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# Analysis of B/C ratio and nuclei energy spectra

NASA GSFC / CRESST-UMBC

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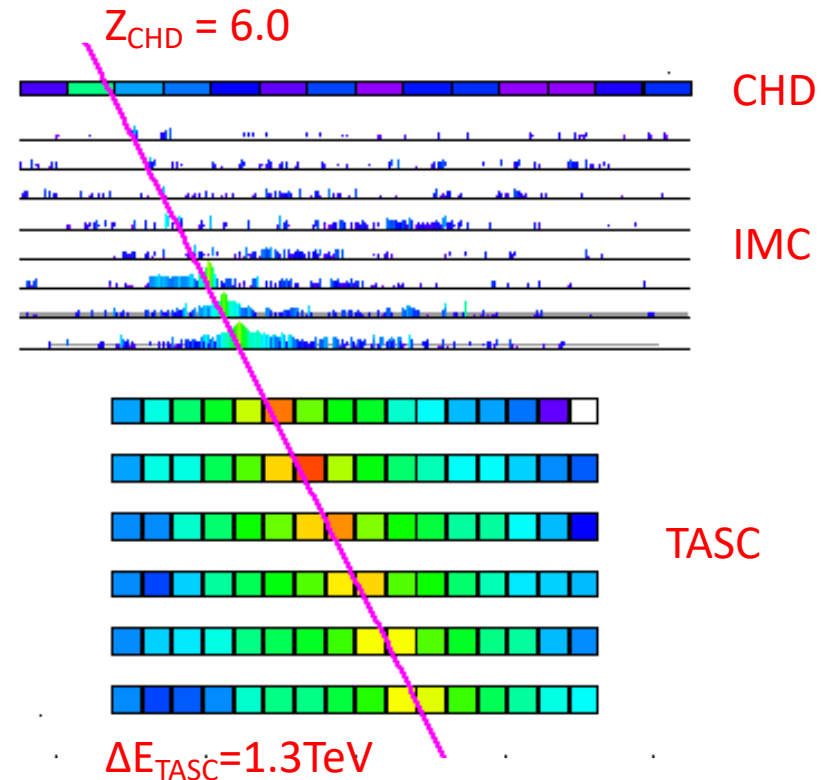
# Contents

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- B/C ratio
- Carbon spectra
  - Energy scale
  - Trigger efficiency
  - Direct pair production
- Heavy nuclei spectra
  - Ne, Mg, Si, S, Ca and Fe

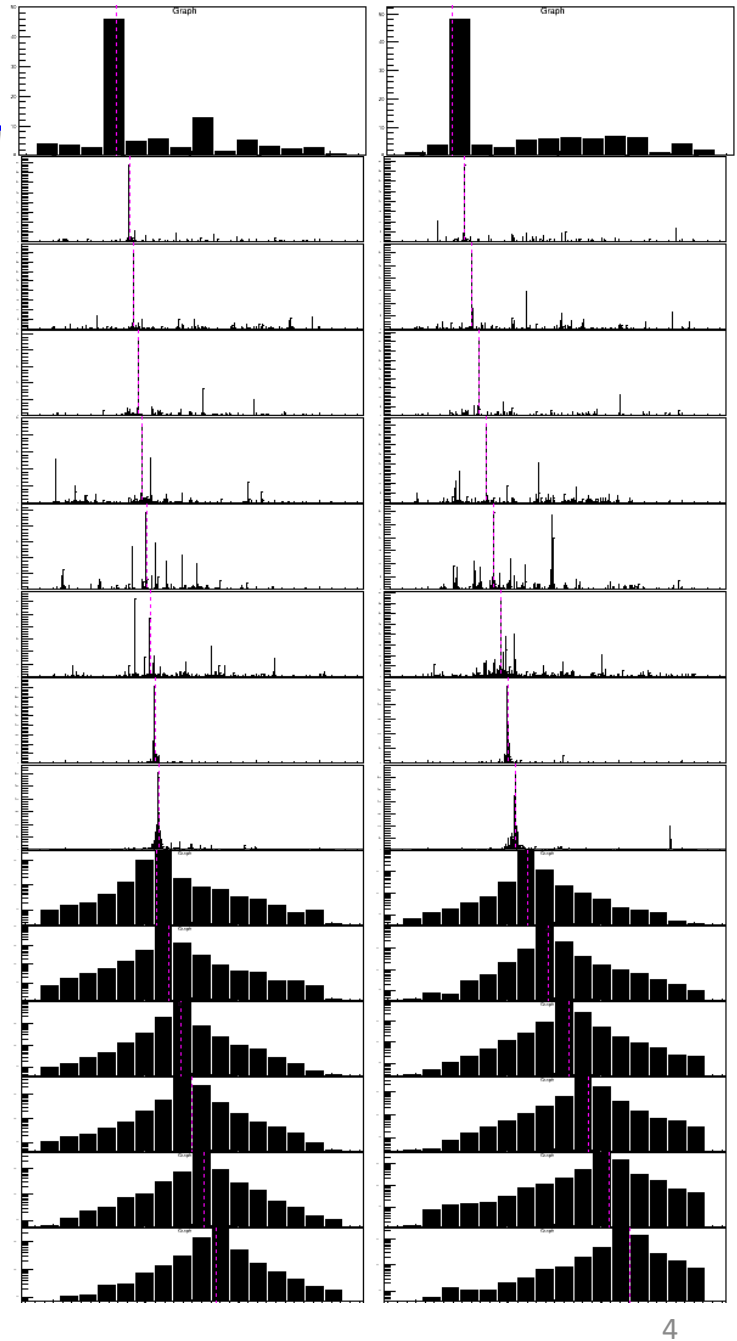
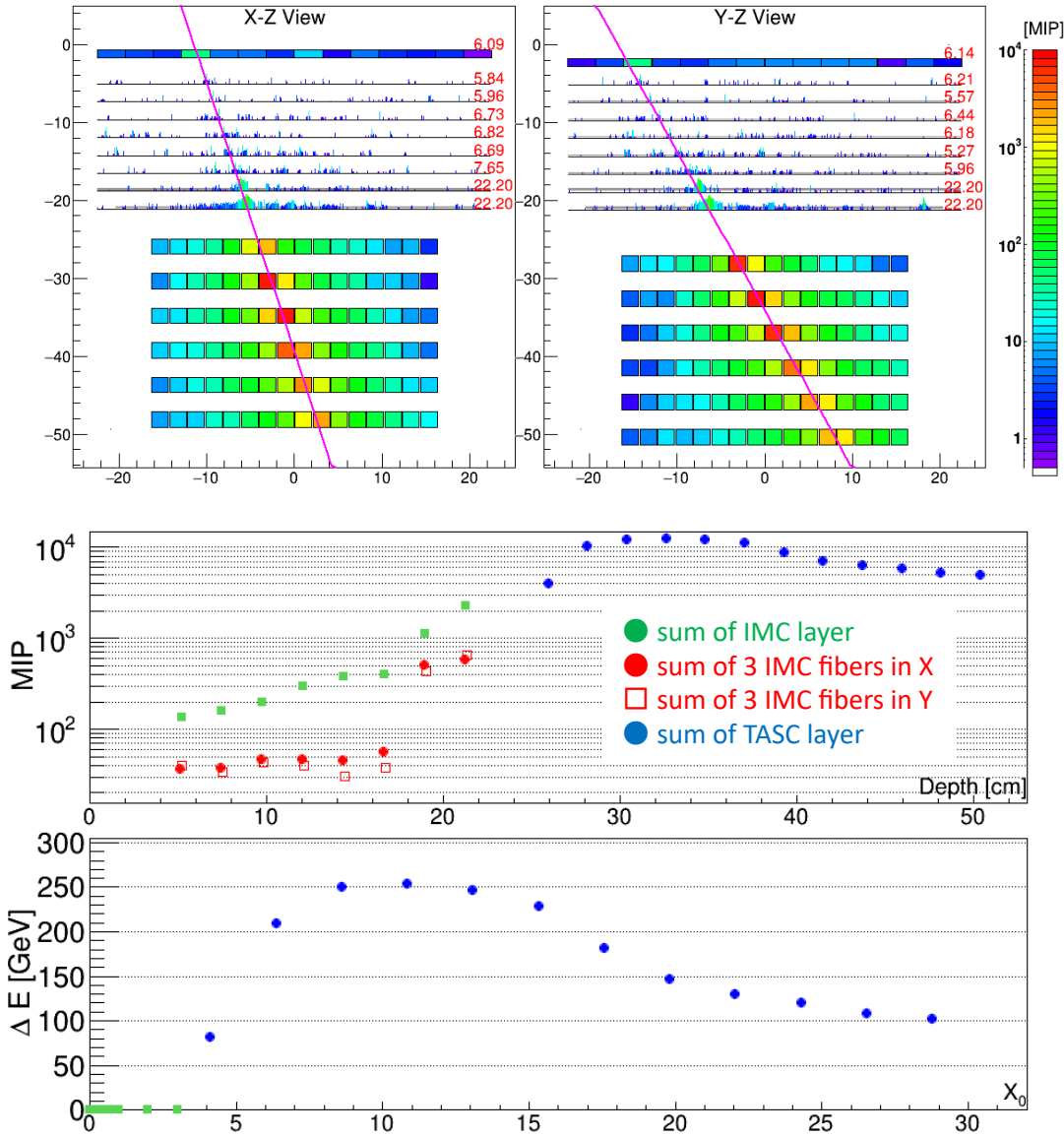
# Analysis Procedure of Nuclei

- HE trigger
  - 151013 – 190731 (1389 days)
  - FOV cut (for analysis of B/C)
- Offline Shower Trigger
  - $N_{\text{IMC-7X+8X}} > 50\text{MIP}$  &  $N_{\text{IMC-7Y+8Y}} > 50\text{MIP}$   
&  $N_{\text{TASC-X1}} > 100\text{MIP}$
- Tracking by UH track with Geom.B
- Charge consistency with CHD and IMC
  - $Z_{\text{CHD-Y}} < 1.10 Z_{\text{CHD-X}}$  &  $Z_{\text{CHD-Y}} > 1/1.10 Z_{\text{CHD-X}}$
  - $Z_{\text{IMC12}} < 1.15 Z_{\text{CHD}}$  &  $Z_{\text{CHD}} > 1/1.15 Z_{\text{IMC12}}$
  - $Z_{\text{IMC34}} < 1.15 Z_{\text{CHD}}$  &  $Z_{\text{CHD}} > 1/1.15 Z_{\text{IMC34}}$
- Track width selection
- Charge selection
  - $|Z_{\text{CHD}} - Z| < 0.4$
- Energy unfolding
- Flux calculation

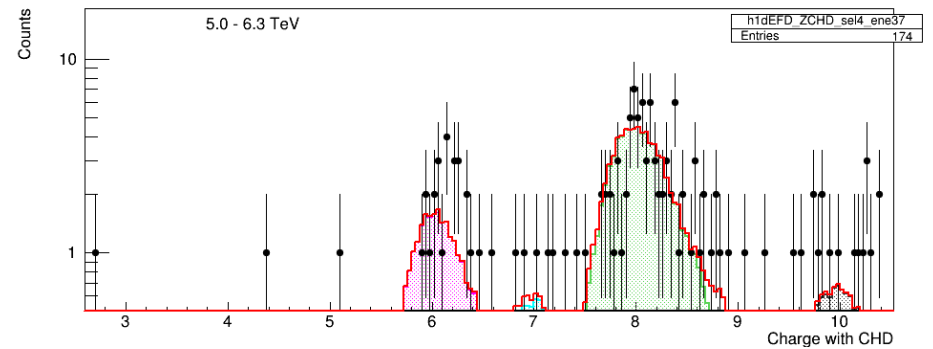
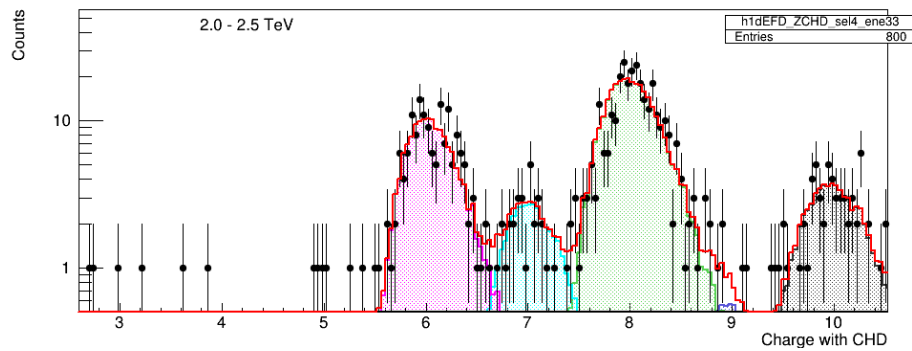
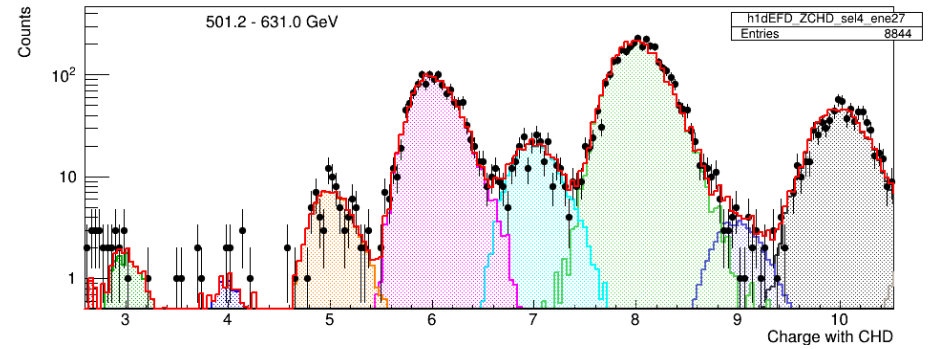
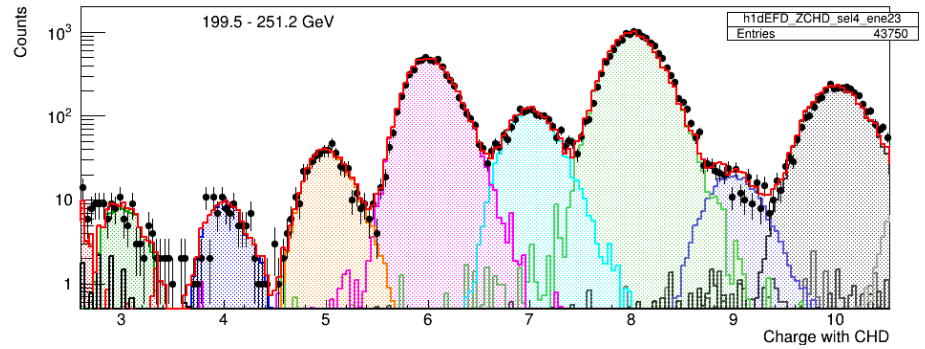
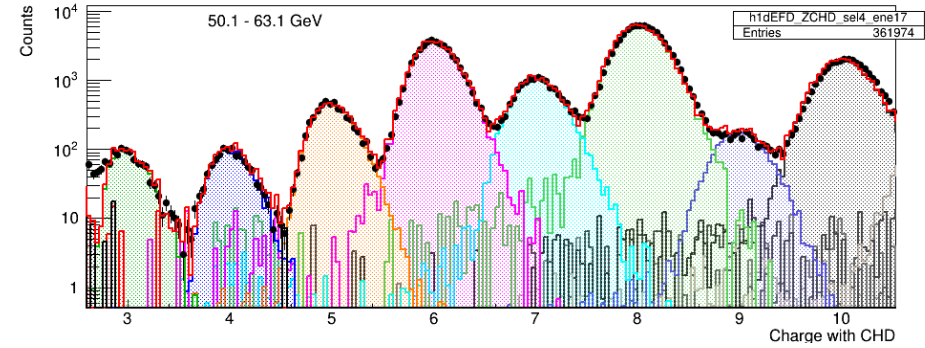
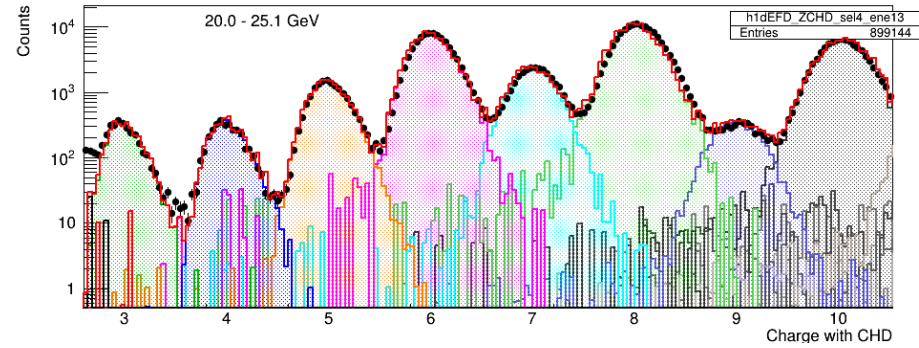


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MDTime 1163099894  
EventID: 040210  
 $\Delta E=2062.6\text{GeV}$

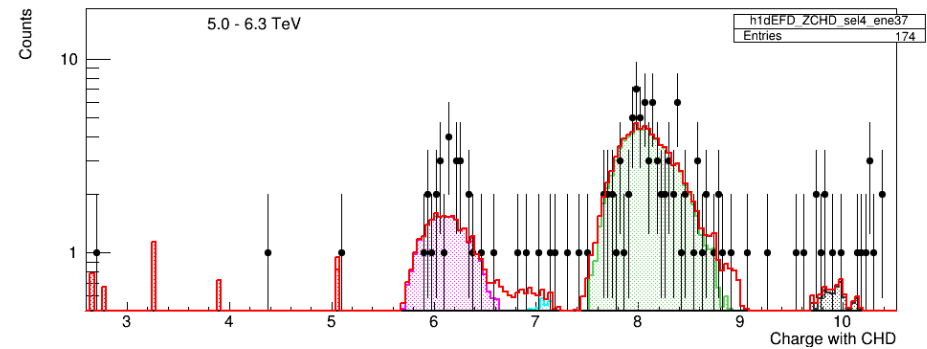
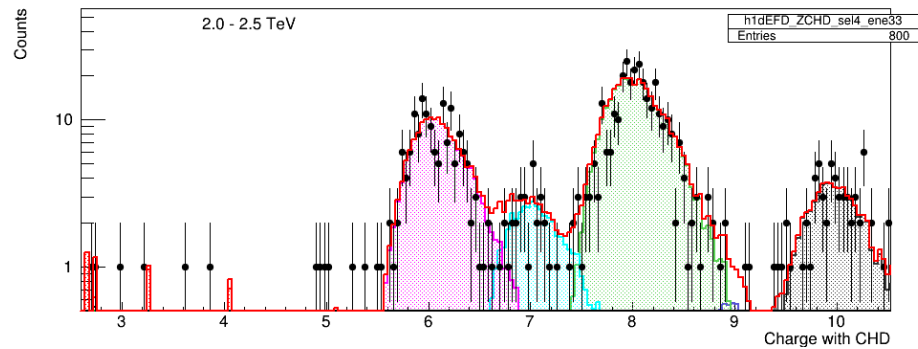
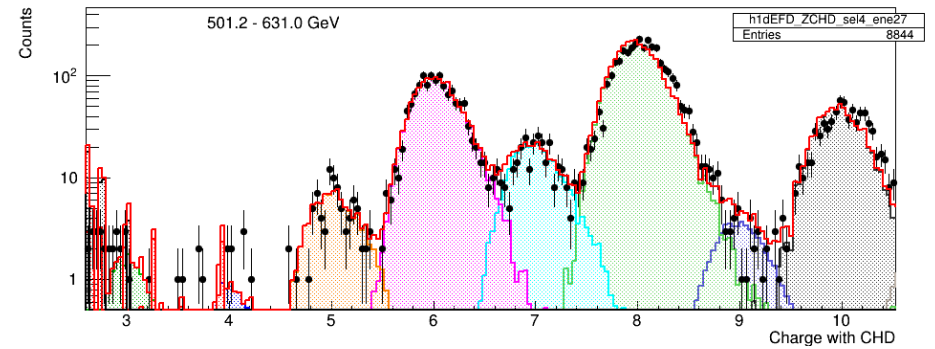
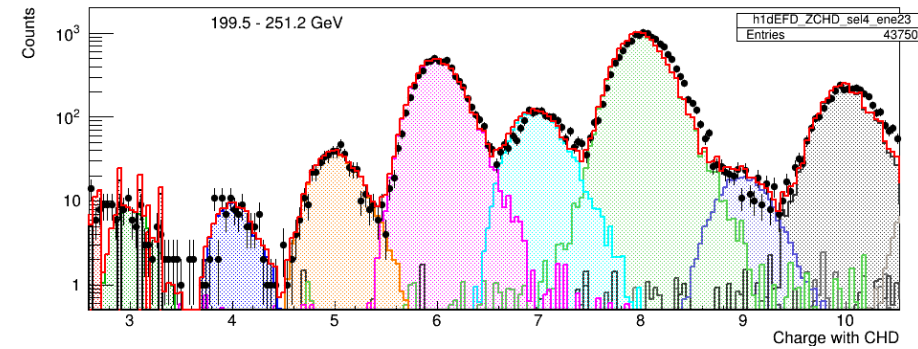
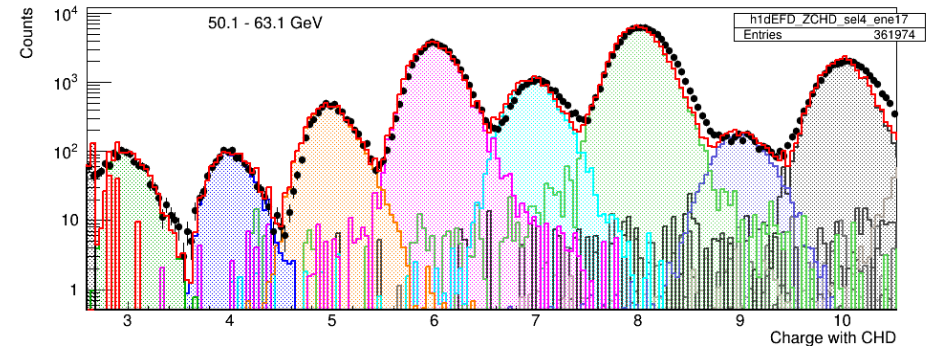
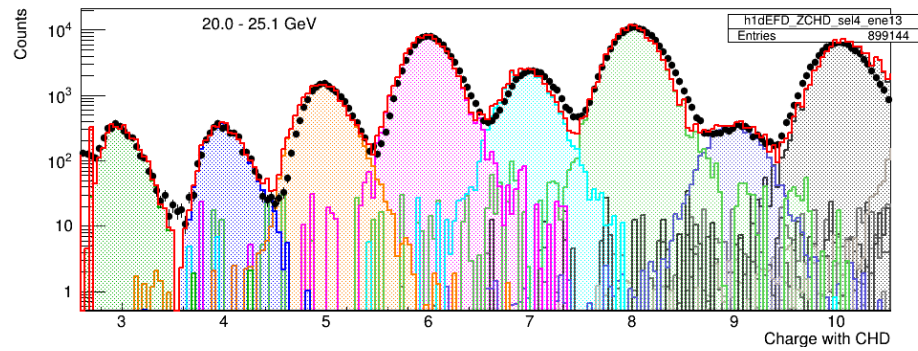
# Carbon event from flight data



# CHD Charge – EPICS –



# CHD Charge – FLUKA –



# Event selection: charge-consistency in CHD and IMC

## ● Charge identification

-  $Z_{\text{CHD}} = (Z_{\text{CHDX}} + Z_{\text{CHDY}})/2$

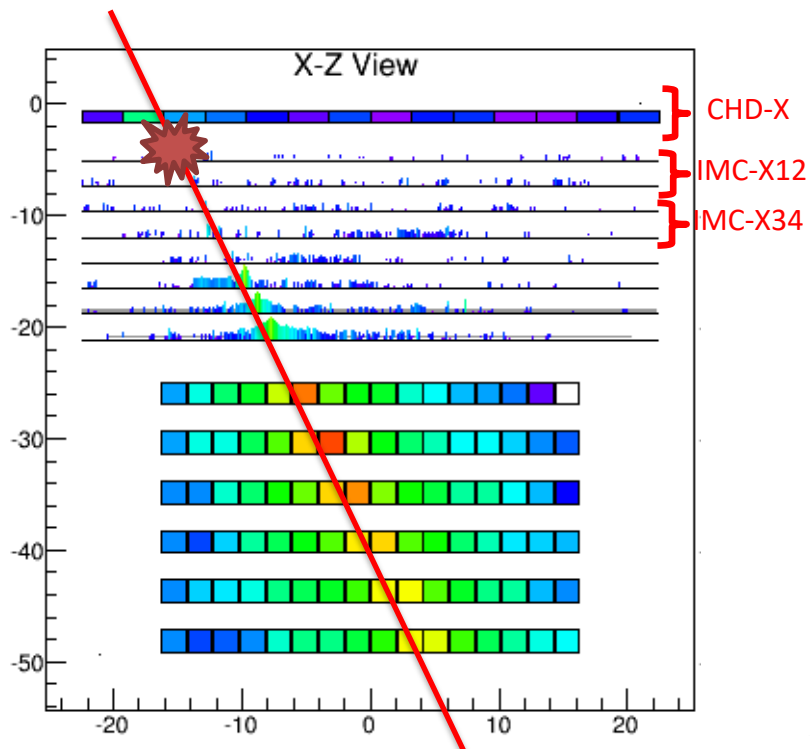
-  $Z_{\text{IMC}}$  : truncated mean of upper 4 layers

$$Z_{\text{IMC}12} = \frac{1}{2}(\sum_{i=1}^4 Z_{\text{IMC}i} - Z_{\text{IMC}i\text{MAX}} - Z_{\text{IMC}i\text{MIN}})$$

## ● Background source

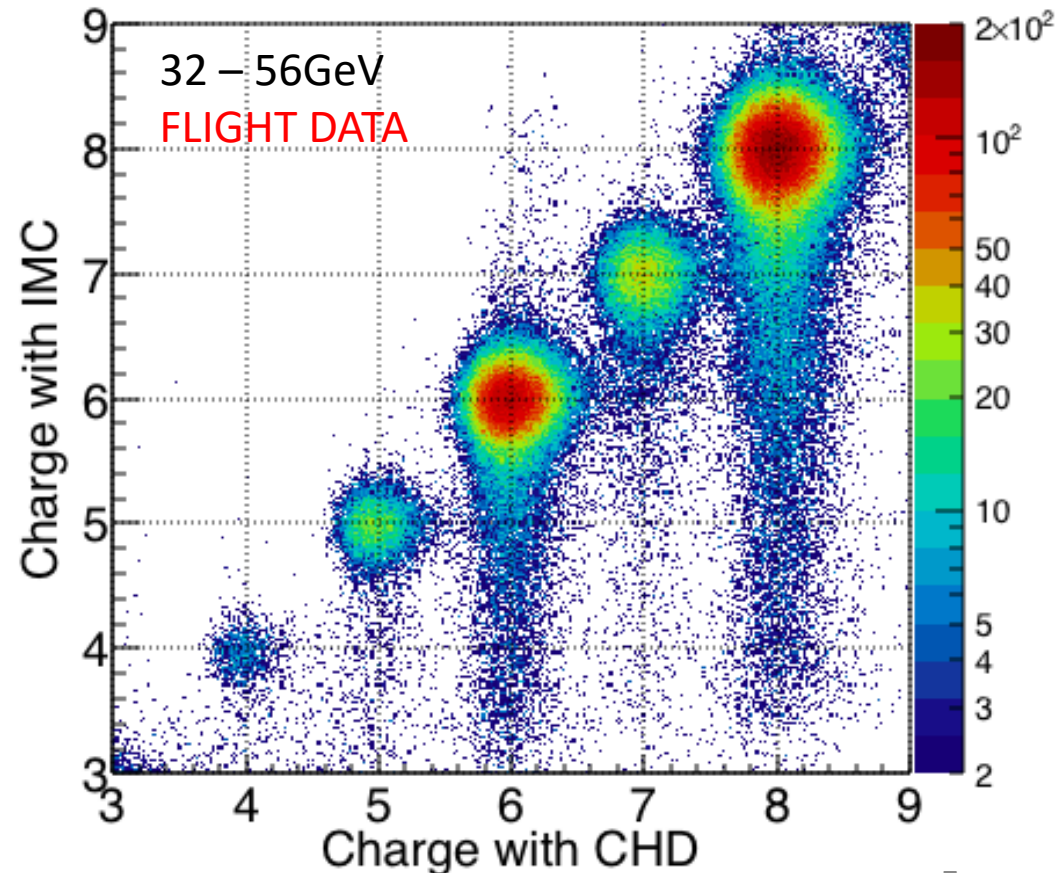
- mis-reconstructed events

- interacted events in CHD or upper IMC layers



## Pre-selection

- HE trigger
- Tracking + geometrical condition
- Z-consistency with CHD-X and CHD-Y



# Event selection: charge-consistency in CHD and IMC

## ● Charge identification

-  $Z_{\text{CHD}} = (Z_{\text{CHDX}} + Z_{\text{CHDY}})/2$

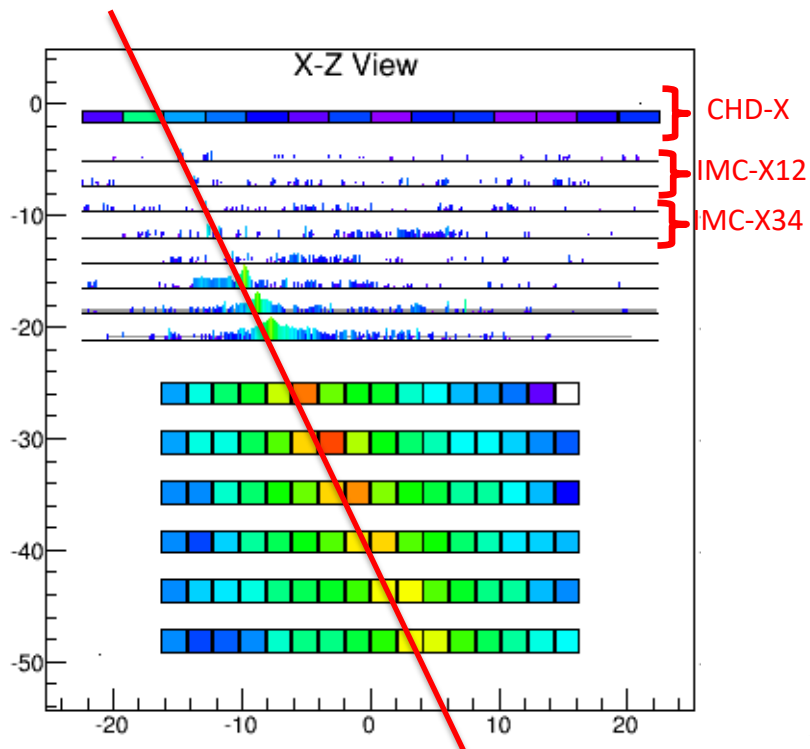
-  $Z_{\text{IMC}}$  : truncated mean of upper 4 layers

$$Z_{\text{IMC12}} = \frac{1}{2} \left( \sum_{i=1}^4 Z_{\text{IMCi}} - Z_{\text{IMCiMAX}} - Z_{\text{IMCiMIN}} \right)$$

## ● Background source

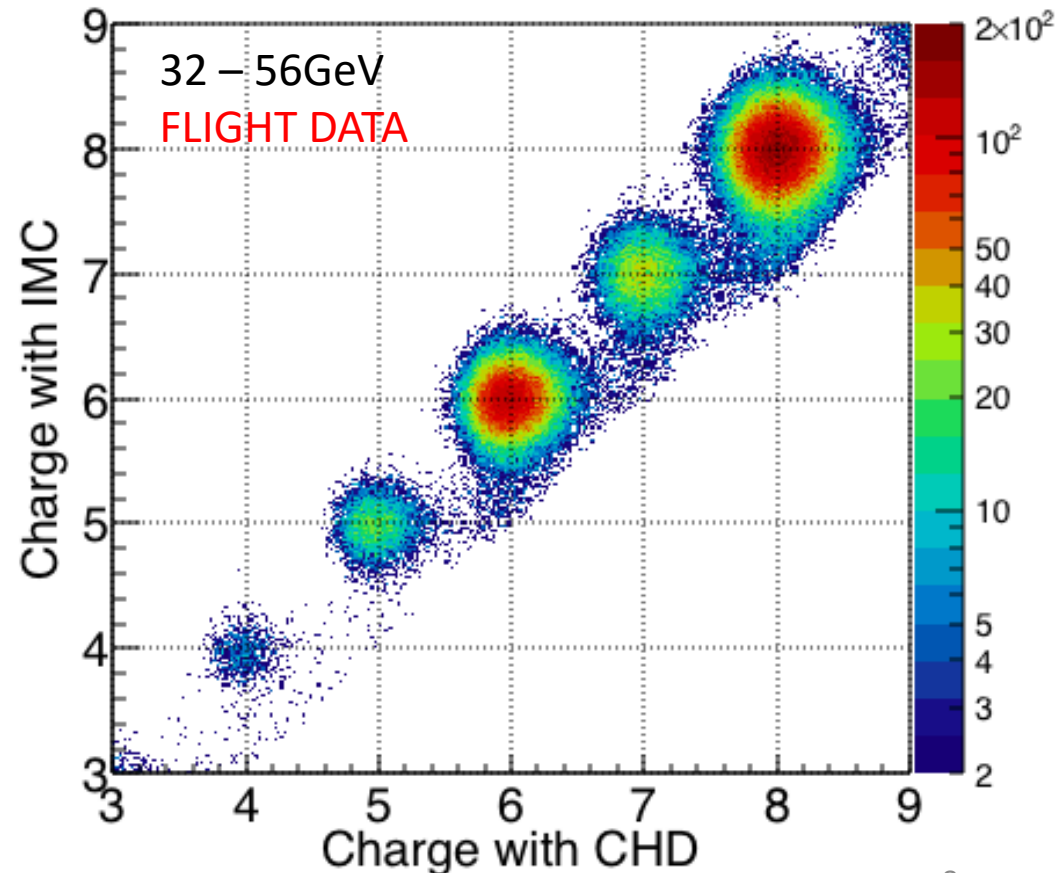
- mis-reconstructed events

- interacted events in CHD or upper IMC layers



## Pre-selection

- HE trigger
- Tracking + geometrical condition
- Z-consistency with CHD-X and CHD-Y
- Z-consistency with IMC12 and IMC34





# Event selection: track width in IMC

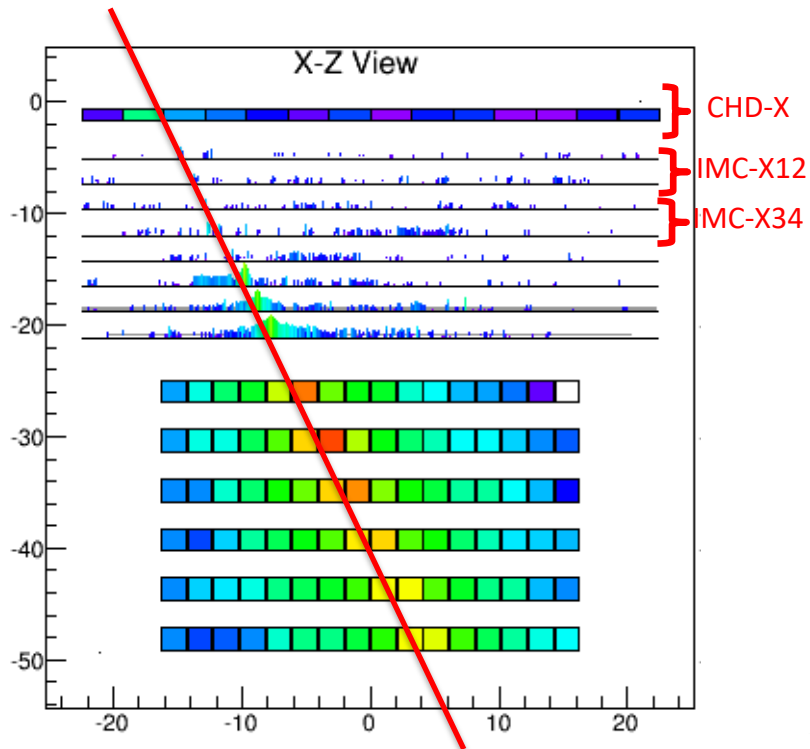
## ● Charge identification

- $Z\_CHD = (ZCHDX + ZCHDY)/2$
- $Z\_IMC$  : truncated mean of upper 4 layers

$$Z_{IMC12} = \frac{1}{2} \left( \sum_{i=1}^4 Z_{IMCi} - Z_{IMC_{iMAX}} - Z_{IMC_{iMIN}} \right)$$

## ● Background source

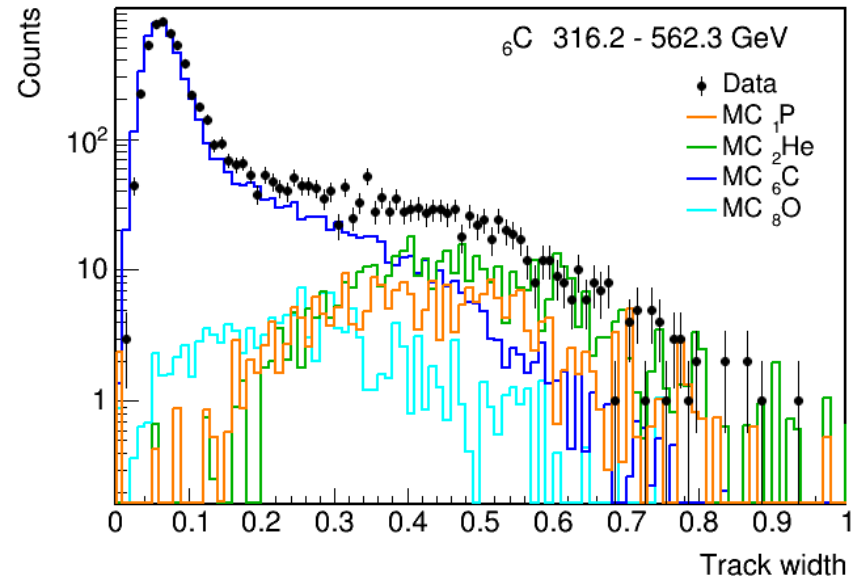
- mis-reconstructed events
- interacted events in CHD or upper IMC layers



## Pre-selection

- HE trigger
- Tracking + geometrical condition
- Z-consistency with CHD-X and CHD-Y
- Z-consistency with IMC12 and IMC34
- Track width selection

$$B_{IMCi} = \left( \underbrace{\sum_{j=-k}^k N_{IMCi,j}}_{\text{Sum of 7 SciFis}} - \underbrace{\sum_{j=-1}^1 N_{IMCi,j}}_{\text{Sum of 3 SciFis}} \right) \frac{1}{Z_{IMCi}^2}$$



# Charge identification and Background estimation

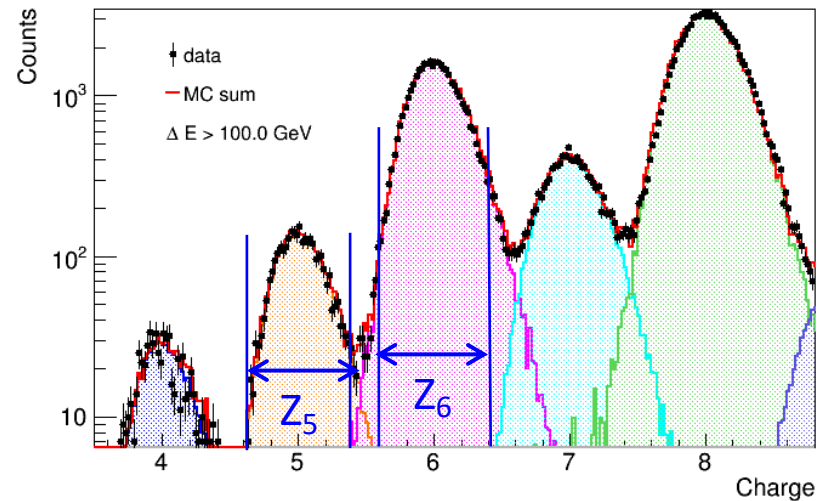
- Particle charge is identified with CHD
- Background is estimated by means of MC

MC data:

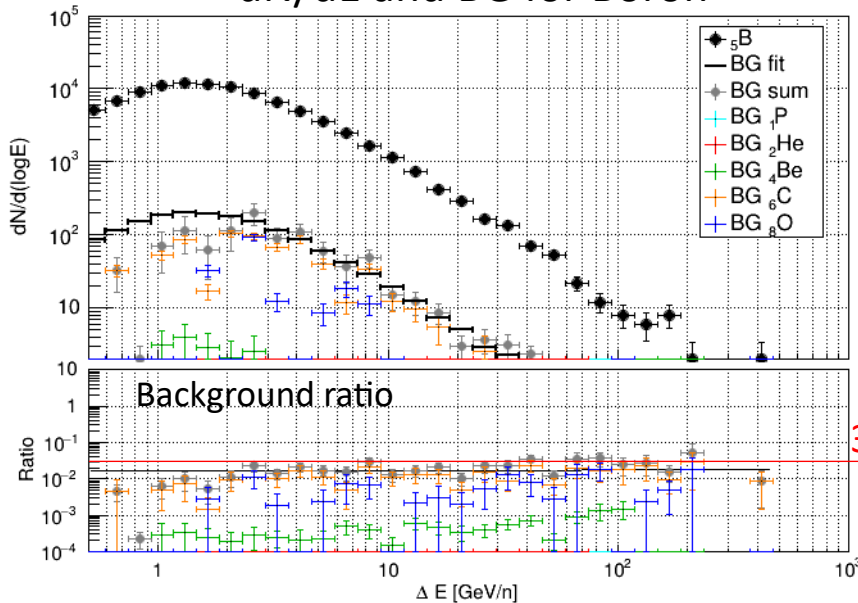
- EPICS v9.21 (Cosmos8.01)
- DPMJET-III

Consider quenching, noise and etc.

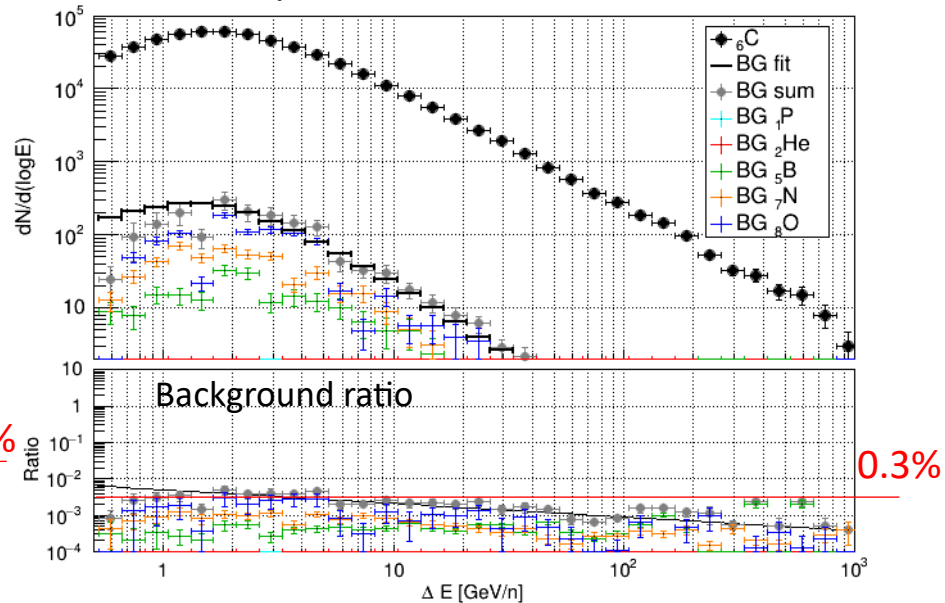
Apply the same selection with flight data.



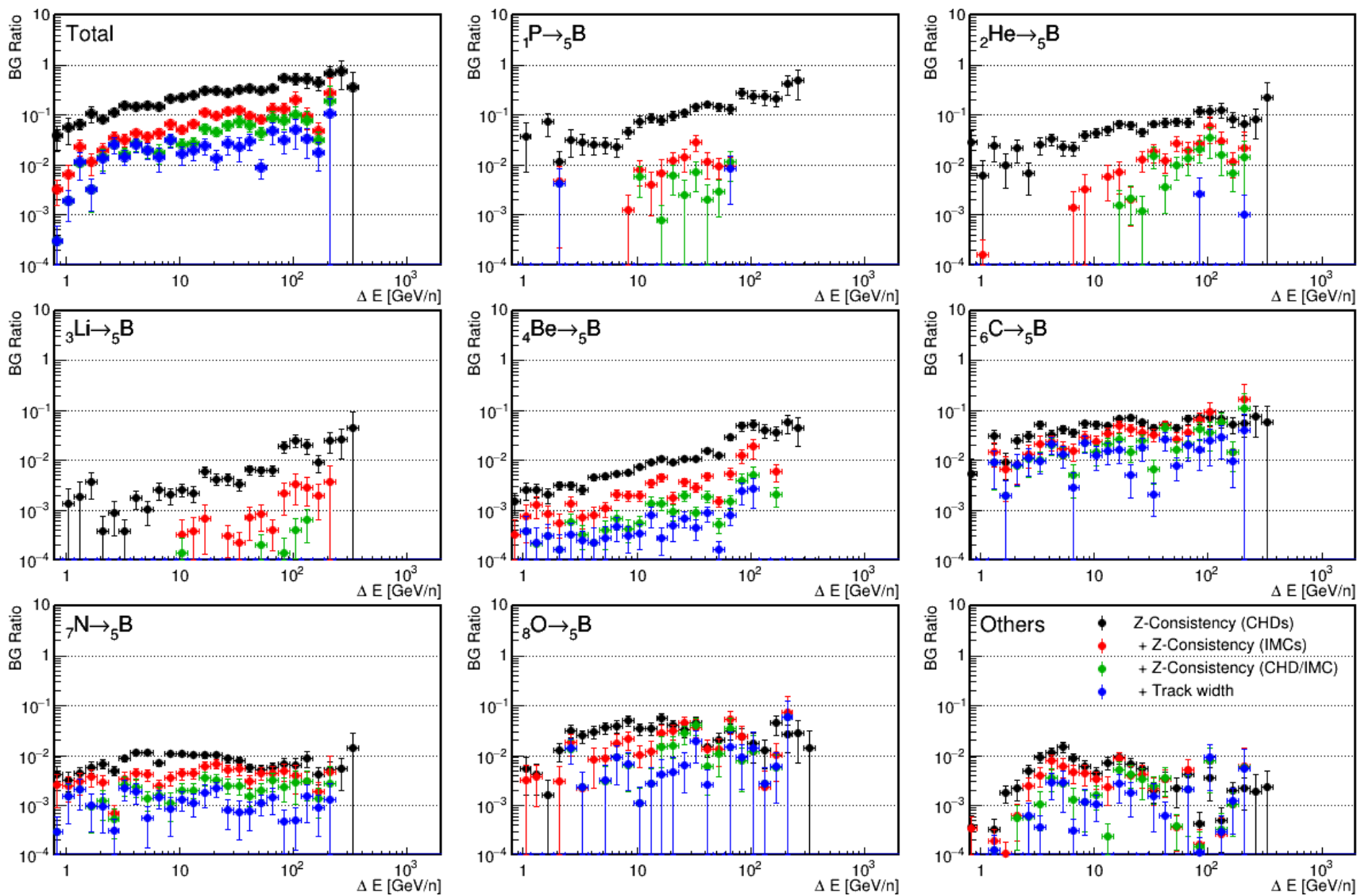
dN/dE and BG for Boron



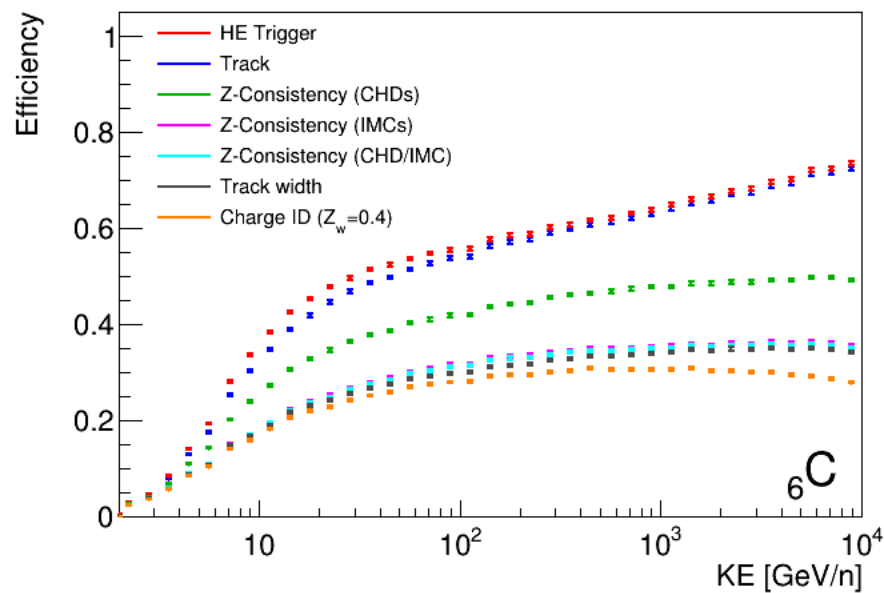
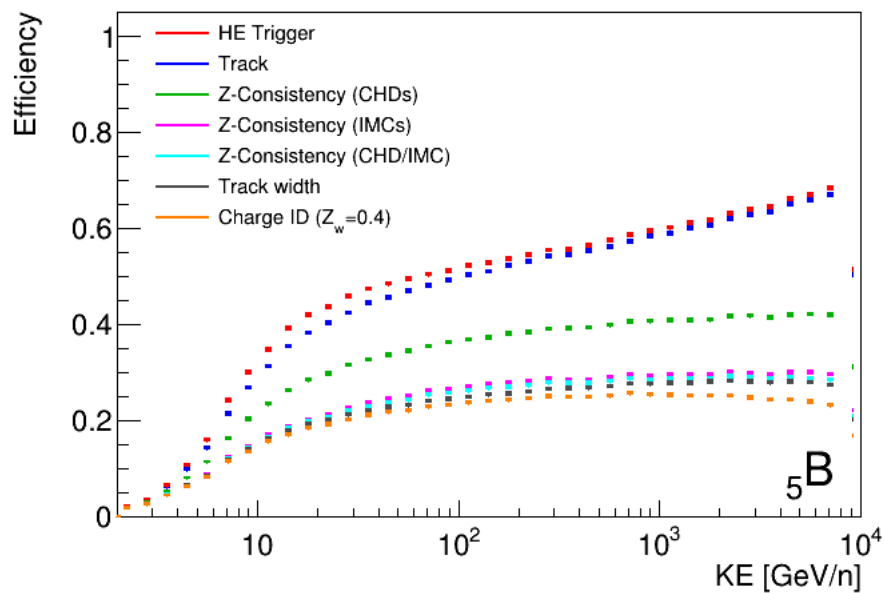
dN/dE and BG for Carbon



# Background components for Boron



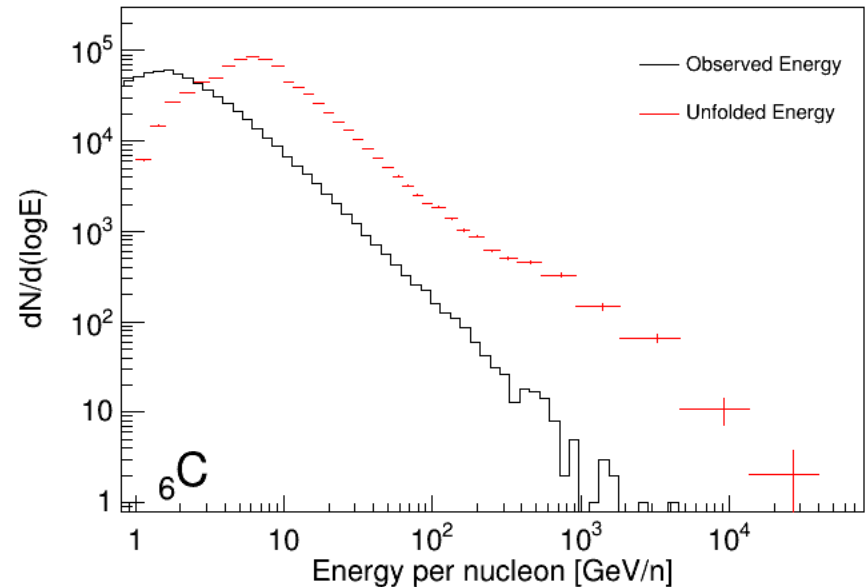
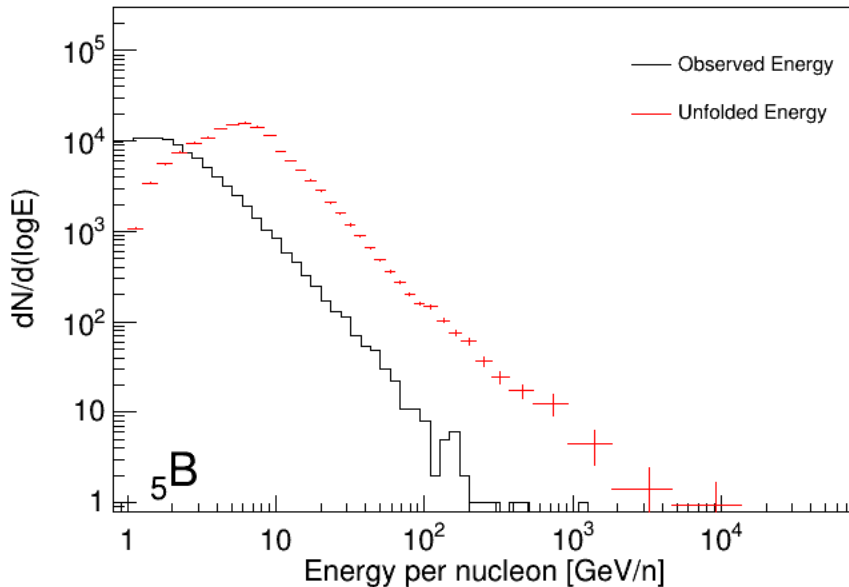
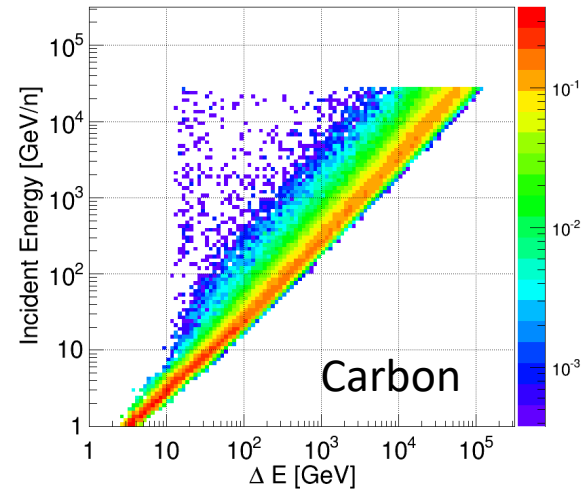
# Efficiency



# Energy unfolding

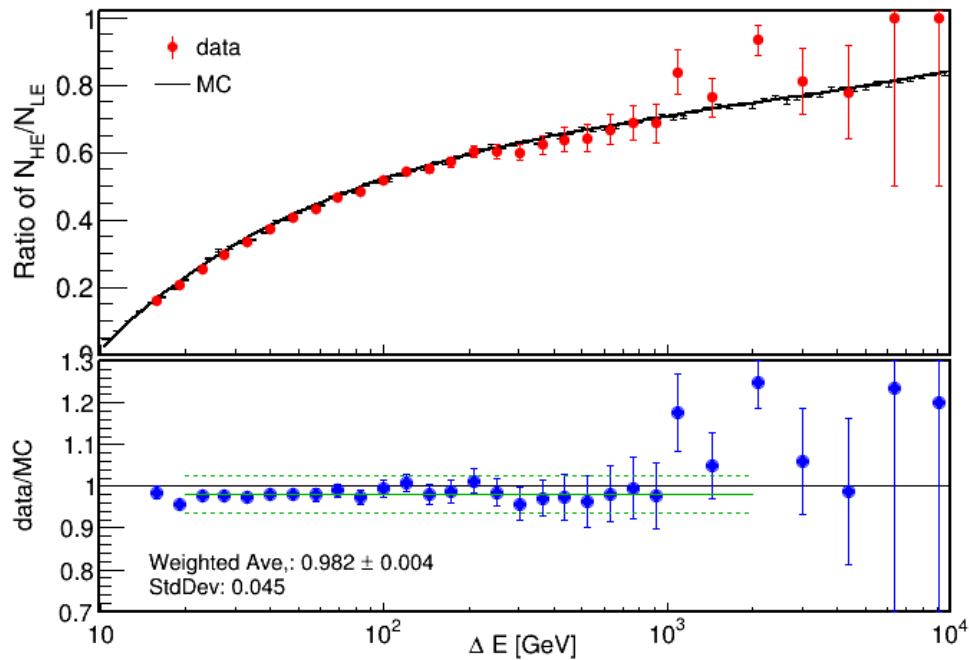
- Iterative Bayesian unfolding

- Initial assuming spectra:  $f(E)=A \times E^{-2.60}$   
A is normalized by charge distribution in CHD
- Response function:  
 $\Delta E$  [GeV] vs  $E_0$  [GeV/n]
- 2 iterations



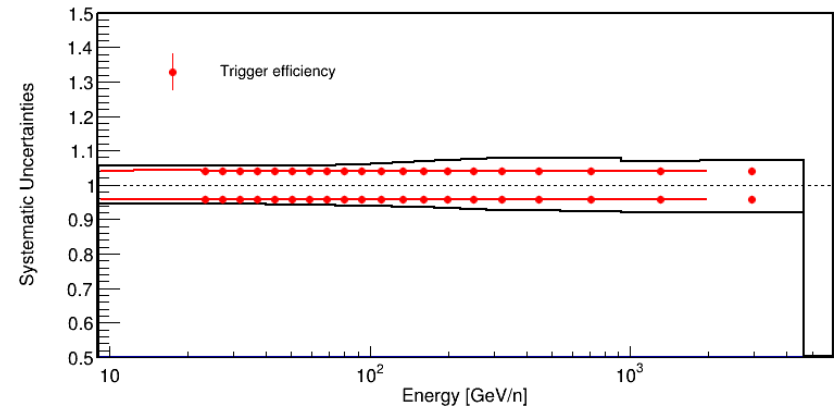
# Systematic uncertainty: Trigger efficiency

$N_{LE\&HE} / N_{LE}$  for Carbon

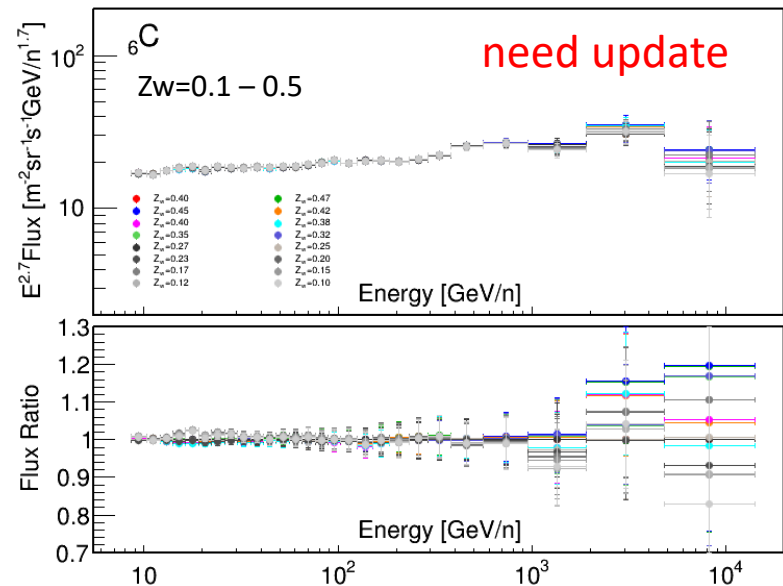
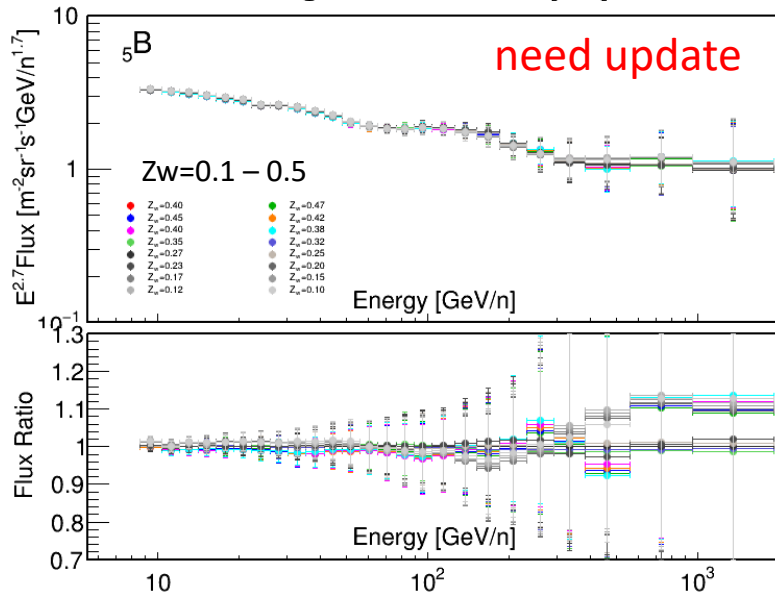
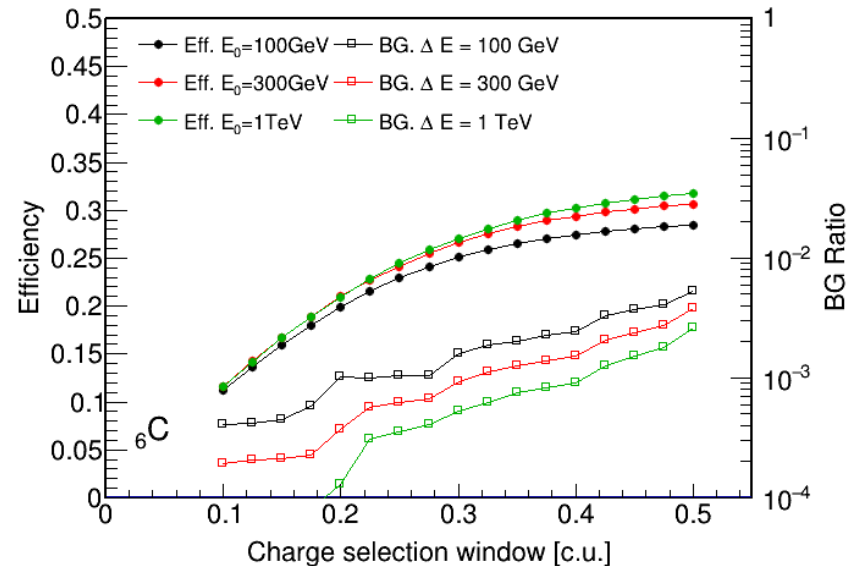
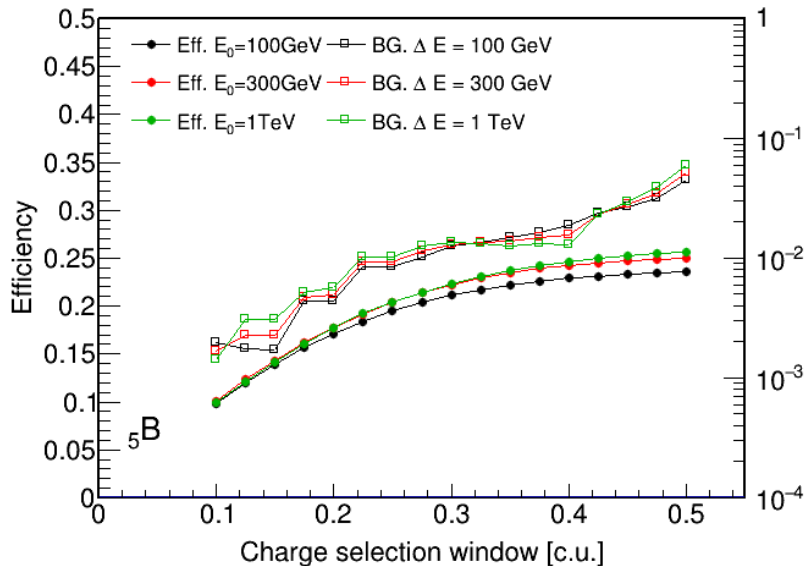


For Boron, same value with carbon is used.

Systematic uncertainty of B/C



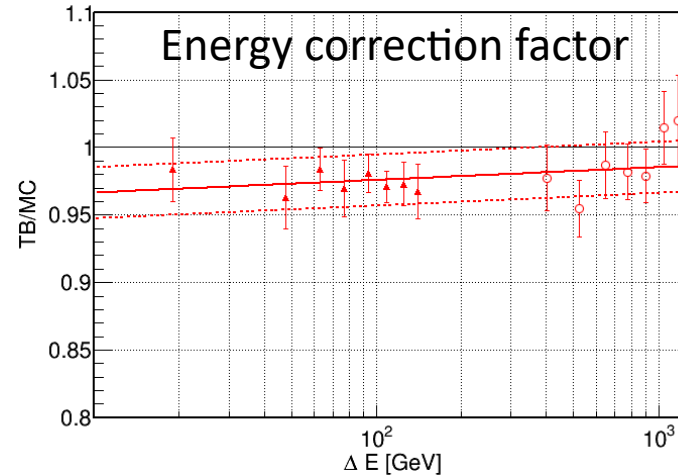
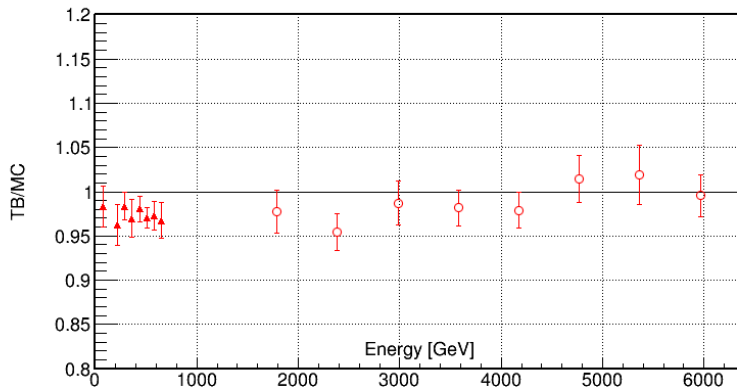
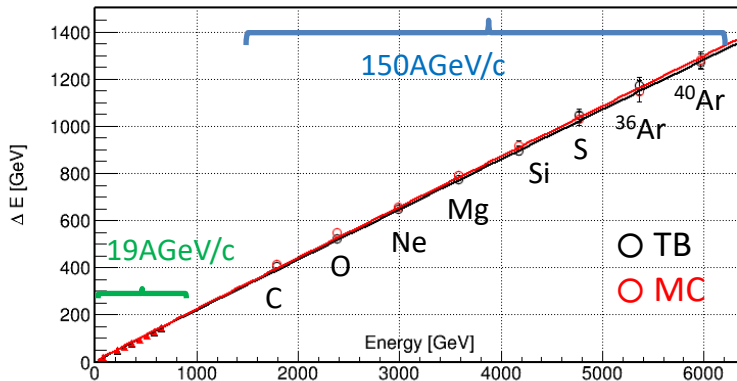
# Systematic uncertainty: Charge selection



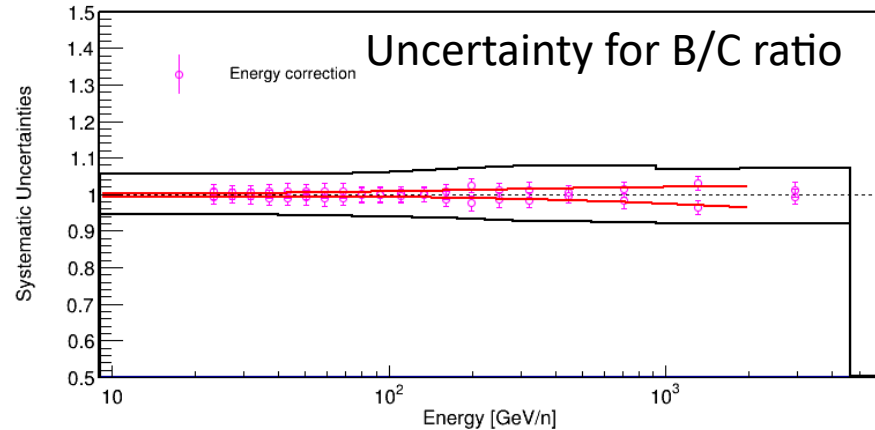
# Systematic uncertainty: Energy scale

Results of CERN 2015 beam test

Deposit energy vs incident energy

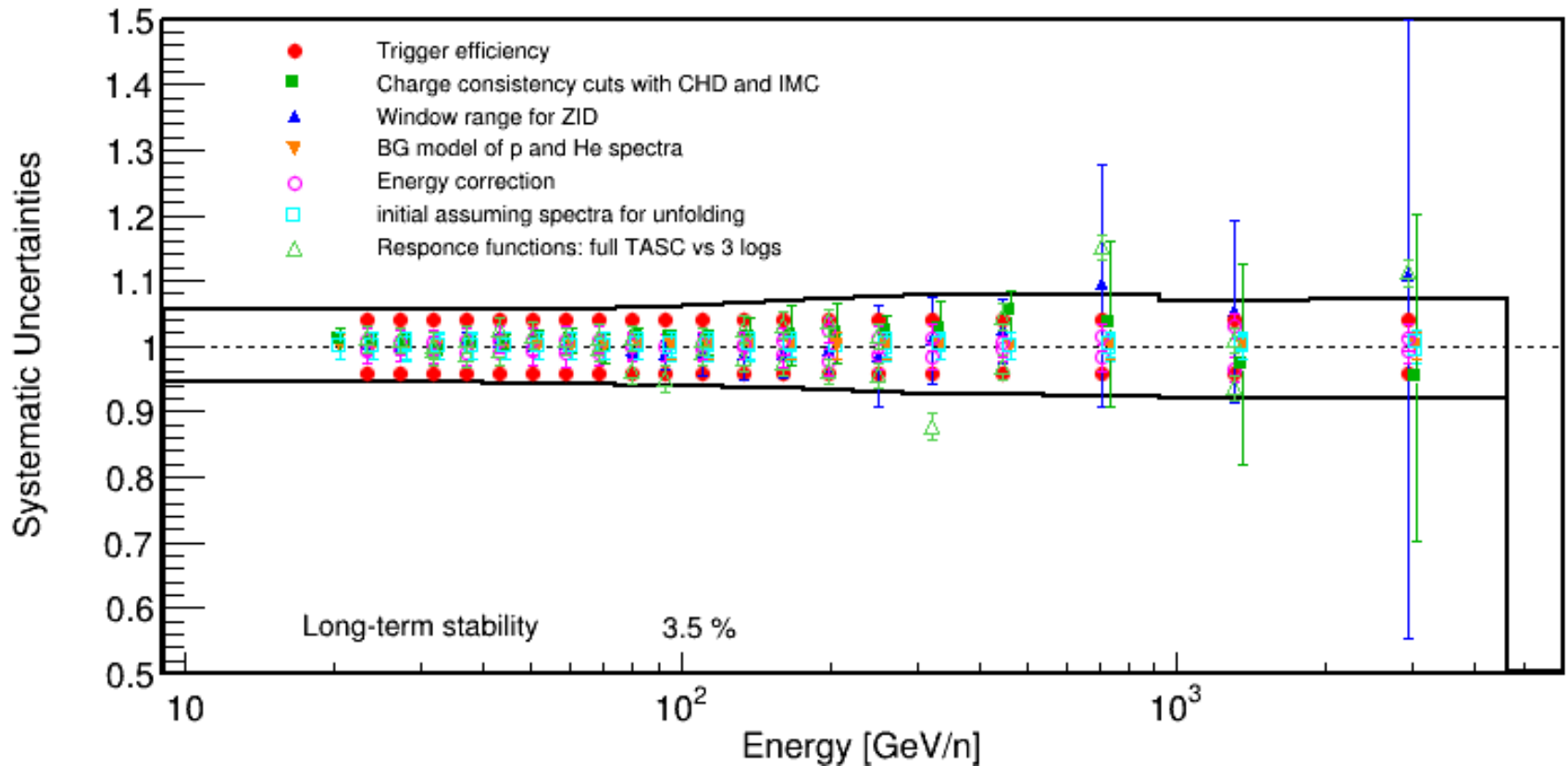


- Apply the correction to flight data analysis
- Error of energy correction is considered



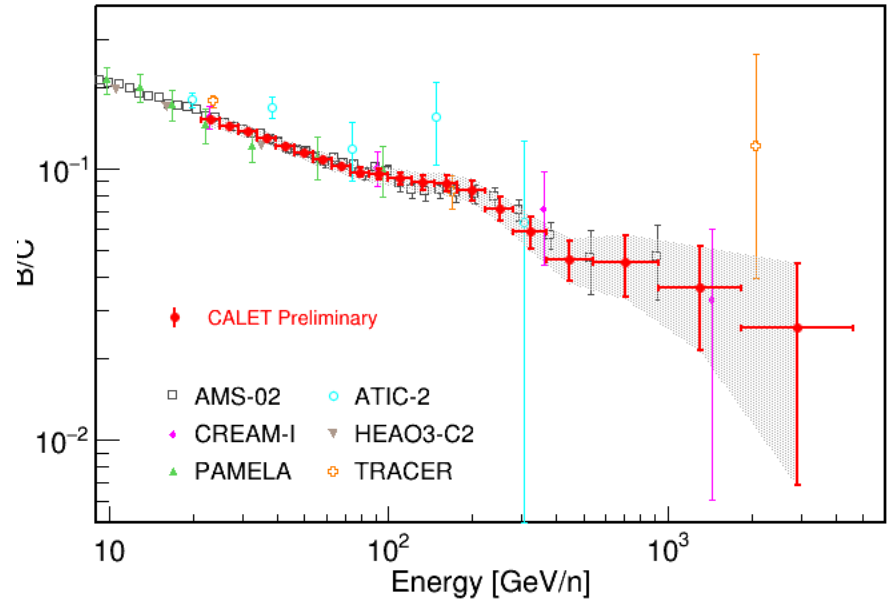
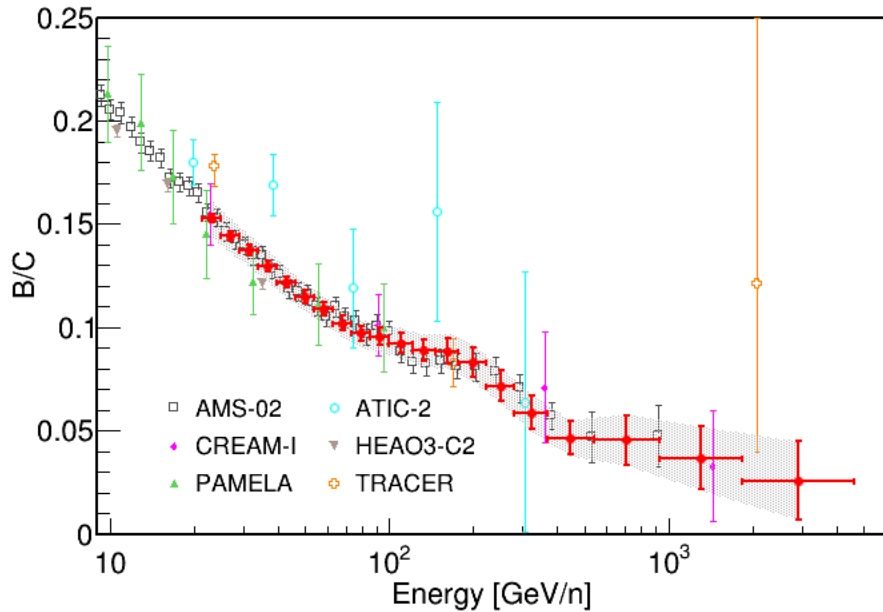


# Systematic uncertainties for B/C ratio

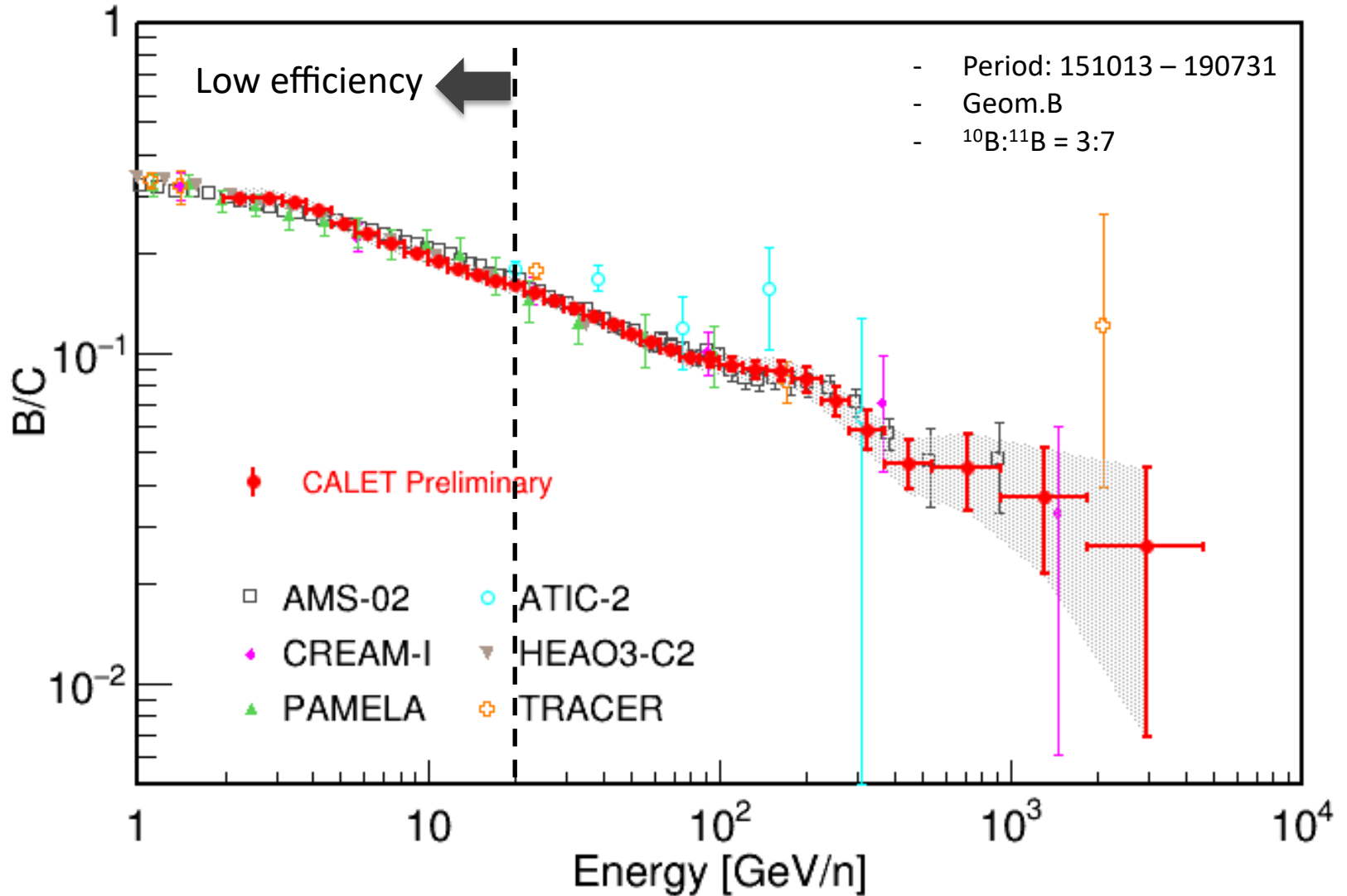


# Boron-to-carbon ratio

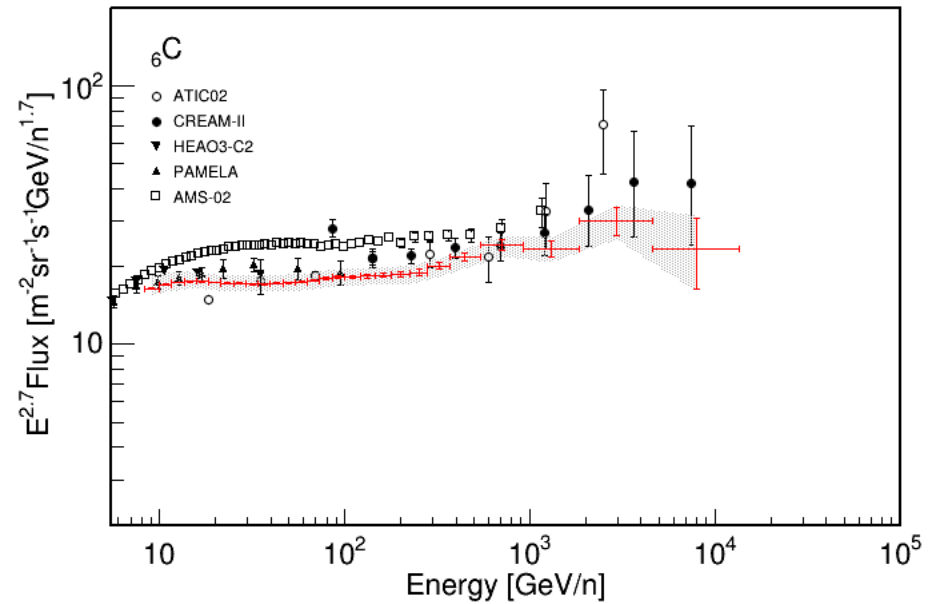
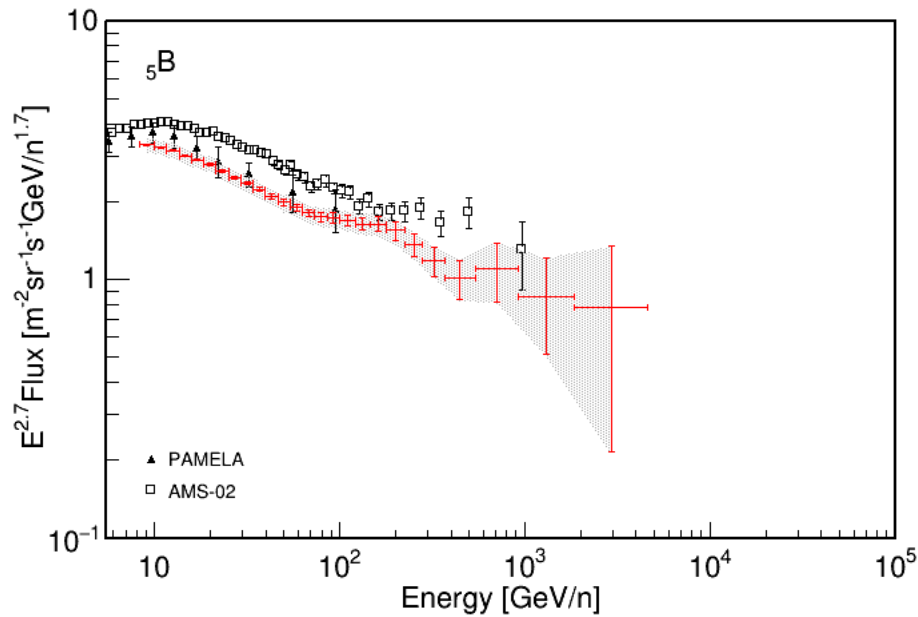
- Period: 151013 – 190731
- Geom.B
- $^{10}\text{B}:^{11}\text{B} = 3:7$



# B/C ratio

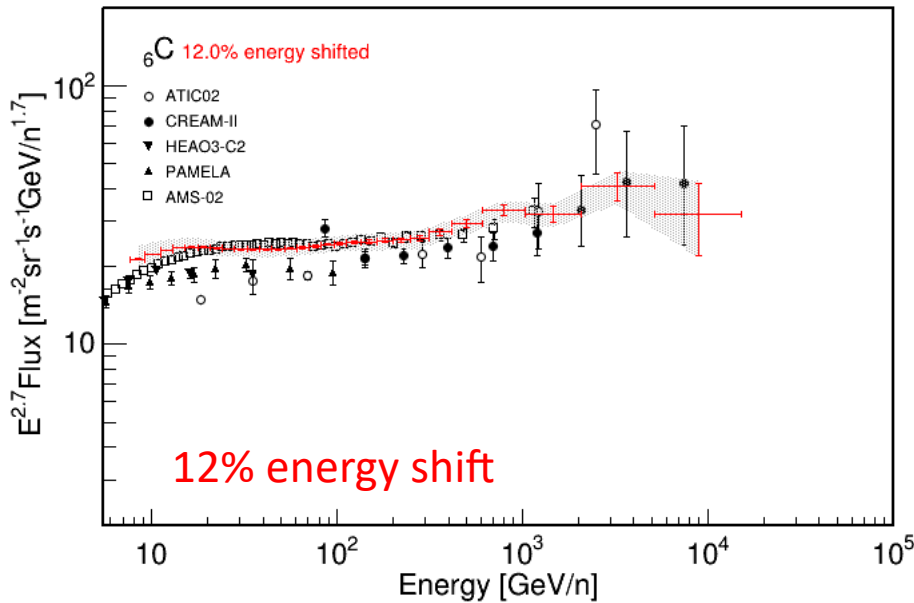


# Boron and Carbon spectra



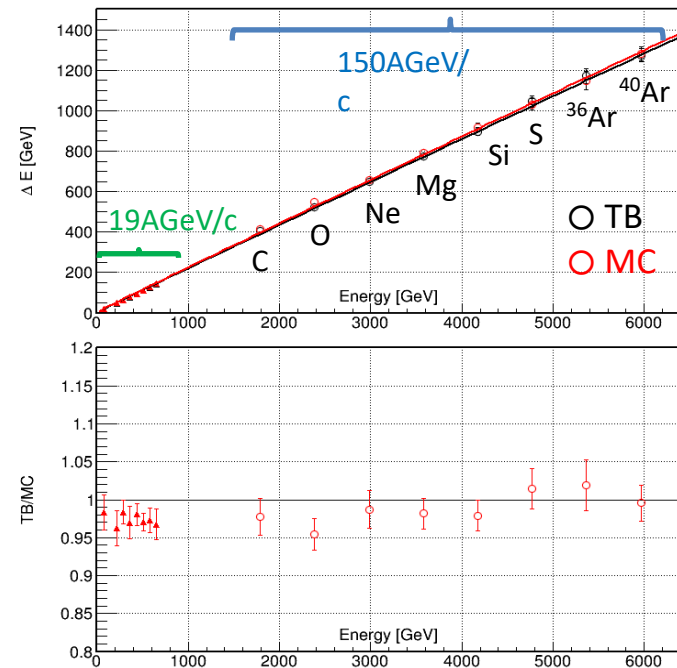
# Compare to AMS-02

## ■ Energy correction?



## Results of CERN 2015 beam test

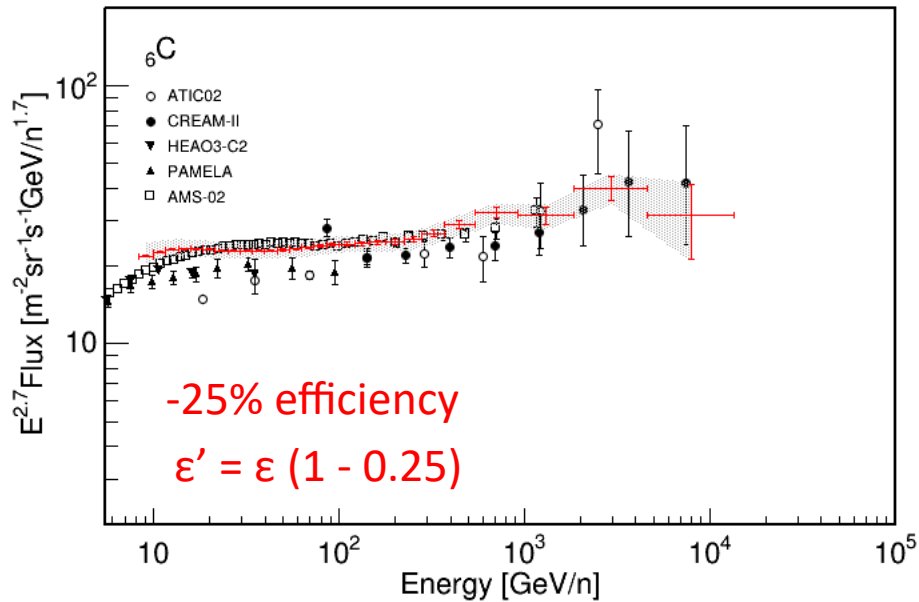
### Deposit energy vs incident energy



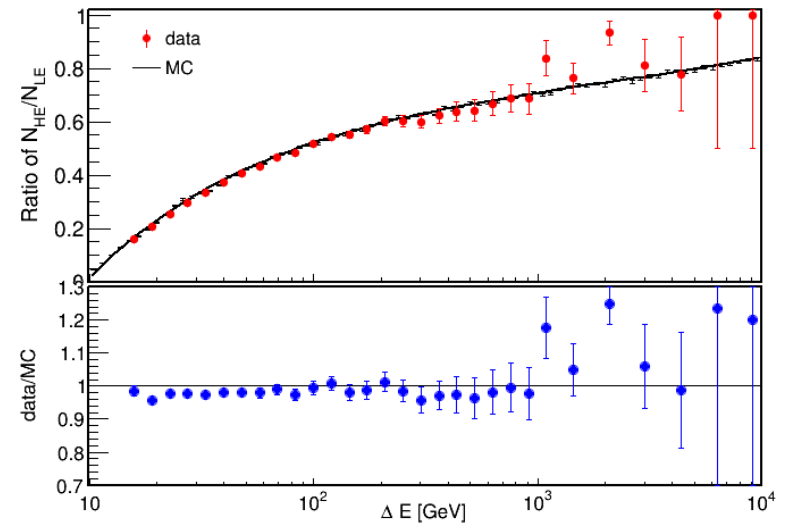
CERN 2015 beam test: ~a few %

# Compare to AMS-02

## ■ Efficiency correction?



## $N_{LE\&HE} / N_{LE}$ for Carbon



X-axis is deposit energy in TASC, not primary energy  
Does this study really confirm the absolute efficiency?

# Trigger efficiency and cross-section

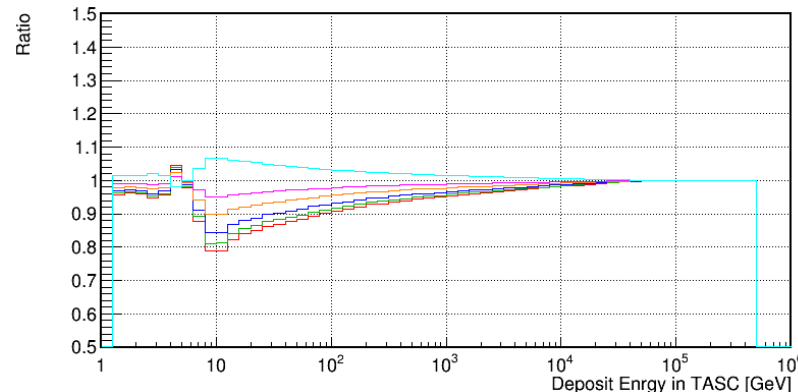
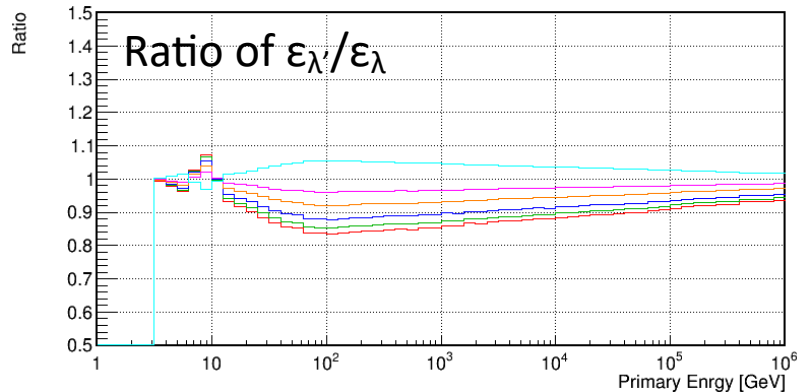
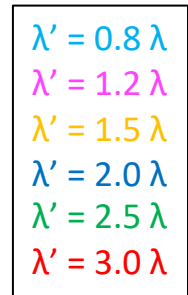
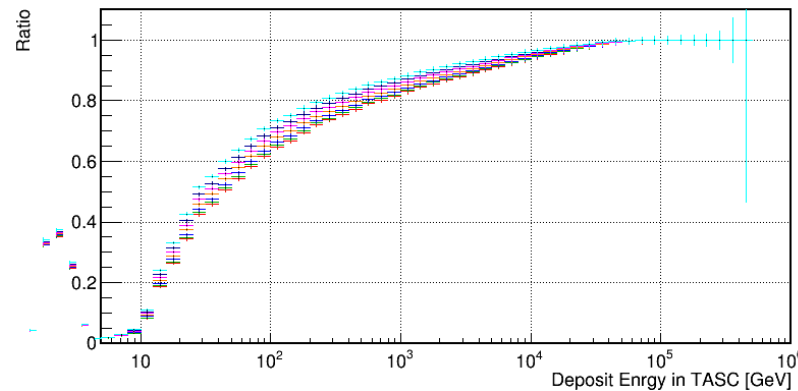
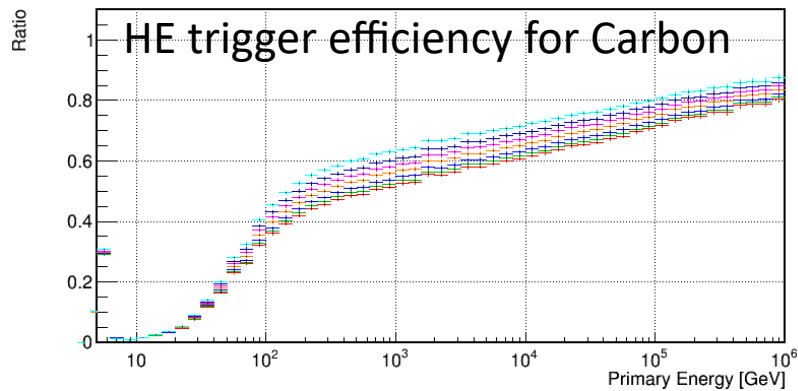
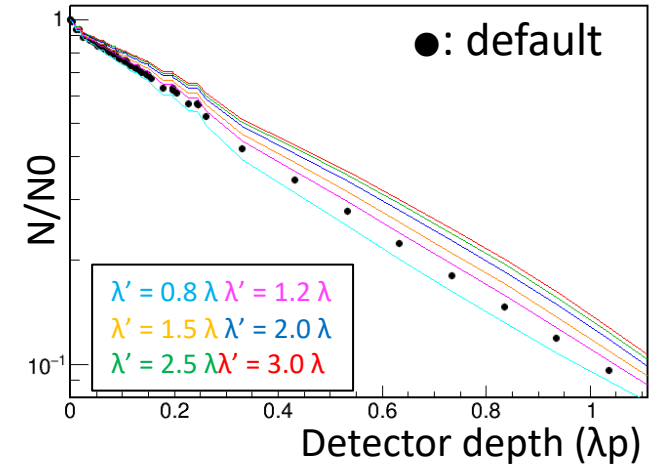
Trigger efficiency is depending on the cross-section.  
 To study the difference of the cross-section in MC,  
 MC events are reweighted by the interaction point

$$w = \exp(-t/\lambda) / \exp(-t/\lambda'),$$

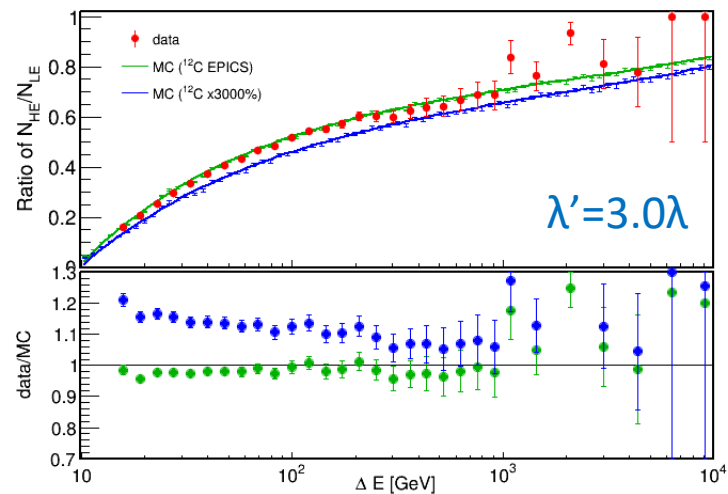
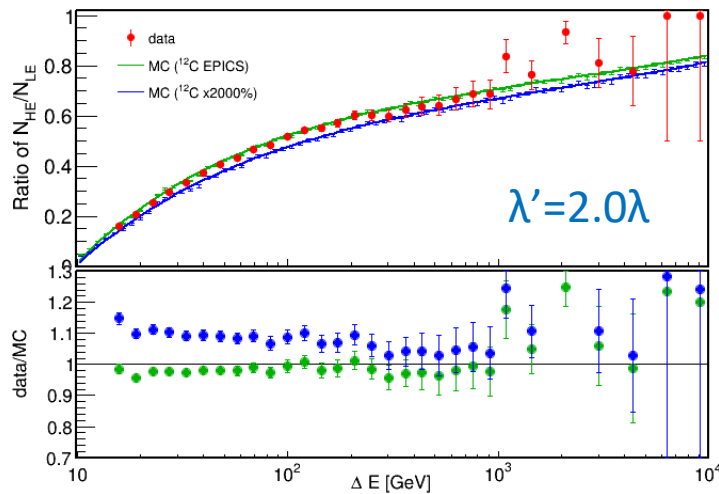
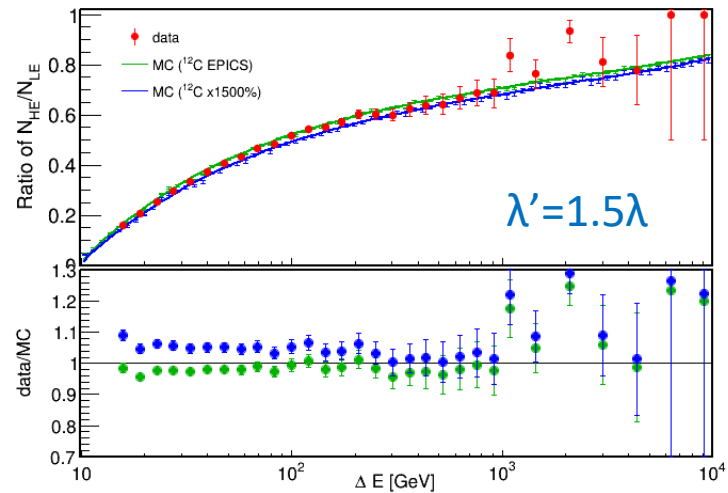
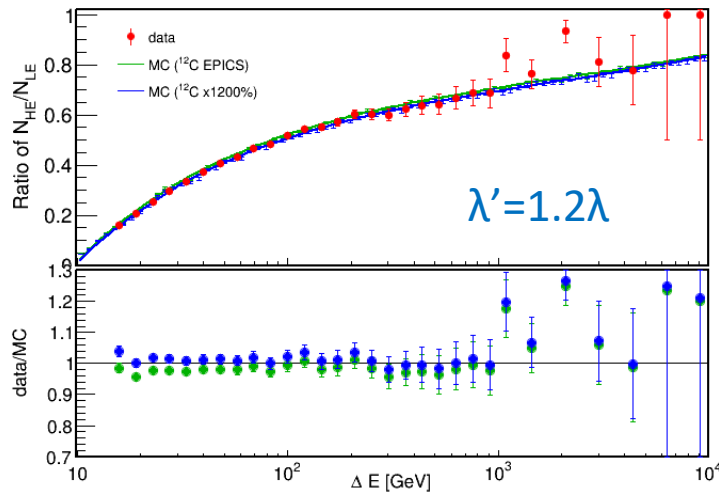
$t$ : depth of 1<sup>st</sup> interaction point (MC true info)

$\lambda$ : interaction length

$\lambda'$ : modified interaction length



# Trigger efficiency with LE triggered events



$\lambda' > 1.5\lambda$  is inconsistent with flight data.

Adjusting cross-section cannot explain the difference from AMS-02

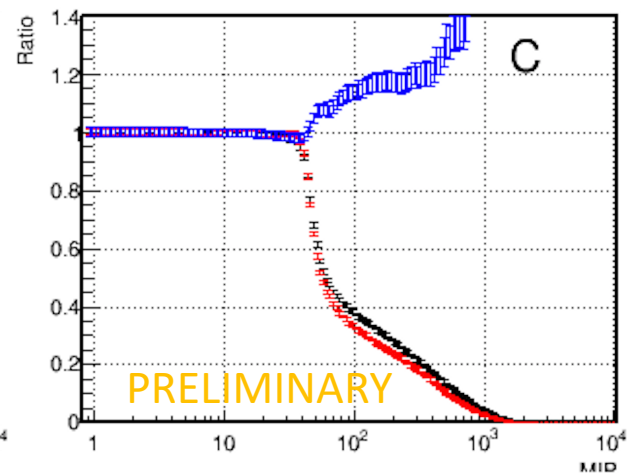
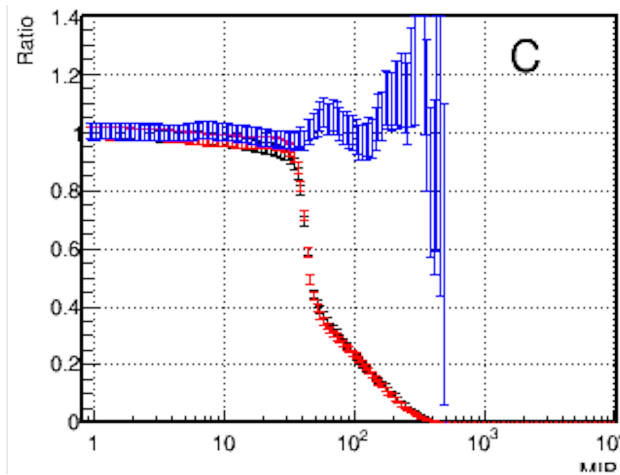
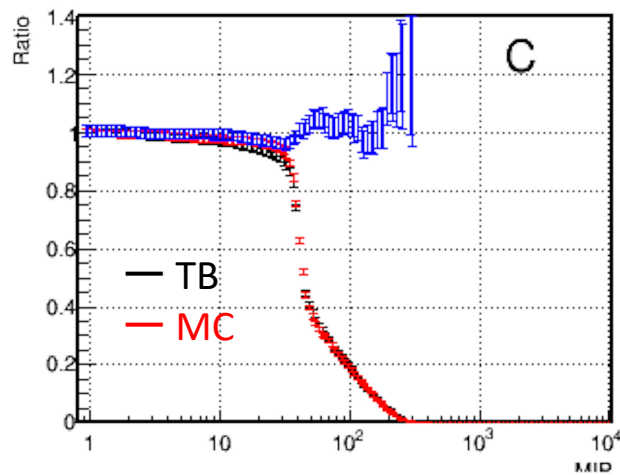
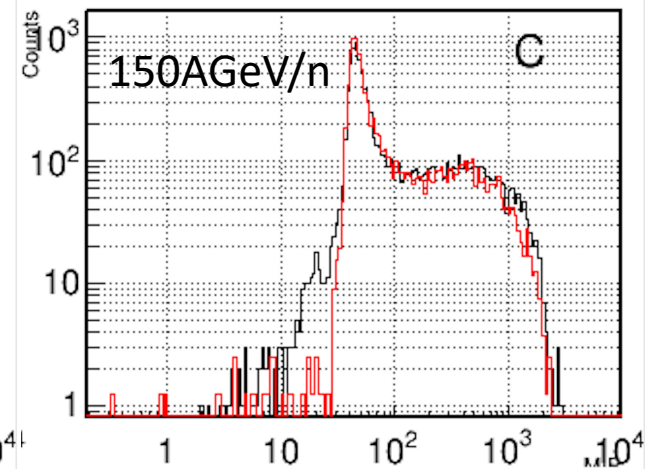
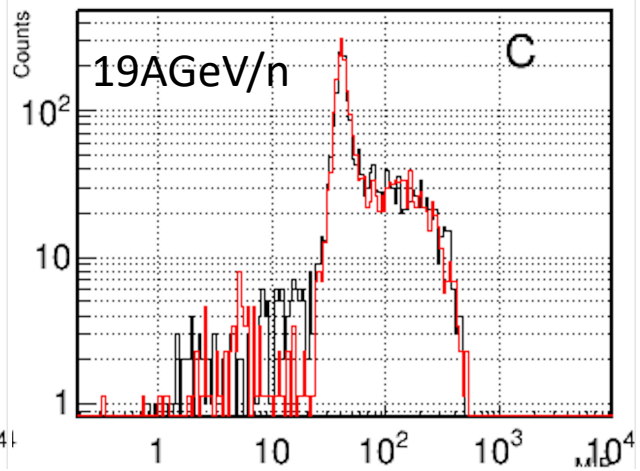
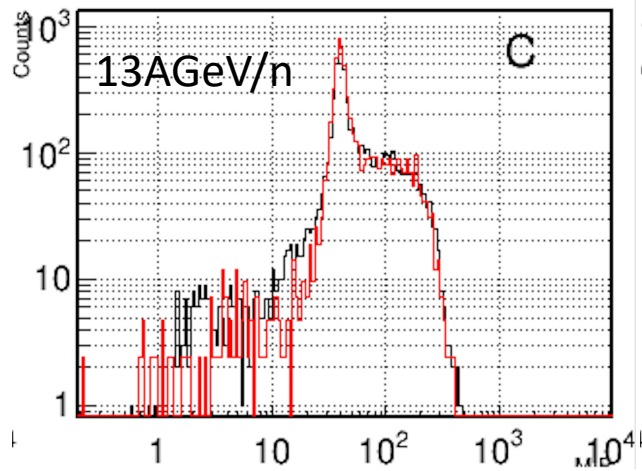


# Trigger efficiency with beam test

Data of 13, 19 GeV/n are required the charge consistency of IC-tracker and CHD

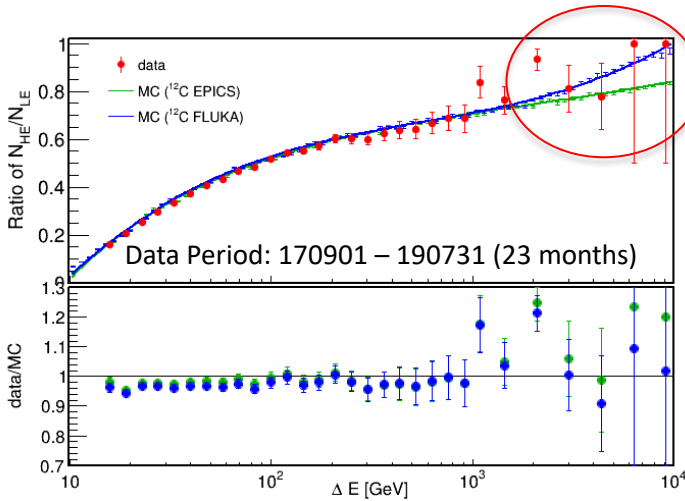
Data of 150 GeV/n is used only charge with CHD

➔ Contaminant events remain for 150 AGeV/n?

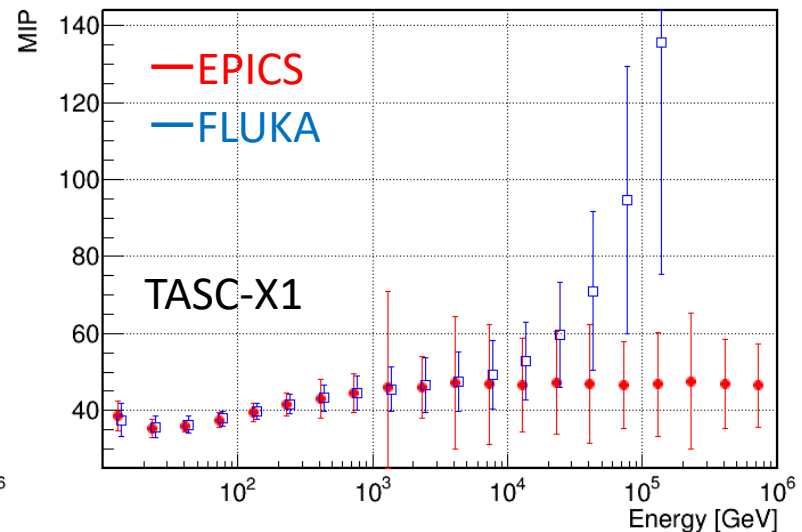
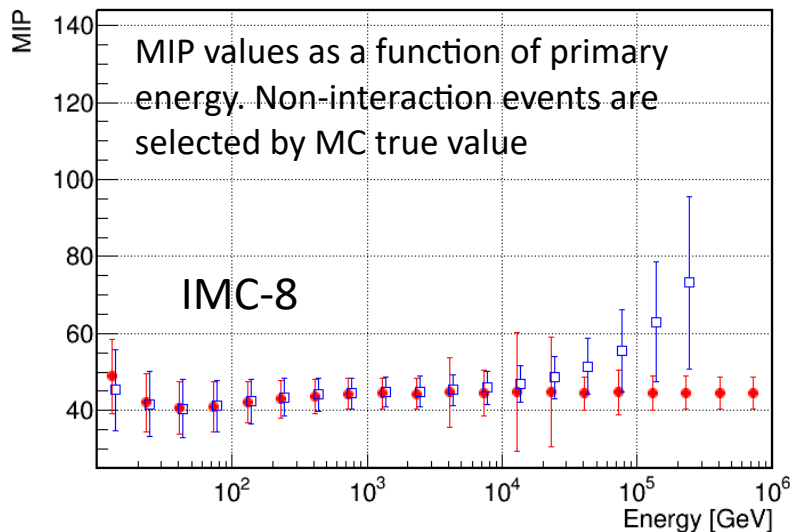
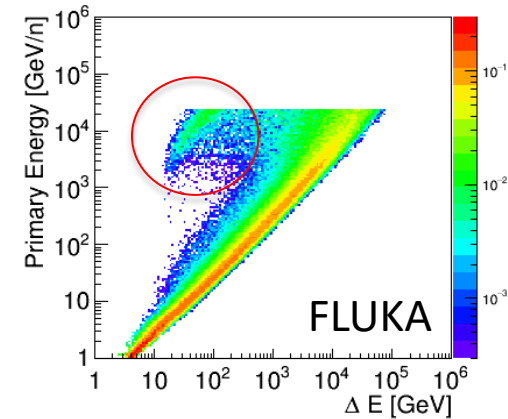
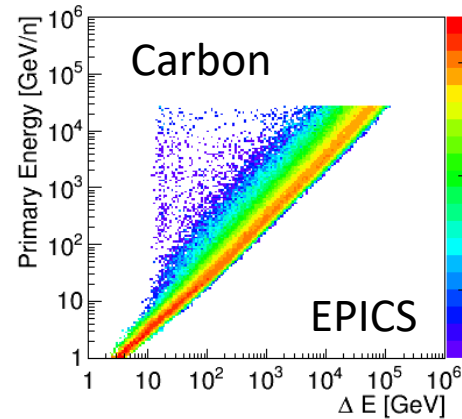


# Difference of EPICS and FLUKA

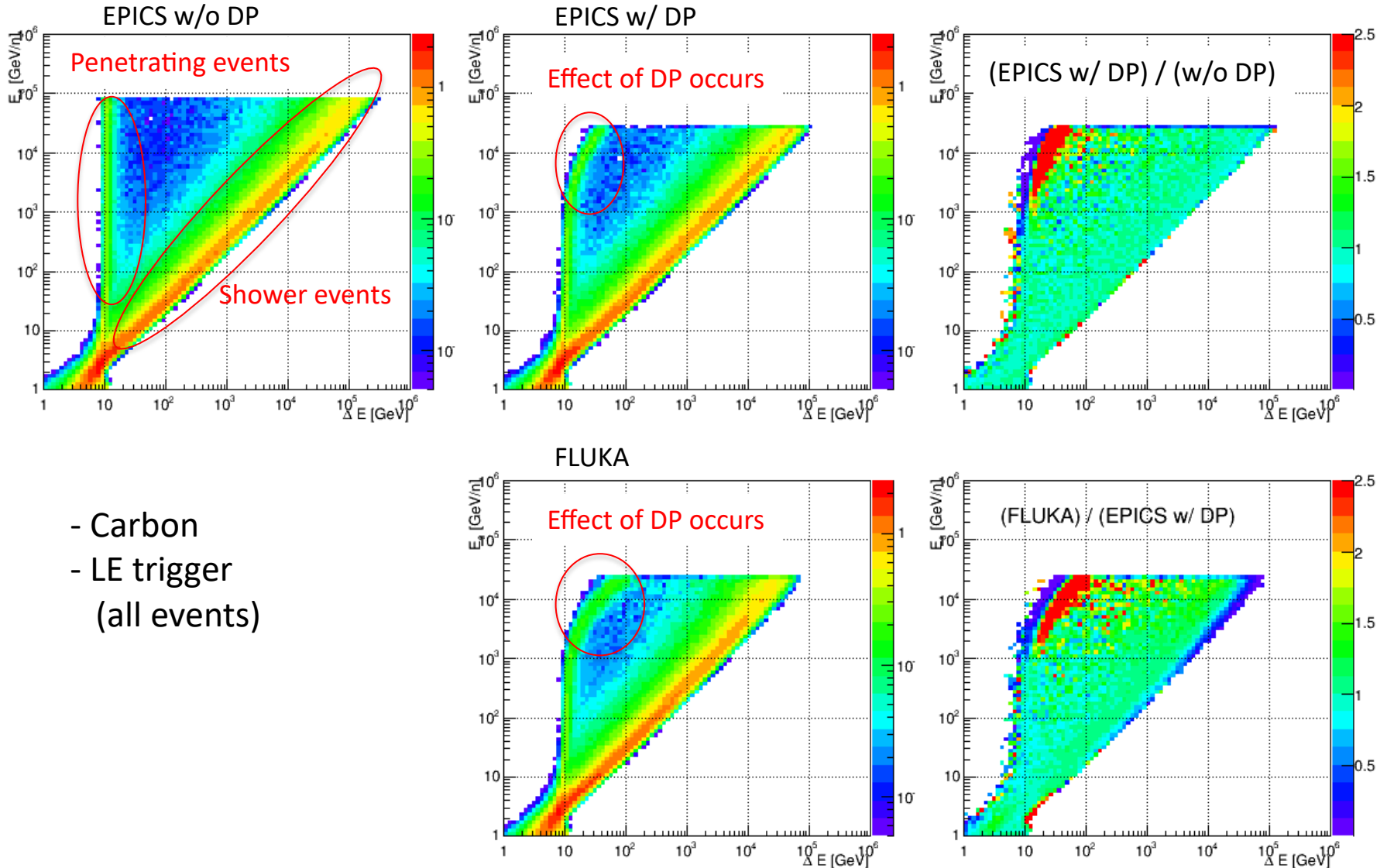
Efficiency of Fluka increase?



Events do not make large shower, but are triggered.

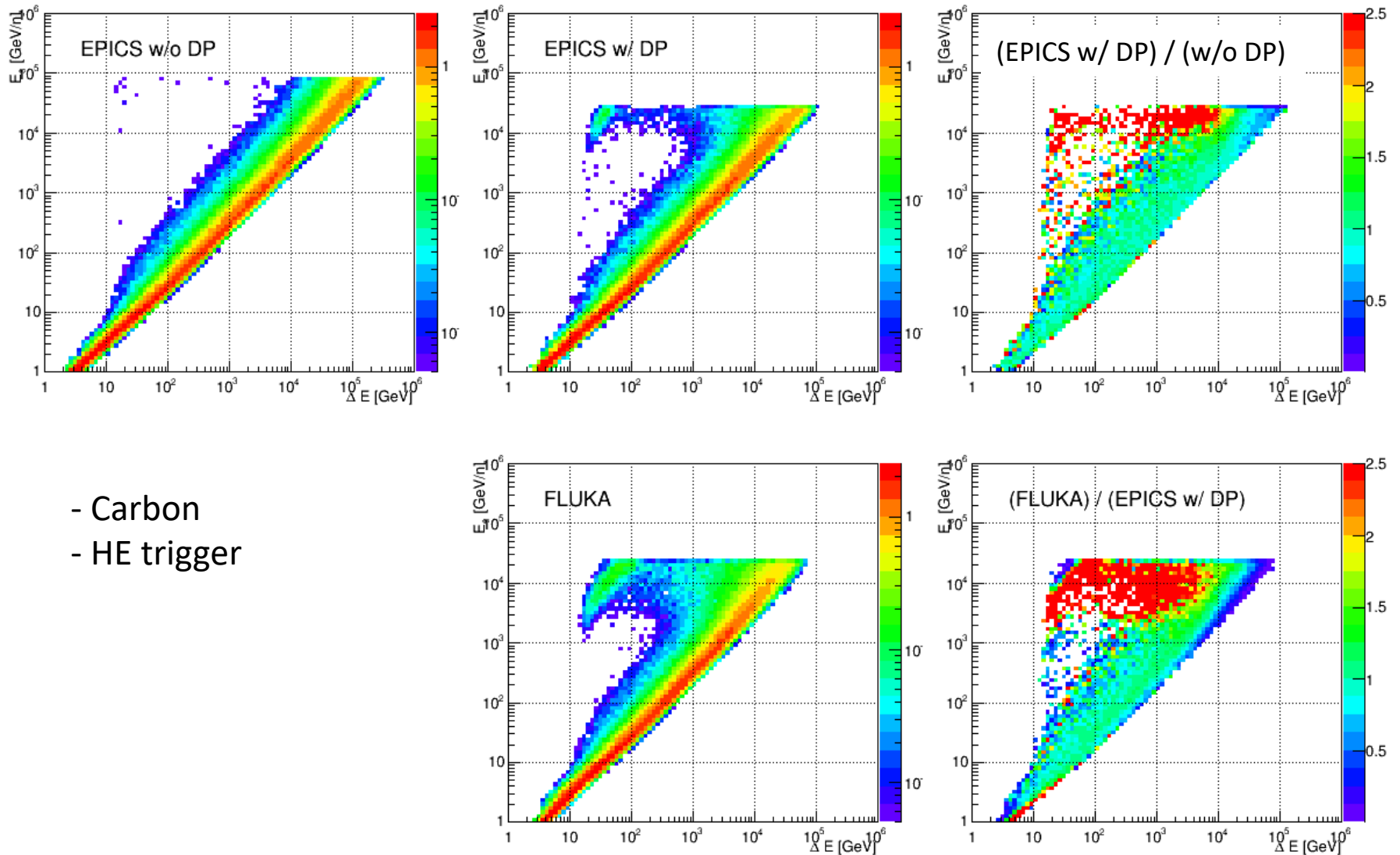


# Primary energy vs observed energy in TASC



- Carbon
- LE trigger (all events)

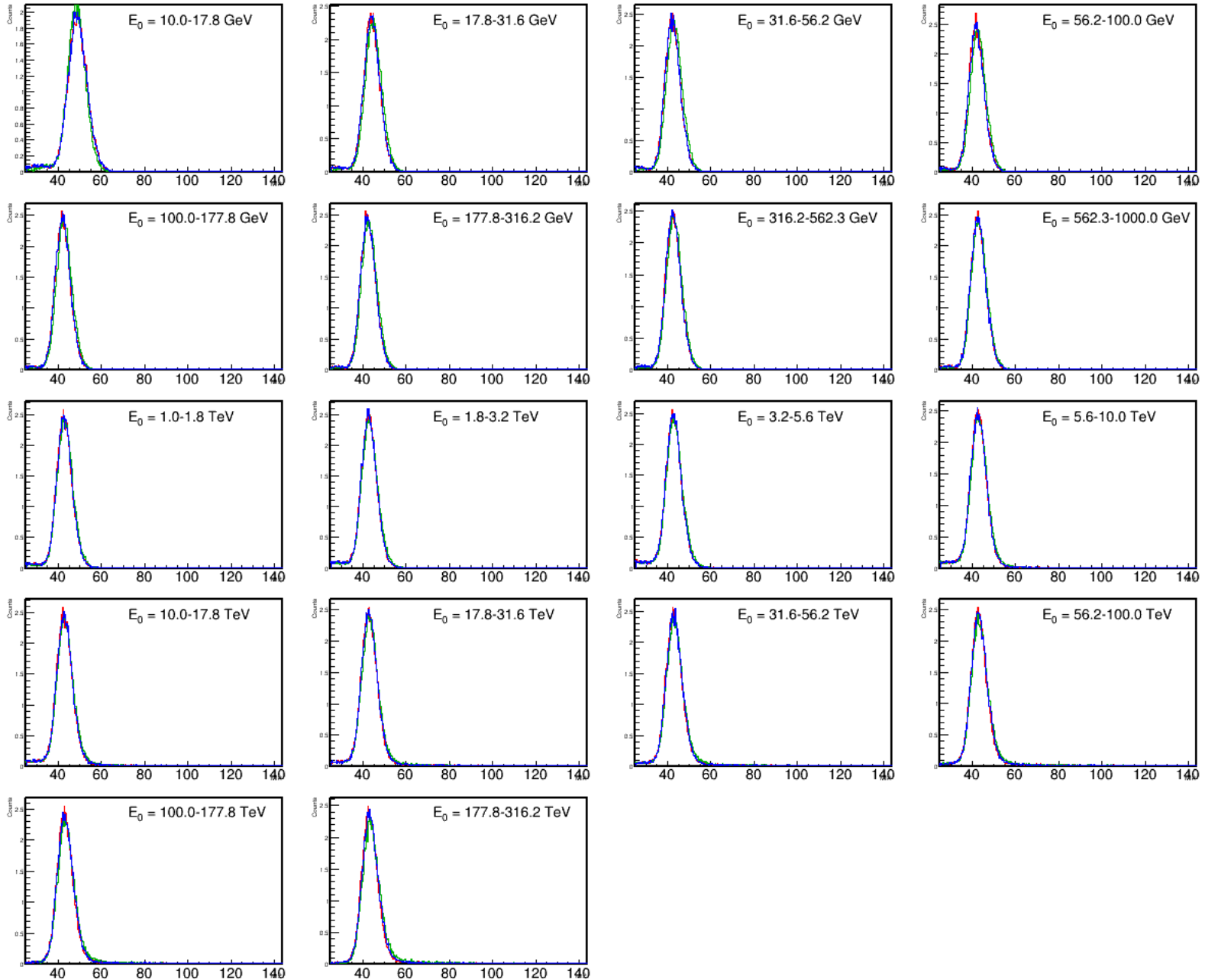
# Primary energy vs observed energy in TASC



- Carbon
- HE trigger

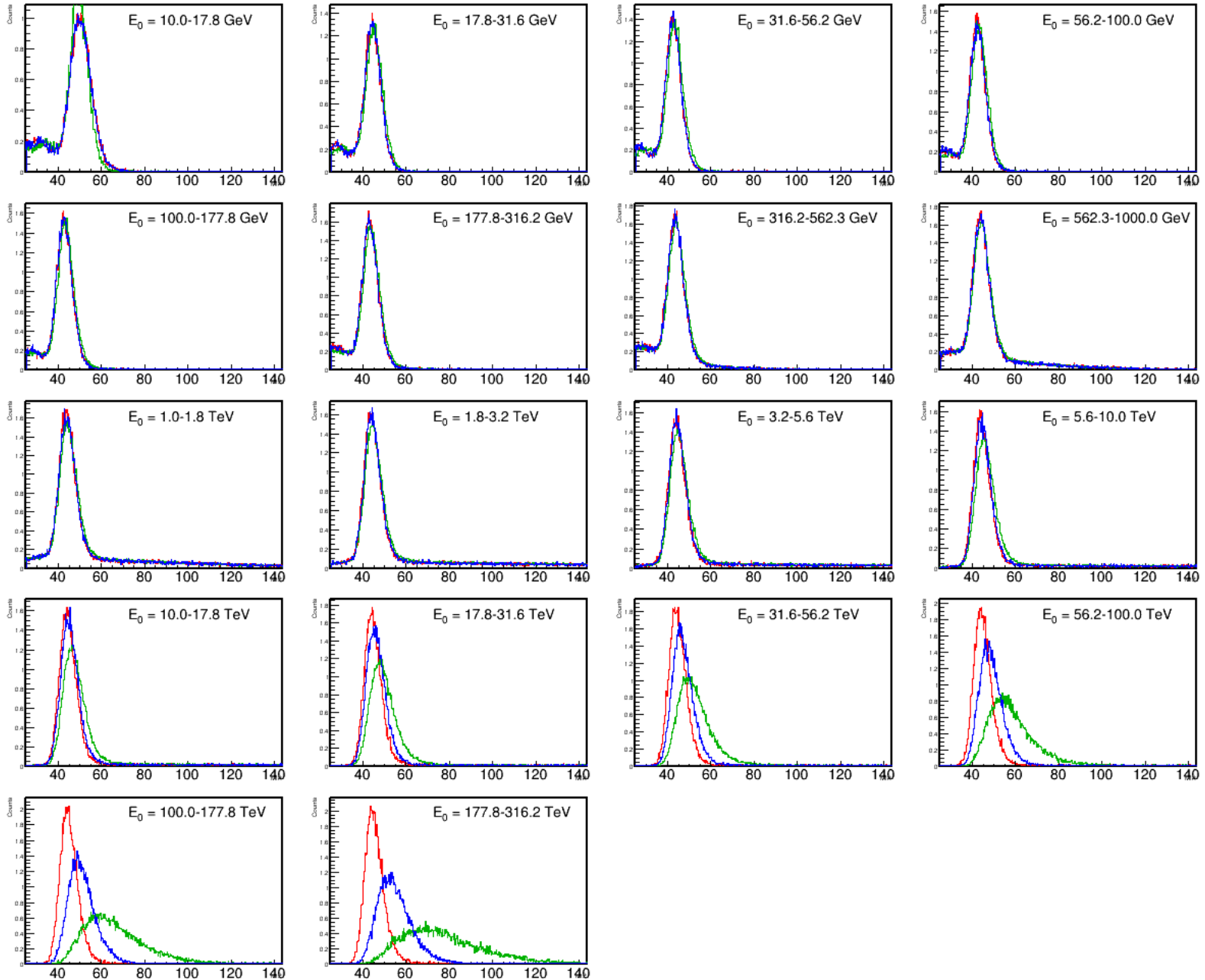
# IMC 1<sup>st</sup> layer all carbon events

— EPICS w/o DP — EPICS w/ DP — FLUKA



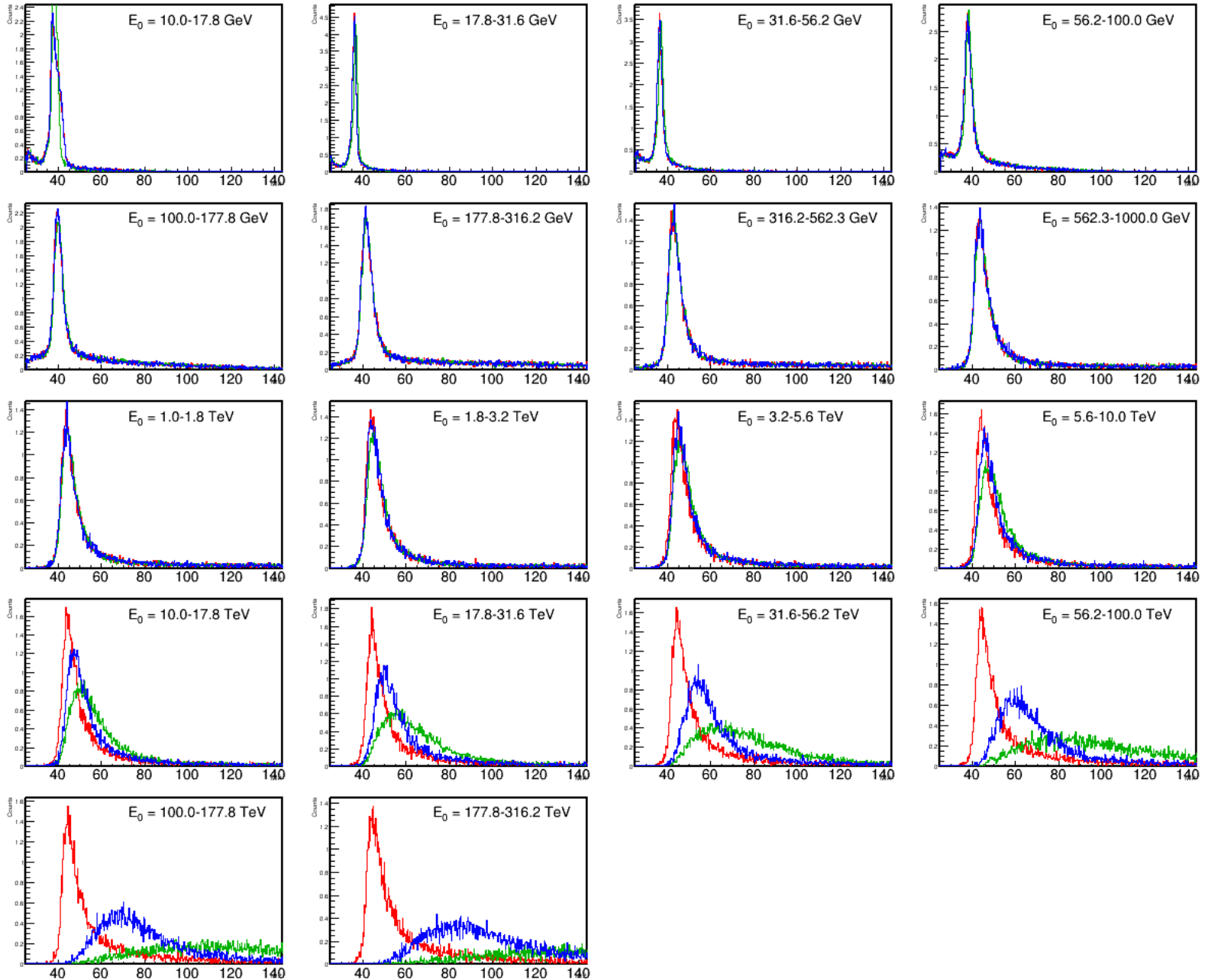
# IMC 8<sup>th</sup> layer all carbon events

— EPICS w/o DP — EPICS w/ DP — FLUKA

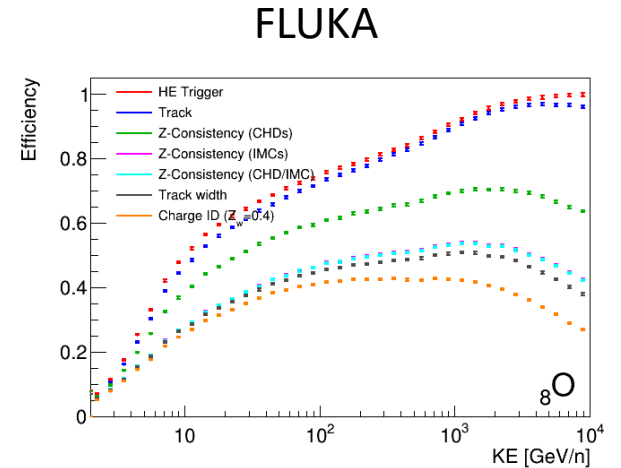
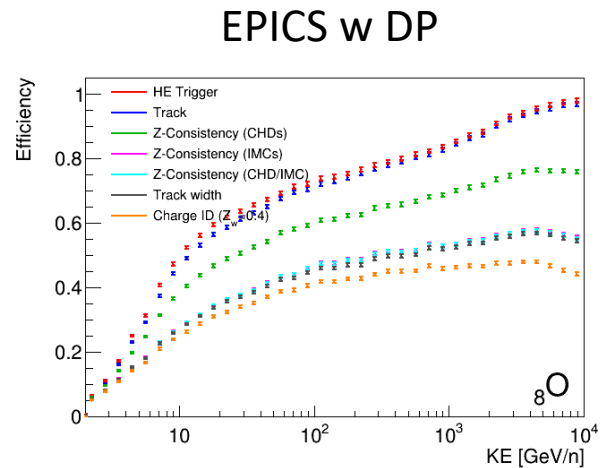
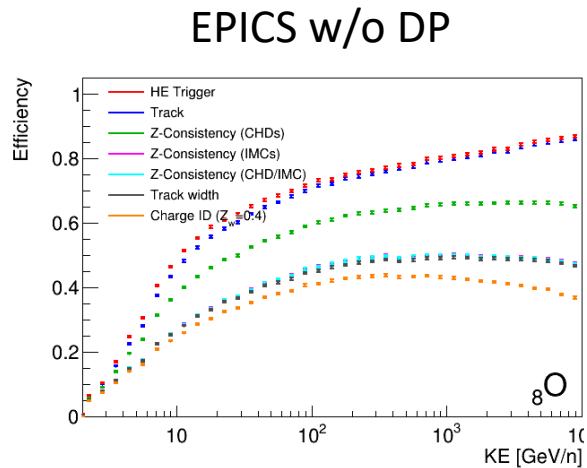
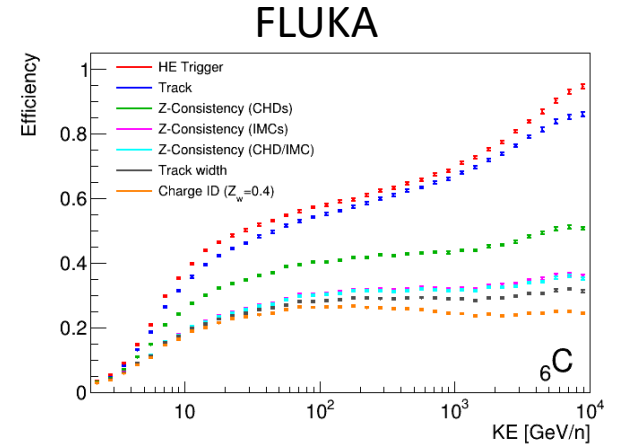
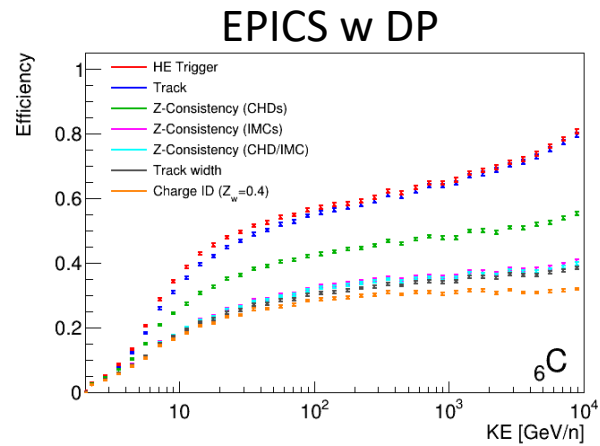
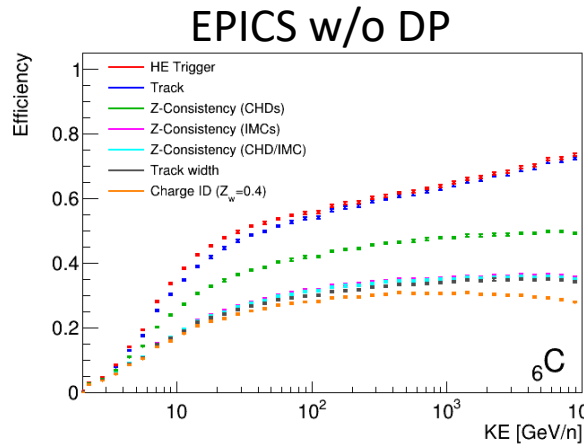


# TASC X1 all carbon events

— EPICS w/o DP — EPICS w/ DP — FLUKA

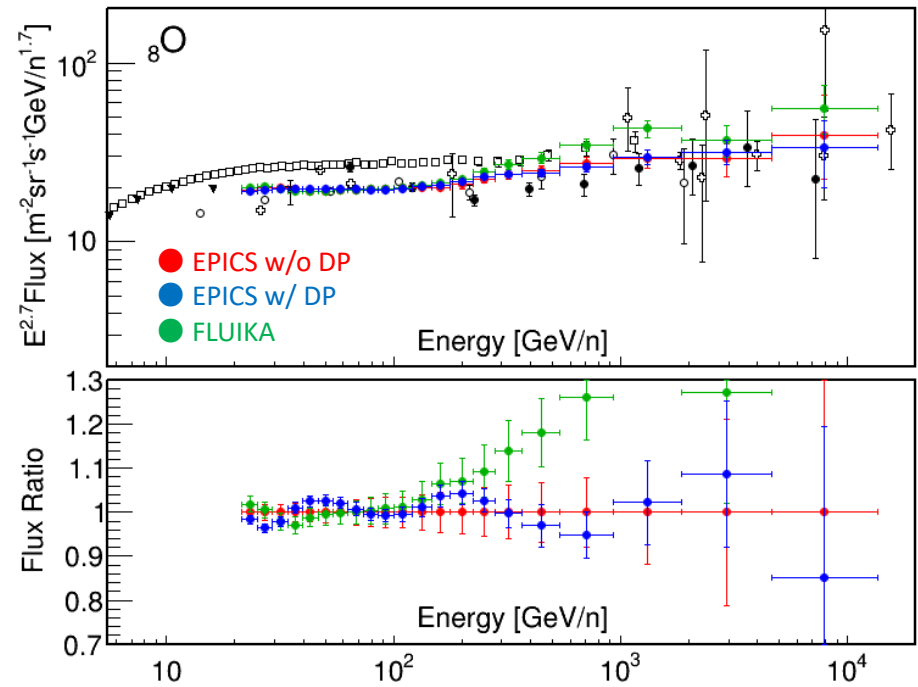
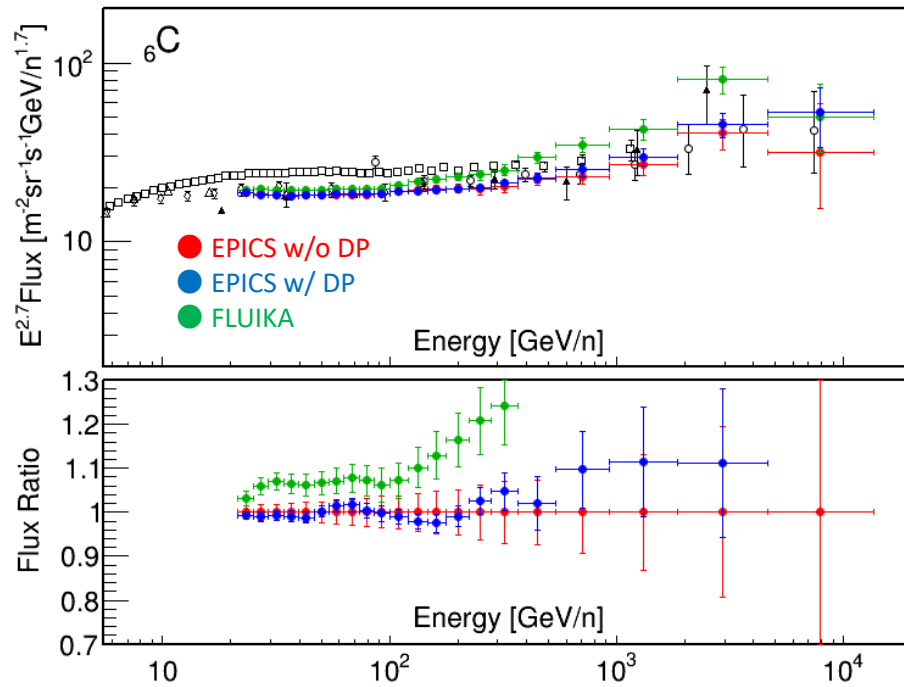


# Efficiencies





# Difference of Carbon and Oxygen spectra



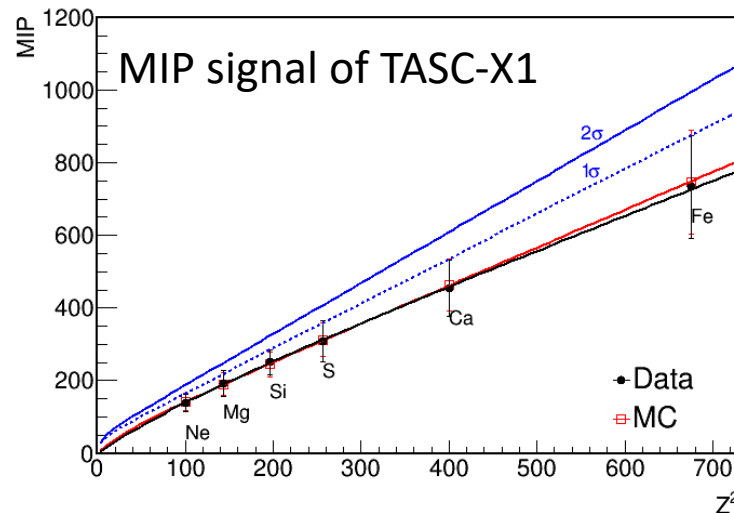
# Analysis of Heavy Nuclei Spectra ( $Z > 8$ )

- Difference of heavy nuclei analysis from light nuclei analysis
  - Charge consistency cut with IMC does not applied because IMC signal is saturated ( $Z > \sim 12$ )  
(track width selection is used because it does not related with the saturation)
  - Shower event selection with TASC is applied because the HE trigger efficiency is  $\sim 100\%$  due to the large  $dE/dx$

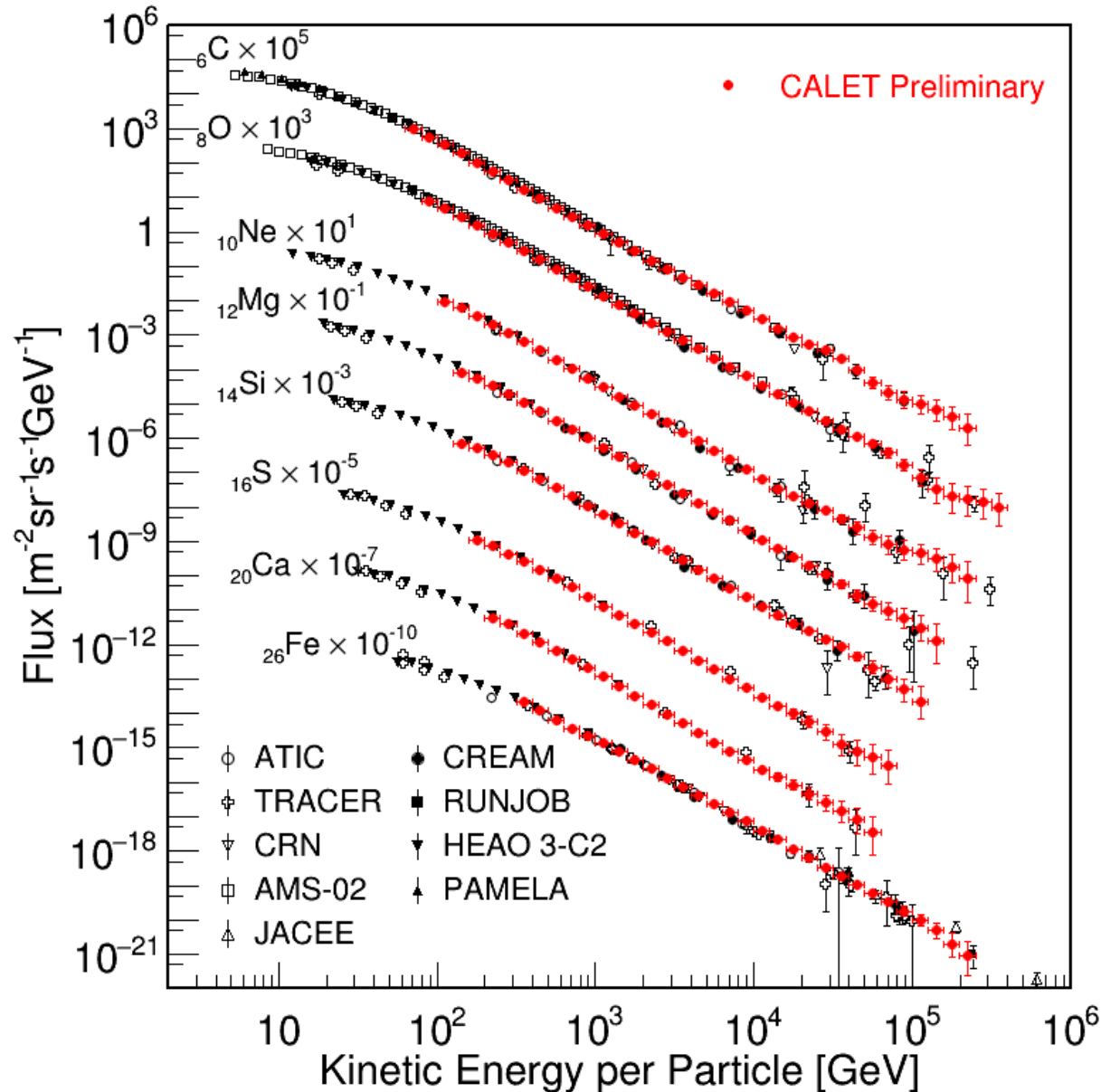
Shower event selection:

- $N_{TASCX1} > Nt(Z)$
- or  $N_{TASCY1} > Nt(Z)$
- or  $N_{TASCX2} > Nt(Z)$
- or  $N_{TASCY2} > Nt(Z)$

$Nt(Z)$  is  $2\sigma$  from the MIP peak of  $Z$

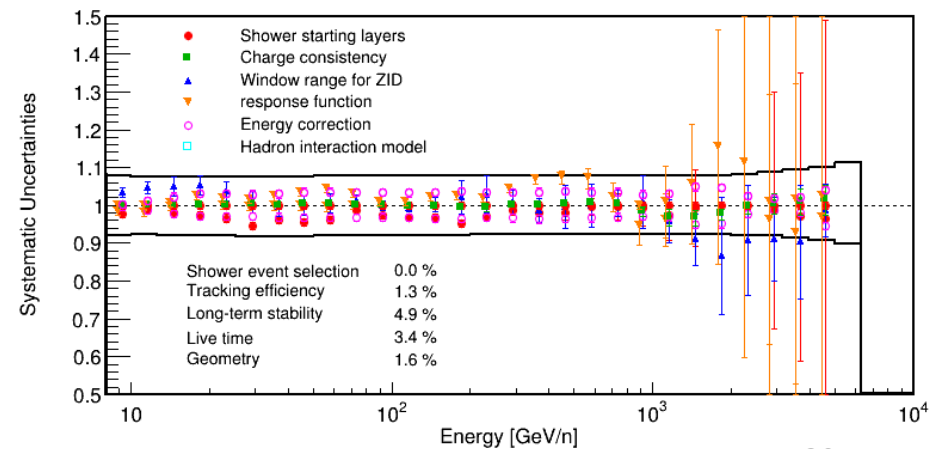
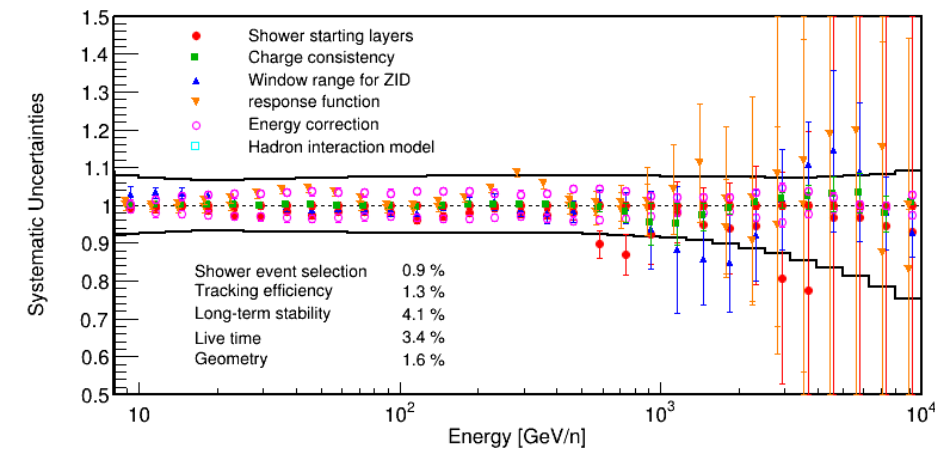
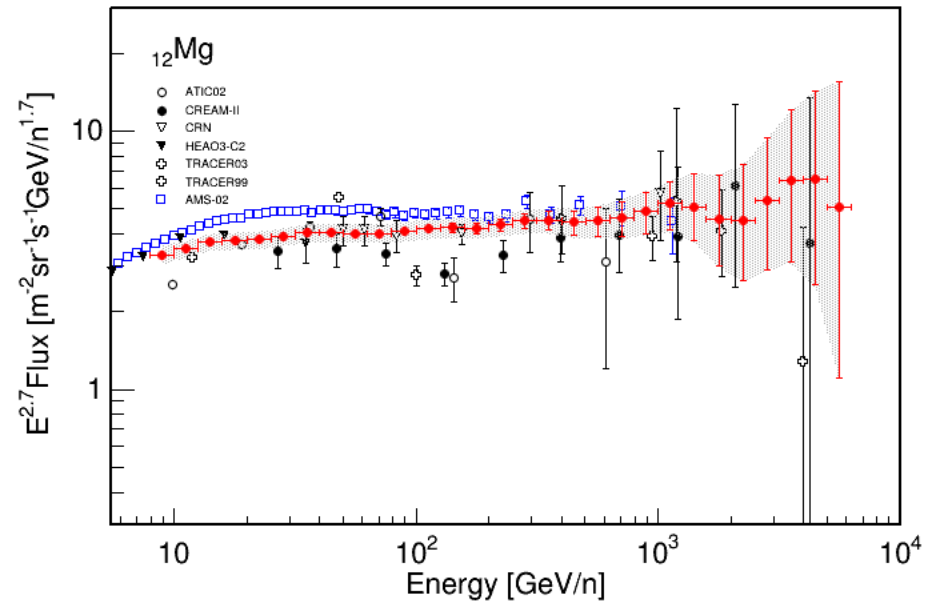
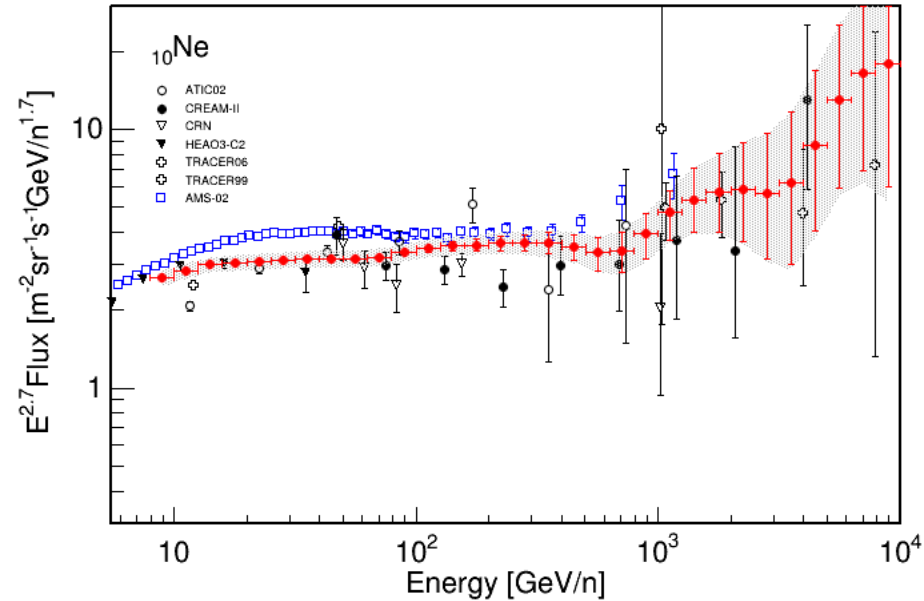


# Heavy nuclei spectra

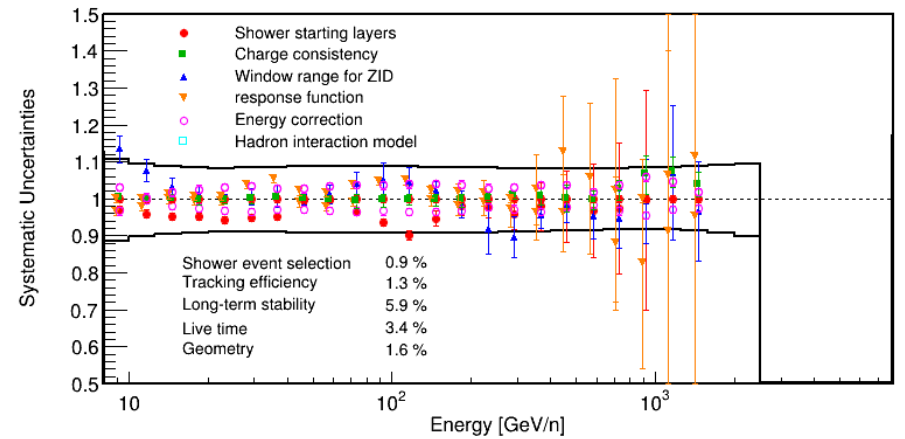
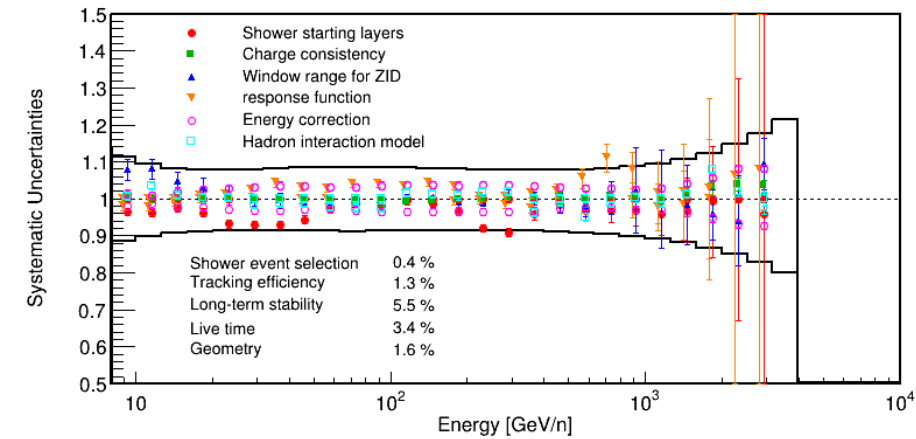
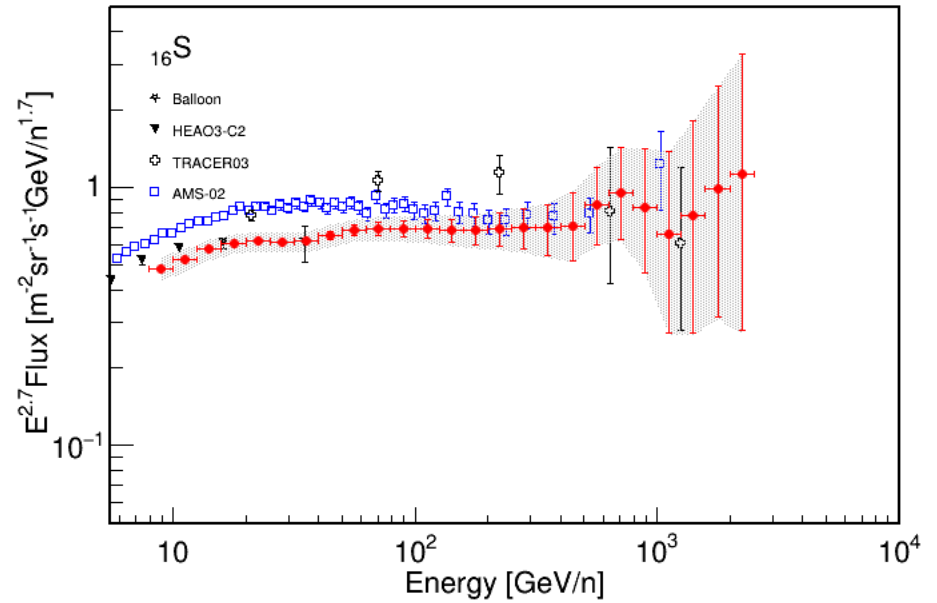
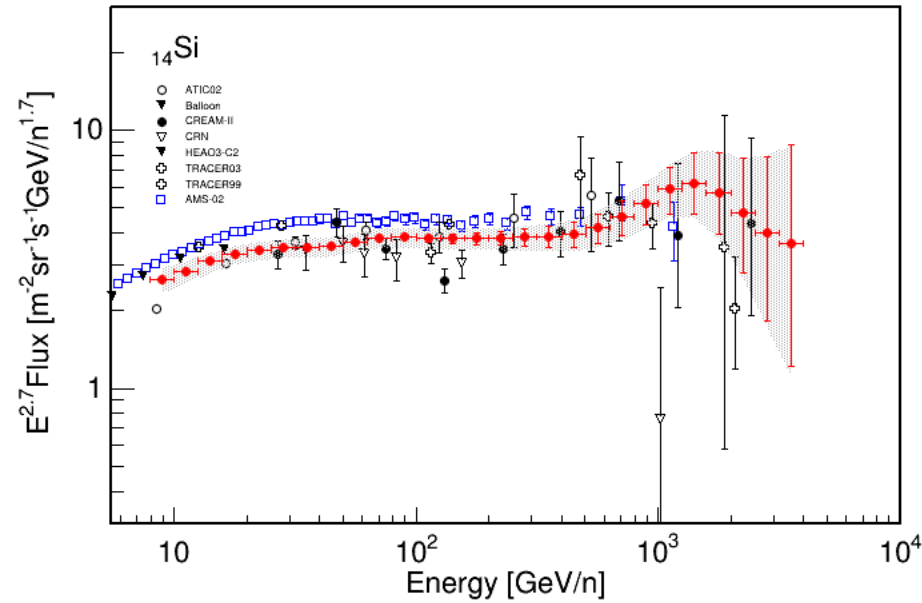


Period: 151013 – 190731  
1389 days

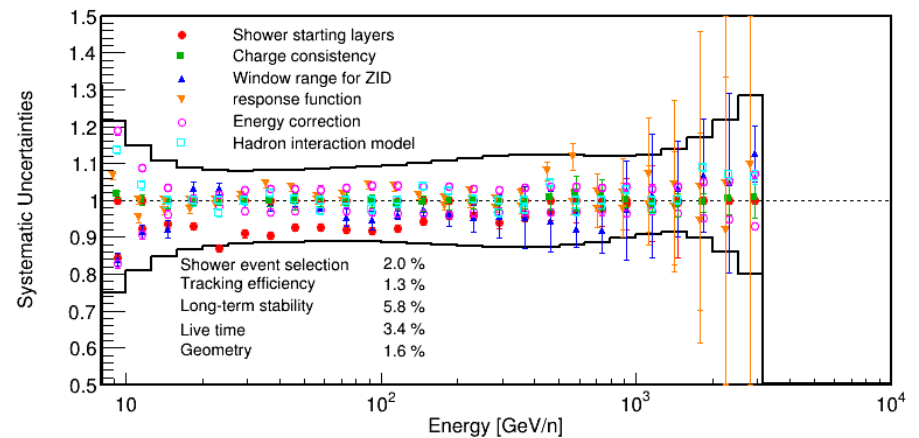
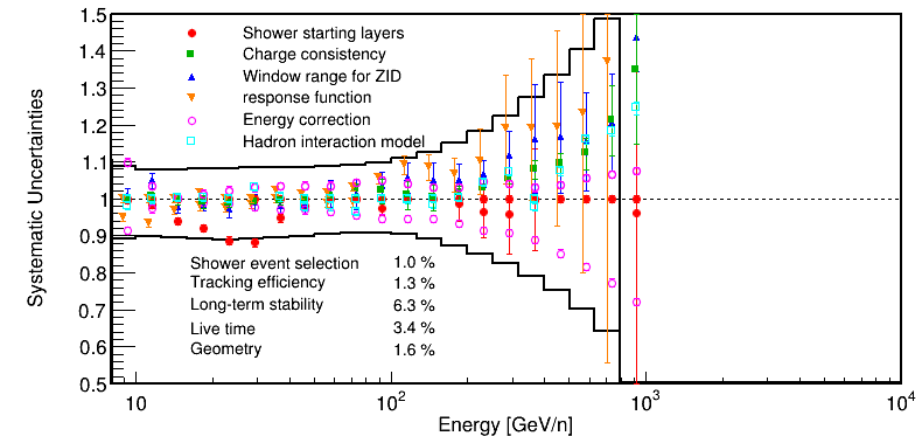
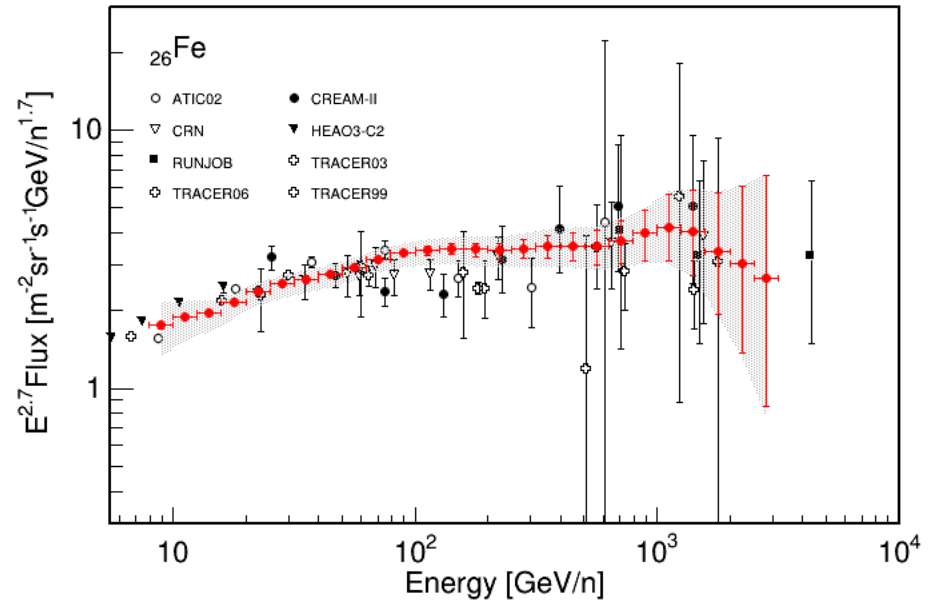
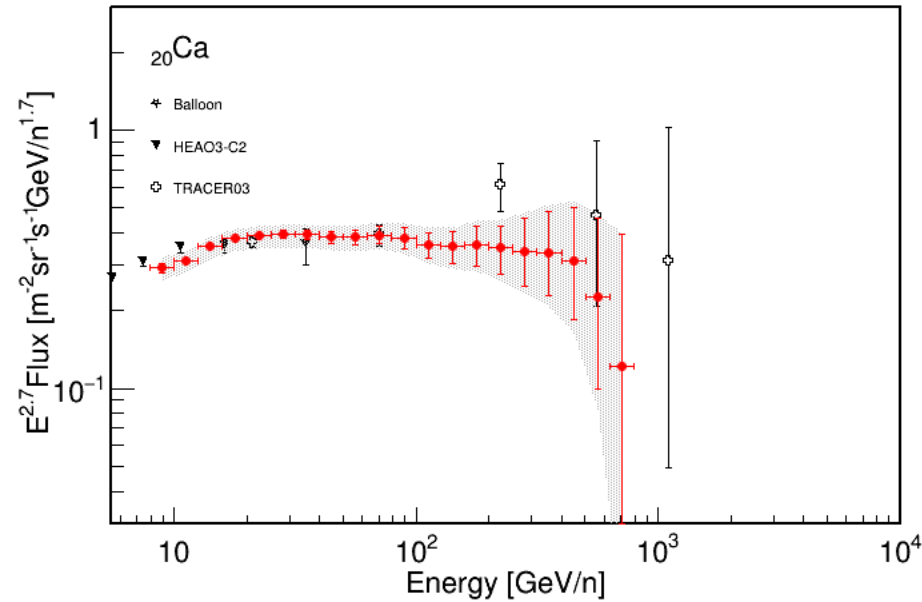
# Neon and Magnesium spectra



# Silicon and Sulfur spectra



# Calcium and Iron spectra



# Conclusion

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- B/C ratio is obtained from 1389 days of operation from 20 GeV/n to 5 TeV/n
  - Spectral index will be studied
- Carbon and Oxygen spectra are obtained
  - FLUKA spectrum is harder than EPICS -> check efficiency
- Difference from AMS-02 for nuclei spectra:
  - Although 12% energy correction or 25% efficiency correction are required, they are unlikely
- Nuclei spectra are obtained over 100TeV