

# Neutrino-Nucleus $CC0\pi$ cross-section tuning in GENIE

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on behalf of the GENIE Collaboration

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# Overview

Neutrino interaction modeling is essential to **estimate the neutrino energy and calculate backgrounds and systematic errors** in neutrino measurements

Modeling of neutrino interactions is one of the dominant systematic uncertainties

- Need to improve the models and quantify and reduce systematic uncertainties associated to neutrino cross sections modeling

There is no available model that allows us to simulate the whole region of interest for neutrino experiments

- Event generators merge models together with ad-hoc prescriptions

**We rely on empirical approaches that need to be tuned**

# Neutrino-nucleus cross section data

## Huge effort from the experimental neutrino community

Many datasets are available for different neutrino energies and targets:

- ★ Inclusive measurements (  $\nu_{\mu} A \rightarrow \mu^{-} X$  )
- ★ Exclusive measurements
  - Topologies: CC0 $\pi$ , CC0p0 $\pi$ , CCNp0 $\pi$ , CC1 $\pi$ , ...
  - Single differential as well as double- and triple-differential measurements are available
  - Big effort on new measurements on STKI variables and proton kinematics

**These datasets are crucial to validate and tune models in neutrino MC event generators**

# GENIE's global analysis with Professor

GENIE develops a global analysis of scattering data for **tuning** and **uncertainty characterization** of comprehensive neutrino interaction models

GENIE's tuning program is based on the **Professor** tool:

- Tuning software tool from LHC community
  - Efficient implementation of complex multi-parameter brute-force scans
  - Applied to neutrinos for the first time by GENIE
  - Decoupled from event reweighting procedures
- Our goal is to perform a global tune that improves agreement with data
  - a. Tune GENIE's free nucleon models, including hadronization, with available data on hydrogen and deuterium targets
  - b. Tune nuclear models with modern neutrino data
  - c. Include electron-scattering data

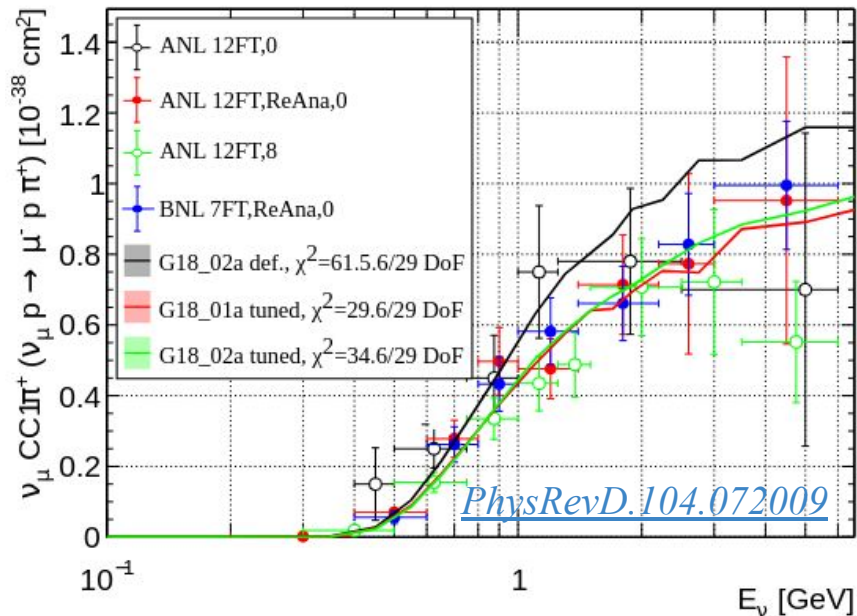


<https://professor.hepforge.org/>

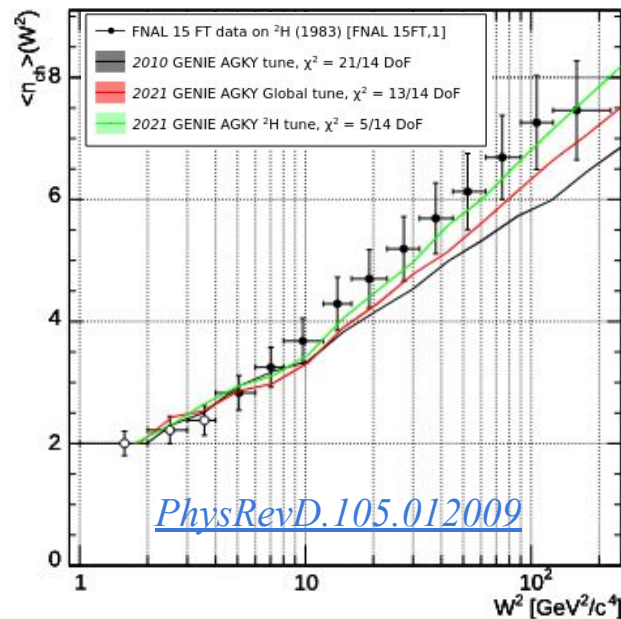
# The GENIE global analysis effort

The first goal was to tune  $\nu$ -N models - Core of  $\nu$ -A simulations!

Global fit to CC inclusive,  $1\pi$ , and  $2\pi$  data sets



First neutrino-induced hadronization tune on charged averaged multiplicity data on H and D



# The GENIE global analysis effort

The latest effort focuses on tuning of  $\nu$ -A data:

*Accepted for Publication at  
Phys.Rev.D*

*Neutrino-Nucleus  $CC0\pi$  cross-section tuning [ArXiv:2206.11050](#)*

This analysis incorporates new challenges with respect to previous free-nucleon tunes. It

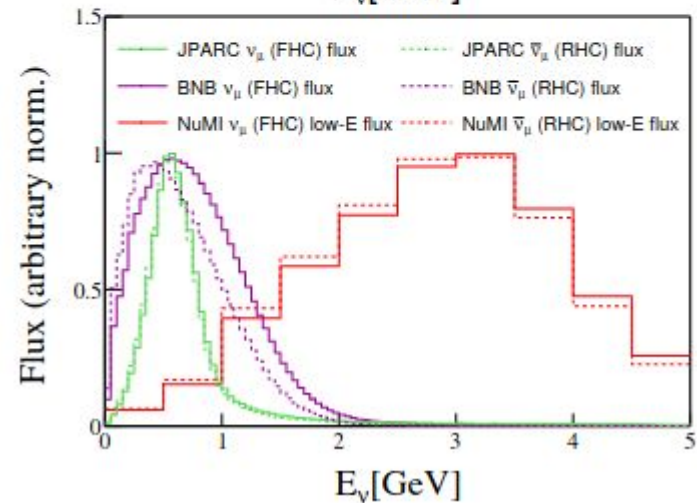
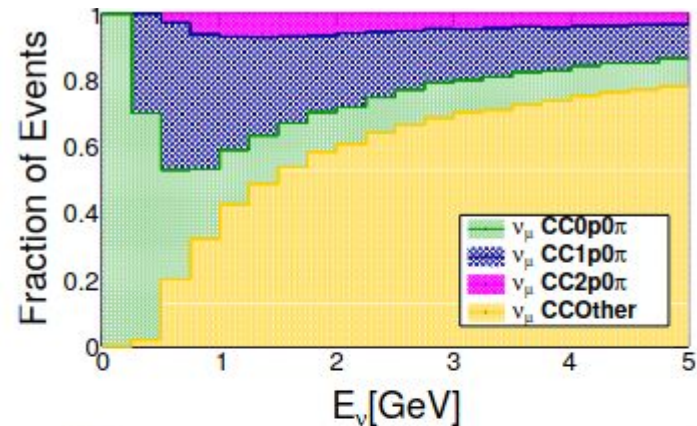
- Deals with the complexity of modeling the nuclear environment
- Consolidates the main elements of the tuning methodology with nuclear data
- Explores avenues for improving the agreement of GENIE and nuclear data
  - New parameterizations that encapsulate our lack of understanding of  $\nu$ A must be developed within GENIE

This work is the first step towards a global tune using all data on nuclei

**This work focuses on the tuning of modern neutrino  $CC0\pi$  cross-section data on carbon**

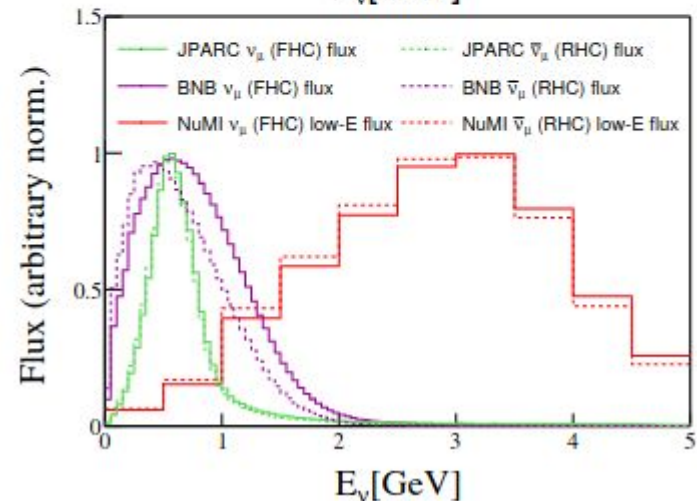
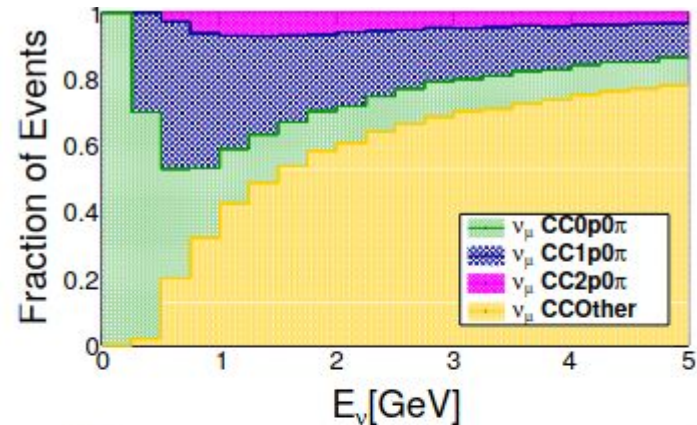
# The $CC0\pi$ Event topology

- **Dominant event topology at  $E_\nu < 1$  GeV**
- Usually defined as an event with a muon and no mesons in the final state
  - $CC0\pi$ : no mesons, any number of protons
  - $CC0p0\pi$ : no protons above detection threshold
  - $CCNp0\pi$ : at least one proton above detection threshold



# The $CC0\pi$ Event topology

- The contribution from different channels depends on the neutrino flux:
  - The  $CC0\pi$  topology is dominated by CCQEL events
  - Inelastic channels are also important due to FSI
    - Small RES contribution for T2K and MicroBooNE with respect to MINERvA
    - For T2K  $CC0p0\pi$ , most  $2p2h$  events have  $W \sim M_N$  whilst for MINERvA,  $W \sim M_\Delta$
    - Negligible contribution of DIS events



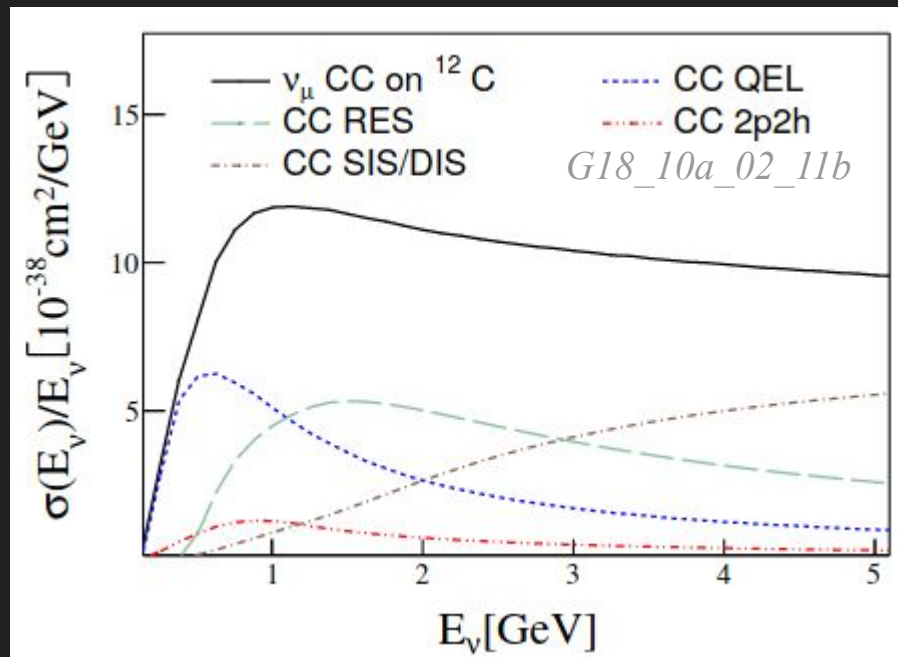


# Modeling of CC0 $\pi$ events with GENIE

GENIE has different models available to simulate neutrino-nucleus interactions:

- **CC QEL:** Llewellyn-Smith, Valencia or SuSAv2 model
- **CC 2p2h:** Empirical, Valencia or SuSAv2 model
- **CC RES:** Rein-Sehgal or Berger-Sehgal model
- **FSI:** hA2018, hN2018, INCL++
- **Nuclear model:** Relativistic, Local or Correlated Fermi Gas model

The models are grouped into different Comprehensive Model Configurations = self-consistent collections of the primary models



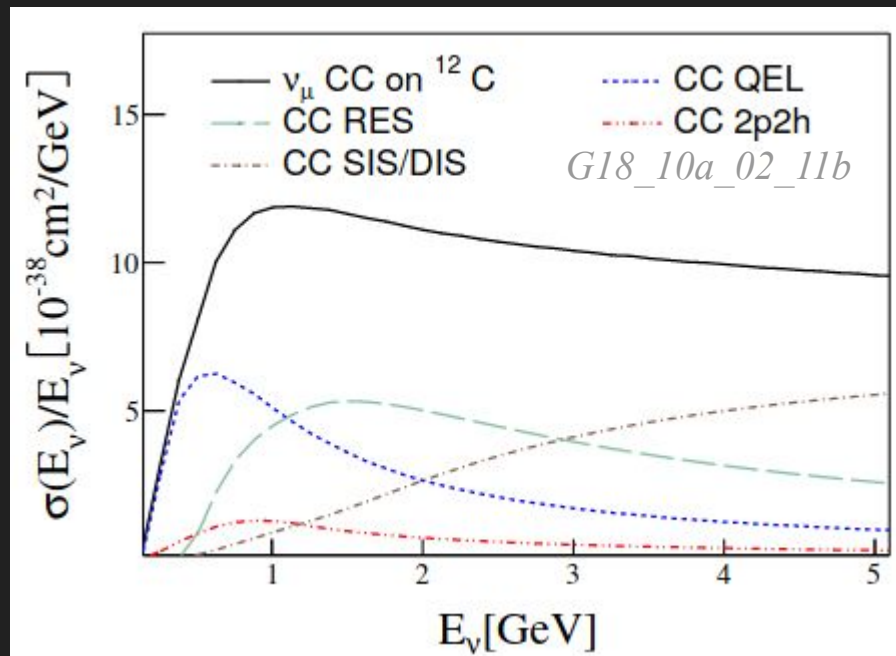
# Modeling of CC0 $\pi$ events with GENIE

In this work we use the  
G18\_10a\_02\_11b CMC:

- QEL+2p2h: Valencia model
- RES: Berger-Sehgal model
- FSI: GENIE hA2018
- Nuclear model: LFG

This CMC is tuned against free nucleon  
data on H and  $^2\text{H}$ ,

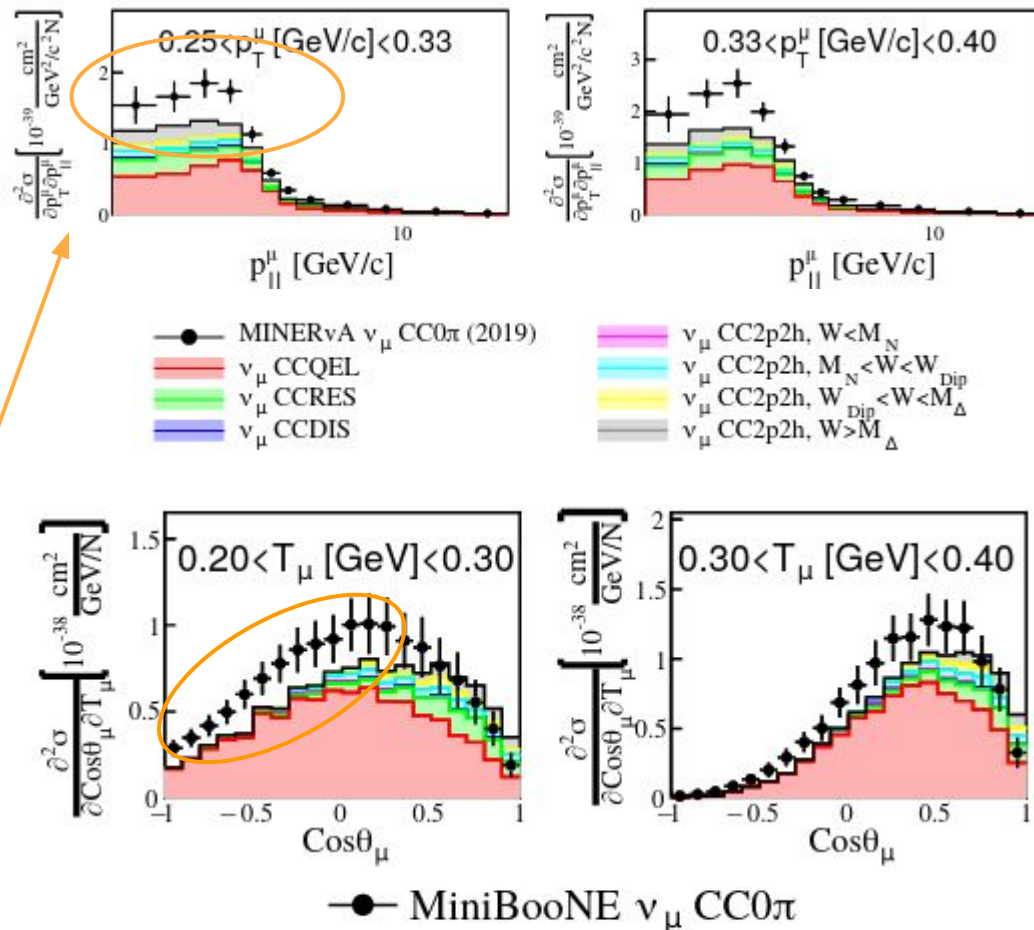
[PhysRevD.104.072009](https://arxiv.org/abs/1407.2009)



# Current description of CC0 $\pi$ data

The G18\_10a\_02\_11b CMC underpredicts all data on CC0 $\pi$  and CC0 $p_0\pi$

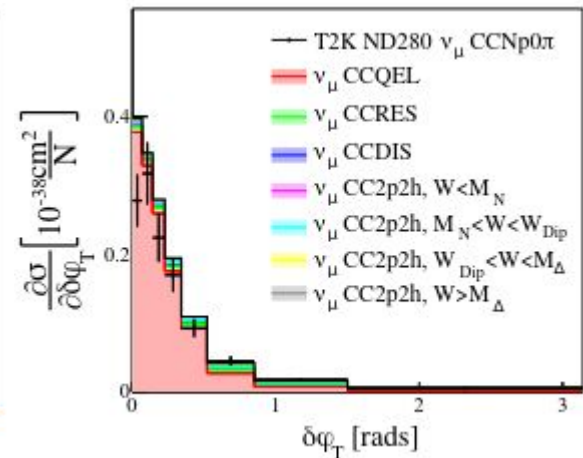
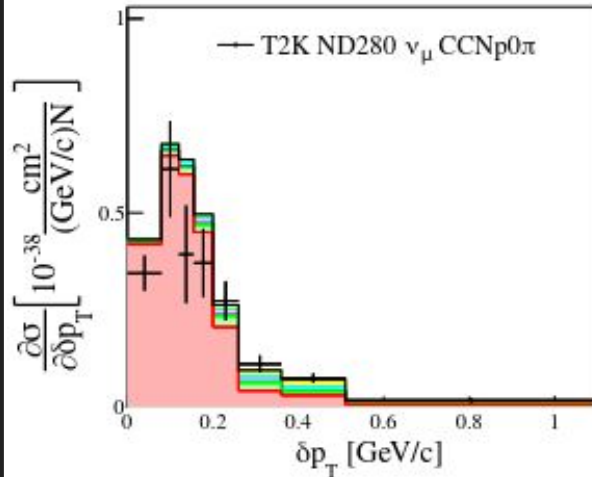
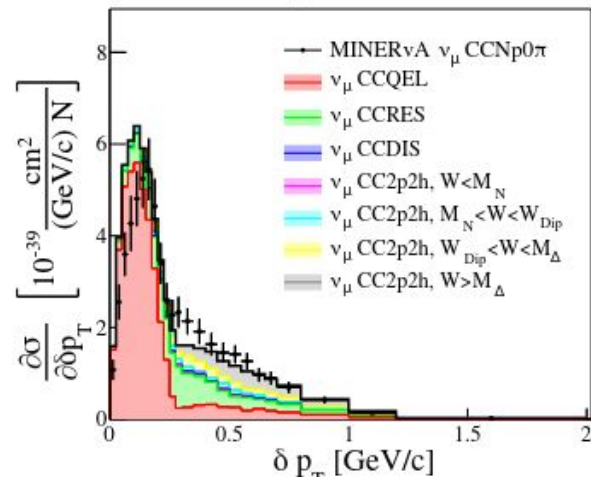
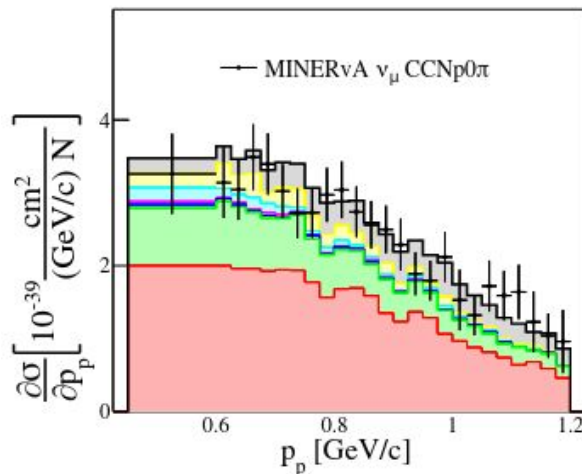
- Kinematic region where 2p2h events dominate
- But also for  $\cos\theta_\mu < 0$ , where QEL events dominate



# Current description of CCNp0 $\pi$ data

The G18\_10a\_02\_11b CMC has better agreement with CCNp0 $\pi$  data:

- It cannot describe CC0 $\pi$  and CCNp0 $\pi$  data at the same time
- CCNp0 $\pi$  data is not directly used in this analysis due to this tension



# Tuned parameters (1)

At the free nucleon level, the QEL cross section is well understood

- Base model was tuned to hydrogen and deuterium data
- We use correlated priors from free nucleon tune to constrain the nucleon axial mass ( $M_A^{\text{QEL}}$ ) and RES normalization factor ( $S_{\text{RES}}$ )

TABLE IV: Priors (a) and covariance matrix (b) for  $M_A^{\text{QEL}}$  and  $S_{\text{RES}}$  obtained to the free-nucleon tune from Ref. [5].

Parameter	Prior
$M_A^{\text{QEL}}$	$1.00 \pm 0.01 \text{ GeV}/c^2$
$S_{\text{RES}}$	$0.84 \pm 0.028$

(a)

	$M_A^{\text{QEL}}$	$S_{\text{RES}}$
$M_A^{\text{QEL}}$	$1.8 \times 10^{-4}$	$1.5 \times 10^{-4}$
$S_{\text{RES}}$	$1.5 \times 10^{-4}$	$6.0 \times 10^{-4}$

(b)

Note:  $M_A^{\text{QEL}}$  was originally also constrained by a prior from  $\nu_\mu d$  and pion electroproduction [*The European Physical Journal C* 53, 349–354 (2008)]

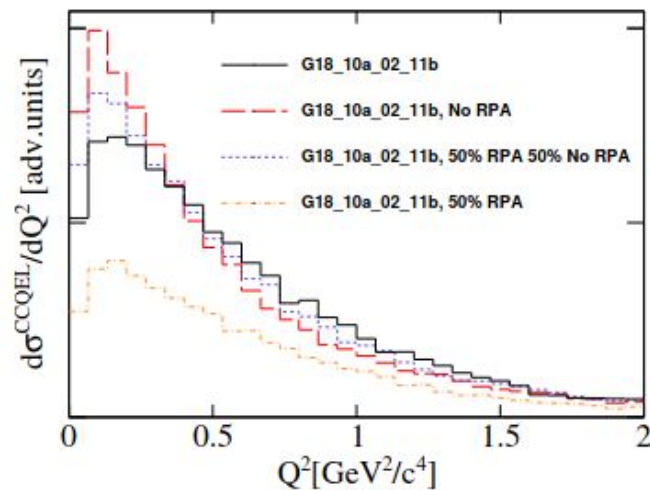
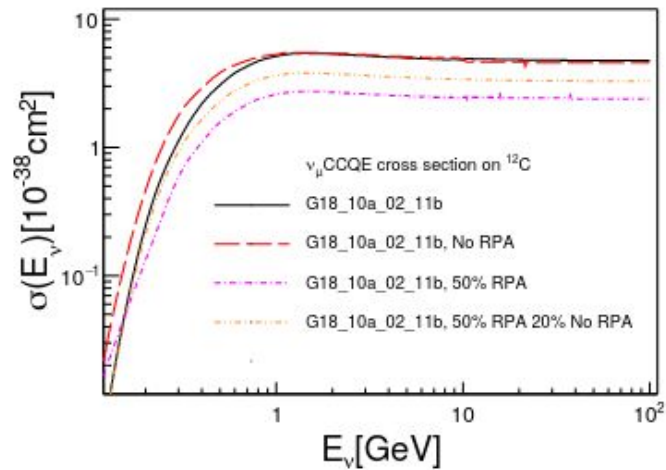
## Tuned parameters (2)

The QEL cross section is affected by the dynamics of the nuclear medium

- We include long-range NN correlations with the RPA correction
- Suppression of the QEL cross section at low  $Q^2$
- We parameterize the RPA correction as:

$$\sigma^{\text{QEL}} = \omega_{\text{RPA}} \cdot \sigma_{\text{RPA}}^{\text{QEL}} + \omega_{\text{No RPA}} \cdot \sigma_{\text{No RPA}}^{\text{QEL}}$$

$\omega_{\text{RPA}}/\omega_{\text{No RPA}}$  scales the cross section w/wo RPA (black/red line)



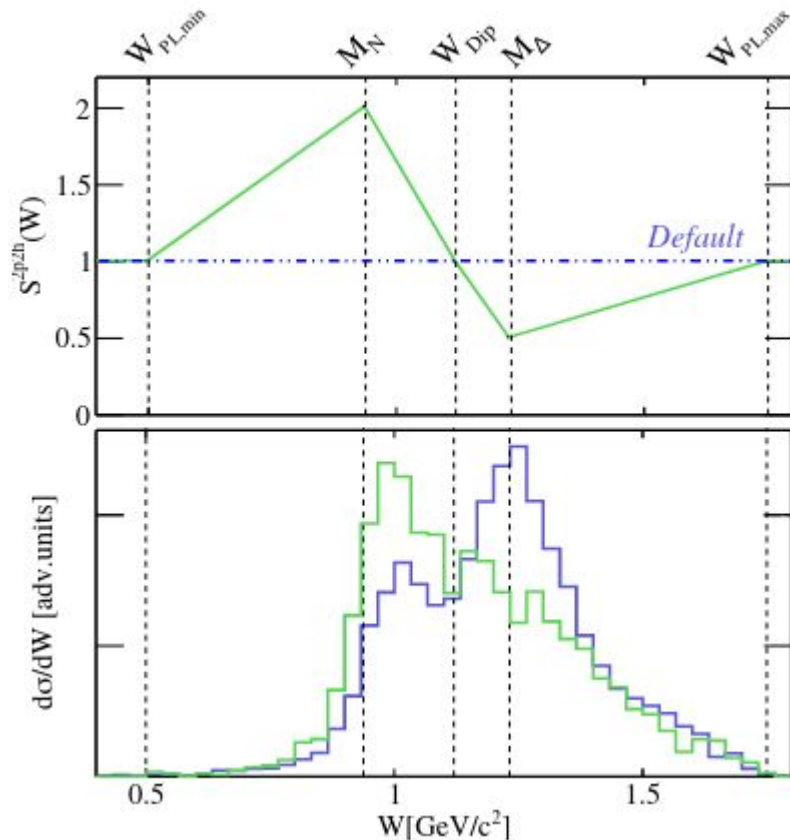
## Tuned parameters (3)

- The different available 2p2h models predict a different shape and strength for the 2p2h cross section
  - The Valencia model predicts two peaks in  $W$  at  $M_N$  and  $M_\Delta$
- We scale the 2p2h cross section as:

$$\frac{d^2 \sigma^{2p2h}}{dq_0 dq_3} \rightarrow S(W) \cdot \frac{d^2 \sigma^{2p2h}}{dq_0 dq_3}$$

$S(W)$  depends linearly on  $W$  between the different regions:

- $S_N^{2p2h} = S(M_N)$
- $S_\Delta^{2p2h} = S(M_\Delta)$
- $S_{PL}^{2p2h} = S(W_{PL,Max})$
- $S(W_{PL,Min}) = S(W_{Dip}) = 1$



**This empirical parameterization can be used to change the shape and scale the 2p2h cross section**

# Neutrino $CC0\pi$ datasets used for tuning

- All hydrocarbon targets
- Distinct fluxes prove  $E_\nu$  dependence
  - MiniBooNE and T2K ND280 flux's peaks below 1 GeV
  - MINERvA's low-E flux peaks at 3 GeV
- Identical base model, G18\_10a\_02\_11b:
  - QEL+2p2h: Valencia model
  - RES: Berger-Sehgal
  - FSI: GENIE hA2018
- A total of 5 partial tunes are performed within the same framework

Experiment	Probe	Event Topology	Partial Tune
MiniBooNE	$\nu_\mu$	$CC0\pi$	G10a
	$\bar{\nu}_\mu$	$CC0\pi$	G11a
T2K ND280	$\nu_\mu$	$CC0p0\pi$	G20a
MINERvA	$\nu_\mu$	$CC0\pi$	G30a
	$\bar{\nu}_\mu$	$CC0p0\pi$	G31a

**This approach provides with a common ground for the discussion of tensions between experiments**



## Partial tune results

- ❖ G10a: GENIE tune to MiniBooNE  $\nu$  CC0 $\pi$  data
- ❖ G11a: GENIE tune to MiniBooNE anti- $\nu$  CC0 $\pi$  data
- ❖ G20a: GENIE tune to T2K ND280  $\nu$  CC0 $\pi$  data
- ❖ G30a: GENIE tune to MINERvA  $\nu$  CC0 $\pi$  data
- ❖ G31a: GENIE tune to MINERvA anti- $\nu$  CC0 $\pi$  data

## All tunes

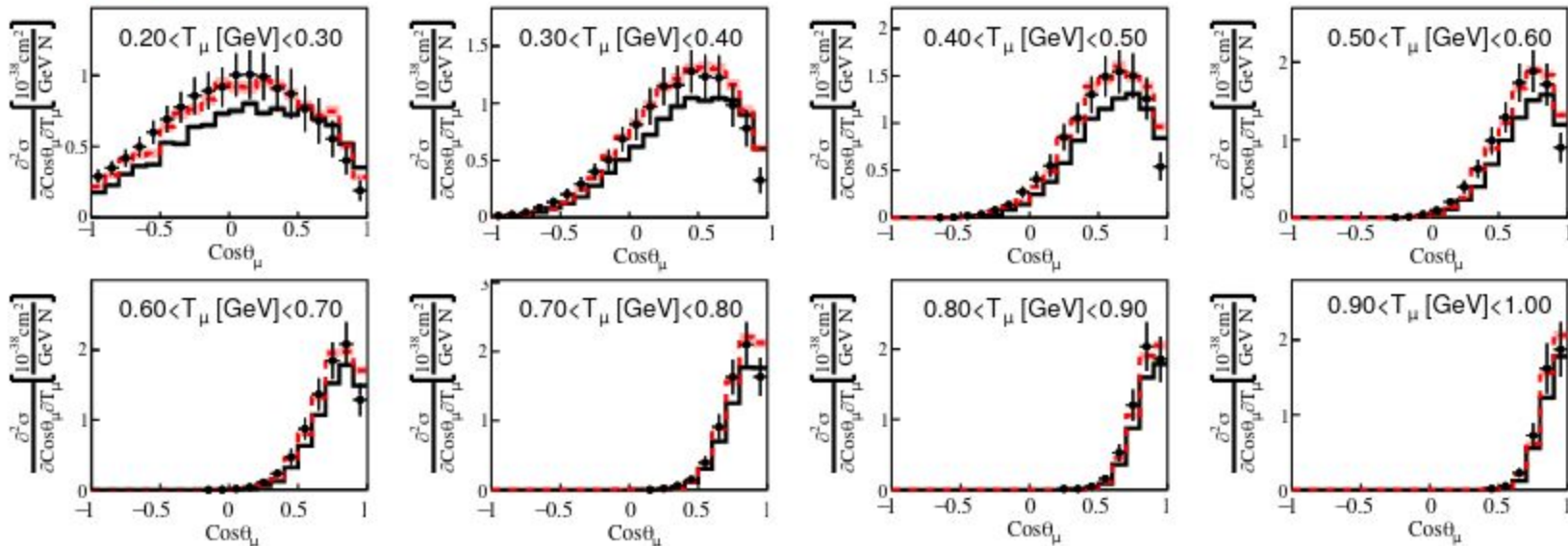
- Respect free nucleon priors imposed on  $M_A^{\text{QEL}}$  and  $S_{\text{RES}}$
- Have a preference for RPA corrections
- **Enhance the CCQEL and CC2p2h cross section**

Parameters	G10a Tune	G11a Tune	G20a Tune	G30a Tune	G31a Tune
$M_A^{\text{QEL}} (\text{GeV}/c^2)$	$1.02 \pm 0.01$	$1.01 \pm 0.01$	$1.00 \pm 0.01$	$1.00 \pm 0.02$	$1.00 \pm 0.01$
$\omega_{\text{RPA}}$	$1.20 \pm 0.03$	$1.14 \pm 0.06$	$1.2 \pm 0.2$	$0.9 \pm 0.1$	$1.3 \pm 0.2$
$\omega_{\text{No RPA}}$	$0.05 \pm 0.02$	$0.09 \pm 0.05$	$-0.1 \pm 0.1$	$0.2 \pm 0.1$	$0.2 \pm 0.2$
$S_{\text{RES}}$	$0.85 \pm 0.02$	$0.86 \pm 0.05$	$0.84 \pm 0.02$	$0.84 \pm 0.03$	$0.84 \pm 0.02$
$S_N^{2p2h}$	$1.5 \pm 0.4$	$2.3 \pm 0.01$	$1.7 \pm 0.3$	$1.2 \pm 0.4$	$1.7 \pm 0.5$
$S_\Delta^{2p2h}$	$0.7 \pm 0.2$	$0.7 \pm 0.3$	(1.00)	$2.1 \pm 0.2$	$2.3 \pm 0.2$
$S_{PL}^{2p2h}$	$0.4 \pm 0.1$	$0.4 \pm 0.1$	(1.00)	$0.9 \pm 0.2$	$0.4 \pm 0.1$
$\chi^2$	89/130	77/71	60/55	61/137	67/53

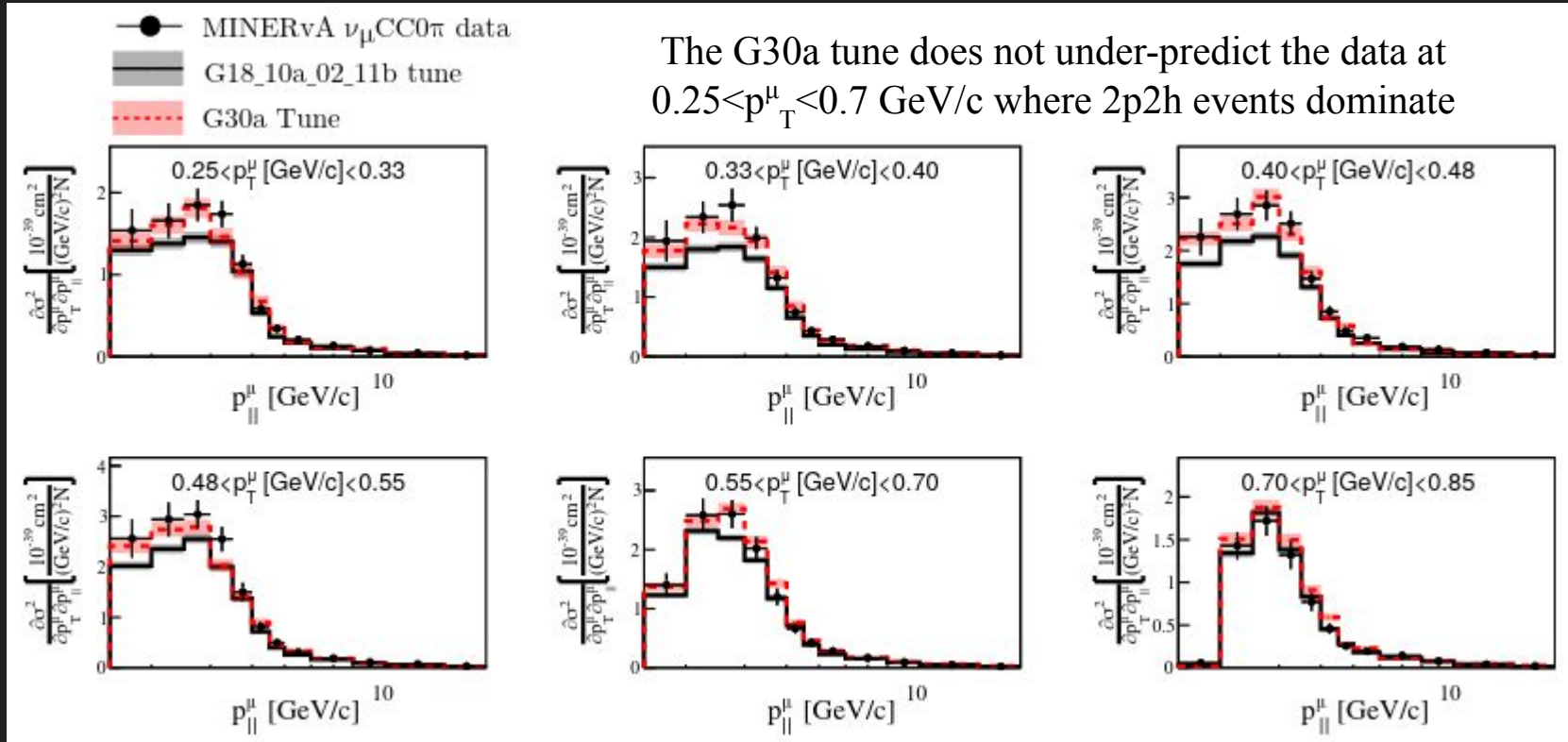
# Post-fit agreement with MiniBooNE CC0 $\pi$ data

- MiniBooNE  $\nu_\mu$ CC0 $\pi$  data
- G18\_10a\_02\_11b tune
- ⋯ G10a Tune

The enhancement of QEL and 2p2h cross sections is crucial for the correct shape and normalization description of the data



# Post-fit agreement with MINERvA CC0 $\pi$ data



# Partial tune results Tensions

However, differences between the partial tune results exist:

- MiniBooNE and T2K's tunes enhance the 2p2h cross section at  $W=M_N$ , whilst suppressing it at  $W=M_\Delta$
- MINERvA's tune enhances both peaks:  $S_{\Delta}^{2p2h} > S_N^{2p2h} > 1$

- ❖ G10a: GENIE tune to MiniBooNE  $\nu$  CC0 $\pi$  data
- ❖ G11a: GENIE tune to MiniBooNE anti- $\nu$  CC0 $\pi$  data
- ❖ G20a: GENIE tune to T2K ND280  $\nu$  CC0p0 $\pi$  data
- ❖ G30a: GENIE tune to MINERvA  $\nu$  CC0 $\pi$  data
- ❖ G31a: GENIE tune to MINERvA anti- $\nu$  CC0p0 $\pi$  data

Parameters	G10a Tune	G11a Tune	G20a Tune	G30a Tune	G31a Tune
$M_A^{QEL}(\text{GeV}/c^2)$	$1.02 \pm 0.01$	$1.01 \pm 0.01$	$1.00 \pm 0.01$	$1.00 \pm 0.02$	$1.00 \pm 0.01$
$\omega_{RPA}$	$1.20 \pm 0.03$	$1.14 \pm 0.06$	$1.2 \pm 0.2$	$0.9 \pm 0.1$	$1.3 \pm 0.2$
$\omega_{NoRPA}$	$0.05 \pm 0.02$	$0.09 \pm 0.05$	$-0.1 \pm 0.1$	$0.2 \pm 0.1$	$0.2 \pm 0.2$
$S_{RES}$	$0.85 \pm 0.02$	$0.86 \pm 0.05$	$0.84 \pm 0.02$	$0.84 \pm 0.03$	$0.84 \pm 0.02$
$S_N^{2p2h}$	$1.5 \pm 0.4$	$2.3 \pm 0.01$	$1.7 \pm 0.3$	$1.2 \pm 0.4$	$1.7 \pm 0.5$
$S_\Delta^{2p2h}$	$0.7 \pm 0.2$	$0.7 \pm 0.3$	(1.00)	$2.1 \pm 0.2$	$2.3 \pm 0.2$
$S_{PL}^{2p2h}$	$0.4 \pm 0.1$	$0.4 \pm 0.1$	(1.00)	$0.9 \pm 0.2$	$0.4 \pm 0.1$
$\chi^2$	89/130	77/71	60/55	61/137	67/53

**There's a clear energy dependence on the cross section shape**

# Conclusions of the nuclear tune

*Now available [on ArXiv](#)*

This work is the **first nuclear tune effort** performed with the GENIE global analysis framework:

- The goal is to tune against **CC0 $\pi$  data from MiniBooNE, T2K and MINERvA**
- Seven parameters are included to encapsulate CC0 $\pi$  modelling uncertainties
  - Correlated priors from the G18\_02a\_02\_11b tune are included
- A partial tune is performed for each experiment, providing with a common ground for the discussion of tensions
- All CC0 $\pi$  partial tunes **increase the CCQEL and CC2p2h** cross section
- Differences between the tunes exist, highlighting a **clear energy dependency on the cross section shape**

# Thank you for your interest

## The GENIE Collaboration

**Luis Alvarez-Ruso [4], Costas Andreopoulos [7,10], Adi Ashkenazi [11], Joshua Barrow [8,11], Steve Dytman [9], Hugh Gallagher [12], Alfonso Andres Garcia Soto [3,4] Steven Gardiner [2], Matan Goldenberg [11], Robert Hatcher [2], Or Hen [8], Timothy Hobbs [2], Igor Kakorin [6], Konstantin Kuzmin [5,6], Anselmo Meregaglia [1], Vadim Naumov [6], Afroditi Papadopoulou [8], Gabriel Perdue [2], Marco Roda [7], Beth Slater [7], Alon Sportes [11], Noah Steinberg [2], Vladyslav Syrotenko [12], Jeremy Wolcott [12], Júlia Tena Vidal [11]**

**1. CENBG, Université de Bordeaux, CNRS/IN2P3, 2. Fermi National Accelerator Laboratory, 3. Harvard University, 4. Instituto de Física Corpuscular (IFIC), 5. Alikhanov Institute for Theoretical and Experimental Physics (ITEP) of NRC "Kurchatov Institute", 6. Joint Institute for Nuclear Research (JINR), 7. University of Liverpool, 8. Massachusetts Institute of Technology (MIT), 9. University of Pittsburgh, 10. STFC Rutherford Appleton Laboratory, 11. Tel Aviv University, 12. Tufts University**

# Backup slides

# GENIE's global analysis with Professor

Advantages of GENIE's tuning approach:

- Not limited to reweightable parameters
- Allows massive parallelisation
- Reduces exponentially expensive brute force tuning to a scaling closer to the power law.
- Advanced system which can handle
  - Correlations between data bins
  - Correlated Gaussian priors
  - Nuisance parameters
  - Weights for specific data bins



<https://professor.hepforge.org/>



## Other nuclear uncertainties

Other parameterizations were considered initially but not used in the final analysis.

These were affecting:

- FSI pion absorption
- Binding energy correction for QEL and 2p2h events

They were found to be highly correlated with other aspects of this tune

These can be used in future global tunes where we include additional data such as:

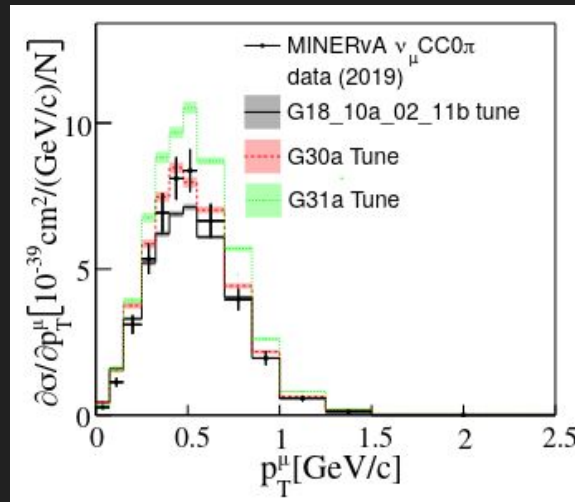
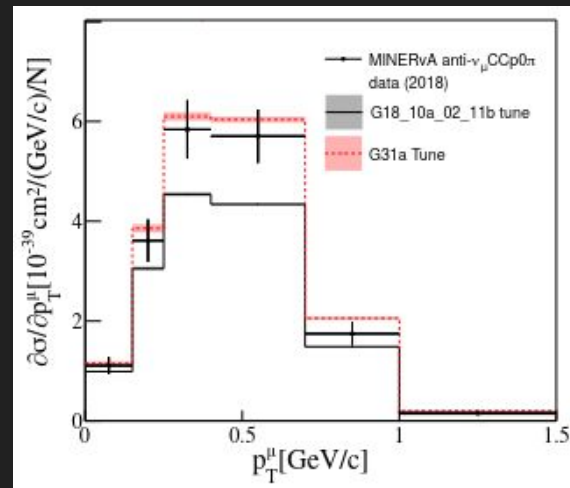
- ❖ CCNp0 $\pi$
- ❖ CC1 $\pi$  ...

# Tensions

The G31a tune (MINERvA anti- $\nu_{\mu}$  CC0p0 $\pi$ ) overpredicts all the other datasets

- MINERvA's  $\nu_{\mu}$  CC0 $\pi$  topology does not impose conditions on the proton multiplicity
- MINERvA's anti- $\nu_{\mu}$  CC0p0 $\pi$  topology requires no protons with  $T_p > 120$  MeV

This tension can be due to the different event topology definition and the neutrino type



# Tension between CC0 $\pi$ and CCNp0 $\pi$ datasets

Tune to MINER $\nu$ A  $\nu_{\mu}$  CCNp0 $\pi$  data:

- The G30a and G35a best-fit values are contradictory
- G35a suppresses the QEL and 2p2h cross sections to better describe the data
- **It highlights the need to improve the nuclear model**

- ❖ G30a: GENIE tune to MINER $\nu$ A  $\nu_{\mu}$  CC0 $\pi$
- ❖ G35a: GENIE tune to MINER $\nu$ A  $\nu_{\mu}$  CCNp0 $\pi$

Parameters	G30a Tune	G35a Tune
$M_A^{\text{QEL}} (\text{GeV}/c^2)$	$1.00 \pm 0.02$	$0.99 \pm 0.01$
$\omega_{\text{RPA}}$	$0.9 \pm 0.1$	$0.75 \pm 0.3$
$\omega_{\text{No RPA}}$	$0.2 \pm 0.1$	$0.09 \pm 0.3$
$S_{\text{RES}}$	$0.84 \pm 0.03$	$0.84 \pm 0.02$
$S_N^{2p2h}$	$1.2 \pm 0.4$	$0.33 \pm 0.2$
$S_{\Delta}^{2p2h}$	$2.1 \pm 0.2$	$0.5 \pm 0.4$
$S_{PL}^{2p2h}$	$0.9 \pm 0.2$	$1.5 \pm 0.4$
$\chi^2$	61/137	17/19

Dataset	$\chi_{\text{Nominal}}^2$	$\chi_{\text{G10a}}^2$	$\chi_{\text{G11a}}^2$	$\chi_{\text{G20a}}^2$	$\chi_{\text{G30a}}^2$	$\chi_{\text{G31a}}^2$	$\chi_{\text{G35a}}^2$	DoF
MINER $\nu$ A CCNp0 $\pi$ data								
$d\sigma/dp_p$	21	22	25	32	36	58	27	25
$d\sigma/d\theta_p$	58	153	150	113	129	226	20	26
$d\sigma/d\delta p_T$	102	637	568	360	352	625	42	24
$d\sigma/d\delta\phi_T$	87	505	467	314	354	566	18	23
$d\sigma/d\delta\alpha_T$	15	21	29	24	30	57	17	12
$d\sigma/d\delta p_{Tx}$	159	727	710	467	555	768	62	32
$d\sigma/d\delta p_{Ty}$	127	832	776	553	599	792	51	33

# Neutrino $CC0\pi$ cross-section data

This analysis focuses on:

**MiniBooNE:**  $\nu_{\mu}$ - $^{12}\text{C}$   $CC0\pi$  and  $\text{anti-}\nu_{\mu}$ - $^{12}\text{C}$   $CC0\pi$

**T2K ND280:**  $\nu_{\mu}$ - $^{12}\text{C}$   $CC0p0\pi$

**MINERvA:**  $\nu_{\mu}$ - $^{12}\text{C}$   $CC0\pi$  and  $\text{anti-}\nu_{\mu}$ - $^{12}\text{C}$   $CC0p0\pi$

T2K and MINERvA's data releases provide with information on the bin-to-bin correlation due to systematic uncertainties

➤ This information is included in our analysis

MiniBooNE does **not** provide with such information - Instead they quote a 10.7% normalization uncertainty for the neutrino case, which we add to our database.

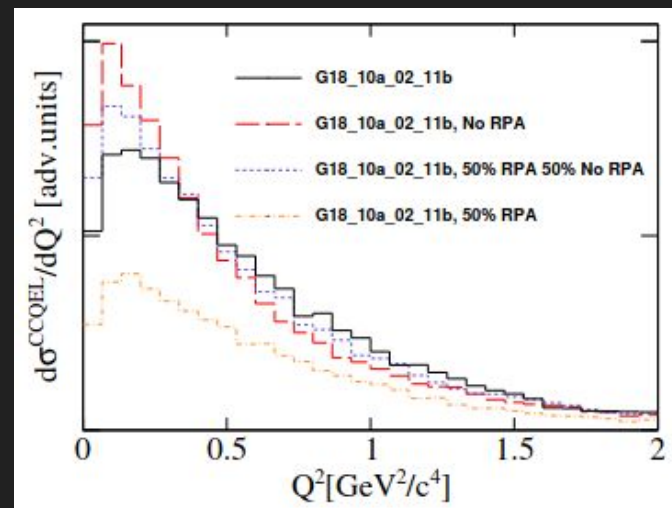
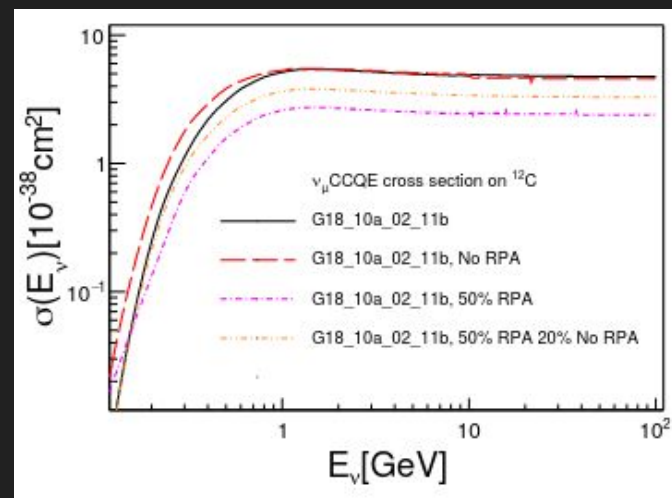
# Goals of the Neutrino-Nucleus $CC0\pi$ cross-section analysis

- Consolidate the main elements of the  $CC0\pi$  tuning methodology
- Explore avenues for improving the agreement of GENIE and nuclear data
- Provide with a common ground for the discussion of tensions between experiments

# Parameterization of nuclear uncertainties of the CCQEL cross section

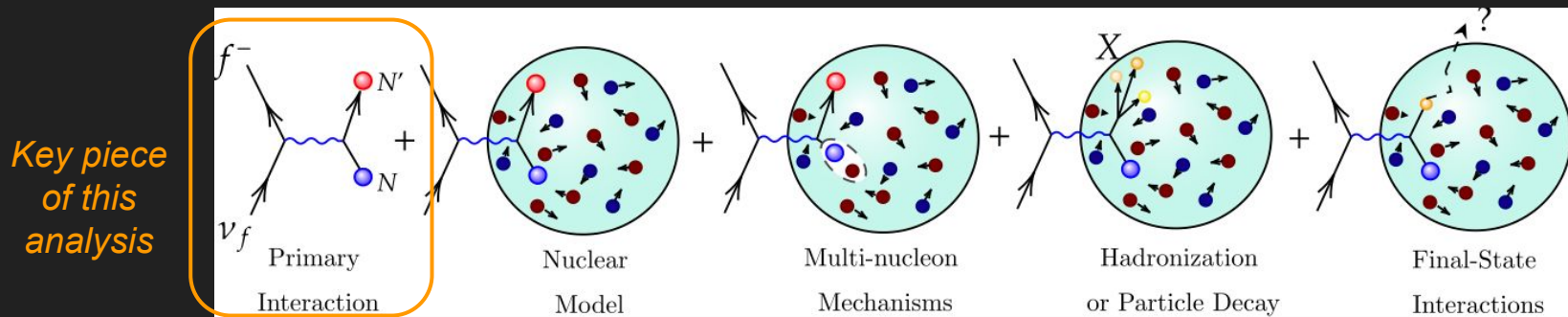
- At the free nucleon level, the CCQEL cross section is well understood
  - We impose a prior on the sum:  $w_{\text{RPA}} + w_{\text{No RPA}} = 1 \pm \sigma_S$
  - Nuclear effects might include some uncertainty on the scaling:  $\sigma_S = 0.2$
- We also need to impose a prior to the difference:
  - $w_{\text{RPA}} - w_{\text{No RPA}} = 1 \pm \sigma_\Delta$
  - $\sigma_\Delta = 5$
- This information is included in our tune using correlated priors

$$\Sigma_{\text{RPA}} = \frac{1}{4} \begin{pmatrix} \sigma_S^2 + \sigma_\Delta^2 & \sigma_S^2 - \sigma_\Delta^2 \\ \sigma_S^2 - \sigma_\Delta^2 & \sigma_S^2 + \sigma_\Delta^2 \end{pmatrix}$$



# Uncertainties related to the free nucleon modeling

We model  $\nu A$  interactions by adding additional layers to the  $\nu N$  modeling



The **free nucleon tune** provides with data-driven constraints for the free-nucleon parameters

- Tuned against exclusive  $1\pi$  and  $2\pi$  data from ANL, BNL, BEBC and FNAL
- This information is included using correlated priors for  $M_A^{\text{QEL}}$  and  $S_{\text{RES}}$

# Parameterization of the CC2p2h cross section

- GENIE has three 2p2h models available
- Each model predicts a different shape and strength for the 2p2h cross section

