Status of the Short-Baseline Near Detector [SBND] at Fermilab

Supraja Balasubramanian,

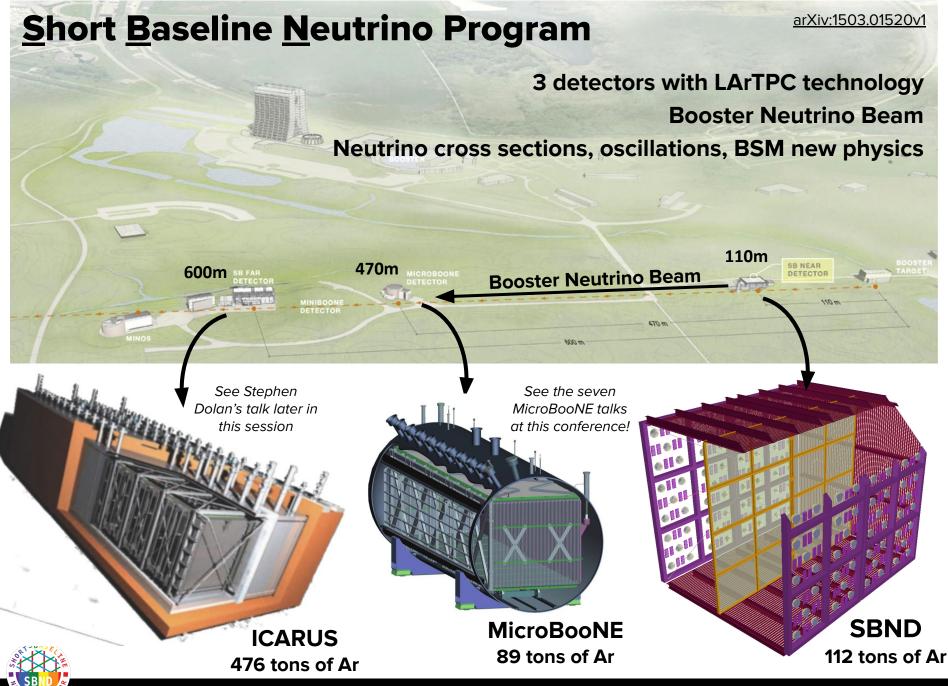
on behalf of the SBND collaboration

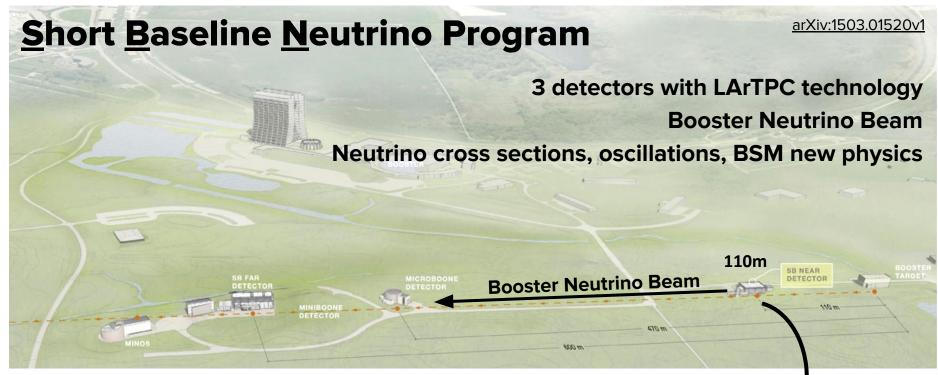
28 October 2022 | NuINT 2022











Short Baseline Near Detector

Last LArTPC to begin operating before DUNE

=> last prototype for various design elements, operations, data reconstruction, etc.

The **near detector** of the SBN program [110 m from target].

High-intensity neutrino beam + proximity to target

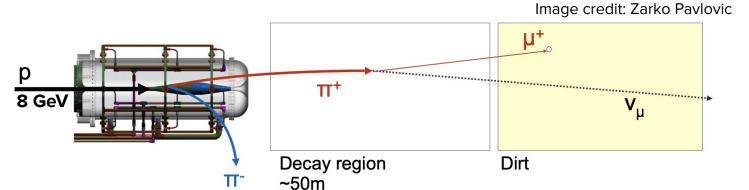
=> large statistics of neutrino-argon interactions, off axis fluxes.

Start of operations planned for **2023**.

SBND

112 tons of Ar

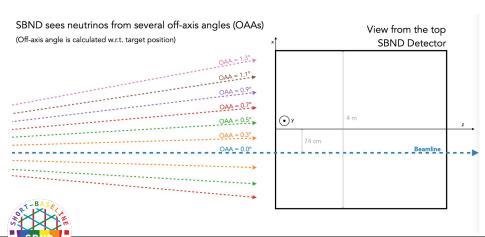
The Booster Neutrino Beam @ SBND

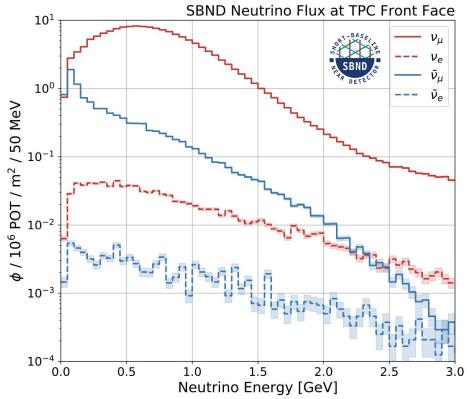


Beam Composition:

$$v_{\mu} = 93.6\%$$
 $\bar{v}_{\mu} = 5.9\%$
 $v_{e} + \bar{v}_{e} = 0.5\%$

- <E_v> ~ 800 MeV.
- Projected to take 10-18 X 10²⁰ POT of data in total => large statistics on Argon.
- Close to the target + slightly off-axis => SBND can sample off-axis fluxes (hear more about this in the next talk by Lauren Yates).





SBND Physics: Neutrino-Argon Cross Sections

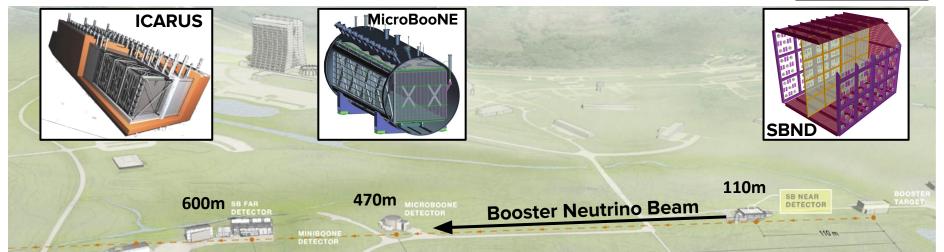
See next talk, "Neutrino Interaction Measurement"

Capabilities of the SBND Experiment"

by Dr. Lauren Yates

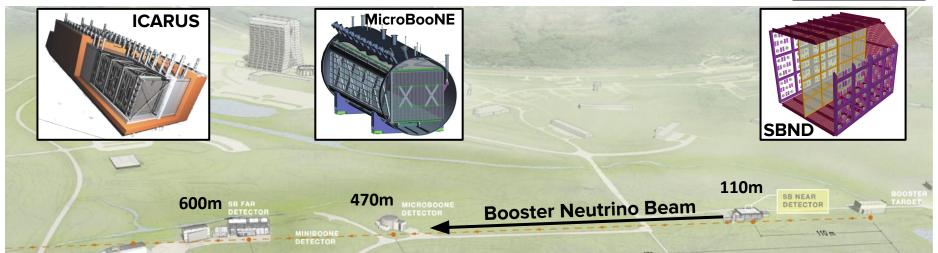


arXiv:1503.01520v1



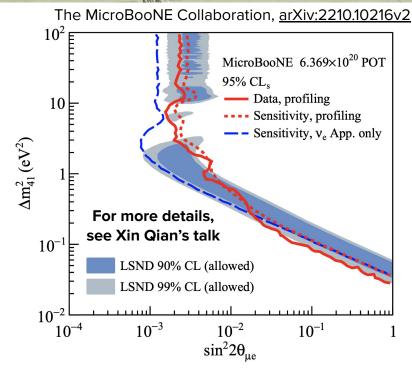


arXiv:1503.01520v1

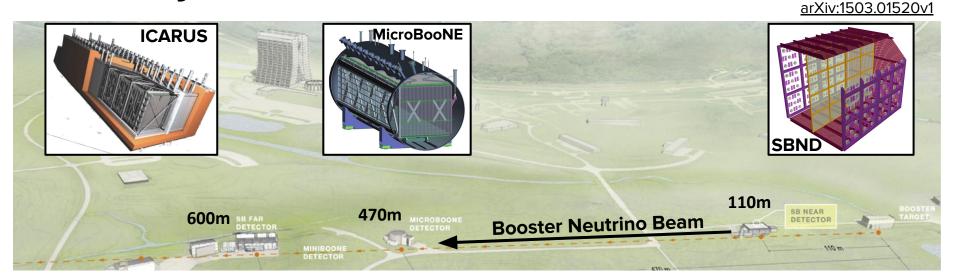


MicroBooNE did not see a "low-energy excess" of electromagnetic events as observed by MiniBooNE, which could have been evidence of sterile neutrino oscillations.

But this does not fully rule out sterile neutrinos.







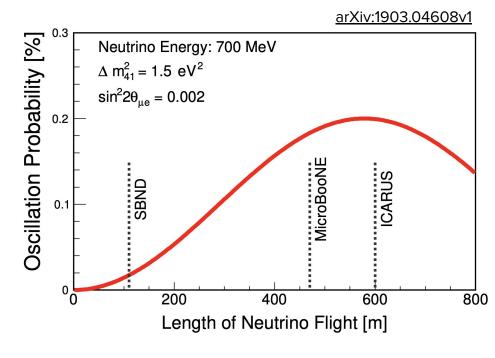
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SBN Program:

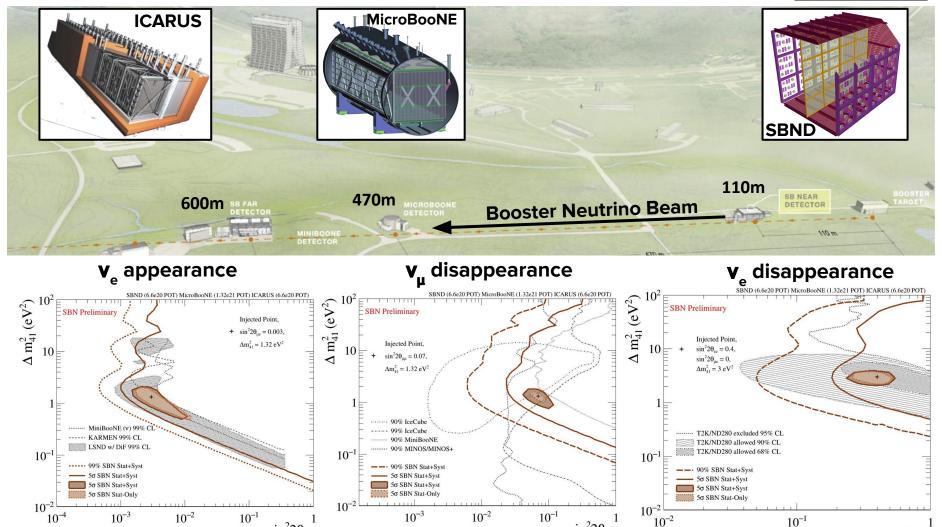
- Near Detector (SBND) with large statistics to constrain systematics.
- Far Detector (ICARUS) with large mass for increased exposure.

Substantially improve the global dataset regarding sterile neutrinos.



 $\sin^2 2\theta_{\text{He}}$

arXiv:1503.01520v1



The SBN program will probe all three sterile neutrino oscillation channels:

 $\boldsymbol{v}_{_{\boldsymbol{e}}}$ appearance, $\boldsymbol{v}_{_{\boldsymbol{\mu}}}$ disappearance, and $\boldsymbol{v}_{_{\boldsymbol{e}}}$ disappearance. Expect $\boldsymbol{5\sigma}$ sensitivity.

External data: <u>lceCube **v**,,,(2020)</u>, <u>MINOS **v**,, (2017)</u>, <u>T2K **v**, (2014), <u>MiniBooNE **v**, (2013)</u> & <u>**v**,,(2011)</u>, <u>KARMEN **v**, (2002), LSND **v**, (2001)</u></u>

 $\sin^2 2\theta_{\text{min}}$

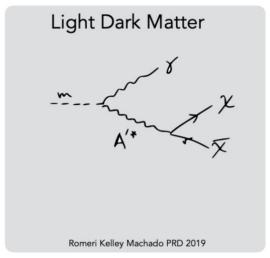
 $\sin^2 2\theta_{aa}$

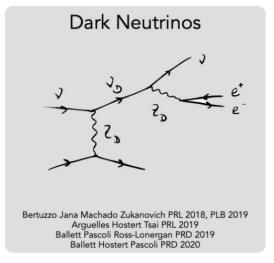
SBND Physics: BSM New Physics

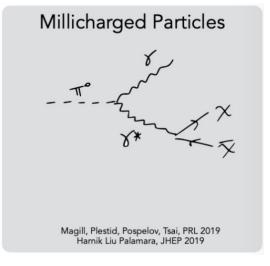
- **High-intensity** proton beam **3 mm resolution** 3D event reconstruction
- Large-mass LArTPC

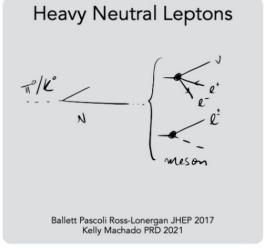
Excellent particle identification with low thresholds

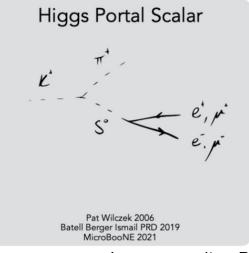
=> SBND can search for a variety of BSM phenomena.











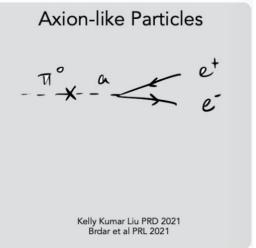




Image credits: Pedro Machado, Marco Del Tutto

SBND Physics: BSM New Physics

We introduce the model into our simulation & run reconstruction.

- High-intensity proton beam 3 mm resolution 3D event reconstruction
- Large-mass LArTPC

Excellent particle identification with low thresholds

=> SBND can search for a variety of BSM phenomena.

Light Dark Matter



single e⁻ scattering or e⁺e⁻ pair with no hadronic activity

Dark Neutrinos



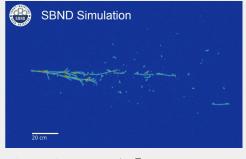
e⁺e⁻ pair with or without hadronic activity

Millicharged Particles



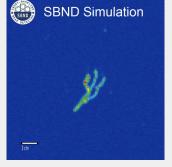
blips or faint tracks

Heavy Neutral Leptons



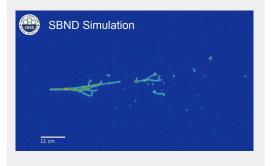
 e^+e^- , $\mu^+\mu^-$, or $\mu^\pm\pi^\mp$ pair with no hadronic activity

Higgs Portal Scalar



 e^+e^- or $\mu^+\mu^-$ pair with no hadronic activity

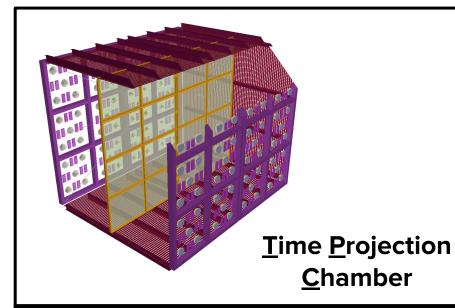
Axion-Like Particles

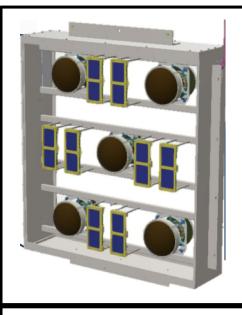


high-energy e^+e^- or $\mu^+\mu^-$ pair



SBND: The Detector Subsystems

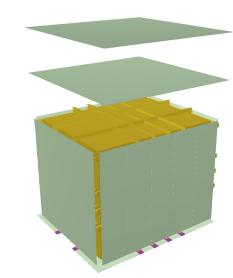




Photon Detection System



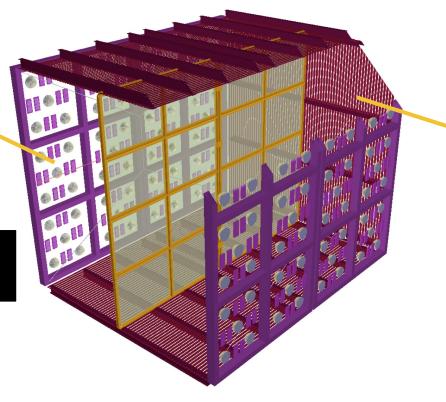
Cryostat



<u>Cosmic Ray</u> <u>Tagger</u>



2 Time Projection Chambers [total 4 X 4 X 5 m]





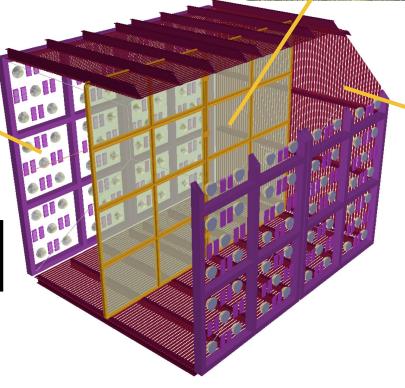




2 Time Projection Chambers [total 4 X 4 X 5 m]



Cathode Plane
in the middle;
divides the detector
into 2 TPCs. Will be
supplied with
-100 kV.

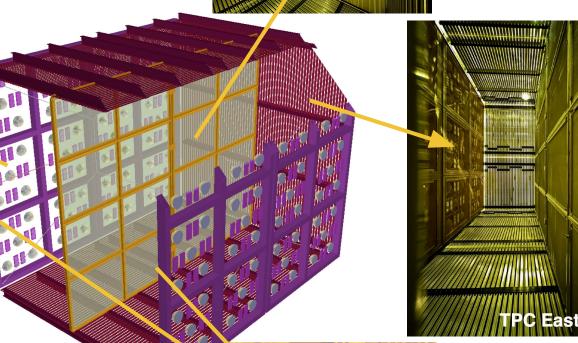






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Anode Plane on either side. Each consists of 3 planes of wires with 3 mm spacing and different angle per plane.





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-100 kV.

2 Time Projection Chambers [total 4 X 4 X 5 m]

Field Cage that wraps around the 2 LArTPCs to step down the voltage & ensure uniform electric field of 500 V/cm.



Anode Plane on either side. Each consists of 3 planes of wires with 3 mm spacing and different



Cold
Electronics to
pre-amplify &
digitize
signals in the
cold



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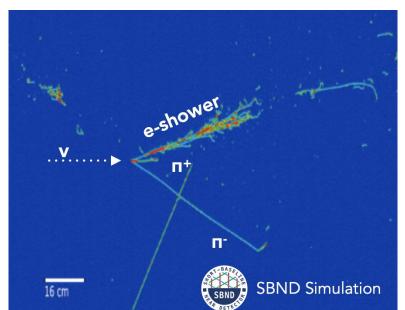


3D event reconstruction.

3 mm position resolution.

Low momentum thresholds.

Particle ID: e/γ separation μ , π , p,etc identification.



2 Time Projection C [total 4 X 4 X 5

Field Cage

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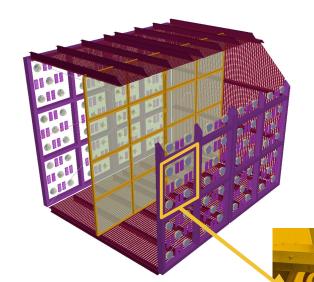


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SBND Photon Detection System



24 Anode Plane boxes

4x24 = 96 **PMTs** (TPB coated)

coated PMT

X-ARAPUCA

uncoated

PMT /

1X24 = 24 **PMTs** (uncoated)

8x24 = 192

X-ARAPUCAs*

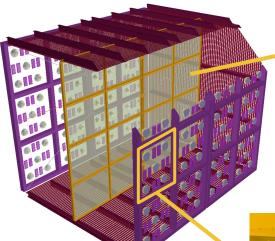
*sensitive to UV + visible light



SBND Photon Detection System

Cathode Plane

with TPB coated reflective foils mounted between mesh panels.





coated PMT

X-ARAPUCA

uncoated

PMT ↓

24 Anode Plane boxes

4x24 = 96 **PMTs**

(TPB coated) 1X24 = 24 **PMTs**

(uncoated)

8x24 = 192

X-ARAPUCAs*

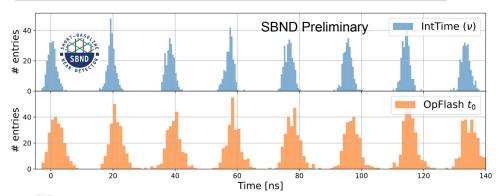
*sensitive to UV + visible light



SBND Photon Detection System

- Primary scintillation and reflected light: improved and more uniform total light yield.
- **Triggering:** recent improvements in timing resolution to resolve the beam structure & identify interaction time.
- Cosmic background tagging: based on amount of light + 3D position reconstruction.
- Calorimetry: light information can supplement TPC information. arXiv:2203.00740

Experiment	Average light yield	Uniform light
	(PE/MeV)	collection?
MicroBooNE	~ 5	no
LArIAT	~ 18	yes
pDUNE-SP	1.9 at 3.3m	no
SBND	$\sim 80 \ (> 50 \ {\rm min})$	yes
DUNE: Vertical Drift	$\sim 38 \ (> 16.5 \ \text{min})$	ves



Simulated (top) and reconstructed (bottom) light flashes showing the neutrino beam structure.

with TPB coated

reflective foils mounted



24 Anode Plane boxes

4x24 = 96 **PMTs** (TPB coated) 1X24 = 24 PMTs(uncoated)

8x24 = 192X-ARAPUCAs*

*sensitive to UV + visible light



SBND Cosmic Ray Tagger

 3-4 cosmic muons in TPC per readout window.

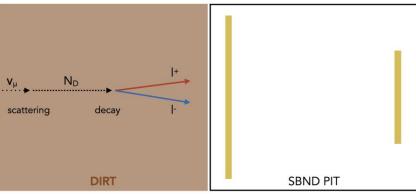
> Cosmic \ ray

- 4π coverage important for surface detectors
- Can act as a "beam telescope," e.g. to look for BSM new physics particles decaying in the dirt around SBND (dark neutrino analysis development underway).

Top CRT panels

Cryostat

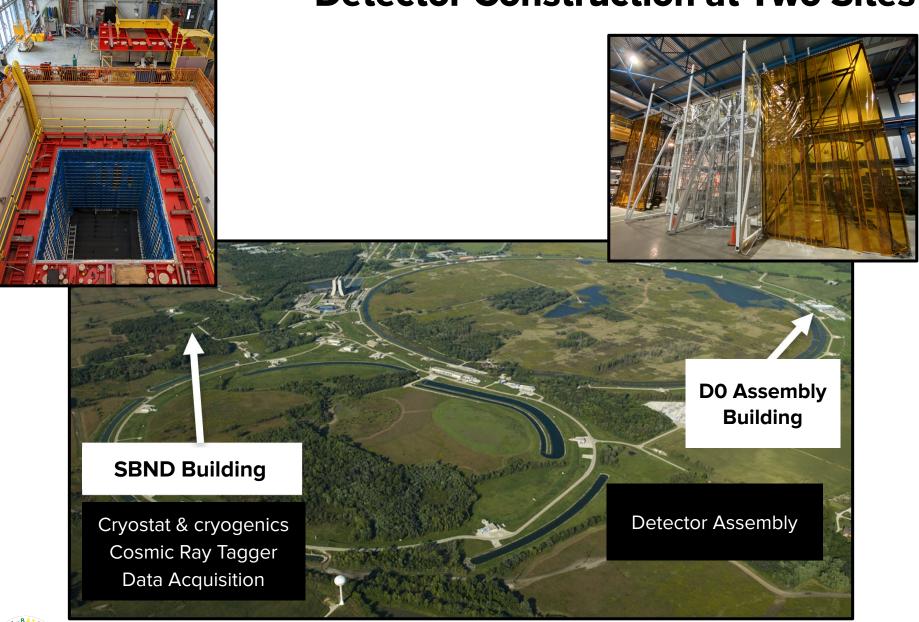






- Prototype for DUNE elements similar to DUNE.
- The detector is suspended from top caps and held inside the cryostat.
- Two top caps through which all cabling and instrumentation Side CRT panels will be passed.

Detector Construction at Two Sites







Empty Assembly Transport Frame, December 2019





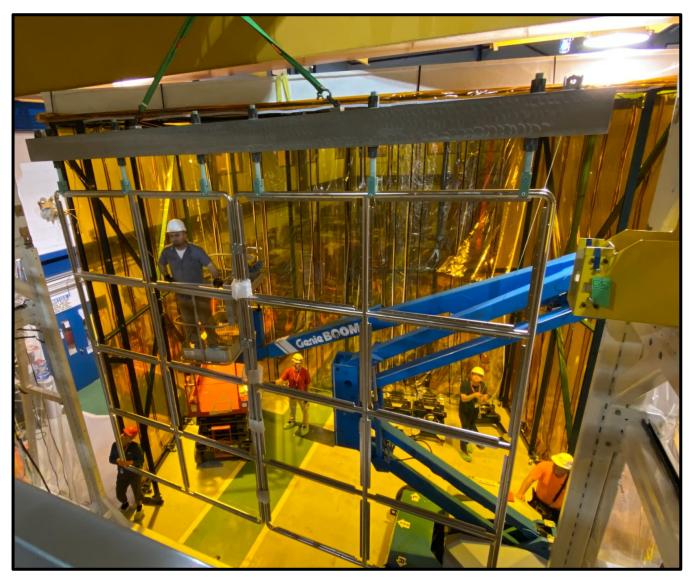


Clean tent with UV filters fully constructed, April 2021









Cathode Plane structure being installed, July 2021











Anode Plane Assembly with wires being brought into place, October 2021













Installation of Cold Electronics, December 2021 & May 2022



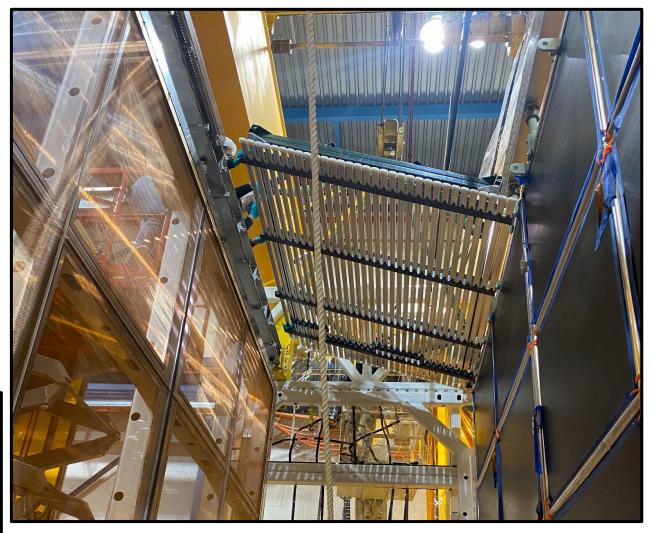












Field Cage top module being lowered,
January 2022

















Photon Detection System boxes fully installed behind anode wire planes, September 2022





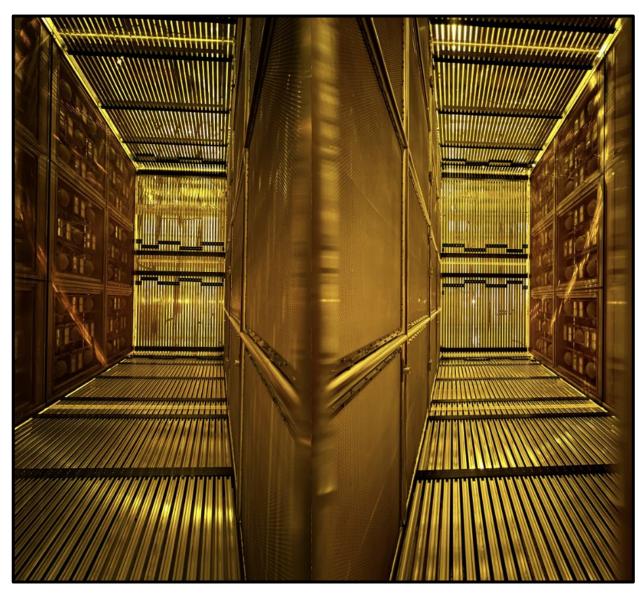






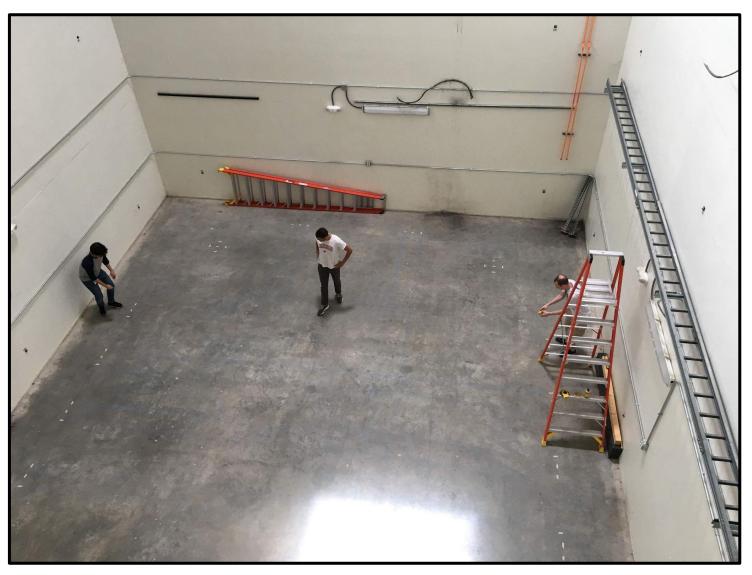






Fully completed assembly of 2 TPCS + Photon Detection System, September 2022





Empty pit where the future detector will go, July 2019







Bottom Cosmic Ray Tagger modules installed, September 2019









Cryostat outer shell completed, November 2019











Data Acquisition System electronic racks installed in the mezzanine level, February 2020

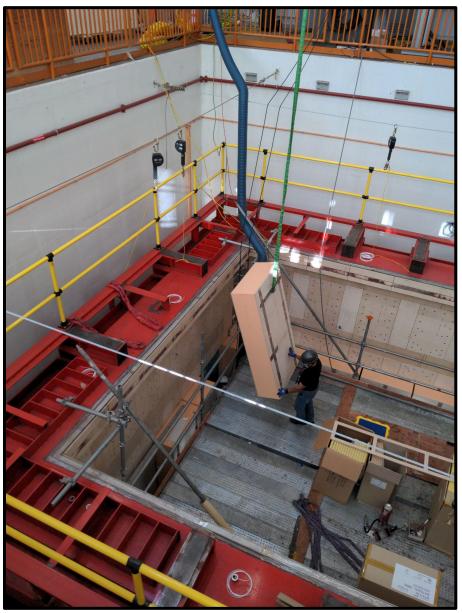


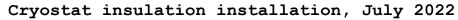














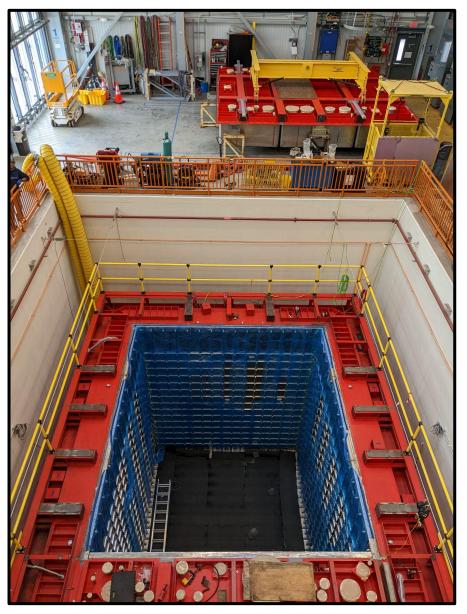












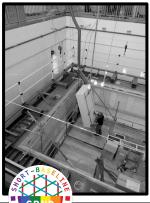
Finished cryostat, September 2022



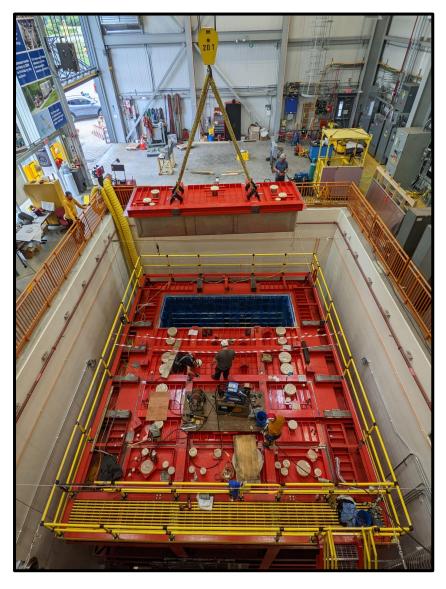












Top caps with ports for cabling & instrumentation test installation, October 2022

The Road to Physics

Detector move: We will move the detector in the Assembly Transport Frame from assembly site to SBN-ND Building (permanent location of SBND).

Planned for late November 2022.

Activities at SBN-ND: Quality check of all detector subsystems, installation of instrumentation, finalizing cryogenics, lowering into position inside the cryostat, etc. **December 2022-early Summer 2023.**

Filling of liquid argon & cold commissioning

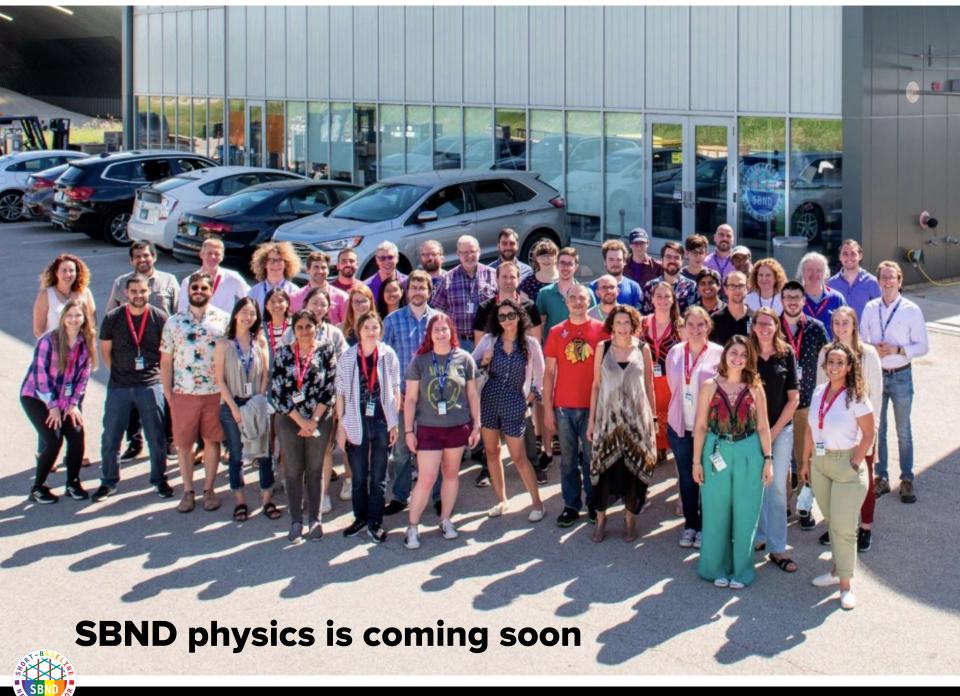
Summer 2023.



Critical Areas for data recording

gh Traffic Zones

Proposed route



Backup



SBND Light Collection & Triggering

X-Arapuca technology uses dichroic filters and waveguides to guide photons to SiPM

