

# Computing requirements of the HotQCD Collaboration

C. Schmidt for the Collaboration



# Plan

- **The Collaboration**
- **The physics program**
- **Some research highlights**
- **Current resources**
- **Future plans**
- **Exascale readiness**
- **Conclusion**

# The HotQCD Collaboration

Alexei Bazavov, MSU East Lansing, MI, USA  
Heng-Tong Ding, CCNU Wuhan, China  
Jishnu Goswami, BU Bielefeld, Germany  
Prasad Hegde, IISc Bangalore, India  
Olaf Kaczmarek, BU Bielefeld, Germany  
Frithjof Karsch, BU Bielefeld, Germany  
Nikhil S Karthik, BNL Upton, NY, USA  
Anirban Lahiri, BU Bielefeld, Germany  
Sheng-Tai. Li, CCNU Wuhan, China  
Swagato Mukherjee, BNL Upton, NY, USA  
Hiroshi Ohno, TU, Tsukuba, Japan  
Peter Petreczky, BNL Upton, NY, USA  
Christian Schmidt, BU Bielefeld, Germany  
Sayantan Sharma, IMS, Chennai, India  
Patrick Steinbrecher, BNL Upton, NY, USA



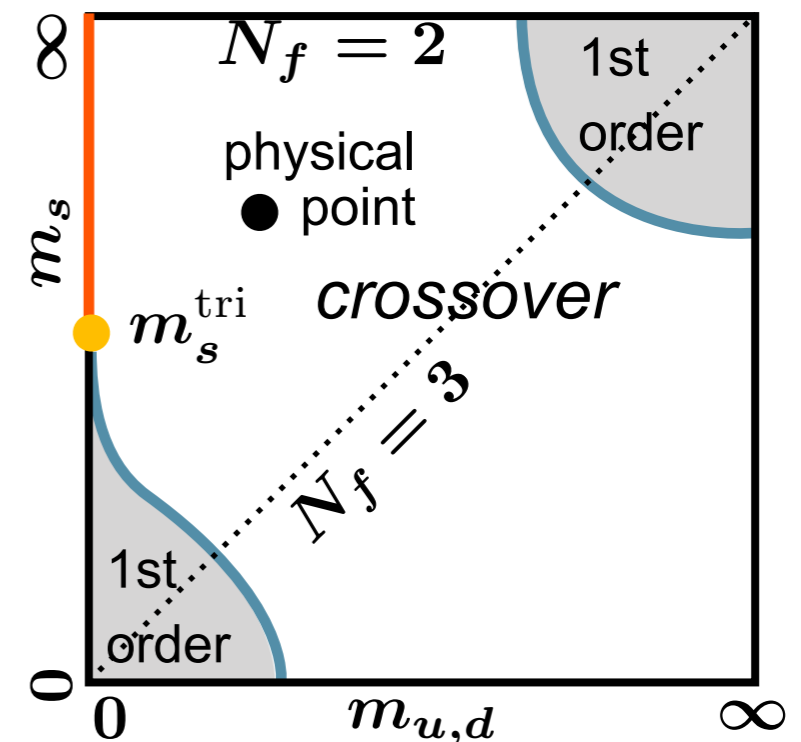
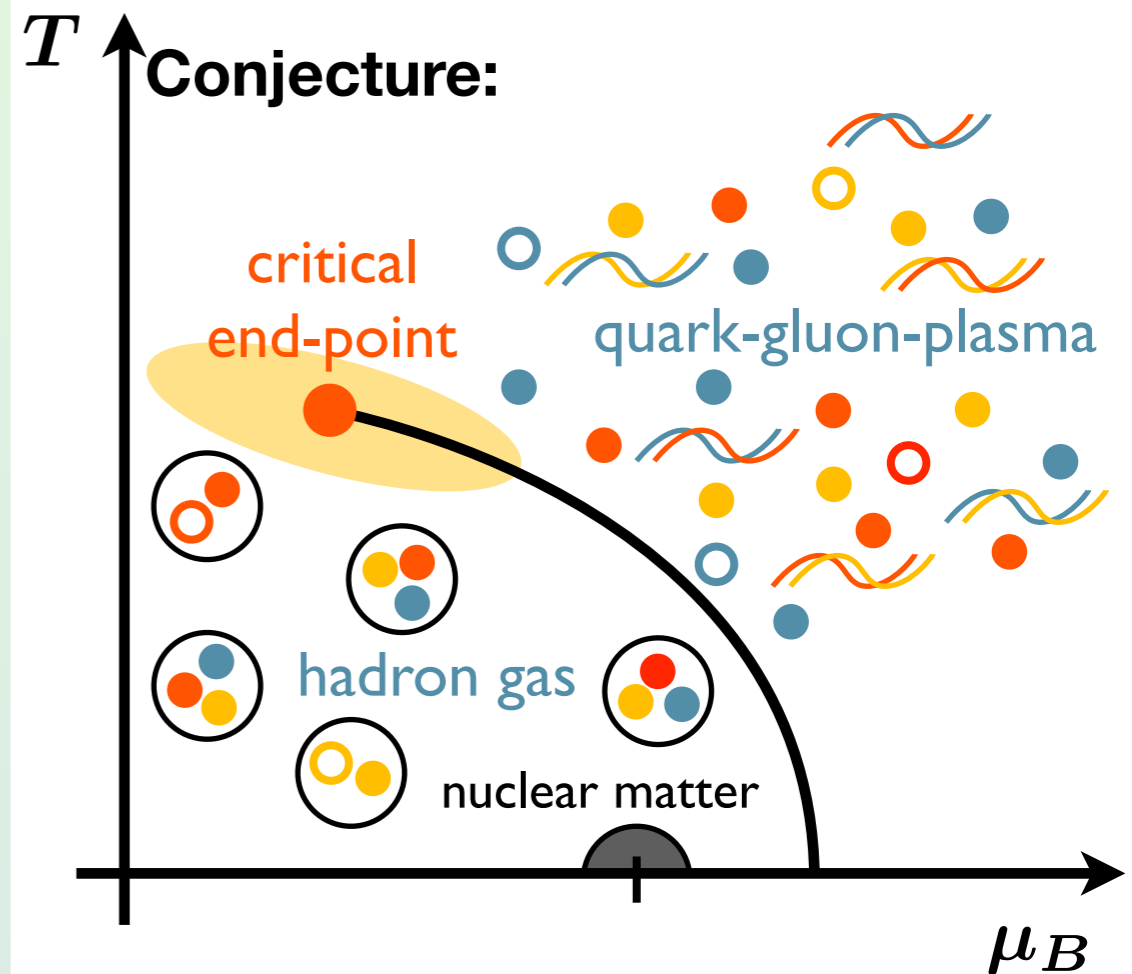
# Topics of interest

- **The QCD Phase diagram**

- The existence/location of the QCD critical point
- The nature of the chiral transition in the chiral limit and the role of the chiral anomaly
- QCD in external fields (constant magnetic background field, imaginary chemical potential)

- **The QCD equation of state (EoS)**

- Isentropic EoS as input for hydrodynamic calculations
- The relevant degrees of freedom at low and high temperatures, thermal modifications of hadrons



Pisarski, Wilczek (1984)

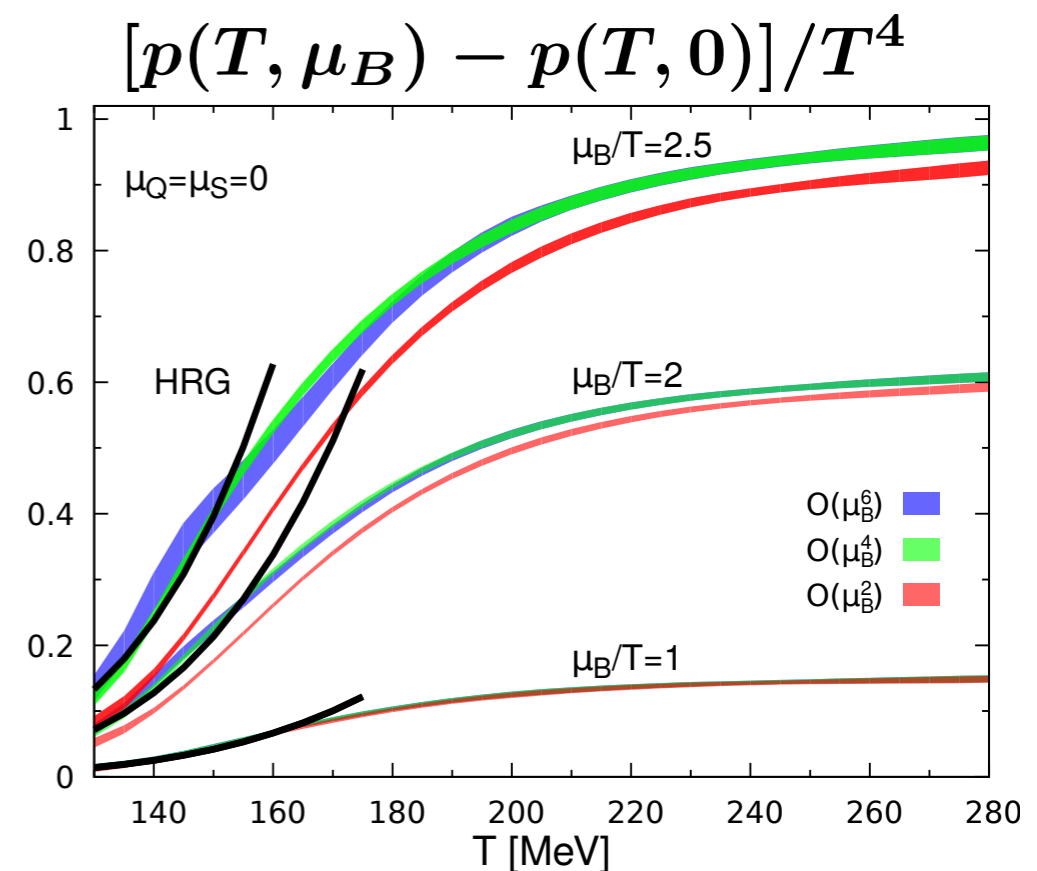
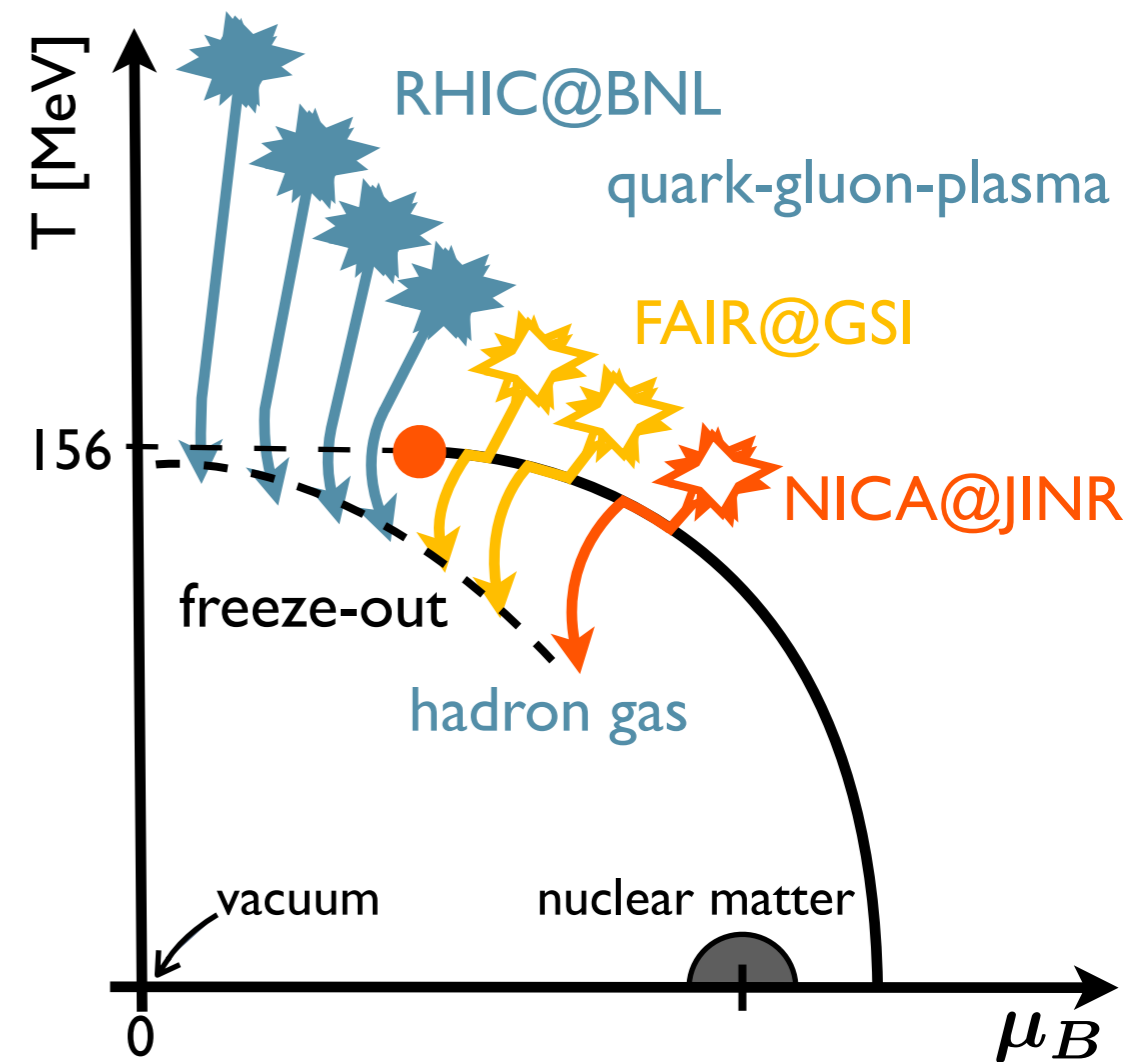
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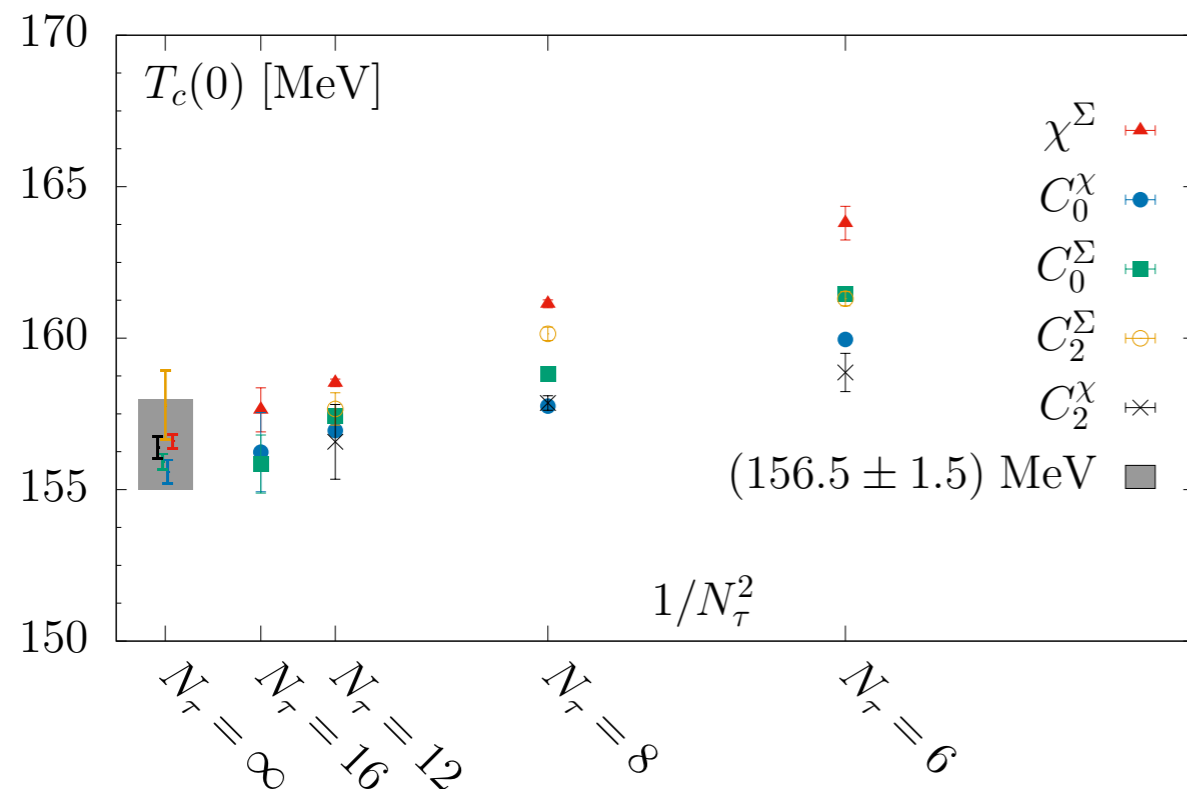
# Highlighting some results

- The chiral crossover of (2+1)-flavor QCD at physical quark masses

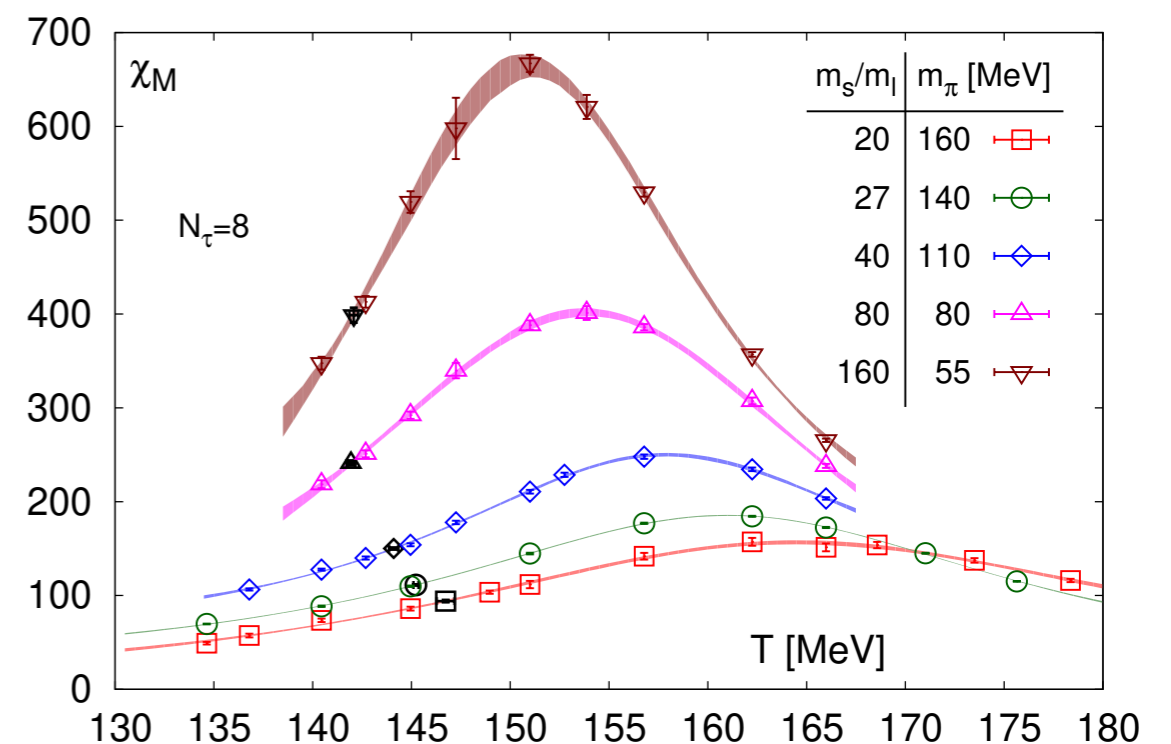
- New estimate of the pseudo critical temperature  $T_{pc} = 156.5 \pm 1.5$  MeV
- The curvature of the pseudo critical line  $\kappa_2^B = 0.012(4)$

- Universal scaling towards the chiral limit of (2+1)-flavor QCD

- Consistency with an O(4) critical point in the chiral limit is found
- The chiral transition temperature is estimated to  $T_c = 132_{-6}^{+3}$  MeV



HotQCD, PLB 795 (2019)



HotQCD, PRL 123 (2019)



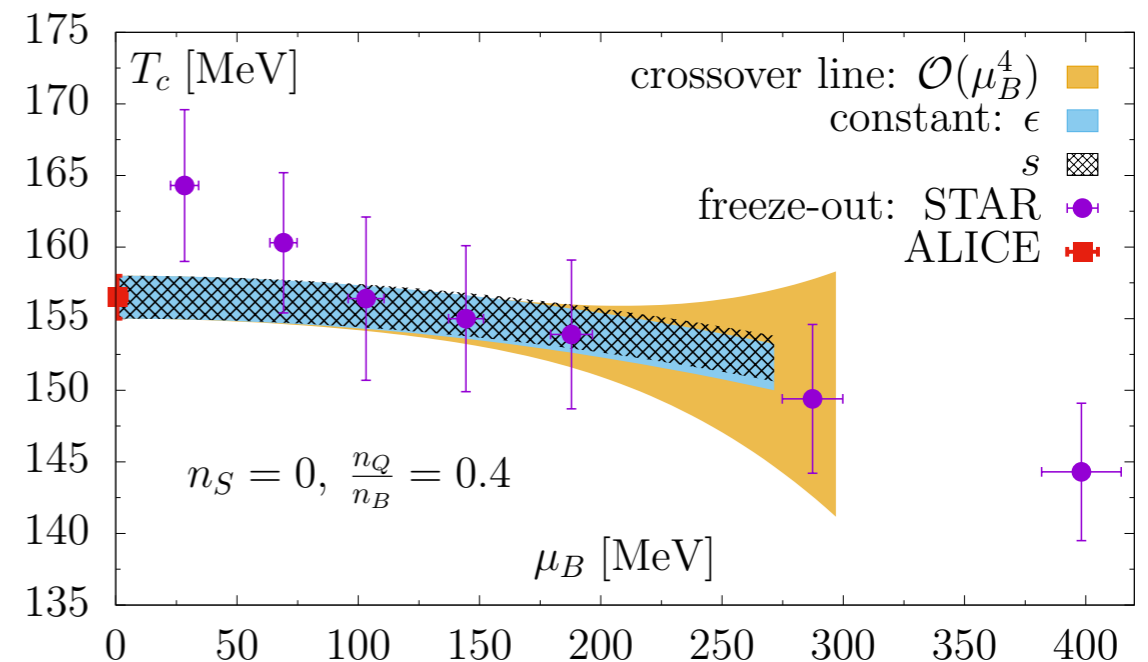
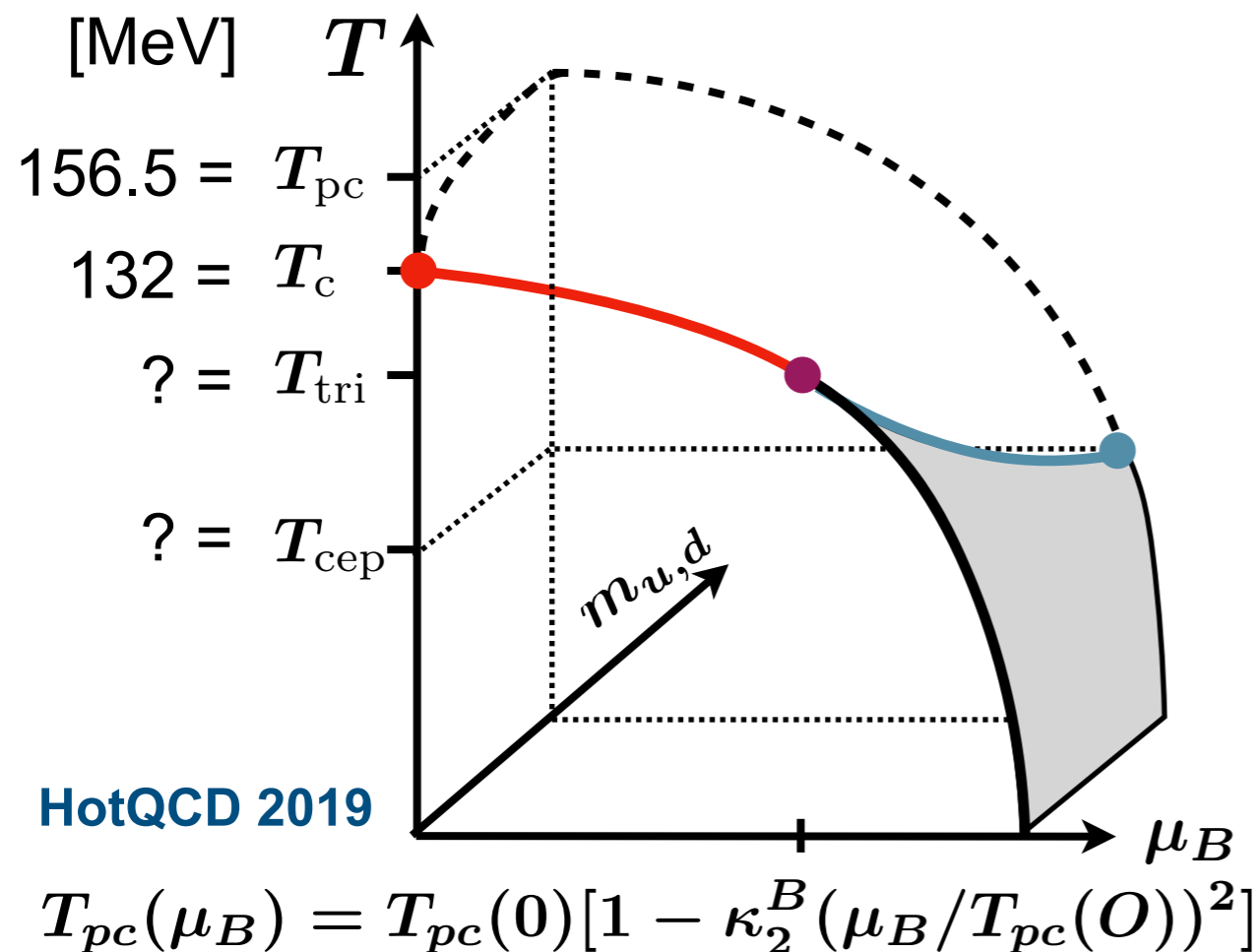
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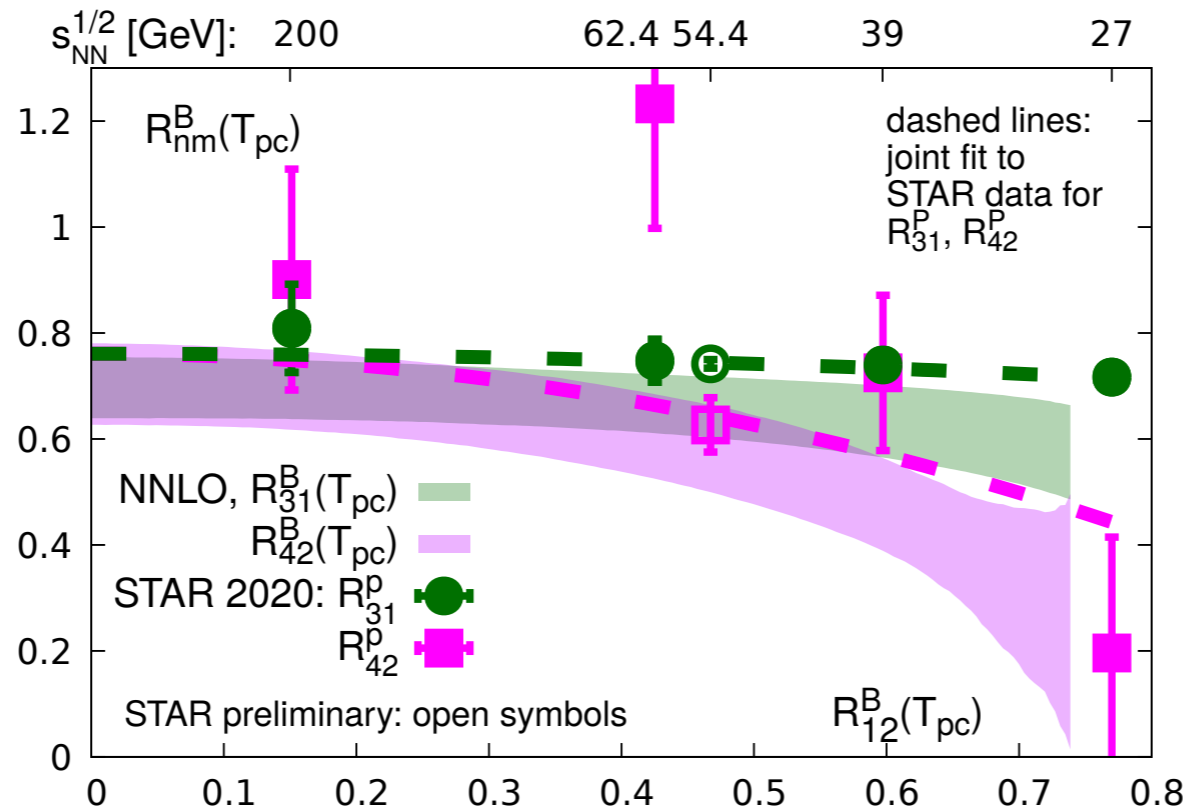
- The pressure expansion up to  $\mathcal{O}(\mu_B^8)$ 
  - NNLO calculation of Skewness and Kurtosis
  - NLO calculation of Hyper-skewness and Hyper-kurtosis

$$R_{nm}^B = \frac{\chi_n^B}{\chi_m^B}$$

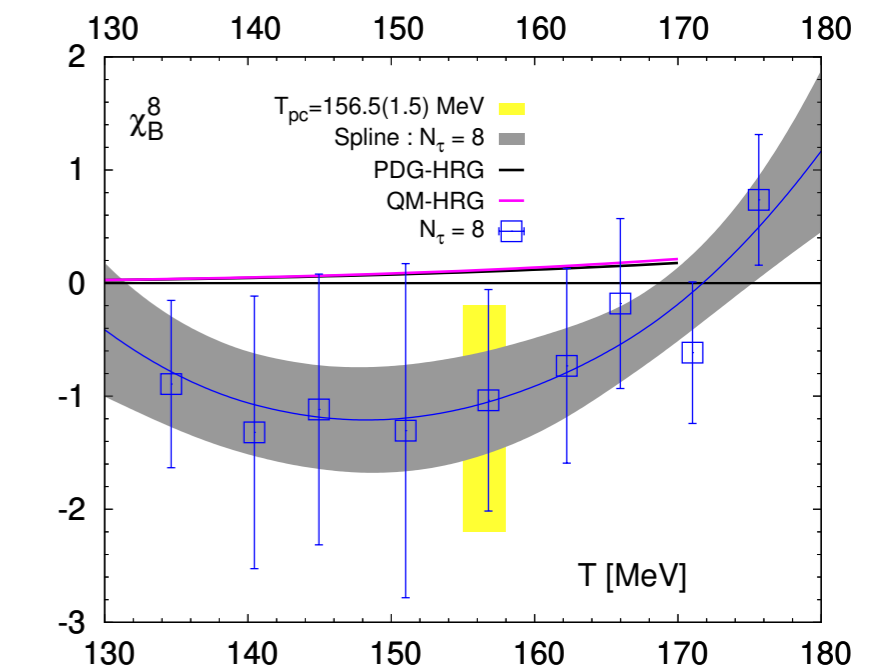
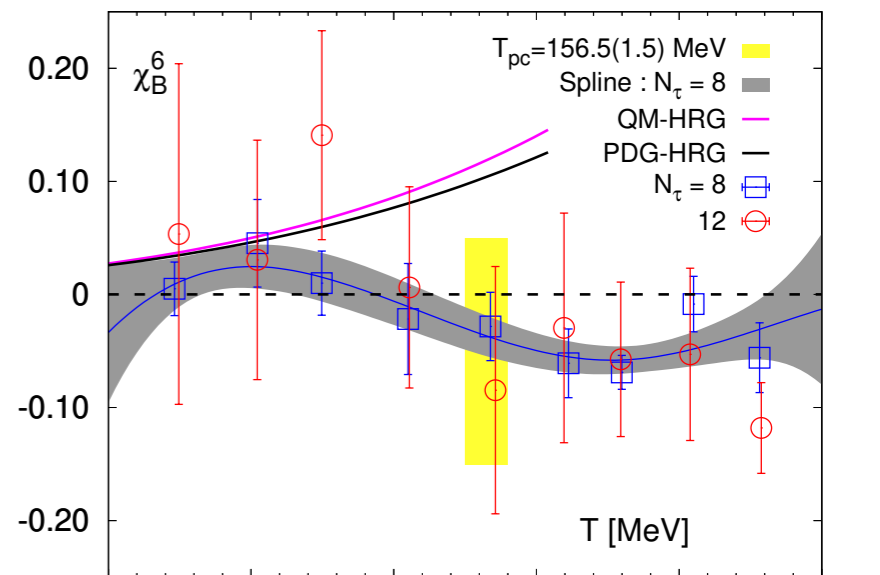
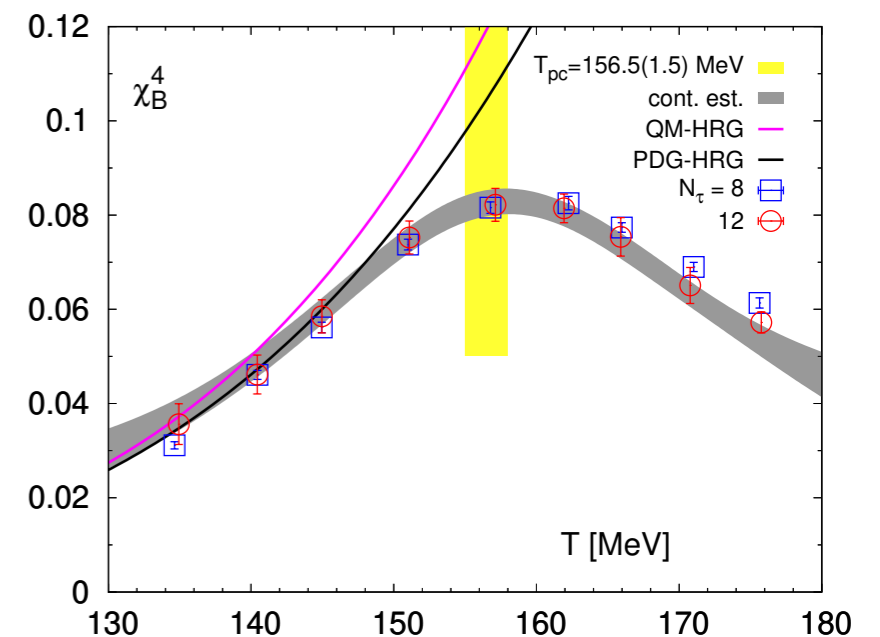
$$\chi_n^B = \frac{\partial^n \ln Z}{\partial \hat{\mu}_B^n T^3 V}$$

$$= \frac{\partial^n p}{\partial \hat{\mu}_B^n T}$$

$$\hat{\mu}_B = \mu_B / T$$



HotQCD, arXiv:2001.08530



HotQCD, arXiv:2001.08530



## Lattice methods

- **Generation of gauge ensembles with dynamical HISQ fermions**
  - Rational HMC
  - Multi-scale Omelyan integrator
  - Hasenbusch mass preconditioning
  - Multiple pseudo-fermion fields
  - Et cetera ...
- **Measurements (not necessary with HISQ, could be Domain Wall, ... )**
  - Susceptibilities and cumulants: statistical noise method with deflation
  - Eigenvalue spectrum: Lanczos/ Kalkreuter-Simma
  - Hadronic correlation functions

## Typical volumes

Physical mass:

$$48^3 \times 12$$

$$64^3 \times 16$$

Lighter than physical:

$$56^3 \times 8$$

$$80^3 \times 8$$

## Typical statistics

High order cumulants:

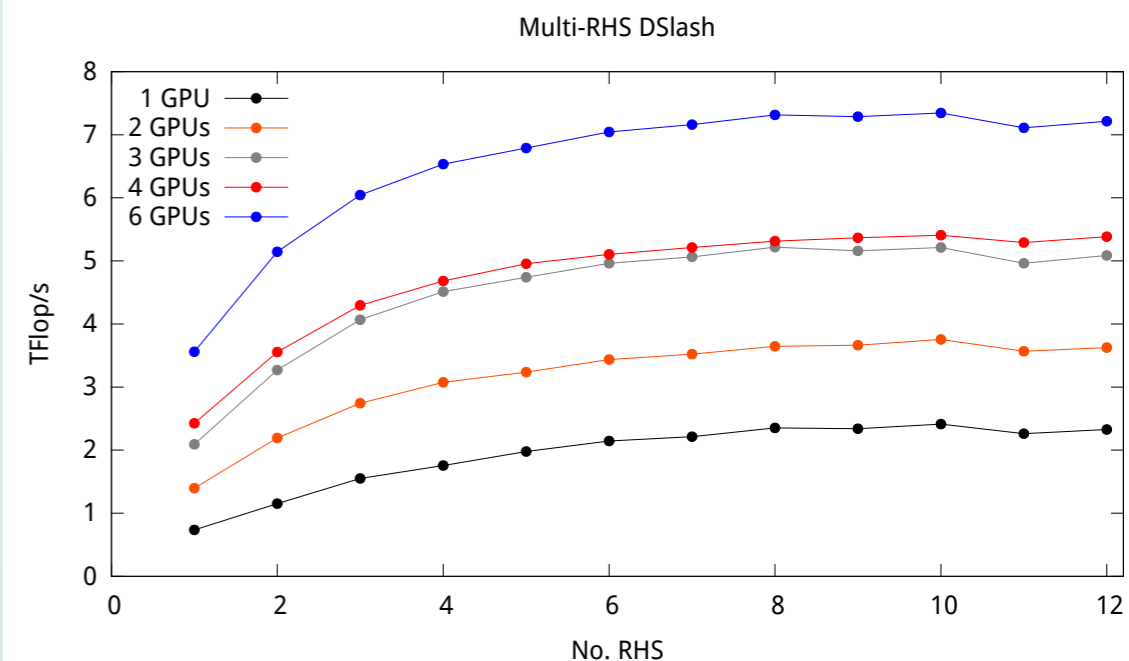
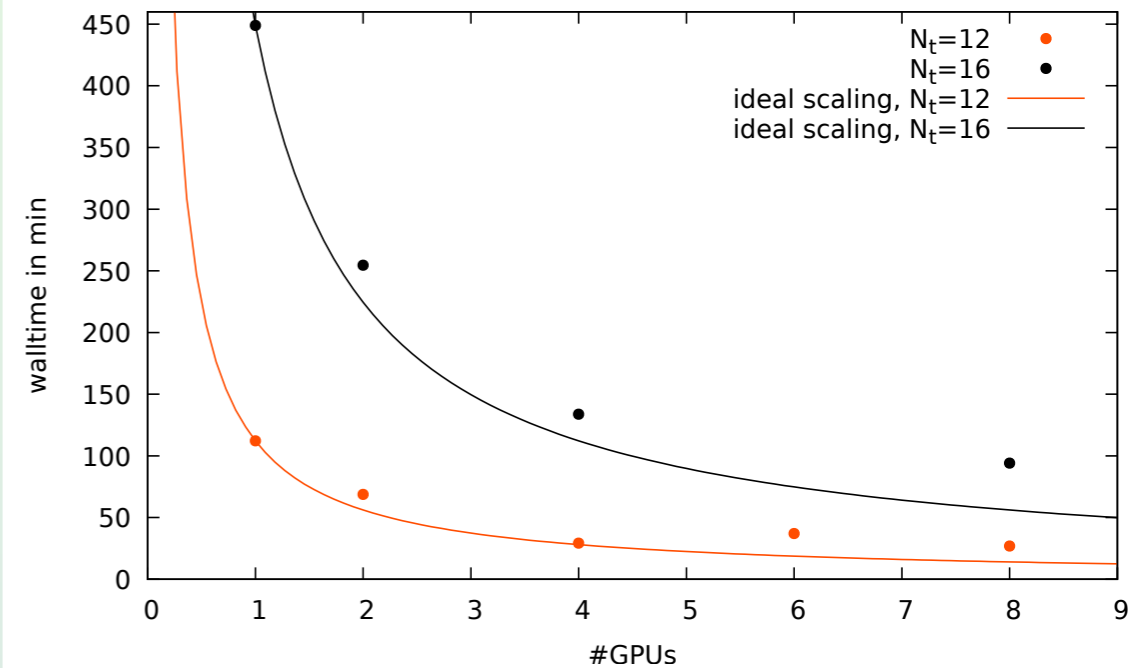
$$\mathcal{O}(10^6) \text{ per ensemble}$$

Chiral susceptibility:

$$\mathcal{O}(10^4) \text{ per ensemble}$$

# Hardware and performance

- **Developed massively parallel CPU (including KNL) and GPU code**
  - For generation and measurements
  - Use C++/CUDA
  - Scales good on few GPUs/nodes
  - Gain performance with appropriate memory fetching, multi-rhs solver, vectorization on CPUs, ...
- **Use existing libraries for some measurements**
  - QUDA for the calculation of high order cumulants
  - Grid for Domain Wall on HISQ



## Computing time in the past

- **Computing time on own Hardware**
  - Bielefeld:
    - 2006-2011 APEnext (5TFlops)
    - 2011-2019 GPU-Cluster 500 TFlops,
    - Since 2019 GPU-Cluster 3PFlops
  - CCNU:
    - Since 2018 GPU-Cluster 6PFlops
- **Computing time in the US**
  - Applications via INCITE, ALCC and USQCD on Titan, Summit, Cori, JLab-Cluster
- **Computing time in the EU**
  - Applications via PRACE, Gauss-Alliance and NIC on JUWELS, Pitz Daint, Marconi

## Typical resource mixture in the EU

TIER 3 (local): 46%  
TIER 1 (national): 18%  
TIER 0 (PRACE): 36%

## Future directions

- High order cumulants need a factor 10 more statistics
- Electric charge fluctuations are very sensitive to the light pion spectrum

→ Need finer lattices

→ Need larger volume

4<sup>th</sup> order:  $96^3 \times 16$

6<sup>th</sup> order:  $48^3 \times 8$

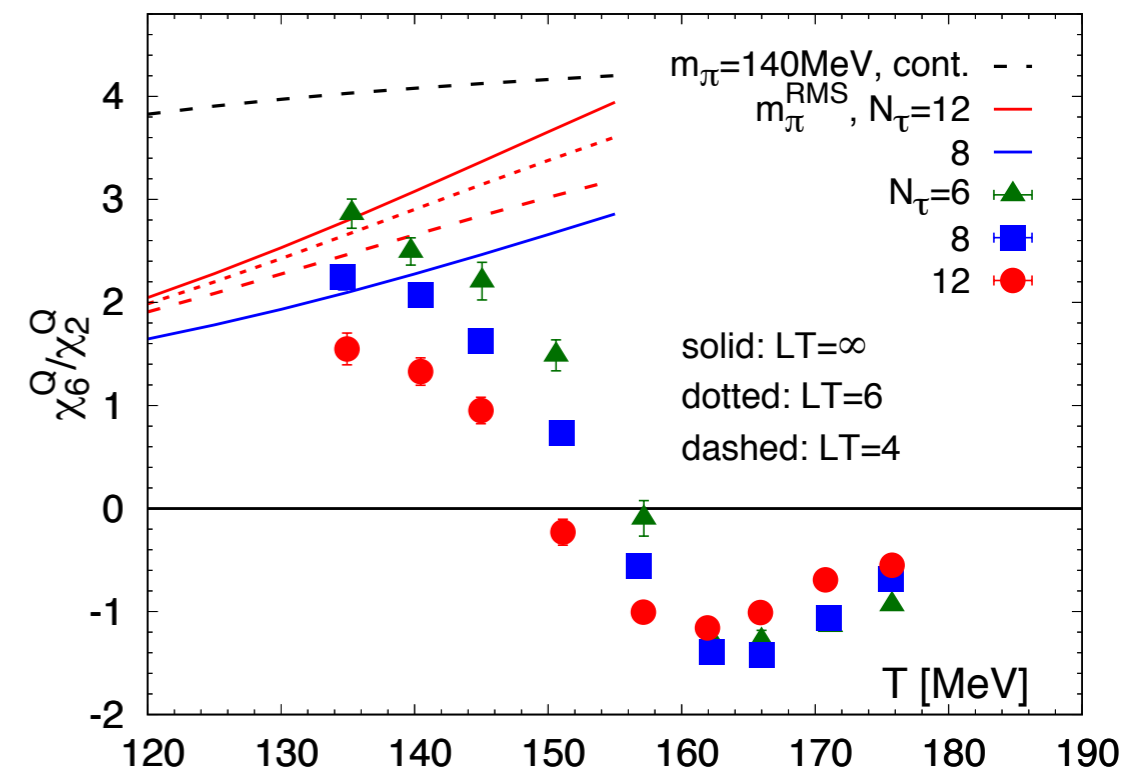
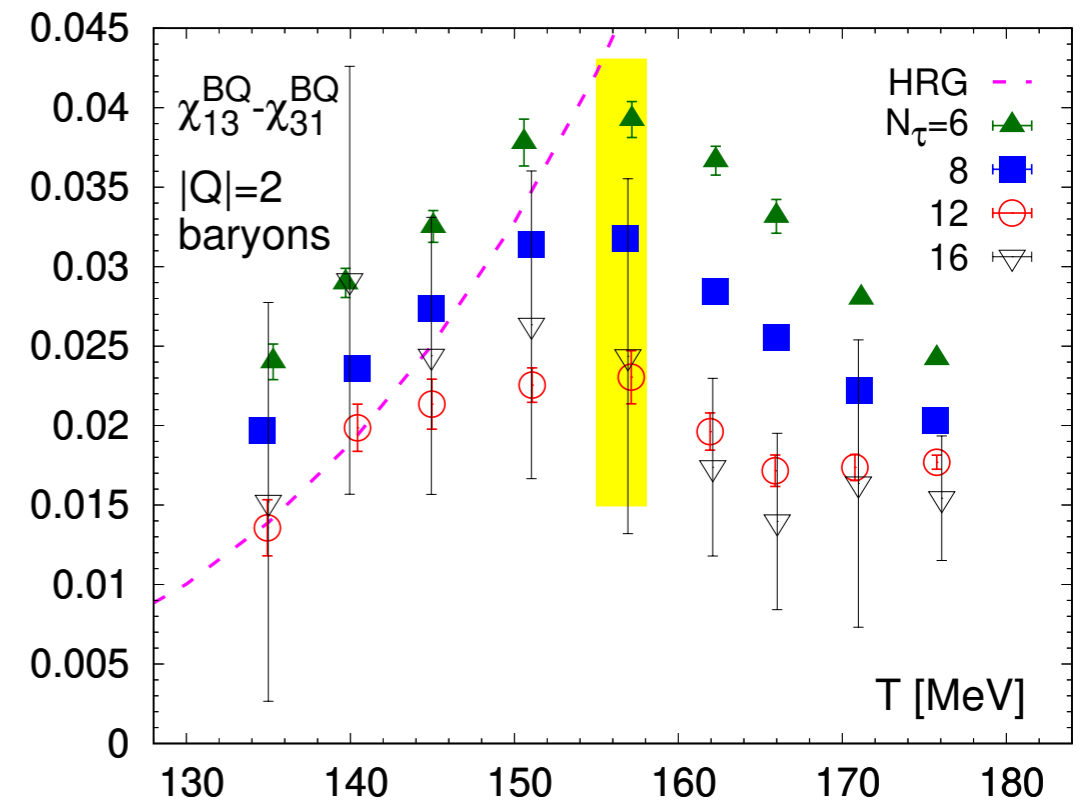
$72^3 \times 12$

→ More chiral actions

- Scaling studies require smaller quark masses and larger spacial volumes

$m_l/m_s=1/320$ :  $80^3 \times 8$

$120^3 \times 12$



# Exascale readiness

- **The number of ensembles is huge**
  - Set of parameters is large
  - Large statistics enables splitting the run in  $\mathcal{O}(1000)$  streams
  - Combine many runs into big jobs

## Conclusion

- **A factor 10 increase in computing time can be very well justified**

