## High-energy neutrino measurements with FASER $\nu$ at the LHC

# NuINT 2022

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### MONSJ UNDATIONC



## Introduction



- Large Hadron Collider (LHC): 27 km ring collider, <u>13.6 TeV</u> proton-proton collisions
- Energetic particles produced in the <u>far-forward direction</u> of the collisions
- FASER(ForwArd Search ExpeRiment) is a new experiment at the LHC to search for new, light, long-lived particles (LLPs) and study neutrinos

**FASER** $\nu$ : neutrino program of the FASER experiment

- Three flavor neutrino cross section measurements at TeV energy ranges





- Muon charge identification by FASER spectrometer with 0.55 T magnets ( $\nu_{\mu}$ )



## Neutrino Interaction Events (MC)



- Generator: GENIE
- MC simulation with FASER simulation framework based on Geant4
- Reconstructed particle hits in emulsion detector

### framework based on Geant4 n detector





Expected CC interaction events (250 fb<sup>-1</sup>)

Gen	erators	$FASER\nu$			
light hadrons	heavy hadrons	$\nu_e + \bar{\nu}_e$	$ u_{\mu} + ar{ u}_{\mu}$	$ u_{ au} + ar{ u}_{ au}$	
SIBYLL	BYLL SIBYLL		7971	24.5	
DPMJET	MJET DPMJET		11813	161	
EPOSLHC	Pythia8 (Hard)	2521	9841	57	
QGSJET	Pythia8 (Soft)	1616	8918	26.8	
Combin	ation (all)	$2850^{+2910}_{-1348}$	$9636^{+2176}_{-1663}$	$67.5^{+94}_{-43}$	
Combination	(w/o DPMJET)	$1880^{+641}_{-378}$	$8910^{+930}_{-938}$	$36^{+20.8}_{-11.5}$	

 Discrepancy between generators for charm production

 ~10,000 ν interactions expected in LHC Run 3 (2022-2025)



## **Cross-Section Sensitivity**





(150 fb<sup>-1</sup>)

Three flavors neutrino cross-section measurements for <u>unexplored</u> energy ranges Neutrino energy reconstruction with resolution 30% expected from simulation studies





- Statistical significance:  $2.7 \sigma$  from null hypothesis
- Phys. Rev. D 104, L091101









### Assembly Film production





## Film Production





- 200 nm diameter crystals
- Double sided emulsion coating
- Total area of 730 films: ~55 m<sup>2</sup>
- Produce emulsion gel and film a few months before module assembly







- Screening of tungsten target plates
  - Require thickness difference < 80  $\mu$ m in each plate (1562/1622 plates)
- Sub-module: vacuum packed 10 films + 10 tungsten plates Apply external force (equivalent to 1 bar) to 73 sub-modules
- in the FASER  $\nu$  box





## Installation and Data-taking



- In order to keep the track multiplicity in the detector manageable, the detector is exchanged during LHC Technical Shutdowns (3 times per year)
- The measured multiplicity corresponds to ~2.3×10<sup>4</sup> cm<sup>-2</sup>/fb<sup>-1</sup>

- Temperature monitoring
  - ~0.1 °C/1 month variation will not affect to the position alignment of the emulsion films

## Development





- Installed new development chains
- 10-12 days 1

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Developed 2





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### : CERN darkroom facility



## Readout







25 cm

Transport films to Japan after development

- Readout by Hyper Track Selector-1 (HTS-1)
- Field of view: 5.1 mm × 5.1 mm
- ~3 months to readout 1 FASER  $\nu$  module (5 hours/day)
- Fast enough to readout in parallel with data taking



30 cm









		Integrated luminosity per module (/fb)	N $\nu$ int. expected
2022 1 <sup>st</sup> emulsion	Mar 15 – Jul 26	0.5	25 ×29%
2022 2 <sup>nd</sup> emulsion	Jul 26 – Sep 13	10.6	530
2022 3 <sup>rd</sup> emulsion	Sep 13 – Nov 29	(~20)	(~1000)



### LHC Run 3 stable collisions

Junoct	Jul	Aug	Se	p	Oct	Nov	Dec
3 m²)		2 <sup>nd</sup> (55 r	m²)		3 <sup>rd</sup> (55	m²)	

- Assembled 3 modules in 2022
- The 2<sup>nd</sup> module should have enough statistics to observe neutrinos!





- Installed the 1<sup>st</sup> FASER  $\nu$  module on March for detector performanc
  - Integrated luminosity: 0.5 fb<sup>-1</sup>
  - 21 sub-modules (29% of full detector)
  - 2 special sub-modules for alignment study
- Performed development on 30<sup>th</sup>-31<sup>st</sup> of July. Film scan is ongoing
- Reconstructed tracks in 20 emulsion films look very good

• Successfully performed the whole sequence of the 1st FASER  $\nu$  detector





## **Position Accuracy**



- Dataset: most downstream 10 emulsion films of the 1<sup>st</sup> FASER  $\nu$  module
- Observed ~0.2  $\mu$ m position accuracy with dedicated alignment using high momentum muon tracks

Position deviation between hits and the straight-line fits to the reconstructed tracks



### Angular Distributions 1<sup>st</sup> module (2022)



In-situ measurement (2018)











- Observed large angle particles
  - Cosmic rays
  - Backward tracks from the beam line
- Observed double-peak structure
  - Consistent with particles arriving from the LHC beam line in the vertical plane
  - The origin of the structure is under investigation with simulation studies

### Pilot run (2018)









### Angular Distributions 1<sup>st</sup> module (2022)



The angular spread of the projected peaks is approximately 0.5 mrad due to MCS through 100 meters of rock

(Angular resolution is expected to be ~0.1 mrad with 10 films)





## $2^{nd}/3^{rd}$ FASER $\nu$ Module



- Installed the 2<sup>nd</sup> module on 25<sup>th</sup> of July
  - Data-taking and development were successfully done
  - Physics analysis in parallel with film scan
  - Looking forward to observing neutrinos! -



- Installed the 3<sup>rd</sup> module on 14<sup>th</sup> of September
  - Development will be done by the end of this year





## Conclusions

- FASER  $\nu$  studies three flavor neutrinos at the high energy frontier
  - $~10,000 \nu$  interactions in LHC Run 3 (2022-2025, 250 fb<sup>-1</sup>)
- Data-taking is on track
  - 3 FASER  $\nu$  modules were assembled and installed this year
  - Performed development of the 2 modules. Film scan is being done
- Successfully performed the whole sequence of the 1<sup>st</sup> FASER  $\nu$  detector
  - $\sim 0.2 \ \mu$ m position accuracy observed on the 10 emulsion films
  - First results with the LHC beam show an excellent performance of the FASER  $\nu$  detector
- Physics analysis has began towards the first physics results
- part of the proposed Forward Physics Facility (FPF)
  - Ref: <u>https://arxiv.org/abs/2203.05090</u>

Starting to discuss a larger, O(10 tonne), FASER  $\nu$  2 detector for the HL-LHC, as









## Collaboration



























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The University of Manchester

UNIVERSITY OF SUSSEX







## Neutrino Production



passing through the detector











## Neutrino Event Rates

Detector			Number of CC Interactions			
Name	Mass	Coverage	Luminosity	$\nu_e + \bar{\nu}_e$	$ u_{\mu}\!+\!ar{ u}_{\mu}$	$ u_{ au} + ar{ u}_{ au} $
$FASER\nu$	1 ton	$\eta\gtrsim 8.5$	$150 {\rm ~fb^{-1}}$	901 / 3.4k	4.7k / 7.1k	15 / 97
SND@LHC	800kg	$7 < \eta < 8.5$	$150 {\rm ~fb^{-1}}$	137 / 395	790 / 1.0k	7.6 / 18.6
$FASER\nu 2$	20  tons	$\eta\gtrsim 8.5$	$3 \mathrm{~ab^{-1}}$	178k / 668k	943k / 1.4M	2.3k / 20k
FLArE	10 tons	$\eta\gtrsim7.5$	$3 \mathrm{~ab^{-1}}$	36k / 113k	203k / 268k	1.5k / 4k
AdvSND	$2  ext{ tons}$	$7.2 \lesssim \eta \lesssim 9.2$	$3 \mathrm{~ab^{-1}}$	6.5k / 20k	41k / 53k	190 / 754

Table 7.1: Detectors and neutrino event rates: The left side of the table summarizes the detector specifications in terms of the target mass, pseudorapidity coverage and assumed integrated luminosity for both the LHC neutrino experiments operating during Run 3 of the LHC as well as the proposed FPF neutrino experiments. On the right, we show the number of charged current neutrino interactions occurring the detector volume for all three neutrino flavors as obtained using two different event generators, Sibyll 2.3d and DPMJet 3.2017.













## New Physics

## constrain SM EFT coefficients



SM neutrino oscillations are expected to be negligible at FASERv. However, sterile neutrinos with mass ~40eV can cause oscillations. FASERv could act as a short-baseline neutrino experiment.

![](_page_26_Figure_5.jpeg)

![](_page_26_Picture_6.jpeg)

![](_page_27_Picture_0.jpeg)

![](_page_27_Figure_1.jpeg)

![](_page_27_Picture_2.jpeg)

## **Input for astroparticle experiments**

Studies of high-energy astrophysical neutrinos with large-scale neutrino telescopes (e.g. IceCube), suffer from backgrounds from atmospheric neutrinos from charm-decay (charm produced in hadronic shower initiated by cosmic rays hitting the atmosphere).

At ultra high-energy light hadrons travel far through the atmosphere, losing energy, and hence produce lower energy neutrinos. Neutrinos produced in charm decay ("prompt neutrinos") are therefore the key background at high energy. This prompt background has a large associated uncertainty which limits the study of astrophysical neutrinos. Measurements of neutrinos from charm at the FPF can provide important information to constrain this background.

![](_page_28_Figure_3.jpeg)

![](_page_28_Picture_5.jpeg)

## Physics

![](_page_29_Figure_1.jpeg)

![](_page_29_Figure_2.jpeg)

![](_page_29_Figure_3.jpeg)

![](_page_29_Picture_4.jpeg)

## **Detector Environment**

### FLUKA simulations

![](_page_30_Figure_2.jpeg)

![](_page_30_Figure_3.jpeg)

- Muon flux simulations/measurements
  - MC prediction is in good agreement with data
- The expected muon flux is low enough to use the emulsion detector in the tunnel

![](_page_30_Picture_7.jpeg)

![](_page_31_Picture_0.jpeg)

![](_page_31_Picture_1.jpeg)

In order to measure the muon rate away from the LOS in Run 3, we recently installed 20 small emulsion detectors within 2m of FASER. These were installed on 23/7 and removed on 2/9, having been exposed to ~10/fb of data. They will provide a useful validation of the FLUKA estimate further from the LOS.

![](_page_31_Figure_3.jpeg)

![](_page_31_Picture_4.jpeg)

![](_page_31_Picture_5.jpeg)

![](_page_31_Picture_6.jpeg)

## Sweeper Magnet: Ongoing Studies

- Preliminary design of sweeper magnet by TE-MSC
  - Based on permanent magnet to avoid power converter in radiation area
  - Consider 7m long (20x20cm<sup>2</sup> in transverse plane) magnet, 7Tm bending power lacksquare
- To install such a magnet would require some modifications to cryogenic lines in relevant area
  - Possibility of modifications to be investigated with LHC cryo  $\bullet$
  - Integration/installation aspects to be studied  $\bullet$
- FLUKA and BDSIM studies ongoing to assess effectiveness of such a magnet in reducing the muon background in the FPF

![](_page_32_Picture_8.jpeg)

![](_page_32_Picture_10.jpeg)

External ring (construction steel)

![](_page_32_Picture_12.jpeg)

![](_page_32_Picture_14.jpeg)