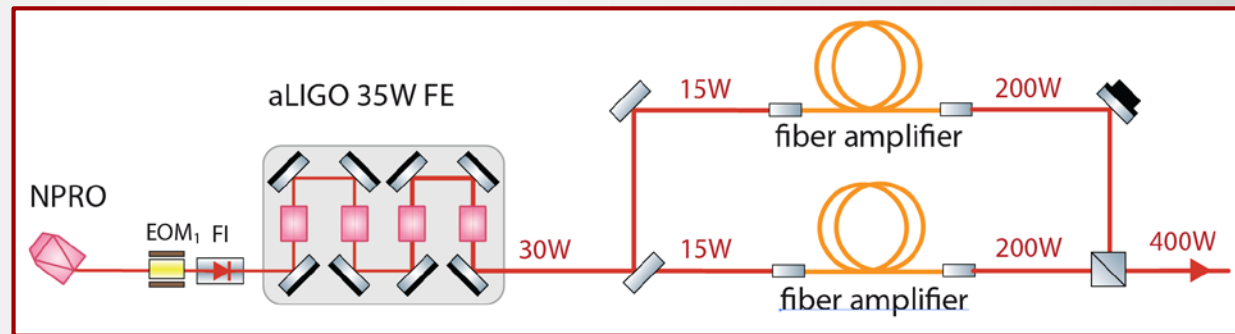
Combine aLIGO 35W laser with fiber amplifier



Coherent beam combination of two neoVAN-4S-HP



Coherent beam combination of two fiber amplifiers



Challenges

- Design and fabricate robust and reliable 250-300 W (fiber) amplifier for 1064 nm, 1.5 μ m and 2 μ m
- Low-noise coherent combination of such laser with pure spatial beam profile, low beam jitter and low power noise
- Power stabilization at relative power noise level of $RPN = 10^{-10} \text{ } 1/\sqrt{\text{Hz}}$
- Transfer laser stability to suspended reference frame of GWD

