

#### **UFO observations with IC and diamond BLMs**

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# **Study of UFO dynamics**

- Requires blown-up bunches, and ideally displaced bunches
- Goals:
  - Understand the release mechanism
    - Do UFOs **statistically drop** from the beam screen?
    - Can they be **charged** and **attracted** by the beam?
  - Benchmarking of the dynamics simulation model
  - Identifying dust particle size and material (benchmarked by measurements)
  - How conditioning/deconditioning mechanism works
  - Identify their origin and how to best operate in their presence
- Triggering algorithm for recording UFOs with dBLMs in IR7 was implemented and blown-up bunches used during standard operation



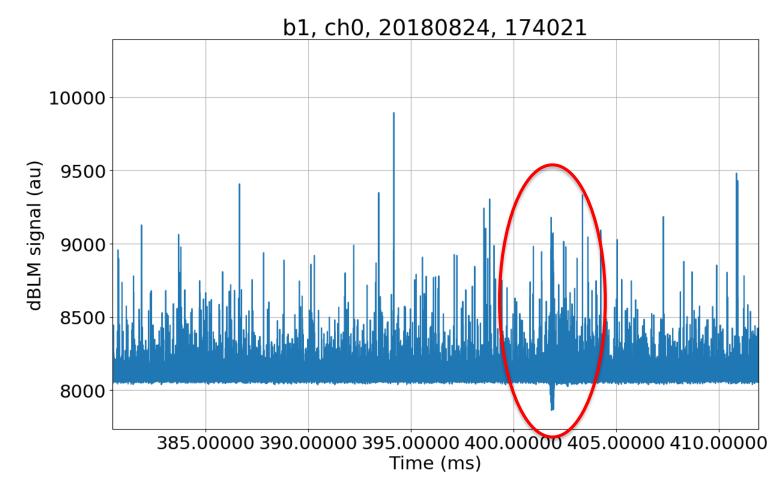
# **Readout system and Trigger algorithm**

- VFC FPGA based readout system installed in parallel to ROSY in 2018
- Allows flexible and advanced triggering
- Two algorithms in parallel:
  - Signal above 8200 bits from min 100 bunches in one/several turns in a 5 ms window AND condition not fulfilled in the next 5 ms window
  - Total integrated signal in 5 ms window above threshold AND condition not fulfilled in preceding and following 5 ms windows
- ~1 UFO recorded per beam day (~2 from UFO buster)



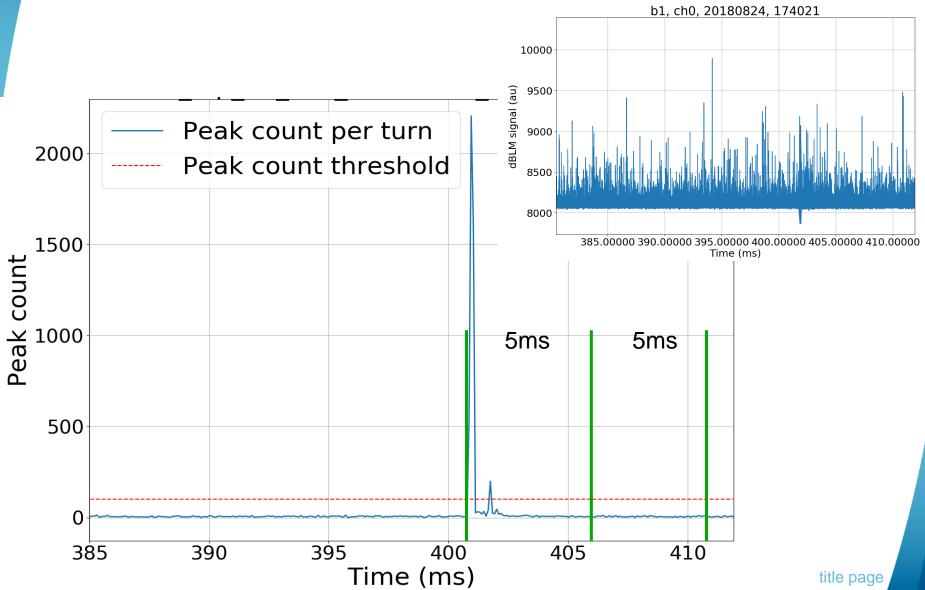
## **Trigger example**

UFOs hidden by background losses in TCP region



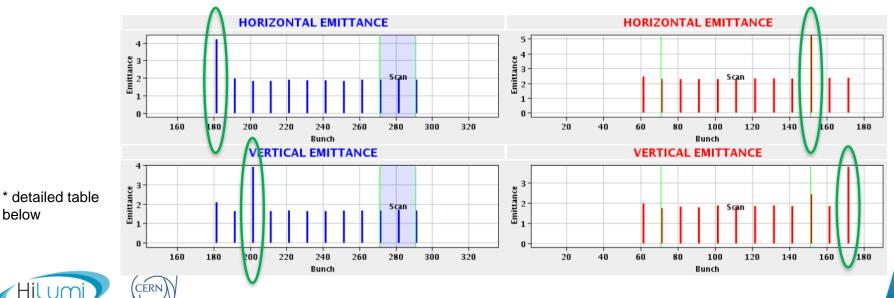


#### **Trigger example**



# UFOs at end of run 2 proton physics

- Automatic script for blowing up bunches implemented by ADT expert
  - Manually triggered at injection, manual confirmation of successful result, no problems encountered
- Routinely used during physics fills from Sep 29 (f7234) until Oct 23 (f7334)
  - **24 fills** with blown-up bunches (2 oo 12 non-colliding)
  - 205.2 hours of SB, 16.9 hours of RAMP
  - 13\* UFOs detected by dBLM (2 during MD period)
  - 3 at top energy, all others during ramp 1096-4522 GeV
  - 6 coincident triggers dBLM/UFO buster (out of 33 detected by UFO buster)
- **No issues** in machine operation due to ADT blow-up were observed



## dBLM UFO events with blown-up bunches

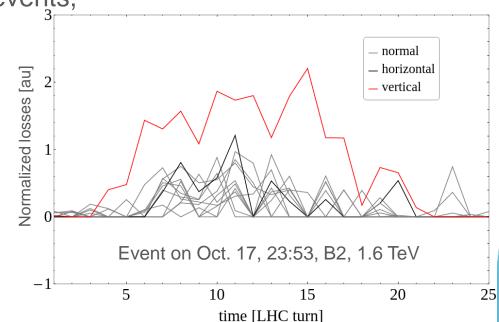
| time (local)      | fill | beam | location (ICBLM with max signal) | energy [TeV] |
|-------------------|------|------|----------------------------------|--------------|
| 20180930_224752   | 7240 | b2   | BLMQI.15L1                       | 6.5          |
| 20181003_121604   | 7252 | b1   | BLMAI.11R8                       | 6.5          |
| 20181003_152726   | 7253 | b2   | n/a                              | 2.9          |
| 20181007_012459   | 7264 | b2   | n/a                              | 3.4          |
| 20181009_174201   | 7271 | b1   | n/a                              | 1.1          |
| 20181016_095144   | 7308 | b1   | BLMBI.27L4                       | 4.5          |
| 20181016_141328   | 7309 | b1   | n/a                              | 2.1          |
| 20181017_235320   | 7314 | b2   | n/a                              | 1.6          |
| 20181017_235320_2 | 7314 | b2   | n/a                              | 1.6          |
| 20181019_143039   | 7319 | b1   | BLMQI.05L1                       | 3.4          |
| 20181020_094316   | 7321 | b1   | BLMQI.08R3                       | 2.6          |
| 20181026_231622   | 7365 | b1   | BLMMI.16L2                       | 6.5          |
| 20181026_231623   | 7365 | b1   | BLMMI.16L2                       | 6.5          |



## **Challenges of dBLM measurements**

Useful signal for several turns in only 4 events

- 1 has horizontal preference / 2 have vertical preference
- 1 was without blown-up bunches
- Useful signal during 2 turns in 3 events, and 1 turn in 5 events
  - Difficult to conclude about the dynamics in very fast events
- Even useful signal suffers from **fluctuations**
- Multi-turn losses in IR7 distort the falling edge



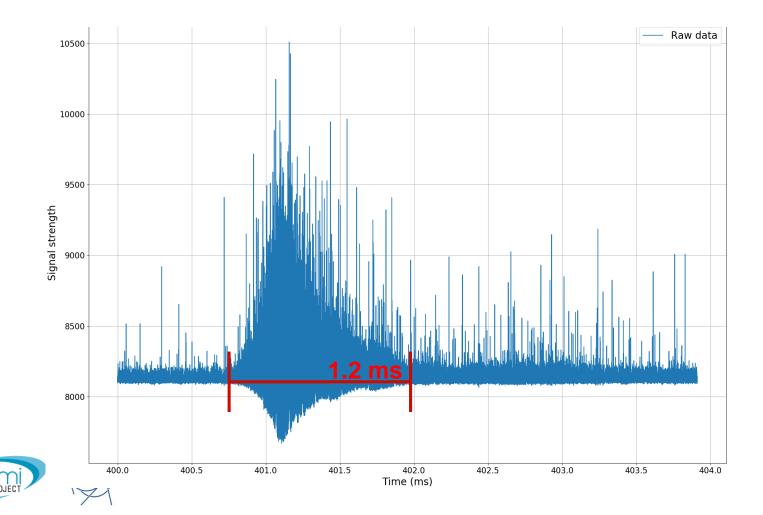
- 10 out of 13 events during ramp
  - Is there a detection bias or an increased UFO rate?
- During validation period, before blown-up bunches, 8 out of 16 events were at 6.5 TeV and consequently signal to noise ratio much better



To draw conclusions, need to analyse dBLM data, UFO buster recordings and perform simulations

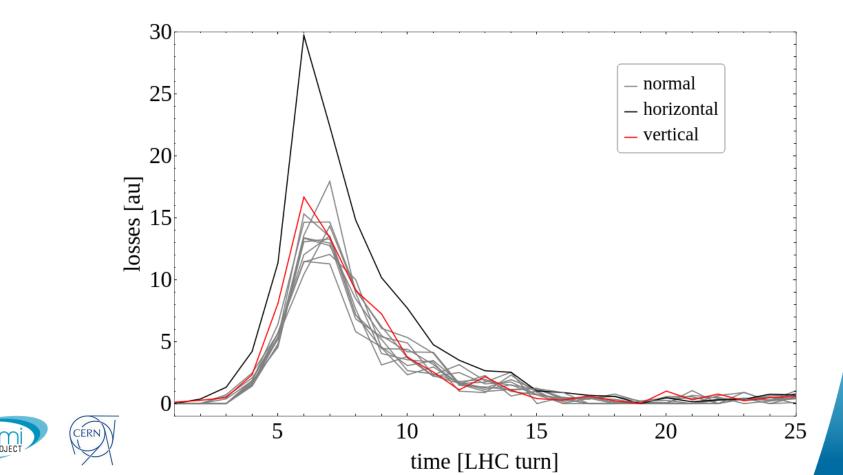
#### **UFO measurement example**

- Event on Sep 30, 22:47, (Q15L1), B2, 6.5 TeV
- Raw waveform as measured in IR7 by dBLM
- Best recording (with blown-up bunches)



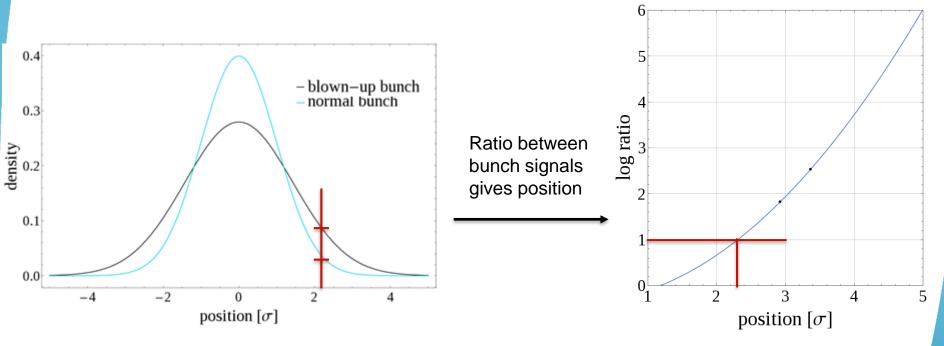
### UFO measurement – 12b train

- Event on Sep 30, 22:47, (Q15L1), B2, 6.5 TeV
- 12b train (bunches shown separately)
- Clear signal, but significant fluctuations between similar bunches
- Bunch-by-bunch signals can be used to estimate UFO position



## **Ratio of bunches – Method**

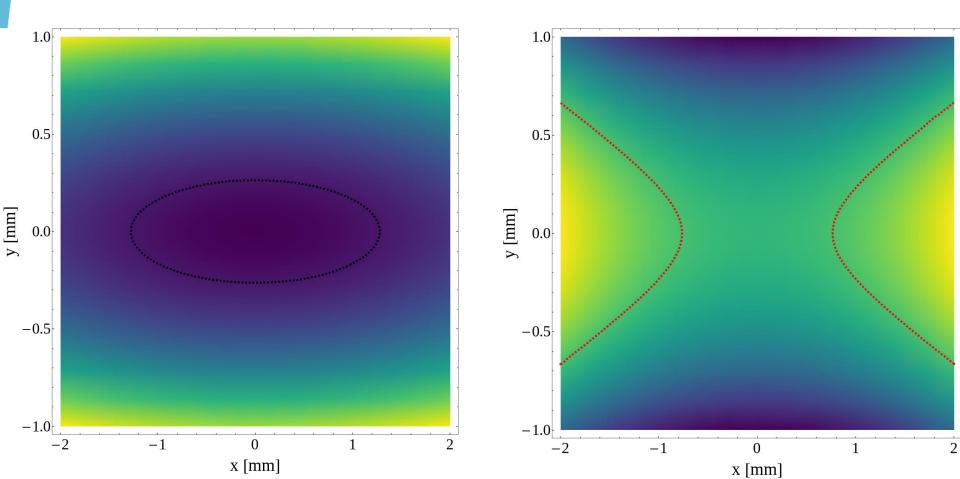
- Bunch profiles assumed gaussian
- Losses proportional to particle density at interaction point
- Ratio of bunch profiles ~= ratio of losses





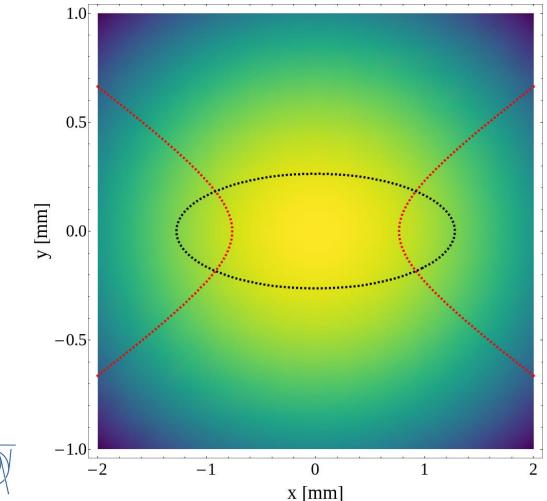
#### **Ratio of bunches – Example**

- Horizontally blown-up bunch divided by reference bunch (left)
- Horizontally blown-up bunch divided by Vertically blown-up bunch (right)
- Dashed lines are contours where horizontal bunch density 2 times reference and vertical bunches



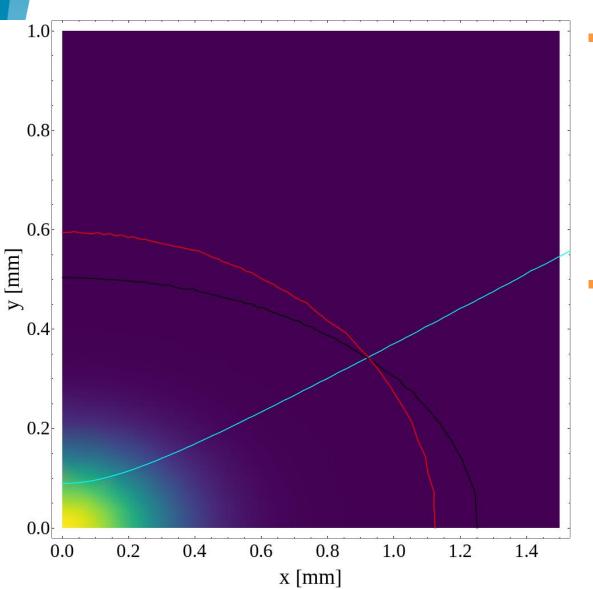
### **Ratio of bunches – Example**

- Combining both ratios gives the estimated UFO position at the intersections
- UFO position is uniquely determined (with a four-fold symmetry)





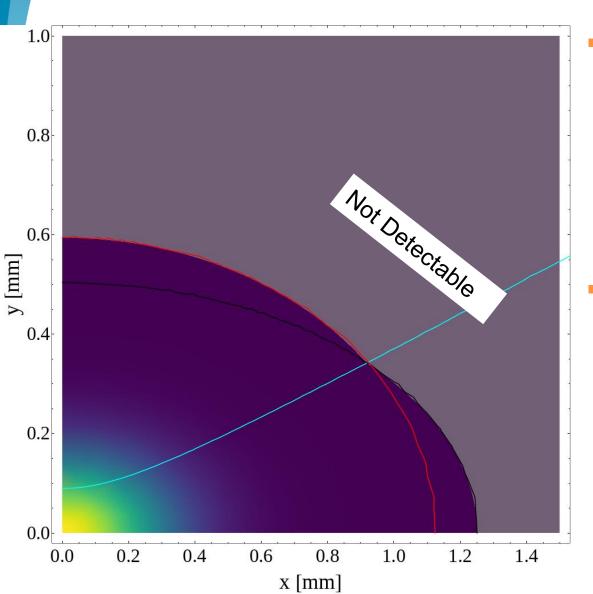
## **Exclusion regions**



Proton density necessary for visible losses within the red (ver. bunch) and black (hor. bunch) ellipses

- includes pessimistic margin
- Cyan line: contour
  where ratio of horizontal
  and vertical bunch
  densities equal 1

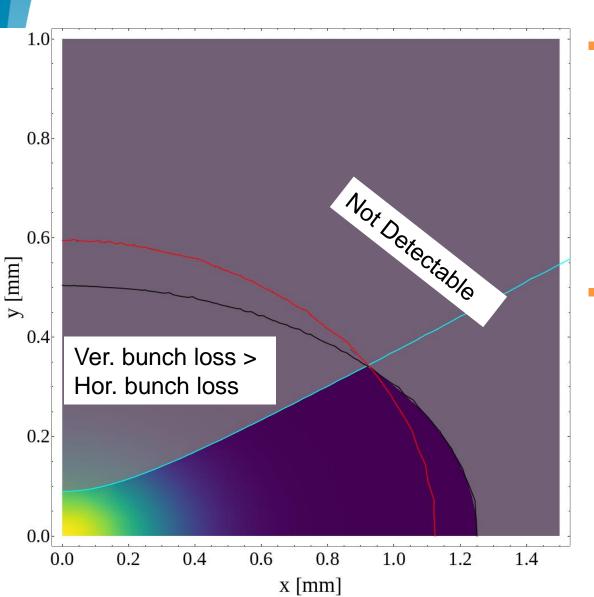
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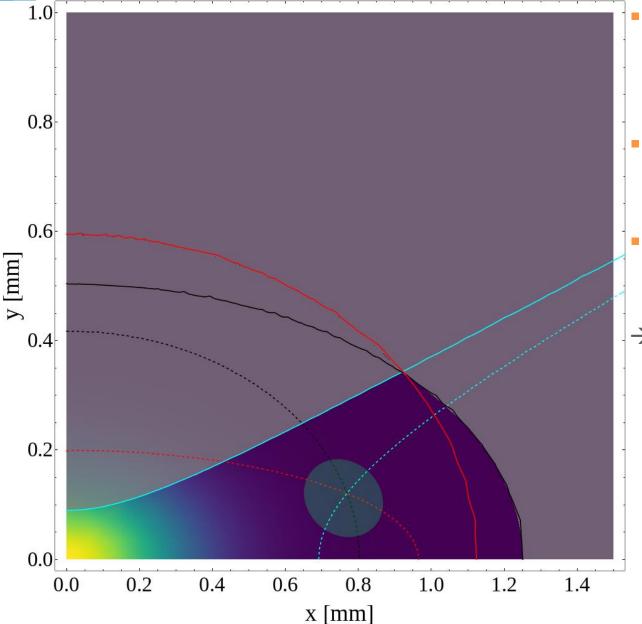
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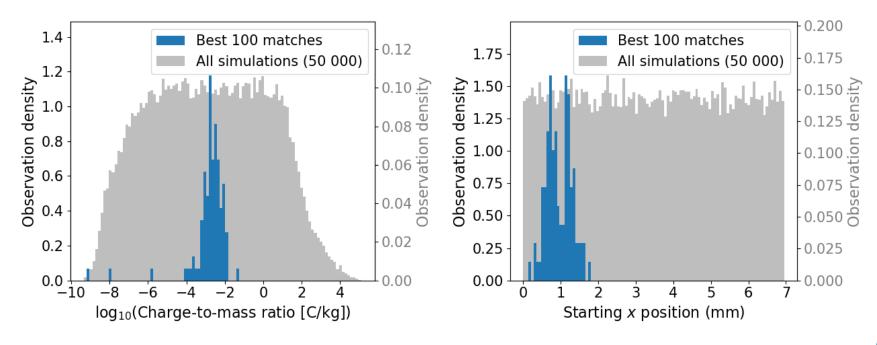
#### **UFO position estimate**



- Filled ellipse shows position during peak losses, with a 1  $\hat{\sigma}$  error on measurements
- Below the cyan line (more losses from hor. bunch)
- Consistently more losses from hor. bunch throughout the event
- $\rightarrow$  Horizontal Movement

# **Monte Carlo simulations**

- Different UFO candidates could explain the measurements (ICBLM, dBLM)
  - 7 input parameters  $\rightarrow$  1 output signal, however many events
  - Different scenarios can lead to the same simulated output (ICBLM, dBLM)
- Nevertheless, the important physical quantities (in order to understand UFOs release mechanism) converge quite well

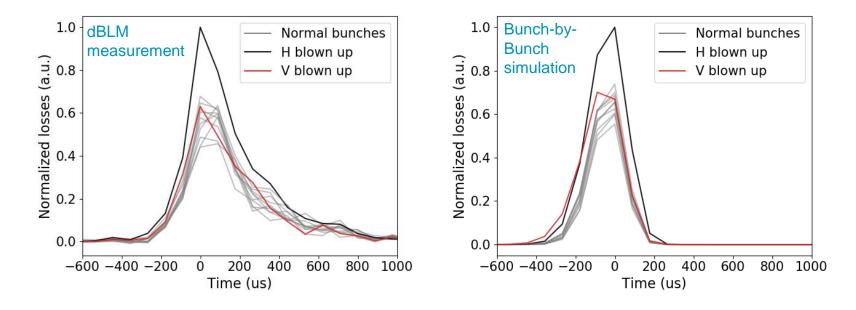




"Best matches" between simulations and the event presented on previous slides

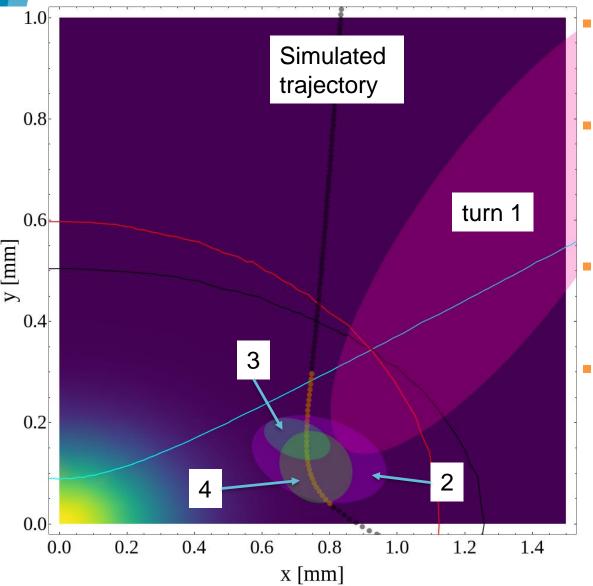
# **Comparison with simulations – Example**

- Simulations were matched to local ICBLM measurements (no bunch-by-bunch input!)
- Good agreement of the rising edge for the blown-up bunches.
  - dBLM measurement suffers from delayed losses, distorting the falling edge
- Important for understanding the simulated UFO trajectory





#### **Simulated vs Measured trajectory**



- Orange dots correspond to the 4 peak turns in simulation
- Ellipses correspond to measured position, 4 is the turn of peak losses
- Closest approach fits well between simulation and measurement
- Simulated trajectory in region where Ver. bunch > Hor. bunch. This is not seen in measurements, better fit should exist

# Conclusions

- Estimated UFO position together with consistently larger losses from horizontal bunches requires horizontal movement
  - UFOs are pre-charged or released with an initial speed
- Good agreement between simulations and measurements in center part, where resolution is best
  - Trajectory and Bunch-by-bunch losses
- Simulations fit despite only taking ICBLM measurements into account
  - Hints that overall time profiles are unique to certain trajectories

#### **Outlook:**

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- Include bunch-by-bunch data to find a better fitting trajectory
- Combine simulations with loss tracking to "deconvolute" the delayed TCP losses
- Reiterate on the other recorded events



## **Acknowledgements**

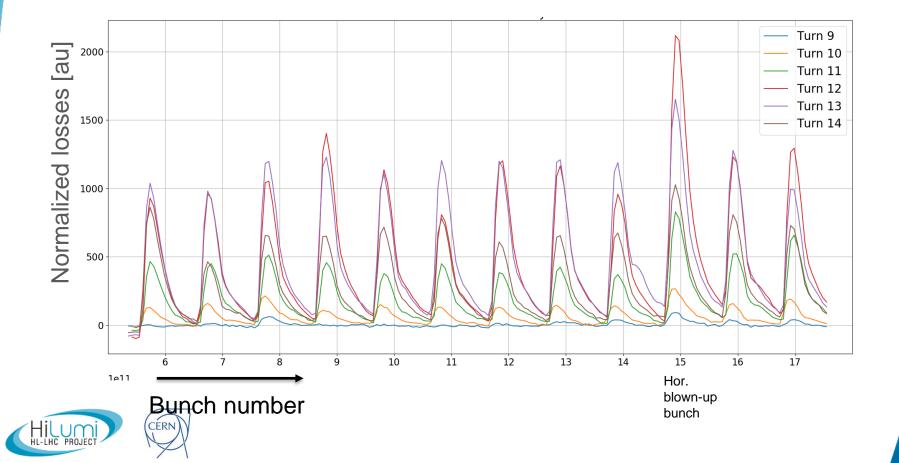
Many thanks to A. Lechner for continuous support and discussions

Many thanks to D. Valuch for implementation of bunch excitation and to OP team for consistent execution



#### **UFO measurements zoom**

Event on Sep 30, 22:47, (Q15L1), B2, 6.5 TeV



# dBLM UFO events without blown-up bunches

| time (local)    | fill | beam | confirmed by UFO buster | energy [TeV] |
|-----------------|------|------|-------------------------|--------------|
| 20180712_180259 |      | b2   | yes                     | 6.5          |
| 20180716_004600 |      | b1   | yes                     | 6.5          |
| 20180717_220721 |      | b1   | yes – 16L2              | 6.5          |
| 20180724_040228 |      | b1   | yes                     | 2.732        |
| 20180728_024804 |      | b1   | no                      | 6.5          |
| 20180806_220338 |      | b1   | yes                     | 6.5          |
| 20180807_210027 |      | b1   | yes                     | 6.5          |
| 20180814_045856 |      | b1   | no                      | 1.211        |
| 20180823_150941 |      | b2   | no                      | 6.5          |
| 20180824_152558 |      | b2   | yes                     | 6.274        |
| 20180824_174021 |      | b1   | yes                     | 6.5          |
| 20180830_110744 |      | b1   | no                      | 1.263        |
| 20180831_052212 |      | b2   | no                      | 2.651        |
| 20180903_080452 |      | b1   | no                      | 3.501        |
| 20180903_181837 |      | b2   | yes                     | 6.326        |
| 20180904_214859 |      | b1   | no                      | 0.465        |



## dBLM UFO events Ion Run

| time (local)    | fill | beam | confirmed by UFO buster | energy [TeV] |
|-----------------|------|------|-------------------------|--------------|
| 20181111_112533 |      | b2   | yes – 16L2              | 5.2          |
| 20181122_101510 |      | b2   | yes                     | 6.4          |
| 20181123_031804 |      | b2   | no                      | 6.4          |
| 20181123_060221 |      | b2   | no                      | 6.4          |
| 20181123_234255 |      | b2   | no                      | 6.4          |
| 20181128_033815 |      | b2   | yes                     | 6.4          |



## **UFO position estimate error**

- Red points show calculated position including error, turn with smallest error
- Error evaluated from std dev.  $\hat{\sigma}$  of turn-by-turn loss value from reference bunches
- The points are calculated by varying the measured values by  $\pm 1 \hat{\sigma}$

