

# Flavor Physics at LHC, LFUV in B-decays

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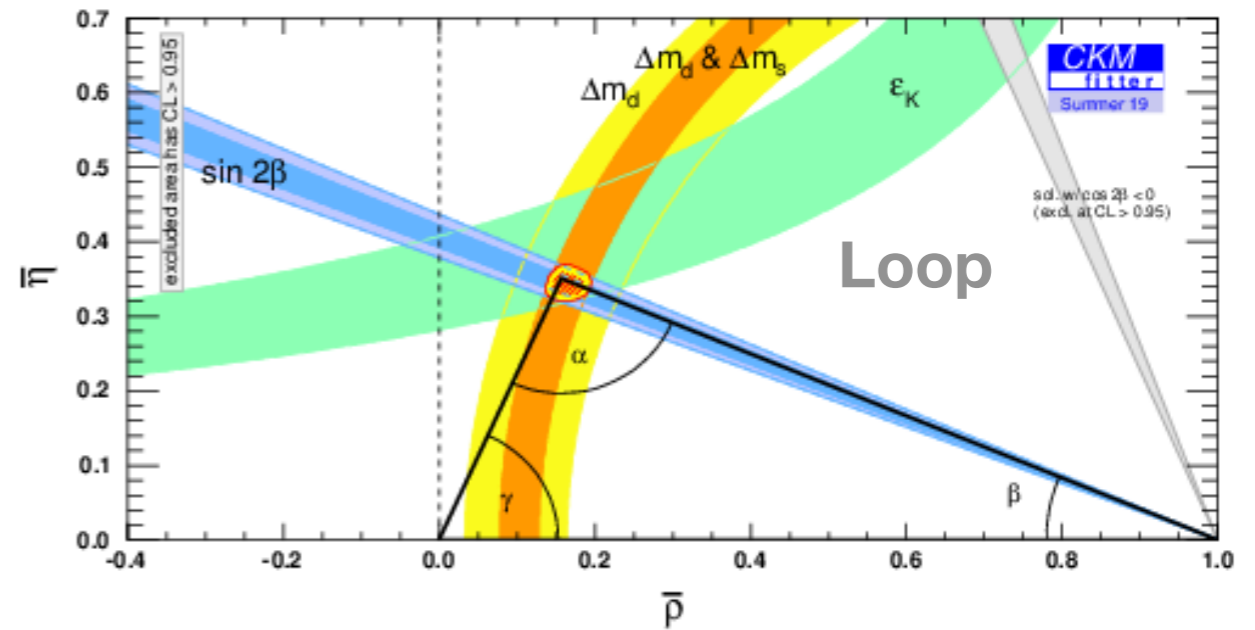
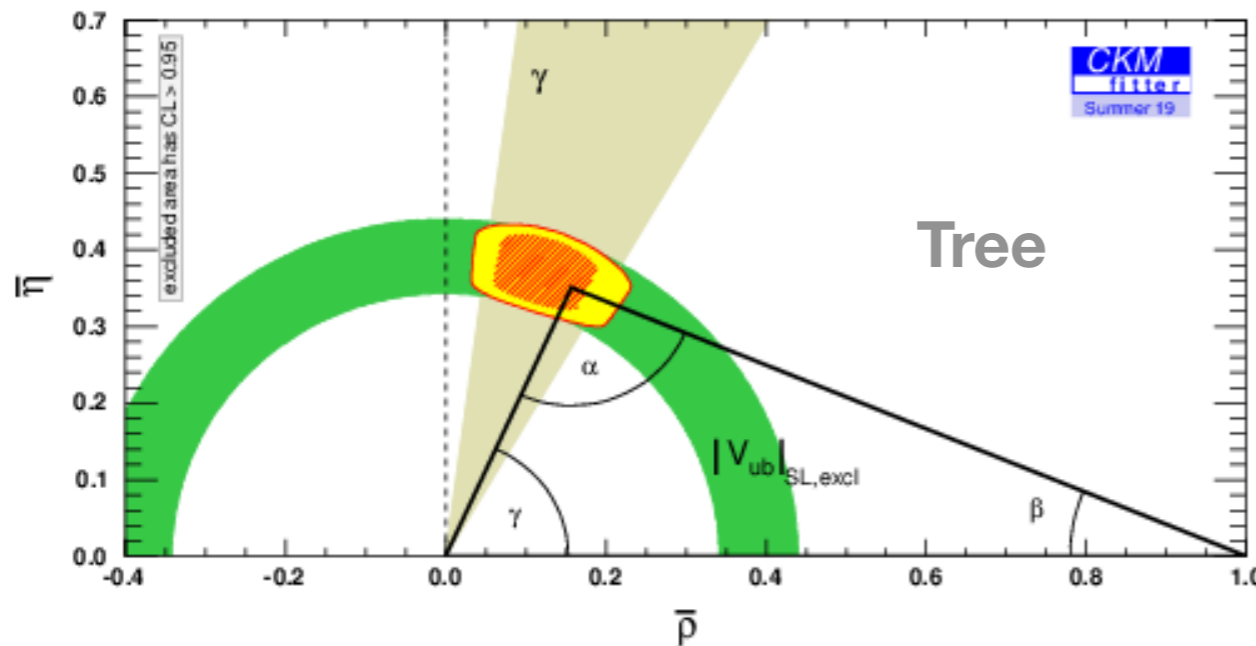


based on works done with A. Angelescu, I. Doršner, S. Fajfer, D. Faroughy,  
F. Jaffredo, N. Košnik, O. Sumensari

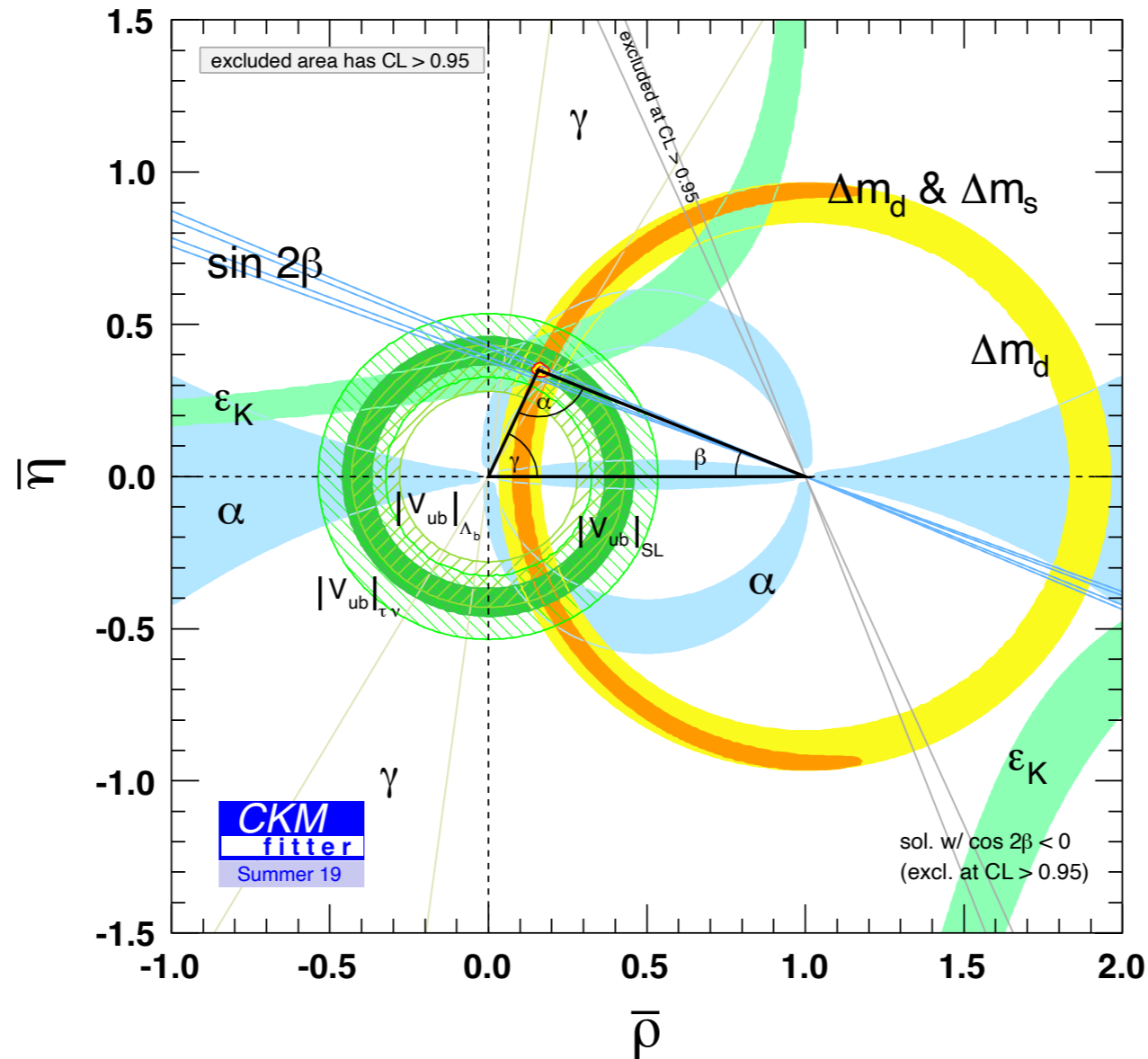
# CKM-ology

$$V_{CKM} = \begin{pmatrix} \text{large} & \text{medium} & \text{small} \\ \text{medium} & \text{large} & \text{small} \\ \text{small} & \text{small} & \text{large} \end{pmatrix} \begin{matrix} d \\ s \\ b \end{matrix} \begin{matrix} u \\ c \\ t \end{matrix}$$

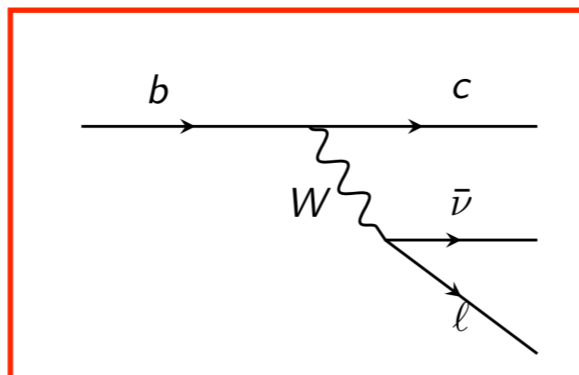
- ✗ Fix CKM entries through tree level processes & over-constrain by loop-induced ones [progress through precision!]



# CKM-ology

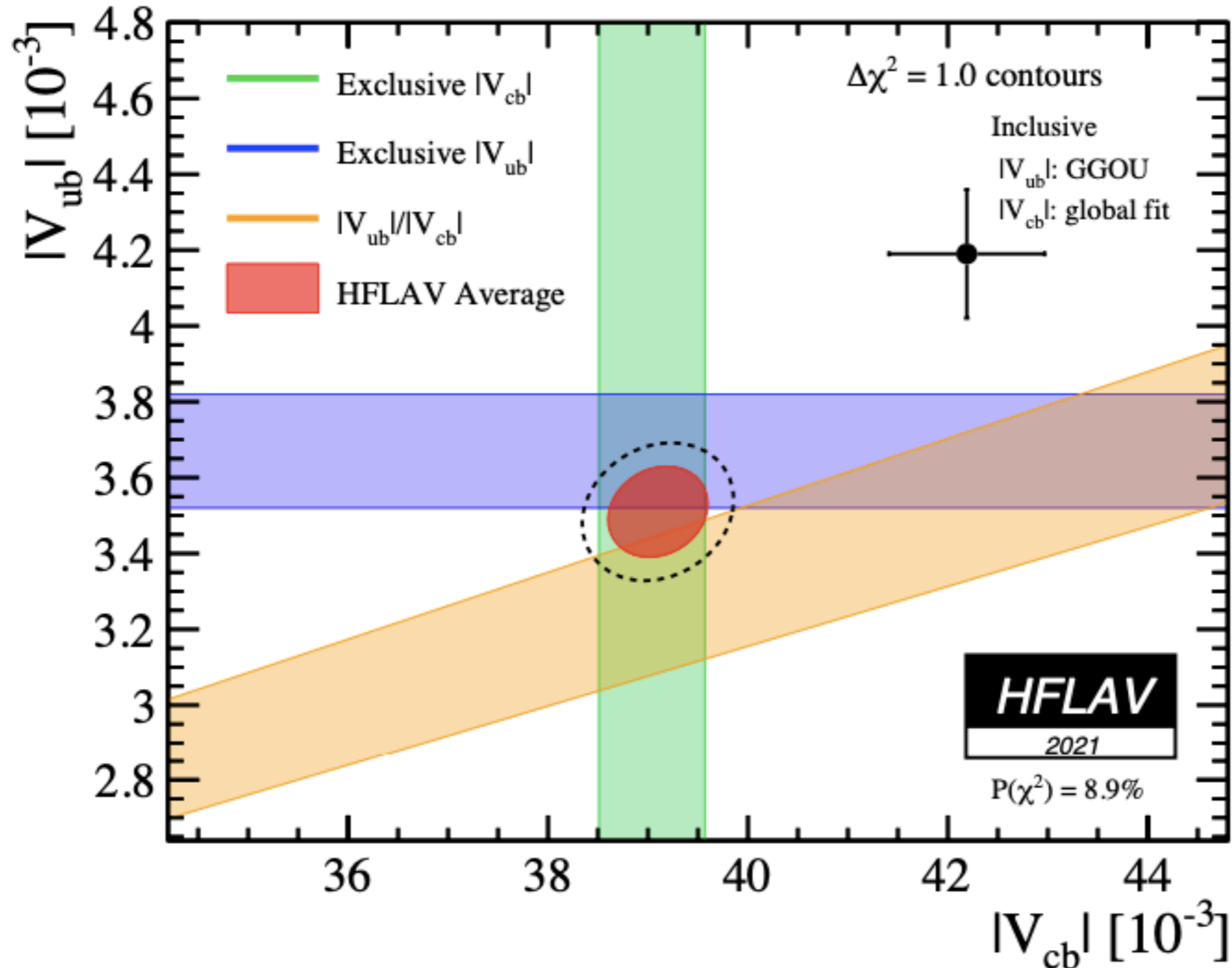


- x Still open: inclusive  $\nu$  exclusive  $V_{ub}$  and  $V_{cb}$ ?  
 Is  $V_{ud}$  well controlled?  $V_{us}$  keeps coming back (EM)...



# CKM-ology - Small flavor 'anomaly'

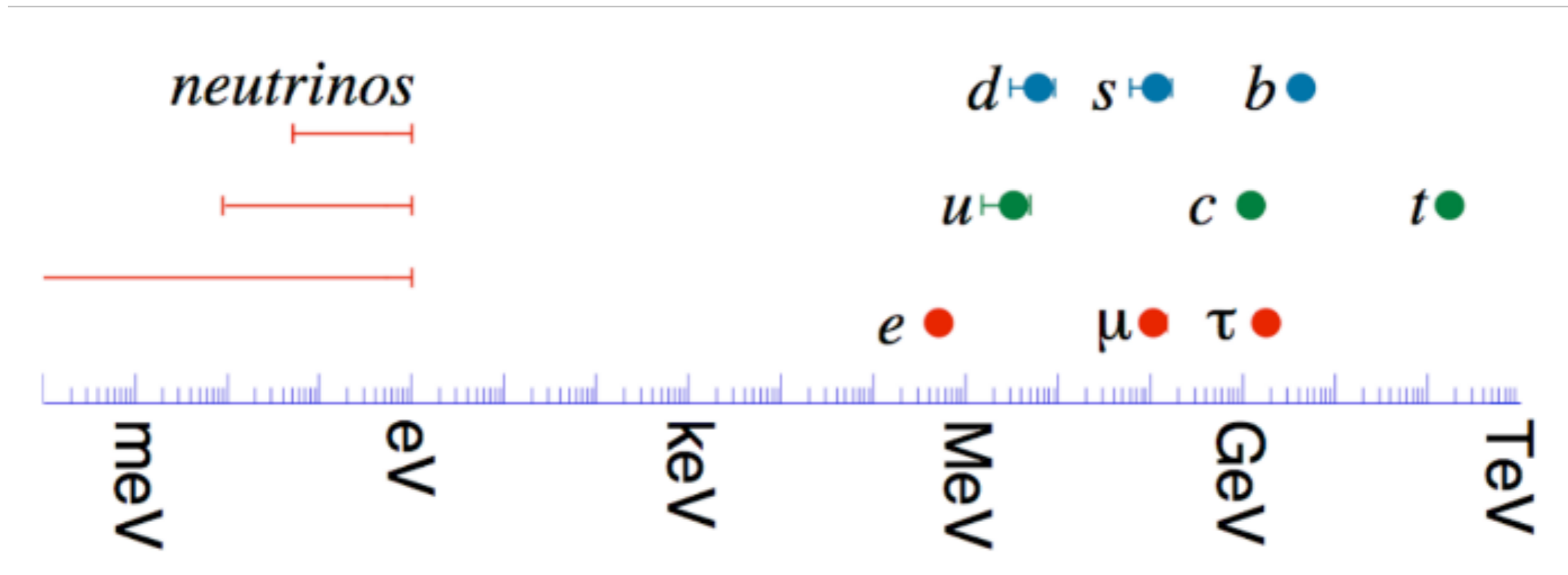
✗ Still open: inclusive v exclusive  $V_{ub}$  and  $V_{cb}$ ?



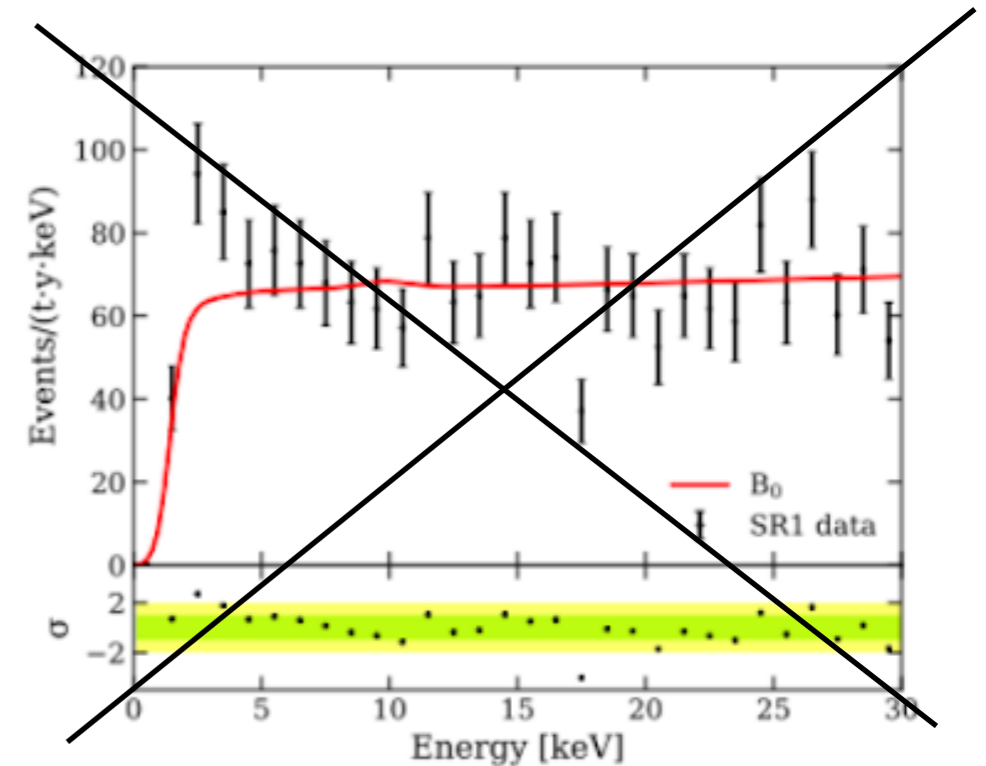
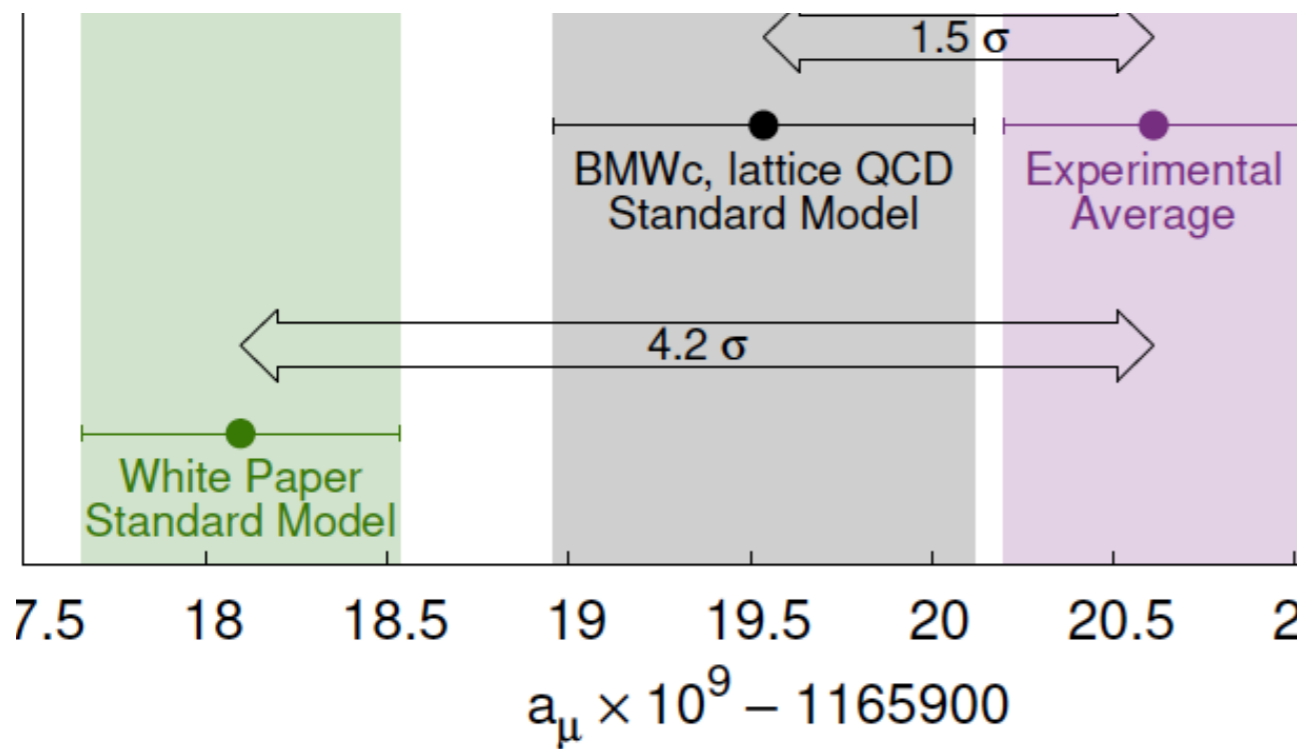
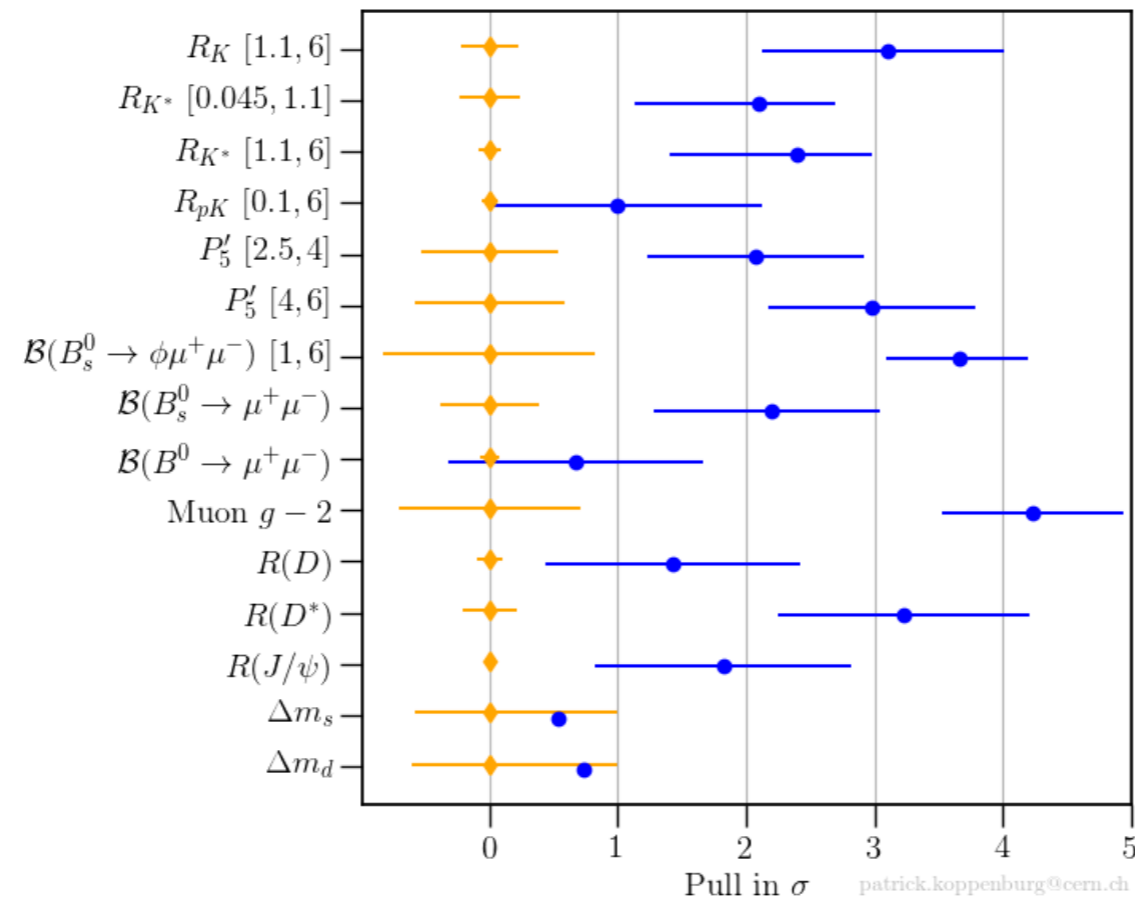


# Flavor Physics

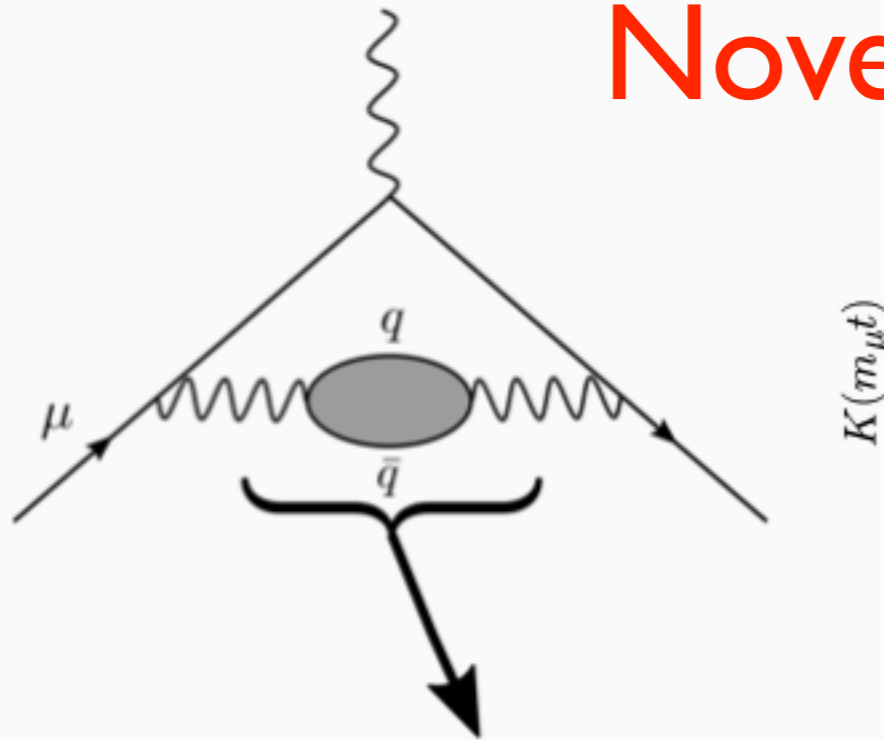
- ✗ Why three generations?
- ✗ Why such hierarchy of masses and mixing?
- ✗ Why so small CPV phase?



# Flavor Anomalies



# Novelties re $(g-2)_\mu$



$$\Pi_{\mu\nu}(Q) = \int d^4x e^{iQ \cdot x} \langle J_\mu(x) J_\nu(0) \rangle = (\delta_{\mu\nu} Q^2 - Q_\mu Q_\nu) \Pi(Q^2)$$

$$\Downarrow$$
$$a_\mu^{\text{LO-HVP}} = 4\alpha_{em}^2 \int_0^\infty dQ^2 \frac{1}{m_\mu^2} f\left(\frac{Q^2}{m_\mu^2}\right) \cdot (\Pi(Q^2) - \Pi(0)).$$

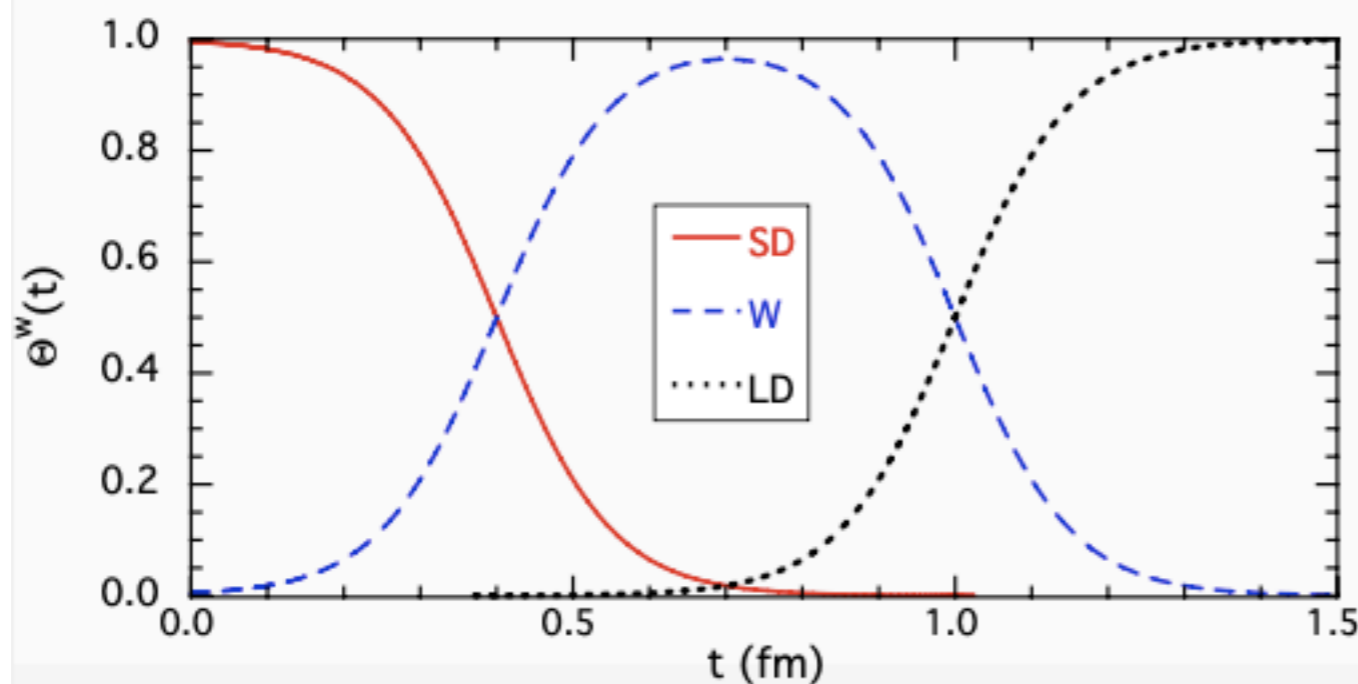
$$a_\mu^{\text{LO-HVP}} = 2\alpha_{em}^2 \int_0^\infty dt t^2 K(m_\mu t) V(t), \quad V(t) \equiv \frac{1}{3} \sum_{i=1,2,3} \int d\vec{x} \langle J_i(\vec{x}, t) J_i(0) \rangle$$

# Novelties re $(g-2)_\mu$

$$a_\mu^{SD} = 2\alpha_{em}^2 \int_0^\infty dt t^2 K(m_\mu t) V(t) \cdot \Theta^{SD}(t)$$

$$a_\mu^W = 2\alpha_{em}^2 \int_0^\infty dt t^2 K(m_\mu t) V(t) \cdot \Theta^W(t)$$

$$a_\mu^{LD} = 2\alpha_{em}^2 \int_0^\infty dt t^2 K(m_\mu t) V(t) \cdot \Theta^{LD}(t)$$



$$\Theta^{SD}(t) + \Theta^W(t) + \Theta^{LD}(t) = 1$$

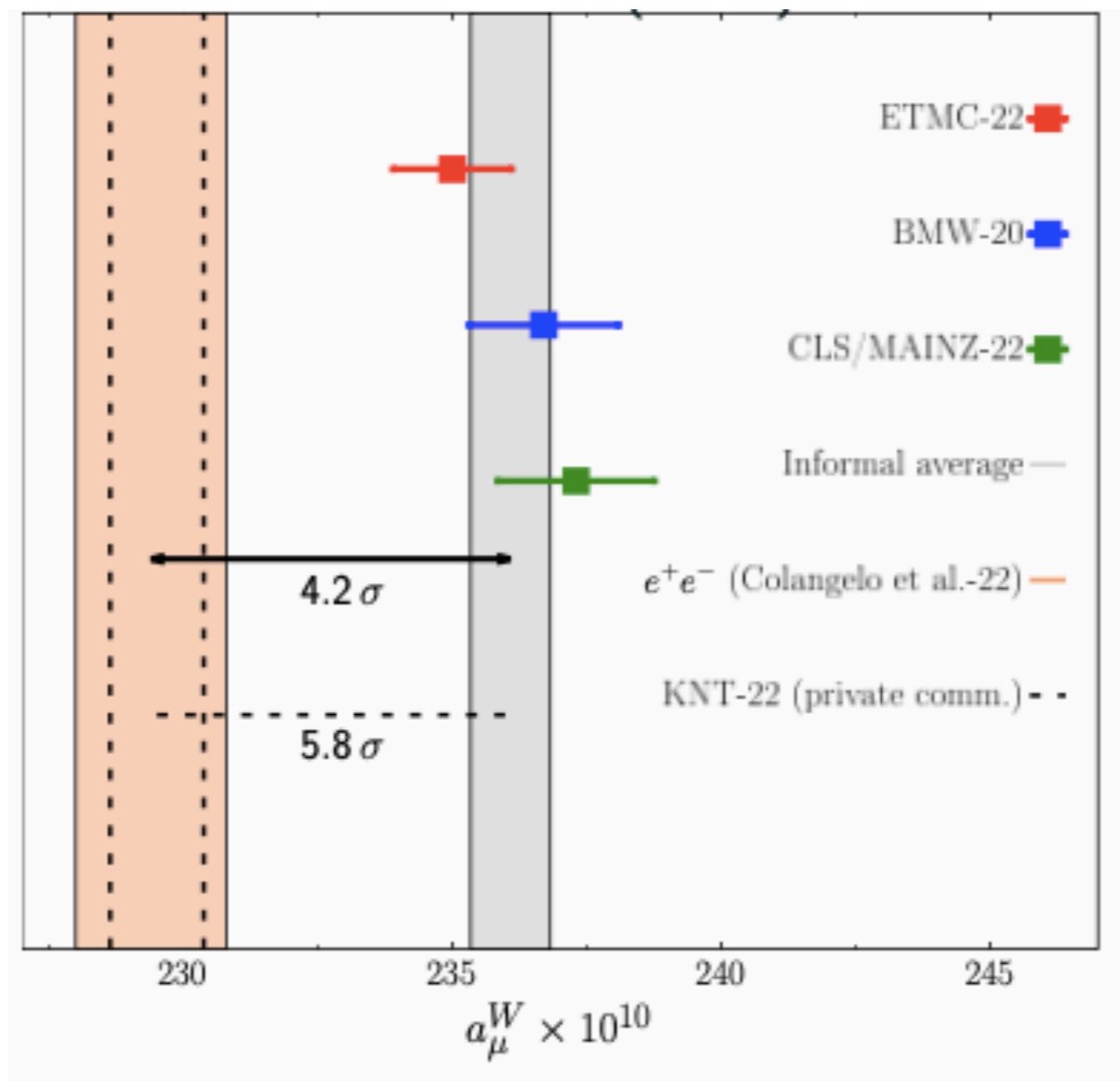


$$a_\mu^{\text{LO-HVP}} = a_\mu^{\text{SD}} + a_\mu^{\text{W}} + a_\mu^{\text{LD}}$$

$$\underbrace{2\alpha_{em}^2 \int_0^\infty dt t^2 K(m_\mu t) V(t)}_{\text{lattice, SM}} = a_\mu^{\text{LO-HVP}} = \underbrace{\frac{2\alpha_{em}^2 m_\mu^2}{9\pi^2} \int_{2M_\pi}^\infty \frac{dE}{E^2} \tilde{K}(E) R^{\text{had}}(E)}_{\text{dispersive, experimental}}$$



# Novelties re $(g-2)_\mu$



2205.12963

$$\underbrace{2\alpha_{em}^2 \int_0^\infty dt t^2 K(m_\mu t) V(t)}_{\text{lattice, SM}} = a_\mu^{\text{LO-HVP}} = \underbrace{\frac{2\alpha_{em}^2 m_\mu^2}{9\pi^2} \int_{2M_\pi}^\infty \frac{dE}{E^2} \tilde{K}(E) R^{\text{had}}(E)}_{\text{dispersive, experimental}}$$

# Lepton Flavor Universality Violation

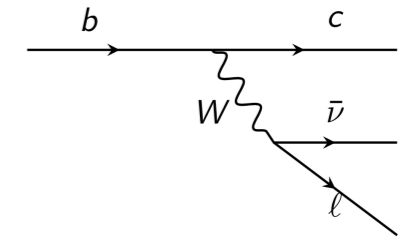
$$R_{D^{(*)}} = \frac{\mathcal{B}(B \rightarrow D^{(*)} \tau \bar{\nu})}{\mathcal{B}(B \rightarrow D^{(*)} \ell \bar{\nu})_{\ell \in (e, \mu)}} \quad \& \quad R_{D^{(*)}}^{\text{exp}} > R_{D^{(*)}}^{\text{SM}}$$

$$R_{K^{(*)}} = \frac{\mathcal{B}(B \rightarrow K^{(*)} \mu \mu)}{\mathcal{B}(B \rightarrow K^{(*)} e e)} \Bigg|_{q^2 \in [q_{\text{min}}^2, q_{\text{max}}^2]} \quad \& \quad R_{K^{(*)}}^{\text{exp}} < R_{K^{(*)}}^{\text{SM}}$$

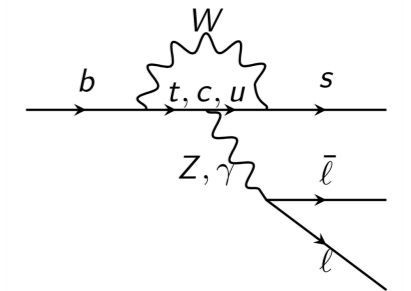
- CKM factor cancels
- Bulk of hadronic uncertainties cancel

# Lepton Flavor Universality Violation

$$R_{D^{(*)}} = \frac{\mathcal{B}(B \rightarrow D^{(*)} \tau \bar{\nu})}{\mathcal{B}(B \rightarrow D^{(*)} \ell \bar{\nu})_{\ell \in (e, \mu)}} \quad \& \quad R_{D^{(*)}}^{\text{exp}} > R_{D^{(*)}}^{\text{SM}}$$



$$R_{K^{(*)}} = \frac{\mathcal{B}(B \rightarrow K^{(*)} \mu \mu)}{\mathcal{B}(B \rightarrow K^{(*)} e e)} \Bigg|_{q^2 \in [q_{\text{min}}^2, q_{\text{max}}^2]} \quad \& \quad R_{K^{(*)}}^{\text{exp}} < R_{K^{(*)}}^{\text{SM}}$$



$$R_{D^{(*)}}^{\text{exp}} > R_{D^{(*)}}^{\text{SM}} \quad \Rightarrow \quad \Lambda_{\text{NP}} \lesssim 3 \text{ TeV}$$

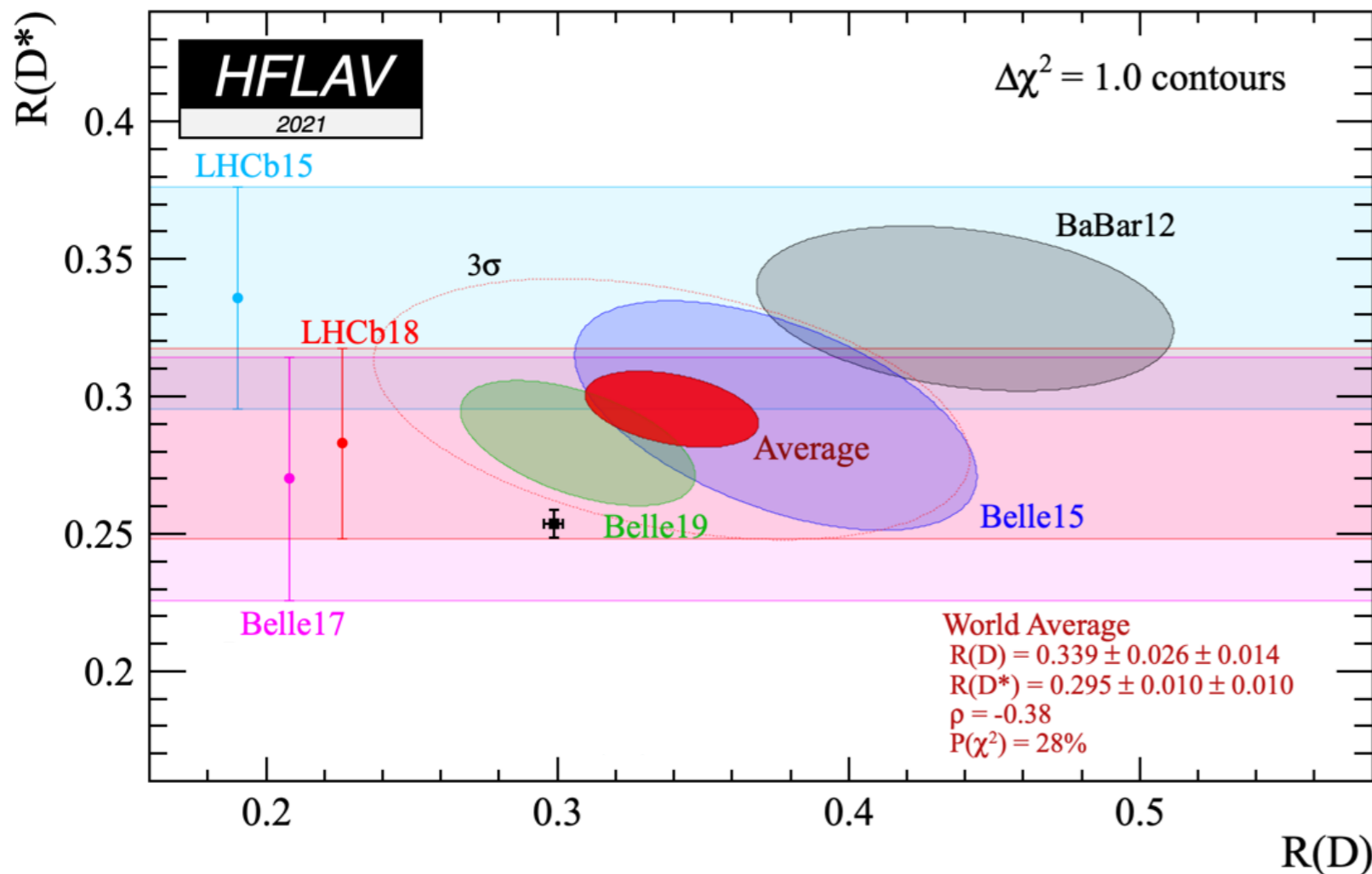
$$R_{K^{(*)}}^{\text{exp}} < R_{K^{(*)}}^{\text{SM}} \quad \Rightarrow \quad \Lambda_{\text{NP}} \lesssim 30 \text{ TeV}$$

# LFUV

## need studying NP effects

$$R_{D^{(*)}} = \frac{\mathcal{B}(B \rightarrow D^{(*)} \tau \bar{\nu})}{\mathcal{B}(B \rightarrow D^{(*)} \ell \bar{\nu})_{\ell \in (e, \mu)}} \quad \& \quad R_{D^{(*)}}^{\text{exp}} > R_{D^{(*)}}^{\text{SM}}$$

$$R_{D^{(*)}}^{\text{exp}} > R_{D^{(*)}}^{\text{SM}} \Rightarrow \Lambda_{\text{NP}} \lesssim 3 \text{ TeV}$$



### LHCb

- $B_c \rightarrow J/\psi \ell \bar{\nu}$

$$R_{J/\psi}^{\text{exp}} > R_{J/\psi}^{\text{SM}}$$

- $\Lambda_b \rightarrow \Lambda_c \ell \bar{\nu}$

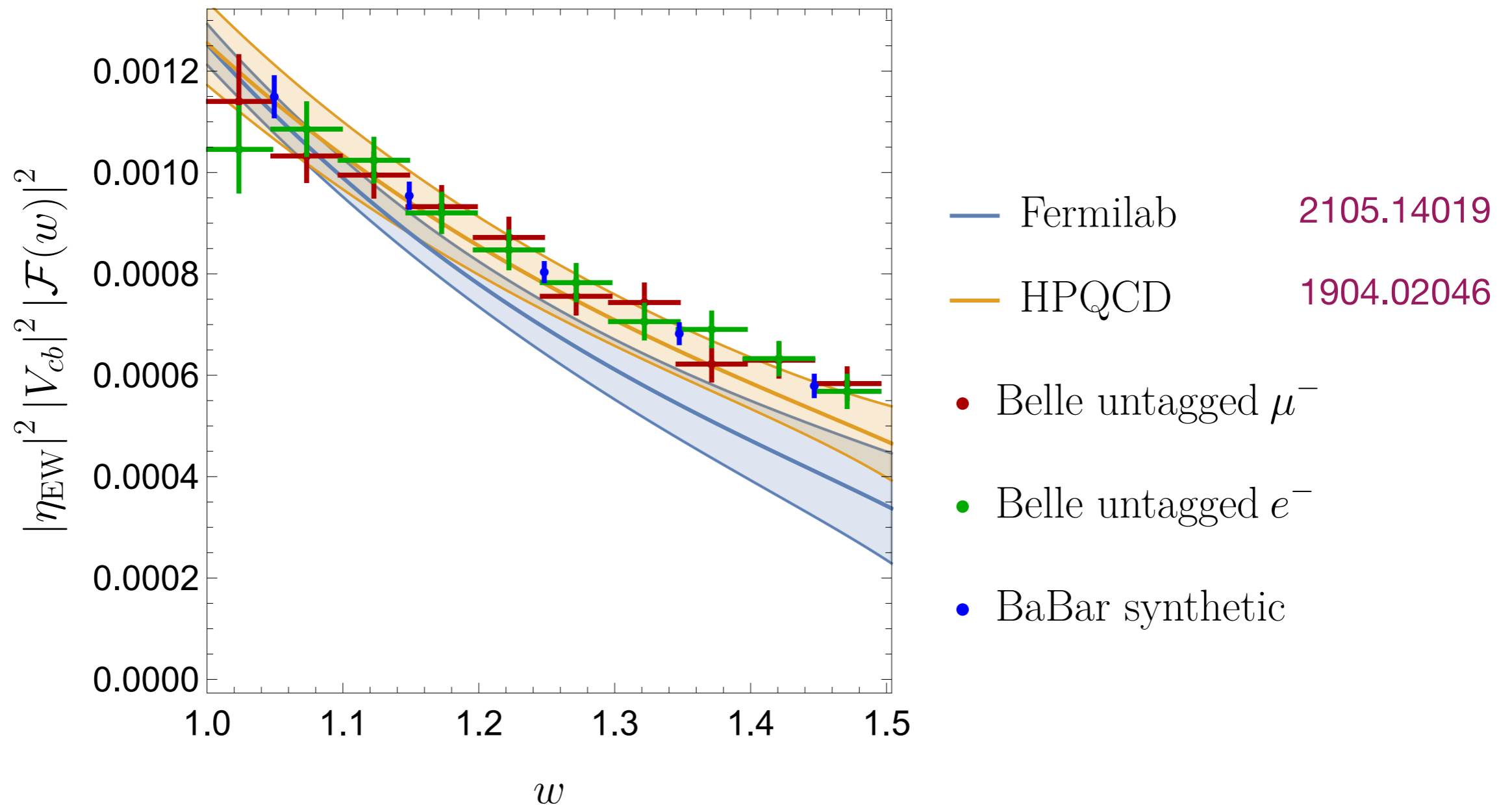
$$R_{\Lambda_c}^{\text{exp}} \approx R_{\Lambda_c}^{\text{SM}}$$

NEW and can be improved... a lot



# We still do not have a control over hadronic uncertainties with

$$\frac{d\mathcal{B}(B \rightarrow D^* \ell \bar{\nu})}{dq^2} = \frac{1}{2m_B m_{D^*}} \frac{d\mathcal{B}(B \rightarrow D^* \ell \bar{\nu})}{dw} \propto |V_{cb}|^2 |\mathcal{F}(w)|^2$$



• Assuming with HPQCD that  $\mathcal{F}(w)^{B_s \rightarrow D_s^*} = \mathcal{F}(w)^{B \rightarrow D^*}$

# We still do not have a full/good control over hadronic uncertainties

Mode	$B \rightarrow D\ell\bar{\nu}$	$B \rightarrow D^*\ell\bar{\nu}$	$\Lambda_b \rightarrow \Lambda_c\ell\bar{\nu}$
$\langle V_\mu \rangle$	<b>2</b> ✓	<b>1</b> ✓	<b>3</b> ✓
$\langle A_\mu \rangle$		<b>3</b> ✓	<b>3</b> ✓
$\langle T_{\mu\nu} \rangle$	<b>1</b> ✗	<b>3</b> ✗	<b>4</b> ✓

1503.07237

1904.02046

1503.01421

1505.03925

2105.14019

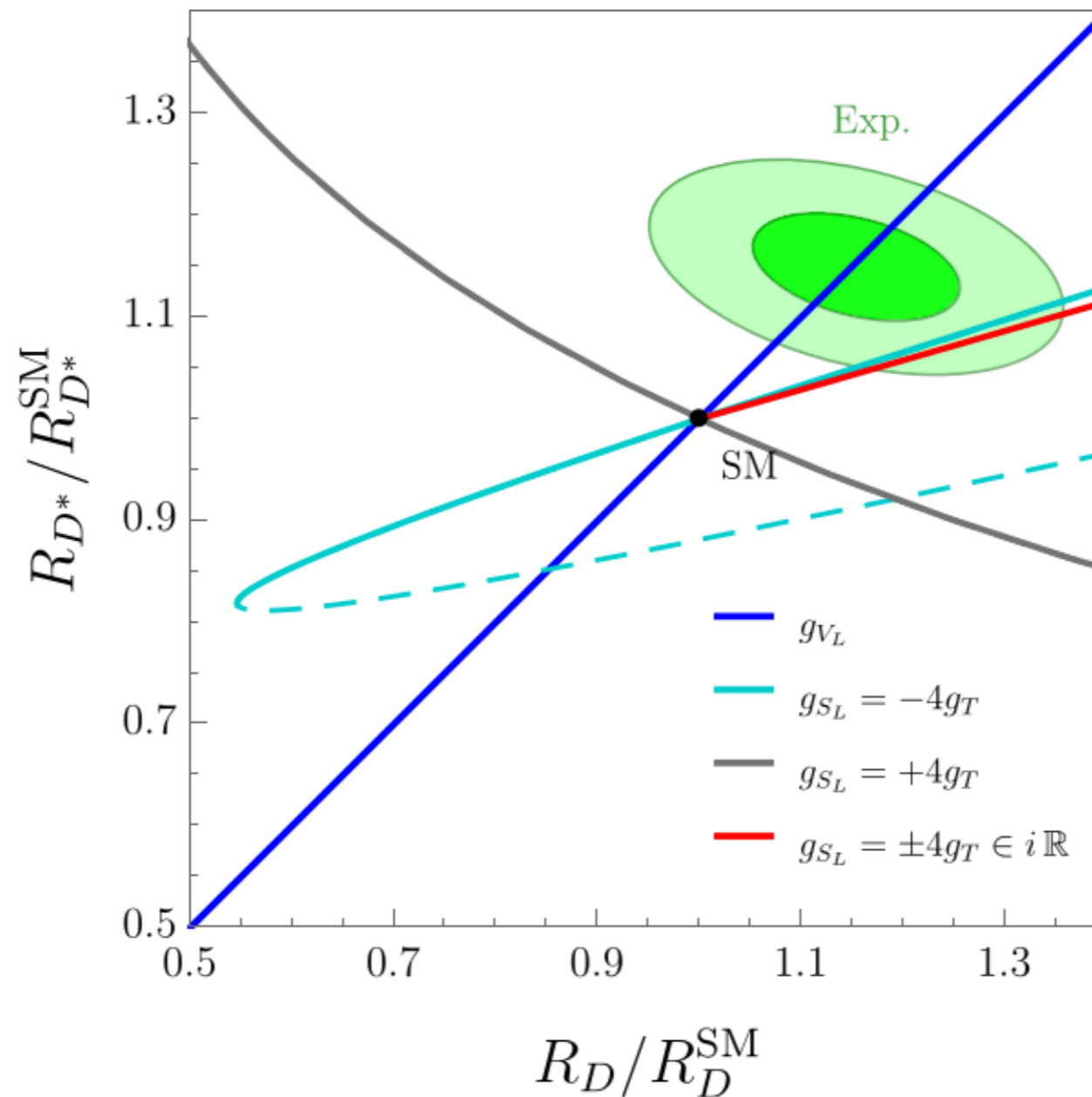
1702.02243

# EFT - exclusive $b \rightarrow c l \nu$

$$\mathcal{L}_{\text{eff}} = -2\sqrt{2}G_F V_{cb} \left[ (1 + g_{V_L})(\bar{c}_L \gamma_\mu b_L)(\bar{\ell}_L \gamma^\mu \nu_L) + g_{V_R}(\bar{c}_R \gamma_\mu b_R)(\bar{\ell}_L \gamma^\mu \nu_L) \right. \\ \left. + g_{S_R}(\bar{c}_L b_R)(\bar{\ell}_R \nu_L) + g_{S_L}(\bar{c}_R b_L)(\bar{\ell}_R \nu_L) + g_T(\bar{c}_R \sigma_{\mu\nu} b_L)(\bar{\ell}_R \sigma^{\mu\nu} \nu_L) \right] + \text{h.c.}$$

# EFT - exclusive $b \rightarrow cl\nu$

$$\mathcal{L}_{\text{eff}} = -2\sqrt{2}G_F V_{cb} \left[ (1 + g_{V_L})(\bar{c}_L \gamma_\mu b_L)(\bar{\ell}_L \gamma^\mu \nu_L) + g_{V_R}(\bar{c}_R \gamma_\mu b_R)(\bar{\ell}_L \gamma^\mu \nu_L) \right. \\ \left. + g_{S_R}(\bar{c}_L b_R)(\bar{\ell}_R \nu_L) + g_{S_L}(\bar{c}_R b_L)(\bar{\ell}_R \nu_L) + g_T(\bar{c}_R \sigma_{\mu\nu} b_L)(\bar{\ell}_R \sigma^{\mu\nu} \nu_L) \right] + \text{h.c.}$$





# EFT - exclusive $b \rightarrow cl\nu$

$g_{V_L}(m_b)$	$0.07 \pm 0.02$	0.02/1	✓
$g_{S_R}(m_b)$	$-0.31 \pm 0.05$	5.3/1	✗
$g_{S_L}(m_b)$	$0.12 \pm 0.06$	8.8/1	✗
$g_T(m_b)$	$-0.03 \pm 0.01$	3.1/1	✓
$g_{S_L} = +4g_T \in \mathbb{R}$	$-0.03 \pm 0.07$	12.5/1	✗
$g_{S_L} = -4g_T \in \mathbb{R}$	$0.16 \pm 0.05$	2.0/1	✓
$g_{S_L} = \pm 4g_T \in i\mathbb{R}$	$0.48 \pm 0.08$	2.4/1	✓

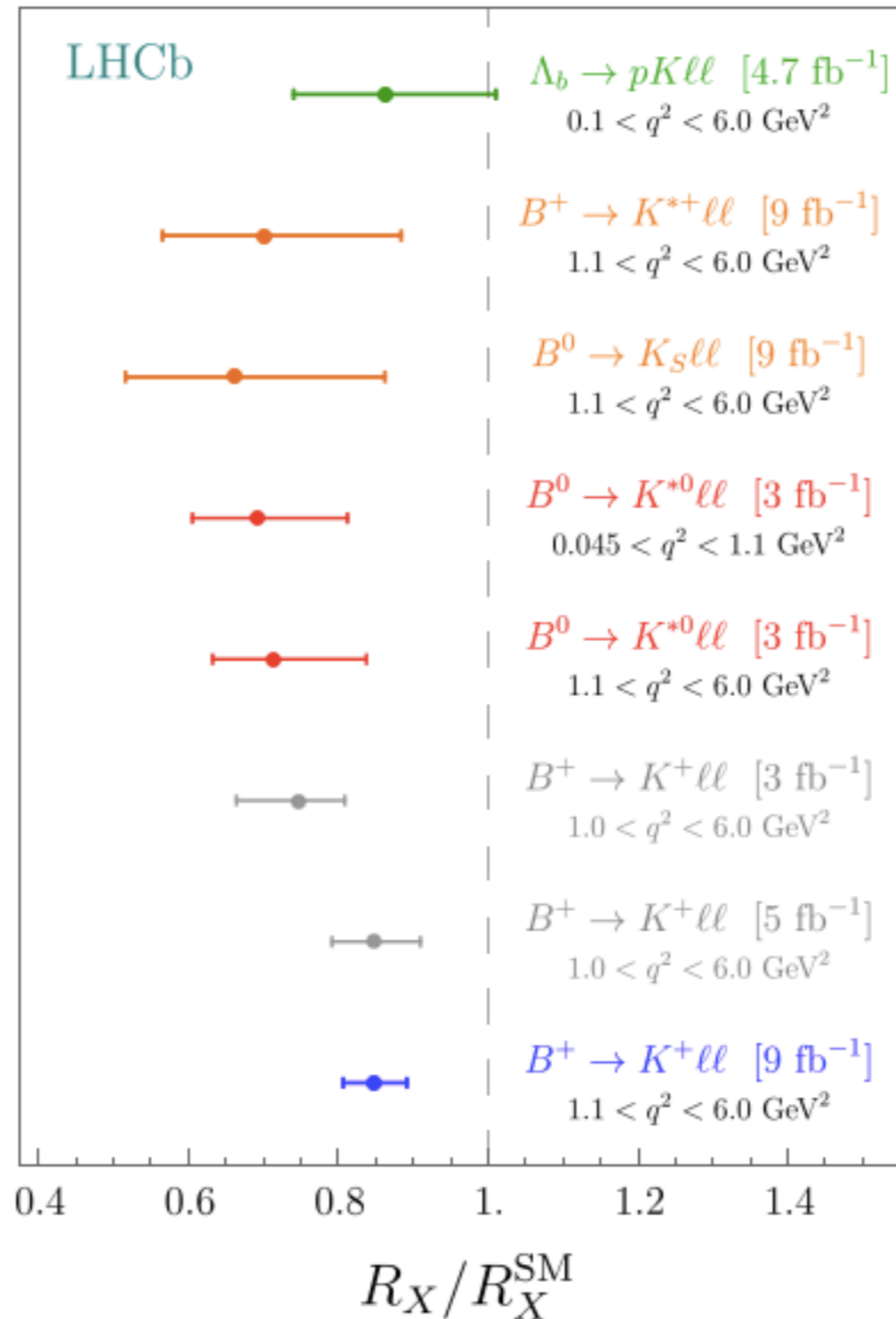
# EFT - exclusive $b \rightarrow c l \nu$

$g_{V_L}(m_b)$			$0.07 \pm 0.02$	$0.02/1$	✓
Model	$g_{\text{eff}}^{b \rightarrow c \tau \bar{\nu}}(\mu = m_\Delta)$	$R_{D^{(*)}}$	$\pm 0.05$	$5.3/1$	✗
$S_1 = (\bar{3}, 1, 1/3)$	$g_{V_L}, g_{S_L} = -4 g_T$	✓	$\pm 0.06$	$8.8/1$	✗
$R_2 = (3, 2, 7/6)$	$g_{S_L} = 4 g_T$	✓			
$S_3 = (\bar{3}, 3, 1/3)$	$g_{V_L}$	✗	$\pm 0.01$	$3.1/1$	✓
...	...	...			
$U_1 = (3, 1, 2/3)$	$g_{V_L}, g_{S_R}$	✓			
$U_3 = (3, 3, 2/3)$	$g_{V_L}$	✗	$\pm 0.07$	$12.5/1$	✗
...	...	...			
$g_{S_L} = -4g_T \in \mathbb{R}$			$0.16 \pm 0.05$	$2.0/1$	✓
$g_{S_L} = \pm 4g_T \in i\mathbb{R}$			$0.48 \pm 0.08$	$2.4/1$	✓

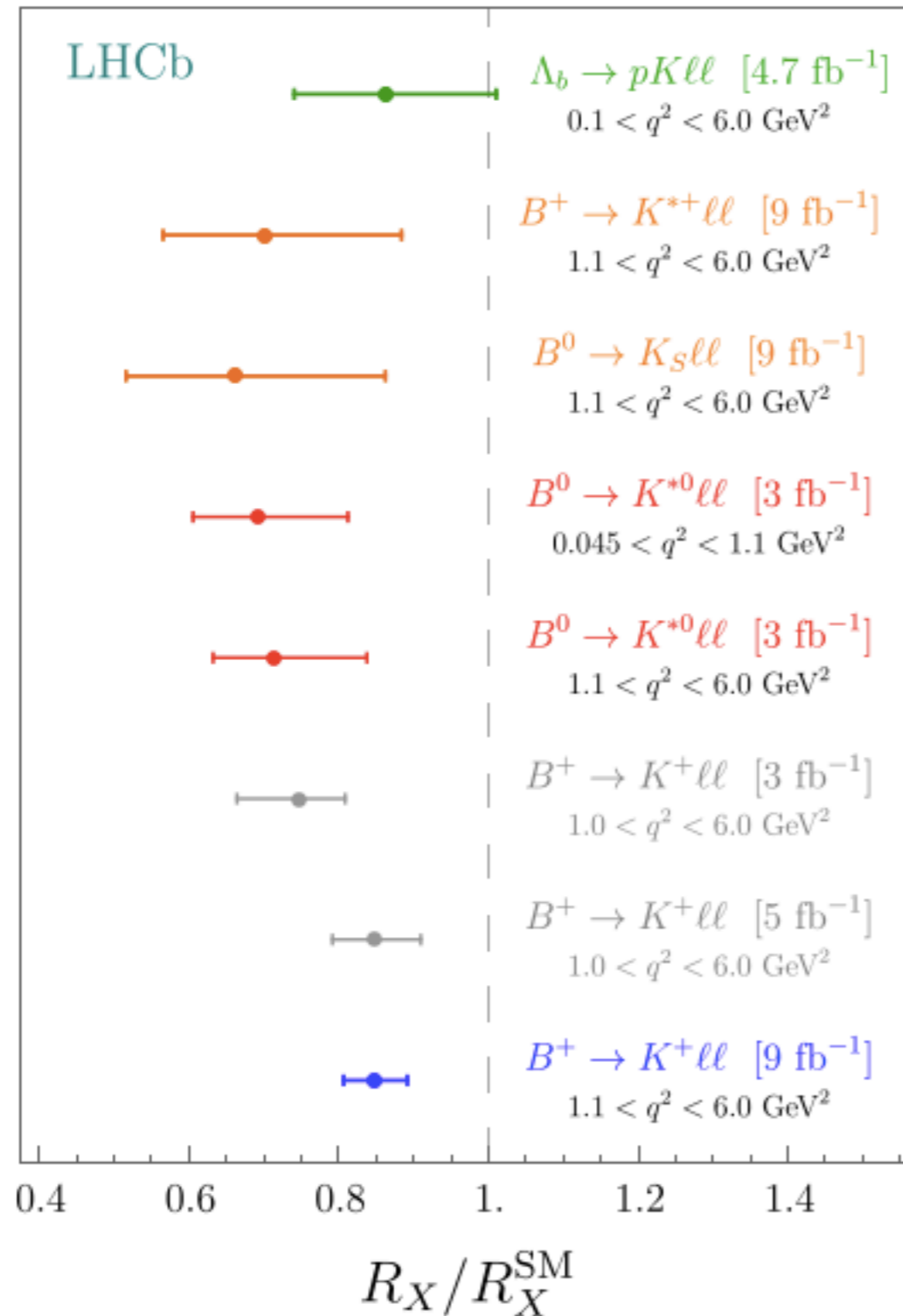
Main worry remain the hadronic uncertainties in the  $D^*$  case:  
 No clear LQCD info regarding the shapes of FFs  
 Keep also in mind the SD part of the soft photon problem is missing

# MORE LFUV

## need studying NP effects



# MORE LFUV need studying NP effects



$$R_{K^*}^{[1.1,6]} = 0.71(10)^{\text{LHCb}} \quad \text{vs} \quad R_{K^*}^{[1,6]} = 1.00(1)^{\text{SM}}$$

$$R_K^{[1.1,6]} = 0.847(42)^{\text{LHCb}} \quad \text{vs} \quad R_K^{[1,6]} = 1.00(1)^{\text{SM}}$$



# EFT - exclusive $b \rightarrow s \ell \ell$

$$\mathcal{H}_{\text{eff}} = -\frac{4G_F}{\sqrt{2}} V_{tb} V_{ts}^* \left[ \sum_{i=1}^6 C_i(\mu) \mathcal{O}_i(\mu) + \sum_{i=7,8,9,10,P,S,\dots} \left( C_i(\mu) \mathcal{O}_i + C'_i(\mu) \mathcal{O}'_i \right) \right] + \text{h.c.}$$

$$\mathcal{O}_9^{(\prime)} = (\bar{s} \gamma_\mu P_{L(R)} b) (\bar{\ell} \gamma^\mu \ell)$$

$$\mathcal{O}_{10}^{(\prime)} = (\bar{s} \gamma_\mu P_{L(R)} b) (\bar{\ell} \gamma^\mu \gamma^5 \ell)$$

$$\mathcal{O}_S^{(\prime)} = (\bar{s} P_{R(L)} b) (\bar{\ell} \ell)$$

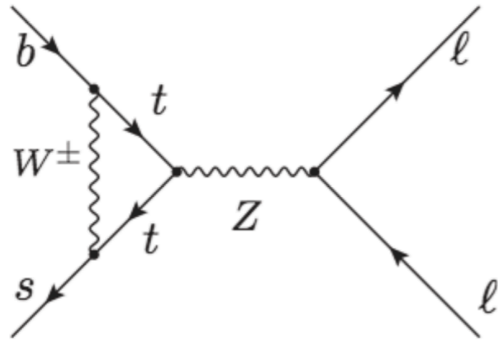
$$\mathcal{O}_P^{(\prime)} = (\bar{s} P_{R(L)} b) (\bar{\ell} \gamma_5 \ell)$$

$$\mathcal{O}_7^{(\prime)} = m_b (\bar{s} \sigma_{\mu\nu} P_{R(L)} b) F^{\mu\nu}$$

$$\text{Exp} : \mathcal{B}(B_s \rightarrow \mu\mu) = (2.85 \pm 0.33) \times 10^{-9}$$

$$\text{SM} : \mathcal{B}(B_s \rightarrow \mu\mu) = (3.66 \pm 0.14) \times 10^{-9}$$

$$\langle 0 | \bar{s} \gamma^\mu \gamma_5 b | B_s(p) \rangle = i f_{B_s} p^\mu$$



Exp :  $\mathcal{B}(B_s \rightarrow \mu\mu) = (2.85 \pm 0.33) \times 10^{-9}$   
12%

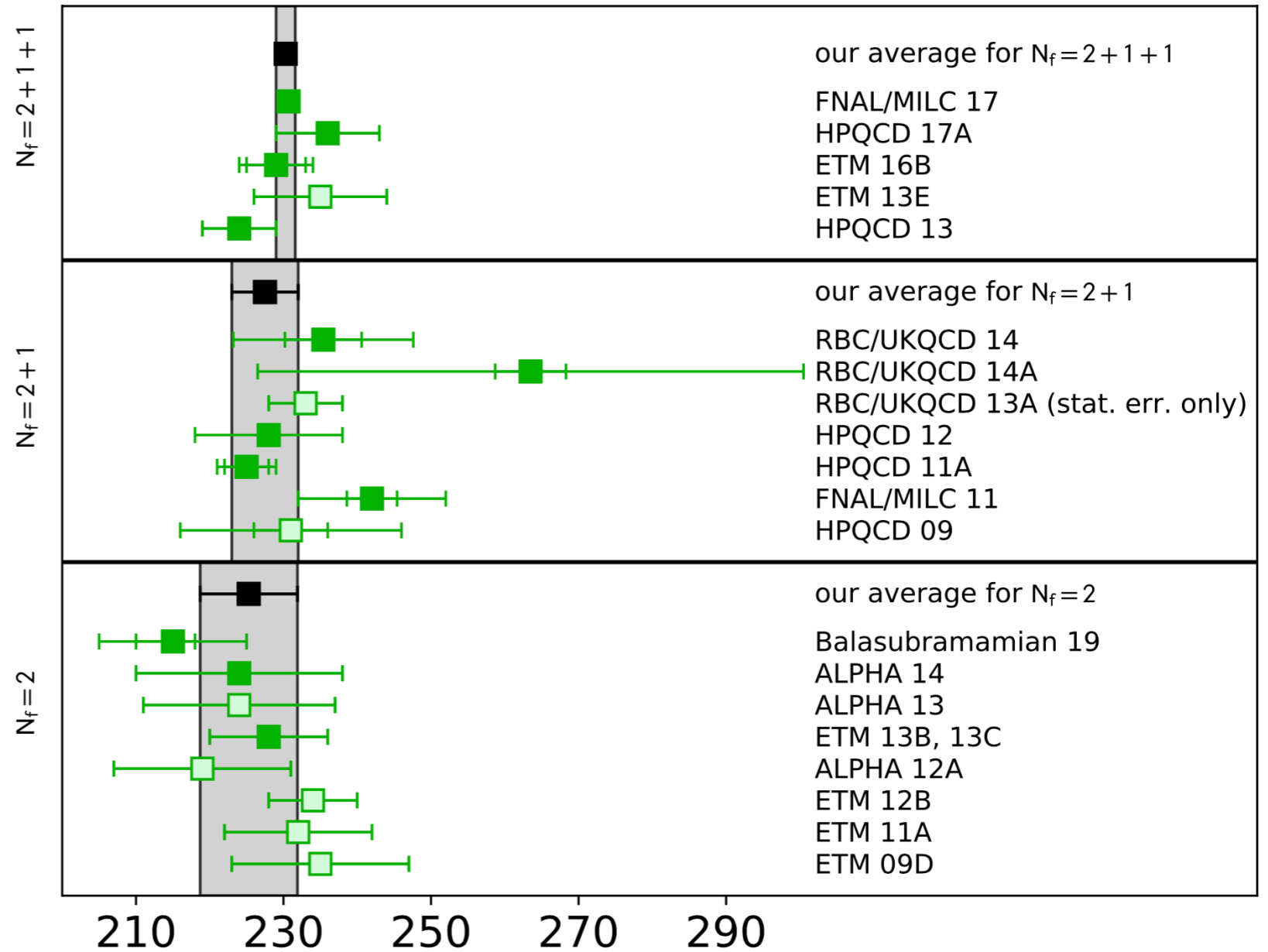
SM :  $\mathcal{B}(B_s \rightarrow \mu\mu) = (3.66 \pm 0.14) \times 10^{-9}$   
4%

$$O = \frac{1}{\Lambda^2} C_{ij} \bar{Q}_i \gamma^\mu Q_j H^\dagger D_\mu H$$

$C_{ij}$	1	$V_{ti} V_{tj}^*$
$B_s \rightarrow \mu^+ \mu^-$	$> 10 \text{ TeV}$	$> 2.5 \text{ TeV}$
$K \rightarrow \pi \nu \bar{\nu}$	$> 100 \text{ TeV}$	$> 1.8 \text{ TeV}$

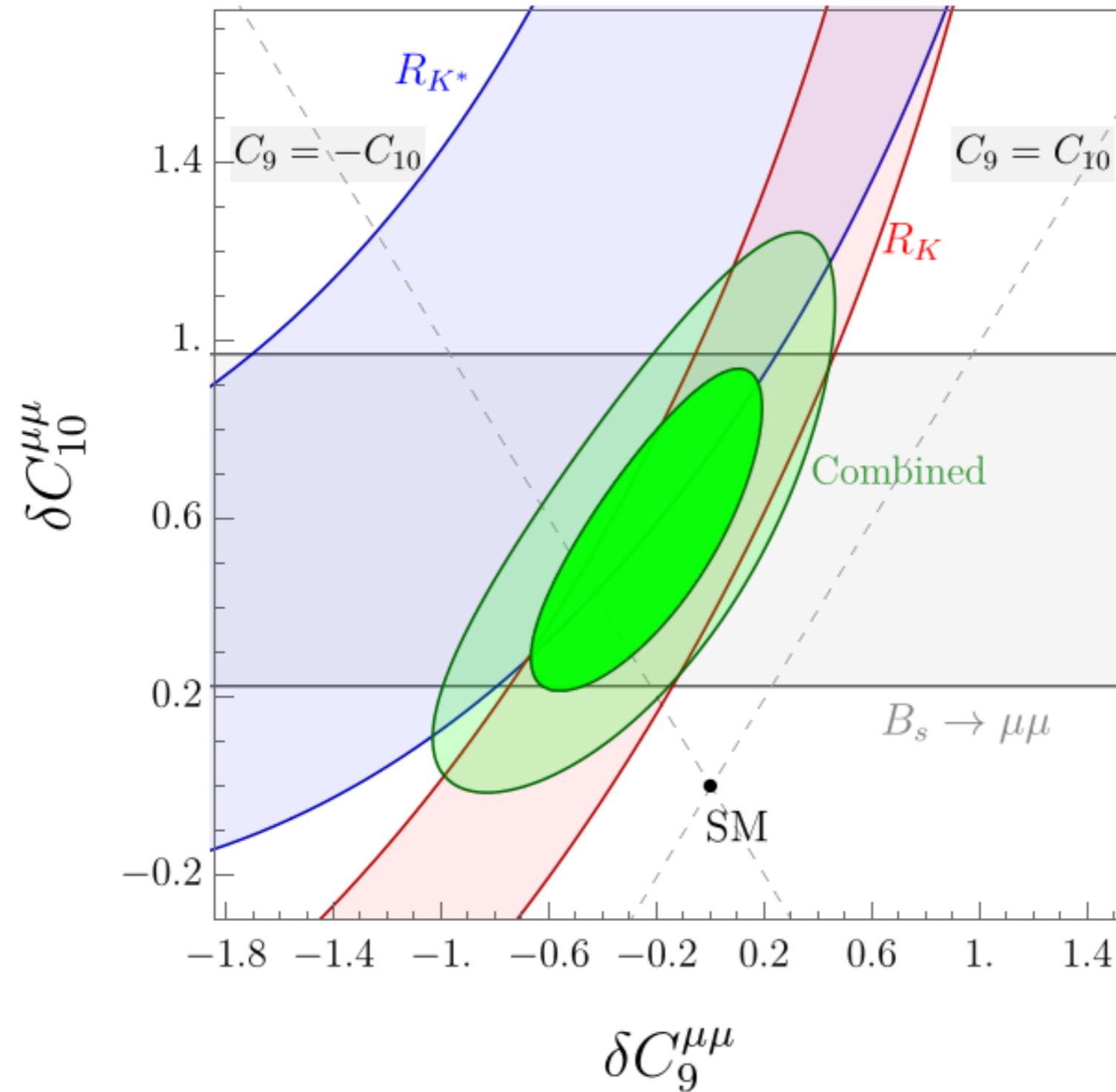
FLAG2021

$f_{B_s}$  [MeV]



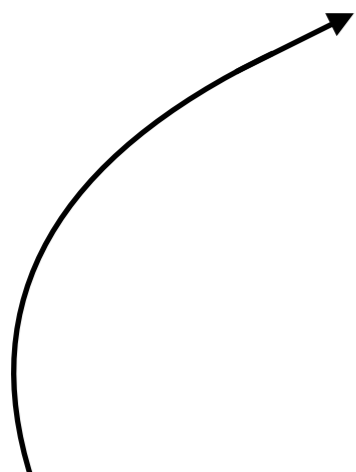
Fit to clean quantities:  $\mathcal{B}(B_s \rightarrow \mu\mu)$  and  $R_{K^{(*)}}$

EFT for  $b \rightarrow sll$



# What LQ scenario?

Model	$R_{D^{(*)}}$	$R_{K^{(*)}}$	$R_{D^{(*)}} \& R_{K^{(*)}}$
$S_1 = (\bar{3}, 1, 1/3)$	✓	✗	✗
$R_2 = (3, 2, 7/6)$	✓	✓*	✗
$S_3 = (\bar{3}, 3, 1/3)$	✗	✓	✗
$U_1 = (3, 1, 2/3)$	✓	✓	✓
$U_3 = (3, 3, 2/3)$	✗	✓	✗

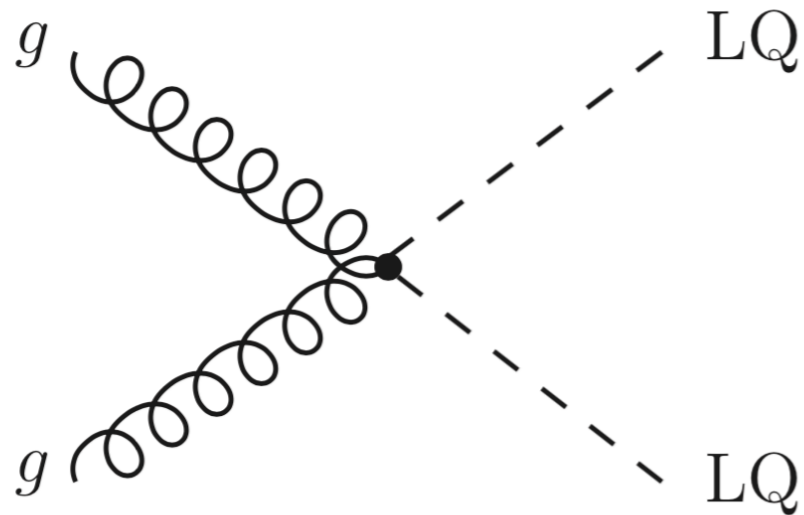


*N.B.  $U_1$  is the only one to accommodate both!*

Observable
$b \rightarrow s\mu\mu$
$b \rightarrow c\tau\nu$
$\mathcal{B}(\tau \rightarrow \mu\phi)$
$\mathcal{B}(B \rightarrow \tau\nu)$
$\mathcal{B}(D_s \rightarrow \mu\nu)$
$\mathcal{B}(D_s \rightarrow \tau\nu)$
$r_K^{c/\mu}$
$r_K^{\tau/\mu}$
$R_D^{\mu/e}$

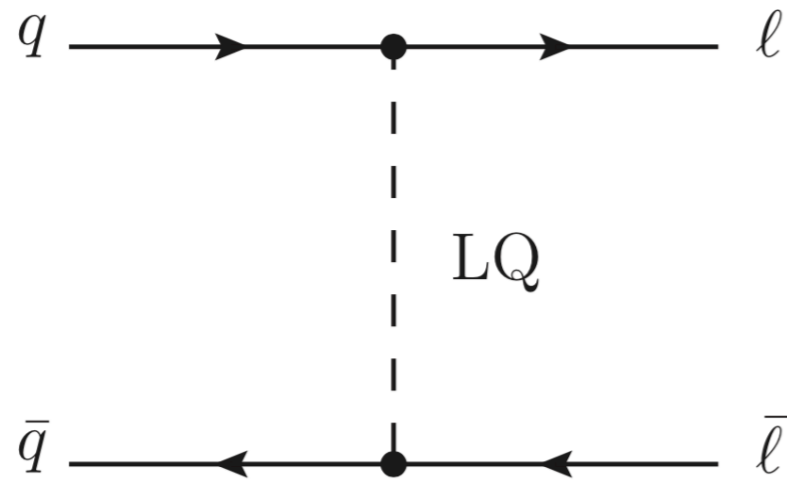
# From direct searches

## Atlas and CMS 2018-2021

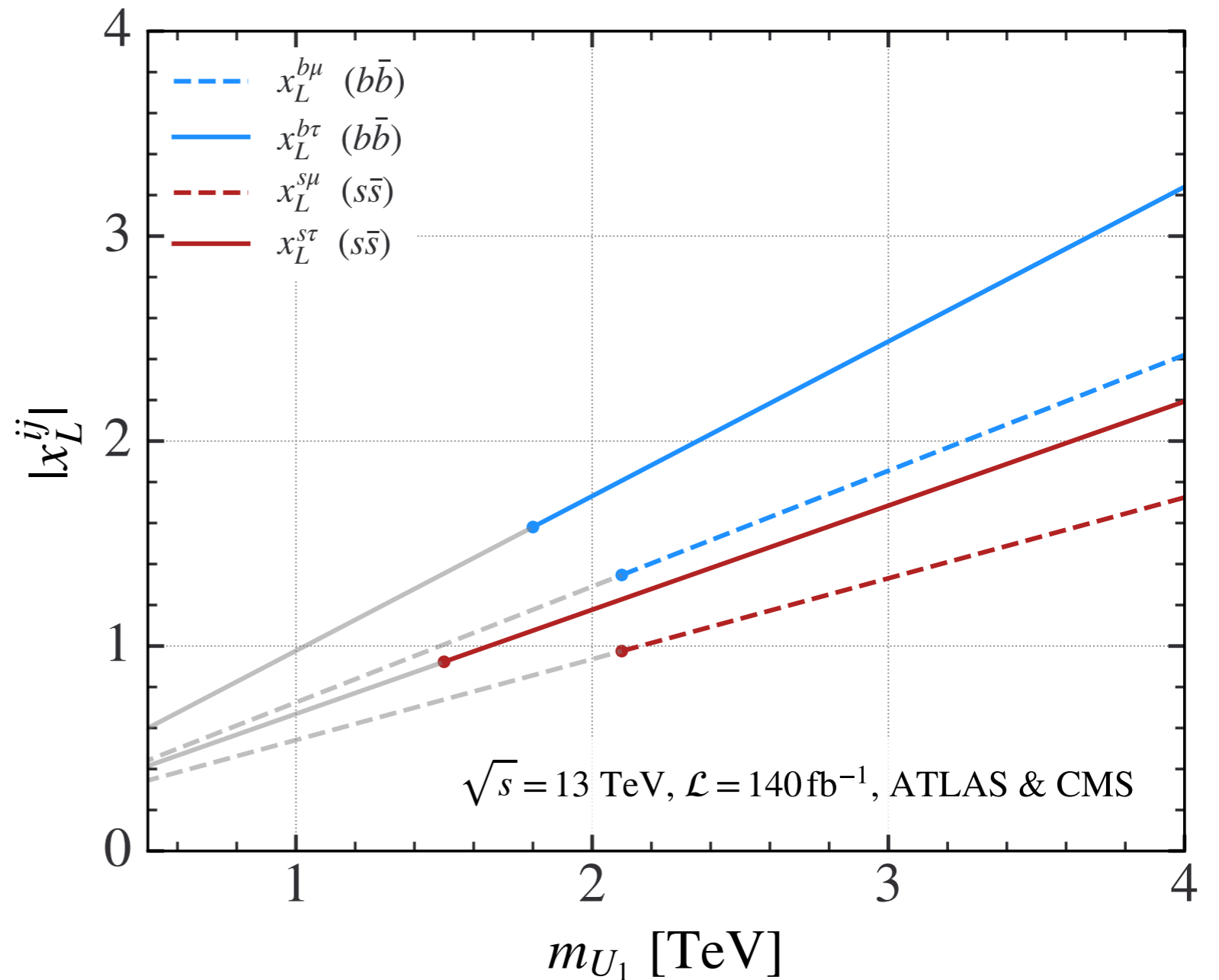


Decays	Scalar LQ limits	Vector LQ limits	$\mathcal{L}_{\text{int}}$
$jj \tau \bar{\tau}$	–	–	–
$b\bar{b} \tau \bar{\tau}$	1.0 (0.8) TeV	1.5 (1.3) TeV	36 fb <sup>-1</sup>
$t\bar{t} \tau \bar{\tau}$	1.4 (1.2) TeV	2.0 (1.8) TeV	140 fb <sup>-1</sup>
$jj \mu \bar{\mu}$	1.7 (1.4) TeV	2.3 (2.1) TeV	140 fb <sup>-1</sup>
$b\bar{b} \mu \bar{\mu}$	1.7 (1.5) TeV	2.3 (2.1) TeV	140 fb <sup>-1</sup>
$t\bar{t} \mu \bar{\mu}$	1.5 (1.3) TeV	2.0 (1.8) TeV	140 fb <sup>-1</sup>
$jj \nu \bar{\nu}$	1.0 (0.6) TeV	1.8 (1.5) TeV	36 fb <sup>-1</sup>
$b\bar{b} \nu \bar{\nu}$	1.1 (0.8) TeV	1.8 (1.5) TeV	36 fb <sup>-1</sup>
$t\bar{t} \nu \bar{\nu}$	1.2 (0.9) TeV	1.8 (1.6) TeV	140 fb <sup>-1</sup>

# From dilepton spectra at high $p_T$ Atlas and CMS 2018-2021

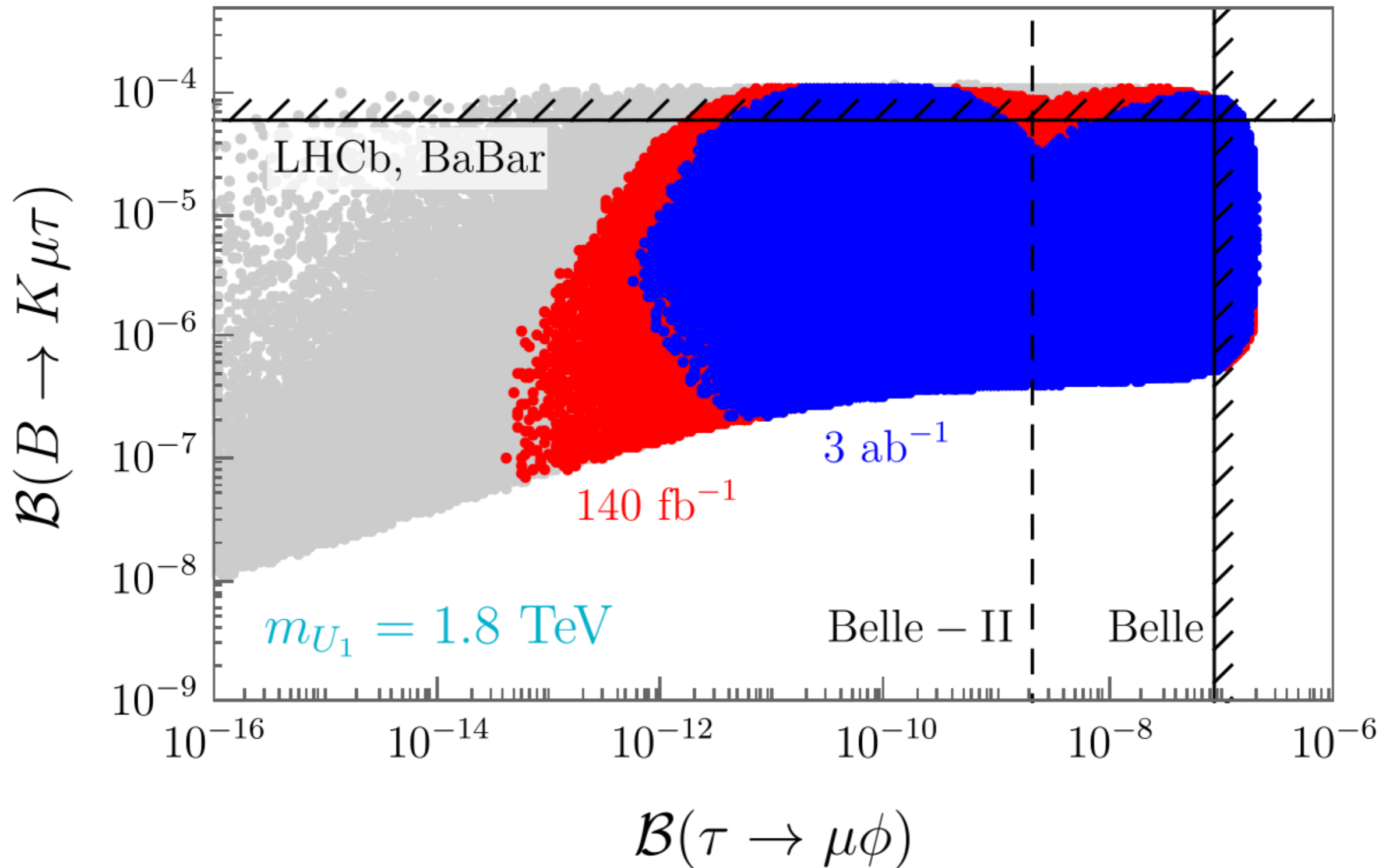


Example U1



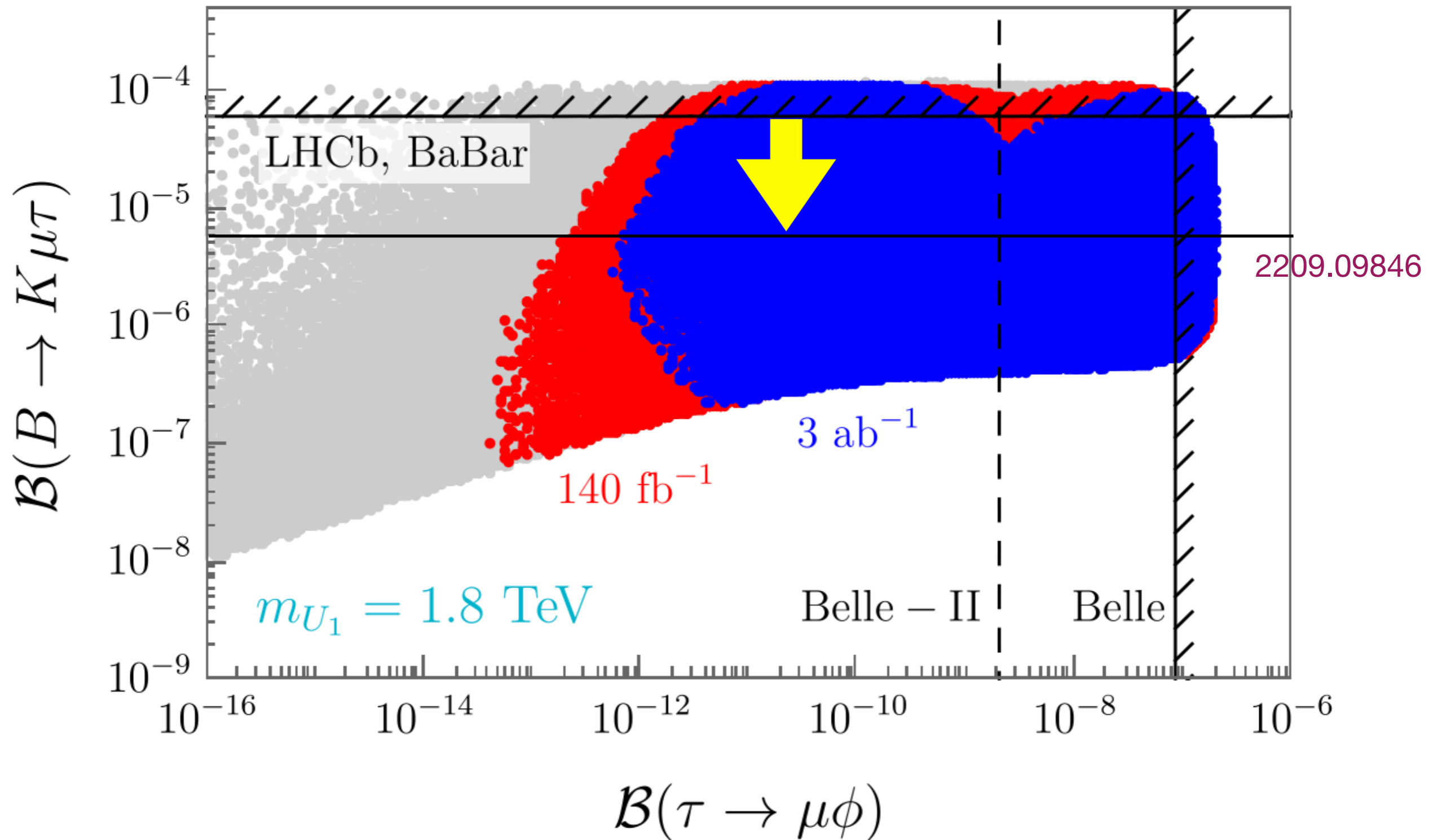
$$\mathcal{L}_{U_1} = x_L^{ij} \bar{Q}_i \gamma_\mu L_j U_1^\mu + x_R^{ij} \bar{d}_{Ri} \gamma_\mu \ell_{Rj} U_1^\mu + \text{h.c.}$$

# LFV predictions

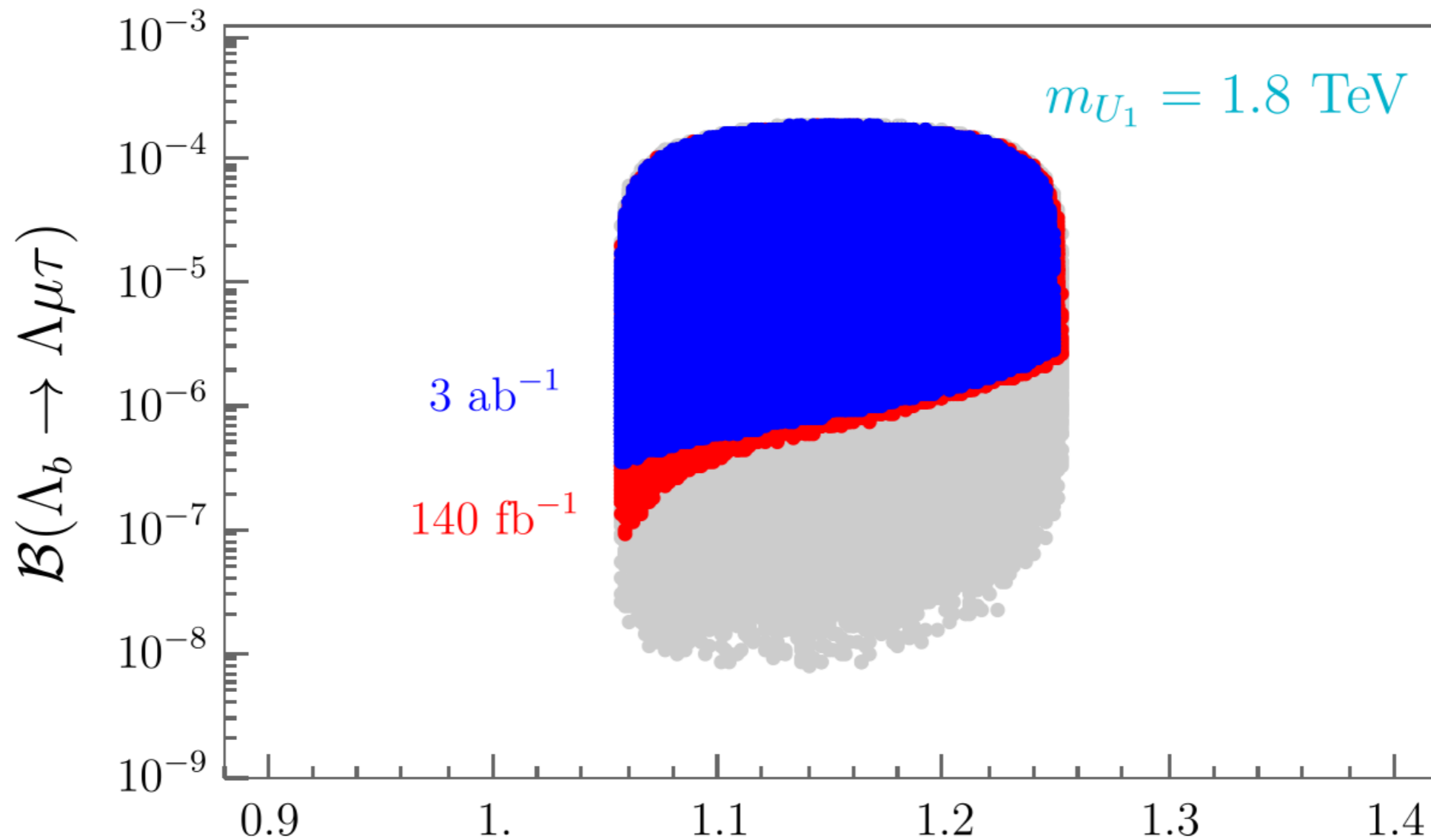




# LFV predictions



# LFV predictions



$$R_{D^{(*)}}/R_{D^{(*)}}^{\text{SM}} = R_{\Lambda_c}/R_{\Lambda_c}^{\text{SM}} = \dots$$

- Way to go 1: Combine two scalar LQs [ $S_1$  with  $S_3$ , or  $R_2$  with  $S_3$ ]
- Way to go 2: Vector LQ ( $U_1$ )  
Non-renormalizable and thus requires UV-completion which can be an opportunity to tackle the hierarchy problem!

# Concerning R2

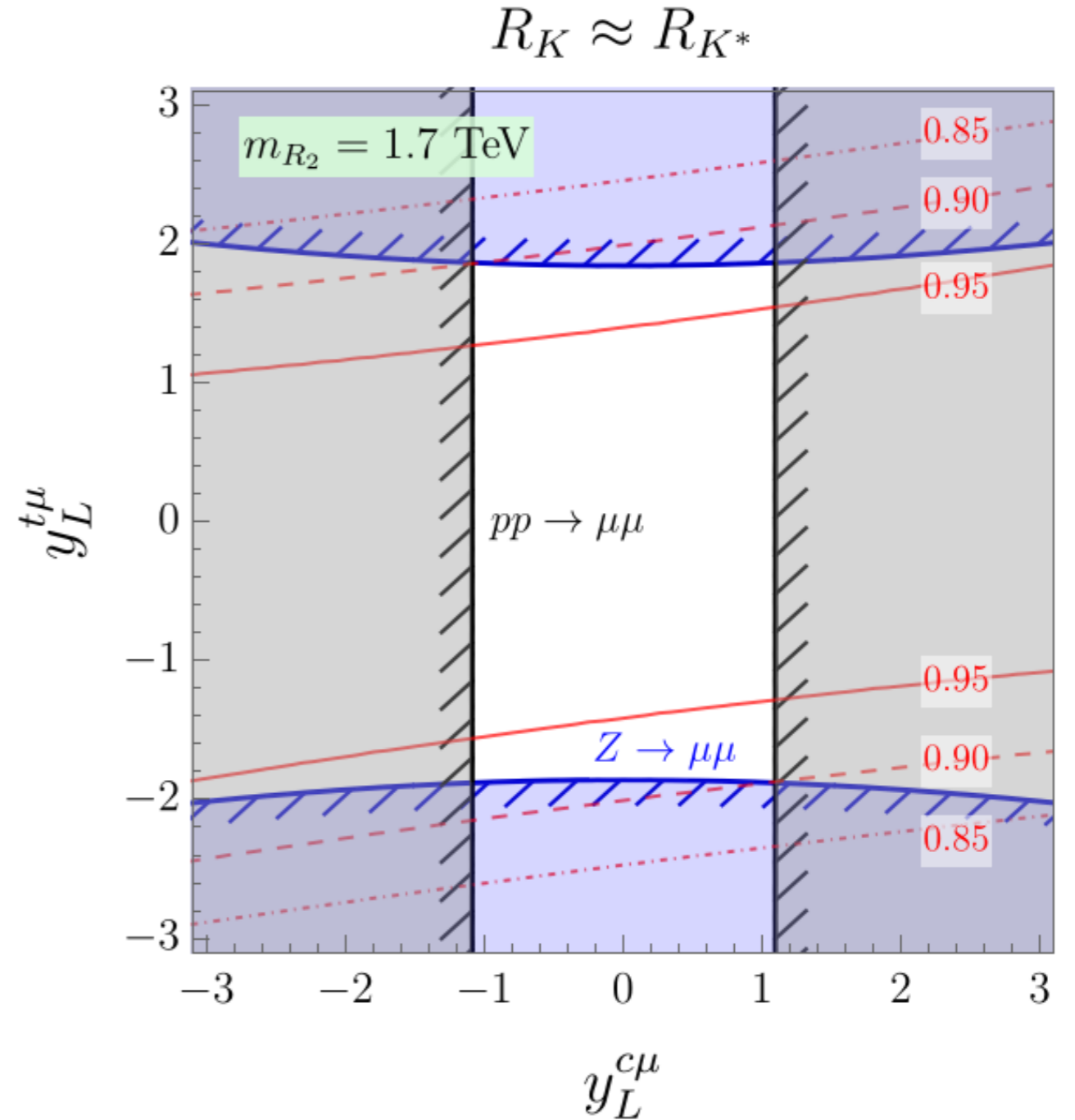
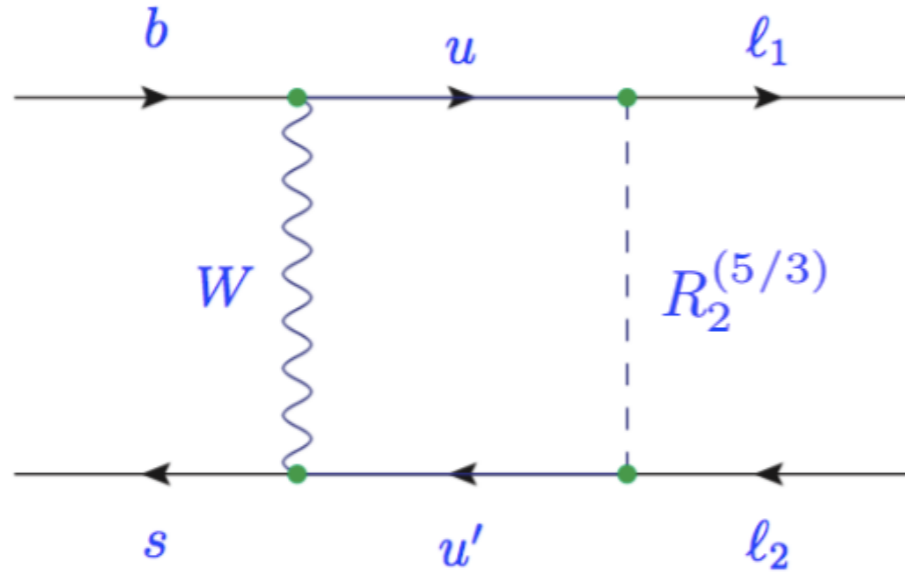
Model	$R_{D(*)}$	$R_{K(*)}$	$R_{D(*)}$ & $R_{K(*)}$
$S_1 = (\bar{3}, 1, 1/3)$	✓	✗	✗
$R_2 = (3, 2, 7/6)$	✓	✓*	✗
$S_3 = (\bar{3}, 3, 1/3)$	✗	✓	✗
$U_1 = (3, 1, 2/3)$	✓	✓	✓
$U_3 = (3, 3, 2/3)$	✗	✓	✗

$$\mathcal{L}_{R_2} = y_R^{ij} \bar{Q}_i \ell_{Rj} R_2 - y_L^{ij} \bar{u}_{Ri} R_2 i\tau_2 L_j + \text{h.c.}$$

$$C_9^{kl} = C_{10}^{kl} \stackrel{\text{tree}}{=} -\frac{\pi v^2}{2V_{tb}V_{ts}^* \alpha_{\text{em}}} \frac{y_R^{sl} (y_R^{bk})^*}{m_{R_2}^2}$$

$$C_9^{kl} = -C_{10}^{kl} \stackrel{\text{loop}}{=} \sum_{u,u' \in \{u,c,t\}} \frac{V_{ub}V_{u's}^*}{V_{tb}V_{ts}^*} y_L^{u'k} (y_L^{ul})^* \mathcal{F}(x_u, x_{u'})$$

$$y_L = \begin{pmatrix} 0 & 0 & 0 \\ 0 & y_L^{c\mu} & 0 \\ 0 & y_L^{t\mu} & 0 \end{pmatrix}, \quad y_R = 0$$



# $S_1$ & $S_3$ Model(s)

- In flavor basis

$$\mathcal{L}_{\text{Yuk}} \supset (y_{S_1}^L)_{ij} \bar{Q}_i^C i\tau_2 L_j S_1 + (y_{S_3}^L)_{ij} \bar{Q}_i^C i\tau_2 L_j (\vec{\tau} \cdot \vec{S}_3) + \text{h.c.}$$

$$S_1 = (\bar{3}, 1, 1/3), \quad S_3 = (\bar{3}, 3, 1/3)$$

- Specifying models

M1:

1703.09226

$$y_{S_1}^L = \begin{pmatrix} 0 & 0 & 0 \\ 0 & \lambda_{s\mu} & \lambda_{s\tau} \\ 0 & \lambda_{b\mu} & \lambda_{b\tau} \end{pmatrix}, \quad y_{S_3}^L = \begin{pmatrix} 0 & 0 & 0 \\ 0 & \lambda_{s\mu} & \lambda_{s\tau} \\ 0 & -\lambda_{b\mu} & -\lambda_{b\tau} \end{pmatrix}$$

M2:

1706.07808

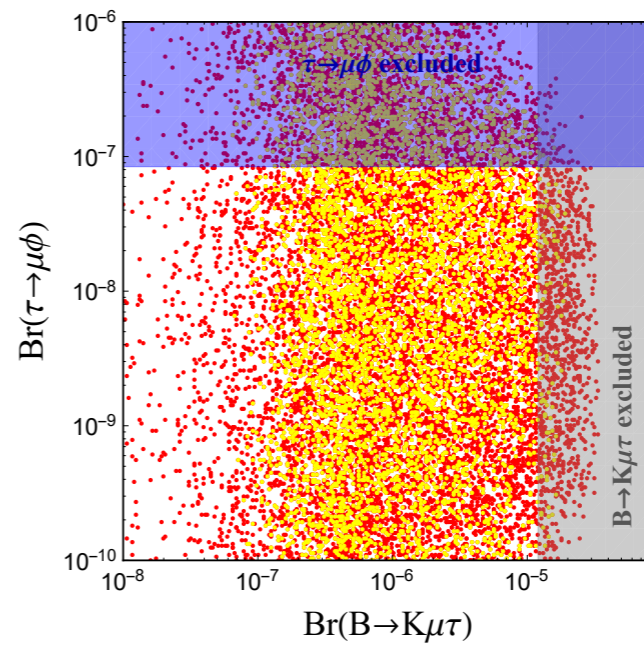
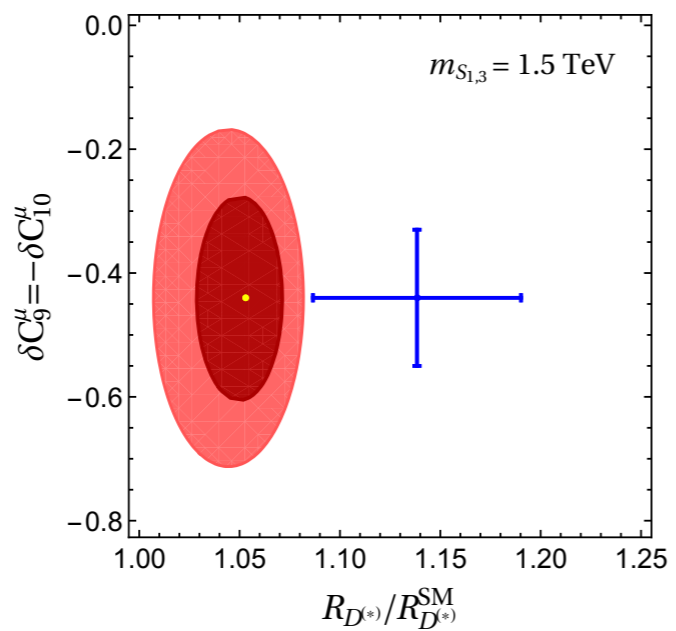
$$y_{S_1}^L = g_{S_1} \begin{pmatrix} 0 & 0 & 0 \\ 0 & \beta_{s\mu} & \beta_{s\tau}^{S_1} \\ 0 & \beta_{b\mu} & 1 \end{pmatrix}, \quad y_{S_3}^L = g_{S_3} \begin{pmatrix} 0 & 0 & 0 \\ 0 & \beta_{s\mu} & \beta_{s\tau}^{S_3} \\ 0 & \beta_{b\mu} & 1 \end{pmatrix}$$

M3:

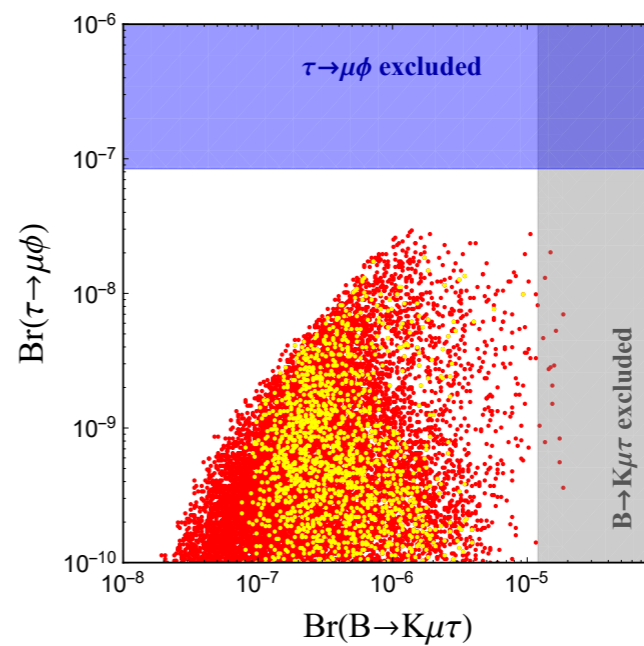
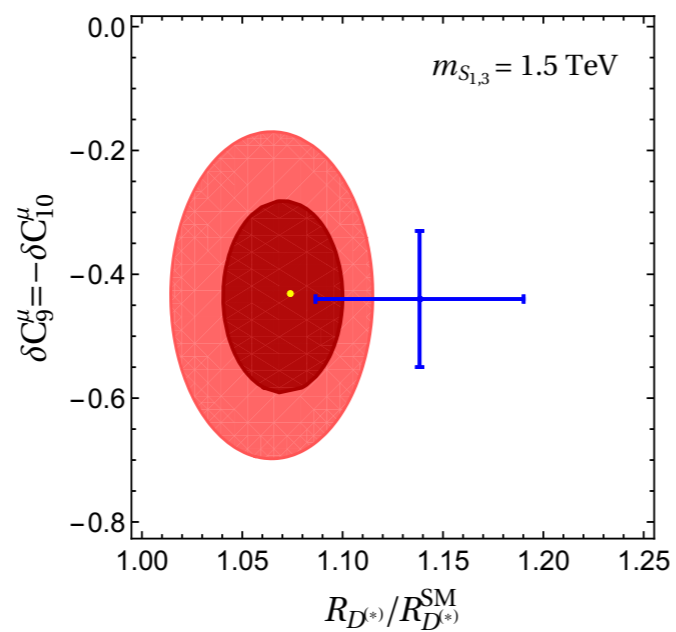
2008.09548

$$y_{S_1}^L = \begin{pmatrix} 0 & 0 & 0 \\ 0 & 0 & y_{s\tau}^{S_1} \\ 0 & 0 & y_{b\tau}^{S_1} \end{pmatrix}, \quad y_{S_3}^L = \begin{pmatrix} 0 & 0 & 0 \\ 0 & y_{s\mu}^{S_3} & y_{s\tau}^{S_3} \\ 0 & y_{b\mu}^{S_3} & y_{b\tau}^{S_3} \end{pmatrix}$$

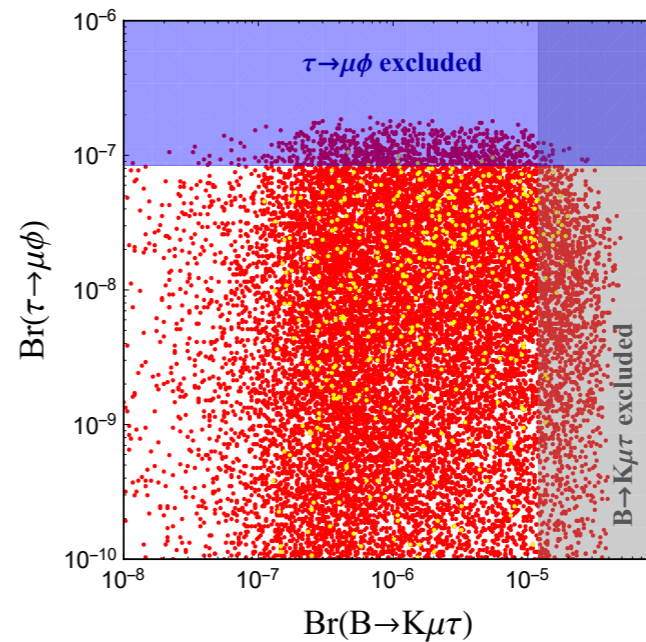
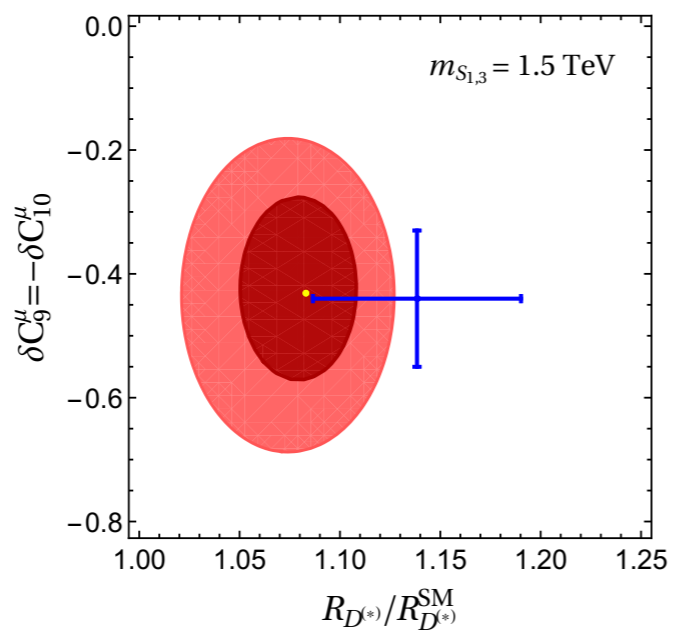
M1:



M2:



M3:



Adding RH couplings considered in 2008.09548 and 1912.04224



# S<sub>3</sub> & R<sub>2</sub> Model

1806.05689

- In flavor basis

$$\mathcal{L} \supset y_R^{ij} \bar{Q}_i \ell_{Rj} R_2 + y_L^{ij} \bar{u}_{Ri} L_j \tilde{R}_2^\dagger + y^{ij} \bar{Q}_i^C i\tau_2 (\tau_k S_3^k) L_j + \text{h.c.}$$

$$R_2 = (3, 2, 7/6), \quad S_3 = (\bar{3}, 3, 1/3)$$

- In mass-eigenstates basis

$$\begin{aligned} \mathcal{L} \supset & (V_{\text{CKM}} y_R E_R^\dagger)^{ij} \bar{u}'_{Li} \ell'_{Rj} R_2^{(5/3)} + (y_R E_R^\dagger)^{ij} \bar{d}'_{Li} \ell'_{Rj} R_2^{(2/3)} \\ & + (U_R y_L U_{\text{PMNS}})^{ij} \bar{u}'_{Ri} \nu'_{Lj} R_2^{(2/3)} - (U_R y_L)^{ij} \bar{u}'_{Ri} \ell'_{Lj} R_2^{(5/3)} \\ & - (y U_{\text{PMNS}})^{ij} \bar{d}'_{Li} \nu'_{Lj} S_3^{(1/3)} - \sqrt{2} y^{ij} \bar{d}'_{Li} \ell'_{Lj} S_3^{(4/3)} \\ & + \sqrt{2} (V_{\text{CKM}}^* y U_{\text{PMNS}})^{ij} \bar{u}'_{Li} \nu'_{Lj} S_3^{(-2/3)} - (V_{\text{CKM}}^* y)^{ij} \bar{u}'_{Li} \ell'_{Lj} S_3^{(1/3)} + \text{h.c.} \end{aligned}$$

and assume

$$\underline{y_R = y_R^T \quad y = -y_L}$$

$$y_R E_R^\dagger = \begin{pmatrix} 0 & 0 & 0 \\ 0 & 0 & 0 \\ 0 & 0 & y_R^{b\tau} \end{pmatrix}, \quad U_R y_L = \begin{pmatrix} 0 & 0 & 0 \\ 0 & y_L^{c\mu} & y_L^{c\tau} \\ 0 & 0 & 0 \end{pmatrix}, \quad U_R = \begin{pmatrix} 1 & 0 & 0 \\ 0 & \cos \theta & -\sin \theta \\ 0 & \sin \theta & \cos \theta \end{pmatrix}$$

Parameters:  $m_{R_2}, m_{S_3}, y_R^{b\tau}, y_L^{c\mu}, y_L^{c\tau}$  and  $\theta$



## Effective Lagrangian at $\mu \approx m_{LQ}$ :

- $b \rightarrow c\tau\bar{\nu}$ :

**NB.**  $\Lambda_{NP}/g_{NP} \approx 1 \text{ TeV}$

$$\propto \frac{y_L^{c\tau} y_R^{b\tau*}}{m_{R_2}^2} \left[ (\bar{c}_R b_L)(\bar{\tau}_R \nu_L) + \frac{1}{4} (\bar{c}_R \sigma_{\mu\nu} b_L)(\bar{\tau}_R \sigma^{\mu\nu} \nu_L) \right] + \dots$$

- $b \rightarrow s\mu\mu$ :

**NB.**  $\Lambda_{NP}/g_{NP} \approx 30 \text{ TeV}$

$$\propto \sin 2\theta \frac{|y_L^{c\mu}|^2}{m_{S_3}^2} (\bar{s}_L \gamma^\mu b_L)(\bar{\mu}_L \gamma_\mu \mu_L)$$

- $\Delta m_{B_s}$ :

$$\propto \sin^2 2\theta \frac{[(y_L^{c\mu})^2 + (y_L^{c\tau})^2]^2}{m_{S_3}^2} (\bar{s}_L \gamma^\mu b_L)^2$$

$\Rightarrow$  Suppression mechanism of  $b \rightarrow s\mu\mu$  wrt  $b \rightarrow c\tau\bar{\nu}$  for **small  $\sin 2\theta$** .

$\Rightarrow$  Phenomenology suggests  $\theta \approx \pi/2$  and  $y_R^{b\tau}$  complex

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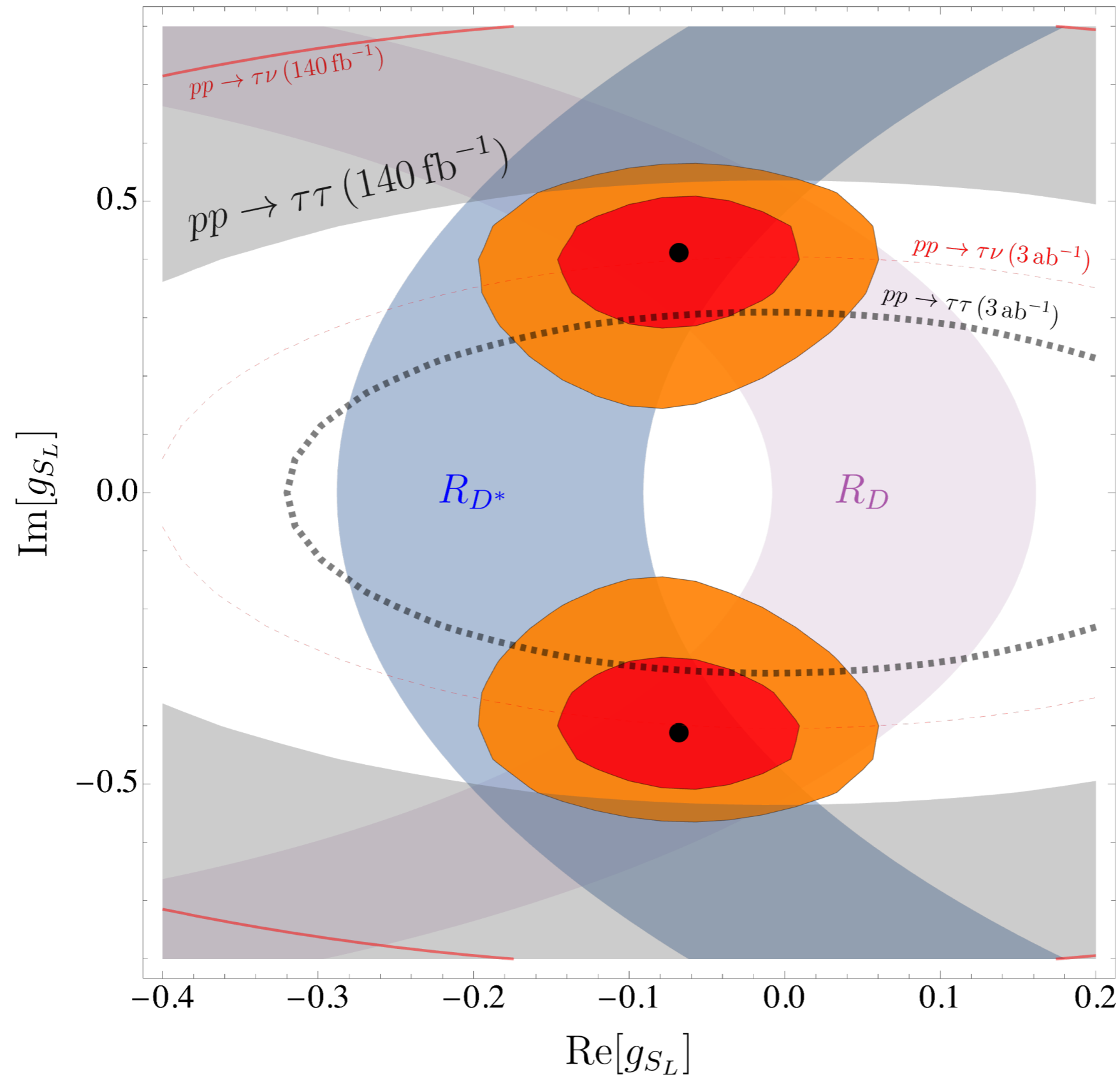
- $\Delta m_{B_s}$ :

$$\propto \sin^2 2\theta \frac{[(y_L^{c\mu})^2 + (y_L^{c\tau})^2]^2}{m_{S_3}^2} (\bar{s}_L \gamma^\mu b_L)^2$$

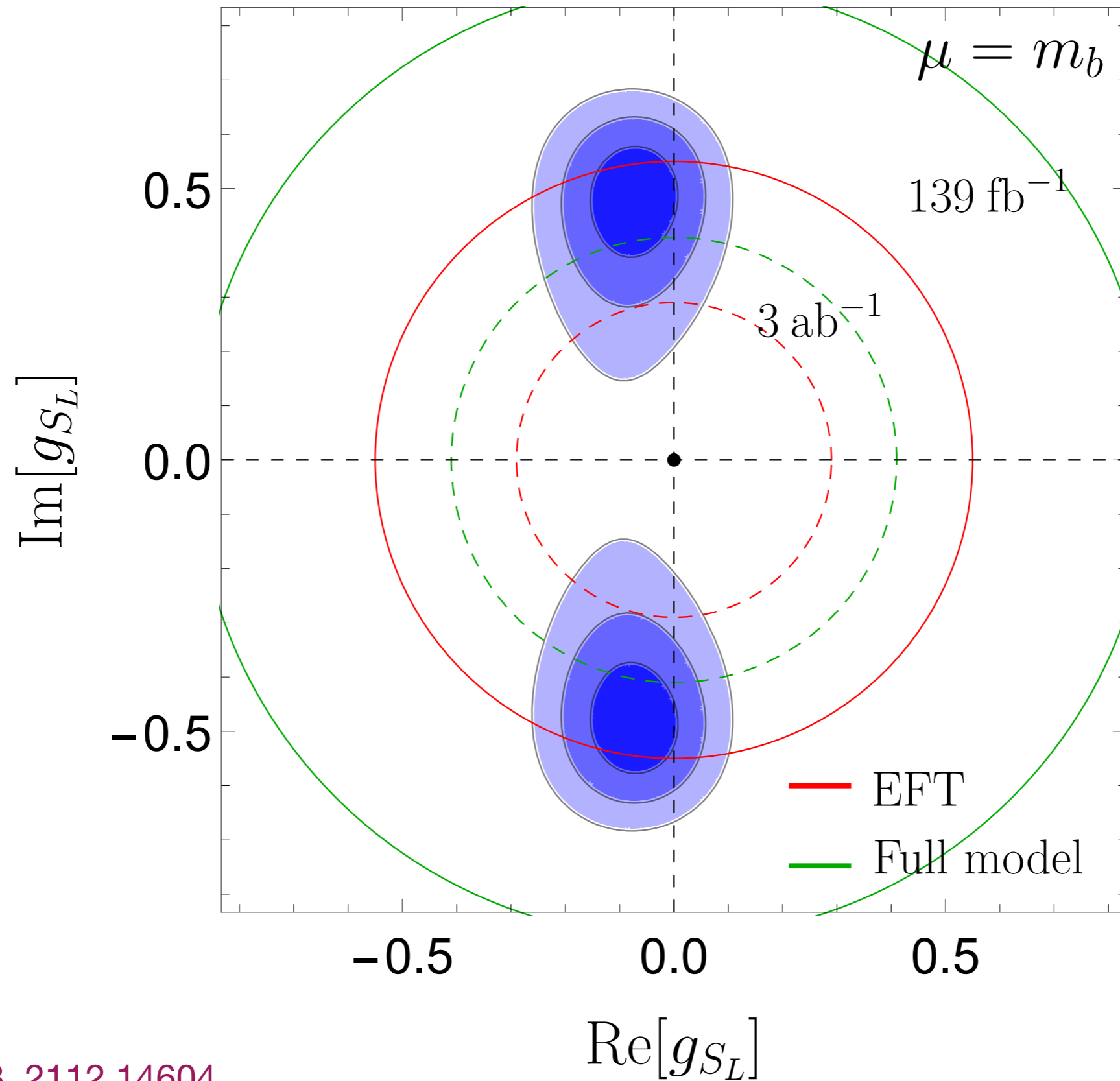
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$\Rightarrow$  Phenomenology suggests  $\theta \approx \pi/2$  and  $y_R^{b\tau}$  complex

$$m_{R_2} = 1.3 \text{ TeV}, m_{S_3} = 2.0 \text{ TeV}$$

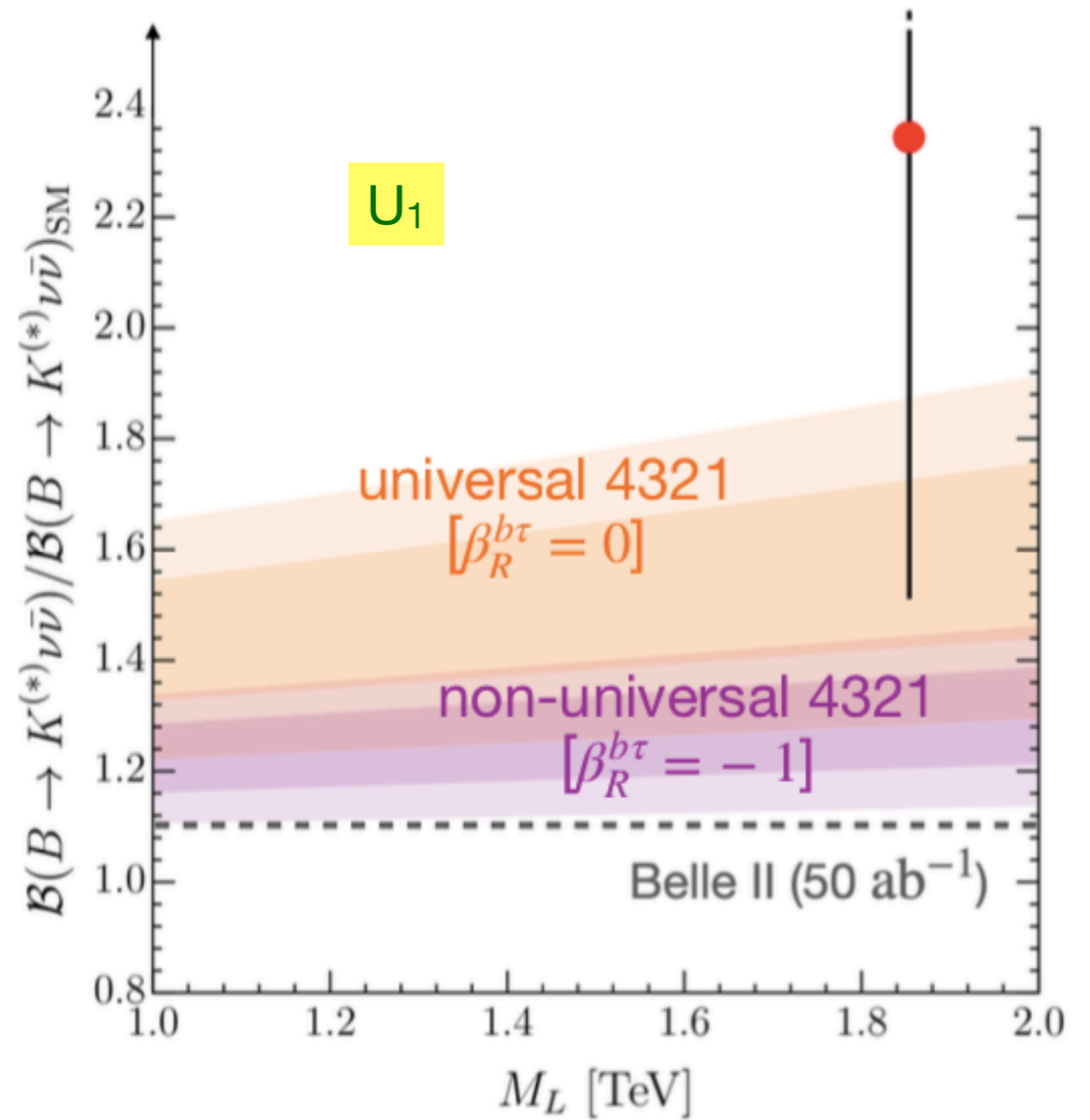
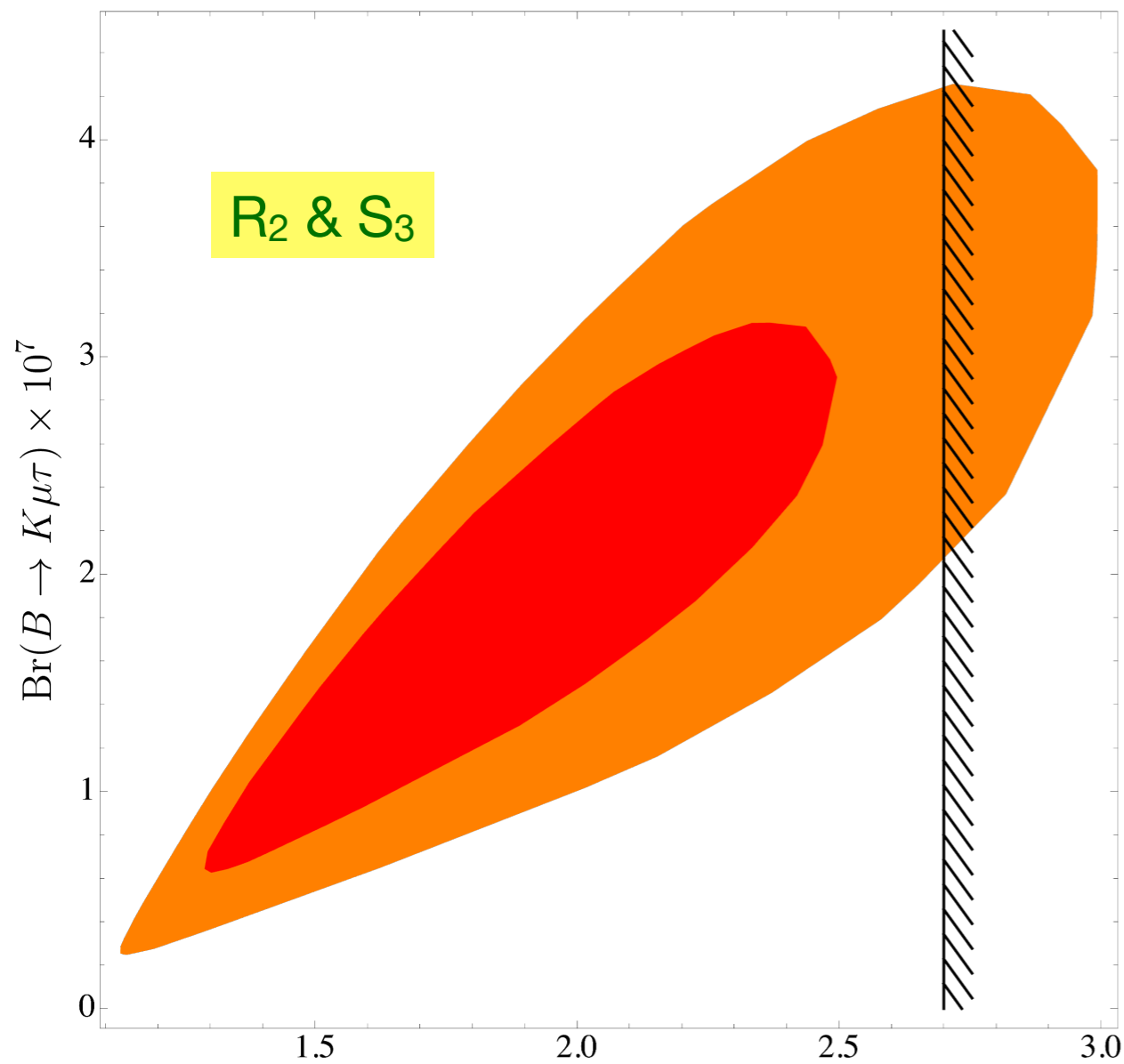


Bounds derived from  $pp \rightarrow \tau\nu$  at high  $p_T$  not useful



# Interesting pheno, ex.

$m_{R_2} = 1.3 \text{ TeV}, m_{S_3} = 2.0 \text{ TeV}$



2206.09717

1806.05689

$$R_{\nu\nu}^{(*)} = \frac{\mathcal{B}(B \rightarrow K^{(*)} \nu\nu)}{\mathcal{B}(B \rightarrow K^{(*)} \nu\nu)^{\text{SM}}}$$

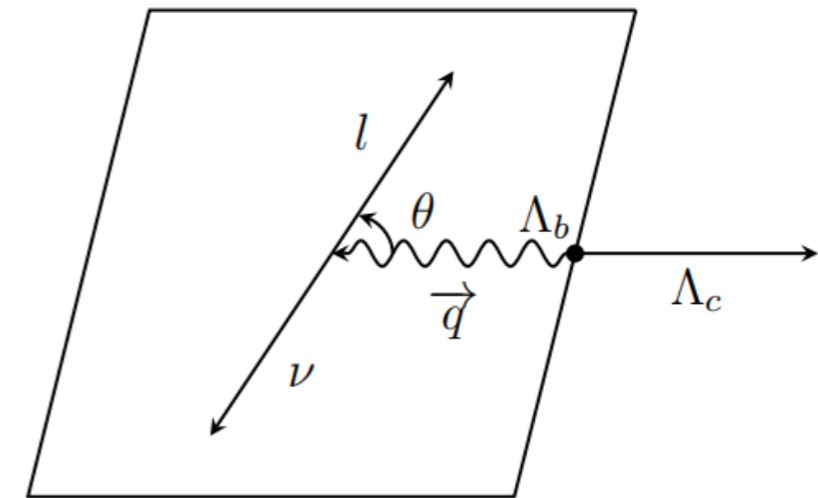
2103.16558

2009.11296

*Belle-II results imminent*

# Angular observables can help disentangling among various NP scenarios

Many works with mesons:  $\mathbf{B} \rightarrow \mathbf{D}\ell\bar{\nu}$        $\mathbf{B} \rightarrow \mathbf{D}^*\ell\bar{\nu}$



Let us now play with baryons:

$$\frac{d^2\Gamma}{dq^2 d\cos\theta} = \frac{\sqrt{\lambda_{\Lambda_b\Lambda_c}(q^2)}}{1024\pi^3 M_{\Lambda_b}^3} \left(1 - \frac{m_l^2}{q^2}\right) \sum_{\lambda_l\lambda_b\lambda_c} \left| \mathcal{M}_{\lambda_c}^{(3)\lambda_b\lambda_l} \right|^2$$

$$\frac{d^2\Gamma(\Lambda_b \rightarrow \Lambda_c^{\lambda_c} \ell^{\lambda_l} \nu)}{dq^2 d\cos\theta} = a_{\lambda_c}^{\lambda_l}(q^2) + b_{\lambda_c}^{\lambda_l}(q^2)\cos\theta + c_{\lambda_c}^{\lambda_l}(q^2)\cos^2\theta$$

Each  $a_{\lambda_c}^{\lambda_l}(q^2)$ ,  $b_{\lambda_c}^{\lambda_l}(q^2)$ ,  $c_{\lambda_c}^{\lambda_l}(q^2)$  is a function of kinematics, form factors and the NP couplings  $g_{V_L}$ ,  $g_{S_L}$ ,  $g_{S_R}$ ,  $g_T$ .

12-2=10 observables

# Angular observables $\Lambda_b \rightarrow \Lambda_c \ell \bar{\nu}$

1907.12554

1908.02328

1909.10769

1702.02243

1502.04864

2209.13409

Three powerful observables:

$$\circ \mathcal{A}_{\text{fb}}(q^2) = \frac{1}{\Gamma} \left[ \int_0^1 - \int_{-1}^0 \right] \frac{d\Gamma}{d \cos \theta} d \cos \theta$$

$$\circ \mathcal{A}_{\pi/3}(q^2) = \frac{1}{\Gamma} \left[ \int_0^{\pi/3} + \int_{2\pi/3}^{\pi} - \int_{\pi/3}^{2\pi/3} \right] \frac{d\Gamma}{d \cos \theta} \sin \theta d\theta$$

$$\circ \mathcal{A}_\lambda(q^2) = \frac{1}{\Gamma} \left[ \frac{d\Gamma^+}{dq^2} - \frac{d\Gamma^-}{dq^2} \right]$$

Examples:

$$U_1 : g_{V_L}$$

$$R_2 : g_{S_L} = 4 g_T$$

$$S_1 : g_{S_L} = -4 g_T$$

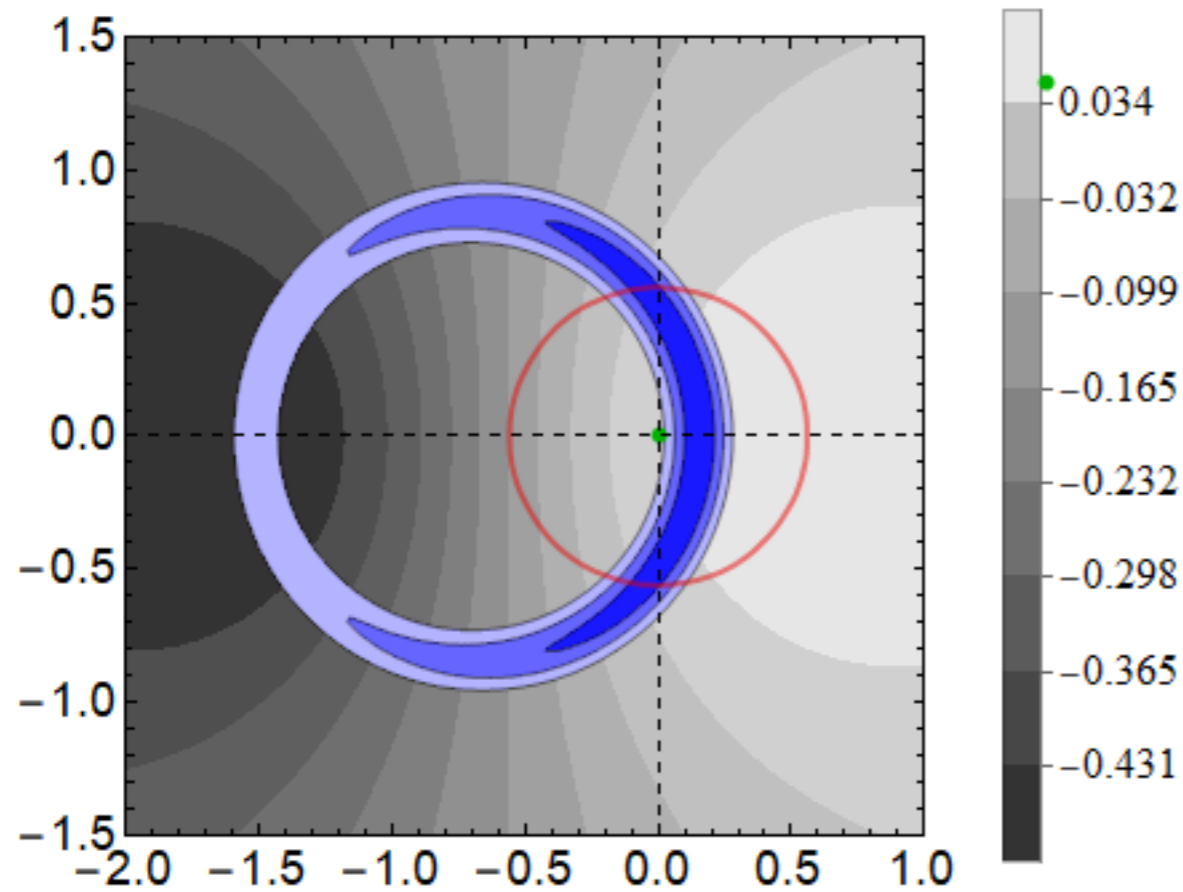


# Angular observables $\Lambda_b \rightarrow \Lambda_c \ell \bar{\nu}$

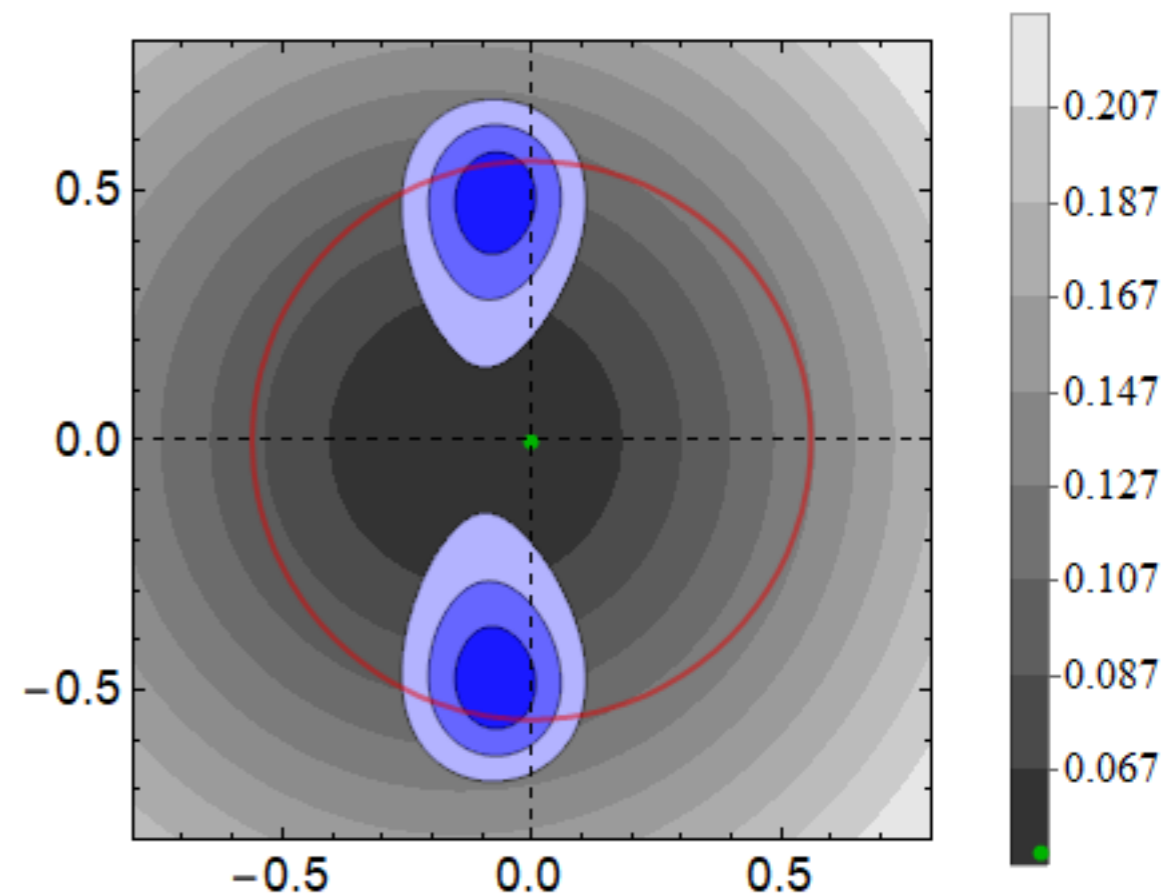
2209.13409

Three powerful observables:

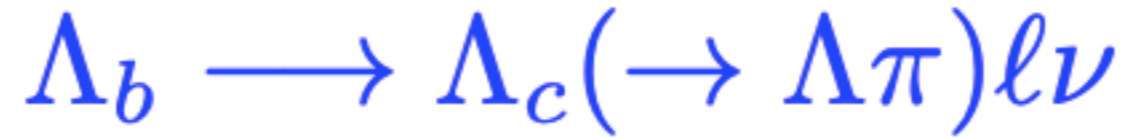
$$\langle A_{\text{fb}}^\tau \rangle$$



$S_1$



$R_2$



2209.13409

NB:  $\mathcal{B}(\Lambda_c \rightarrow \Lambda \pi) = 1.30(7)\%$  or  $\mathcal{B}(\Lambda_c \rightarrow p K_S) = 1.59(8)\%$

Many more angular observables and checking on  $\text{Im}[g_x] \neq 0$

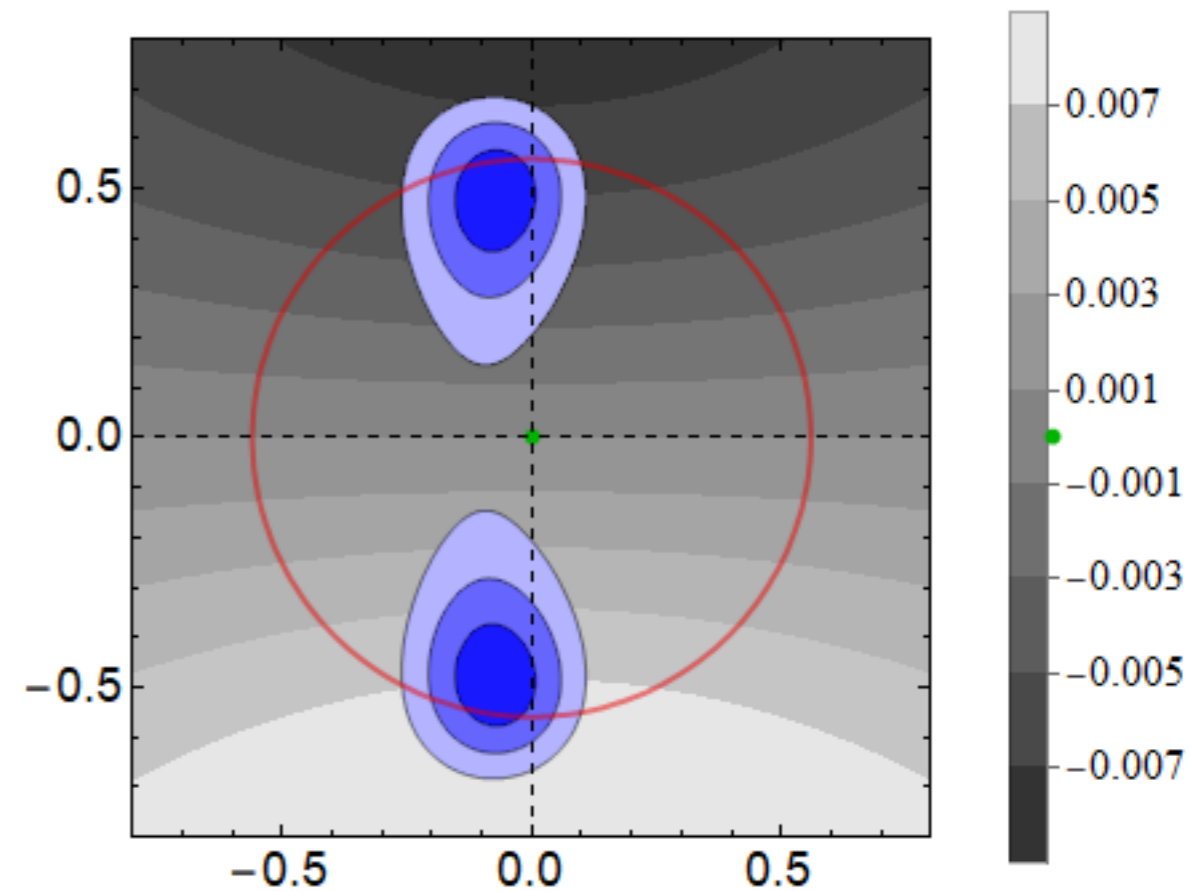
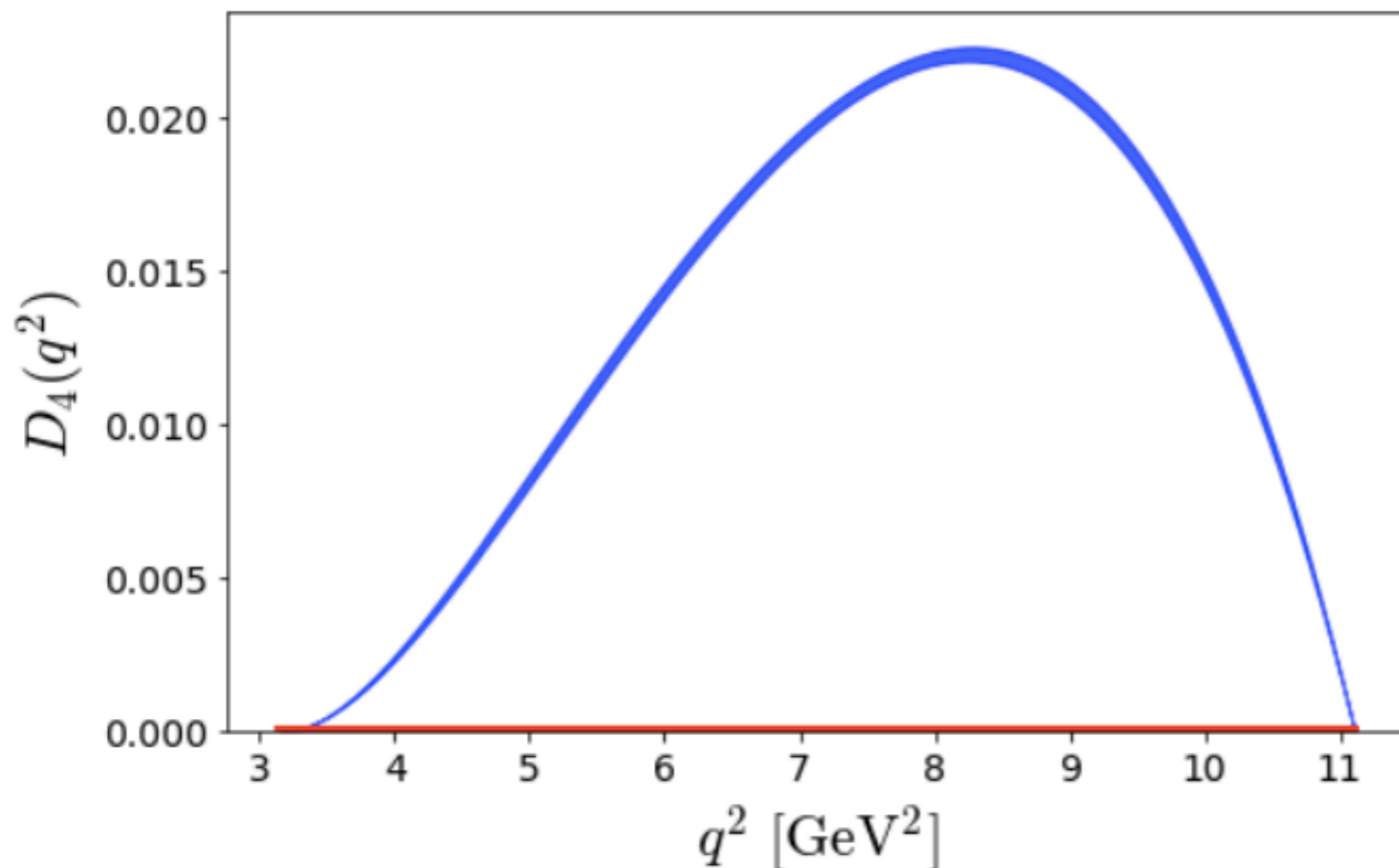
$$\begin{aligned} \dots \\ \frac{d^4 \Gamma^{\lambda_\ell}}{dq^2 d \cos \theta d \cos \theta_\Lambda d \phi} &= A_1^{\lambda_\ell} + A_2^{\lambda_\ell} \cos \theta_\Lambda \\ &- \left( B_1^{\lambda_\ell} + B_2^{\lambda_\ell} \cos \theta_\Lambda \right) \cos \theta && \mathcal{A}_{\text{fb}} \text{ Baryon} \\ &+ \left( C_1^{\lambda_\ell} + C_2^{\lambda_\ell} \cos \theta_\Lambda \right) \cos^2 \theta && \mathcal{A}_{\text{fb}} \text{ Lepton} \\ &+ \left( D_3^{\lambda_\ell} \sin \theta_\Lambda \cos \phi + D_4^{\lambda_\ell} \sin \theta_\Lambda \sin \phi \right) \sin \theta \\ &+ \left( E_3^{\lambda_\ell} \sin \theta_\Lambda \cos \phi + E_4^{\lambda_\ell} \sin \theta_\Lambda \sin \phi \right) \sin \theta \cos \theta \end{aligned}$$

$$\Lambda_b \longrightarrow \Lambda_c (\longrightarrow \Lambda \pi) \ell \nu$$

NB:  $\mathcal{B}(\Lambda_c \rightarrow \Lambda \pi) = 1.30(7)\%$  or  $\mathcal{B}(\Lambda_c \rightarrow p K_S) = 1.59(8)\%$

Many more angular observables and checking on  $\text{Im}[g_x] \neq 0$

$$\langle D_4^\tau \rangle$$



$R_2$

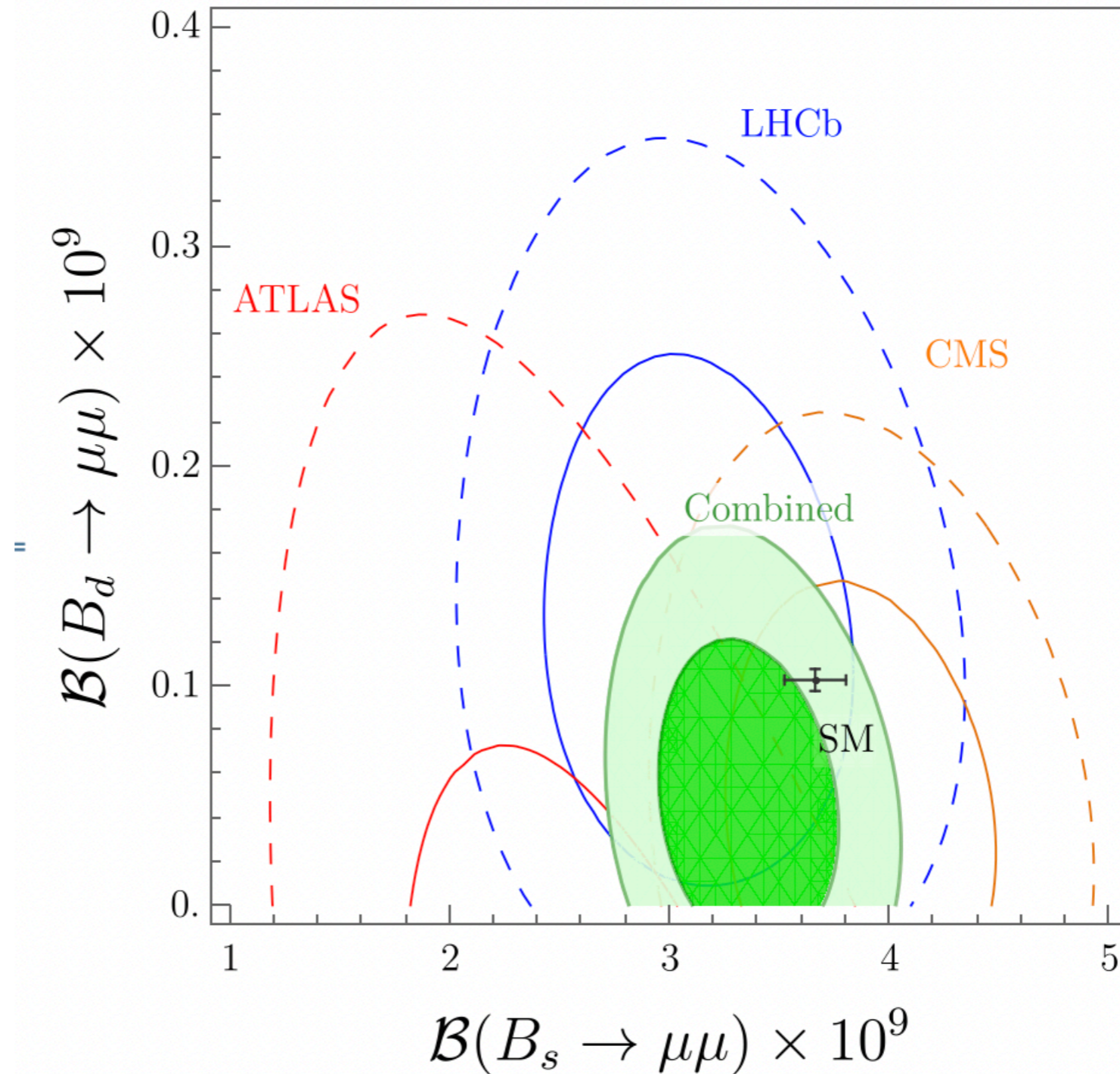
# CONCLUDING REMARKS

- To tackle the flavor issues need more and better data: NP effects through high  $p_T$  tails of  $pp$  scattering are to be combined with constraints obtained from low energy observables.
- EFT treatment is a modern tool to describe deviations wrt SM.
- Current data, and LFUV in particular, favor leptoquark scenarios:  $U_1, S_1$  &  $S_3, R_2$  &  $S_3$ .
- $R_D$  and  $R_{D^*}$  are too few observables to understand the source of LFUV. Too many NP solutions exist and could be filtered through angular  $B \rightarrow D^{(*)} \tau \nu$  and  $\Lambda_b \rightarrow \Lambda_c \tau \nu$  observables.
- Even if  $R_D$  and  $R_{D^*}$  were SM-like, angular observables can help unveiling a presence of BSM physics.
- What about  $R_K$  and  $R_{K^*}$ ? [From rumors to actual results. Please help!]
- $B(B_s \rightarrow \mu\mu)$  this summer...

# CONCLUDING REMARKS

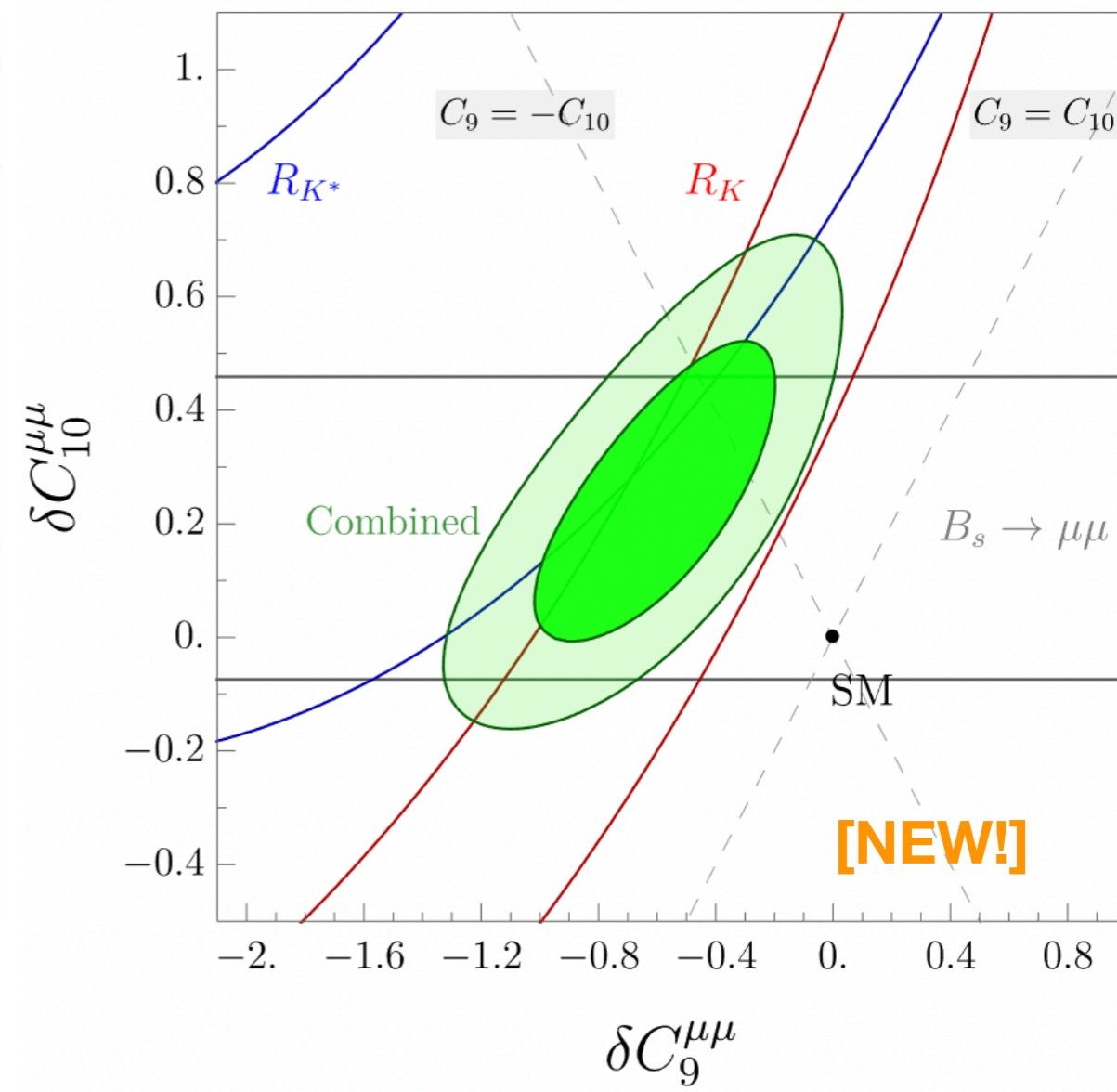
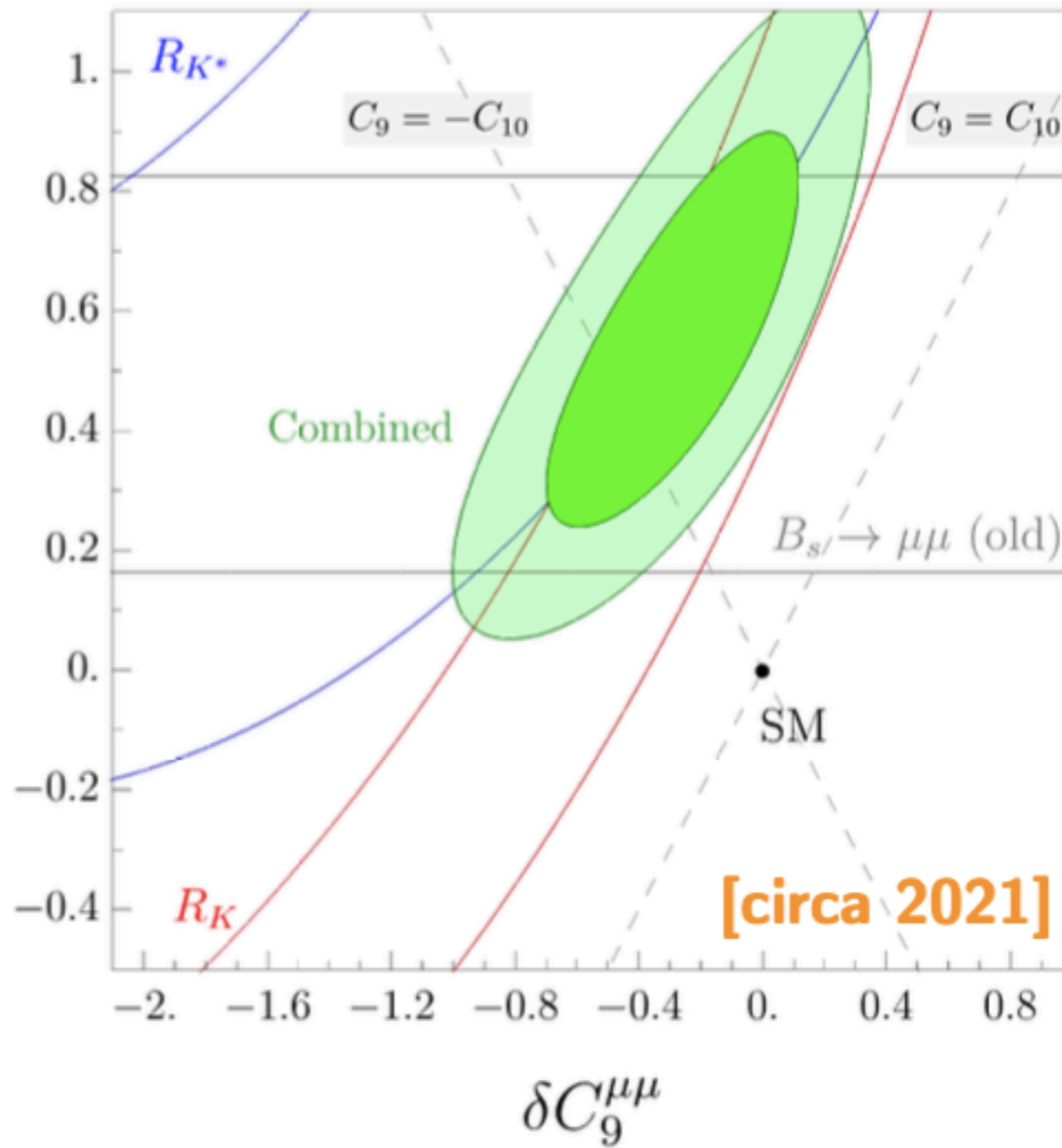
- New LHCb determination of  $f_s/f_d$  shifted their  $B(B_s \rightarrow \mu\mu)$  upwards
- New CMS washed out the deficit of  $B(B_s \rightarrow \mu\mu)$  wrt SM

CMS-PAS-BPH-21-006





# CONCLUDING REMARKS



Quo vadis homini?  
*I šta ćemo sad?*

Experiment is hard...