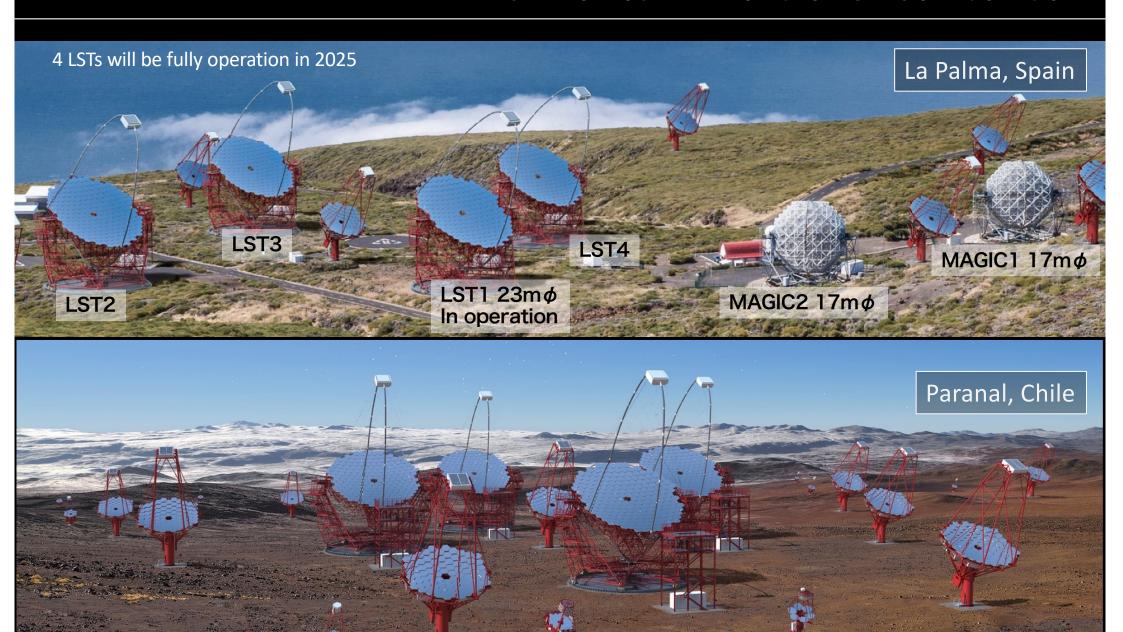
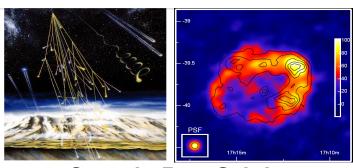
CTA - LST Project

Masahiro Teshima for the LST Collaboration



Science of CTA is very wide

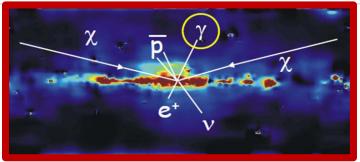
SNRs, PWNe, AGNs, GRBs, Dark Matter



Cosmic Ray Origin

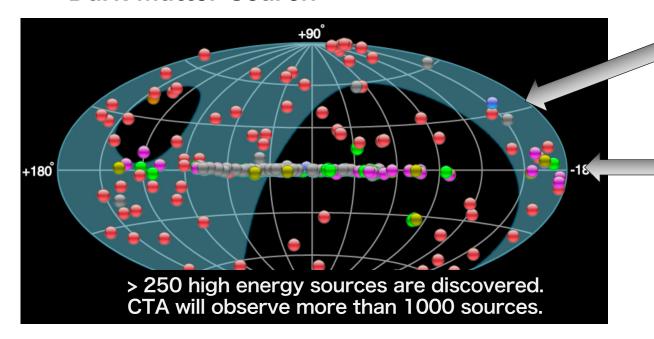


Super Massive Black Holes



Dark Matter Search (Discovery)

- Origin of Cosmic Rays (Big accelerators)
- Black Hole and S.M.B.H.
- Dark Matter Search



Extragalactic Sources



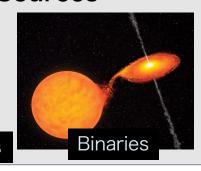
Active Galactic Nuclei



Galactic Sources



Super Nova Remnants

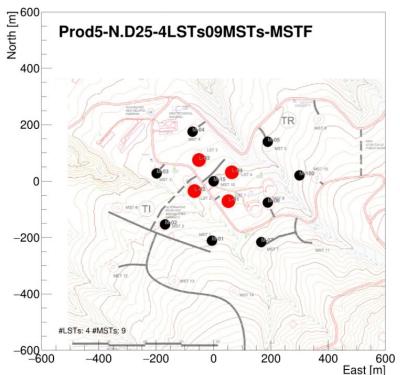




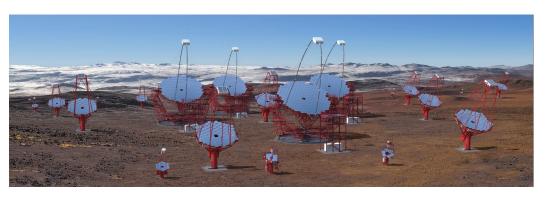
Alpha Configuration is decided with the financial constraints

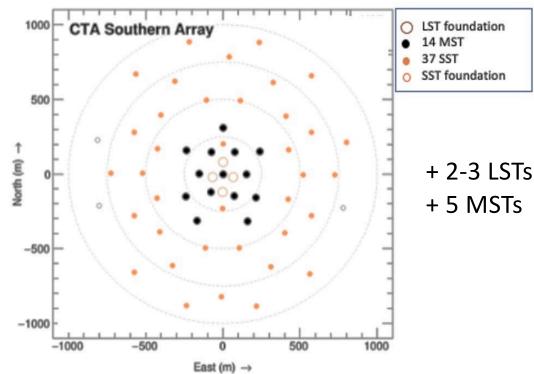
Roque de los Muchachos Observatory La Palma, Spain





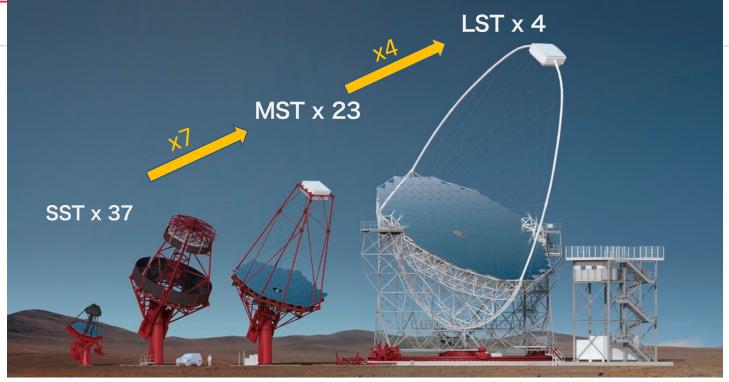
Paranal, Chile







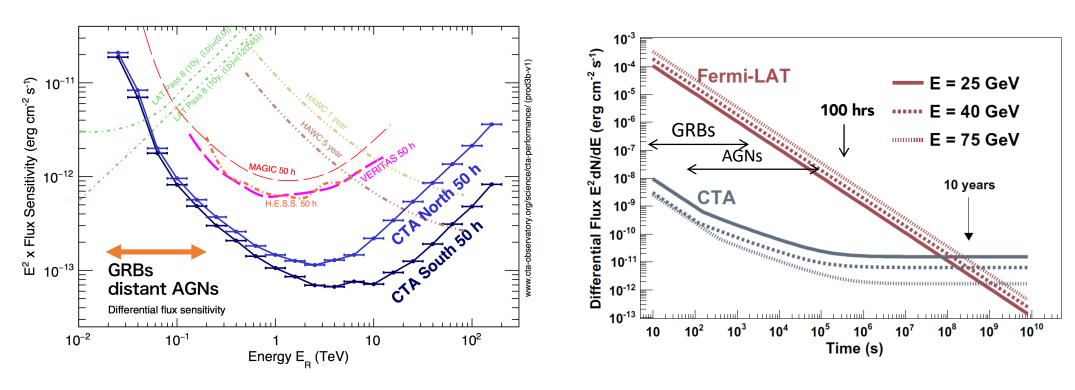
telescope array Telescope Design



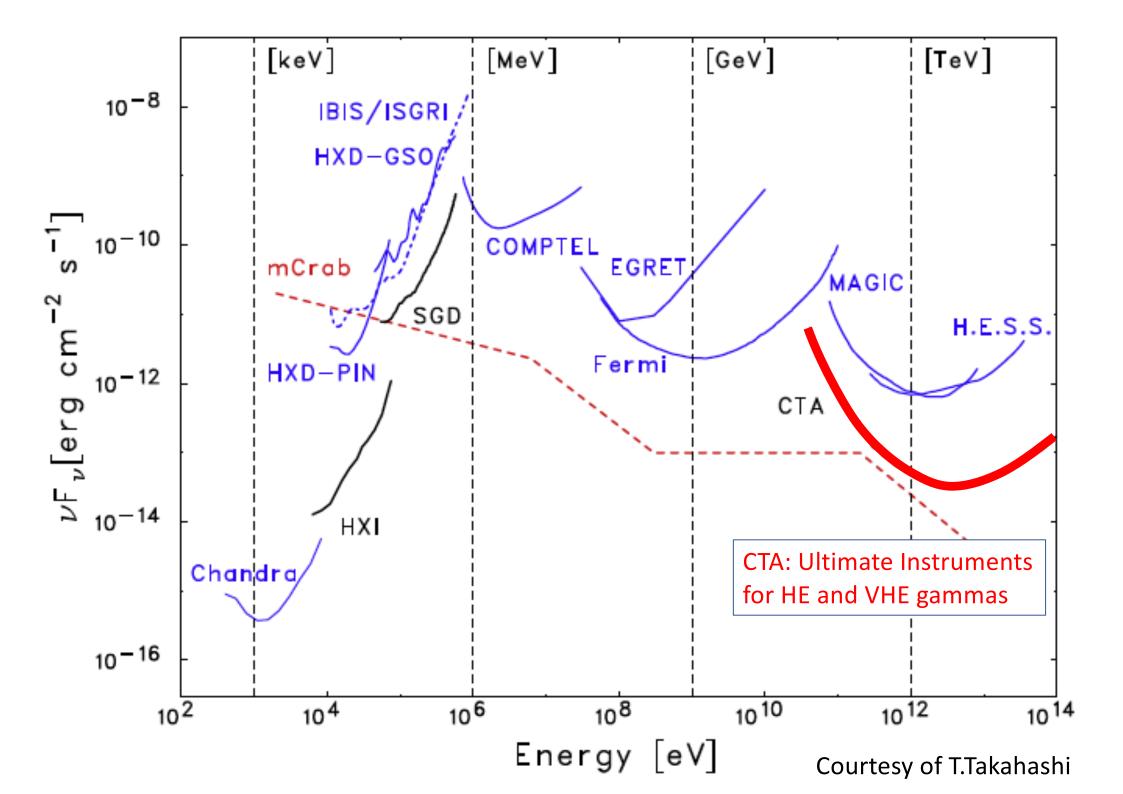
Telescope Types	SST	MST	LST	
Optics	Schwarzschild-Couder	Davies-Cotton	Parabolic (Isochronous)	
FoV and Camera	10.5 deg SiPM	7.5 deg PMT	4.3 deg PMT	
Mirror Diameter	4.3m	11.5m	23m	
Energy Range	3 TeV - 200 TeV	100GeV - 10TeV	20GeV – 2000GeV	
Science Targets	Galactic Sources PeVatron (UHE CR)	Galactic Sources Nearby AGNs (z<0.5) Dark Matter	Transient Sources AGNs(z<2), GRBs(z <4) Dark Matter	



10 times better sensitivity Wide Energy coverage 20GeV~200TeV

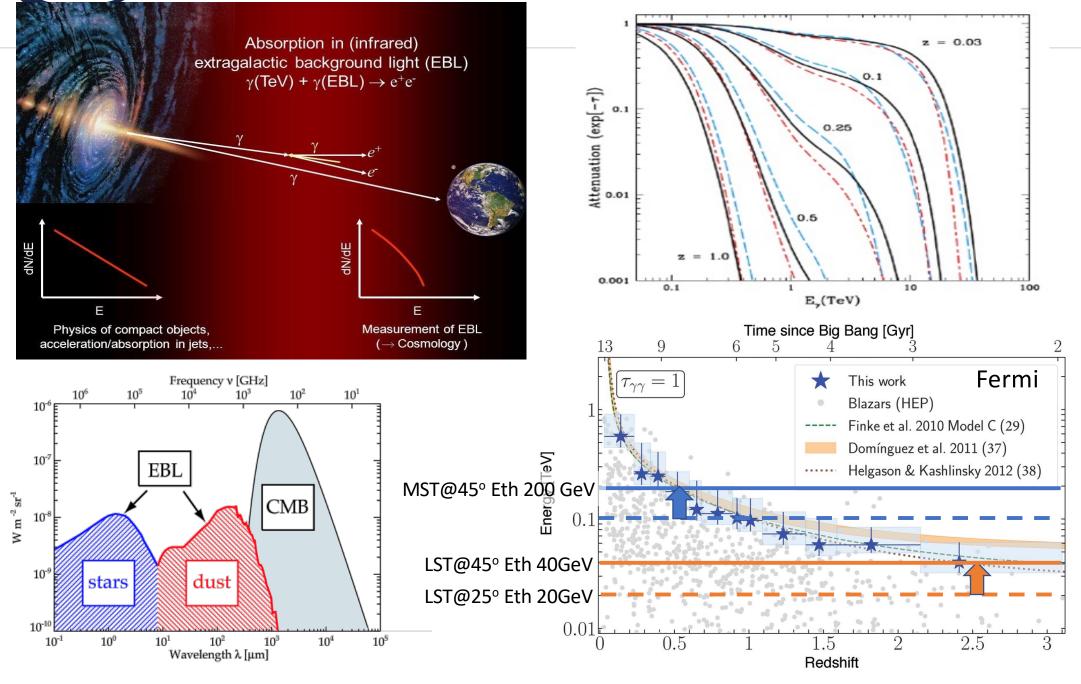


- CTA array has a 10 times better sensitivity than HESS, MAGIC, and VERITAS
- CTA covers wide energy range from 20GeV to 200TeV (4 orders of magnitude)
- LSTs will offer
 - Distant AGNs up to z = 2 and GRBs up to z = 4 are observable with LSTs
 - X10000 sensitivity for GRBs and AGN flares than Fermi
 - The fast rotation (20 sec) offers the observation of GRBs even in prompt phase



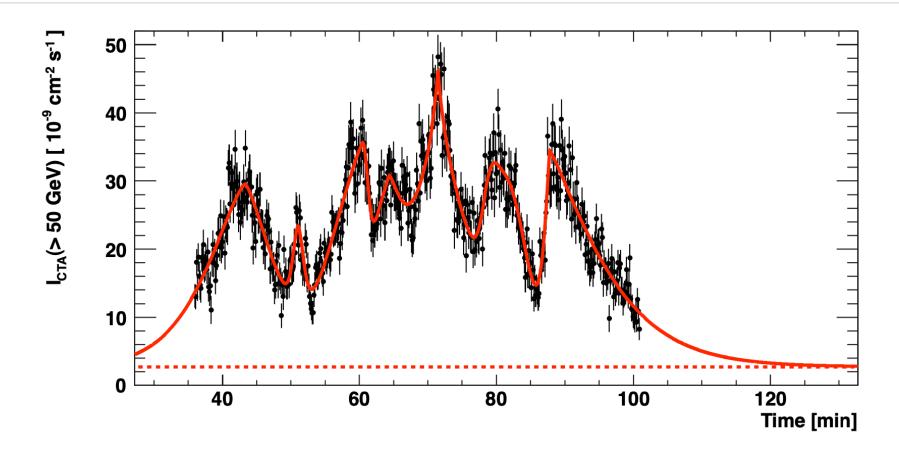


Gamma Ray Horizon 20GeV Low Threshold Energy → z~4





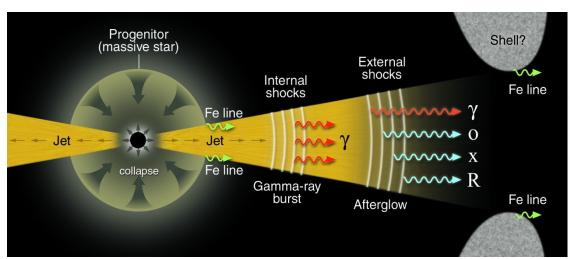
Simulated AGN Flares Template: the 2006 flare of PKS2155-304 Low Threshold Energy → High Precision Light curve

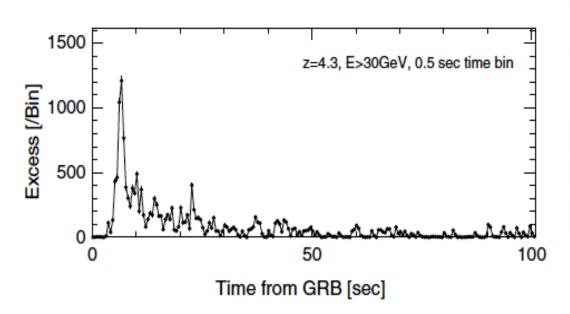


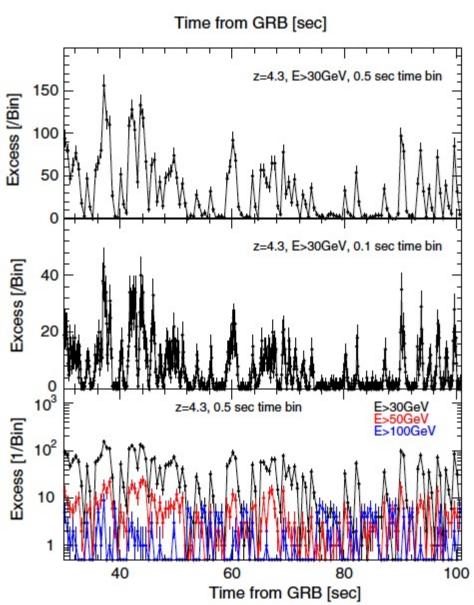
- ☐ Light curve can be examined, a few minutes scale structure → a few 10s of seconds
 - ☐ Particle acceleration mechanism, Cooling process
 - ☐ Light curve vs. Energy dependence → Q.G. Energy scale > Planck Mass scale



GRB: Simulated light curve Newly Born B.H.s or N-N mergers



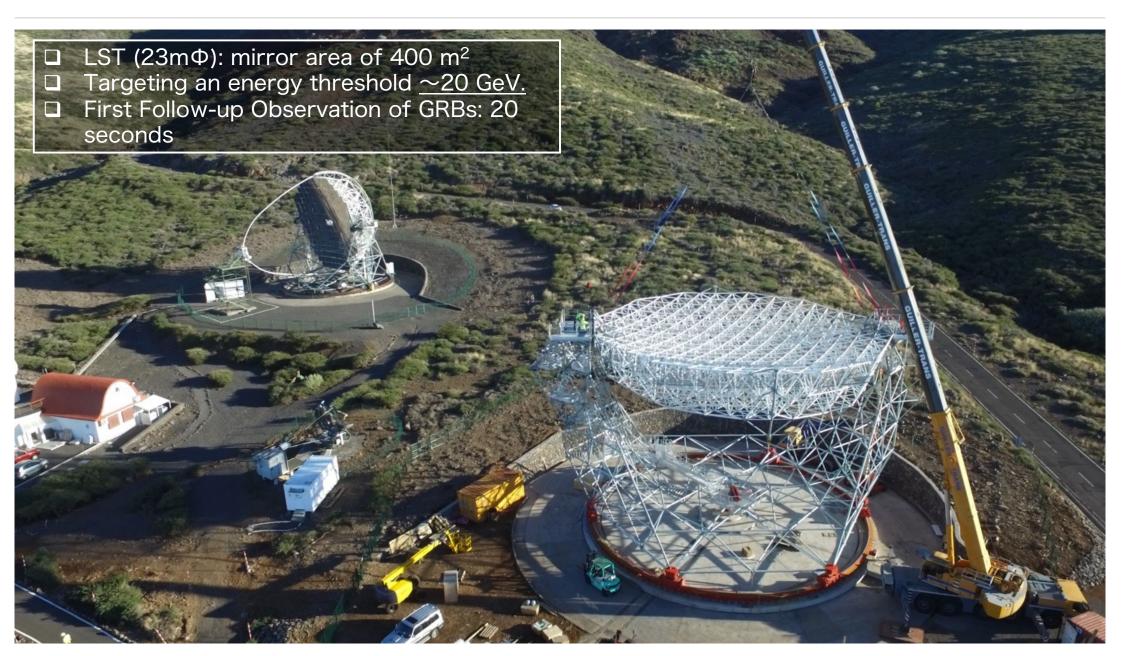






The CTA Large Size Telescope LST1

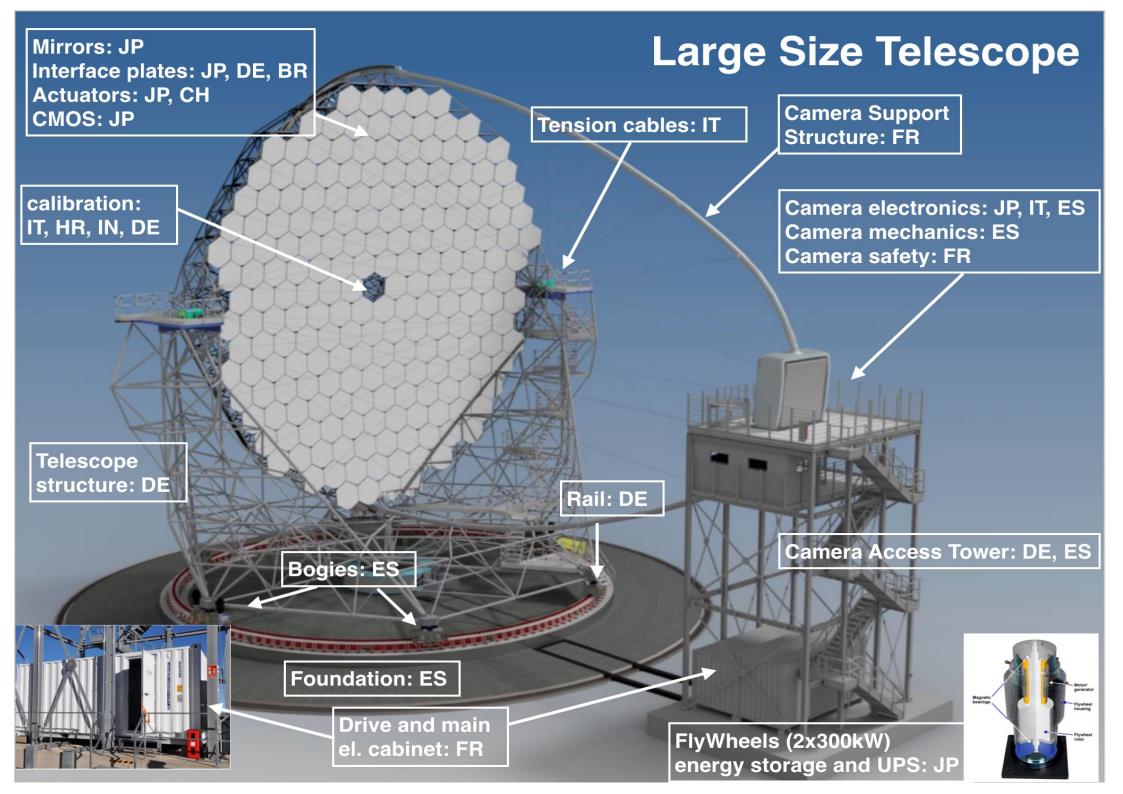
(Photo under construction in December 2017)





Commissioning since 2019 including Science Operation since 2020

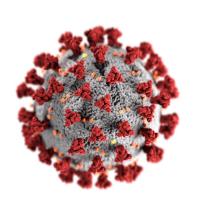


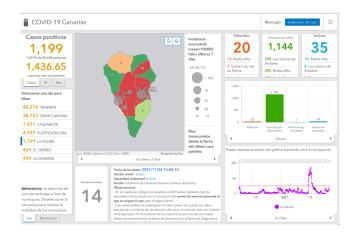




Last three years, we suffered from "" Oh, my God!!

Covid-19



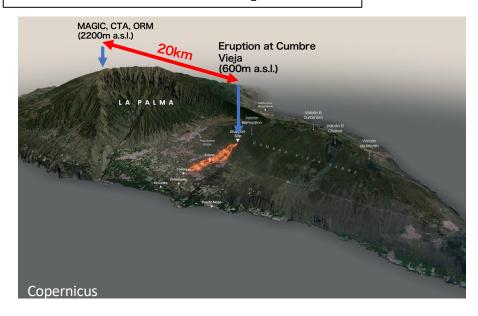


Russian Aggression in Ukraine

Inflation / Cost Increase

Shortage of semiconductors and materials

Volcano Eruption

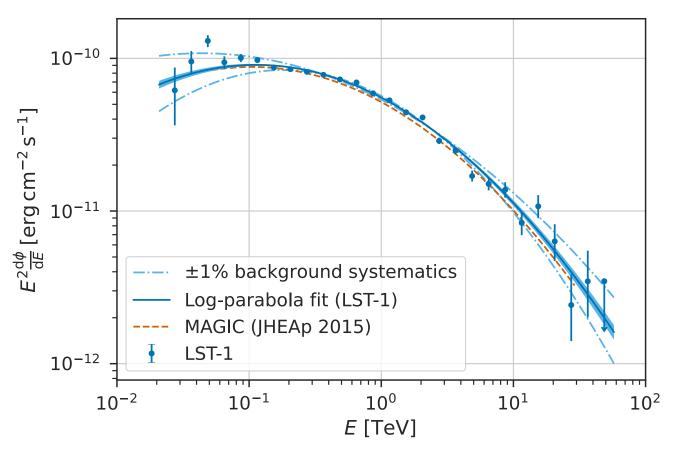




Erupciones históricas en La Palma						
#	Erupción	Año	Días erupción			
1	nombre?	2021	85days ?			
2	Teneguía	1971	24			
3	San Juan	1949	47			
4	Charco	1712	56			
5	San Antonio	1667/1678	66			
6	Tigalate	1646	82			
7	Tehuya	1585	84			
8	Tacande	1430/1440	?			



Performance: Crab Nebula spectrum

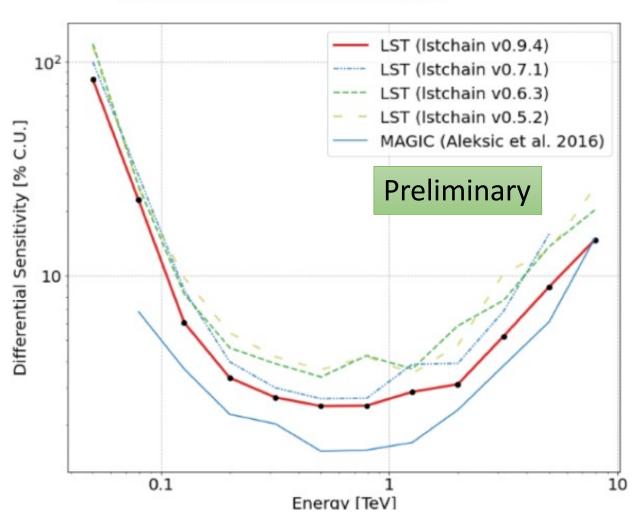


- 34 h effective time, γ -ray efficiency: 70% from gammaness cut and 70% from θ^2 cut
- Error bars are only statistical.
- Systematics: blue lines correspond to effect of ±1% background.
- Consistent with MAGIC and Fermi-LAT.
- Lowest data point at 25 GeV!



Single Telescope LST-1 performance: sensitivity

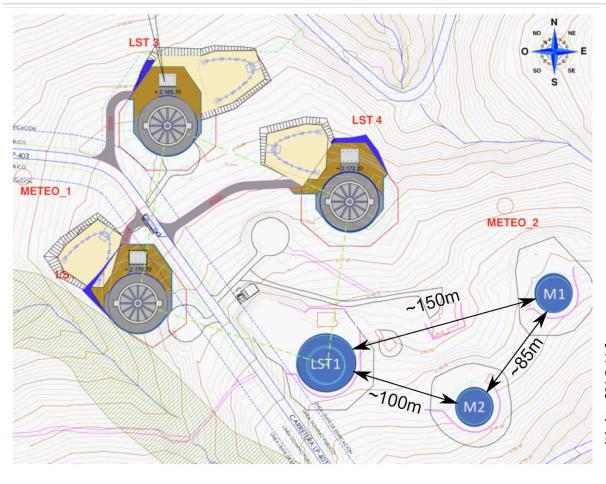


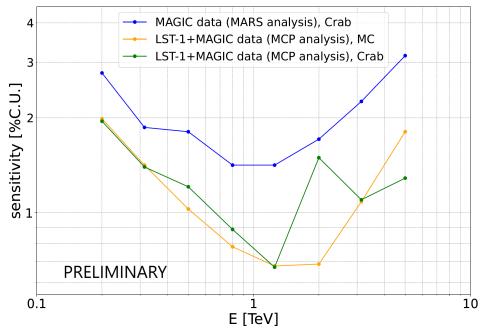


- ☐ Consistent sensitivity for source-dependent and source-independent analyses.
- ☐ The sensitivity is close to MAGIC stereo array.
- ☐ X10 better sensitivity is expected with 4 LST array



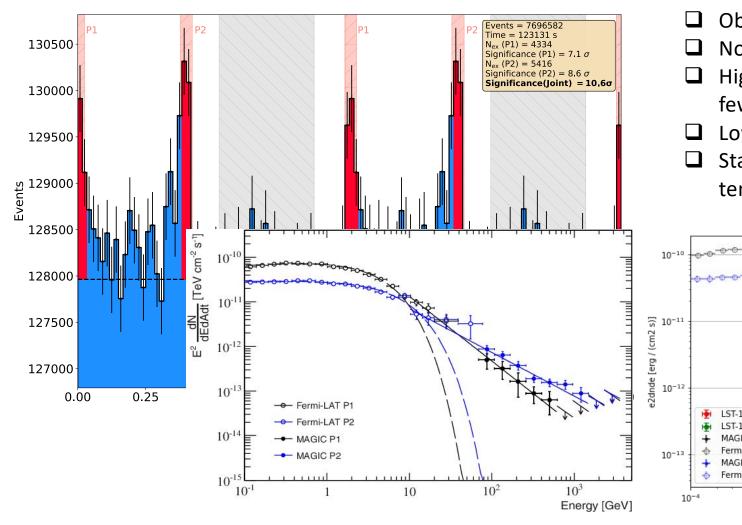
LST1 + MAGIC joint data analysis Stereo Observations improve the sensitivity



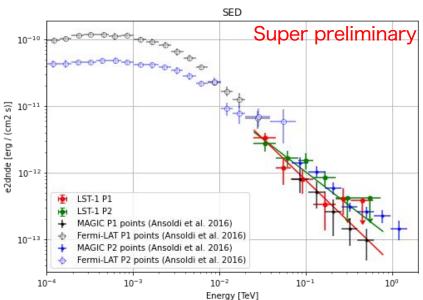




Crab pulsar phaseogram



- Observation time: 34.2 hours
- Nov 2020 March 2022
- Highly significant detection down to few tens of GeV.
- ☐ Low energies: P1/P2 tends to 1.
- ☐ Stay tuned for spectrum down to few tens of GeV...



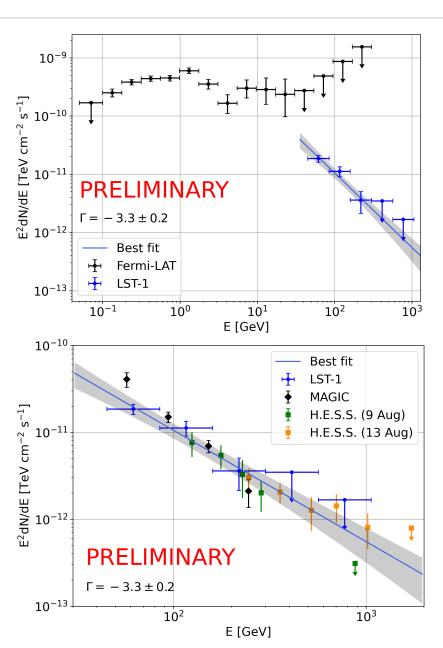


First VHE-detected Recurrent nova: RS Ophiuci

- RS Ophiuchi is a recurrent Nova.
- Explosions, 1898, 1933, 1958, 1985, 2006, 2021
- Mag 12.5 (low state) → Mag 4.7 (~1000 times)
- ☐ Binary System with a White Dwarf and a Red Giant
- Accumulation of material on the WD, and then thermonuclear reaction makes recurrent explosions

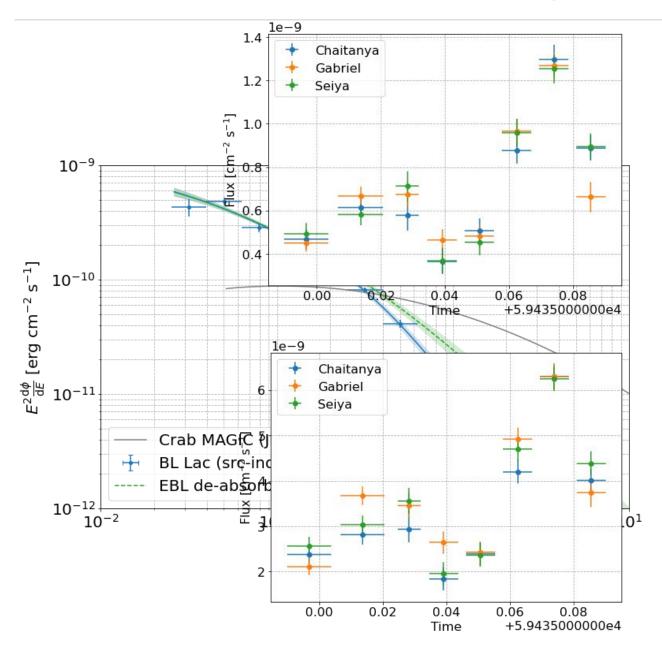


Credit: David A.Hardy/ www.astroart.org & PPARC.

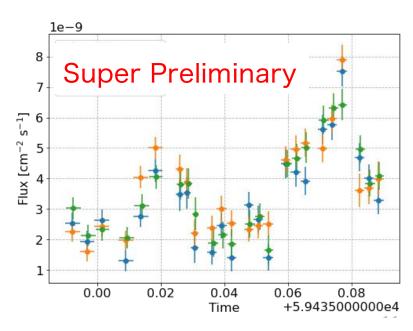




BL Lacertae flare on 8th August 2021



- ☐ IBL at z=0.069
- ☐ In a high emission state since 2020
- □ August 8th 2021: High state >1 crab for E<300 GeV.
- ☐ Soft spectrum allows to extract spectral point at 30 GeV in <2 hour observation.





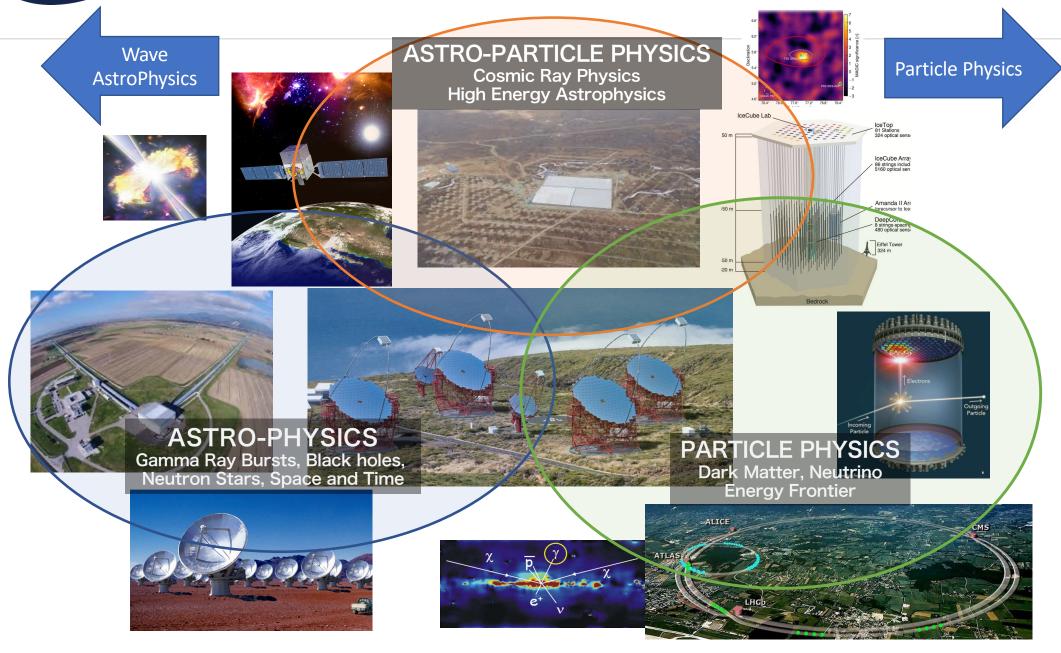
CTA and LST Timeline

- □ 2016 2018 LST1 in construction
- □ 2019 LST1 in commissioning phase
- □ 2022 2024 LST2-4 will be constructed
- □ 2025 LST1-4 in commissioning
- □ 2026 2027 The final Acceptance of LST1-LST4 and IKC process
- □ 2023 2027 LST5-8 construction

	2020	2021	2022	2023	2024	2025	2026	2027	2028	2029	2030	
Organization	CTAO gGmbH (Heidelberg)											
Organization	CTAO ERIC (European Research Infrastructure Consortium)											
Alpha Config	2020	2021	2022	2023	2024	2025	2026	2027	2028	2029	2030	
LST North		Comissionin	ing and Operation of LST1				: A I ST Array	,				
LST NOT IT	C	DR	Depl	oyment of LS	ST2-4	Operation as 4 LST Array						
MST North	Design ar	nd Finance	INFRA			Construction of 9MSTs					Observatory Operation	
CTA South	Array conf	fig, Finance		Construction a		and Deplyment of 14 MSTs						
CTA South	and	CDR	1141	T.A.	Construction and Deployment of 37 SSTs							
Extension	2020	2021	2022	2023	2024	2025	2026	2027	2028	2029	2030	
LST South		Finance	e / CDR	Construction of 4 LSTs ???			c	peration ???				

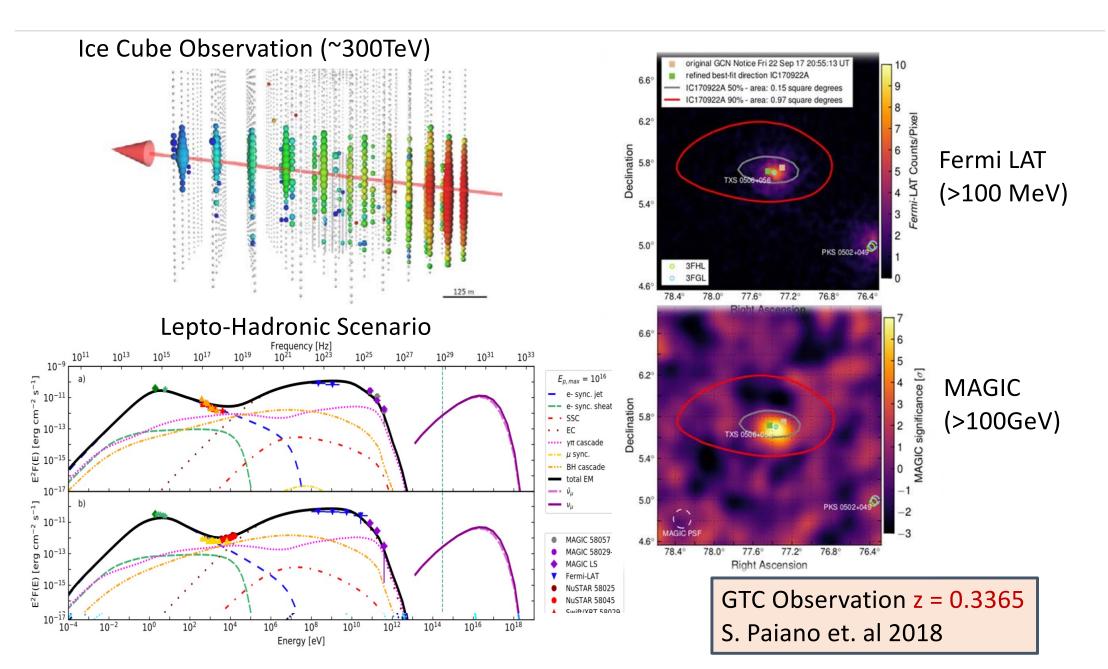


Multi-messenger and Multi-wavelength Astrophysics





Multi Messenger Astronomy IC170922A / TXS 0506+056



MAGIC Highlight, Gamma Ray Burst GRB190114C (z=0.42)

Historical achievement

 10^{-7}

 10^{-8}

 10^{-9}

 10^{-10}

10-7

 10^{-8}

10⁻⁹

 10^{-10}

68-110 s

110-180 s

XRT

GBM

106

BAT

Flux (erg cm $^{-2}$ s $^{-1}$)

Flux (erg cm $^{-2}$ s $^{-1}$)

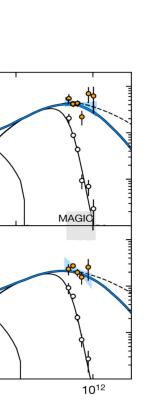
☐ First Detection of the GRB from ground.

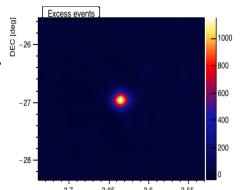
LAT

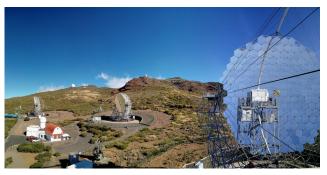
10⁹

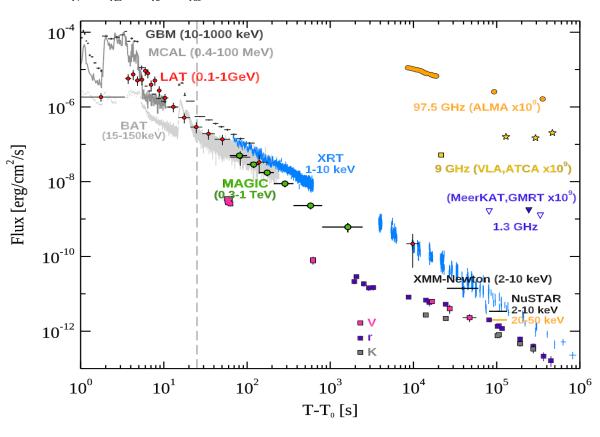
Energy (eV)

- □ ~100 Crab flux in the first minutes.
- ☐ TeV bump has a similar energetics with KeV-GeV bump





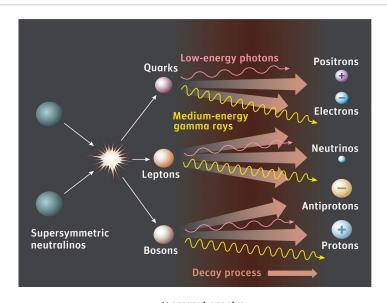


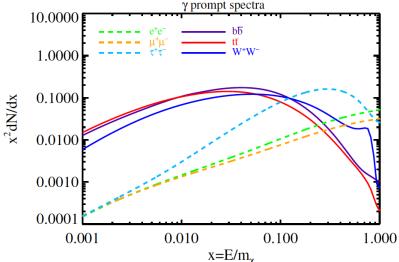




Dark Matter Search

Sensitive M_x : 200GeV - 10TeV

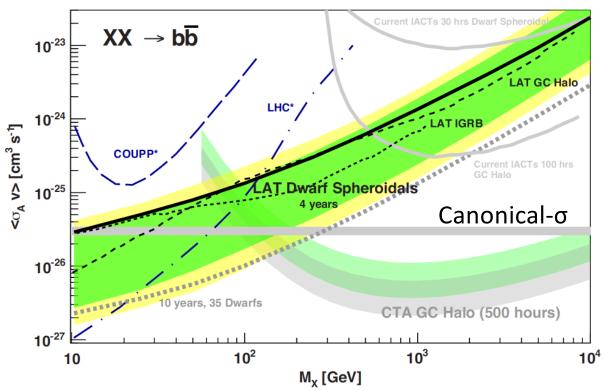




$$\frac{\mathrm{d}\Phi_{\gamma}}{\mathrm{dE}_{\gamma}} = \frac{1}{4\pi} \underbrace{\frac{\langle \sigma_{\mathrm{ann}} v \rangle}{2\mathrm{m}_{\mathrm{WIMP}}^{2}} \sum_{f} \frac{\mathrm{dN}_{\gamma}^{f}}{\mathrm{dE}_{\gamma}} \mathrm{B}_{f}}_{'\mathrm{Particle Physics'}} \times \underbrace{\int_{\Delta\Omega} \mathrm{d}\Omega' \int_{\mathrm{los}} \rho^{2} \mathrm{d}l(r, \theta')}_{'\mathrm{Astrophysics' or }J(E)}$$

Particle Physics

Astrophysics



Gamma rays from Annihilation produce the bump around 1/10 -1/20 M_X → 20GeV-1TeV domain

CTA gives the stringent upper limit. Stefan Funk 2015



- CTA is a big and ambitious project and plays an important role in the MM and MWL astronomy in the next decades
- CTA South construction will start in 2023
- CTA North construction, LST2-4 and MST1 will start in summer 2022 and completed in 2024
- LST5-8 construction in South is on Discussion to enhance the performance of CTA Observatory. INAF successfully got a funding for LST5-6.
- LST1 commissioning →
 - We confirmed LST1 satisfies the design performance
 - GRBs with the redshift up to z = 4 can be seen with LST
- ~10 sources are detected with LST1. Crab Pulsar, BL Lac flare, G.C., RS Oph are very interesting

