



Measurements of Electron-Neutrino Interactions in MicroBooNE

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Outline

- Why do we care about electron-neutrino measurements (and particularly on argon)
- MicroBooNE and its beams
- ν_e CC inclusive measurements using the NuMI beam
- ν_e CC measurement using the BNB
- Future outlook.



Big Questions in Neutrino Physics

- The questions below, are what is currently driving the field of experimental neutrino physics

- *how much do neutrinos weigh?*
- *what is the nature of the ν ?*

- *which neutrino is the heaviest and which is the lightest (MH)?*

- *do neutrinos violate CP?*

- *is our picture correct?*

- *are there more than 3 kinds of neutrinos?*

β decay
and $0\nu\beta\beta$ decay

long-baseline
neutrinos

short-baseline
neutrinos

ν_e
appearance
in 0.1-10GeV
Range!
e.g. in
LArTPCs



Looking back to last NuINT






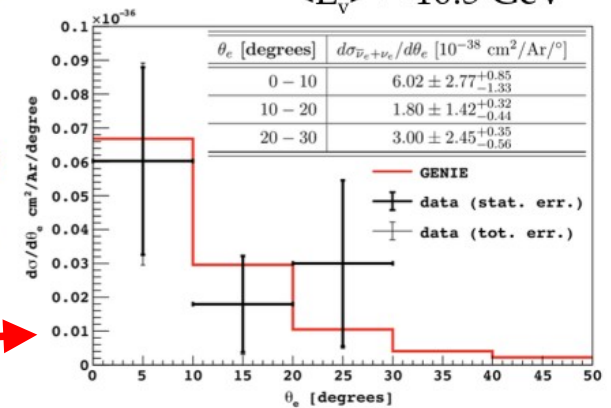
ν_e cross sections on argon circa last NuINT



All ν_e measurements ca. early 2020

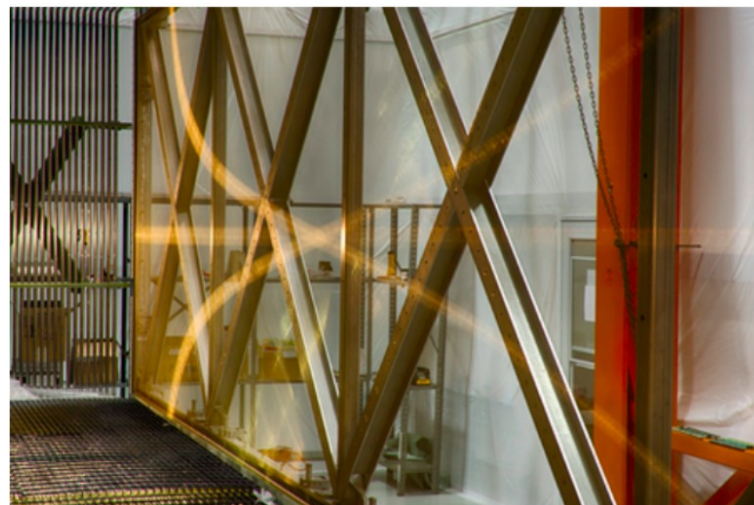
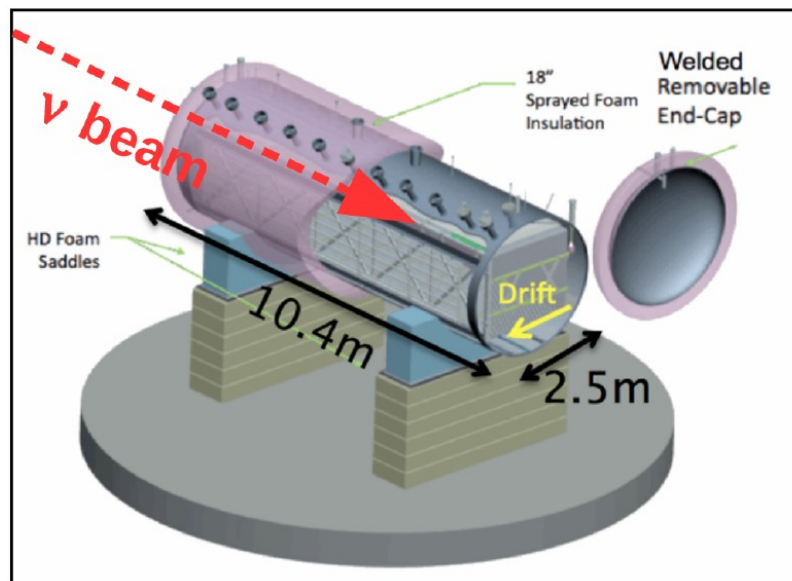
Experiment	Target	Type	Reference
Gargamelle	CF3Br	Inclusive	<i>Nuclear Physics B</i> , vol. 133, no. 2, 1978
T2K	C, H ₂ O	Inclusive	<i>Phys. Rev. Lett.</i> , vol. 113, 2014 <i>Phys. Rev. D</i> , vol. 91, 2015. <i>Journal of HEP</i> , vol. 2020, no. 10, 2020
MINERvA	(C ₈ H ₈) _n	Exclusive	<i>Phys. Rev. Lett.</i> , vol. 116, 2016.
ArgoNeuT	Ar	Inclusive	<i>Phys. Rev. D</i> , vol. 102, 2020. 

Only one on Ar, with 13
 $\nu_e + \bar{\nu}_e$ events
 $\langle E_{\nu} \rangle = 4.3$ GeV
 $\langle E_{\bar{\nu}} \rangle = 10.5$ GeV

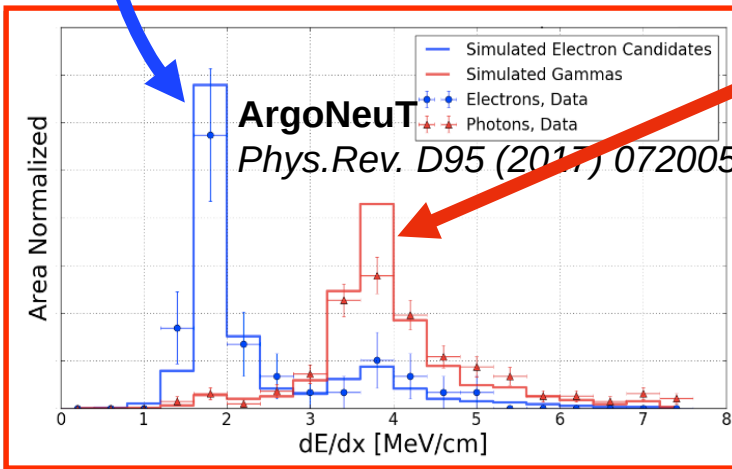
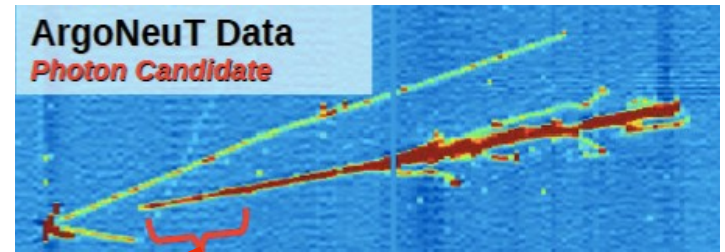
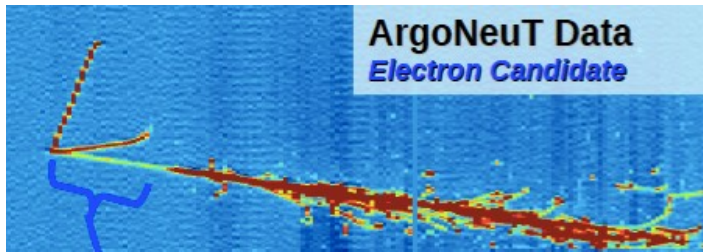


MicroBooNE at a glance

- See Xin's talk on Monday for overview.
- 170 tons of LAr (90 tons active).
- Scintillation light detected by PMTs provides signal timing.
- Longest running LArTPC in a neutrino beam.
- Able to detect neutrinos from two beams.
- LArTPC, and so excellent electron-photon separation capabilities.

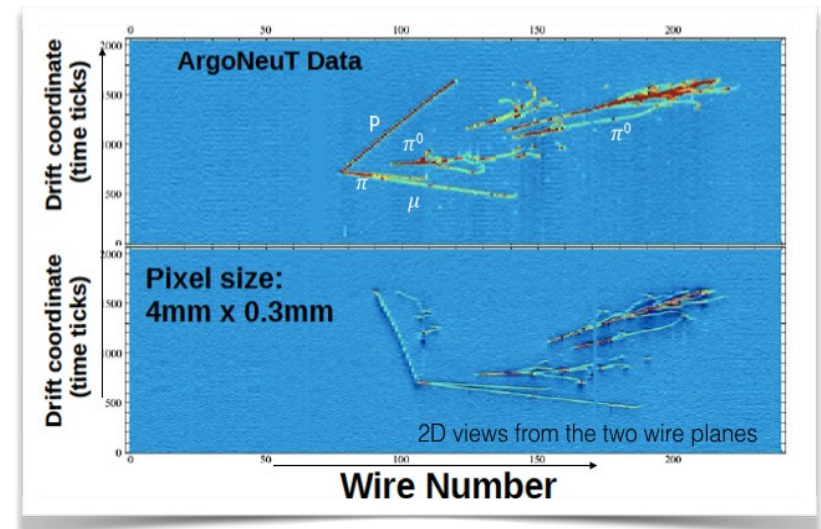


Measuring ν_e in liquid argon



Double handle:
topology and dE/dx

The LArTPC is an excellent tool for electron/photon separation





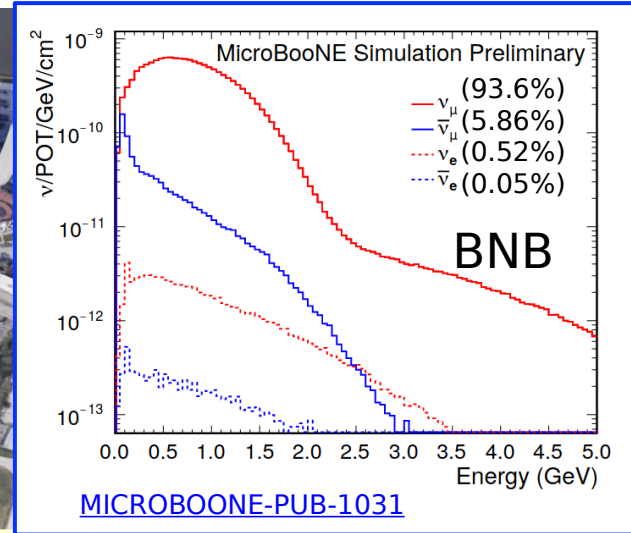
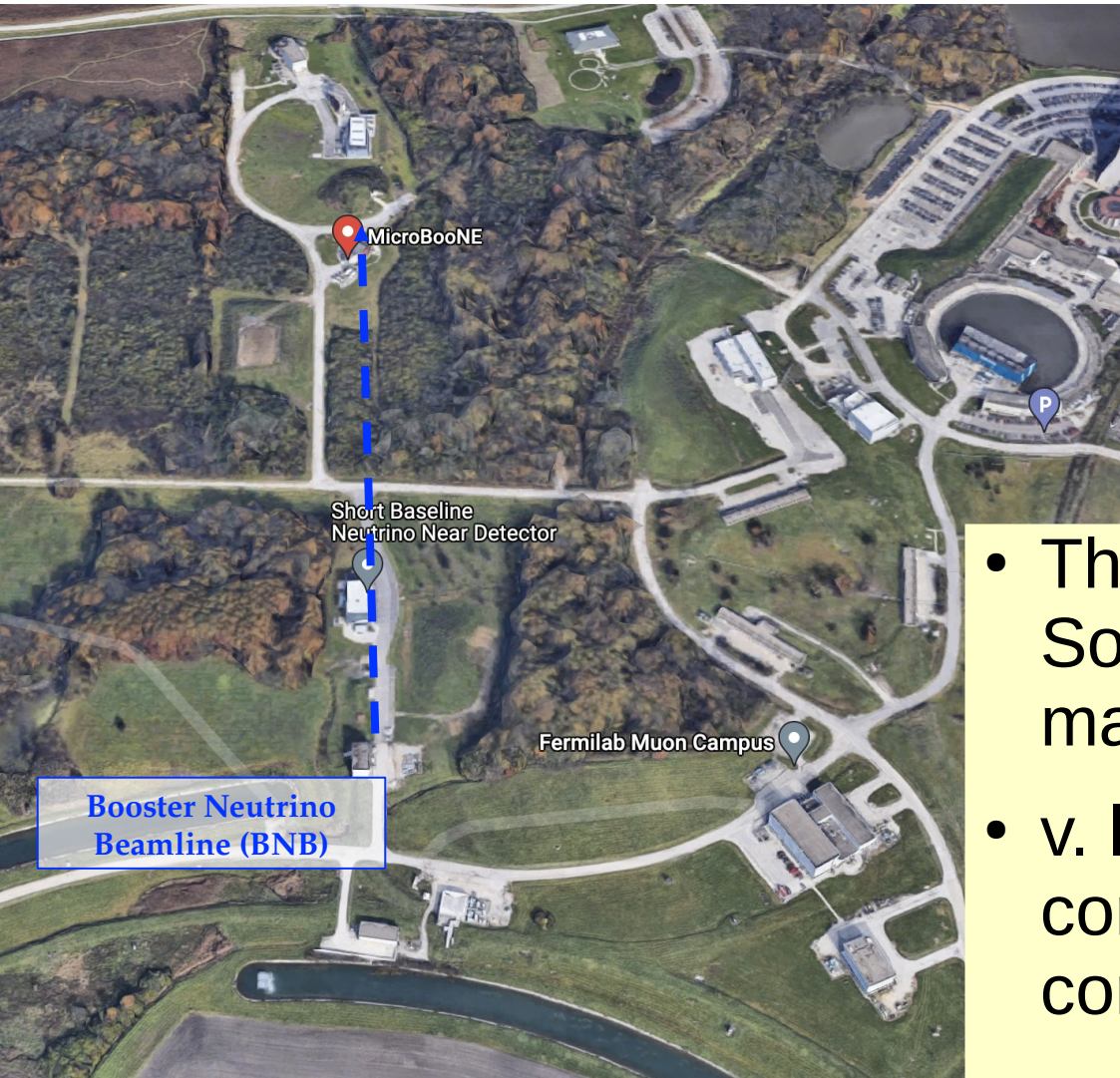
ν_e in MicroBooNE

μ BooNE

23 cm

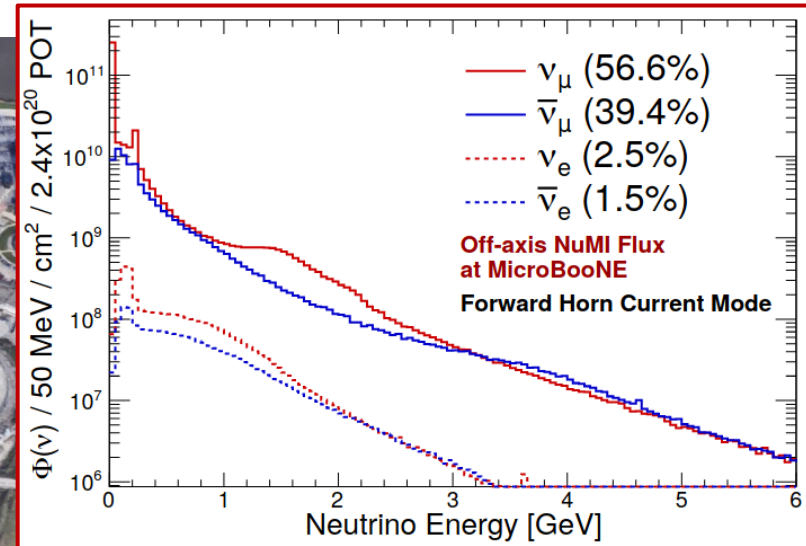
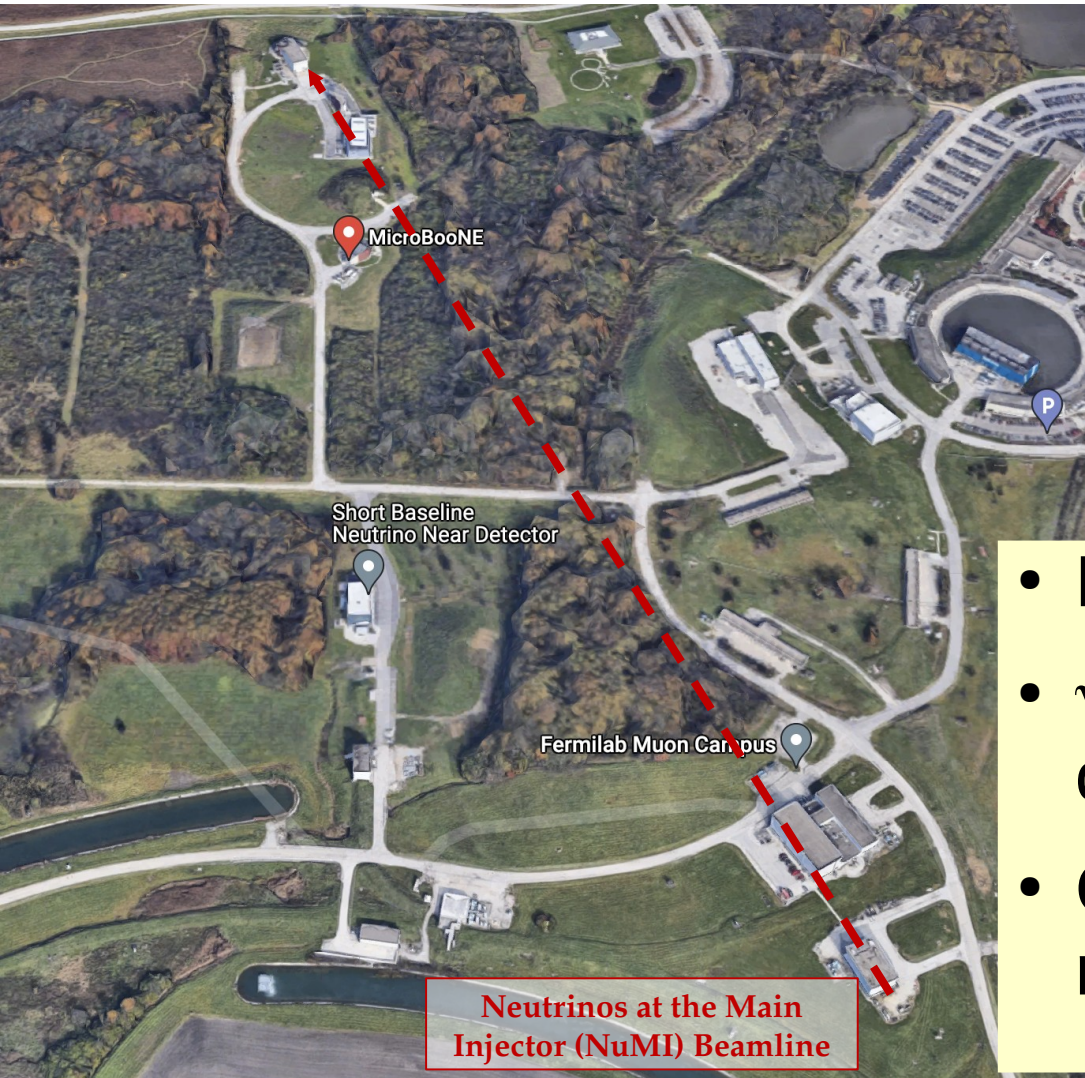
NUMI DATA : RUN 5440 EVENT 2577. MARCH 15, 2016

MicroBooNE and its beams



- The on-axis beam. Source of neutrinos for majority of analyses.
- v. low intrinsic ν_e component (<1%), $\bar{\nu}_e$ component even smaller.

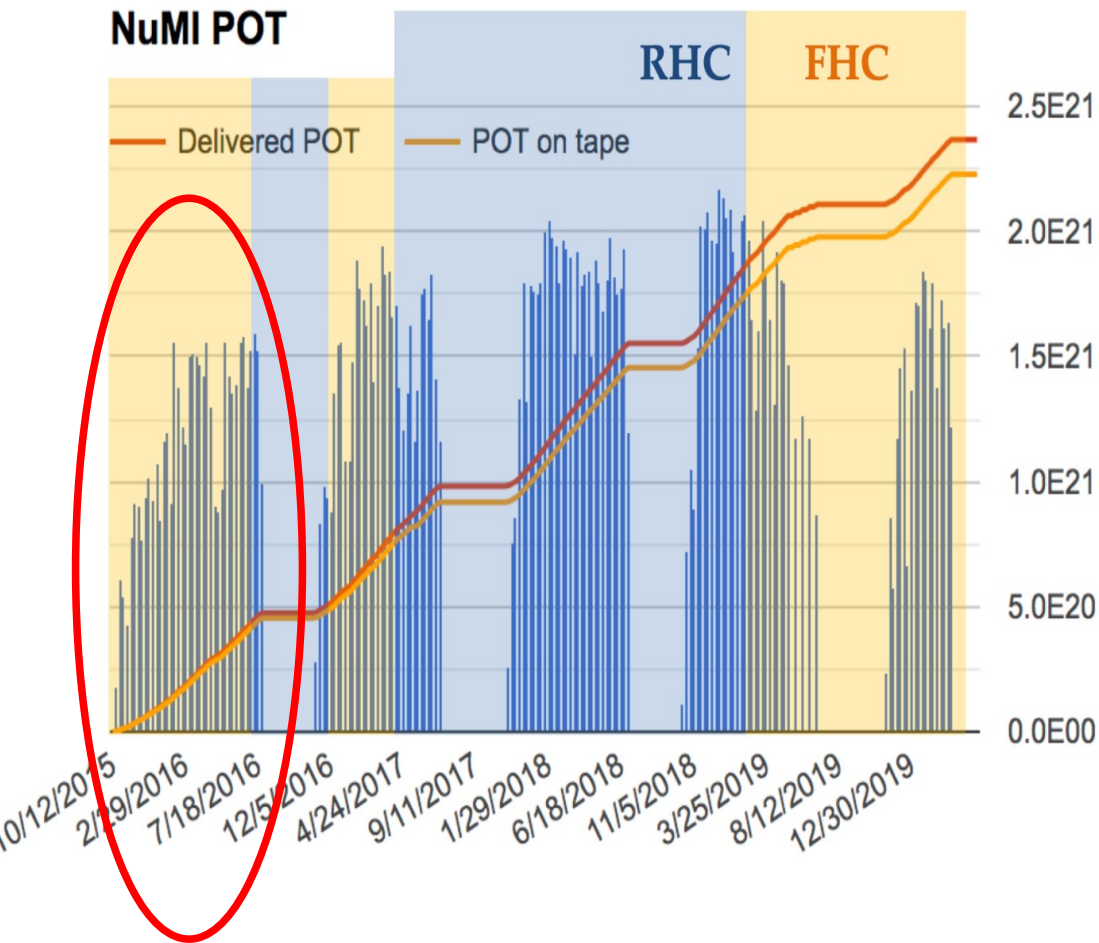
MicroBooNE and its beams (2)



- Highly off-axis (8°).
- ν_e component 5xBNB, $\bar{\nu}_e$ component comparable.
- Great source of electron neutrinos.



NuMI beam in MicroBooNE

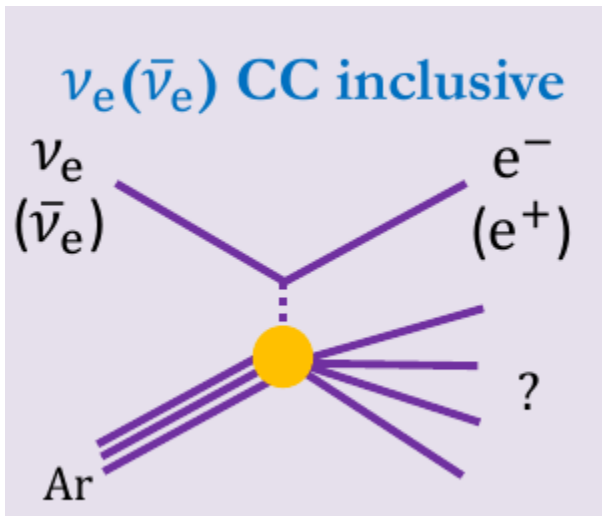


- Total NuMI POT on tape 2.3×10^{21}
- In this talk: uB run 1 (Med. Energy, FHC).
- The Package to Predict FluX (PPFX) constraints hadron production uncertainties

[[Phys.Rev.D](#)
94 (2016) 9, Aliaga et al]

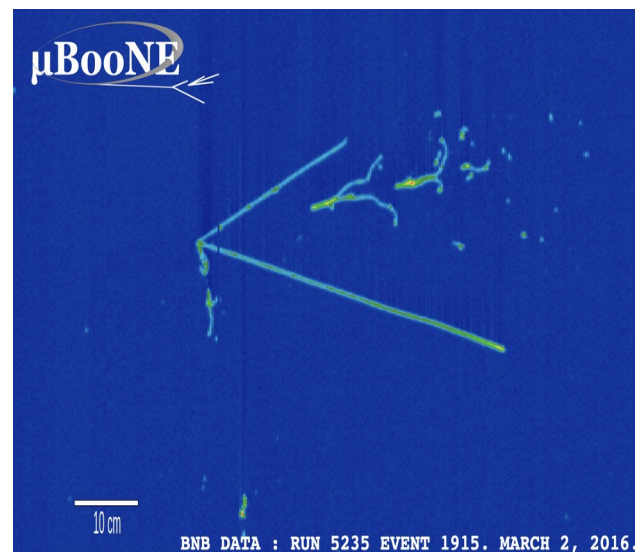
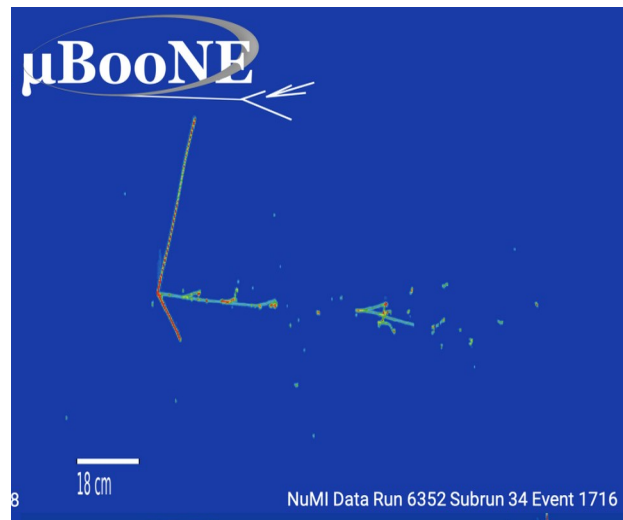
1. Flux Averaged Inclusive $\nu_e + \bar{\nu}_e$ CC

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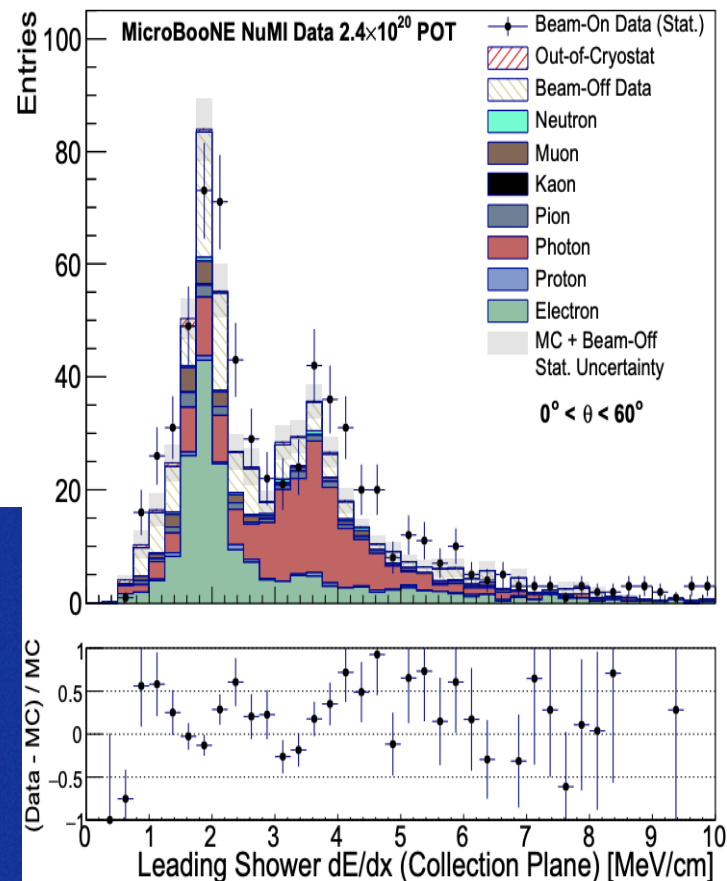
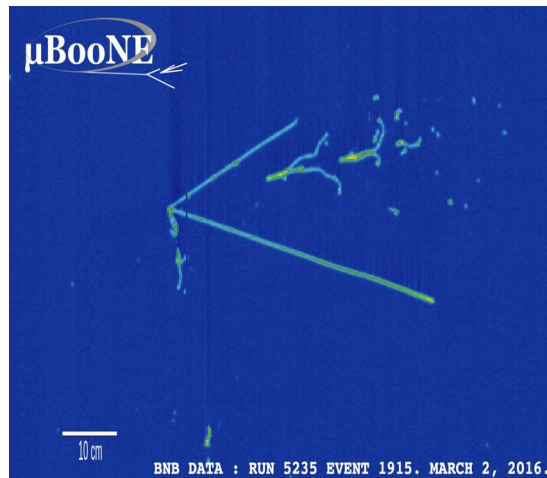
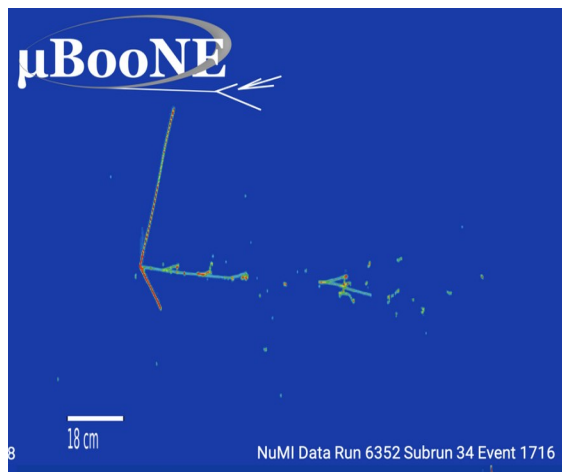
The measurement

- “1st Generation” MicroBooNE measurement
- Beam simulated using FLUGG.
- Using CORSIKA to simulate cosmics.
- Signal: All ν_e and $\bar{\nu}_e$ interactions with energy greater than 250 MeV



Automated e/ γ separation

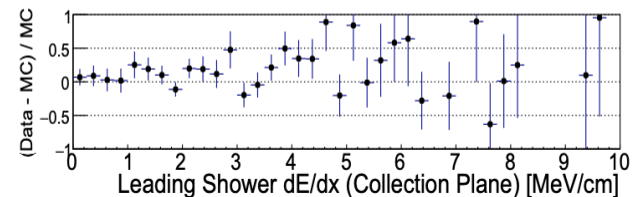
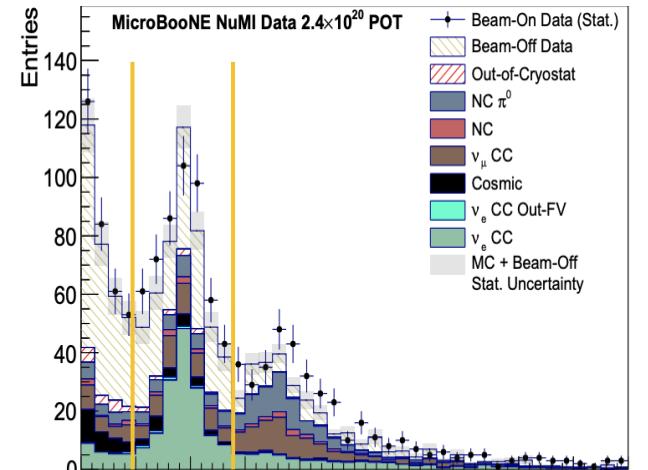
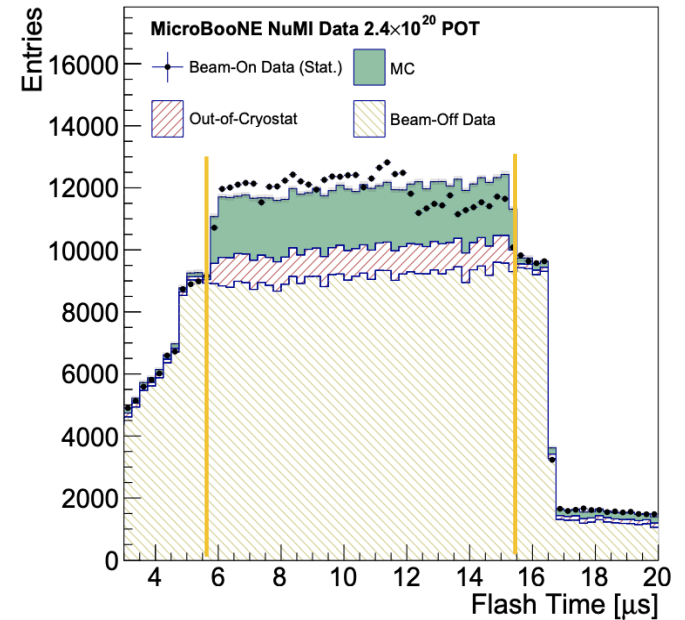
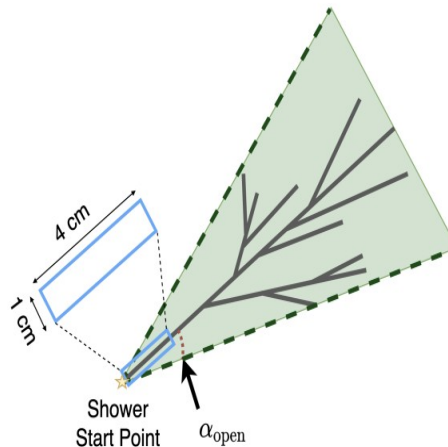
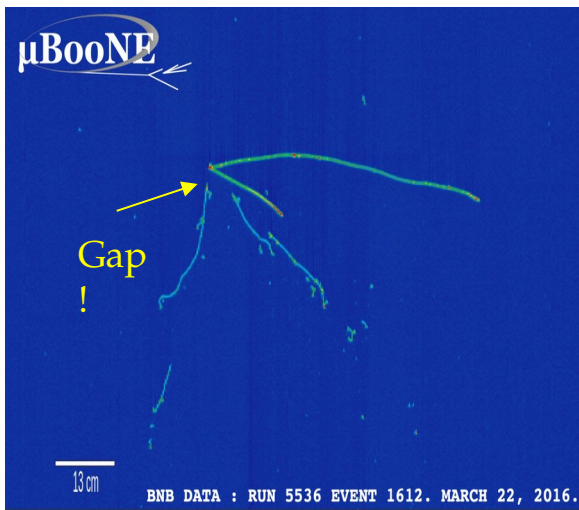
- First demonstration of a fully automated e/ γ discrimination based on shower dE/dx & vertex distance.
- dE/dx observed to provide more discriminating power than vertex distance (analysis specific).



Selection

Selection (main points):

- Require optical beam flash within the beam window
- 1 or more showers
- Vtx-leading shower distance < 4 cm
- Leading shower “physics”:
 - $1.4 \text{ MeV/cm} < \text{trunk } dE/dx < 3 \text{ MeV/cm}$
 - $3^\circ < \text{Opening angle} < 15^\circ$

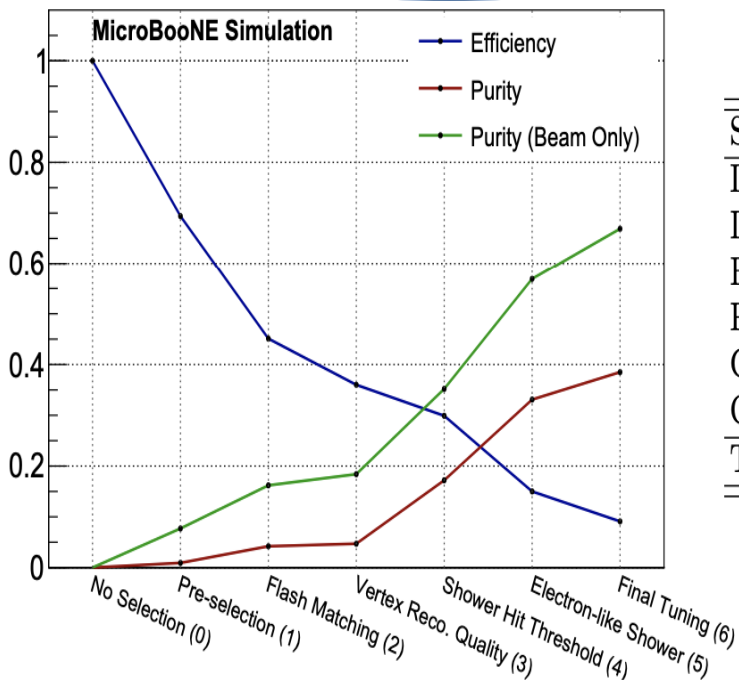




Final Result

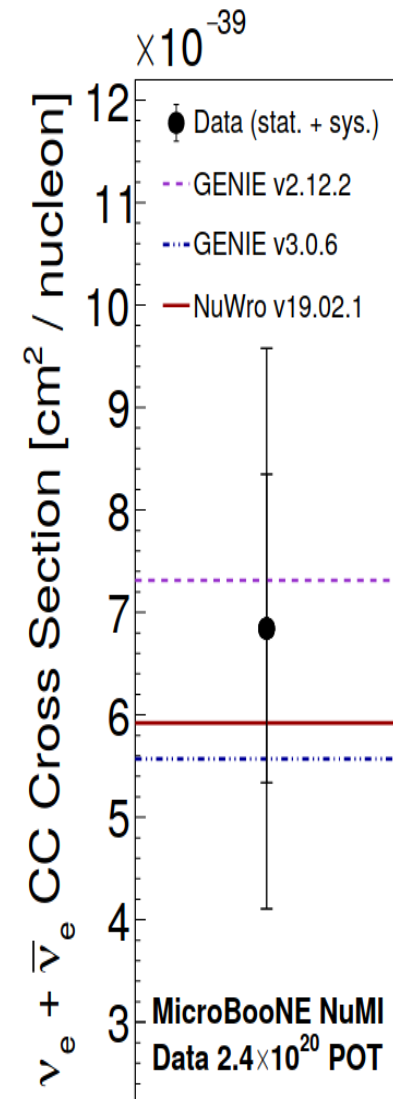
$$\langle \sigma \rangle = 6.84 \pm 1.51 \text{ (stat.)} \pm 2.33 \text{ (sys.)} \times 10^{-39} \frac{\text{cm}^2}{\text{nucleon}}$$

Overall purity: 38%
 Overall efficiency: 9%
 Cosmic ray contamination
 reduced by a factor of 10^5



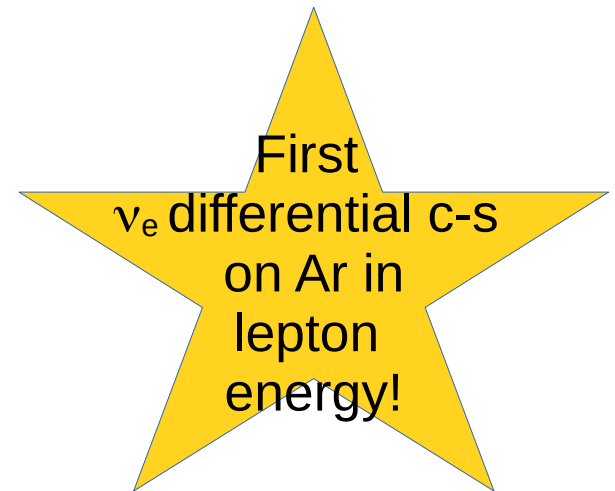
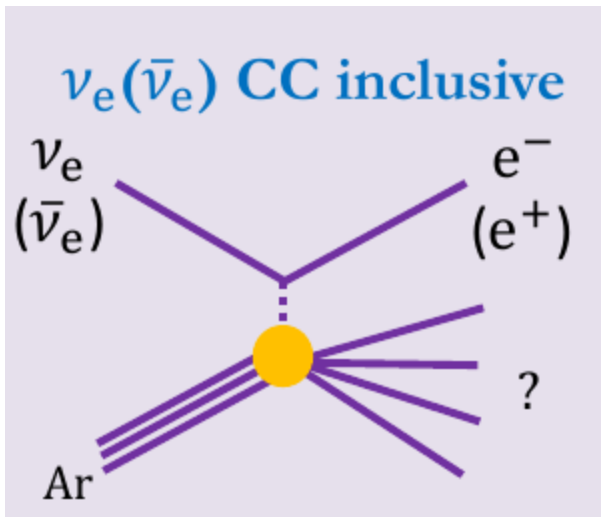
22% statistical uncertainty

Systematic Source	Relative Uncertainty [%]
Interaction	10
Detector Response	23
Beam Flux	22
POT Counting	2
Cosmic Simulation	4
Out-of-Cryostat Simulation	6
Total	34



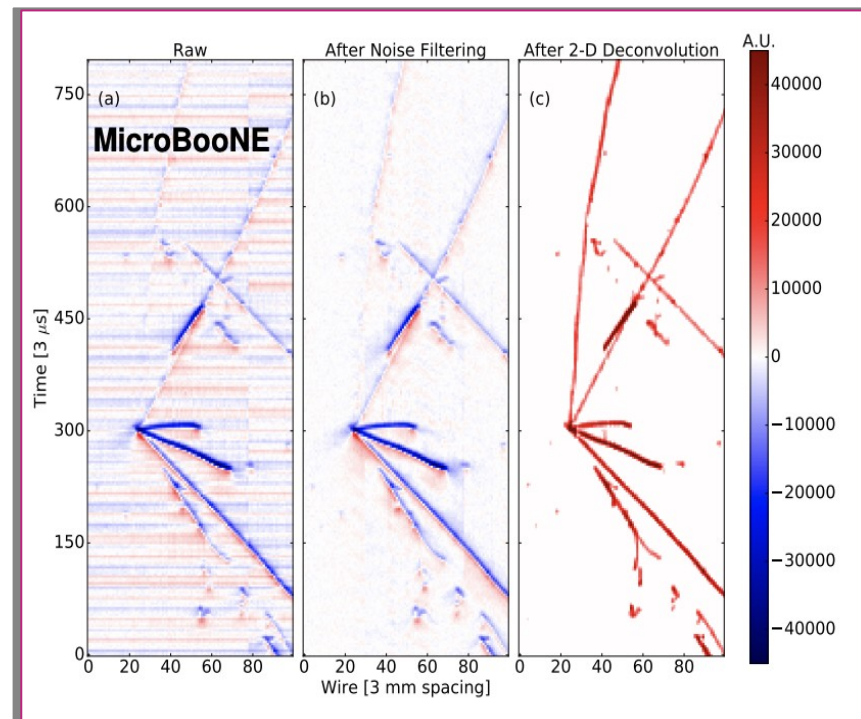
2. Single differential Inclusive $\nu_e + \bar{\nu}_e$ CC

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Improving the analysis

- “2nd generation” MicroBooNE Analysis
- Improved by:
 - Better signal processing and reconstruction, [JINST 13, P07007 \(2018\)](#)
[JINST 13, P07007 \(2018\)](#).
 - Cosmic backgrounds now estimated using overlays.
 - Beam simulation using G4NUMI
 - Using GENIE v.3 tuned to T2K CC0 π data.
 - Improved signal reconstruction.
- Signal: All ν_e and $\bar{\nu}_e$ interactions with energy greater than 60 MeV and charged lepton energy > 120 MeV



New Selection chain

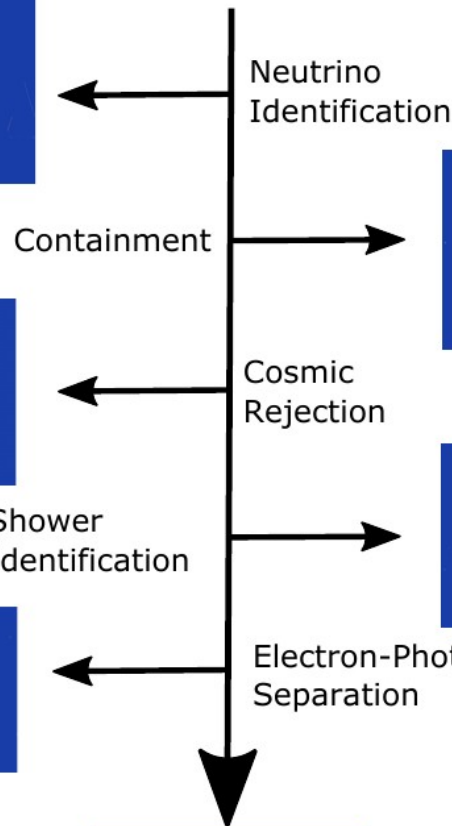
1. Slice ID
2. Shower in Neutrino Slice



5. Topological Score
6. Cosmic IP



9. Moliere Angle
10. 2D: Shower distance, dE/dx
11. dE/dx (no tracks)

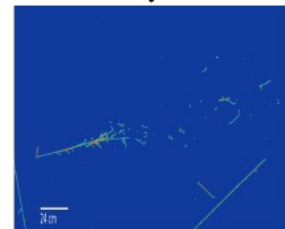


3. Fiducial Volume Boundary
4. Contained Fraction



7. Shower Score
8. Hit Ratio

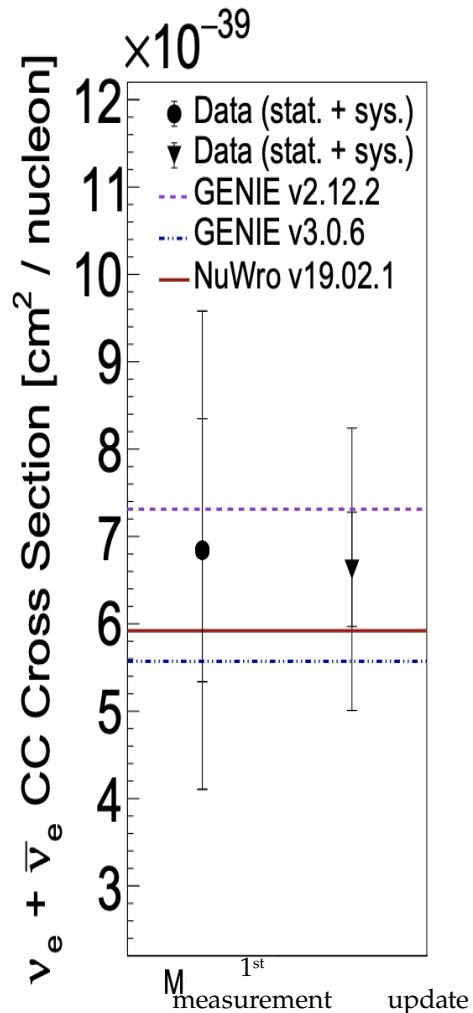
Using tools developed for LEE Pandora analysis



Comparison

After model improvement, reconstruction retuning & update selection:

- Consistent results (CV within 3%),
- halved uncertainties
- Increased purity (38% → 72%)
- Increased efficiency (9 → 21%)



1st measurement

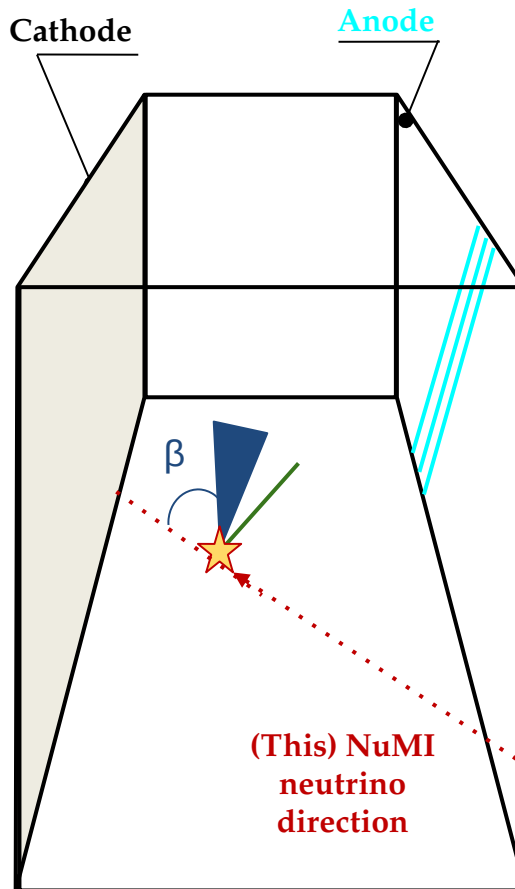
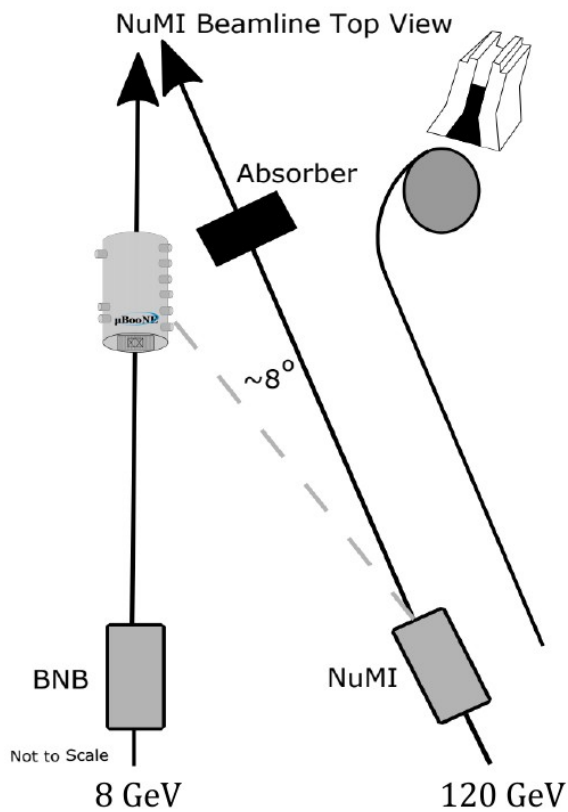
Systematic Source	Relative Uncertainty [%]
Interaction	10
Detector Response	23
Beam Flux	22
POT Counting	2
Cosmic Simulation	4
Out-of-Cryostat Simulation	6
Total	34

Update

Source of Uncertainty	Relative Uncertainty [%]
Beam Flux	17.4
Detector	6.8
Cross Section	5.8
POT Counting	2.0
Out-of-Cryostat	1.8
Proton/Pion Reinteractions	1.2
Beam-off Normalization	0.1
Total Systematic Uncertainty	19.8
MC Statistics	0.8
Data Statistics	10.0
Total Uncertainty	22.2

Differential cross section

- Variables: lepton energy and angle.



Assuming that neutrino comes from the target is correct within 3 degrees from true direction for 95% of selected ν_e

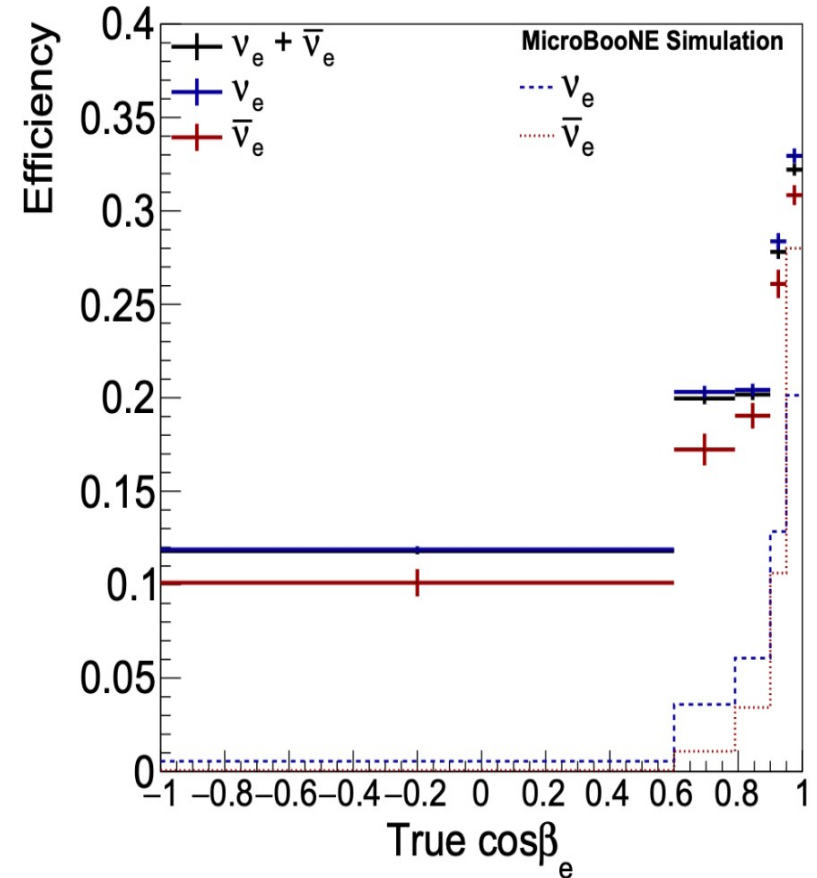
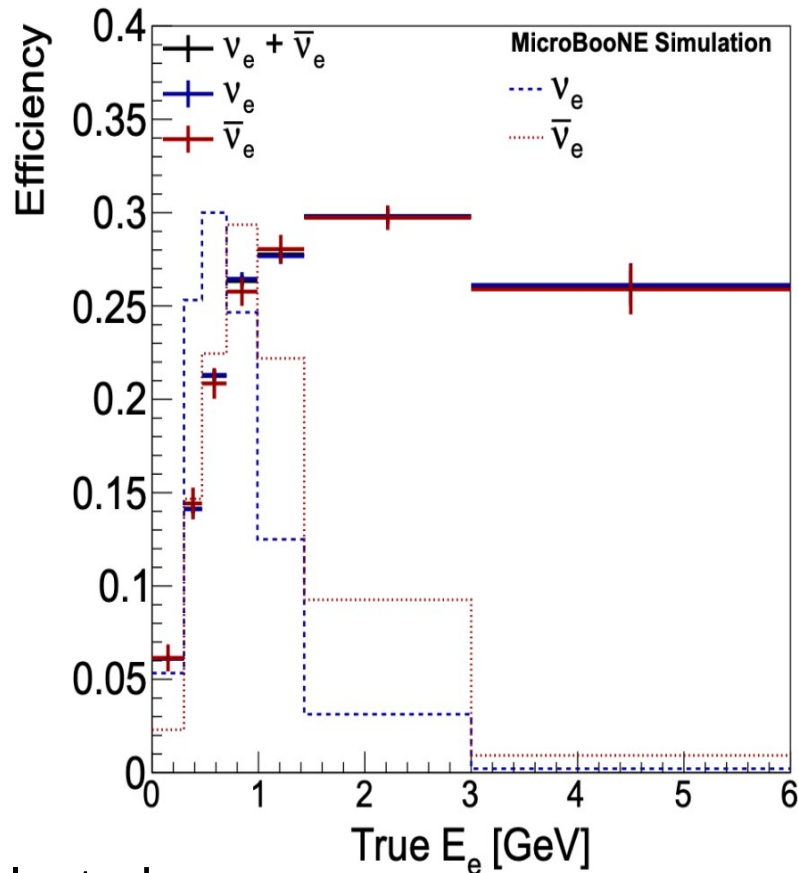
$0.01 < \cos\beta$
resolution < 0.05

NuMI Target





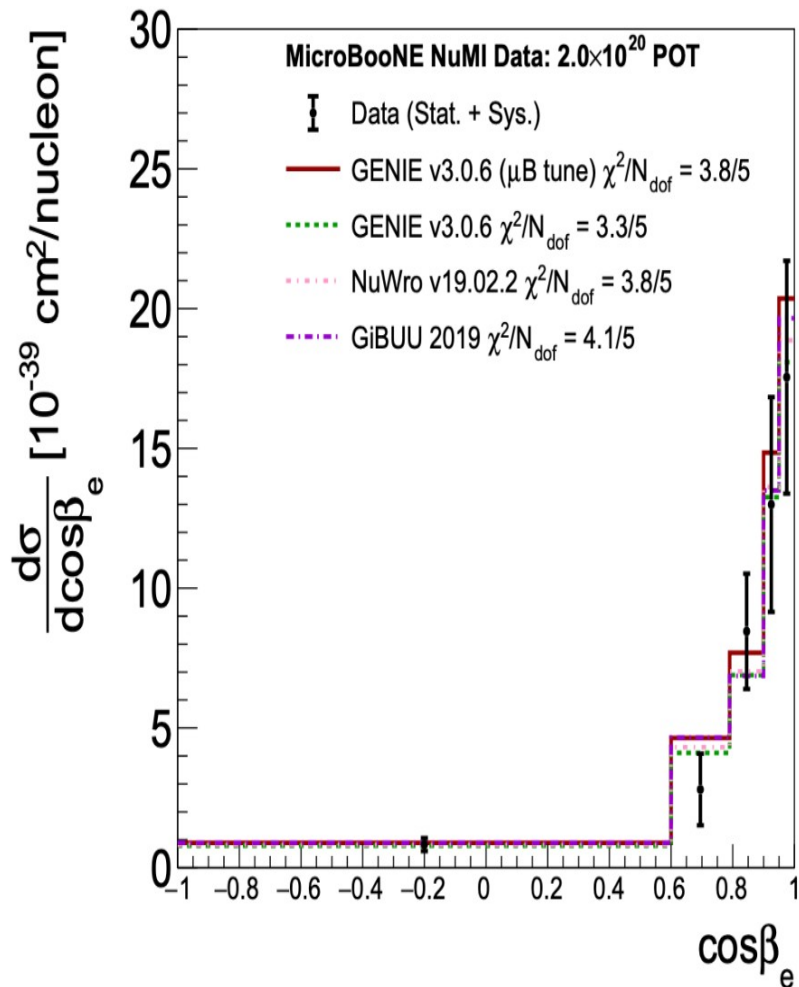
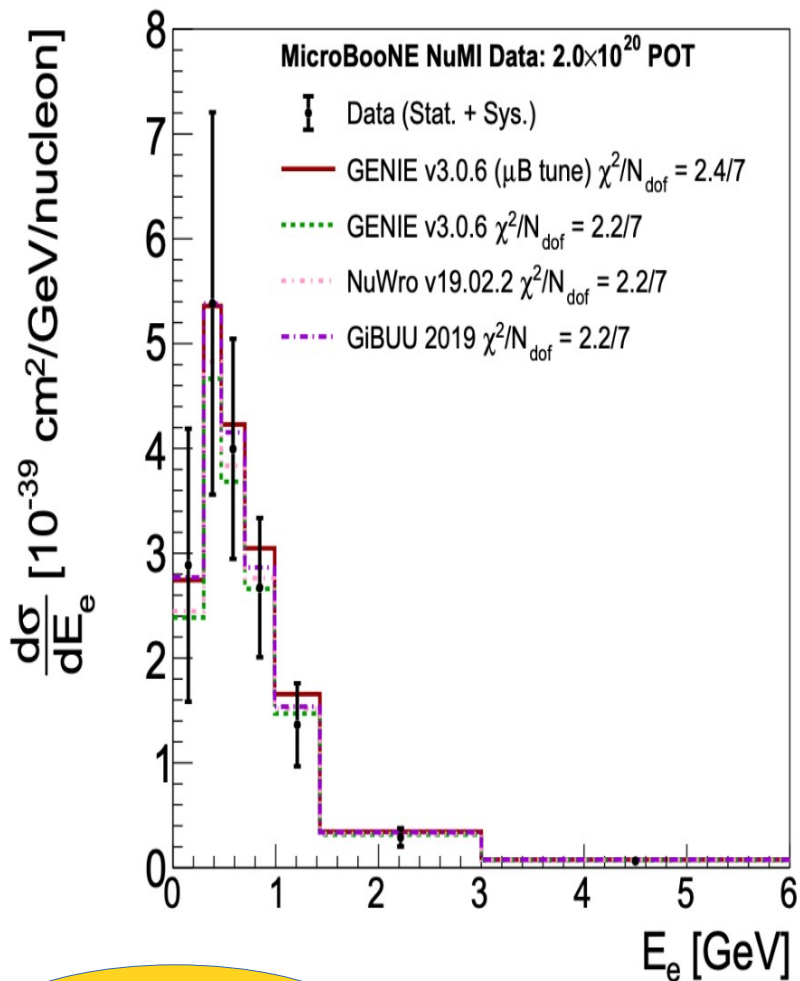
Event Selection and Performance



243 selected candidates.
 purity: 72%
 Avg efficiency: 21%

Main backgrounds:
 → cosmics 8%
 → non- ν_e containing π^0 11%
 → interactions outside FV 7%

Cross Section Results

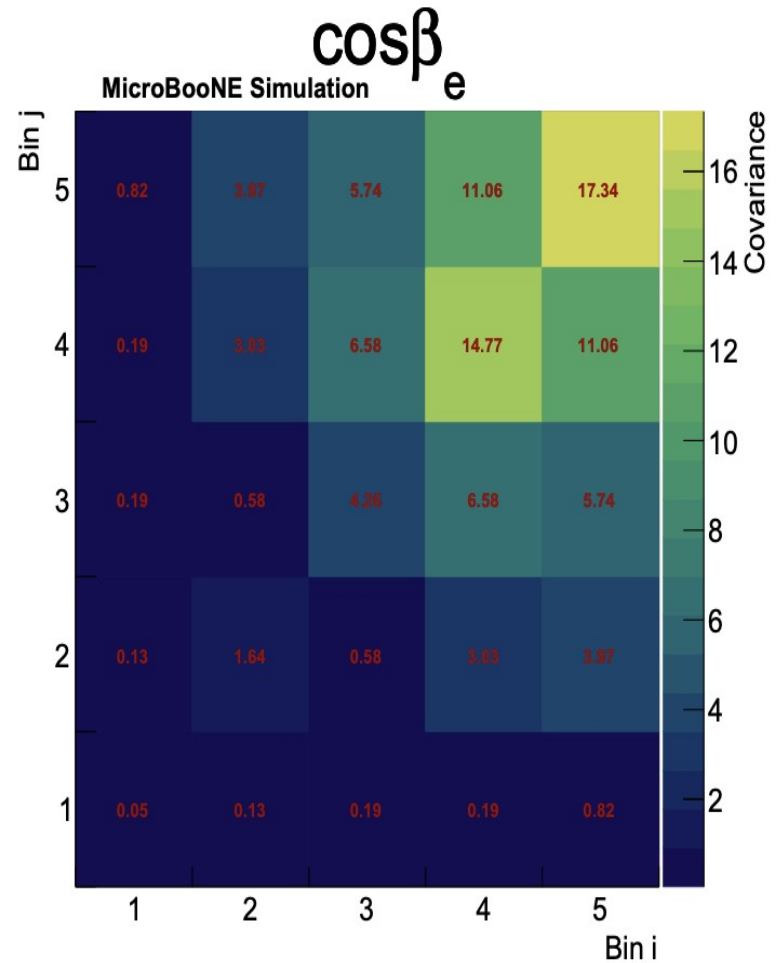
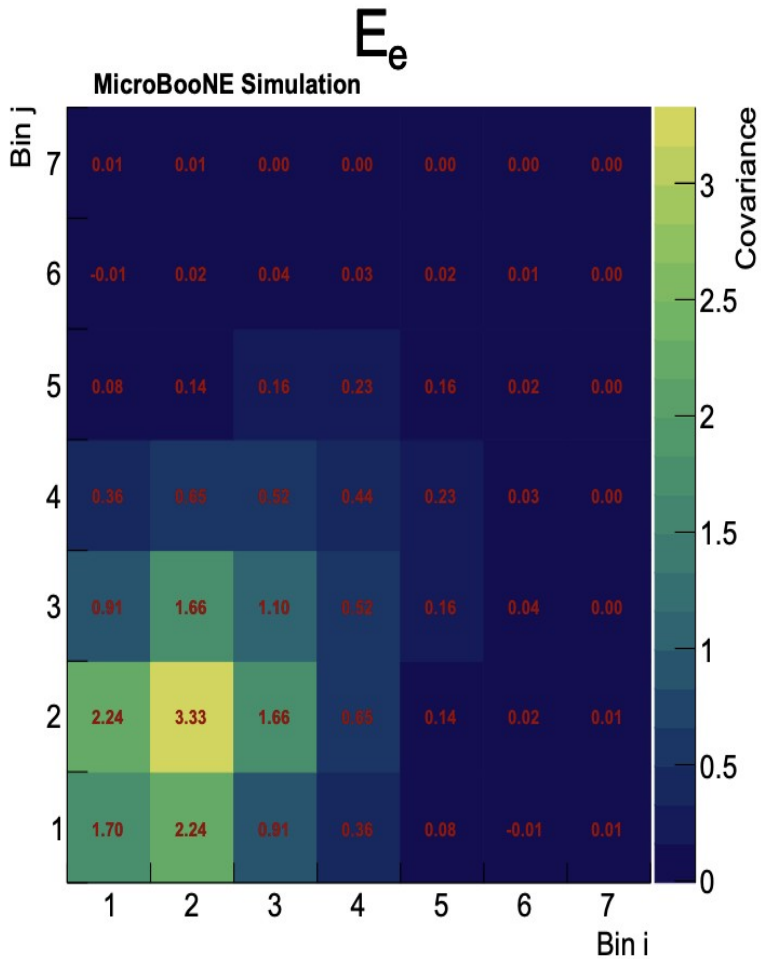


Wiener-SVD
unfolded

Good data – MC agreement



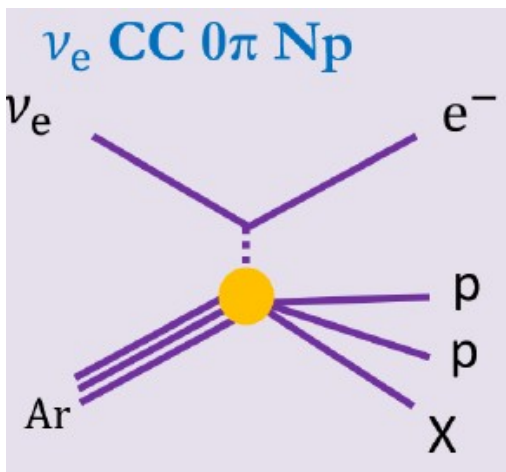
Unfolded data-covariance matrix



supplemental material

3. Differential pionless ν_e CC cross section using BNB

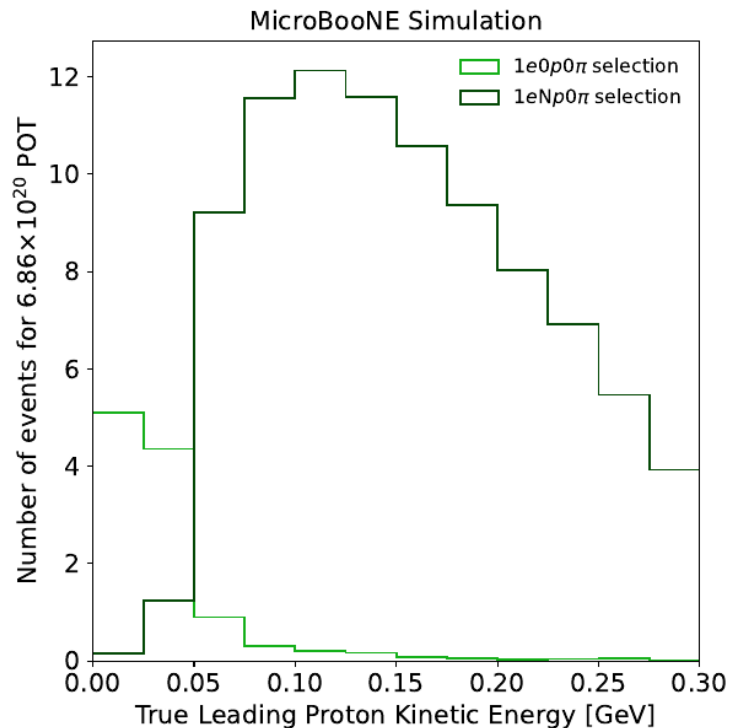
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Measurement

- “2nd generation” MicroBooNE analysis
- Using the BNB and reconstruction framework developed for the LEE.
- Using Run 1-3 data.
- Using overlays for cosmics.
- Signal: CC ν_e interactions with:
 - $KE_e > 30$ MeV, $KE_\pi < 40$ MeV
- Two selection chains depending on presence of protons (i.e. $KE_p > 50$ MeV).
 - For **1e0p0 π** chain phase space restricted to: $\cos\theta_e > 0.6$ & $E_e > 0.51$ GeV to enhance purity.
- ~100+10 events expected (+ ~30+8 background events).



Same Selection chain as NUMI 1d

1. Slice ID
2. Shower in Neutrino Slice



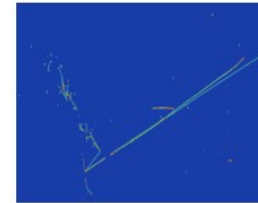
Neutrino Identification



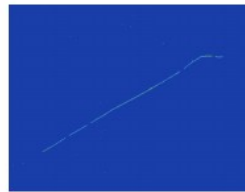
Containment



3. Fiducial Volume Boundary
4. Contained Fraction



5. Topological Score
6. Cosmic IP



Cosmic Rejection



Shower Identification



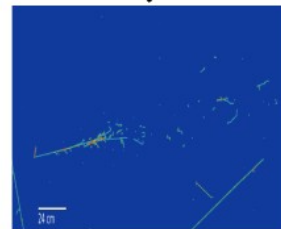
7. Shower Score
8. Hit Ratio



9. Moliere Angle
10. 2D: Shower distance, dE/dx
11. dE/dx (no tracks)



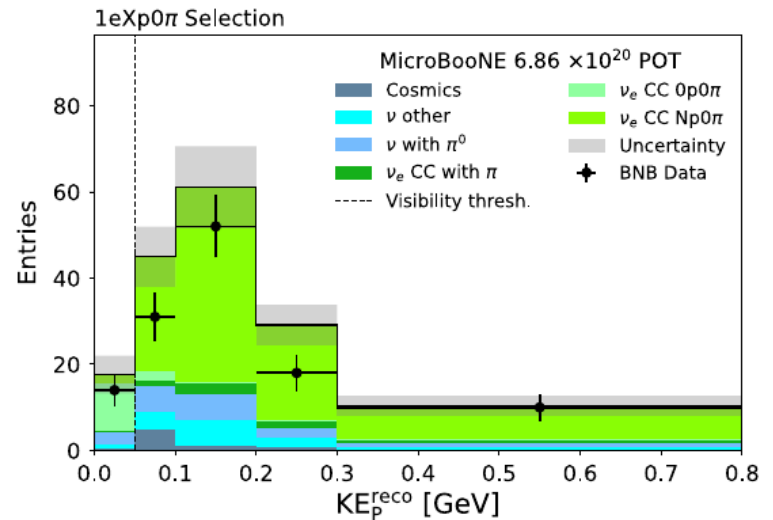
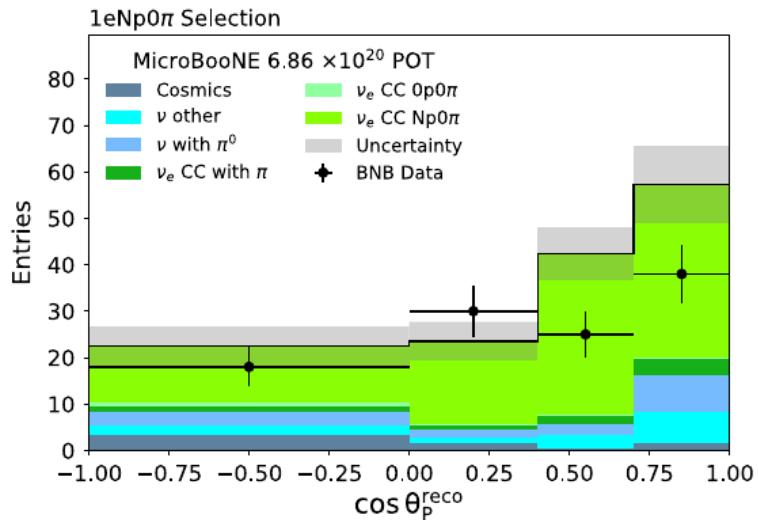
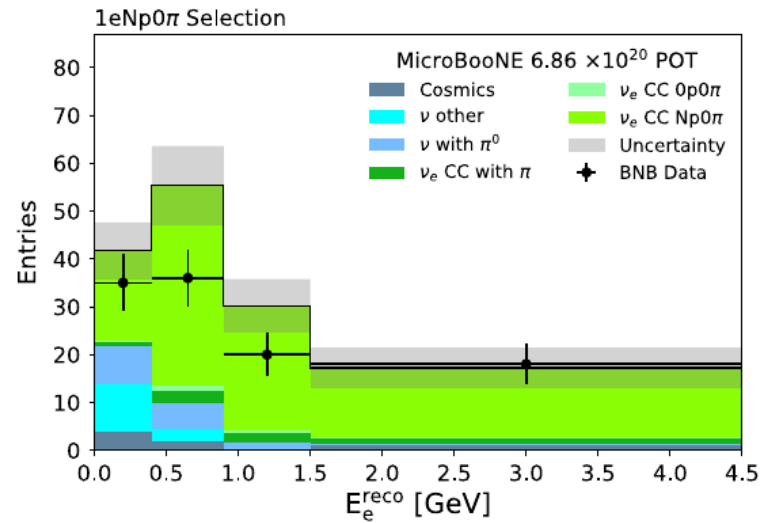
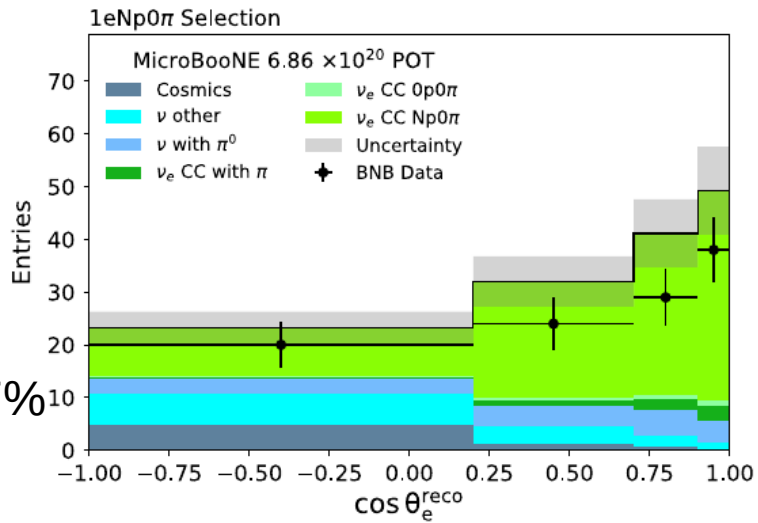
Electron-Photon Separation



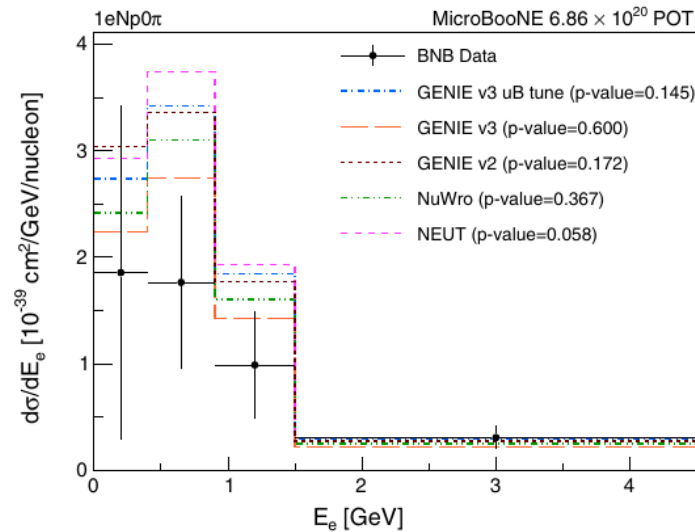
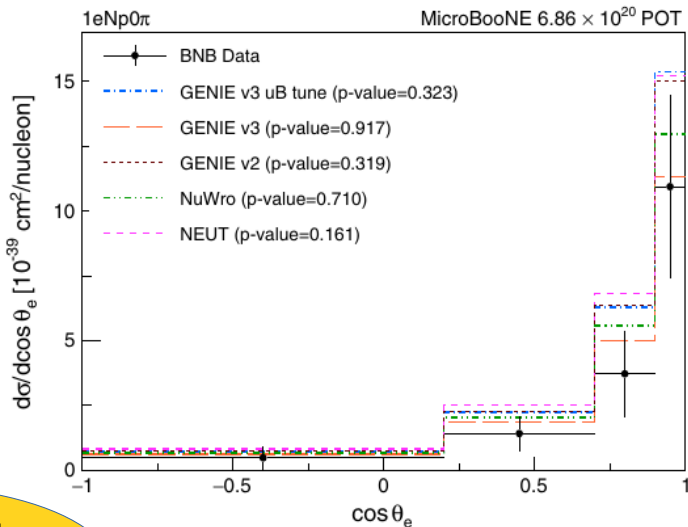


Selection Results

- **1eNp0 π** :
efficiency: 17%
purity: 69%
- **1e0p0 π** :
efficiency: 12%
purity: 65%

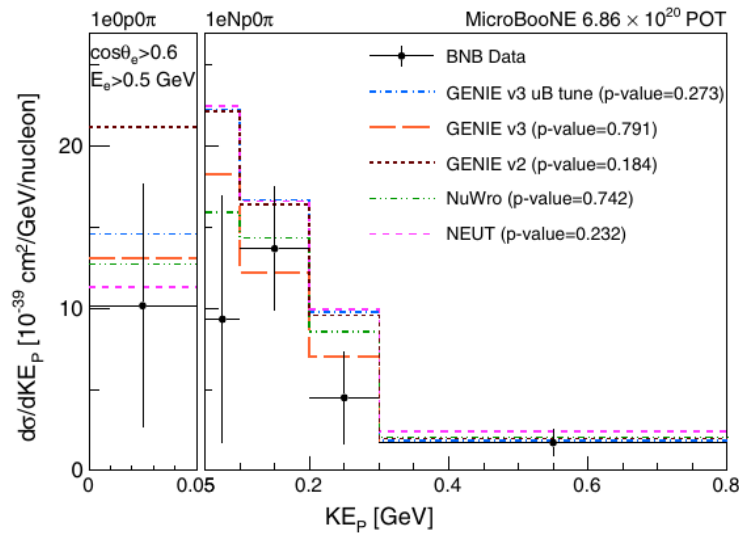
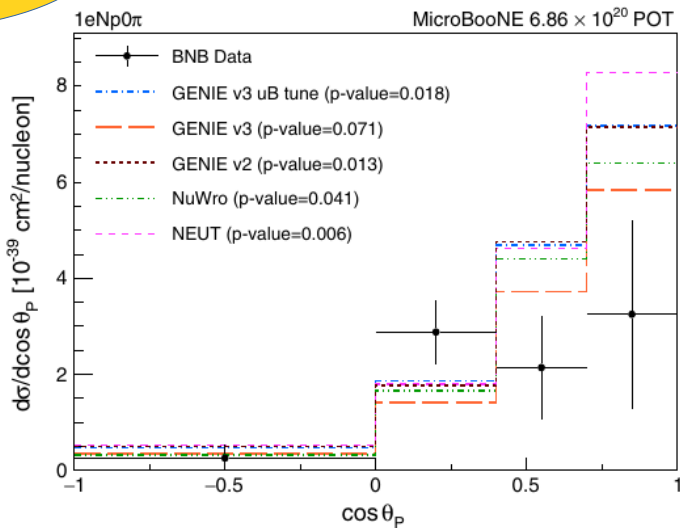


Cross section results



d'Agostini unfolded

Reasonable data – MC agreement



Covariance matrices in supplemental material



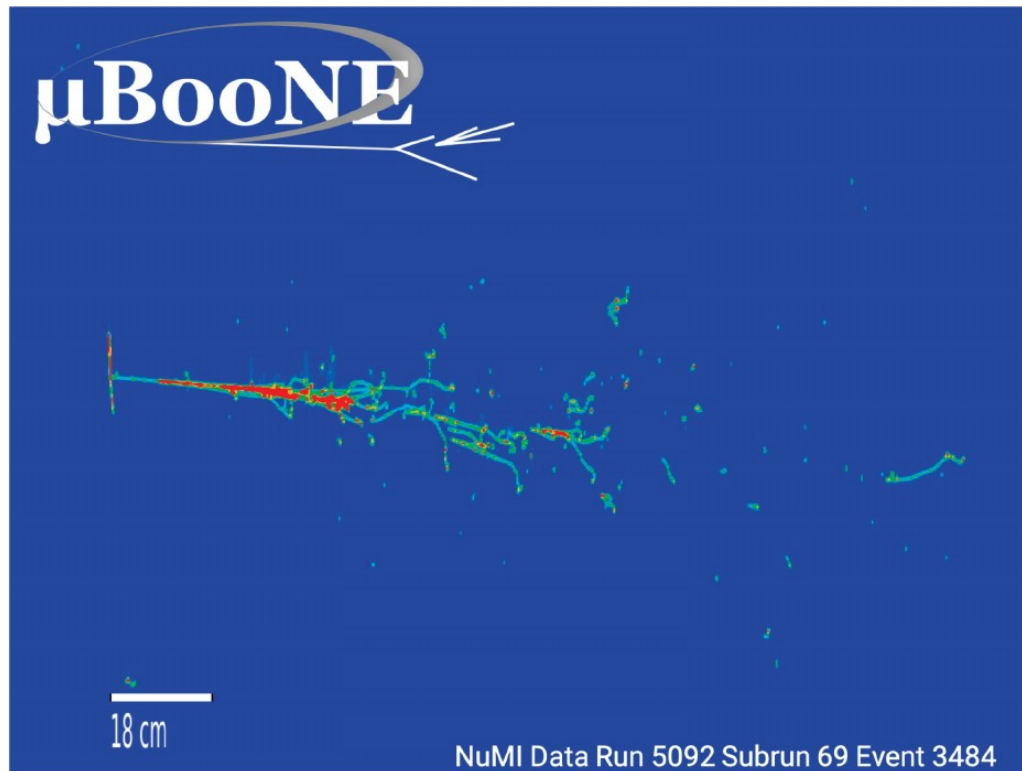
Summary and Outlook

- MicroBooNE is leading the way in electron-neutrino cross sections on argon.
- We have the largest electron-neutrino sample on argon to date.
- 3 measurements since last NuINT.
- More coming soon, including using the whole MicroBooNE data set!



감사합니다

- Upcoming MicroBooNE ν_e analyses:
 - ν_e CC 0π Np (NuMI)
 - $\bar{\nu}_e$ CC Inclusive (NuMI)
 - ν_e/ν_μ ratio (NuMI)
 - Full dataset analyses





Backup slides



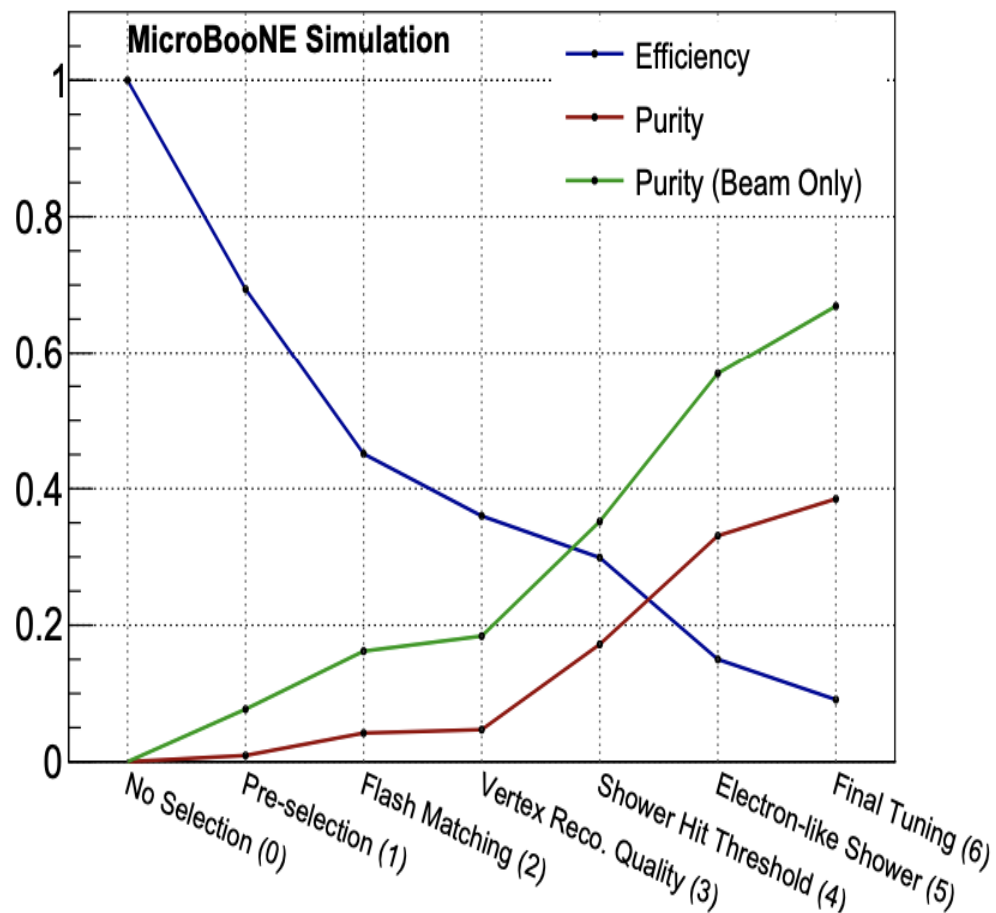
Selection Performance

- Overall purity: 38%

Overall efficiency: 9%

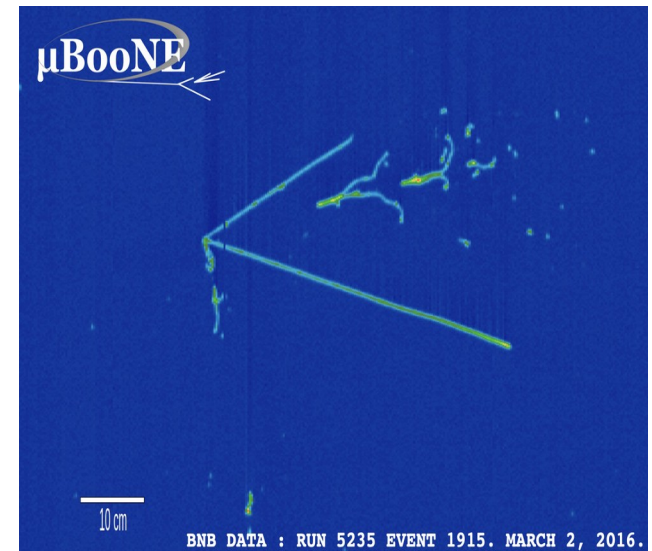
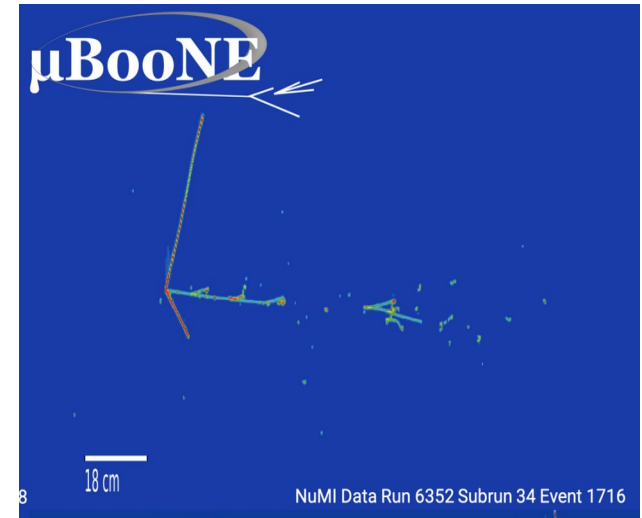
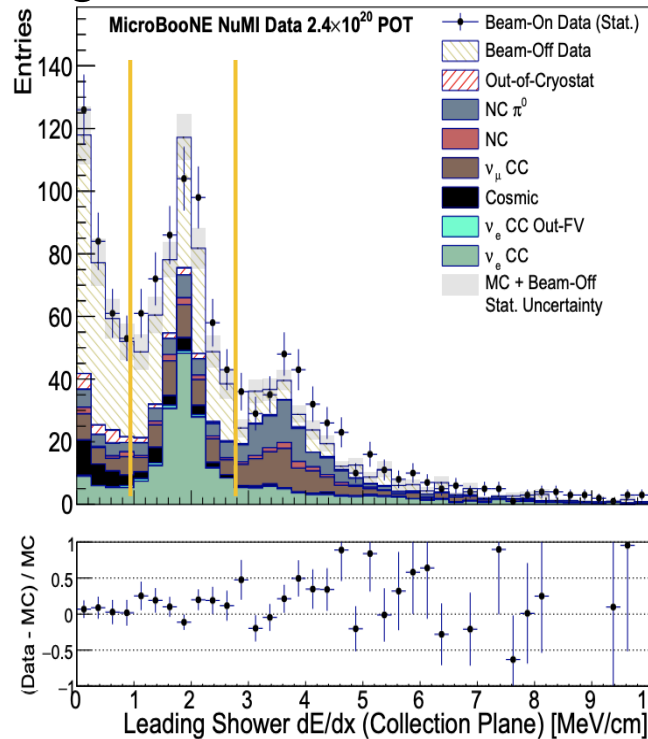
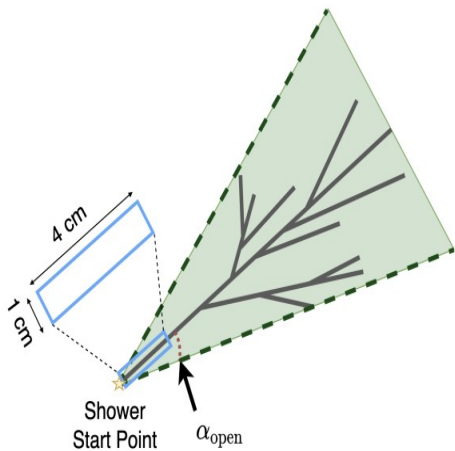
Cosmic ray contamination decreases by a factor of 10^5 wrt initial selection stage... yet, cosmic rays represent the major background contribution to the analysis: 1:1 with signal

→ highlights the need of targeted cosmic ray removal (done in updated analysis)



Selection cont'd

- Leading shower quality: 200+ total & 80+ collection hits
- Leading shower “physics”:
 - $1.4 \text{ MeV/cm} < \text{trunk } dE/dx < 3 \text{ MeV/cm}$
 - $3^\circ < \text{Opening angle} < 15^\circ$



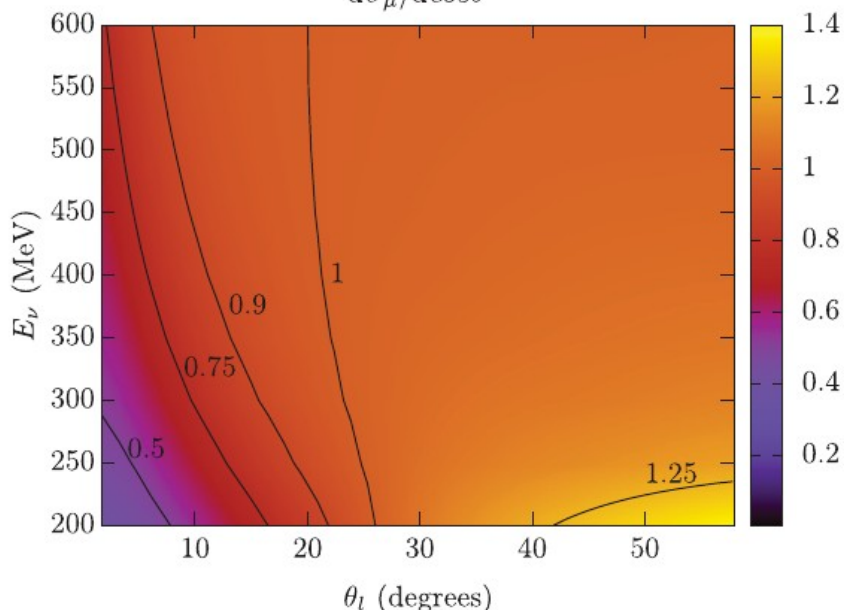
Why measure ν_e cross sections?

Many ν_μ measurements exist – much easier to produce ν_μ at accelerator energies.

Can't we use lepton universality and be done with it?

Yes, but ...

$$\frac{d\sigma_e/d\cos\theta}{d\sigma_\mu/d\cos\theta}$$



- Flavour-dependent effects due to the different lepton mass and uncertainties in nucleon form factors can be up to 5-10%:
[Phys.Rev.D 86 \(2012\) 053003](#),
[PhysRevLett.123.052501](#), [arXiv:2105.07939](#)
<https://arxiv.org/abs/2204.11379>]
- Radiative part of these effects is not implemented in ν -generators.
- Direct measurements of ν_e cross sections could provide tests of lepton universality.
- Testing ground for ν_e reconstruction.