

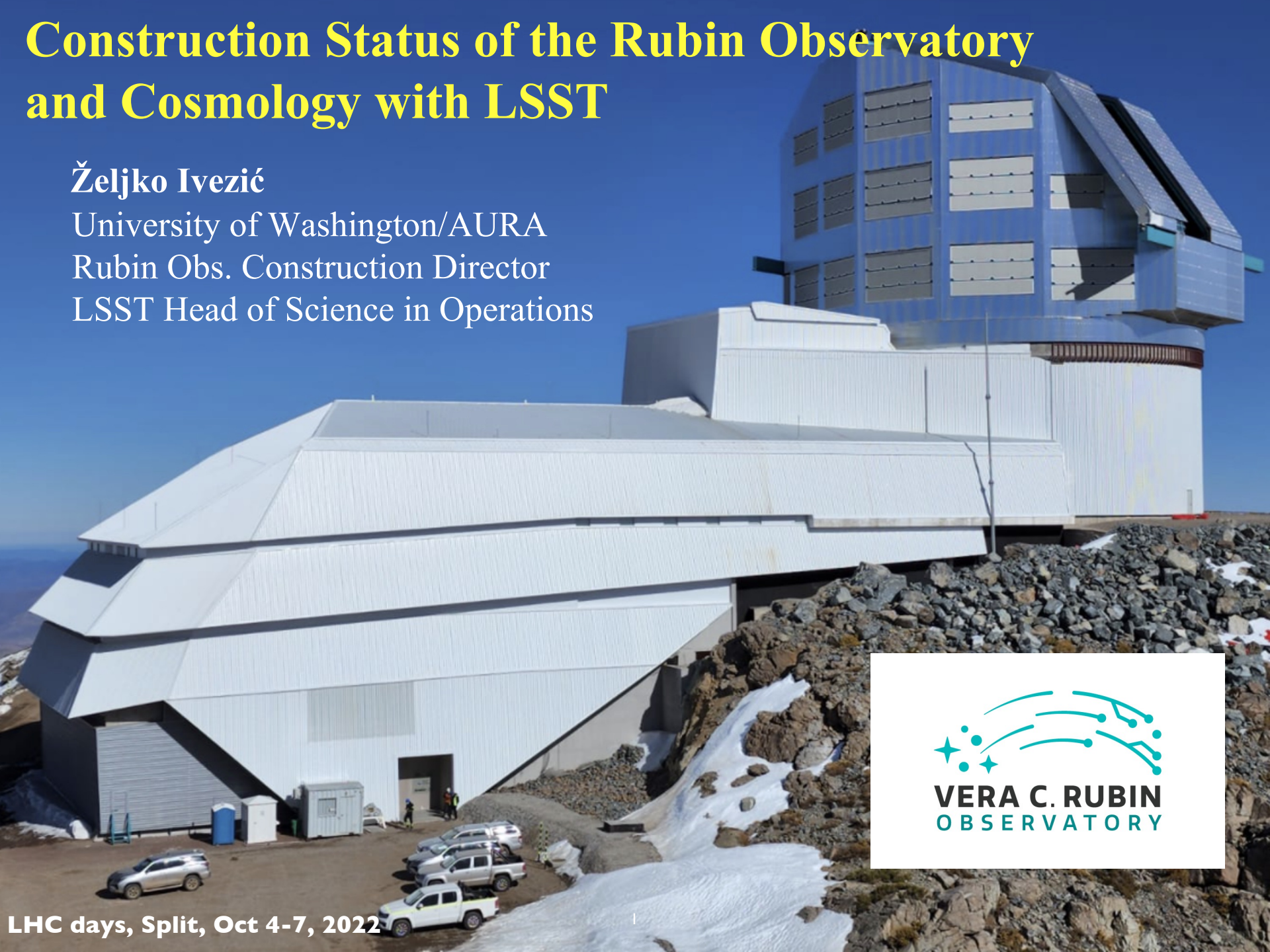
Construction Status of the Rubin Observatory and Cosmology with LSST

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LSST Head of Science in Operations



Construction Status of the Rubin Observatory and Cosmology with LSST



LHC days, Split, Oct 4-7, 2022

Outline

- Brief introduction to science drivers for the Rubin Observatory and LSST
- Rubin Observatory construction status

Legacy Survey of Space and Time (LSST) will be the first 10-year project at the Vera C. Rubin Observatory

What did we learn so far from cosmological measurements?

- the expansion of the Universe is accelerating but we don't know whether there is really a new major (70%) constituent of mass-energy in the Universe, **or perhaps Einstein was wrong?**
- with existing data cannot tell which hypothesis is more likely explanation: to change this, we need to obtain precise data for about **10 billion galaxies**
- to get such an unprecedented dataset, we need a very unique observatory: **Rubin Observatory and LSST**; we need **both** a large telescope mirror to be sensitive enough to faint galaxies, and a large field-of-view for fast sky scanning speed

Cosmology with LSST: dark energy vs. modified gravity

- Even for a model with modified gravity, it is possible to assume that GR is correct and always find "dark energy" with suitable $w(z)$ to explain data for $H(z)$. 😲
- However, the growth of structure will be different and thus when **both** $H(z)$ and $G(z)$ are measured, the degeneracy can be broken and DE vs. modified gravity models distinguished (Jain & Zhu 2008, PhysRevD 78, 063503)

$$ds^2 = -(1 + 2\psi) dt^2 + (1 - 2\phi) a^2(t) d\vec{x}^2$$

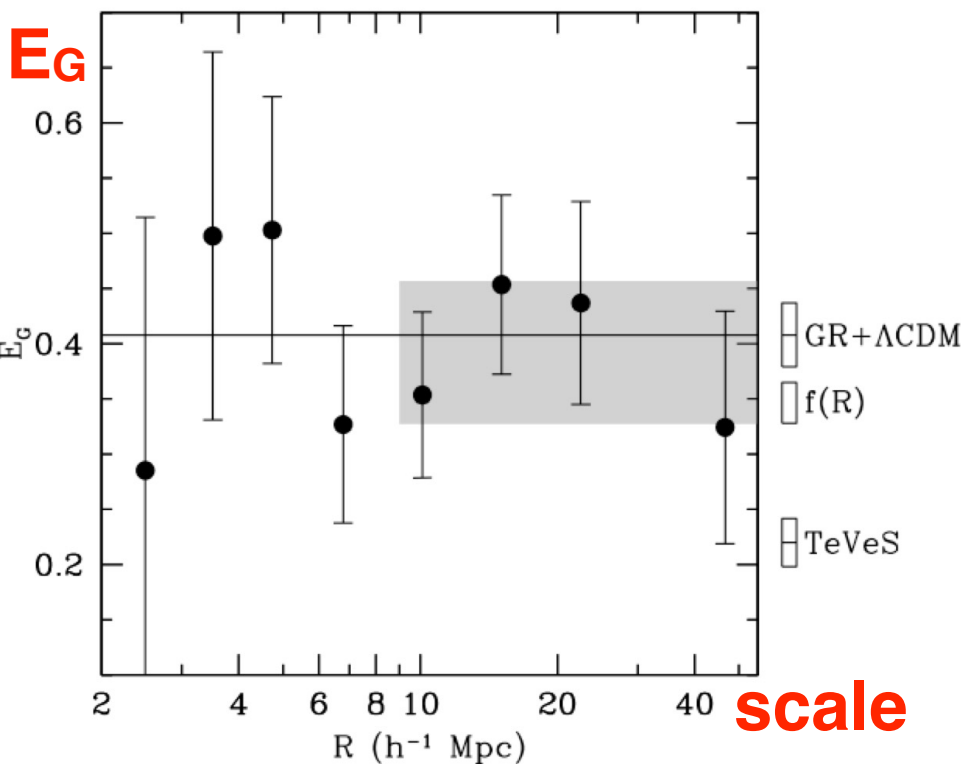
- ϕ is the curvature perturbation and ψ is the potential perturbation
- In General Relativity $\phi = \psi$ in the absence of anisotropic stresses. A metric theory of gravity relates the two potentials above to the perturbed energy-momentum tensor.
 ϕ and ψ can be constrained with astronomical observations 😊

Modern Cosmological Probes

- **Cosmic Microwave Background**
(the state of the Universe at the recombination epoch, at redshift ~ 1000)
- **Weak Lensing:** growth of structure Possible with LSST
- **Galaxy Clustering:** growth of structure
- **Baryon Acoustic Oscillations:** standard ruler
- **Supernovae:** standard candle

Except for CMB, measuring $H(z)$ and growth of structure $G(z)$
 $H(z) \sim d[\ln(a)]/dt$, $G(z) = a^{-1}\delta\rho_m/\rho_m$, with $a(z) = (1+z)^{-1}$

Cosmology with LSST: dark energy vs. modified gravity



- **E_G** combines 3 measures of large-scale structure: galaxy-galaxy lensing ($\phi+\psi$), galaxy clustering (ϕ) and galaxy velocities (from galaxy redshifts; measures $G(z)$)

SDSS data enabled a test of GR at 15% level: it passed!

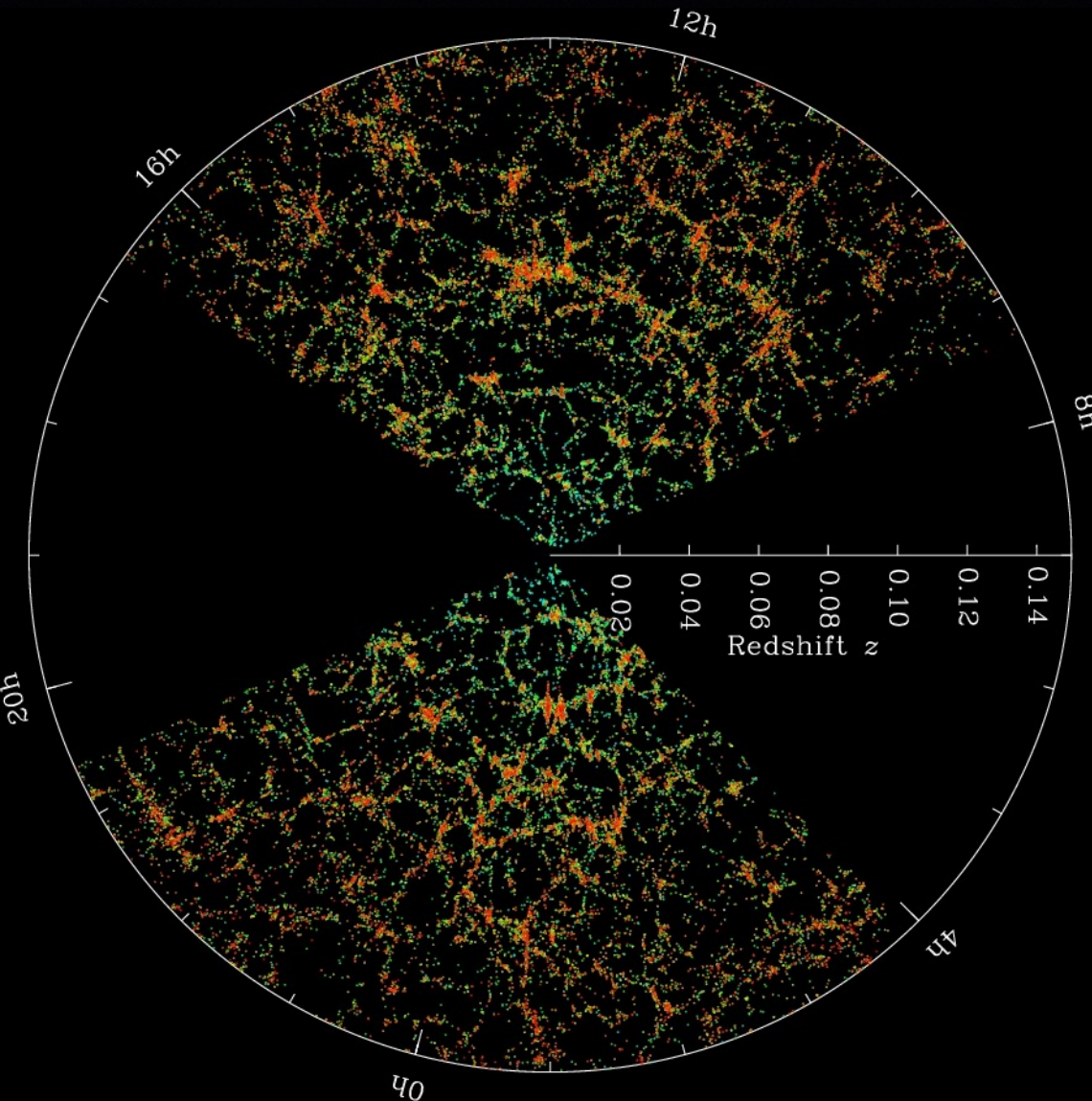
- SDSS data already excludes a model within the tensor-vector-scalar gravity theory, which modifies both Newtonian and Einstein gravity.

Five times better precision needed to rule out $f(R)$

Reyes et al. (2010, Nature 464, 256)

LSST will measure E_G about 10 times more precisely and will be able to rule out a large class of modified gravity theories (or GR!)

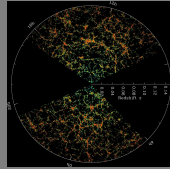
Spatial distribution of galaxies



Left: each dot is one galaxy from SDSS

Note that the galaxy distribution is highly **inhomogeneous**: statistical details of that distribution contain rich cosmological information

Spatial distribution of galaxies



LSST goal:

map galaxies to a ten times larger distance limit than possible today!

20 billion galaxies!

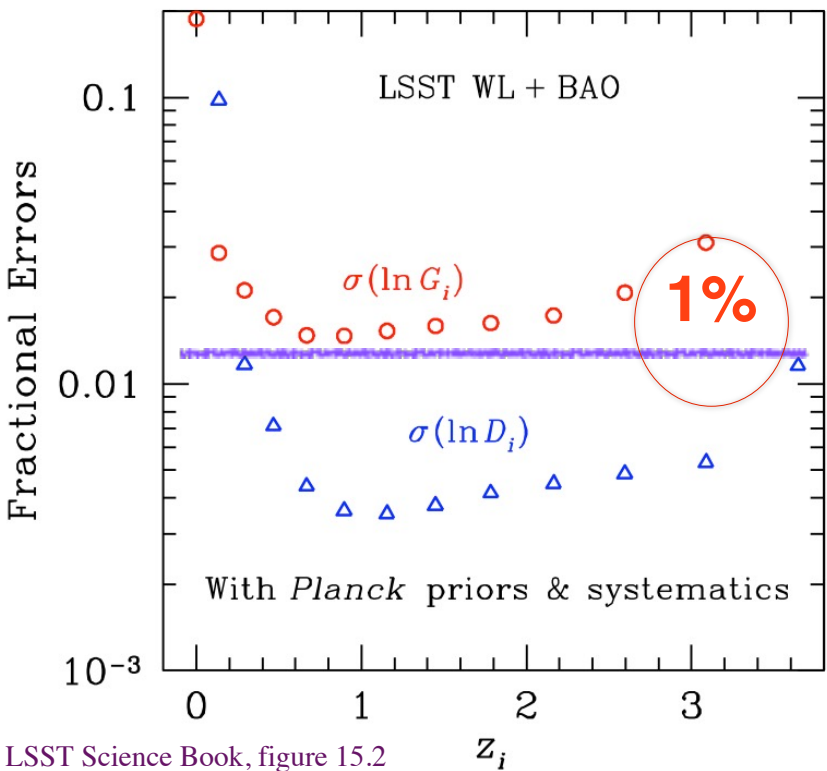
Precise cosmological parameter constraints.

Sum of neutrino masses with accuracy of ~ 0.04 eV from the power spectrum of dark matter fluctuations.

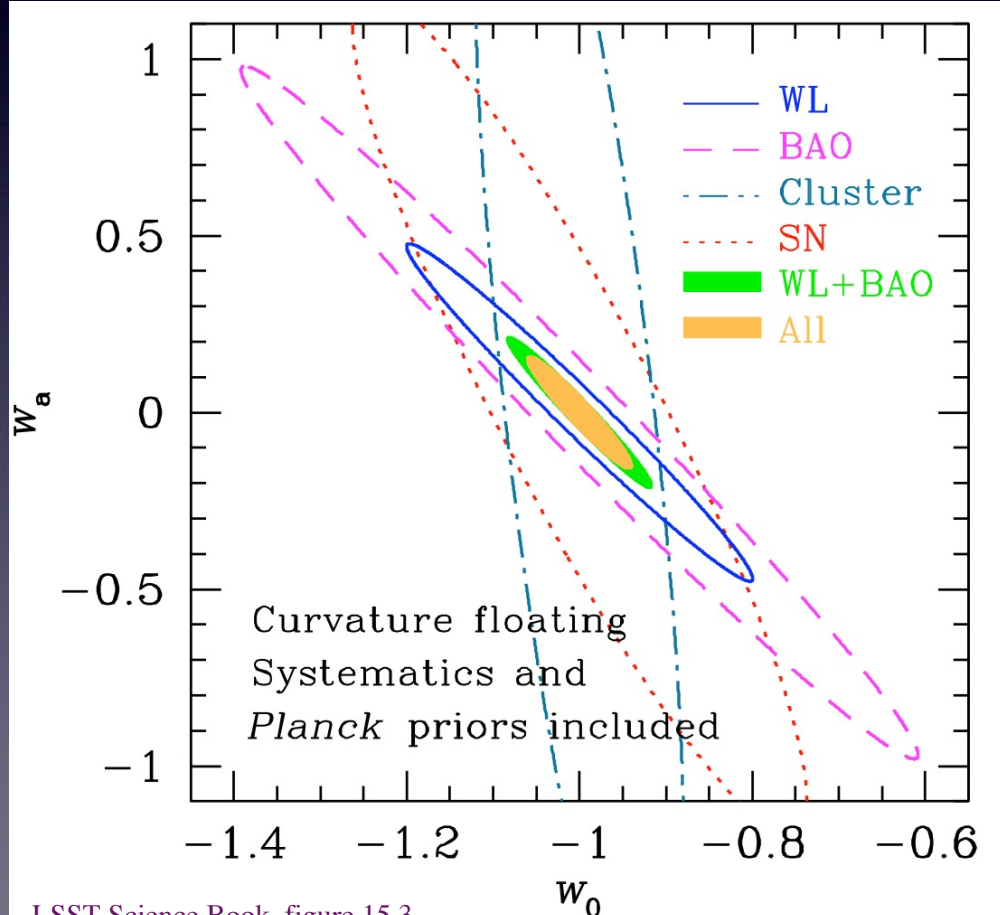
Cosmology with LSST: high precision measurements

- Measuring distances, $H(z)$, and growth of structure, $G(z)$, with a percent accuracy for $0.5 < z < 3$

- Multiple probes is the key!



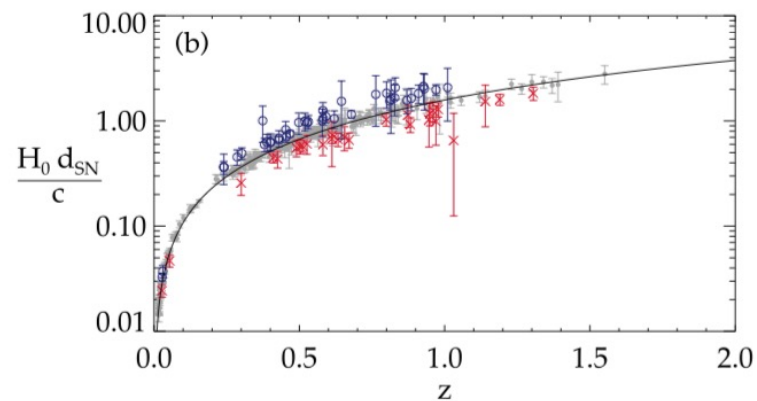
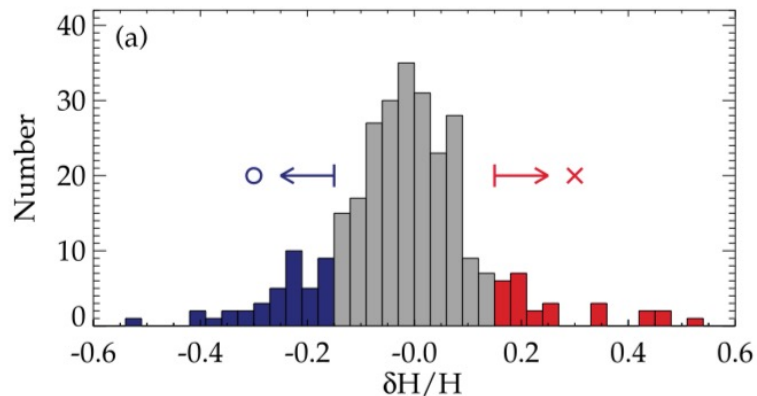
LSST Science Book, figure 15.2



LSST Science Book, figure 15.3

By simultaneously measuring growth of structure and curvature, LSST data will tell us whether the recent acceleration is due to **dark energy** or **modified gravity**.

Cosmology with LSST SNe: is the cosmic acceleration the same in all directions?



Cooke & Lynden-Bell (2009, MNRAS 401, 1409)

Is there spatial structure in the SNe distance modulus residuals for the concordance model?

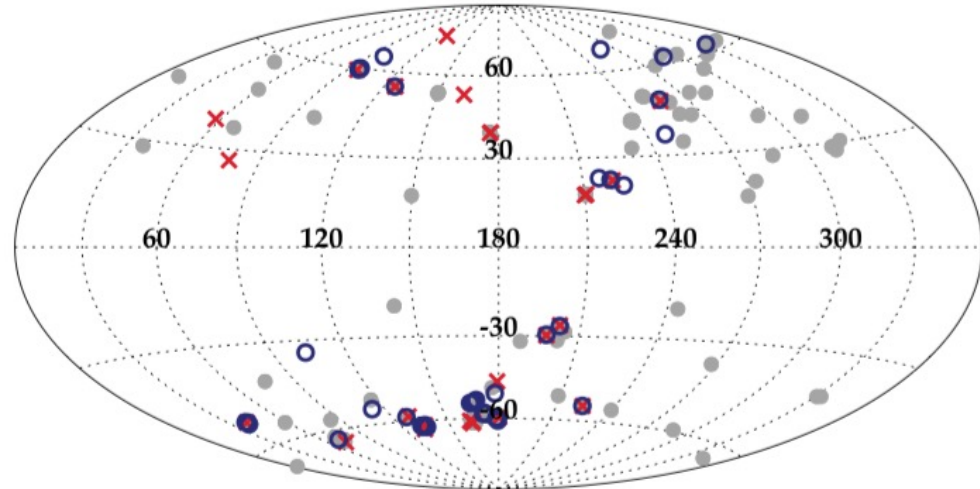


Figure 1. A projection of the spatial distribution of the Union SNe Ia sample in Galactic coordinates. Note the relative uniformity of the points, except around the Galactic plane. The symbols correspond to those in Fig. 2, and are explained in Section 3.1.

- Even a single supernova represents a cosmological measurement!
- LSST will obtain light curves for several million Type Ia supernovae!

Science motivation for undertaking the Legacy Survey of Space and Time:

More details about
science drivers and
system design:
Ivezic et al. (2019)
[ls.st/lop](#)

Expansion and history of the Universe and the growth of structure
(dark matter, dark energy, cosmology, spatial distribution of galaxies, gravitational lensing, supernovae): “[Was Einstein right?](#)“

Time domain: what changes on the sky?
(cosmic explosions, variable stars, [unknown unknowns](#))

The Solar System structure
(near-Earth [hazardous asteroids](#), main-belt asteroids, trans-Neptunian objects, comets)

The Milky Way structure
(stars as tracers of the [structure and evolution of our Galaxy](#), interstellar matter, the physics of stars)

A key point: most of science programs will utilize the same dataset.

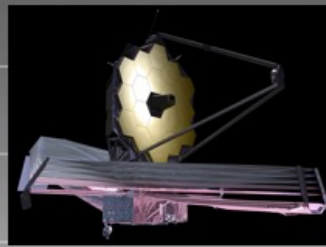
100 m

80 m

60 m

40 m

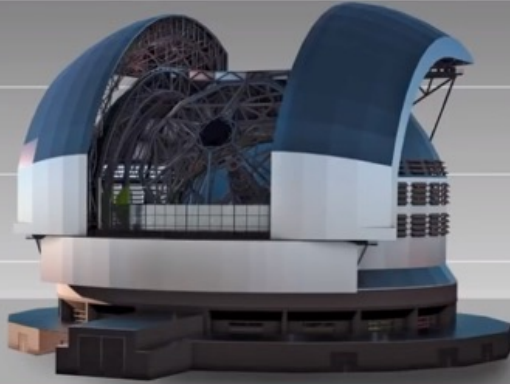
20 m



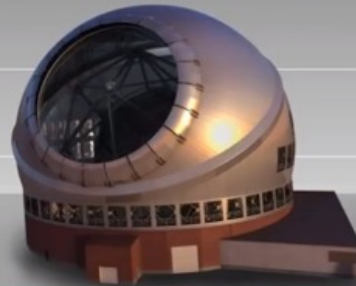
**James Webb
Space Telescope
(6.5m)**



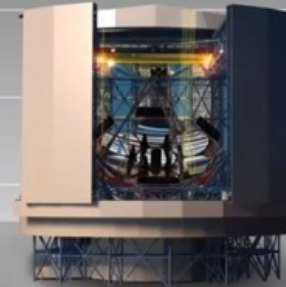
**Nancy Grace
Roman Space
Telescope (2.4m)**



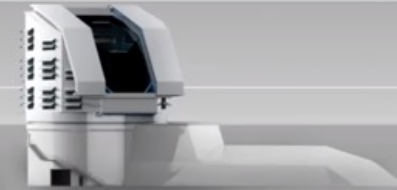
ELT: 40m



TMT: 30m



GMT: 30m



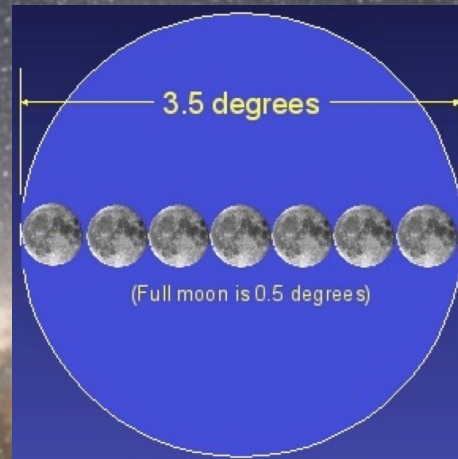
Rubin: 8m

Rubin Obs. will not have the largest mirror but will have by far the largest product of the mirror area and the field-of-view size (etendue or throughput)

LSST will be delivered by the Vera C. Rubin Observatory, as its first, 10-year, project



Every circle
contains
10 million
galaxies



Andy Connolly
University of WA

To deliver LSST, Rubin Observatory will observe about half of the sky close close to 1000 times over 10 years.

Field-of-view comparison: Gemini vs. LSST

Primary Mirror Diameter

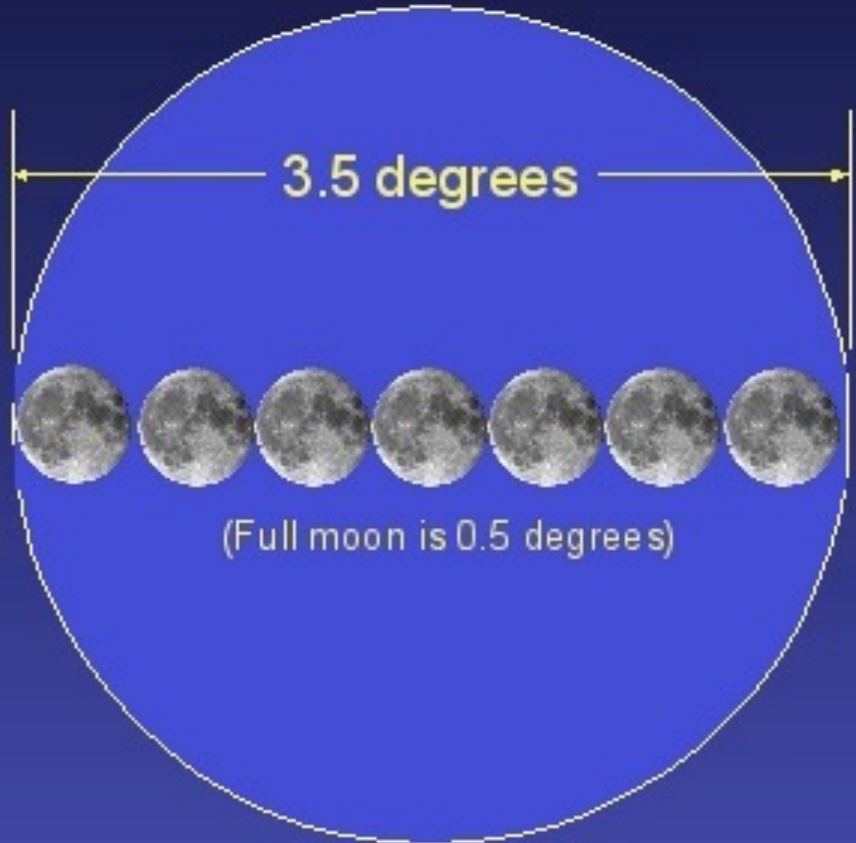
Field of View



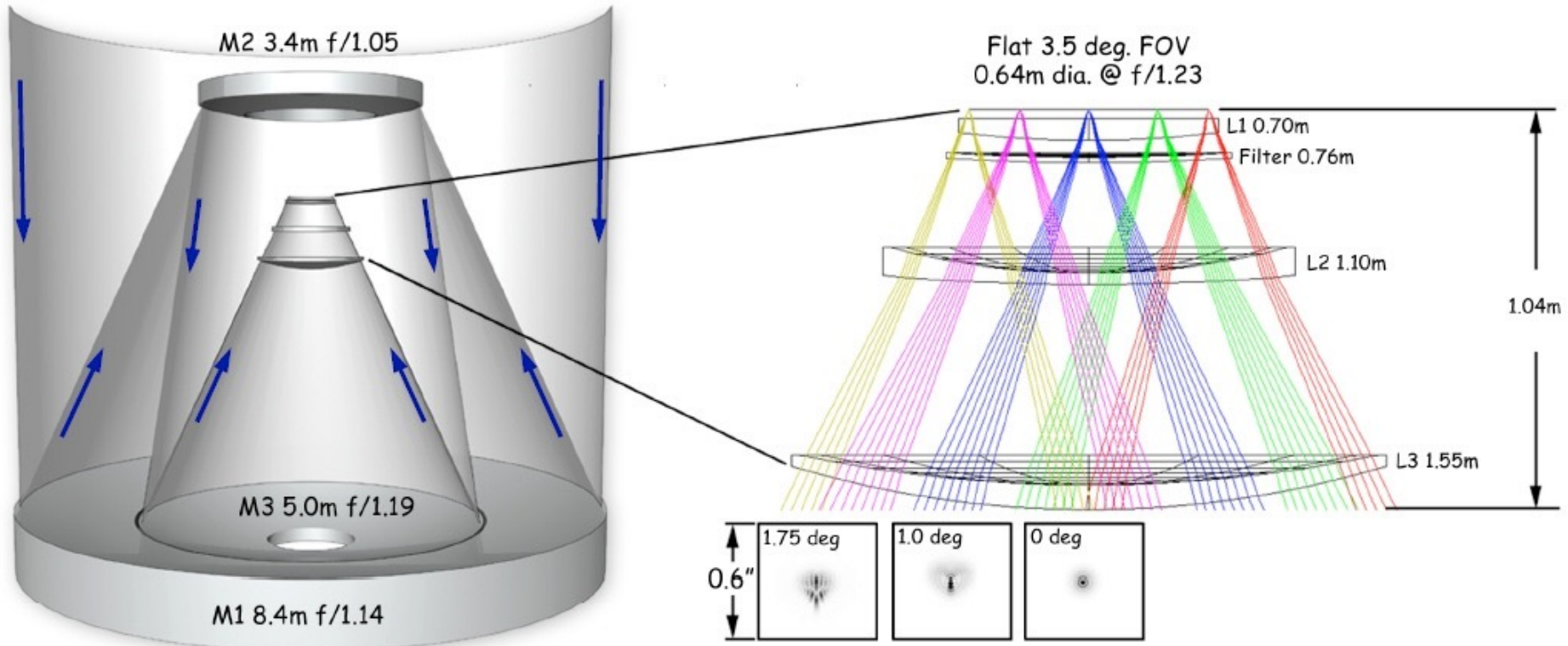
Gemini South Telescope



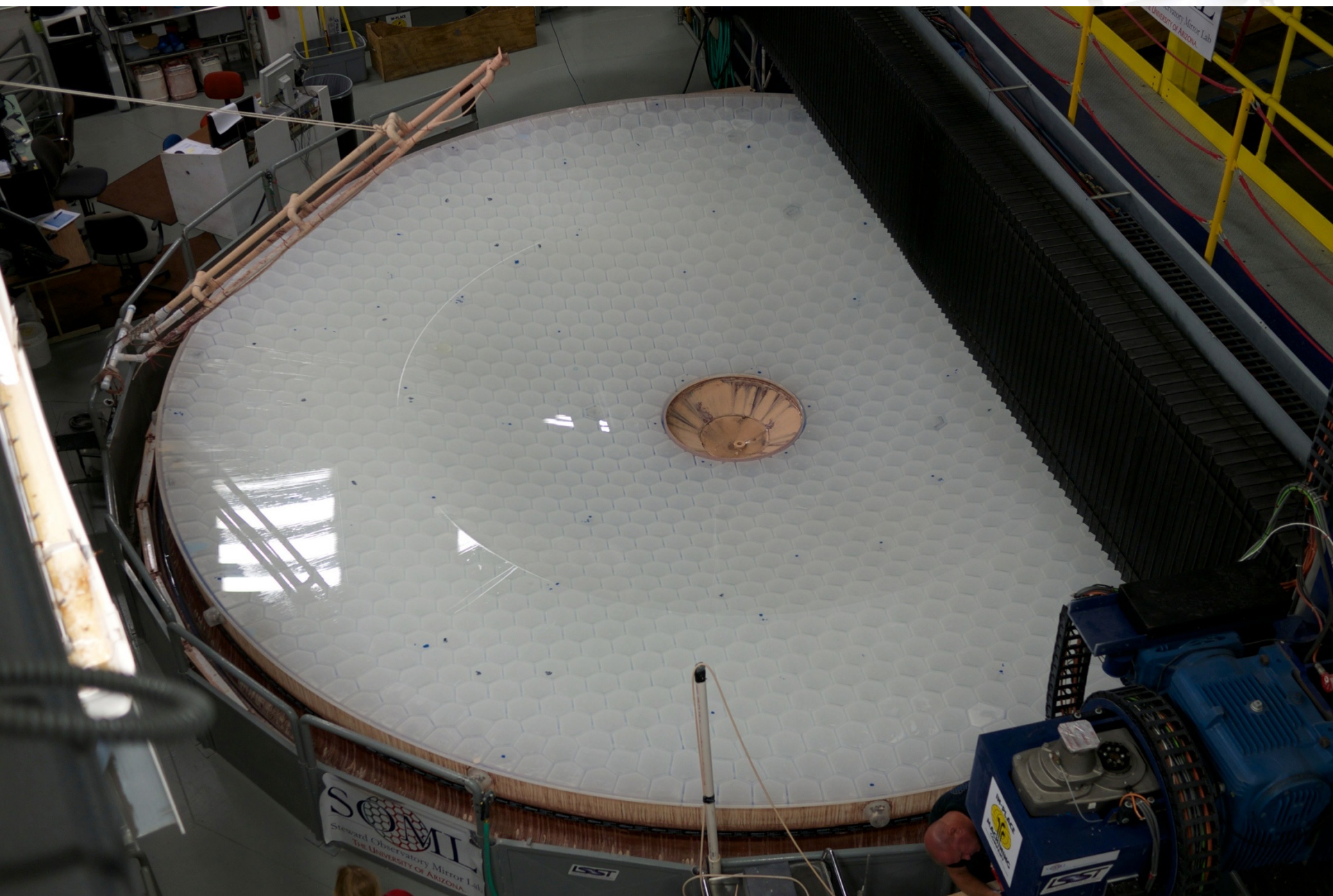
LSST



Optical Design for LSST

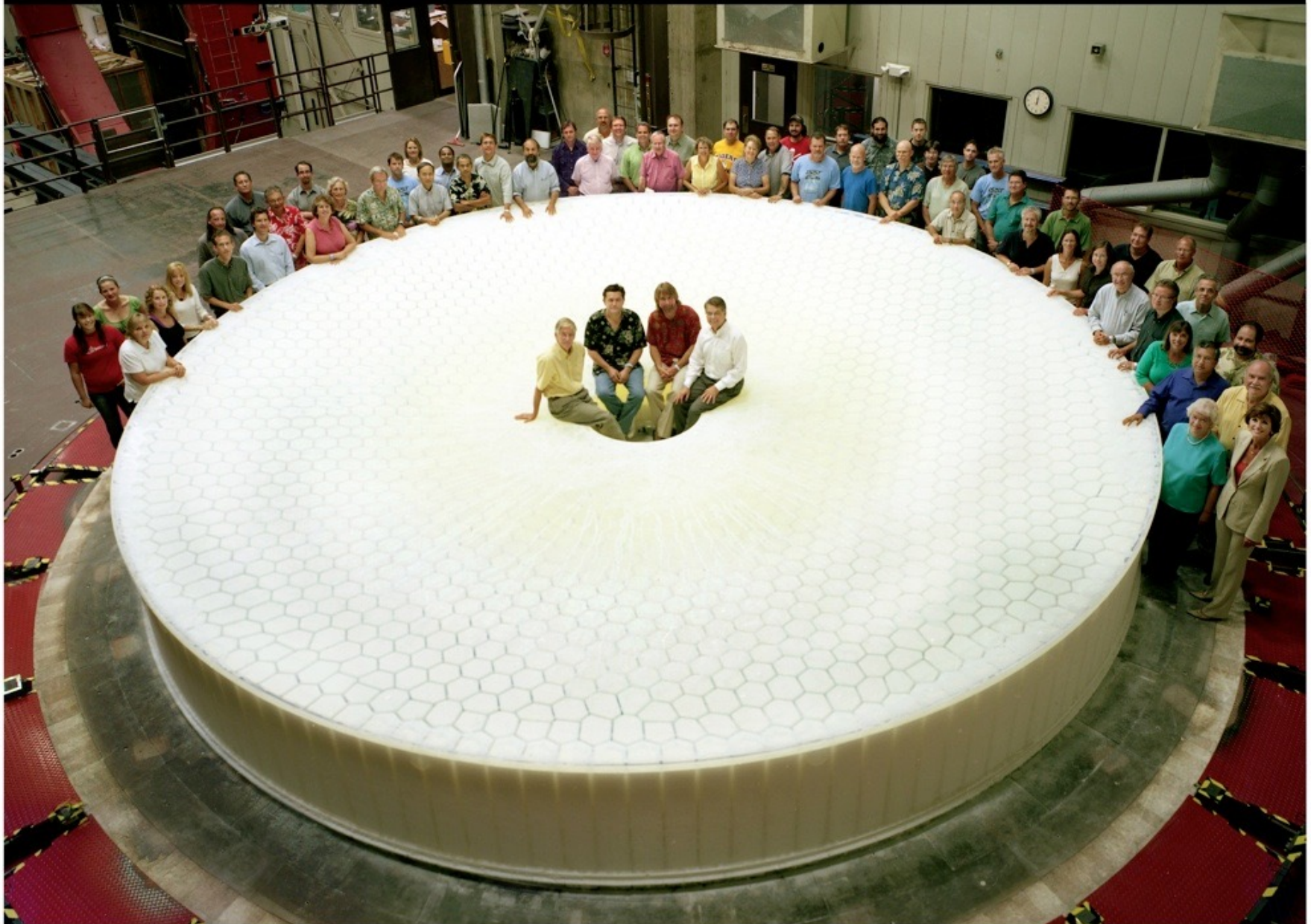


Three-mirror design (Paul-Baker system)
enables large field of view with excellent image quality:
delivered image quality is dominated by atmospheric seeing



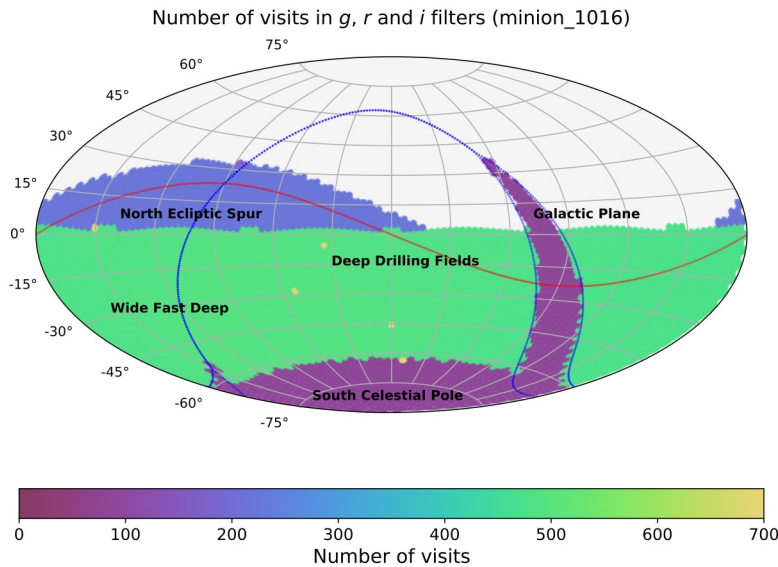


Large Synoptic Survey Telescope



Basic idea behind LSST: a uniform sky survey

- **90% of time will be spent on a uniform survey:** every 3-4 nights, the whole observable sky will be scanned twice per night
- **after 10 years, half of the sky will be imaged about 1000 times (in 6 bandpasses, ugrizy):** a digital color movie of the sky
- **~100 PetaBytes of data:** about a billion 16 Mpix images, enabling **measurements for 40 billion objects**



LSST in one sentence:

An optical/near-IR survey of half the sky in ugrizy bands to $r \sim 27.5$ (36 nJy , $3.6 \times 10^{-31} \text{ erg/s/cm}^2/\text{Hz}$) based on 825 visits over a 10-year period: **deep wide fast.**

Left: a 10-year simulation of LSST survey: the number of visits in the *r* band (Aitoff projection of eq. coordinates)

SDSS

gri

3.5'x3.5'

r~22.5

3 arcmin
is 1/10
of the full
Moon's
diameter



HSC

gri

3.5'x3.5'

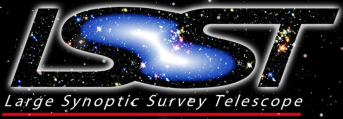
r~27

3 arcmin
is 1/10
of the full
Moon's
diameter

like LSST
depth (but
tiny area)

LSST will
deliver 5
million
such
images





LSST Operations: Sites & Data Flows



and a British site, too!

French Site

Satellite Processing Center
Data Release Production
Long-term Storage (copy 3)

Archive Site

Archive Center
Alert Production
Data Release Production
Calibration Products Production
EPO Infrastructure
Long-term Storage (copy 2)
Data Access Center
Data Access and User Services

HQ Site

Science Operations
Observatory Management
Education & Public Outreach

Base Site

Base Center
Long-term storage (copy 1)
Data Access Center
Data Access & User Services

Summit Site

Telescope & Camera
Data Acquisition
Crosstalk Correction

Google

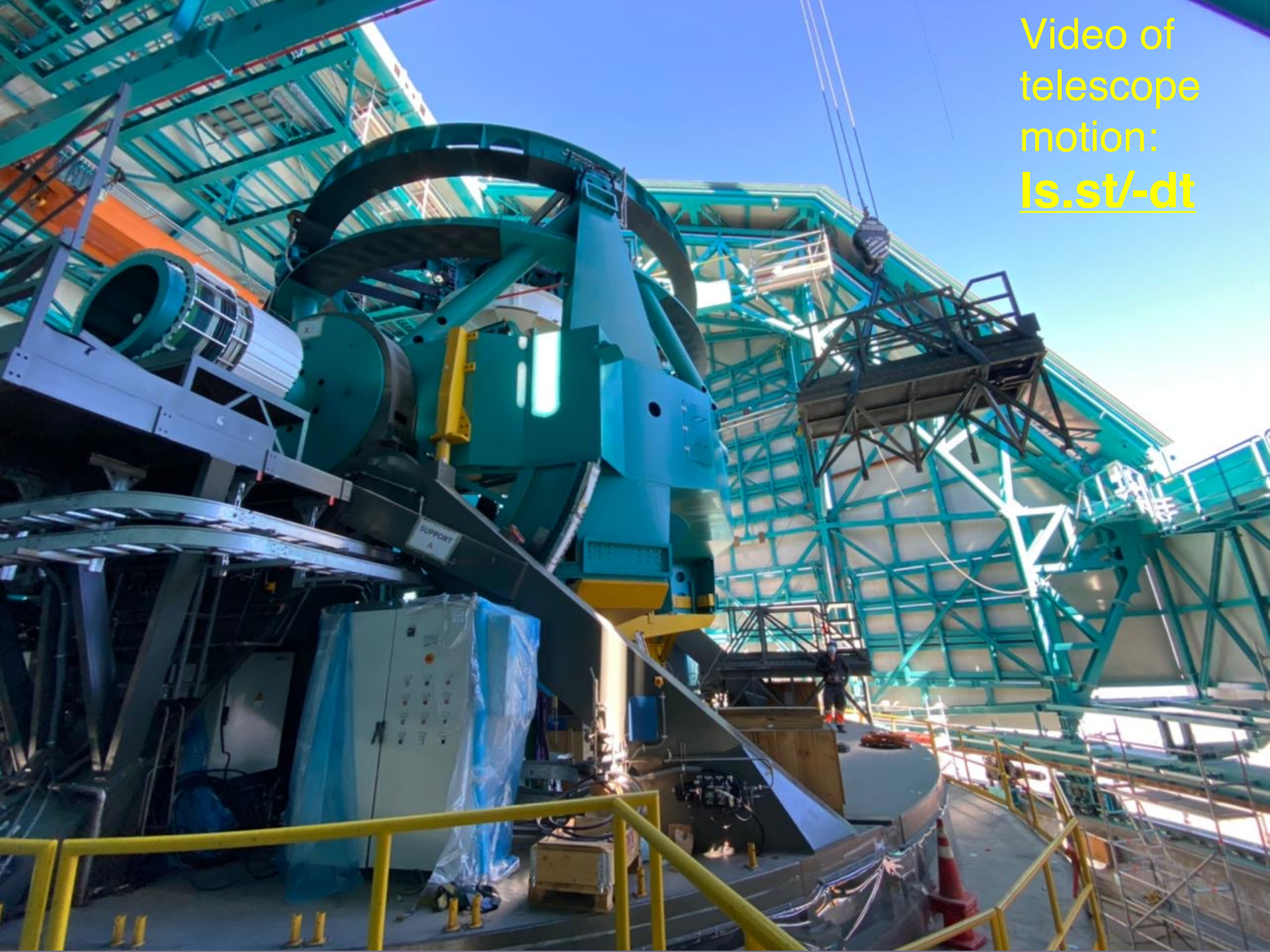
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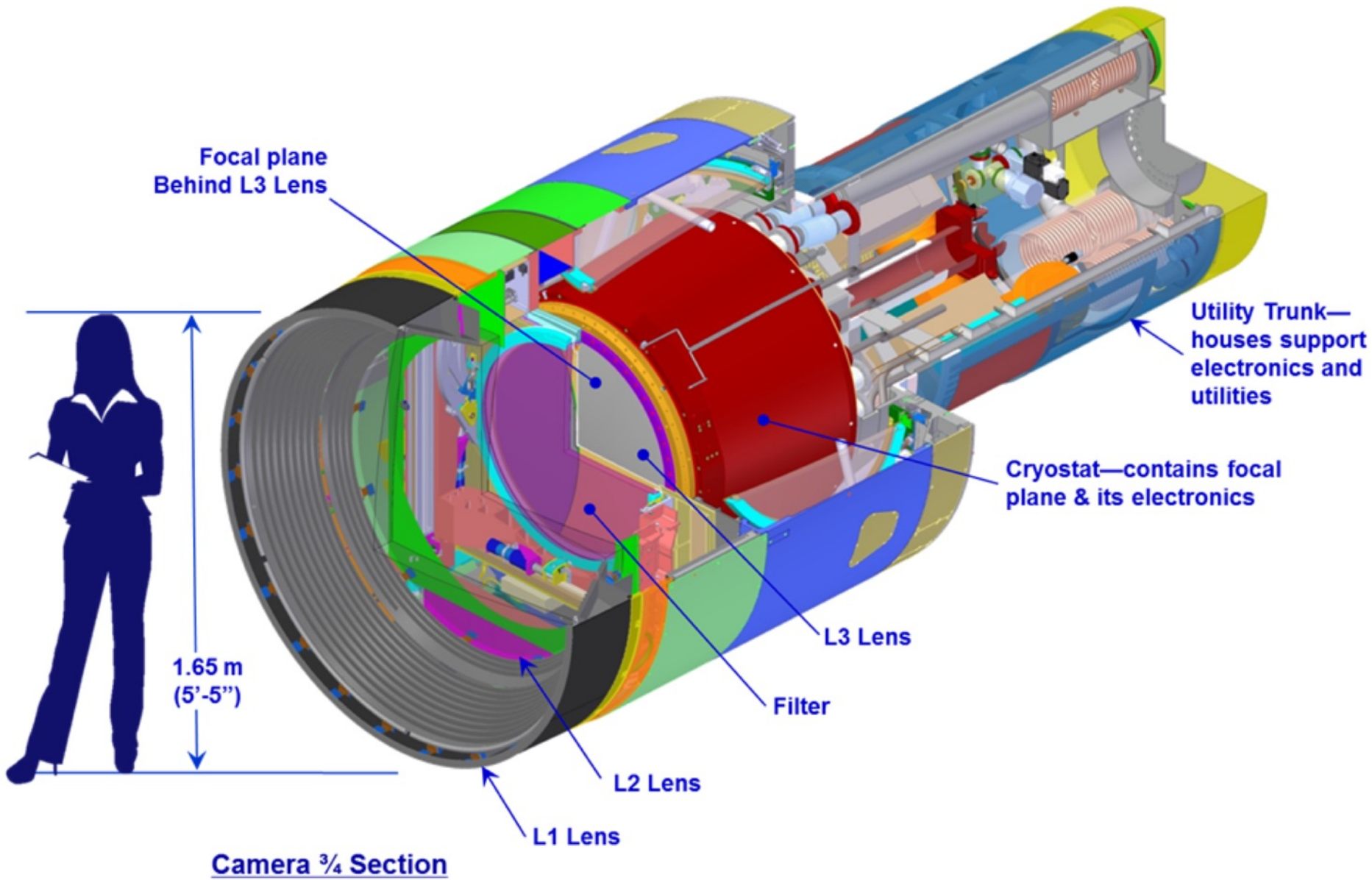
The rise of Rubin Observatory



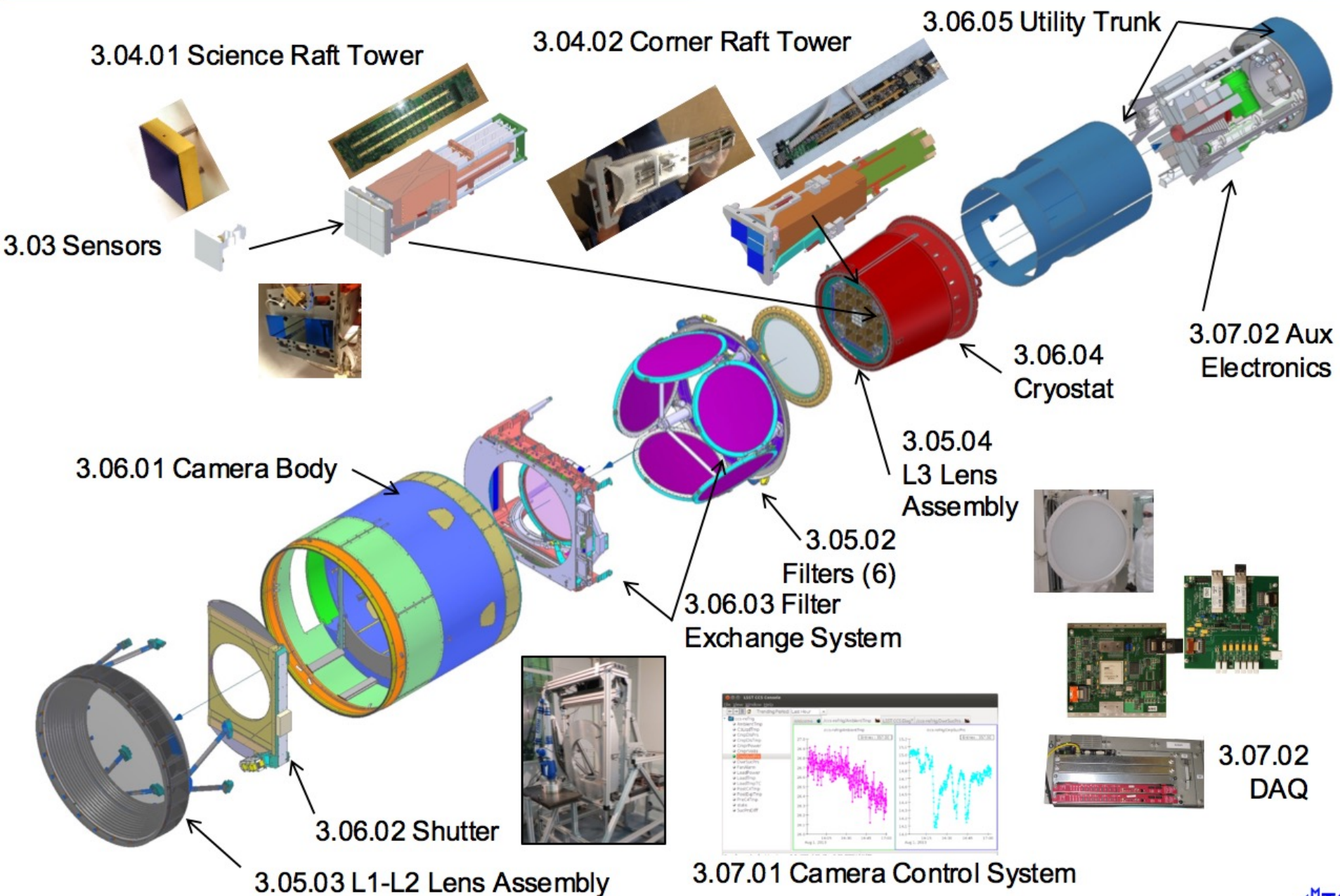


Video of
telescope
motion:
[ls.st/-dt](#)

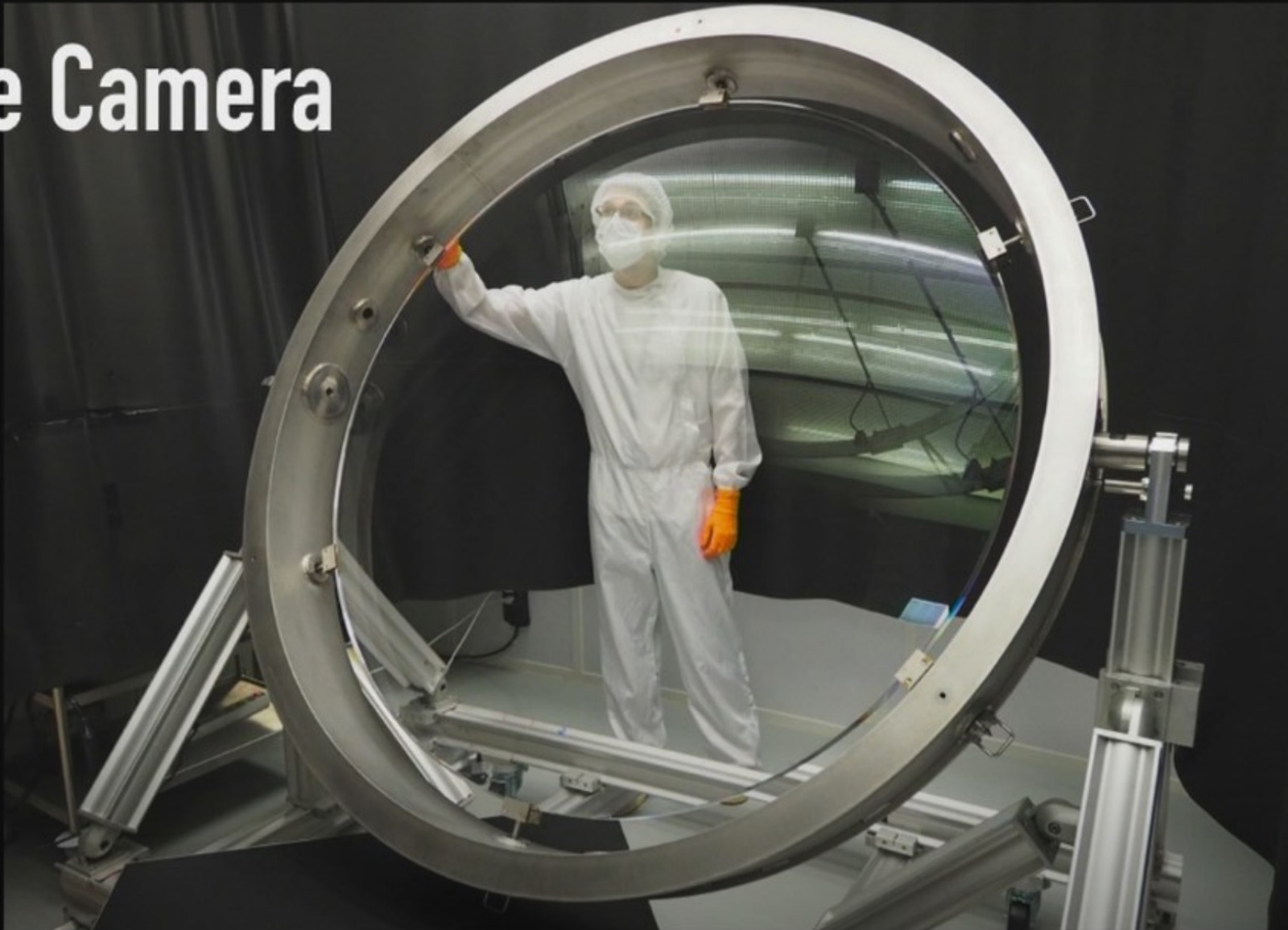




Major Camera Elements

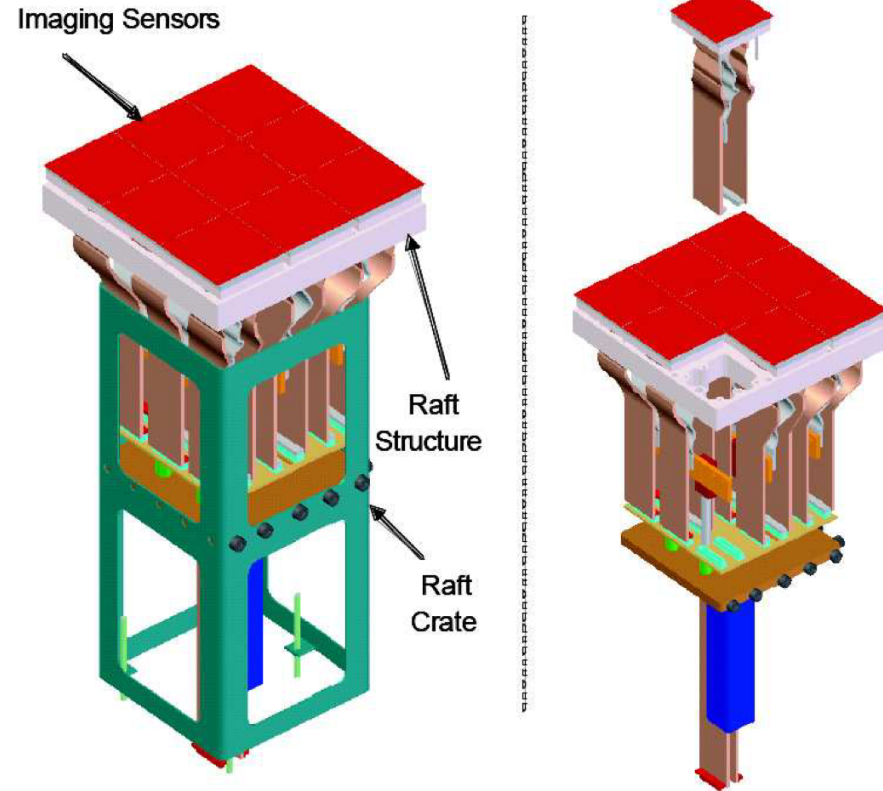
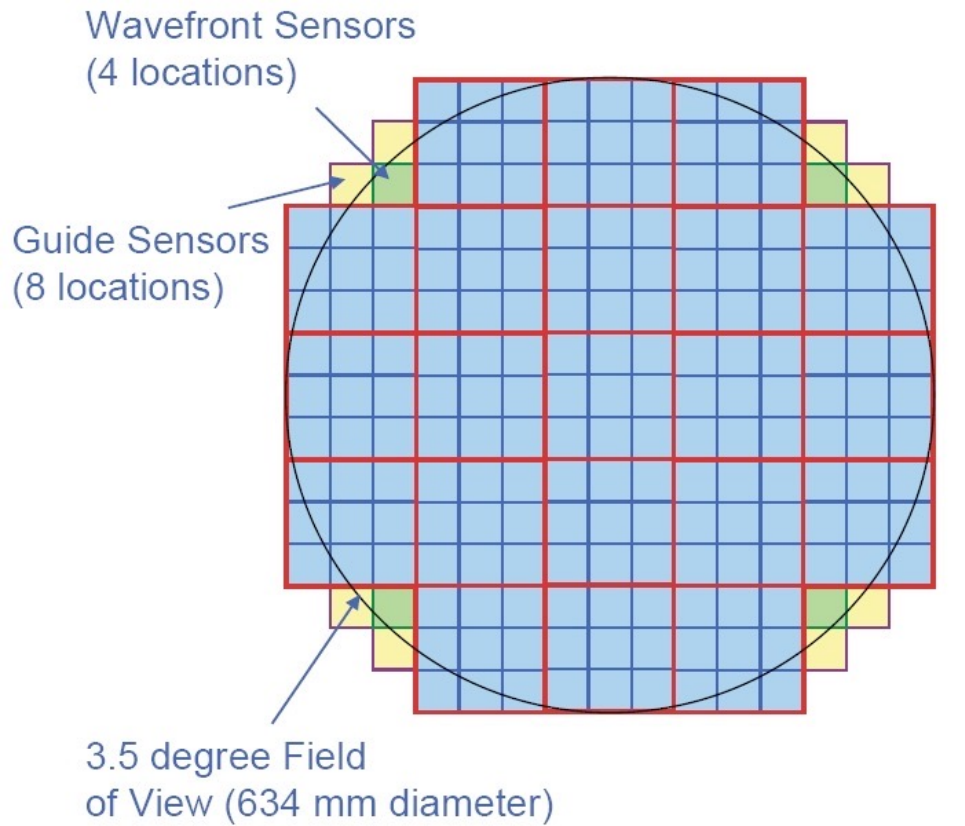


Large Camera

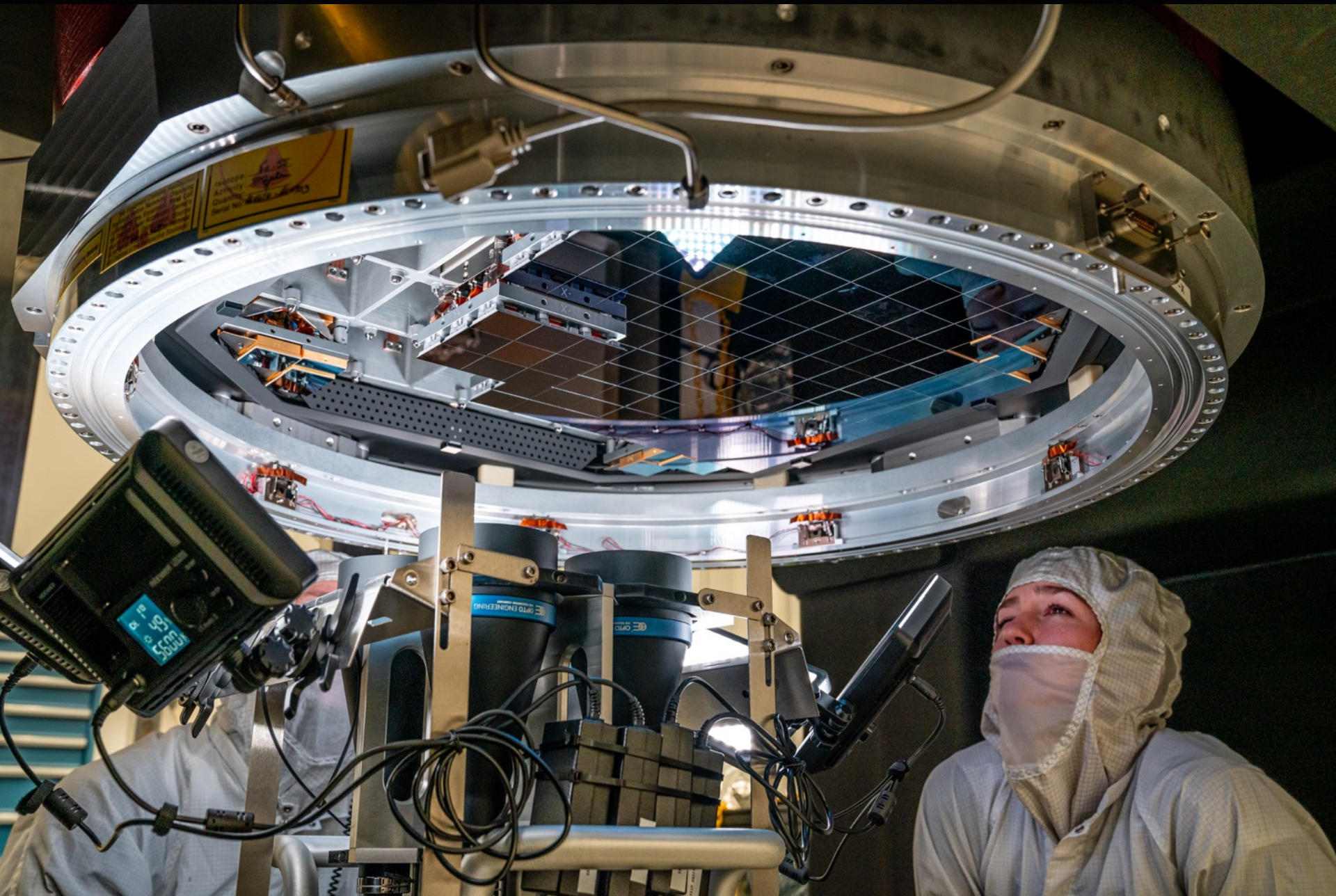


L-1, the largest lens ever produced, is the front lens of the LSST camera $D = 1.57\text{m}$

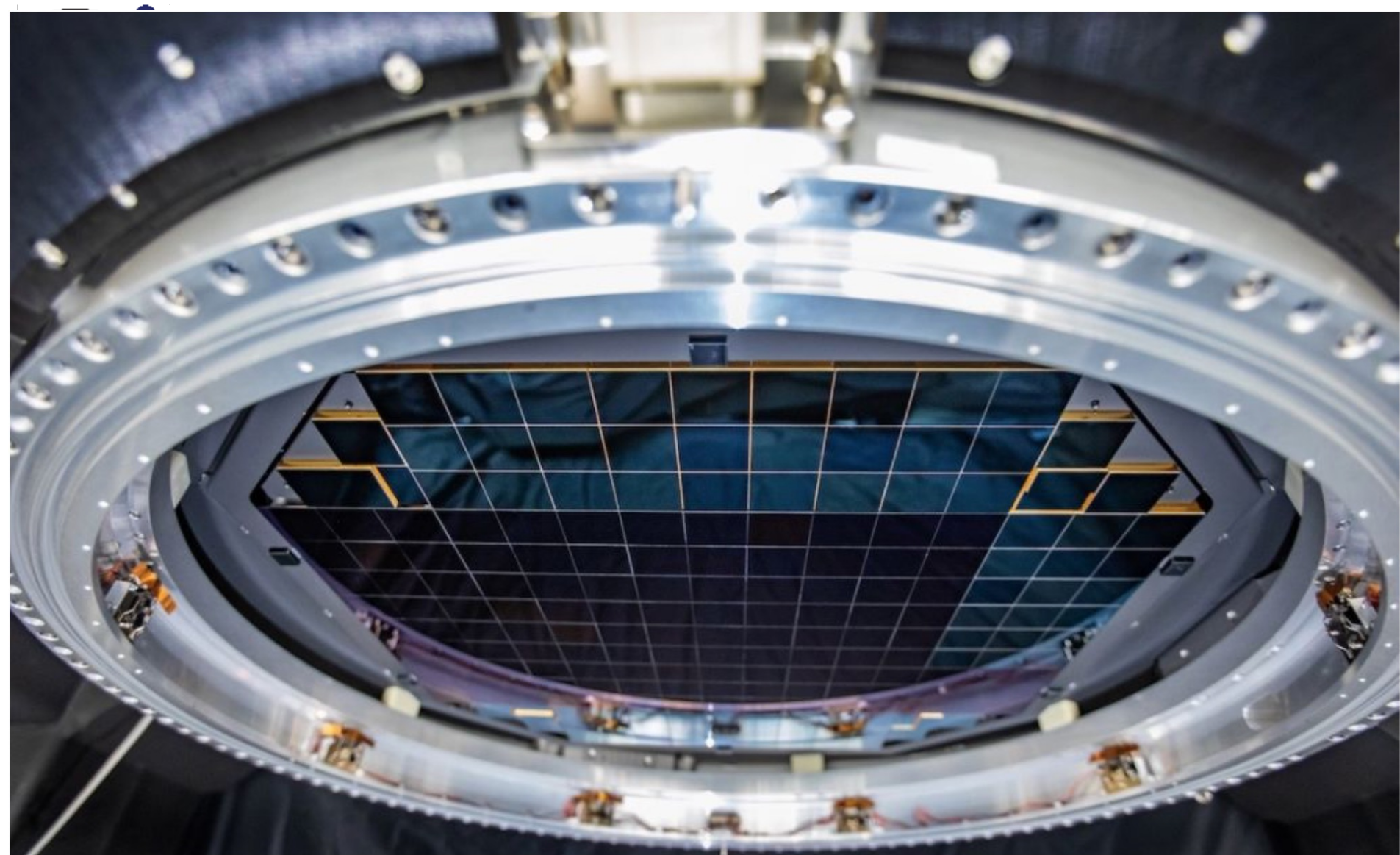
LSST camera



Modular design: 3200 Megapix = 189 x 16 Megapix CCD
9 CCDs share electronics: raft (=camera)
Problematic rafts can be replaced relatively easily



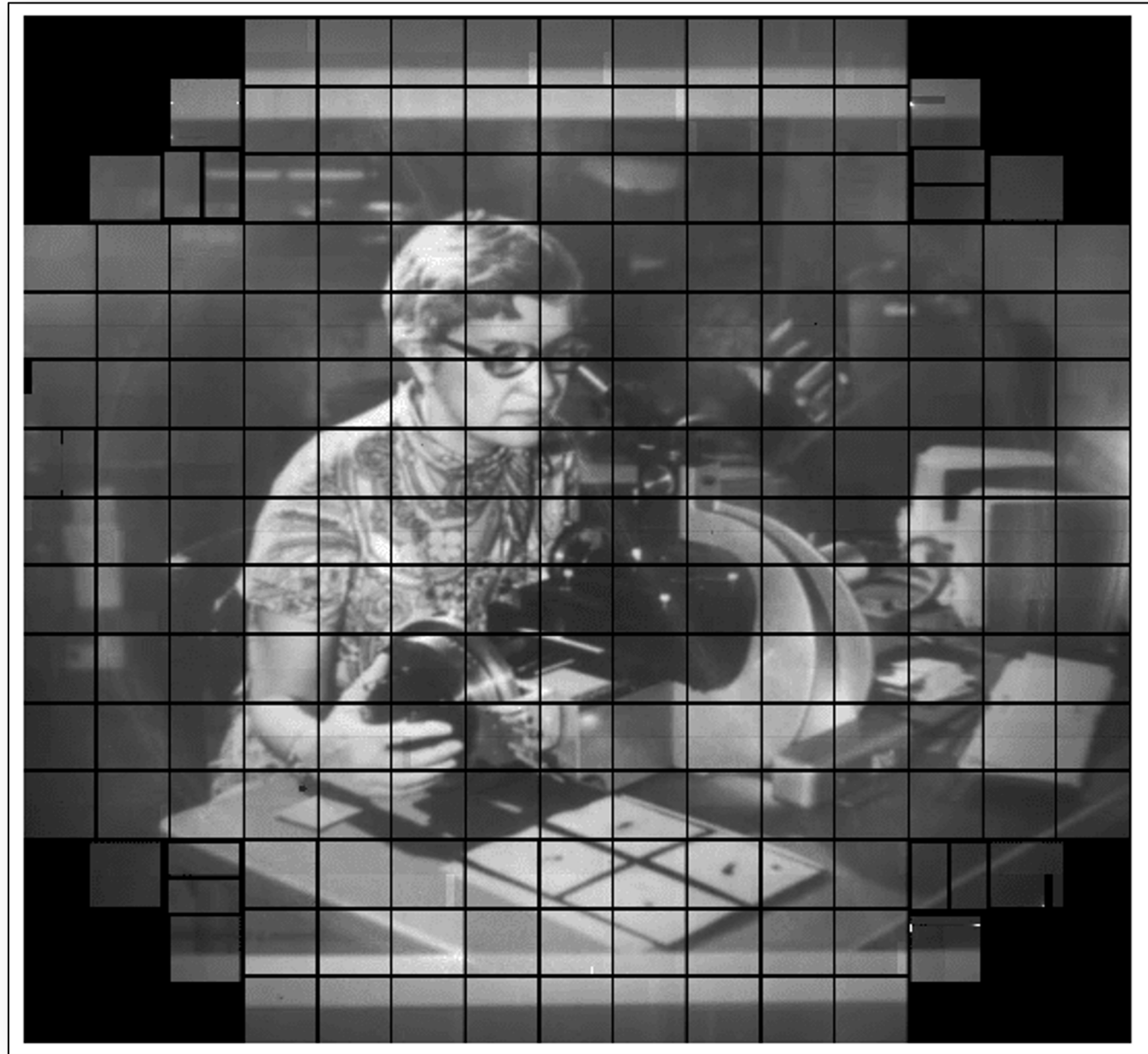
The largest astronomical camera in the world: 3,200 Megapix



The complete focal plane of the future LSST Camera is more than 2 feet wide and contains 189 individual sensors that will produce 3,200-megapixel images.

It works!

(well, 3
amplifiers, out
of 3024, don't...)





At the highest level, LSST objectives are:

- 1) Obtain about 5.5 million images, with 189 CCDs (4k x 4k) in the focal plane; this is about **a billion 16 Megapixel images of the sky**
- 2) Calibrate these images (and provide other metadata)
- 3) Produce catalogs (“model parameters”) of detected objects (37 billion)
- 4) **Serve** images, catalogs and all other metadata, that is, **LSST data products to LSST users**

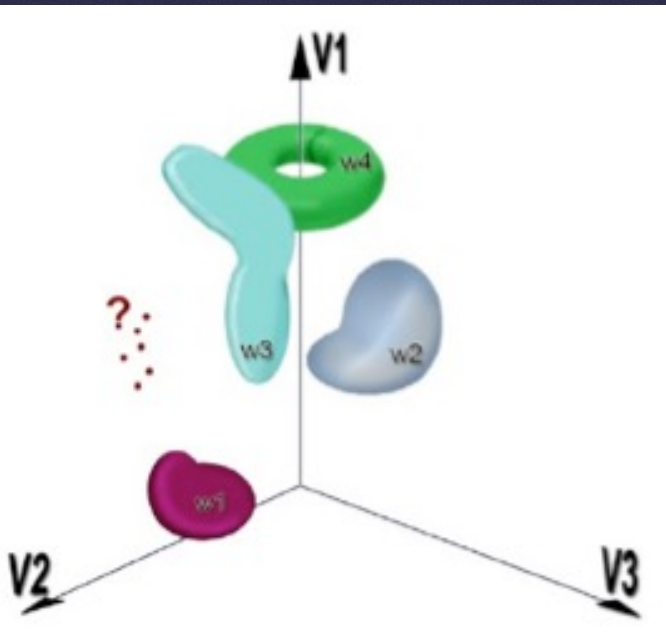
The ultimate deliverable of LSST is not just the telescope, nor the camera, but the fully reduced science-ready data as well. Software!

About a million lines of in-house C++ and python.

Statistical analysis of a massive LSST dataset

- A large (100 PB) database and sophisticated analysis tools: for each of 40 billion objects there will be about 1000 measurements (each with a few dozen measured parameters)

Data mining and knowledge discovery



- 10,000-D space with 40 billion points
 - Characterization of known objects
 - Classification of new populations
 - Discoveries of unusual objects
- Clustering, classification, outliers

Schedule to Rubin Obs. Completion

- Despite the progress during the COVID Pandemic the critical path suffered dramatically
- Schedule continues to adjust to conditions
 - **ComCam: Engineering First Light: July 2023**
 - **LSSTCam: System First Light: March 2024**
 - **Start of Operations and LSST: Fall 2024**

Still months of uncertainty in first light dates, but we will get there!

Legacy Survey of Space and Time: a 10-year survey starting in 2024

More details:
ls.st/lop

multi-color time-resolved faint sky map

- 20 billion galaxies
- 20 billion stars
- 10 billion alerts
- “millions and millions” of SNe, quasars, asteroids...

dark energy vs. modified gravity...

dark matter constraints...

